Fully General Relativistic Structure and Spectrum of ADAFs

Li Yan-Rong

Supported by

Prof. Wang Jian-Min

Prof. Yuan Ye-Fei

December 10, 2008

Contents

1	Basic Usage								
2	The Structure of ADAFs								
	2.1	Equati	ion set for ADAFs		7				
	2.2	The bo	oundary conditions	. 12	2				
	2.3	Dimen	nsion units	. 12	2				
2.4 The transonic point		The tra	ansonic point	. 12	2				
	2.5	Compa	arison	. 13	3				
3	The	The Spectrum of ADAFs 17							
	3.1	Compt	tonization	. 17	7				
		3.1.1	Klein-Nishina formular	. 17	7				
		3.1.2	Scattering rate	. 18	8				
		3.1.3	Mean scattered photon energy	. 18	8				
		3.1.4	Dispersion about $\langle \omega \rangle$. 18	8				
		3.1.5	Scattered-photon distribution	. 19	9				
		3.1.6	Kinetic equation	. 19	9				
		3.1.7	Electron distribution	. 19	9				
	3.2	Radiat	tion mechanisms	. 20	0				

4 CONTENTS

		3.2.1	Synchrotron emission	20			
		3.2.2	Bremstralung emission	20			
		3.2.3	Compton scattering	21			
3.3 ADAF			model	21			
		3.3.1	Self-simular solution	21			
		3.3.2	Global structure	22			
		3.3.3	Radiation transfer	22			
4	TIL.		atara anno de la d	23			
4	The ray tracing method						
	4.1	Jacobi	an Elliptic Function	23			
	4.2	Integra	als	23			
		4.2.1	<i>r</i> -component	24			
		4.2.2	θ -component	26			
4.3		TeV T	rajectory	26			
		4.3.1	Calculating $ au_{\mu}$	27			
		4.3.2	Solving r_e Given $ au_{\mu}$	27			
		4.3.3	Solving β	28			
5	Mat	hematio	cal functions	29			
				20			
	5.1	Bessel	functions	29			

Chapter 1

Basic Usage

- /src: contains the source file for calculation of disk structure, intrinsic spectrum and observed spectrum;
- /data: contains the input/output data;
- /doc: the use manual.

To use the code, change directory to / src, type command make and three executive files are produced,

- disk: for disk structure;
- spec: for intrinsic spectrum;
- obs: for observed spectrum.

Then type ./disk in shell will calculate the disk structure and analogically for ./spec and ./obs.

In the directory /data, the files are

- datain.txt: the input data.
- /spec: the SED at different radius, in which the number in the file name (e.g. spec025.txt) correspond with the line number in file rdisk.dat.
- nrows.txt: the number of rows in rdisk.dat.
- soltot.dat: the solutions of all variables of ADAFs.
- adaf.dat: used for spec.
- sol-for-spec.dat: used for obs.
- rdisk.dat: used for calculation of optical depth.
- spectrum.dat specobs.dat: the intrinsic spectrum and the observed spectrum.

Chapter 2

The Structure of ADAFs

2.1 Equation set for ADAFs

We adopt the geometrical units G = c = 1. The kerr metric expanded around the equatorial plane is

$$ds^{2} = -\frac{r^{2}\Delta}{A}dt^{2} + \frac{A}{r^{2}}(d\phi - \omega dt)^{2} + \frac{r^{2}}{\Delta}dr^{2} + dz^{2}$$
(2.1)

$$= -\left(\frac{r^2\Delta}{A} - \frac{A\omega^2}{r^2}\right)dt^2 - \frac{2A\omega}{r^2}dtd\phi + \frac{A}{r^2}d\phi^2 + \frac{r^2}{\Delta}dr^2 + dz^2$$
 (2.2)

Therefore

$$g_{tt} = -\left(\frac{r^2\Delta}{A} - \frac{A\omega^2}{r^2}\right), \quad g_{t\phi} = -\frac{A\omega}{r^2}, \quad g_{\phi\phi} = \frac{A}{r^2}, \quad g_{rr} = \frac{r^2}{\Delta}, \quad g_{zz} = 1$$
 (2.3)

where

$$\Delta = r^2 - 2Mr + a^2$$
, $A = r^4 + r^2a^2 + 2Mra^2$, $\omega = \frac{2Mar}{A}$, $a = \frac{J}{M}$ (2.4)

The equation set are

• Continuity equation

$$\dot{M} = -2\pi \Delta^{1/2} \Sigma_0 \gamma_r V \tag{2.5}$$

Momentum equation

$$\frac{\dot{M}}{2\pi}(L - L_{in}) = rW_{\phi}^{r} \tag{2.6}$$

$$\gamma_r^2 V \frac{dV}{dr} + \frac{1}{\mu \Sigma_0} \frac{dW}{dr} = -\frac{\gamma_\phi^2 AM}{r^4 \Delta} \frac{(\Omega - \Omega_K^+)(\Omega - \Omega_K^-)}{\Omega_K^+ \Omega_K^-}.$$
 (2.7)

• Energy equation

$$\frac{\dot{M}W_i}{2\pi r \Sigma_0} \frac{1}{\Gamma_i - 1} \left(\frac{d \ln W_i}{dr} - \Gamma_i \frac{d \ln \Sigma_0}{dr} + \frac{\Gamma_i - 1}{r} \right) = (1 - \delta) \frac{\alpha W}{r} \frac{\gamma_\phi^4 A^2}{r^6} \frac{d\Omega}{dr} + \Lambda_{ie}$$
 (2.8)

$$\frac{\dot{M}W_e}{2\pi r \Sigma_0} \frac{1}{\Gamma_e - 1} \left(\frac{d \ln W_e}{dr} - \Gamma_e \frac{d \ln \Sigma_0}{dr} + \frac{\Gamma_e - 1}{r} \right) = \delta \frac{\alpha W}{r} \frac{\gamma_\phi^4 A^2}{r^6} \frac{d\Omega}{dr} - \Lambda_{ie} + F^-$$
 (2.9)

The involved variables:

$$\gamma_r = \frac{1}{\sqrt{1 - V^2}}, \quad \gamma_\phi = \sqrt{1 + \frac{r^2 L^2}{\mu^2 \gamma_r^2 A}}$$
(2.10)

$$\mu = 1 + \frac{W}{\Sigma_0} \left[\left(a_i + \frac{1}{\beta} \right) \frac{W_i}{W} + \left(a_e + \frac{1}{\beta} \right) \frac{W_e}{W} \right]$$
 (2.11)

$$a_i = \frac{1}{\gamma_i - 1} + \frac{2(1 - \beta)}{\beta}, \quad a_e = \frac{1}{\gamma_e - 1} + \frac{2(1 - \beta)}{\beta}$$
 (2.12)

$$\gamma_i = 1 + \theta_i \left[\frac{3K_3(1/\theta_i) + K_1(1/\theta_i)}{4K_2(1/\theta_i)} - 1 \right]^{-1}$$
(2.13)

$$\gamma_e = 1 + \theta_e \left[\frac{3K_3(1/\theta_e) + K_1(1/\theta_e)}{4K_2(1/\theta_e)} - 1 \right]^{-1}$$
(2.14)

$$\Gamma_i = 1 + \left[a_i \left(1 + \frac{d \ln a_i}{d \ln T_i} \right) \right]^{-1}, \qquad \Gamma_e = 1 + \left[a_e \left(1 + \frac{d \ln a_e}{d \ln T_e} \right) \right]^{-1}$$
 (2.15)

$$\Omega_K^{\pm} = \pm \frac{M^{1/2}}{r^{3/2} \pm aM^{1/2}}, \quad \Omega = \omega + \frac{r^3 \Delta^{1/2} L}{\mu \gamma_r \gamma_\phi A^{3/2}}$$
(2.16)

