



Squeezed Vacuum States of Light

Exploring Numerical Simulations Under Multiple Conditions

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Background

Introduction: What is Squeezed Light?

$$\Delta x \Delta p \geq \frac{\hbar}{2} \qquad \text{The Uncertainty Principle}$$

$$\vec{E}_{\mathbf{k},\lambda}(\vec{r},t) = \vec{e}_{\mathbf{k},\lambda}(p\cos{(\mathbf{k}\cdot\vec{r}-ckt)} + q\sin{(\mathbf{k}\cdot\vec{r}-ckt)}) \quad \text{TEM field}$$

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We want to modulate ΔA and $\Delta \phi$ while preserving the uncertainty

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$$\vec{A} = \vec{\alpha} e^{i(\mathbf{k} \cdot \vec{r} - \omega_{\mathbf{k}} t)} \Rightarrow \omega_{\mathbf{k}}^2 = c^2 k^2 \wedge \vec{\alpha} \cdot \mathbf{k} = 0 \qquad \text{Harmonic Oscillator}$$

$$\vec{A}_{\mathbf{k},\lambda} = \vec{e}_{\mathbf{k},\lambda} \text{Re}\{\alpha_{\mathbf{k},\lambda} e^{i(\mathbf{k} \cdot \vec{r} - ckt)}\} (\mathbf{k} = 2\pi (m,n,l)/L, \lambda = 1,2) \quad \text{Per-mode}$$

$$\hat{\alpha}_{\mathbf{k},\lambda} = (2\hbar/\varepsilon_0 L^3 ck)^{1/2} \hat{a}_{\mathbf{k},\lambda} \qquad \text{Annihilation operator for each mode}$$

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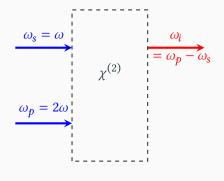
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$$W(x,p) = \frac{1}{\pi\hbar} \int_{-\infty}^{\infty} \psi^*(x+y)\psi(x-y)e^{2ipy/\hbar}dy$$
 Wigner function

$$\frac{1}{\pi\hbar}\int_{-\infty}^{\infty}e^{-i\alpha(x+y)^2}\psi^*(x+y)\psi(x-y)e^{i\alpha(x-y)^2}e^{2ipy/\hbar}dy=W(x,p-2\hbar\alpha x)$$

Intuitive calculation: A linear map with determinant 1

Background: Parametric amplification



$$\hat{H} = \hbar \omega \hat{b}^{\dagger} \hat{b} + \hbar \omega_{P} \hat{a}^{\dagger} \hat{a} + i \hbar \chi (\hat{b}^{2} \hat{a}^{\dagger} - (\hat{b}^{\dagger})^{2} \hat{a})$$
second-order interaction

One pump $\hat{a} \leftrightarrow \mathsf{Two} \; \hat{b}$

Approximate: \hat{a} very strong $\Rightarrow \hbar \omega \hat{b}^{\dagger} \hat{b} + \text{Im}\{ce^{i\omega_p t} \hat{b}^2\}$

Higher-order polarization effect generates multiple frequencies

Background: Lindblad master equation for simulation

Interaction with the environment: Lindblad Equation

$$\partial_t \rho = -\frac{i}{\hbar} [H_{sys}, \rho] + \frac{1}{2} \sum_n (2C_n \rho C_n^{\dagger} - \rho C_n^{\dagger} C_n - C_n^{\dagger} C_n \rho)$$

Methodology

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How to solve

- Projection onto finite dimemsional Hilbert space spanned by $\{|i\rangle\}_{0 \le i \le n}$
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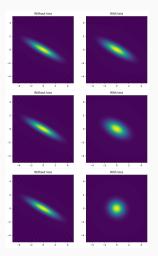
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What to solve

- No interaction with environment
- Interaction with $C = \sqrt{\gamma} \hat{b}$ (Thermal loss)
- Higher order generation

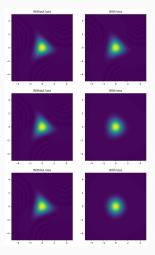
Results

Comparison between lossy and non-lossy probabilities



- Adjust $\sqrt{\gamma}$ in $C = \sqrt{\gamma}\hat{b}$ of the Lindblad equation. $\sqrt{\gamma} = 0.2, 0.5, 1.0$, and the plots are taken after 5 seconds of time evolution.
- Since the environment absorbs the photons, the nonlinear effects are diminished.

Third-order generation?



- The same plots, with H_{sys} containing \hat{b}^3
- Expected the Wigner function to turn into transform into the form $W(2^{nd}, 2^{nd})$

Summary

- Studied: what are squeezed states, how to generate them
- Tried out: the QuTiP library to simulate interaction with the environment and higher order generation
- All results + this slide: available on GitHub

References

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