

Chapter 1

Size Spectrum Theory

In this chapter we consider the models in [M’Kendrick \(1925\)](#) and [Foerster \(1959\)](#) who describe a system indexed by age. We then consider of the advances in [Silvert and Platt \(1978\)](#) who describe new equations indexed by weight and the derive the standard McKendrick-von Foerster Equation for weight indexed systems.

This work serves as the basis for [Datta et al. \(2010\)](#), who constructs a stochastic model for the dynamics of population, which they call the deterministic jump-growth equation. We consider the relevance of this equation in relation to the McKendrick-von Foerster Equation drawing further on the work of [Datta et al. \(2010\)](#).

1.1 Age Indexed Models

[M’Kendrick \(1925\)](#) first posed the idea of modeling biological processes for medicinal science by a single characteristic. In a process of individuals meet they transfer information within the system, and if one considers these as individuals as particles in a system moving according to a dimension indexed by this single characteristic then their movement becomes a study in kinetics.

[Foerster \(1959\)](#) extended the principle equations derived in [M’Kendrick \(1925\)](#) with extensions. [Trucco](#) gives a full rigorous discussion of the advancements made in [Foerster \(1959\)](#). It’s discussion involve considering the steady state solutions of what he calls “The Von Foerster Equation”.

1.1.1 The Von Foerster Equation

The Von Foerster Equation was a major extension to the work of [M’Kendrick \(1925\)](#). It defines the standard in age indexed population models and reads

$$\frac{\partial n}{\partial t} + \frac{\partial n}{\partial a} = -m(a)n. \quad (1.1)$$

for a mortality function dependent on age $m(a)$. Following the work of [Trucco](#), who gives a good derivation of the equation, we now describe Von Foerster's reasoning to determine [Equation 1.1](#).

Suppose that $n(t, a)$ represents the density of individuals at time t in the age category $(a, a + \Delta a)$. Then we have that

$$\begin{aligned} \frac{\partial}{\partial t} (n(a, t)\Delta a) &= + \text{rate of entry of } a \\ &= - \text{rate of departure at } (a + \Delta a) \\ &= - \text{deaths in } (a, a + \Delta a). \end{aligned} \quad (1.2)$$

We can express which in mathematical terms for some flux, $J(t, a)$, which describes that rate of movement of individuals in $(a, a + \Delta a)$ as

$$\frac{\partial n}{\partial t} = \frac{J(t, a) - J(t, a + \Delta a)}{\Delta a} - m(a)n(a, t). \quad (1.3)$$

for some per capita mortality rate m . When dealing with the flux J we consider that this represents the movement of individuals in age. As individuals become older the flux can be assumed to be proportional to the density of individuals with some velocity $v(t, a)$. If the aging corresponds to the parsing of time then we have that

$$v = \frac{\partial a}{\partial t} = 1$$

and so $J(t, a) = n(t, a)$. Substituting this in it is clear that, in the limit as $\Delta a \rightarrow 0$, [Equation 1.3](#) becomes the [Equation 1.1](#).

1.2 The Transport Equation

After we consider [Equation 1.1](#) we consider a more general form of equation: transport, or convection-diffusion, Equations take the general form

$$\frac{\partial u}{\partial t} = \nabla \cdot (D \nabla u) - \nabla \cdot (\vec{v} u) + R \quad (1.4)$$

and describe particles undergoing diffusion and convection. In one dimension [Equation 1.4](#) reduces to

$$\frac{\partial u}{\partial t} = -vu + R + (Du_x)_x. \quad (1.5)$$

Comparing the coefficient definitions from [Stocker \(2011\)](#) to the ideas from the derivation of [Equation 1.1](#) we take u as the quantity of interest, population (density) and then we consider the coefficients D and v and parameter R .

The first term $-vu$ describes the convection (movement due to mass) in the system. v describes the velocity of that the population is moving at, this is analogous to the growth of the individuals in the population the rate at which they grow and move through the weight range.

