

Law of Dulong-Petit  $C = 3k_b$  per atom

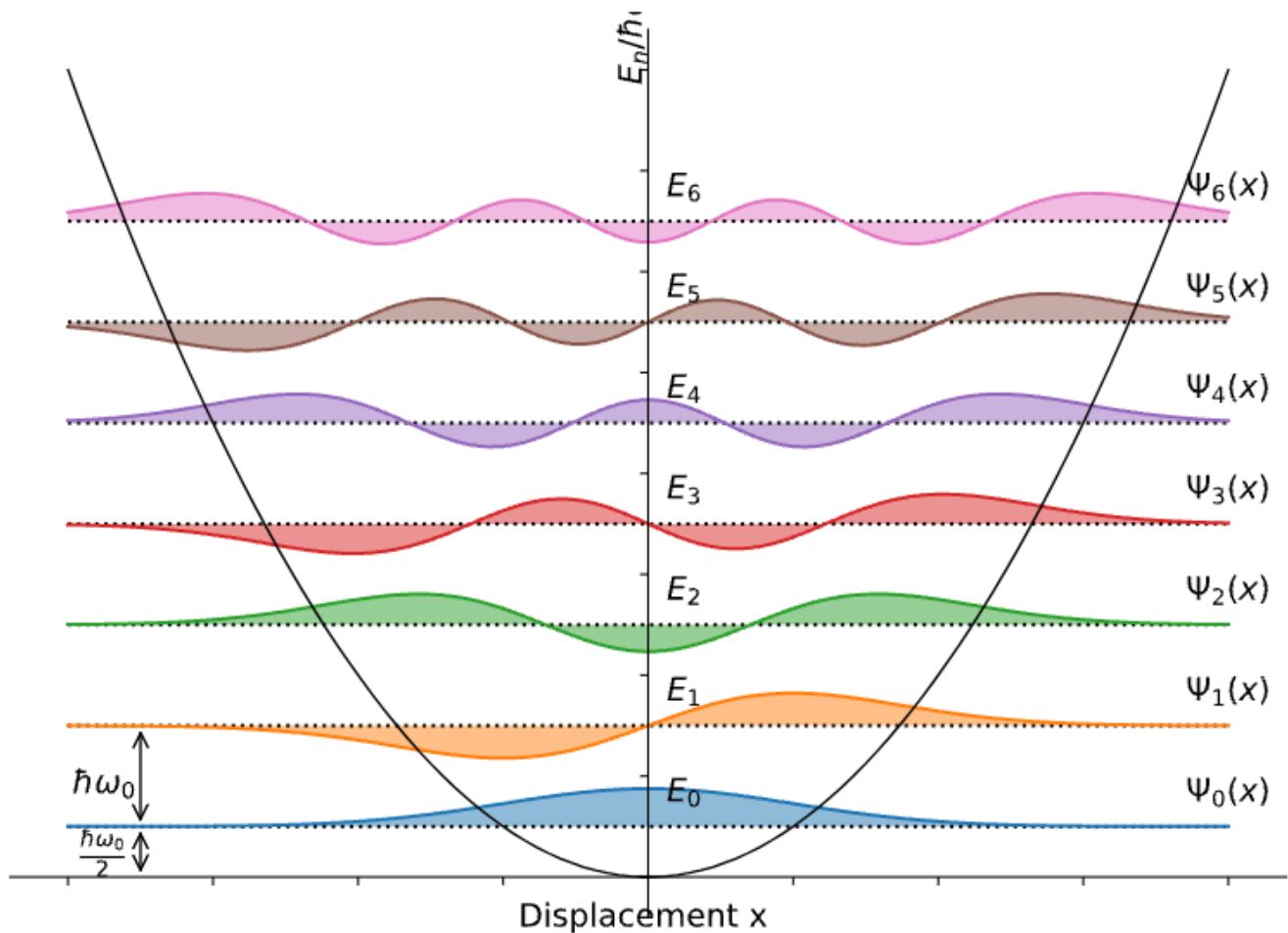
This law holds well at room temperature with deviations at lower temperatures

## Einstein's model

Equipartition theorem states that each atom can be modeled as a classical harmonic oscillator

Starting Point

- Each atom is modeled as independent quantum harmonic oscillators
- Each atom has the same frequency  $\omega_0$