The vertical scale height is

$$h^{2} = \frac{W}{\mu \Sigma_{0}} \frac{r^{4}}{(L/\mu)^{2} - a^{2}[(E/\mu)^{2} - 1]}$$
 (2.17)

where

$$L = \mu u_{\phi}, \quad E = -\mu u_{t}, \quad \frac{u^{\phi}}{u^{t}} = \Omega \tag{2.18}$$

and

$$u_{\phi} = g_{\phi t}u^{t} + g_{\phi\phi}u^{\phi} = \left(\frac{g_{\phi t}}{\Omega} + g_{\phi\phi}\right)u^{\phi}, \quad u_{t} = g_{tt}u^{t} + g_{\phi t}u^{\phi} = \left(\frac{g_{tt}}{\Omega} + g_{t\phi}\right)u^{\phi}$$
(2.19)

Thus we obtain

$$E = -\mu u_t = -\frac{g_{tt} + \Omega g_{t\phi}}{g_{t\phi} + \Omega g_{\phi\phi}} L = \frac{r^4 \Delta / A^2 + \omega \Omega - \omega^2}{\Omega - \omega} L$$
 (2.20)

Note that

$$\Omega - \omega = \frac{r^3 \Delta^{1/2} L}{\mu \gamma_r \gamma_\phi A^{3/2}} \tag{2.21}$$

we have

$$E = \mu \gamma_r \gamma_\phi r \frac{\Delta^{1/2}}{A^{1/2}} + \omega L \tag{2.22}$$

• calculation of VdV/dr

$$\frac{V}{\sqrt{1 - V^2}} = -\frac{\dot{M}}{2\pi\Delta^{1/2}\Sigma_0} \tag{2.23}$$

Since

$$\frac{V^2}{1 - V^2} = \frac{\dot{M}^2}{4\pi^2} \frac{1}{\Delta \Sigma_0^2} = B \tag{2.24}$$

a simple calculation gives

$$V^2 = 1 - \frac{1}{1+B} \tag{2.25}$$

and thereby

$$2V\frac{dV}{dr} = \frac{1}{(1+B)^2}\frac{dB}{dr} \tag{2.26}$$

$$= -\frac{1}{(1+B)^2} \frac{\dot{M}^2}{4\pi^2} \frac{1}{\Delta \Sigma_0} \left[\frac{1}{\Delta} \frac{d\Delta}{dr} + \frac{2}{\Sigma_0} \frac{d\Sigma_0}{dr} \right]$$
 (2.27)

$$= -\frac{2B}{(1+B)^2} \left(\frac{r-M}{\Delta} + \frac{1}{\Sigma_0} \frac{d\Sigma_0}{dr} \right)$$
 (2.28)

• calculation of $d\Omega/dr$

$$\Omega = \omega + \frac{r^3 \Delta^{1/2} L}{\mu \gamma_r \gamma_\phi A^{3/2}}, \quad L = \frac{2\pi r W_\phi^r}{\dot{M}} + L_{in}, \quad W_\phi^r = \alpha \frac{A^{3/2} \Delta^{1/2} \gamma_\phi^3}{r^6} W$$
 (2.29)

$$\frac{d\Omega}{dr} = \frac{d\omega}{dr} + \frac{2\pi\alpha}{\dot{M}} \frac{d}{dr} \left(\frac{\gamma_{\phi}^2 \Delta W}{\mu \gamma_r r^2} \right) + L_{in} \frac{d}{dr} \left(\frac{r^3 \Delta^{1/2}}{\mu \gamma_r \gamma_{\phi} A^{3/2}} \right)$$
(2.30)

To sum up, we write down the equation set

$$\frac{dW_i}{dr} + \frac{dW_e}{dr} - \frac{\beta\mu\gamma_r^2B}{(1+B)^2}\frac{d\Sigma_0}{dr} = \frac{\beta\mu\gamma_r^2B\Sigma_0}{(1+B)^2}\frac{r-M}{\Delta} - \frac{\beta\mu\gamma_\phi^2\Sigma_0AM}{r^4\Delta}\frac{(\Omega-\Omega_K^+)(\Omega-\Omega_K^-)}{\Omega_K^+\Omega_K^-}$$
(2.31)

$$\left[\frac{1}{\Gamma_i - 1} \frac{\dot{M}}{\Sigma_0} - \frac{(1 - \delta)4\pi^2 \alpha^2 \gamma_\phi^6}{\mu \gamma_r \beta \dot{M}} \frac{\Delta A^2 W}{r^8} \right] \frac{dW_i}{dr} - \left[\frac{(1 - \delta)4\pi^2 \alpha^2 \gamma_\phi^6}{\mu \gamma_r \beta \dot{M}} \frac{\Delta A^2 W}{r^8} \right] \frac{dW_e}{dr} - \frac{\Gamma_i}{\Gamma_i - 1} \frac{W_i \dot{M}}{\Sigma_0^2} \frac{d\Sigma_0}{dr} \frac{d\Sigma_0}{dr} + \frac{1}{2} \frac{W_i \dot{M}}{r^8} \frac{d\Sigma_0}{dr} \frac{d\Sigma_0}{r^8} \frac{d\Sigma_0}{dr} + \frac{1}{2} \frac{W_i \dot{M}}{r^8} \frac{d\Sigma_0}{r^8} \frac{d\Sigma_0}{r^8$$

$$= (1 - \delta)2\pi\alpha W \frac{\gamma_{\phi}^4 A^2}{r^6} \left[\frac{2\pi\alpha W}{\dot{M}} \frac{d}{dr} \left(\frac{\gamma_{\phi}^2 \Delta}{\mu \gamma_r r^2} \right) + L_{in} \frac{d}{dr} \left(\frac{r^3 \Delta^{1/2}}{\mu \gamma_{\phi} \gamma_r A^{3/2}} \right) + \frac{d\omega}{dr} \right] - \frac{W_i \dot{M}}{r \Sigma_0} + 2\pi r \Lambda_{ie} \quad (2.32)$$