The second term R describes the creation or destruction of the quantity. Thus in [Equation 1.1](#) this becomes the death (naturally of predatorily) of the population.

Lastly we have the diffusion term $(Du_x)_x$. Imagine that c is the concentration of a chemical. When concentration is low somewhere compared to the surrounding areas (e.g. a local minimum of concentration), the substance will diffuse in from the surroundings, so the concentration will increase. Conversely, if concentration is high compared to the surroundings (e.g. a local maximum of concentration), then the substance will diffuse out and the concentration will decrease. This is analogous to phenomenon that are exhibited in the Jump-Growth Equation which we talk about in the next section.

1.3 Weight Indexed Models

1.3.1 McKendrick-von Foerster Equation

[Silvert and Platt \(1978\)](#) introduced a more general construction of the McKendrick-von Foerster Equation, notwithstanding a change from age indexed population to sized based population, in a model which allowed growth and mortality to be functions of body mass. Their changes are widely used in mathematical biology and a full derivation can be found in [Silvert and Platt \(1978\)](#), however a simple argument is that the flux described in [Equation 1.3](#) is changed for a flux that depends on the growth of individuals

$$J(t, w) = g(t, w)n(t, w) \quad (1.6)$$

and $m(a)$ becomes a mortality function in weight and age $\mu(w)$. Thus the McKendrick-von Foerster Equation reads

$$\frac{\partial n}{\partial t} + \frac{\partial}{\partial w} (g \cdot n) = -\mu \cdot n. \quad (1.7)$$

Considering this equation in the context of [Section 1.2](#) we can see that the [Equation 1.7](#) is a transport equation with no diffusion term. While a transport equation is defined on the region of $\mathbb{R}^+ \times \mathbb{R}$, simply requiring an initial condition, the Von Foerster Equation is only defined on $\mathbb{R}^+ \times \mathbb{R}^+$, thus requiring a left boundary condition along the line $(t, 0)$. Generally this boundary condition is defined as the births across the population which could simply be defined with

$$u(t, 0) = 0 \quad (1.8)$$

however this would be an uninteresting problem since the system would only include growth and death, but no birth. Thus instead of [Equation 1.8](#) we introduce a birth rate $b(a)$ and integrate over the population to give a boundary condition

$$u(t, 0) = \int_0^\infty b(w)u(t, w) dt. \quad (1.9)$$

Of course numerically we would introduce a right hand side boundary since we cannot handle the domain $[0, \infty)$, but we will deal with this in due course.

1.3.2 Jump Growth Equation

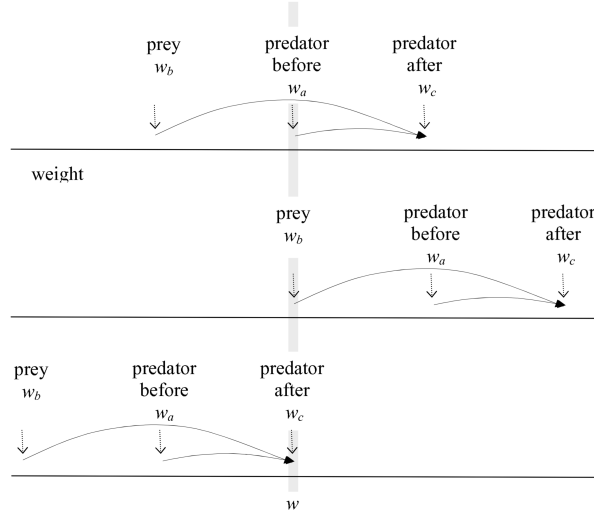
While the mathematical framework presented in [Silvert and Platt \(1978\)](#), [Datta et al. \(2010\)](#) presents a different method of describing weight indexed population under the assumption that predation is a Markov process. The predation event extends the ideas of [Silvert and Platt \(1980\)](#) where predation is modeled the coupled death at one size to the growth of another. This is illustrated in [Figure 1.1](#) where we see there are three types of predation to yield new individuals weight w_c .