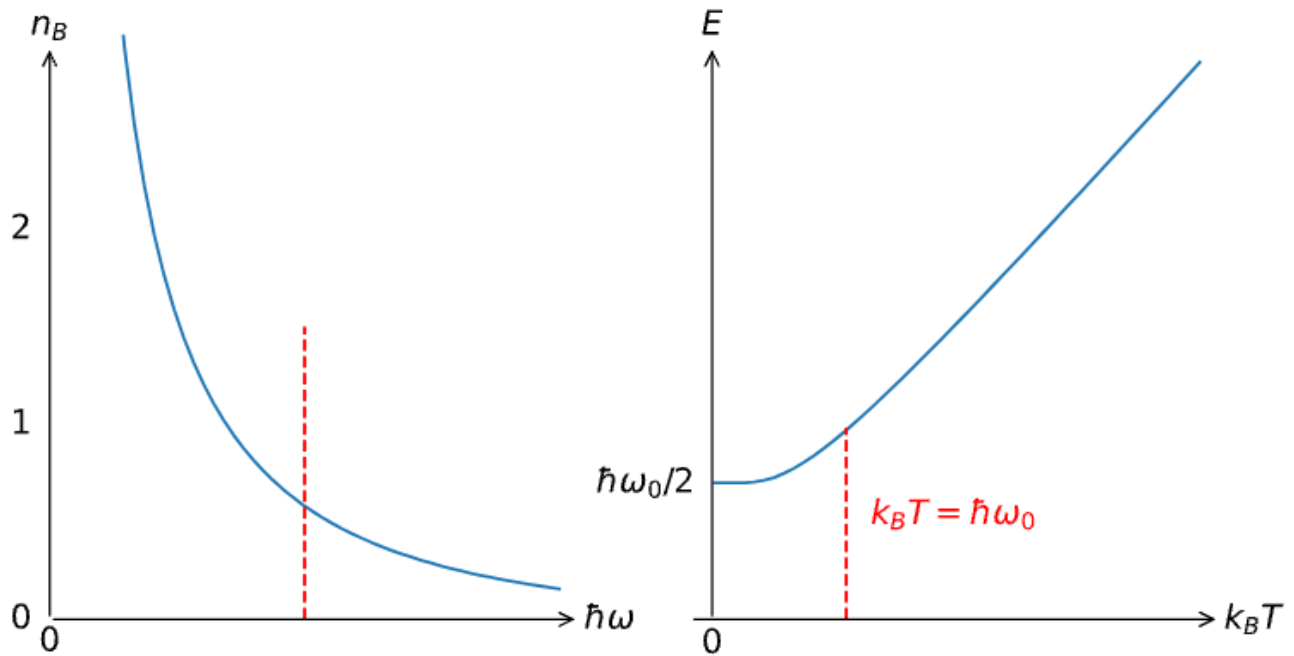
With eigenenergies  $E_n = (n + 1/2)\hbar\omega_0$

The occupation number is given by the Bose-Einstein distribution:

$$\langle n \rangle = n_B(\omega, T) = \frac{1}{e^{\hbar\omega/k_B T} - 1}$$

Resubstitution of this gives

$$E(T) = \frac{\hbar\omega_0}{e^{\hbar\omega_0/k_B T} - 1} + 0.5\hbar\omega_0$$



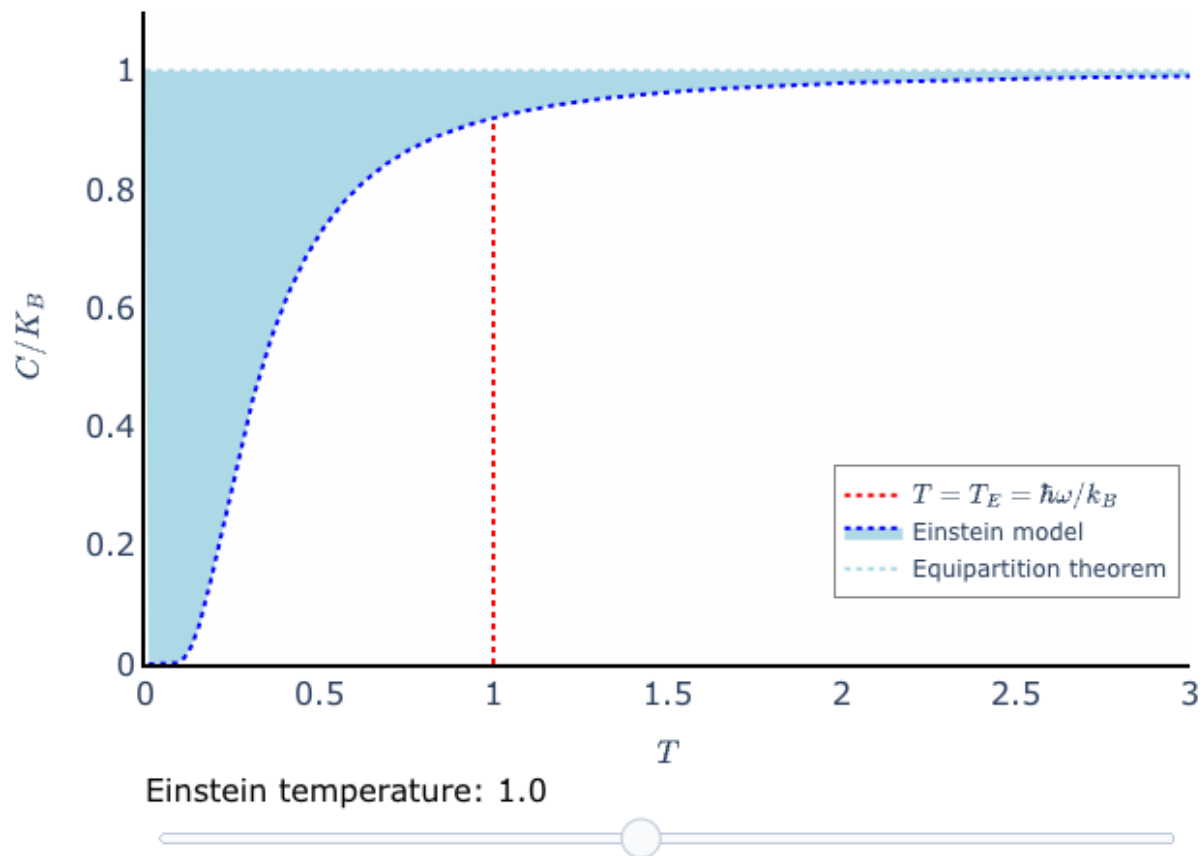
Now one can simply derive the heat capacity per harmonic oscillator by the definition

$$C(T) = \frac{dE(T)}{dT}$$

$$C(T) = k_b \left( \frac{T_E}{T} \right)^2 \frac{e^{T_E/T}}{(e^{T_E/T} - 1)^2},$$

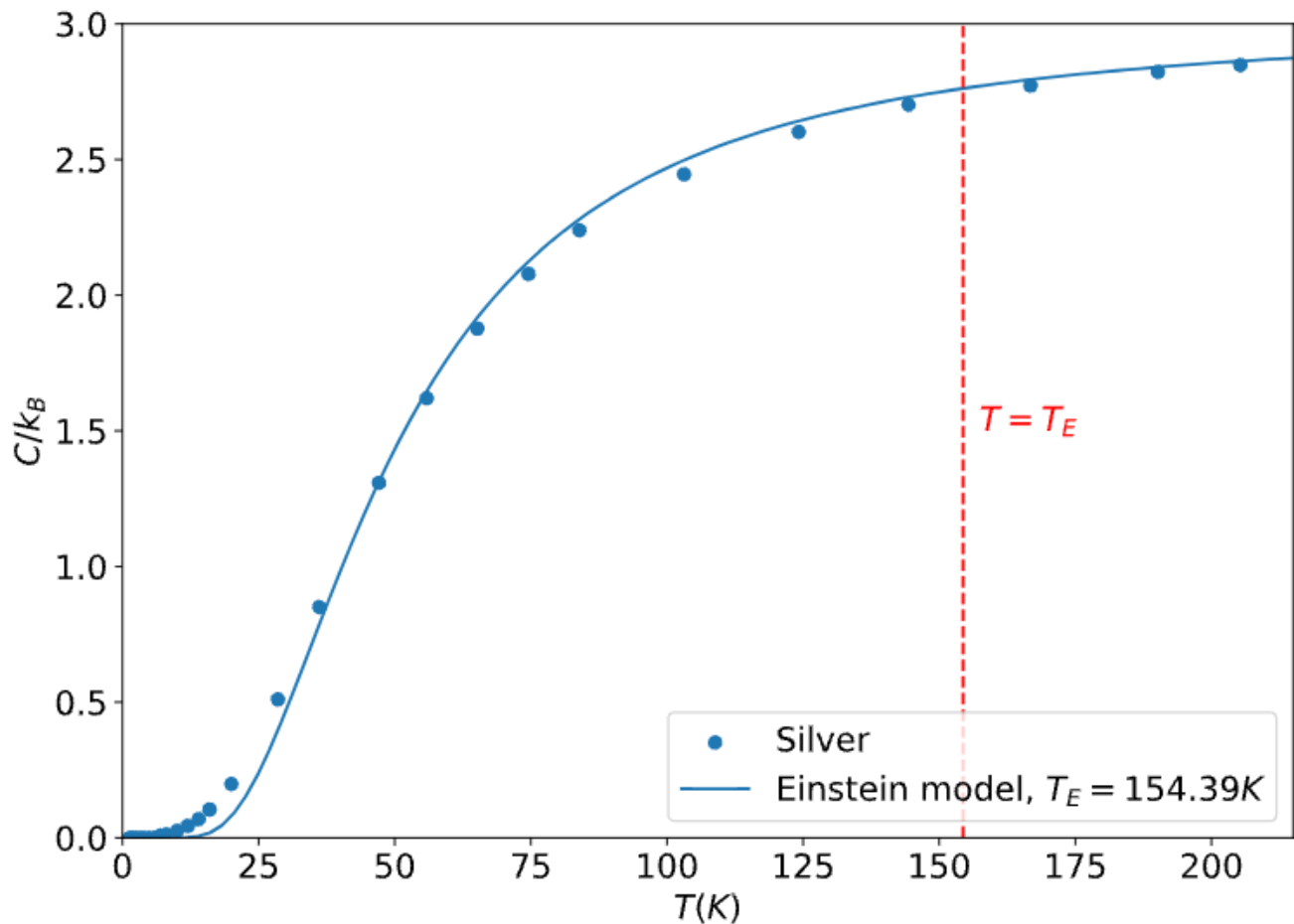
where we introduced the *Einstein temperature*  $T_E \equiv \hbar\omega_0/k_B$ .

