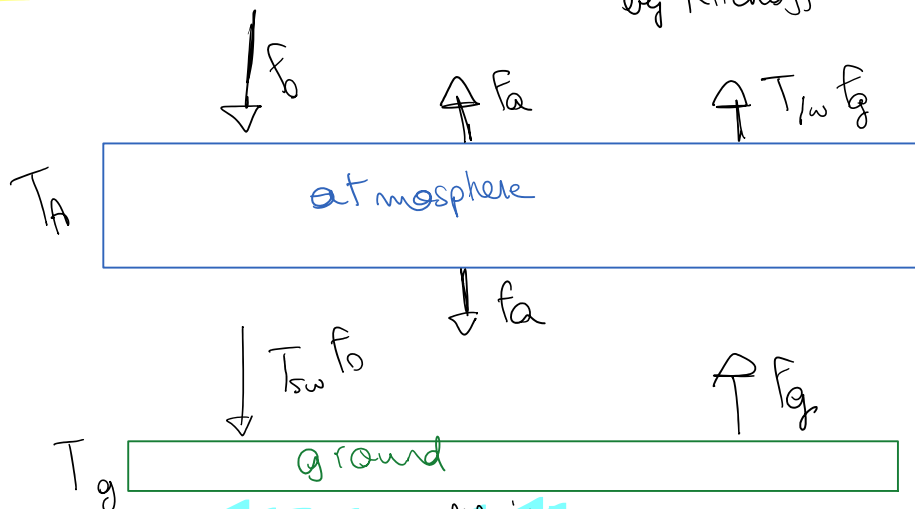


- Chapter 1: model with non-absorbing atmosphere
- total solar irradiance in a tube πa^2 \rightarrow Earth $F_s = 1370 \text{ W m}^{-2}$, contained earth has a planetary albedo $A = 0.3$
- 30% of incoming radiation reflected back to space. \Rightarrow reflects $0.3 F_s \pi a^2$ of incoming solar power.
- earth is assumed to emit as a black body Stefan law: $P = \sigma T^4$ \Rightarrow assume all transmitted to space & none absorbed by atmosphere. \Rightarrow stef-boltz const.
 - Assume earth is in thermal equilibrium $P_{in} = P_{out}$ $(1-A) F_s \pi a^2 = 4\pi a^2 \sigma T^4 \Rightarrow A = 0.3 \Rightarrow T = 255 \text{ K}$
 effective temperature of the earth.
 Real value = 288K

- 1.3.2: adding an atmosphere, temperature
- add one layer of atmosphere T_a . it transmits fraction T_{sw} incident short waves & T_{lw} incident long waves.
- these are called transmittances. \rightarrow low transmittance high absorbance means
- Assume ground at temperature T_g . \rightarrow low transmittance high absorbance means
- \Rightarrow mean unreflected incoming solar irradiance at the top of the atmosphere is $F_0 = \left(\frac{1}{4}\right)(1-A) F_s \Rightarrow 240 \text{ W m}^{-2}$ due to ratio of surfaces.
- of this $T_{sw} F_0$ is absorbed by the ground & $(1-T_{sw}) F_0$ is absorbed by the atmosphere
- ground emits as a black body with irradiance $F_g = \sigma T_g^4$
- $T_{lw} F_g$ reaches atmosphere $(1-T_{lw}) F_g$ is absorbed.
- atmosphere has irradiance $F_a = (1-T_{lw}) \sigma T_g^4$ both up & down. by Kirchhoff's law of emittance



assume the system is in radiative equilibrium. neglect energy transfers due to non-radiative processes.

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$$F_0 = F_A + T_{ew} F_g \quad F_g = T_{sw} F_0 + F_A$$

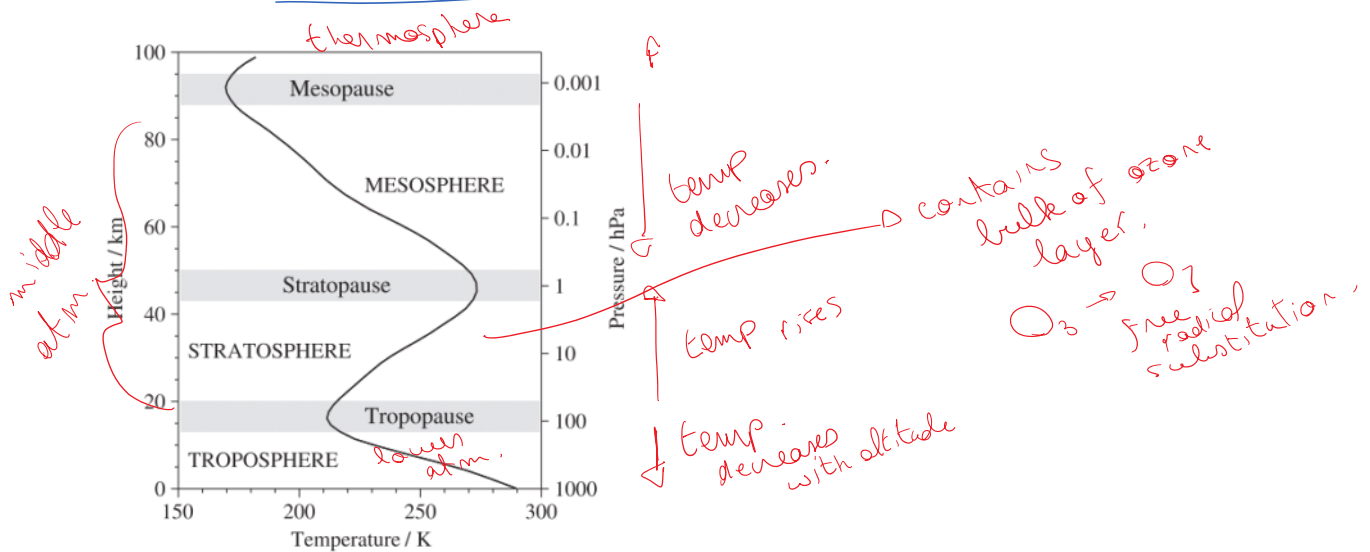
$$\Rightarrow F_g = \sigma T_g^4 = F_0 \frac{1 + T_{sw}}{1 + T_{lw}} \quad \text{rough values for earth's atmosphere}$$

gives $T_g \approx 288 \text{ K}$ which is close to observed 288 K

$T_{sw} = 0.9 \quad T_{lw} = 0.2$

$$F_A = F_0 \frac{1 - T_{sw} T_{lw}}{1 + T_{lw}} \Rightarrow T_a = 295 \text{ K}$$

Greenhouse effect: there is less absorption for solar than thermal radiation.
Quantify the greenhouse effect of a gas: $F_g - F_0$ by which it reduces the outgoing irradiance from its surface value
is emitted upwards & the rest towards the earth.
 $(1 - T_{lw}) \sigma T_g^4$ is absorbed by atmosphere, whereas $(1 - T_{sw}) \sigma T_a^4$



the ozone absorbs solar UV radiation.