$$-\left[\frac{\delta 4\pi^2\alpha^2\gamma_{\phi}^6}{\mu\gamma_r\beta\dot{M}}\frac{\Delta A^2W}{r^8}\right]\frac{dW_i}{dr} + \left[\frac{1}{\Gamma_e-1}\frac{\dot{M}}{\Sigma_0} - \frac{\delta 4\pi^2\alpha^2\gamma_{\phi}^6}{\mu\gamma_r\beta\dot{M}}\frac{\Delta A^2W}{r^8}\right]\frac{dW_e}{dr} - \frac{\Gamma_e}{\Gamma_e-1}\frac{W_e\dot{M}}{\Sigma_0^2}\frac{d\Sigma_0}{dr}$$

$$=\delta 2\pi\alpha W\frac{\gamma_{\phi}^{4}A^{2}}{r^{6}}\left[\frac{2\pi\alpha W}{\dot{M}}\frac{d}{dr}\left(\frac{\gamma_{\phi}^{2}\Delta}{\mu\gamma_{r}r^{2}}\right)+L_{in}\frac{d}{dr}\left(\frac{r^{3}\Delta^{1/2}}{\mu\gamma_{\phi}\gamma_{r}A^{3/2}}\right)+\frac{d\omega}{dr}\right]-\frac{W_{e}\dot{M}}{r\Sigma_{0}}+2\pi r(F^{-}-\Lambda_{ie}) \quad (2.33)$$

Following we substitute

$$M = r_g = \frac{GM}{c^2},\tag{2.34}$$

$$\Delta = 1 - \frac{2r_g}{r} + \frac{a^2}{r^2}, \quad A = 1 + \frac{a^2}{r^2} + \frac{2r_g a^2}{r^3}, \quad \omega = \frac{2r_g a}{r^3 A}$$
 (2.35)

The equations set will be

$$\frac{dW_i}{dr} + \frac{dW_e}{dr} - \frac{\beta\mu\gamma_r^2B}{(1+B)^2}\frac{d\Sigma_0}{dr} = \frac{\beta\mu\gamma_r^2B\Sigma_0}{(1+B)^2}\frac{r - r_g}{r^2\Delta} - \frac{\beta\mu\gamma_\phi^2\Sigma_0Ar_g}{r^2\Delta}\frac{(\Omega - \Omega_K^+)(\Omega - \Omega_K^-)}{\Omega_K^+\Omega_K^-}$$
(2.36)

$$\left[\frac{1}{\Gamma_{i}-1}\frac{\dot{M}}{\Sigma_{0}} - \frac{(1-\delta)4\pi^{2}\alpha^{2}\gamma_{\phi}^{6}}{\mu\gamma_{r}\beta\dot{M}}r^{2}\Delta A^{2}W\right]\frac{dW_{i}}{dr} - \left[\frac{(1-\delta)4\pi^{2}\alpha^{2}\gamma_{\phi}^{6}}{\mu\gamma_{r}\beta\dot{M}}r^{2}\Delta A^{2}W\right]\frac{dW_{e}}{dr} - \frac{\Gamma_{i}}{\Gamma_{i}-1}\frac{W_{i}\dot{M}}{\Sigma_{0}^{2}}\frac{d\Sigma_{0}}{dr}$$

$$= (1 - \delta)2\pi\alpha W \gamma_{\phi}^{4} A^{2} r^{2} \left[\frac{2\pi\alpha W}{\dot{M}} \frac{d}{dr} \left(\frac{\gamma_{\phi}^{2} \Delta}{\mu \gamma_{r}} \right) + L_{in} \frac{d}{dr} \left(\frac{\Delta^{1/2}}{\mu \gamma_{\phi} \gamma_{r} r^{2} A^{3/2}} \right) + \frac{d\omega}{dr} \right] - \frac{W_{i} \dot{M}}{r \Sigma_{0}} + 2\pi r \Lambda_{ie} \quad (2.37)$$

$$-\left[\frac{\delta 4\pi^2\alpha^2\gamma_{\phi}^6}{\mu\gamma_r\beta\dot{M}}r^2\Delta A^2W\right]\frac{dW_i}{dr} + \left[\frac{1}{\Gamma_e-1}\frac{\dot{M}}{\Sigma_0} - \frac{\delta 4\pi^2\alpha^2\gamma_{\phi}^6}{\mu\gamma_r\beta\dot{M}}r^2\Delta A^2W\right]\frac{dW_e}{dr} - \frac{\Gamma_e}{\Gamma_e-1}\frac{W_e\dot{M}}{\Sigma_0^2}\frac{d\Sigma_0}{dr}$$

$$= \delta 2\pi \alpha W \gamma_{\phi}^{4} A^{2} r^{2} \left[\frac{2\pi \alpha W}{\dot{M}} \frac{d}{dr} \left(\frac{\gamma_{\phi}^{2} \Delta}{\mu \gamma_{r}} \right) + L_{in} \frac{d}{dr} \left(\frac{\Delta^{1/2}}{\mu \gamma_{\phi} \gamma_{r} r^{2} A^{3/2}} \right) + \frac{d\omega}{dr} \right] - \frac{W_{e} \dot{M}}{r \Sigma_{0}} + 2\pi r (F^{-} - \Lambda_{ie})$$
(2.38)

where

$$\frac{d\Delta}{dr} = \frac{2r_g}{r^2} - \frac{2a^2}{r^3}, \quad \frac{dA}{dr} = -\frac{2a^2}{r^3} - \frac{6r_g a^2}{r^4}$$
 (2.39)

$$\frac{d}{dr} \left(\frac{\Delta^{1/2}}{r^2 A^{3/2}} \right) = \frac{\Delta^{1/2}}{r^2 A^{3/2}} \left(\frac{1}{2\Delta} \frac{d\Delta}{dr} - \frac{2}{r} - \frac{3}{2A} \frac{dA}{dr} \right)$$
(2.40)

$$\frac{d\omega}{dr} = -\frac{2r_g a}{r^3 A} \left(\frac{3}{r} + \frac{1}{A} \frac{dA}{dr} \right) \tag{2.41}$$

The solution are

$$\frac{d\Sigma_0}{dr} = \frac{(c_3 - a_{31}c_1)(a_{22} - a_{21}) - (c_2 - a_{21}c_1)(a_{32} - a_{31})}{(a_{33} - a_{31}a_{13})(a_{22} - a_{21}) - (a_{23} - a_{21}a_{13})(a_{32} - a_{31})} = S$$
(2.42)

$$\frac{dW_i}{dr} = \frac{a_{22}c_1 - c_2}{a_{22} - a_{21}} - \frac{a_{22}a_{13} - a_{23}}{a_{22} - a_{21}}S$$
(2.43)

$$\frac{dW_e}{dr} = \frac{c_2 - a_{21}c_1}{a_{22} - a_{21}} - \frac{a_{23} - a_{21}a_{13}}{a_{22} - a_{21}}S$$
(2.44)

