The Constantly Changing Hubble Constant

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Find the presentation at
https://tinyurl.com/bycke8v6

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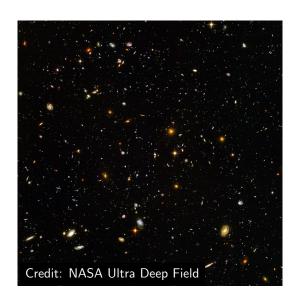
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The Universe is expanding!

- But what does this actually mean?
- How do we know it is expanding?
- Why is it expanding?
- How fast is it expanding?
- ► Are cosmologists completely realistic about the uncertainties in their results?

How do we know?

- Everywhere we look, distant galaxies are receding; more distant galaxies are receding faster.
- ➤ So either we are at the centre of a cosmic conspiracy, or all the space between all the galaxies is expanding.



Is the solar system expanding? Are we expanding?

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- 1. Yes, a lot
- 2. Yes, but only a tiny amount
- 3. No

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Is the solar system expanding? Are we expanding?

- Other forces molecular forces between the molecules in your body, and gravitational forces between the Sun and the planets - are far more than strong enough to overcome the effect of cosmic expansion.
- Gravity is even strong enough to keep the Andromeda Galaxy from receding from us.



▶ It's only the furthest objects - where gravity becomes negligible - that recede.

What does recession velocity actually mean?

- We say 'distant galaxies are moving away from us'. This is informal language.
- ► They aren't really moving, they just appear to be because the intervening space is expanding.
- ➤ Sometimes this makes a difference for example, the recession velocity can exceed the speed of light.

- For every additional distance of one megaparsec, there's an additional recession velocity of about 70 kilometers per second.
- ➤ So the expansion speed is about 70 kilometers per second per megaparsec.
- One megaparsec is about three million light years. It's the typical distance between galaxies.
- ▶ 70 kilometers per second is about 150,000 miles per hour.

Start with:

Start with:

13.5 million years later:

Start with:
13.5 million years later:
•

Continental drift is about six times faster...

- \triangleright The expansion rate is denoted ' H_0 '.
- ► The 'H' commemorates Edwin Hubble (1889-1953), who was one of the first to measure it.



▶ The '0' refers to today. The expansion rate was different in the distant past.



Credit: Johan Hagemeyer (1884-1962), Public domain, via Wikimedia Common