#### The Constantly Changing Hubble Constant

Lorne Whiteway lorne.whiteway@star.ucl.ac.uk

Astrophysics Group Department of Physics and Astronomy University College London

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Find the presentation at
https://tinyurl.com/bycke8v6

#### Interactive content

You are invited to go to

www.menti.com

and enter code

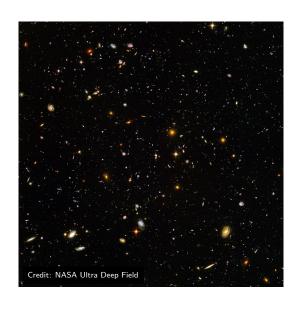
14 11 05 9

# The Universe is expanding!

- But what does this actually mean?
- How do we know it is expanding?
- Why is it expanding?
- How fast is it expanding?
- ► Are cosmologists completely realistic about the uncertainties in their results?

#### How do we know?

- Everywhere we look, distant galaxies are receding; more distant galaxies are receding faster.
- So either we are at the centre of a cosmic conspiracy, or all the space between all the galaxies is expanding.



Is the solar system expanding? Are we expanding?

Go to www.menti.com (code 14  $11\ 05\ 9$ ) and choose one possibility:

- 1. Yes, a lot
- 2. Yes, but only a tiny amount
- 3. No

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# Is the solar system expanding? Are we expanding?

Other forces - molecular forces between the molecules in your body, and gravitational forces between the Sun and the planets - are far more than strong enough to overcome the effect of cosmic expansion.

 Gravity is even strong enough to keep the Andromeda Galaxy from receding from us.



► It's only the furthest objects - where gravity becomes negligible - that recede.

# What does recession velocity actually mean?

- We say 'distant galaxies are moving away from us'. This is informal language.
- ► They aren't really moving, they just appear to be because the intervening space is expanding.
- Sometimes this makes a difference for example, the recession velocity can exceed the speed of light.

Which 'Ed' first had the idea that the Universe is expanding?

Go to www.menti.com (code 14 11 05 9) and choose one possibility:

- 1. Edmond Halley
- 2. Edwin Hubble
- 3. Edgar Allan Poe

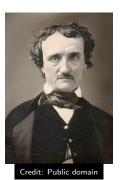
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# History

In 1848 Edgar Allan Poe published Eureka, which included a description of expanding space.



Expansion is not obvious without large telescopes and so isn't usually part of pre-modern cosmologies. Full understanding only came in the 20th century.

- For every additional distance of one megaparsec, there's an additional recession velocity of about 70 kilometers per second.
- ➤ So the expansion speed is about 70 kilometers per second per megaparsec.
- One megaparsec is about three million light years. It's the typical distance between galaxies.
- ▶ 70 kilometers per second is about 150,000 miles per hour.

Start with a distance:

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13.5 million years later it will be 1% longer:

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Continental drift is about six times faster...

# $H_0$

- ▶ The current expansion rate is called the *Hubble constant* or *Hubble parameter* and is denoted  ${}^{\prime}H_0{}^{\prime}$ .
- ➤ The 'H' commemorates Edwin Hubble (1889-1953), who was one of the first to measure it.



► The '0' refers to today. The expansion rate was different in the distant past.

# Why does the Universe expand?

- Science is not so good with 'why?' questions...
- ► There's an *initial condition*: the Universe started expanding at the Big Bang.
- ► The later behaviour of the expansion (does it slow down? speed up?) then depends, essentially via gravity, on what's in the Universe.

General relativity, our modern theory of gravity, is due to Einstein (1916).



Credit: Paul Ehrenfest

▶ Remember mass and energy are the same  $(E = mc^2)$ .

Mass/Energy bends spacetime, essentially changing distances and angles.



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- But it also works on the Universe as a whole mass/energy can cause distances to change everywhere in the Universe - and in particular can lead to increasing distances everywhere. This is the expansion that we see.

#### Contents of Universe control expansion

It was Alexander
 Friedmann (Александр
 Александрович
 Фридман) (1888-1925)
 who first realised this
 (1922).



Credit: Public domain

# What if we go backwards in time?

► George Lamaître (1894-1966) realised that if the Universe was expanding then it must, at an earlier stage, have been very small; he thereby invented the idea of the 'Big Bang'.



#### How do we measure the expansion rate?

- ► In theory it's easy: find a distant galaxy, measure its recession velocity and its distance, and take the ratio.
- ► Example: a galaxy is receding at 1600 kilometers per second and is 20 megaparsecs away; then H<sub>0</sub> is 80 kilometers per second per megaparsec.

# For a distant galaxy, which is harder to measure?

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# Can use redshift to measure recession velocity

- ► Light from distant galaxies gets streeeetched by the expansion; this makes it turn redder.
- It's fairly easy to measure the amount of red-shifting, as spectral lines are a convenient reference point. The redshift then immediately gives the velocity.