Resolving the macroscopic behavior of these predation events yields the deterministic equation

$$\frac{\partial \varphi_i}{\partial t} = \sum_j (-k_{ij}\varphi_i\varphi_j - k_{ji}\varphi_j\varphi_i + k_{mj}\varphi_m\varphi_j), \quad (1.10)$$

where m is the index satisfying the weight bracket $w_m \leq w_i - Kw_j < w_{m+1}$ and k_{xy} is the indexed rate of predation events, predator before prey. [Datta et al. \(2010\)](#) explains that these three terms correspond to the three types of predation event seen in [Figure 1.1](#): losses from bracket i (negative terms) occur because individuals in this bracket eat prey and become heavier, and because these individuals are themselves eaten. Gains into weight bracket i (the positive term) occur through smaller predators growing into this bracket by eating prey.

Figure 1.1: Figure 1 from [Datta et al. \(2010\)](#). The death of an individual with weight w_b is attributed to the growth of another individual weight w_a , who consequently dies to yield a new individual with weight $w_c = w_a + Kw_b$ for some predation efficiency K that could be a function of weight.



We can create an analytic equation for use in calculations by taking the continuum limit of [Equation 1.10](#) and using a continuous feeding rate function $k(w, w')$ to replace the rate constants k_{xy} . Thus [Equation 1.10](#) becomes

$$\begin{aligned} \frac{\partial \phi(w)}{\partial t} = & \int (- k(w', w) \phi(w) \phi(w') \\ & - k(w, w') \phi(w') \phi(w) \\ & + k(w - Kw', w') \phi(w - Kw') \phi(w')) dw'. \end{aligned} \quad (1.11)$$

[Datta et al. \(2010\)](#) labels this the “Deterministic Jump-Growth Equation” and again the three terms represent the methods of transferring weight within in the system through predation corresponding to the methods in [Figure 1.1](#): Predation on prey to become larger, being predated upon and being fed with the correct weight to increase.

1.3.3 McKendrick-von Foerster Equation with Diffusion

The McKendrick-von Foerster Equation can be shown to be an approximation to the Jump-Growth Equation, but only to first order. A second order approximation can be yielded by including a diffusion term. Consider the third term in the Jump-Growth Equation as a Taylor Series:

$$\begin{aligned} k(w - Kw', w') \phi(w - Kw') = & k(w, w') \phi(w) \\ & + (-Kw') \frac{\partial}{\partial w} (k(w, w') \phi(w') \phi(w)) \end{aligned}$$

$$+ \frac{1}{2}(-Kw')^2 \frac{\partial^2}{\partial w^2} (k(w, w')\varphi(w)) + \dots \quad (1.12)$$

Once this is substituted back in to [Equation 1.11](#) gives

$$\begin{aligned} \frac{\partial \phi(w)}{\partial t} = & \int -k(w', w)\phi(w)\phi(w') \, dw' \\ & - \frac{\partial}{\partial w} \int Kw'k(w, w')\varphi(w)\varphi(w') \, dw' \\ & + \frac{1}{2} \frac{\partial^2}{\partial w^2} \int (Kw')^2 k(w, w')\varphi(w)\varphi(w') \, dw' + \dots \end{aligned} \quad (1.13)$$

Which can obviously be seen as an approximation to the McKendrick-von Foerster Equation, with a diffusion term added. Thus for an appropriate choice of coefficients we can restrict our attention to equations in a Transport Equation form

$$\frac{\partial u}{\partial t} = -(g \cdot u)_w - \mu \cdot u + \frac{1}{2}(Du)_{xx}. \quad (1.14)$$

1.4 Summary

In this chapter we have introduced the mathematical framework that will be used to model population and shown that the macroscopic stochastic Jump-Growth equation is comparable to the Transport Equation forms of the McKendrick-von Foerster Equation. In [Chapter 2](#) we introduce a numerical construction of the coefficients giving arise to the true models that will be solved in [Chapter 4](#).