The Einstein Temperature is the temperature below which the thermal excitations of the quantum harmonic oscillator start to freeze out. i.e there is not enough thermal energy to excite the harmonic oscillator above the ground state



## Debye Model

## Heat capacity of silver and a fit of the Einstein model



The low-temperature heat capacity of silver is underestimated by the Einstein model.

Debye considered the sound waves in a material as independent harmonic oscillators

## Recap Sound wave

The displacement of an atom at position  $r$  and time  $t$  is described by

$$\delta r = \delta r_0 e^{i(kr - \omega t)}$$

Where  $k = (k_x, k_y, k_z)$  is the wave vector the wavelength is related to the wavevector  $k$  through  $\lambda = 2\pi/|k|$

Thus the wave depends on time only through the factor  $e^{-i\omega t}$  and they are normal modes

A wave has three normal modes

- Two transverse (perpendicular to  $k$ ) and two longitudinal (parallel to  $k$ )

Waves are characterized by their frequency, wavevector and polarization. The frequency is related to the wavevector with the dispersion relation

$$\omega = v_s |k|$$

The quantum excitations of harmonic oscillators are called phonons, the expected number of phonons in the oscillator at temperature  $T$  is given by the Bose-Einstein distribution. We now have  $3N$  oscillators with frequencies depending on  $k$  through the dispersion relation

$$\omega(k) = v_s |k|$$

$$\begin{aligned} E &= 3 \sum_{\mathbf{k}} \left( \frac{1}{2} \hbar \omega(\mathbf{k}) + \hbar \omega(\mathbf{k}) n_B(\beta \hbar \omega(\mathbf{k})) \right) \\ &= 3 \sum_{\mathbf{k}} \left( \frac{1}{2} \hbar \omega(\mathbf{k}) + \frac{\hbar \omega(\mathbf{k})}{e^{\hbar \omega(\mathbf{k})/k_B T} - 1} \right). \end{aligned}$$

## Periodic Boundary Conditions

The heat capacity is a macroscopic property: it doesn't depend on the material's shape just on the volume

Let  $V = L^3$  be a box with periodic boundary conditions

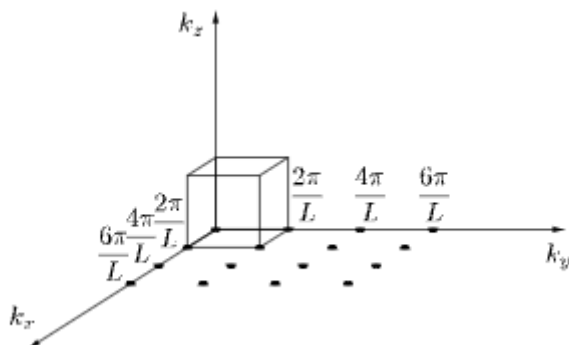
$$\delta r(r + Lx) = \delta r(r)$$

Substituting in the plane wave one gets

$$\delta r_0 e^{i(k \cdot r + k_x L - \omega t)} = \delta r_0 e^{i(kr - \omega t)}$$