Once the solution for Σ_0 , W_e and W_i are obtained,

$$B = \frac{\dot{M}^2}{4\pi^2} \frac{1}{r^2 \Delta \Sigma_0^2}, \quad V = \sqrt{\frac{B}{1+B}}, \quad \gamma_r = \frac{1}{\sqrt{1-V^2}}$$
 (2.45)

$$L = L_{in} + \frac{2\pi r^2}{\dot{M}} \alpha A^{3/2} \Delta^{1/2} W \gamma_{\phi}^3, \qquad \gamma_{\phi} = \sqrt{1 + \frac{L^2}{\mu^2 \gamma_r^2 r^2 A}}$$
 (2.46)

$$T_i = \frac{\mu_i W_i m_H}{k \Sigma_0}, \quad T_e = \frac{\mu_e W_e m_H}{k \Sigma_0}$$
 (2.47)

2.2 The boundary conditions

$$\Omega = 0.8\Omega_K \tag{2.48}$$

$$T_i = T_e = 0.1 T_{vir}, \quad T_{vir} = (\gamma - 1) \frac{GMm_H}{kr}$$
 (2.49)

2.3 Dimension units

$$r: R_s, \quad L: cR_s, \quad \Omega: \frac{c}{R_s}, \quad \dot{M}: \dot{M}_{Edd}, \quad \Sigma: \frac{M_{Edd}}{cR_s}$$
 (2.50)

$$W: \frac{\dot{M}_{Edd}c}{R_s}, \quad Q_{rad}: \frac{\dot{M}_{Edd}c^2}{R_s^2}$$
 (2.51)

2.4 The transonic point

$$a_{22} - a_{21} = -\frac{1}{\Gamma_i - 1} \frac{\dot{M}}{\Sigma_0}, \quad a_{32} - a_{31} = \frac{1}{\Gamma_e - 1} \frac{\dot{M}}{\Sigma_0}$$
 (2.52)

$$a_{33} - a_{31}a_{13} = -\frac{\Gamma_e}{\Gamma_e - 1} \frac{\dot{M}W_e}{\Sigma_0^2} + \delta Y \frac{\beta \mu \gamma_r^2 B}{(1+B)^2}$$
 (2.53)

$$a_{23} - a_{21}a_{13} = -\frac{\Gamma_i}{\Gamma_i - 1} \frac{\dot{M}W_i}{\Sigma_0^2} + \left[\frac{1}{\Gamma_i - 1} \frac{\dot{M}}{\Sigma_0} + (1 - \delta)Y \right] \frac{\beta \mu \gamma_r^2 B}{(1 + B)^2}$$
(2.54)

Then we have

$$\mathcal{D} = (a_{33} - a_{31}a_{13})(a_{22} - a_{21}) - (a_{23} - a_{21}a_{13})(a_{32} - a_{31})$$

$$= -\frac{1}{\Gamma_{i} - 1} \frac{\dot{M}}{\Sigma_{0}} \left[-\frac{\Gamma_{e}}{\Gamma_{e} - 1} \frac{\dot{M}W_{e}}{\Sigma_{0}^{2}} + \delta Y \frac{\beta \mu \gamma_{r}^{2} B}{(1 + B)^{2}} \right]$$

$$-\frac{1}{\Gamma_{e} - 1} \frac{\dot{M}}{\Sigma_{0}} \left\{ -\frac{\Gamma_{i}}{\Gamma_{i} - 1} \frac{\dot{M}W_{i}}{\Sigma_{0}^{2}} + \left[\frac{1}{\Gamma_{i} - 1} \frac{\dot{M}}{\Sigma_{0}} + (1 - \delta)Y \right] \frac{\beta \mu \gamma_{r}^{2} B}{(1 + B)^{2}} \right\}$$
(2.55)

where

$$\beta \mu \gamma_r^2 BY = -\frac{\alpha^2 \gamma_r \gamma_\phi^6 A^2 W \dot{M}}{\Sigma_0^2}$$
 (2.56)

2.5. COMPARISON

$$-\frac{\Sigma_{0}}{\dot{M}}\mathcal{D} = -\frac{1}{(\Gamma_{i}-1)(\Gamma_{e}-1)} \left(\Gamma_{i}\frac{W_{i}}{W} + \Gamma_{e}\frac{W_{e}}{W}\right) \frac{\dot{M}W}{\Sigma_{0}^{2}} - \left(\frac{\delta}{\Gamma_{i}-1} + \frac{1-\delta}{\Gamma_{e}-1}\right) \frac{\alpha^{2}\gamma_{r}\gamma_{\phi}^{6}A^{2}}{(1+B)^{2}} \frac{\dot{M}W}{\Sigma_{0}^{2}} + \frac{1}{(\Gamma_{i}-1)(\Gamma_{e}-1)} \frac{\beta\mu\gamma_{r}^{4}}{(1+B)^{2}} \frac{B}{\gamma_{r}^{2}} \frac{\dot{M}}{\Sigma_{0}} = \frac{\dot{M}}{\Sigma_{0}} [b_{1}V^{2} - (b_{2}+b_{3})c_{s}^{2}]$$
 (2.57)

Therefore, the transonic point satisfies,

$$\frac{V^2}{c_s^2} = \frac{b_2 + b_3}{b_1} \tag{2.58}$$

where

$$b_1 = \frac{1}{(\Gamma_i - 1)(\Gamma_e - 1)} \frac{\beta \mu \gamma_r^4}{(1 + B)^2}$$
 (2.59)

$$b_2 = \frac{1}{(\Gamma_i - 1)(\Gamma_e - 1)} \left(\Gamma_i \frac{W_i}{W} + \Gamma_e \frac{W_e}{W} \right)$$
 (2.60)

$$b_3 = \left(\frac{\delta}{\Gamma_i - 1} + \frac{1 - \delta}{\Gamma_e - 1}\right) \frac{\alpha^2 \gamma_r \gamma_\phi^6 A^2}{(1 + B)^2}$$
 (2.61)

2.5 Comparison

The differences of equation set under the Newtonian theory and General relativity theory.

1. Continuity equation.

Newronian:

$$\dot{M} = -2\pi r \Sigma v_r \tag{2.62}$$

GR:

$$\dot{M} = -2\pi \Delta^{1/2} \Sigma \gamma_r V \tag{2.63}$$

In the Newtonian approximation, $\Delta = r^2 - 2Mr + a^2 \sim r^2$, $\gamma_r \sim 1$, the above equation will be

$$\dot{M} = -2\pi r \Sigma V. \tag{2.64}$$

2. Radial momentum equation.

Newtonian:

$$v_r \frac{dv_r}{dr} + \frac{1}{\Sigma} \frac{dW}{dr} = r(\Omega^2 - \Omega_K^2) - \frac{W}{\Sigma} \frac{d \ln \Omega_k}{dr}$$
 (2.65)

GR:

$$\gamma_r^2 V \frac{dV}{dr} + \frac{1}{\mu \Sigma} \frac{dW}{dr} = -\frac{\gamma_\phi^2 AM}{r^4 \Delta} \frac{(\Omega - \Omega_K^+)(\Omega - \Omega_K^-)}{\Omega_K^+ \Omega_K^-}.$$
 (2.66)