Chapter 2

Model Equations

In this chapter we build on [Chapter 1](#) to construct a fully derived equation to model in [Chapter 4](#). We begin by considering numerical interpretations of feeding kernels and prey selection then construct the various predation coefficients for [Equation 1.13](#).

2.1 Feeding Kernel

Central to the models that we are studying is the idea of a feeding kernel which is the rate at which a predator eats prey of a different size. Centralized around the idea that ‘big fish eat smaller fish’ we assume that make predators select their prey based around their own size and thus have a preferred size that they primarily seek out and thus the preference function must be based only on the weight of the individual and their prey.

Taking this key idea of prey size selection we consider the work of [Benoit and Rochet \(2004\)](#) who showed that the feeding kernel for an individual takes the form

$$k(w, w') = sw^\alpha S\left(\frac{w}{w'}\right) \quad (2.1)$$

for some parameters s, α to be determined and for the size selection function S . This says that the rate at which a predator of weight w feeds is a product of the volumetric search rate sw^α and the feeding preference function S . In this thesis we make the assumption that $s = 1$ for convenience.

2.1.1 Prey Selection

As discussed the prey selection will peak around some fixed ratio w/w' for predator / prey, which we call β . Assuming that there is some spread for which individuals will vary from this

ratio, noted as σ then we can say that the prey selection function S will take the form

$$S(z) = \kappa \exp\left(-\frac{(z - \beta)^2}{2\sigma^2}\right) \quad (2.2)$$

for some κ that we discuss later.

2.1.2 Coefficient Functions

When we consider the form of the McKendrick-von Foerster in it's transport equation form we can extract coefficients. Taking [Equation 1.13](#) in the form of [Equation 1.14](#) it reads

$$\begin{aligned} \frac{\partial u}{\partial t} = & - \left(\left(\int (Kw')k(w, w')u(w') \, dw' \right) \cdot u \right)_w \\ & - \left(\int k(w', w)u(w') \, dw' \right) \cdot u \\ & + \frac{1}{2} \left(\left(\int (Kw')^2 k(w, w')u(w') \, dw' \right) \cdot u \right)_{xx}. \end{aligned} \quad (2.3)$$

Consider the growth term in [Equation 2.3](#),

$$g(w) = \int (Kw')k(w, w')u(w'). \quad (2.4)$$

One can think of this as the expected value of the mass eaten by individuals of weight w . Thus we have that the prey selection function can be thought of as a probability distribution and so we take

$$\kappa = \frac{1}{\sigma\sqrt{2\pi}}.$$

Thus, if we take [Equation 2.4](#) and let $v = g$, and define

$$R = \mu(w) = \int k(w', w)u(w') \, dw' \quad (2.5)$$

$$D(w) = \int (Kw')^2 k(w, w')u(w') \, dw' \quad (2.6)$$

then we have [Equation 1.13](#) in the form of [Equation 1.5](#).

2.2 Steady State Solutions

In marine ecosystems it's been found that the abundance of organisms within weight classes is roughly constant ([Sheldon et al. \(1972\)](#)) if those weight classes are distributed

logarithmically. Further, averaged over time, [Datta et al. \(2011\)](#) observes that this abundance changes rather little suggesting that the system is near steady state.

[Benoit and Rochet \(2004\)](#) found that the McKendrick-von Foerster Equation has a power law steady state of the form $\phi(w) \propto w^{-\gamma}$, for $\gamma \in \mathbb{R}^+$. We note though that in the real world such a power law steady state must break down, since it predicts an infinite number of individuals with weight $w = 0$.

