

Dynamo Equations

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April 16, 2023

1 Dynamo Equations

Starting with the dimensionless dynamo equations, (5), (6), (7), and (9) as given in Jones [2]:

$$R_1 \left[\frac{\partial \mathbf{u}(r, \theta, \phi, t)}{\partial t} + (\mathbf{u}(r, \theta, \phi, t) \cdot \nabla) \mathbf{u}(r, \theta, \phi, t) \right] = 2\mathbf{u}(r, \theta, \phi, t) \times \hat{z} - \nabla p(r, \theta, \phi, t) + (\nabla \times \mathbf{B}(r, \theta, \phi, t)) \times \mathbf{B}(r, \theta, \phi, t) + E \nabla^2 \mathbf{u}(r, \theta, \phi, t) + R_2(T(r, \theta, \phi, t) \frac{\mathbf{r}}{r_{cmb}}), \quad (1)$$

or we may write

$$R_1 \left[\frac{\partial \mathbf{u}(r, \theta, \phi, t)}{\partial t} + (\mathbf{u}(r, \theta, \phi, t) \cdot \nabla) \mathbf{u}(r, \theta, \phi, t) \right] = \mathbf{F}(r, \theta, \phi, t) - \nabla p(r, \theta, \phi, t)$$

where

$$\mathbf{F}(r, \theta, \phi, t) = 2\mathbf{u}(r, \theta, \phi, t) \times \hat{z} + (\nabla \times \mathbf{B}(r, \theta, \phi, t)) \times \mathbf{B}(r, \theta, \phi, t) + E \nabla^2 \mathbf{u}(r, \theta, \phi, t) + R_2(T(r, \theta, \phi, t) \frac{\mathbf{r}}{r_{cmb}}). \quad (2)$$

To the fluid equation (1) we have an equation for $T(r, \theta, \phi, t)$

$$\frac{\partial T(r, \theta, \phi, t)}{\partial t} + [\mathbf{u}(r, \theta, \phi, t) \cdot \nabla] T(r, \theta, \phi, t) = R_3 \nabla^2 T(r, \theta, \phi, t), \quad (3)$$

and the requirements

$$\nabla \cdot \mathbf{u}(r, \theta, \phi, t) = \nabla \cdot \mathbf{B}(r, \theta, \phi, t) = 0. \quad (4)$$

Using the definition and vector identities

$$\boldsymbol{\omega}(r, \theta, \phi, t) = \nabla \times \mathbf{u}(r, \theta, \phi, t), \quad (5)$$

and

$$\frac{1}{2} \nabla (\mathbf{u}(r, \theta, \phi, t)^2) = (\mathbf{u}(r, \theta, \phi, t) \cdot \nabla) \mathbf{u}(r, \theta, \phi, t) + (\mathbf{u}(r, \theta, \phi, t) \times \boldsymbol{\omega}(r, \theta, \phi, t)), \quad (6)$$

and

$$\nabla \times (\mathbf{u}(r, \theta, \phi, t) \times \boldsymbol{\omega}(r, \theta, \phi, t)) = \boldsymbol{\omega}(r, \theta, \phi, t) \cdot \nabla \mathbf{u}(r, \theta, \phi, t) - \mathbf{u}(r, \theta, \phi, t) \cdot \nabla \boldsymbol{\omega}(r, \theta, \phi, t). \quad (7)$$

This eliminates the pressure term in Eq. (1) which now reads

$$\begin{aligned} \frac{\partial \boldsymbol{\omega}(r, \theta, \phi, t)}{\partial t} + \mathbf{u}(r, \theta, \phi, t) \cdot \nabla \boldsymbol{\omega}(r, \theta, \phi, t) = \\ [\boldsymbol{\omega}(r, \theta, \phi, t) \cdot \nabla] \mathbf{u}(r, \theta, \phi, t) + \nabla \times \mathbf{F}(r, \theta, \phi, t). \end{aligned} \quad (8)$$

Now we have equations for $\mathbf{u}(r, \theta, \phi, t)$ and $\mathbf{B}(r, \theta, \phi, t)$ and $T(r, \theta, \phi, t)$ where the pressure term has been eliminated.

R_1 , R_2 , and R_3 are dimensionless constants given in [2].

2 Representing a Solenoidal Vector Field

In [1] it is shown that if $\nabla \cdot A(r, \theta, \phi) = 0$, one may write

$$A(r, \theta, \phi) = \nabla \times (\boldsymbol{\Lambda} a(r, \theta, \phi)) + \boldsymbol{\Lambda} b(r, \theta, \phi), \quad (9)$$

where we have the operator

$$\boldsymbol{\Lambda} = \mathbf{r} \times \nabla. \quad (10)$$

This yields

$$\mathbf{r} \cdot A(r, \theta, \phi) = \boldsymbol{\Lambda}^2 a(r, \theta, \phi), \quad (11)$$

and

$$\boldsymbol{\Lambda} \cdot A(r, \theta, \phi) = \boldsymbol{\Lambda}^2 b(r, \theta, \phi) \quad (12)$$

By inverting the operator $\boldsymbol{\Lambda}^2$, we may solve for $a(r, \theta, \phi)$ and $b(r, \theta, \phi)$.

If we have $\boldsymbol{\Lambda}^2 f(\Omega) = g(\Omega)$, then

$$\begin{aligned} f(\Omega) &= \frac{1}{\boldsymbol{\Lambda}^2} g(\Omega) \\ f(\Omega) &= \frac{1}{4\pi} \int d\Omega' g(\Omega') \log_e(1 - \Omega \cdot \Omega'). \end{aligned} \quad (13)$$

In this Ω is a vector of length 1 on the surface of a 2D-sphere.

$$\begin{aligned} a(r, \theta, \phi) &= \frac{1}{\boldsymbol{\Lambda}^2} (\mathbf{r} \cdot A(r, \theta, \phi)) \\ a(r, \theta, \phi) &= \frac{1}{4\pi} \int d\Omega' (\mathbf{r} \cdot A(r, \theta', \phi')) \log_e(1 - \Omega \cdot \Omega'). \\ b(r, \theta, \phi) &= \frac{1}{4\pi} \int d\Omega' (\boldsymbol{\Lambda} \cdot A(r, \theta', \phi')) \log_e(1 - \Omega \cdot \Omega') \end{aligned} \quad (14)$$

References

- [1] George Backus. A class of self-sustaining dissipative spherical dynamics. *Annals of Physics*, 4:372–447, 1958.
- [2] Chris A. Jones. Planetary magnetic fields and fluid dynamos. *Annual Reviews of Fluid Mechanics*, 43:583–614.