In the Newtonian approximation, $\gamma_r \sim 1$, $\mu \sim 1$, $A = r^4 + r^2 a^2 + 2Mra^2 \sim r^4$, and

$$\Omega_K^+ = -\Omega_K^- = \frac{M^{1/2}}{r^{3/2}} \tag{2.67}$$

the above equation will be

$$V\frac{dV}{dr} + \frac{1}{\Sigma}\frac{dW}{dr} = r(\Omega^2 - \Omega_K^2)$$
 (2.68)

3. Azimuthal momentum equation.

Newtonian:

$$\frac{\dot{M}}{2\pi}(L - L_{in}) = r^2 \alpha W. \tag{2.69}$$

GR:

$$\frac{\dot{M}}{2\pi}(L - L_{in}) = \frac{A^{3/2} \Delta^{1/2} \gamma_{\phi}^{3}}{r^{5}} \alpha W$$
 (2.70)

In the Newtonian approxiamtion,

$$\frac{\dot{M}}{2\pi}(L - L_{in}) = r^2 \alpha W. \tag{2.71}$$

4. Energy equation.

Newtonian:

$$\frac{\dot{M}W_{i}}{\Sigma} \left[\frac{1}{\gamma - 1} \frac{d \ln W_{i}}{dr} - \frac{\gamma}{\gamma - 1} \frac{d \ln \Sigma}{dr} + \frac{d \ln H}{dr} \right] = 2\pi r^{2} \alpha W \frac{d\Omega}{dr} + 2\pi r \Lambda_{ie}$$
 (2.72)

2.5. COMPARISON 15

$$\frac{d\Omega}{dr} = \frac{2\pi\alpha}{\dot{M}} \frac{dW}{dr} - \frac{2L_{in}}{r^3} \tag{2.73}$$

GR:

$$\frac{\dot{M}W_i}{\Sigma} \left[\frac{1}{\Gamma_i - 1} \frac{d \ln W_i}{dr} - \frac{\Gamma_i}{\Gamma_i - 1} \frac{d \ln \Sigma}{dr} + \frac{1}{r} \right] = 2\pi \alpha W \frac{\gamma_{\phi}^4 A^2}{r^6} \frac{d\Omega}{dr} + 2\pi r \Lambda_{ie}$$
 (2.74)

$$\frac{d\Omega}{dr} = \frac{2\pi\alpha\gamma_{\phi}^{2}\Delta}{\mu\gamma_{r}\dot{M}r^{2}}\frac{dW}{dr} + \frac{2\pi\alpha W}{\dot{M}}\frac{d}{dr}\left(\frac{\gamma_{\phi}^{2}\Delta}{\mu\gamma_{r}r^{2}}\right) + L_{in}\frac{d}{dr}\left(\frac{r^{3}\Delta^{1/2}}{\mu\gamma_{r}\gamma_{\phi}A^{3/2}}\right) + \frac{d\omega}{dr}$$
(2.75)

Chapter 3

The Spectrum of ADAFs

3.1 Comptonization

3.1.1 Klein-Nishina formular

In the electron rest frame,

$$\sigma_{KN} = \frac{3}{8}\sigma_T \left(\frac{\nu}{\nu_i}\right)^2 \left(\frac{\nu}{\nu_i} + \frac{\nu_i}{\nu} - 1 + \alpha^2\right),\tag{3.1}$$

where α is cosine of the scattering angle. Now for a motive electron, using coordinate conversion formular in the special relativity can easily obtains,

$$\sigma_{KN} = \frac{3}{8}\sigma_T \left(\frac{x'}{x}\right)^2 \left(\frac{x'}{x} + \frac{x}{x'} - 1 + \alpha^2\right),\tag{3.2}$$

where $x = \gamma \omega' (1 - \beta \mu), x' = x/[1 + (1 - \alpha)x]$. Keep in mind that α in equ (2) is still defined in the electron rest frame while ω' is energy of the incoming photon in the lab frame. The energy of outgoing photon reads,

$$\omega = \frac{\gamma [\gamma (1 - \beta \mu) + \gamma \beta \cos \theta (\mu - \beta) + \beta \sin \theta (1 - \mu^2)^{1/2} \cos \phi}{1 + \gamma (1 - \beta \mu) (1 - \cos \theta) \omega'} \omega'$$
(3.3)

3.1.2 Scattering rate

The angle-averaged scattering rate is given by

$$R(\omega, \gamma) = c \int_{-1}^{1} \frac{d\mu}{2} (1 - \beta\mu) \int_{-1}^{1} d\alpha \sigma_{KN}$$
 (3.4)

Standard asympotic forms for the rate R are,

• $\gamma\omega << 1$

$$R(\omega, \gamma) = c\sigma_T \left[1 - \frac{2\gamma\omega}{3} (3 + \beta^2) \right]$$
 (3.5)

• $\gamma\omega >> 1$

$$R(\omega, \gamma) = \frac{3c\sigma_T}{8\gamma\omega} \left\{ \left[1 - \frac{2}{\gamma\omega} - \frac{2}{(\gamma\omega)^2} \right] \ln(1 + \gamma\omega) + \frac{1}{2} - \frac{4}{\gamma\omega} - \frac{1}{2(1 + 2\gamma\omega)^2} \right\}$$
(3.6)

3.1.3 Mean scattered photon energy

$$\langle \omega \rangle (\omega', \gamma) = \frac{c}{R(\omega', \gamma)} \int_{-1}^{1} \frac{d\mu}{2} (1 - \beta \mu) \int_{-1}^{1} d\alpha \omega \sigma_{KN}$$
 (3.7)

3.1.4 Dispersion about $\langle \omega \rangle$

$$\langle \omega^2 \rangle (\omega', \gamma) = \frac{c}{R(\omega', \gamma)} \int_{-1}^1 \frac{d\mu}{2} (1 - \beta\mu) \int_{-1}^1 d\alpha \omega^2 \sigma_{KN}$$
 (3.8)

where

$$\omega^{2} = \frac{\gamma^{2} \omega^{2} \{ [\gamma (1 - \beta \mu) + \gamma \beta \alpha (\mu - \beta)]^{2} + \frac{1}{2} \beta^{2} (1 - \alpha^{2}) (1 - \beta^{2}) \}}{[1 + \gamma (1 - \beta \mu) (1 - \alpha) \omega]^{2}}$$
(3.9)

The dispersion is then given by

$$\langle \Delta \omega^2 \rangle = \langle \omega^2 \rangle - \langle \omega \rangle^2 \tag{3.10}$$

19

3.1.5 Scattered-photon distribution

$$P(\omega; \omega', \gamma) = \frac{1}{2D(\omega', \gamma)} H[D(\omega', \gamma) - |\omega - \langle \omega \rangle|]$$
(3.11)

where H(x) is the Heaviside function, and $D(\omega', \gamma)$ is defined as

$$D(\omega', \gamma) = \min\{ [\sqrt{3\langle \Delta\omega^2 \rangle}, (\langle \omega \rangle - \omega_{min}) \}$$
 (3.12)