Combining this research to now show that the [Equation 2.3](#) can be transformed using the logarithmic change of variables and appropriate change of function to have a constant solution as [Benoit and Rochet \(2004\)](#) discusses.

2.2.1 Logarithmic Scale

[Sheldon and Parsons \(1967\)](#), who coincidentally first introduces the idea of a size spectrum to organize counts of particles in the ocean, stipulates the fundamental conjecture that the total mass within logarithmically spaced size groups was constant over the size range from bacteria to whales.

Interpreting this conjecture biologically [Silvert and Platt \(1980\)](#) poses that the total mass of prey is the same as the mass of its predators, but can also be written in terms of the number spectrum, where the number of individuals with size m per volume is described by a density function $N(w) \propto w^{-\gamma}$ (the same as the work of [Benoit and Rochet \(2004\)](#)).

Since these conjectures assume a logarithmic weight scale, we now perform a change of variables from [Equation 2.3](#) using the dimensionless variable $x = \log w$ assuming that $\phi(x)dx = u(w)dw$ (a more comprehensive derivation can be found in [Appendix A](#)). Following [Equation 1.14](#) that

$$u_t = -(g \cdot u)_w - \mu \cdot u + \frac{1}{2}(D \cdot u)_{ww} \quad (2.7)$$

then after changing variables using a logarithmic weight scale we gain a new equation for ϕ reading

$$\phi_t = -(\hat{g} \cdot \phi)_x - \hat{\mu} \cdot \phi + \frac{e^{-x}}{2} \left(\left(\hat{D} \cdot \phi \right)_{xx} - \left(\hat{D} \cdot \phi \right)_x \right). \quad (2.8)$$

However if we are to maintain a constant steady state as stipulated by [Sheldon and Parsons \(1967\)](#) then we must perform a simple change of function. The steady state $u(w) \propto w^\gamma$ reduces to $\phi(x) \propto e^{x(\gamma+1)}$, which is clearly not constant. However if we reduce to a constant by taking $e^{x(\gamma+1)} \rho(x) = \phi(x)$ then we have our constant but ρ is governed by the more complicated equation

$$e^{x(\gamma+1)} \rho_t = -(\bar{g} \cdot \rho) - \bar{\mu} e^{x(\gamma+1)} \rho + \frac{e^{-x}}{2} \left((\bar{D} \cdot \rho)_{xx} - (\bar{D} \cdot \rho)_x \right). \quad (2.9)$$

2.2.2 Constant Steady State

2.3 Logarithmic Coefficients

2.4 Summary

In this chapter we have outlined the realized equations that will be used in [Chapter 4](#). Next we will discuss some of the methods that we can use to approach these equations analytically and numerically.

Chapter 3

Mathematics of Discretization

The major benefit of studying equations which hold the form of a Transport equation is the large body of research of their solutions and behavior. We begin this chapter by considering the most basic transport equation, the Advection Equation which will allow the introduction of finite difference techniques.

We then follow this through with a discussion of the approach used by [Hartvig et al. \(2011\)](#) on the food web framework that they have developed using semi-implicit methods.

3.1 Advection Equation

The Advection Equation is a simple hyperbolic partial differential equation, in one dimension we take

$$\frac{\partial u}{\partial t} + v \frac{\partial u}{\partial x} = du, \tag{3.1}$$

for some transfer velocity v and destruction rate d on the domain $\mathbb{R}^+ \times R$ for some $R \subseteq \mathbb{R}$. The linear advection has been studied extensively and thus we can gain some insight in to potential issues that might occur in our more complicated equations such as [Equation 2.8](#).

