



## **Squeezed Vacuum States of Light**

Exploring Numerical Simulations Under Multiple Conditions

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## Background

#### Introduction: What is Squeezed Light?

$$\Delta x \Delta p \geq \frac{\hbar}{2} \qquad \text{The Uncertainty Principle}$$
 
$$\vec{E}_{\mathbf{k},\lambda}(\vec{r},t) = \vec{e}_{\mathbf{k},\lambda}(p\cos{(\mathbf{k}\cdot\vec{r}-ckt)} + q\sin{(\mathbf{k}\cdot\vec{r}-ckt)}) \quad \text{TEM field}$$
 
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We want to modulate  $\Delta p$  and  $\Delta q$  while preserving the uncertainty

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$$\vec{A} = \vec{\alpha} e^{i(\mathbf{k} \cdot \vec{r} - \omega_{\mathbf{k}} t)} \Rightarrow \omega_{\mathbf{k}}^2 = c^2 k^2 \wedge \vec{\alpha} \cdot \mathbf{k} = 0 \qquad \text{Harmonic Oscillator}$$
 
$$\vec{A}_{\mathbf{k},\lambda} = \vec{e}_{\mathbf{k},\lambda} \text{Re}\{\alpha_{\mathbf{k},\lambda} e^{i(\mathbf{k} \cdot \vec{r} - ckt)}\} (\mathbf{k} = 2\pi (m,n,l)/L, \lambda = 1,2) \quad \text{Per-mode}$$
 
$$\hat{\alpha}_{\mathbf{k},\lambda} = (2\hbar/\varepsilon_0 L^3 ck)^{1/2} \hat{a}_{\mathbf{k},\lambda} \qquad \text{Annihilation operator for each mode}$$

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$$W(x,p) = \frac{1}{\pi\hbar} \int_{-\infty}^{\infty} \psi^*(x+y)\psi(x-y)e^{2ipy/\hbar}dy$$
 Wigner function

$$\frac{1}{\pi\hbar}\int_{-\infty}^{\infty}e^{-i\alpha(x+y)^2}\psi^*(x+y)\psi(x-y)e^{i\alpha(x-y)^2}e^{2ipy/\hbar}dy=W(x,p-2\hbar\alpha x)$$

Intuitive calculation: A linear map with determinant 1

#### Background: SHG, Lindblad Equation

Higher-order polarization effect generates multiple frequencies

$$\hat{H} = \hbar \omega \hat{b}^{\dagger} \hat{b} + \hbar \omega_P \hat{a}^{\dagger} \hat{a} + i \hbar \chi (\hat{b}^2 \hat{a}^{\dagger} - (\hat{b}^{\dagger})^2 \hat{a})$$

Second-order interaction: One  $\hat{a} \leftrightarrow \mathsf{Two} \; \hat{b}$ 

Approximate:  $\hat{a}$  very strong  $\Rightarrow \hbar \omega \hat{b}^{\dagger} \hat{b} + \text{Im}\{ce^{i\omega_p t} \hat{b}^2\}$ 

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Interaction with the environment: Lindblad Equation

$$\partial_t \rho = -\frac{i}{\hbar} [H_{sys}, \rho] + \frac{1}{2} \sum_n (2C_n \rho C_n^{\dagger} - \rho C_n^{\dagger} C_n - C_n^{\dagger} C_n \rho)$$

Methodology

#### Methodology

#### How to solve

- Projection onto finite dimensional Hilbert space spanned by  $\{|i\rangle\}_{0 \le i \le n}$
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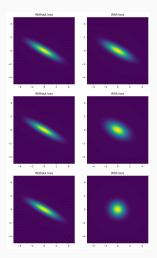
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#### What to solve

- No interaction with environment
- Interaction with  $C = \sqrt{\gamma} \hat{b}$  (Thermal loss)
- Higher order generation

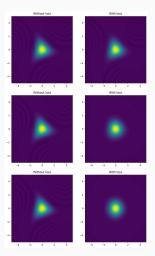
## **Results**

#### Comparison between lossy and non-lossy probabilities



- Adjust  $\sqrt{\gamma}$  in  $C = \sqrt{\gamma}\hat{b}$  of the Lindblad equation.  $\sqrt{\gamma} = 0.2, 0.5, 1.0$ , and the plots are taken after 5 seconds of time evolution.
- Since the environment absorbs the photons, the nonlinear effects are diminished.

#### Third-order generation?



■ The same plots, with  $H_{svs}$ containing  $\hat{b}^3$ 

Methodology

■ Expected the Wigner function to turn into transform into the form  $W(2^{nd}, 2^{nd})$ 

#### **Summary**

- Studied: what are squeezed states, how to generate them
- Tried out: the QuTiP library to simulate interaction with the environment and higher order generation

#### References

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