3.1.6 Kinetic equation

$$\frac{dn(\omega)}{dt} = -n(\omega) \int d\gamma N(\gamma) R(\omega, \gamma) + \int \int d\omega' d\gamma P(\omega; \omega', \gamma) R(\omega', \gamma) N(\gamma) n(\omega')$$
(3.13)

Acctually, in the calculation we only handle the second integeration,

$$\frac{dn(\omega)}{dt} = \int \int d\omega' d\gamma P(\omega; \omega', \gamma) R(\omega', \gamma) N(\gamma) n(\omega')$$

$$= (\ln 10)^2 \int \int d\log \omega' d\log \gamma P(\omega; \omega', \gamma) R(\omega', \gamma) N(\gamma) n(\omega') \omega' \gamma \qquad (3.14)$$

For simplicity, we take dt = H/c, consequently,

$$n(\omega) = (\ln 10)^2 \frac{H}{c} \int \int d\log \omega' d\log \gamma P(\omega; \omega', \gamma) R(\omega', \gamma) N_e(\gamma) n(\omega') \omega' \gamma$$
 (3.15)

$$= (\ln 10)^{2} N_{e} H \sigma_{T} \int \int d \log \omega' d \log \gamma P(\omega; \omega', \gamma) \frac{R(\omega', \gamma)}{c \sigma_{T}} \frac{N(\gamma)}{N_{e}} n(\omega') \omega' \gamma \qquad (3.16)$$

$$= (\ln 10)^{2} \tau_{e} \int \int d\log \omega' d\log \gamma P(\omega; \omega', \gamma) \frac{R(\omega', \gamma)}{c\sigma_{T}} \frac{N(\gamma)}{N_{e}} n(\omega') \omega' \gamma$$
 (3.17)

3.1.7 Electron distribution

The relativistic Maxwell-Boltzmann distribution is gien by (Özel et al 2000)

$$N(\gamma) = N_{th} \gamma^2 \beta \frac{\exp\left(-\gamma/\theta_e\right)}{\theta_e K_2(1/\theta_e)}$$
(3.18)

3.2 Radiation mechanisms

3.2.1 Synchrotron emission

$$\chi_{syn} = 4.43 \times 10^{-30} \frac{4\pi n_e \nu}{K_2 (1/\theta_e)} I\left(\frac{4\pi m_e c \nu}{3e B\theta_e^2}\right)$$
(3.19)

where

$$I(x) = \frac{4.0505}{x^{1/6}} \left(1 + \frac{0.4}{x^{1/4}} + \frac{0.5316}{x^{1/2}} \right) \exp(-1.8899x^{1/3})$$
 (3.20)

3.2.2 Bremstralung emission

$$\chi_{br} = q_{br}^{-} G \exp\left(\frac{h\nu}{kT_e}\right) \tag{3.21}$$

where G is the Gaunt factor, q_{br}^- is the bremsstrahlung emission per unit volume , which reads

$$q_{br}^{-} = 1.48 \times 10^{-22} n_e^2 F(\theta)$$
 (3.22)

$$\begin{split} F(\theta) &= 4 \left(\frac{2\theta_e}{\pi^3}\right)^{1/2} (1 + 1.781\theta_e^{1.34}) + 1.73\theta_e^{3/2} (1 + 1.1\theta_e + \theta_e^2 - 1.25\theta_e^{5/2}) & \theta_e < 1 \\ &= \left(\frac{9\theta_e}{2\pi}\right) \left[\ln(0.48 + 1.123\theta_e) + 1.5\right] + 2.30\theta_e (\ln 1.123\theta_e + 1.28) & \theta_e > 1 \end{split} \tag{3.23}$$

Gaunt factor is

$$G = \frac{h}{kT_e} \left(\frac{3}{\pi} \frac{kT_e}{h\nu} \right)^{1/2} \qquad \frac{kT_e}{h\nu} < 1 \tag{3.24}$$

$$G = \frac{h}{kT_e} \frac{\sqrt{3}}{\pi} \ln \left(\frac{4}{\xi} \frac{kT_e}{h\nu} \right)^{1/2} \qquad \frac{kT_e}{h\nu} > 1$$
 (3.25)

3.3. ADAF MODEL 21

3.2.3 Compton scattering

The energy enhancement factor η is

$$\eta = \exp[s(A-1)][1 - P(j_m+1,As)] + \eta_{max}P(j_m+1,s)$$
(3.26)

where *P* is the incomplete gamma function and

$$A = 1 + 4\theta_e + 16\theta_e^2, \quad s = \tau_{es} + \tau_{es}^2$$
 (3.27)

$$\eta_{max} = \frac{3kT_e}{h\nu}, \qquad j_m = \frac{\ln \eta_{max}}{\ln A}, \qquad \tau_{es} = 2n_e \sigma_T H$$
(3.28)

With the energy enhancement factor η , the local radiative cooling rate Q_{rad}^- is given by

$$Q_{rad}^{-} = \int d\nu 2\eta F_{\nu} \tag{3.29}$$

3.3 ADAF model

3.3.1 Self-simular solution

$$v = -2.12 \times 10^{1}0\alpha c_{1}r^{-1/2} \text{ cm s}^{-1}$$

$$\Omega = 7.19 \times 10^{4}c_{2}m^{-1}r^{-3/2} \text{ s}^{-1}$$

$$c_{s}^{2} = 4.50 \times 10^{20}c_{3}r^{-1} \text{ cm s}^{-1}$$

$$\rho = 3.79 \times 10^{-5}\alpha^{-1}c_{1}^{-1}c_{3}^{-1/2}m^{-1}\dot{m}r^{-3/2} \text{ g cm}^{-3}$$

$$p = 1.71 \times 10^{16}\alpha^{-1}c_{1}^{-1}c_{3}^{1/2}m^{-1}\dot{m}r^{-5/2} \text{ g cm}^{-1} \text{ s}^{-2}$$

$$q^{+} = 1.84 \times 10^{21}\epsilon'c_{3}^{1/2}m^{-2}\dot{m}r^{-4} \text{ ergs cm}^{-3} \text{ s}^{-1}$$

$$\tau_{es} = 2n_{e}\sigma_{T}H = 12.4\alpha^{-1}c_{1}^{-1}\dot{m}r^{-1/2}$$

3.3.2 Global structure

3.3.3 Radiation transfer

$$F_{\nu} = \frac{2\pi}{\sqrt{3}} B_{\nu} [1 - \exp(-2\sqrt{3}\tau_{\nu}^{*})]$$
 (3.31)

$$\tau_{\nu}^* = \frac{\sqrt{\pi}}{2} \kappa_{\nu}(0) H \tag{3.32}$$

$$\kappa_{\nu} = \frac{\chi_{\nu}}{4\pi B_{\nu}} = \frac{\chi_{\nu,syn} + \chi_{\nu,br}}{4\pi B_{\nu}} \tag{3.33}$$

$$B_{\nu} = \frac{2h\nu^3}{c^2} \frac{1}{\exp(h\nu/kT) - 1}$$
 (3.34)