3.1.1 Analytic Solution

As with ordinary differential equations, partial differential equations can be solved with the “method of characteristics” (or method of lines). Say we consider [Equation 3.1](#) on the domain $\mathbb{R}^+ \times \mathbb{R}^+$, as to line up with the McKendrick-von Foerster Equation. Then we define some boundary value and initial value problem (BVP and IVP):

$$\begin{aligned}
u_t + v u_x &= +d u \\
u(0, x) &= I(x) \\
u(t, 0) &= B(x)
\end{aligned} \tag{3.2}$$

Using the method of characteristics we consider a point in the domain $\{(t, x) : t, x > 0\}$ and solve the equation along some characteristic line $(t(s), x(s)) = y(s)$ stemming from an $y_0 \in \Gamma = \{(0, x) : x \in \mathbb{R}^+\} \cup \{(t, 0) : t \in \mathbb{R}^+\}$. The full details of the method are left to the reader however we find that

$$u(t, x) = \begin{cases} I(x - vt) e^{-dt} & x - vt > 0 \\ B(x - vt) e^{-dt} & x - vt \leq 0 \end{cases} \tag{3.3}$$

along any characteristic $x = vt$. We note that the McKendrick-von Foerster will take a form similar to this if the growth and death terms are constants, but the problem will become more complicated if the coefficients are functions of t, x or non-linear. However this problem does allow us to introduce the idea of domain of dependence.

Definition 3.1 (Analytic Domain of dependence). Given any BVP and/or IVP on $U \subseteq \mathbb{R}^n$ the with solution φ , the domain of dependence for for any $x \in U$ is the set $V \subseteq U$ that $\varphi(x)$ depends on to be calculated.

[Figure 3.1](#) shows the characteristic lines for, the blue lines represent characteristics which depend on the time axis boundary condition while the red arrows depend on the spatial axis initial condition. Looking at the analytic solution to [Equation 3.2](#) we see that the initial condition is carried along the characteristic.

While it is not obvious how this definition will be useful now, it will become extremely important in the analysis of stability of the numerical methods that we discuss in the next section.

3.1.2 Numerical Solutions to the Advection Equation

Suppose now that we want to solve the advection numerically. Writing the equation using limits we have that

$$\lim_{\Delta t \rightarrow 0} \frac{u(t + \Delta t, x) - u(t, x)}{\Delta t} + v \lim_{\Delta x \rightarrow 0} \frac{u(t, x + \Delta x) - u(t, x)}{\Delta x} = du(t, x). \tag{3.4}$$

However what if we simply ignored the limits and used Newtonian infinitesimals $\Delta t, \Delta x$? This would give us the equation

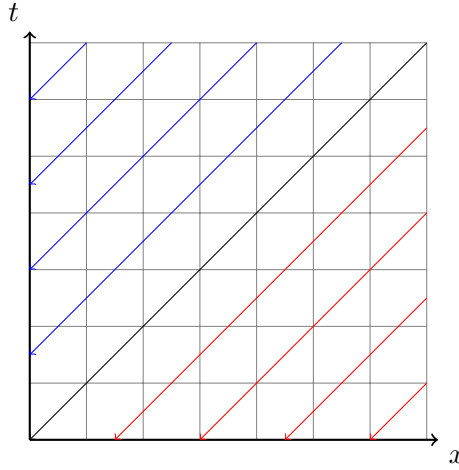


Figure 3.1: The characteristic lines for the solution of Equation 3.3 with constant coefficients. There is no overlap between the lines which means that the solution can be smooth if the boundary conditions along Γ are smooth. The domain of dependence for traverses the direction of each arrowhead line.

$$\frac{u(t + \Delta t, x) - u(t, x)}{\Delta t} + v \frac{u(t, x + \Delta x) - u(t, x)}{\Delta x} = du(t, x) \quad (3.5)$$

which can be rearranged to solve for $u(t + \Delta t, x)$. This would give a discrete way of calculating a next time step of the solution to the equation ($u(t + \Delta t, R)$) if we knew the solution $u(t, R)$. This forms the fundamental idea for finite difference equations. These equations approximate a problem by considering a discrete mesh on the domain and then solving the equation using a consistent (see Subsection 3.3.2) difference equations.