In order to satisfy periodic boundary conditions  $k_x = n_x \frac{2\pi}{L}$  damit ergeben sich die erlaubten Werte für  $k$  als

$$k = \frac{2\pi}{L} (n_x, n_y, n_z)$$



There is exactly one allowed  $k$  value per  $(\frac{2\pi}{L})^3$  consider with large box sizes  $L \rightarrow \infty$  that the volume per allowed mode becomes smaller and smaller

$$\sum_k \approx \frac{L^3}{(2\pi)^3} \int dk$$

With this approximation the total energy can be rewritten

$$E = 3 \sum_{\mathbf{k}} \left( \frac{1}{2} \hbar \omega(\mathbf{k}) + \frac{\hbar \omega(\mathbf{k})}{e^{\hbar \omega(\mathbf{k})/k_B T} - 1} \right) \\ \approx 3 \left( \frac{L}{2\pi} \right)^3 \int d\mathbf{k} \left( \frac{1}{2} \hbar \omega(\mathbf{k}) + \frac{\hbar \omega(\mathbf{k})}{e^{\hbar \omega(\mathbf{k})/k_B T} - 1} \right).$$

Because through the dispersion relation we now that  $\omega(k) = v_s |k|$  thus this is only dependent on the magnitude, we can rewrite the integral in spherical coordinates

$$\int d\mathbf{k} \rightarrow \int_0^{2\pi} d\varphi \int_0^\pi \sin(\theta) d\theta \int_0^\infty k^2 dk = 4\pi \int_0^\infty k^2 dk.$$

$$E = 3 \frac{L^3}{(2\pi)^3} \int_0^\infty 4\pi k^2 \left( \frac{1}{2} \hbar \omega(k) + \frac{\hbar \omega(k)}{e^{\hbar \omega(k)/k_B T} - 1} \right) dk \\ = 3 \frac{L^3}{(2\pi)^3} \frac{4\pi}{v_s^3} \int_0^\infty \omega^2 \left( \frac{1}{2} \hbar \omega + \frac{\hbar \omega}{e^{\hbar \omega/k_B T} - 1} \right) d\omega.$$

This integral can be split into two parts, the factor in the brackets describes the average energy of a phonon mode with frequency  $\omega$ . The other factor is the density of states  $g(\omega)$

The density of states  $g(\omega)$  counts the total number of available normal modes inside a frequency interval  $\omega + d\omega$

With this definition, the integral becomes

$$E = \int_0^{\infty} d\omega \left( \frac{1}{2} \hbar \omega + \frac{\hbar \omega}{e^{\hbar \omega / k_B T} - 1} \right) g(\omega),$$

with

$$g(\omega) = 3 \left( \frac{L}{2\pi} \right)^3 \frac{4\pi \omega^2}{v_s^3}.$$

The density of states thus is not dependent of the temperature, one can split the integral now into temperature dependent and independent parts

$$E = E_Z + \frac{3L^3}{2\pi^2 v_s^3} \int_0^{\infty} \left( \frac{\hbar \omega}{e^{\hbar \omega / k_B T} - 1} \right) \omega^2 d\omega.$$

The term  $E_Z$  is the temperature-independent zero-point energy of the phonon modes. Despite  $E_Z$  diverging towards infinity, does not contribute to  $C$ .

The integral depends on the temperature through the  $e^{\hbar \omega / k_B T}$  term. In order evaluate the integral, we substitute  $x \equiv \frac{\hbar \omega}{k_B T}$  and remove the temperature dependence of the integrand:

$$\begin{aligned} E &= E_Z + \frac{3L^3}{2\pi^2 v_s^3} \frac{(k_B T)^4}{\hbar^3} \int_0^{\infty} \frac{x^3}{e^x - 1} dx \\ &= E_Z + \frac{3L^3}{2\pi^2 v_s^3} \frac{(k_B T)^4}{\hbar^3} \frac{\pi^4}{15}, \end{aligned}$$

We can now see by calculating the derivative that

$$C = \frac{dE}{dT} \approx T^3$$

- At temperature  $T$ , only phonon modes with an energy below the thermal energy  $E_T = k_b T$  become thermally excited. These modes have  $\hbar\omega(k) \leq k_b T$

## Conclusions

- The Debye model assumes atoms in a material move in a collective fashion, described by quantized normal modes with the dispersion relation  $\omega = v_s |k|$
- The phonon modes have a constant density of  $(L/2\pi)^3$  in the reciprocal  $k$ -space
- The total energy and heat capacity are obtained via the integration
- The density of states  $g(\omega)$  is the number of states per frequency
- At low temperature the phonon heat capacity is proportional to  $T^3$