Chapter 4

The ray tracing method

4.1 Jacobian Elliptic Function

We denote $F(\phi, k)$ as the Incomplete elliptic integral of first kind.

$$sn^{-1}(\sin\phi|k^2) = F(\phi, k)$$
 (4.1)

$$cn^{-1}(\cos\phi|k^2) = F(\phi, k)$$
 (4.2)

$$tn^{-1}(\tan\phi|k^2) = F(\phi, k)$$
 (4.3)

4.2 Integrals

$$R(r) = r^4 + (a^2 - \lambda^2 - Q)r^2 + 2[Q + (\lambda - a)^2]r - a^2Q$$
(4.4)

$$\Theta(\mu) = Q + (a^2 - \lambda^2 - Q)\mu^2 - a^2\mu^4$$
(4.5)

4.2.1 r-component

The four roots for R(r) = 0 are

$$r_a = \frac{1}{2}N - \frac{1}{2}\sqrt{-M + 2D - 4C/N} \tag{4.6}$$

$$r_b = \frac{1}{2}N + \frac{1}{2}\sqrt{-M + 2D - 4C/N} \tag{4.7}$$

$$r_c = -\frac{1}{2}N - \frac{1}{2}\sqrt{-M + 2D + 4C/N}$$
(4.8)

$$r_d = -\frac{1}{2}N + \frac{1}{2}\sqrt{-M + 2D + 4C/N}$$
(4.9)

where

$$C = (a - \lambda)^2 + Q, \quad D = \frac{2}{3}(Q + \lambda^2 - a^2)$$
 (4.10)

$$E = \frac{9}{4}D^2 - 12a^2Q, \quad F = -\frac{27}{4}D^3 - 108a^2QD + 108C^2$$
 (4.11)

$$M = \frac{1}{3} \left(\frac{F - \sqrt{F^2 - 4E^3}}{2} \right)^{1/3} + \frac{1}{3} \left(\frac{F + \sqrt{F^2 - 4E^3}}{2} \right)^{1/3}$$
(4.12)

$$N = \sqrt{M + D} \tag{4.13}$$

1. Case I: four real roots (Byrd & Friedman 1954, eq [258.00]).

We denote the four real roots by r_1 , r_2 , r_3 and r_4 , with $r_1 \ge r_2 \ge r_3 \ge r_4$.

$$\int_{r_1}^{r} \frac{dr}{\sqrt{R(r)}} = \int_{r_2}^{r} \frac{dr}{\sqrt{(r-r_1)(r-r_2)(r-r_3)(r-r_4)}}$$
(4.14)

$$= \frac{2}{\sqrt{(r_1 - r_3)(r_2 - r_4)}} sn^{-1} \left(\sqrt{\frac{(r_2 - r_4)(r - r_1)}{(r_1 - r_4)(r - r_2)}} \right| m_4 \right)$$
(4.15)

where

$$m_4 = \frac{(r_1 - r_4)(r_2 - r_3)}{(r_1 - r_3)(r_2 - r_4)} \tag{4.16}$$

4.2. INTEGRALS 25

2. Case II: two real roots and two complex roots (Byrd & Friedman 1954, eq [260.00]).

We asume two complex roots are r_1 and r_2 ($r_1 = \overline{r_2}$), two real roots are r_3 and r_4 with $r_3 > r_4$.

$$\int_{r_1}^{r} \frac{dr}{\sqrt{R(r)}} = \int_{r_1}^{r} \frac{dr}{\sqrt{(r-r_1)(r-\overline{r_1})(r-r_3)(r-r_4)}}$$
(4.17)

$$= \frac{2}{\sqrt{AB}}cn^{-1} \left[\sqrt{\frac{(A-B)r + r_3A - r_4A}{(A+B)r - r_3A - r_4A}} \right] m_2$$
 (4.18)

where

$$u = \frac{r_1 + \overline{r_1}}{2} \qquad v^2 = -\frac{(r_1 - \overline{r_1})^2}{4} \tag{4.19}$$

$$A^{2} = (r_{3} - u)^{2} + v^{2}, \qquad B^{2} = (r_{4} - u)^{2} + v^{2}$$
 (4.20)

$$m_2 = \frac{(A+B)^2 - (r_3 - r_4)^2}{4AB} \tag{4.21}$$

3. Case III: four complex roots (Byrd & Friedman 1954, eq [267.00]).

The four complex roots are r_1 , r_2 , r_3 and r_4 ($r_1 = \overline{r_2}$, $r_3 = \overline{r_4}$). We introduce denotations:

$$u_1 = \frac{r_1 + \overline{r_1}}{2}$$
 $v_1^2 = -\frac{(r_1 - \overline{r_1})^2}{4}$ $u_2 = \frac{r_3 + \overline{r_3}}{2}$ $v_2^2 = -\frac{(r_3 - \overline{r_3})^2}{4}$ (4.22)

$$A^{2} = (u_{1} - u_{2})^{2} + (v_{1} + v_{2})^{2} \qquad B^{2} = (u_{1} - u_{2})^{2} + (v_{1} - v_{2})^{2}$$
(4.23)

$$g_1^2 = \frac{4v_1^2 - (A - B)^2}{(A + B)^2 - 4v_1^2} \qquad r_0 = u_1 - v_1 * g_1$$
 (4.24)

$$\int_{r_0}^{r} \frac{dr}{\sqrt{R(r)}} = \int_{r_0}^{r} \frac{dr}{\sqrt{(r-r_1)(r-\overline{r_1})(r-r_3)(r-\overline{r_3})}}$$
(4.25)

$$= \frac{2}{A+B}tn^{-1}\left(\frac{r-u_1+v_1g_1}{v_1+g_1u_1-g_1r}\middle|m_c\right) \tag{4.26}$$

where

$$m_c = \frac{4AB}{(A+B)^2} \tag{4.27}$$

4.2.2 θ -component

1. $Q \ge 0$ (Byrd & Friedman 1954, eq [213.00])

$$\Theta(\mu) = a^2(\mu_-^2 + \mu^2)(\mu_+^2 - \mu^2) \quad (0 \le \mu^2 \le \mu_+^2)$$
(4.28)

$$\mu_{\pm}^{2} = \frac{1}{2a^{2}} \left\{ \left[\left(\lambda^{2} + Q - a^{2} \right)^{2} + 4a^{2}Q \right]^{1/2} \mp \left(\lambda^{2} + Q - a^{2} \right) \right\}$$
(4.29)

$$\int_{\mu}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} = \frac{1}{\sqrt{a^{2}(\mu_{+}^{2} + \mu_{-}^{2})}} cn^{-1} \left(\frac{\mu}{\mu_{+}} \middle| m_{\mu}\right)$$
(4.30)

where

$$m_{\mu} = \frac{\mu_{+}^{2}}{\mu_{+}^{2} + \mu_{-}^{2}} \tag{4.31}$$

2. Q < 0 (Byrd & Friedman 1954, eq [218.00])

$$\Theta(\mu) = a^2(\mu^2 - \mu_-^2)(\mu_+^2 - \mu^2) \quad (\mu_-^2 \le \mu^2 \le \mu_+^2)$$
(4.32)