3.2 Difference Equations

In general, most partial differential equations do not have explicit analytic solutions. Consequently we must find accurate numerical approximations and methods for solving these problems. In this section we introduce the required mathematics for understanding the numerical methods used to solve problems in partial differential equation theory.

Boole and Moulton (1880) writes “Differential Calculus is occupied about the limits to which such ratios approach as the increments are indefinitely diminished.” However, “The Calculus of Finite Differences may be strictly defined as the science which is occupied about the ratios of the simultaneous increments of quantities mutually dependent.” In simple terms the calculus of finite differences is only concerned about ratios of infinitesimals, in line with the ideas of Newtonian calculus.

In Equation 3.4 we broke down the limits in the derivatives by simply ignoring the limit, in

this section we reconstruct this equation using difference equations derived from [Boole and Moulton \(1880\)](#) and [Hildebrand \(1987\)](#).

Definition 3.2 (First Forward Difference). For a function $u : U \rightarrow \mathbb{R}$ the first forwards difference of u is

$$\Delta u = u(x + \Delta x) - u(x). \quad (3.6)$$

Definition 3.3 (First Backwards Difference). For a function $u : U \rightarrow \mathbb{R}$ the first backwards difference of u is

$$\Delta u = u(x) - u(x - \Delta x). \quad (3.7)$$

These two definitions form the basis for finite difference theory. Just as with the differential operator we have that it is linear and satisfies the Leibniz rule. The proof of this is left to the reader, but follows exactly the same method as for derivatives. It is easy to construct the n^{th} forward difference $(\Delta^n[f](x))$ by simply considering $\Delta[\Delta^{n-1}[f]](x)$.

When we consider the partial derivative of $u : \mathbb{R} \rightarrow \mathbb{R}$ we simply have that

$$u'(x) = \lim_{\Delta x \rightarrow 0} \frac{\Delta[u](x)}{\Delta x} \quad (3.8)$$

which can easily be extended to higher dimensions. Using Taylor's theorem, we know that

$$u(x + \Delta x) = u(x) + \Delta x u'(x) + O(x) \quad (3.9)$$

and so we can take the difference equation

$$\frac{\Delta[u](x)}{\Delta x} = u'(x) + O(\Delta x) \quad (3.10)$$

as a first order approximation to a first derivative of u . By combining n^{th} forward and backward differences we can construct higher order approximations to derivatives at higher orders and higher orders of accuracy.

3.3 Finite Difference Equation Errors

Since finite difference equations have become such a staple in numerical analysis there are many texts that supply the different coefficients required to approximate derivatives

accurately. [Fornberg \(1988\)](#) provides a details derivation and set of tables for almost all derivatives required in general PDE problems up to high orders in both accuracy and order.

However as with all numerical problems issues arise when constructing approximations due to errors that occur with the generation of those approximations themselves. In this section we begin to outline the two main types of error that occur in approximations and introduce the idea of consistency and convergence.

3.3.1 Notation

Before we begin talking about the errors generated by finite difference equations we begin by introducing the notation that will be used going forward. Suppose that we are given a partial differential equation of the form

$$L(u) = \frac{\partial u}{\partial t} - \mathcal{L}u \quad (3.11)$$

for some linear differential operator \mathcal{L} which is a function $\mathcal{L}(x, u, x_x, \dots, u_{x^m})$ for some $m \in \mathbb{N}$, that is that $\mathcal{L}, L : \mathcal{C}^m \rightarrow \mathcal{C}$. We consider L as a function of continuous differentiable functions with root solution Ψ , that is $L(\Psi) = 0$.