$$\mu_{\pm}^{2} = \frac{1}{2a^{2}} \left\{ \left(|Q| + a^{2} - \lambda^{2} \right) \pm \left[\left(|Q| + a^{2} - \lambda^{2} \right)^{2} - 4a^{2} |Q| \right]^{1/2} \right\}$$
(4.33)

$$\int_{\mu}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} = \frac{1}{a\mu_{+}} sn^{-1} \left(\sqrt{\frac{\mu_{+}^{2} - \mu^{2}}{\mu_{+}^{2} - \mu_{-}^{2}}} \middle| m_{\mu} \right)$$
(4.34)

where

$$m_{\mu} = \frac{\mu_{+}^{2} - \mu_{-}^{2}}{\mu_{+}^{2}} \tag{4.35}$$

4.3 TeV Trajectory

$$\tau_r = \pm \int_{r_o}^{\infty} \frac{dr}{\sqrt{R(r)}} = \pm \int_{\mu_o}^{\mu_{obs}} \frac{d\mu}{\sqrt{\Theta(\mu)}} = \tau_{\mu}$$
 (4.36)

where

$$R(r) = r^4 + (a^2 - \lambda^2 - Q)r^2 + 2[Q + (\lambda - a)^2]r - a^2Q$$
(4.37)

$$\Theta(\mu) = Q + (a^2 - \lambda^2 - Q)\mu^2 - a^2\mu^4 \tag{4.38}$$

Now we know the initial condition r_e , μ_e , μ_{obs} and assume $\alpha = 0$ (that is to say $\lambda = 0$), we need solve the β to calculate another constant of motion $Q = \beta^2 + (\alpha^2 - a^2)\cos^2\theta_{obs}$.

27

4.3.1 Calculating τ_{μ}

$$\tau_{\mu} = \int_{\mu_{e}}^{\mu_{\text{obs}}} \frac{d\mu}{\sqrt{\Theta(\mu)}} = \int_{\mu_{e}}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} - \int_{\mu_{\text{obs}}}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} \quad \text{without turning point}$$

$$= \int_{\mu_{e}}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} + \int_{\mu_{\text{obs}}}^{\mu_{+}} \frac{d\mu}{\sqrt{\Theta(\mu)}} \quad \text{with turning point}$$

$$(4.39)$$

4.3.2 Solving r_e Given τ_{μ}

$$\tau_{r} = \int_{r_{e}}^{\infty} \frac{dr}{\sqrt{R}} = \int_{r_{+}}^{\infty} \frac{dr}{\sqrt{R}} + \int_{r_{+}}^{r_{e}} \frac{dr}{\sqrt{R}} = \tau_{\infty} + \tau_{e} \quad \text{with turning point}$$

$$= \int_{r_{+}}^{\infty} \frac{dr}{\sqrt{R}} - \int_{r_{+}}^{r_{e}} \frac{dr}{\sqrt{R}} = \tau_{\infty} - \tau_{e} \quad \text{without turning point}$$

$$(4.41)$$

where $r_+ = r_1$ for Case I, $r_+ = r_3$ for Case II and $r_+ = r_0$ for Case III.

1. Case I: four real roots.

$$\tau_{\infty} = \frac{2}{\sqrt{(r_1 - r_3)(r_2 - r_4)}} sn^{-1} \left(\sqrt{\frac{r_2 - r_4}{r_1 - r_4}} \middle| m_4 \right)$$
(4.43)

$$\tau_e = \frac{2}{\sqrt{(r_1 - r_3)(r_2 - r_4)}} s n^{-1} \left[\frac{(r_2 - r_4)(r_e - r_1)}{(r_1 - r_4)(r_e - r_2)} \right] m_4$$
(4.44)

The solution is

$$r_e = \frac{r_1(r_2 - r_4) - r_2(r_1 - r_4)sn^2(\xi_4|m_4)}{(r_2 - r_4) - (r_1 - r_4)sn^2(\xi_4|m_4)}$$
(4.45)

where

$$\xi_4 = \frac{1}{2} (\tau_\mu - \tau_\infty) \sqrt{(r_1 - r_3)(r_2 - r_4)}$$
(4.46)

2. Case II: two real roots and two complex roots.

$$\tau_{\infty} = \frac{1}{\sqrt{AB}} c n^{-1} \left(\frac{A - B}{A + B} \middle| m_2 \right) \tag{4.47}$$

The solution is

$$r_e = \frac{r_4 A - r_3 B - (r_4 A + r_3 B) cn(\xi_2 | m_2)}{(A - B) - (A + B) cn(\xi_2 | m_2)}$$
(4.48)

where

$$\xi_2 = (\tau_\mu - \tau_\infty)\sqrt{AB} \tag{4.49}$$

3. Case III: four complex roots.

$$\tau_{\infty} = \frac{2}{A+B}tn^{-1}\left(-\frac{1}{g_1}\left|m_c\right.\right) \tag{4.50}$$

$$\tau_e = \frac{2}{A+B} t n^{-1} \left(\frac{r_e - u_1 + v_1 g_1}{v_1 + g_1 u_1 - g_1 r_e} \middle| m_c \right)$$
(4.51)

The solution is

$$r_e = \frac{u_1 - v_1 g_1 + (v_1 + u_1 g_1) t n(\xi_c | m_c)}{1 + g_1 t n(\xi_c | m_c)}$$
(4.52)

where

$$\xi_c = \frac{1}{2} (\tau_{\mu} - \tau_{\infty}) (A + B) \tag{4.53}$$

4.3.3 Solving β

Specifying a β , then we get λ and Q. We can calculate directly τ_{μ} , and go on to solve the r_e . If r_e matches the initial condition, we find out the proper β .

Chapter 5

Mathematical functions

5.1 Bessel functions

The approximation of Bessel functions for large x is

$$K_n(x) \approx \sqrt{\frac{\pi}{2x}} e^{-x} \left(1 + \frac{4n^2 - 1}{8x} + \dots \right)$$
 (5.1)

Therefore for small θ ,

$$K_3(1/\theta) = \sqrt{\frac{\pi}{2}} \theta^{1/2} e^{-1/\theta} \left(1 + \frac{35\theta}{8} \right),$$
 (5.2)

$$K_2(1/\theta) = \sqrt{\frac{\pi}{2}} \theta^{1/2} e^{-1/\theta} \left(1 + \frac{15\theta}{8} \right),$$
 (5.3)

$$K_1(1/\theta) = \sqrt{\frac{\pi}{2}} \theta^{1/2} e^{-1/\theta} \left(1 + \frac{3\theta}{8} \right),$$
 (5.4)

and

$$\gamma = 1 + \theta \left[\frac{3K_3(1/\theta) + K_1(1/\theta)}{4K_2(1/\theta)} - 1 \right]^{-1} = 1 + \frac{8 + 15\theta}{12}.$$
 (5.5)