Given a BVP and/or IVP with L as the operator on some domain $D \subseteq \mathbb{R}^+ \times R$, for $R \subseteq \mathbb{R}$ then we discretize the the domain in to a mesh-grid using a partition on some finite subdomain of time, that is the mesh-grid M represents the continuous domain $[0, T] \times R \subset D$. We index the time steps by n using a $\Delta t = h$ and the spatial steps by j using $\Delta x = k$. Using these indexes we can notate any value of some function on the mesh grid as

$$u(t, x) = u(n\Delta t, j\Delta x) = u(n \cdot h, j \cdot k) = u_j^n. \quad (3.12)$$

3.3.2 Consistency

When considering the appropriateness of a finite difference equation to model a particular problem there are three problems that arise, the most easy of these to understand is consistency. The broad definition of consistency is that a finite difference equation D is consistent with a partial differential equation L if the truncation error of D from L tends to 0 for smaller mesh grid approximations.

To formally introduce consistency consider a PDE in them form of [Equation 3.11](#) with a true solution Φ . Suppose that we take any finite difference equation $D : \mathcal{C}^m \rightarrow \mathcal{C}$ to approximate [Equation 3.11](#) and let ϕ be the solution to $D(u) = 0$.

Definition 3.4 (Truncation Error). For any problem described above the Truncation error at any grid point (n, j) is given as

$$T_j^n(\psi) = D(\psi_j^n) - L(\psi_j^n). \quad (3.13)$$

If Φ , the true solution of L is known then the local truncation error τ_j^n is defined as $T_j^n(\Phi)$.

If $\tau_j^n = O(h^p)$ then D is said to be an order p method, and in reality most finite difference equations are of order p for some $p \in \mathbb{N}$. However this is not always the case and this is why we rely on consistency in a numerical calculation. T_j^n gives an estimate of the error of replacing $L(v_j^n)$ by $F(v_j^n)$. This gives arise to a formal definition of consistency.

Definition 3.5 (Consistency). A finite difference equation D is consistent to a partial differential equation L as described above. Then if v is any function, with a sufficient number of continuous derivatives such that $L(v_j^n)$ can be evaluated, D is considered consistent with L if

$$\lim_{k, h \rightarrow 0} T_j^n(v) = \lim_{k, h \rightarrow 0} F(v_j^n) - L(v_j^n) \rightarrow 0. \quad (3.14)$$

If $v = \Phi$ then $T_j^n(v) = \tau_j^n$ and the truncation error coincides with the local truncation error. So another definition of consistency is that the limiting value of the local truncation error is 0. Suppose that we review [Equation 3.5](#) as an approximation to [Equation 3.1](#), having engineered the scheme using [Fornberg \(1988\)](#) we know that rearranging to gain an equation in the form $D(u) = 0$ the truncation error can be written as

$$T(v) = D(v) - L(v) = O(h) + vO(k) - dO(1) = O(h) + O(k) \quad (3.15)$$

and thus as $h, k \rightarrow 0$ then $T(v) \rightarrow 0$ and we have a consistent difference equation of order $p = 1$.

3.3.3 Stability

Numerical stability theory stems more directly from analytical stability theory. Suppose that $L(u)$ is a PDE problem with two analytic solutions Ψ_1 and Ψ_2 , one can analyze how likely a perturbation in the initial condition is to affect the system as it tends to one of the two analytic solutions. This same principle applies to numerical methods. The essential idea defining stability is that the numerical process should not cause any small perturbations introduced through rounding at any stage to grow and ultimately dominate the solution, similar to the study of whether a small perturbation introduced into an analytic solution will cause the solution to tend to a dominant solution.

Definition 3.6 (Numerical Stability). A numerical method is stable if for every $T > 0$, there exists a constant $c(T) > 0$ such that

$$||\phi_n||_h < c(T) \tag{3.16}$$

for all n , as $h, k \rightarrow 0$. Where ϕ_n represents the vector (ϕ_j)

3.4 Convergence Theory

3.4.1 Linear Theory

3.4.2 Non-Linear Theory

3.5 Conclusion

Chapter 4

Method

Appendix A

Logarithmic Variables

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