

# Attitude Dynamics and Control of a Nano-Satellite Orbiting Mars

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**This project for ASEN5010 Spacecraft Dynamics and Control considers a small satellite orbiting Mars at a low altitude. This spacecraft gathers science data and transfers this data to another satellite orbiting at a higher altitude. Periodically, this spacecraft must transition from nadir-pointing, science gathering mode to sun-pointing mode to recharge the battery system. The three missions goals are nadir-pointing, communicating with the mother spacecraft, and to sun-point. Both of these spacecraft are in circular orbits.**

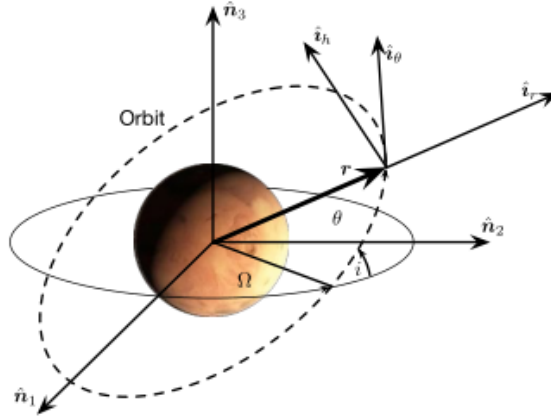
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## Introduction

### Problem Statement

Let us begin with defining the orbit of the nano-satellite with the following figure



**Figure 1: Illustration of the Inertial, Hill, and perifocal geometrical constructions. Taken from ASEN5010 Semester Project sheet.**

### Initial Conditions and Problem Setup

Before we begin with our work on the simulation and controller, we must define our initial conditions and operational parameters.

$$\sigma_{B/N}(t_0) = \begin{bmatrix} 0.3 \\ -0.4 \\ 0.5 \end{bmatrix} \quad (1)$$

$${}^B\omega_{B/N}(t_0) = \begin{bmatrix} 1 \\ 1.75 \\ -2.2 \end{bmatrix} \text{ deg/s} \quad (2)$$

$${}^B[I] = \begin{bmatrix} 10 & 0 & 0 \\ 0 & 5 & 0 \\ 0 & 0 & 7.5 \end{bmatrix} \text{ kg m}^2 \quad (3)$$

We will also use the following Keplerian orbital elements for both spacecraft:

Spacecraft	$r$	$\Omega$	$i$	$\theta(t_0)$
LMO	3796.19 km	20°	30°	60°
GMO	20424.2 km	0°	0°	250°

Additionally, we use a gravitational parameter  $\mu = 42828.3 \text{ km}^3/\text{s}^2$ . We can also calculate each rotation rate using the circular motion assumption of  $\dot{\theta} = \sqrt{\frac{\mu}{r^3}}$ . These parameters will be used for all of the tasks and intermediate simulations in this project.

### Task 1: Orbit Simulation

Our Hill frame is defined by the basis:  $\{\hat{i}_r, \hat{i}_\theta, \hat{i}_h\}$  with the inertial defined as  $\{\hat{n}_1, \hat{n}_2, \hat{n}_3\}$ . Given the inertial and Hill frame definitions, we know that the position vector of the LMO satellite is  $r\hat{i}_r$ . Additionally we know that since it is a circular orbit, it has a time invariant angular rate  $\omega_{H/N} = \dot{\theta}\hat{i}_h$ . Calculating the vectorial inertial derivative:

$$\dot{\mathbf{r}} = \frac{{}^N d}{dt} \mathbf{r} = \frac{{}^H d}{dt} \mathbf{r} + \boldsymbol{\omega}_{H/N} \times \mathbf{r} \quad (4)$$

$$= \dot{\theta} \hat{\mathbf{i}}_h \times r \hat{\mathbf{i}}_r \quad (5)$$

$$= r \dot{\theta} \hat{\mathbf{i}}_\theta \quad (6)$$

Additionally, we can use this information to find the inertial position and velocity vectors by performing transformations using the perifocal frame information. We know that the perifocal frame can be defined by an Euler 3-1-3 rotation defined the set  $\{\Omega, i, \theta\}$

$$C_{ECI} = \begin{bmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos i & \sin i \\ 0 & -\sin i & \cos i \end{bmatrix} \begin{bmatrix} \cos \Omega & \sin \Omega & 0 \\ -\sin \Omega & \cos \Omega & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad (7)$$

Which describes a rotation from Earth Centered Inertial frame. Each portion of the DCM is a single-axis rotation. We can then use this to project scalar values in the Hill frame to inertial vectors with the following:

$${}^N \vec{\mathbf{r}} = C_{ECI}^T \begin{bmatrix} r \\ 0 \\ 0 \end{bmatrix} \quad (8)$$

$${}^N \vec{\mathbf{v}} = C_{ECI}^T \begin{bmatrix} 0 \\ r \dot{\theta} \\ 0 \end{bmatrix} \quad (9)$$

When the ECI direction cosine matrix is calculated,  $\theta$  must be propagated over time, as the true anomaly is the only perifocal parameter that is time variant. It is calculated as such:  $\theta = \theta_0 + t * \dot{\theta}$ .

## Task 2: Orbit Frame Orientation

It is simple to generate bases vectors for the Hill frame, under motion, using our new inertial vectors. As stated before,  $\mathcal{H} = \{\hat{\mathbf{i}}_r, \hat{\mathbf{i}}_\theta, \hat{\mathbf{i}}_h\}$ , which can be constructed with the following:

$$\hat{\mathbf{i}}_r = \frac{\mathbf{r}_{LM}}{\|\mathbf{r}_{LM}\|} \quad (10)$$

$$\hat{\mathbf{i}}_\theta = \hat{\mathbf{i}}_h \times \hat{\mathbf{i}}_r \quad (11)$$

$$\hat{\mathbf{i}}_h = \frac{\mathbf{r}_{LM} \times \dot{\mathbf{r}}_{LM}}{\|\mathbf{r}_{LM} \times \dot{\mathbf{r}}_{LM}\|} \quad (12)$$

If we stack up these vectors into a matrix  $[\hat{\mathbf{i}}_r \ \hat{\mathbf{i}}_\theta \ \hat{\mathbf{i}}_h]$ , this defines the direction cosine matrix which takes vectors in the Hill frame to the inertial frame:  $[NH]$ . We can take the transpose to find the opposite:  $[HN] = [\hat{\mathbf{i}}_r \ \hat{\mathbf{i}}_\theta \ \hat{\mathbf{i}}_h]^T$ .

## Task 3: Sun-Pointing Reference Frame Orientation

The solar panel axis  $\hat{\mathbf{b}}_3$  must be pointed at the sun, and a reference frame  $\mathcal{R}_s$  must be generated such that  $\hat{\mathbf{r}}_3$  points in the sun direction ( $\hat{\mathbf{n}}_2$ ). Given that the solar reference frame is constant with respect to the inertial frame, the  ${}^N \boldsymbol{\omega}_{R_s N} = \mathbf{0}$ . And our DCM is easily constructed using our assumptions with the following:

$$[R_s N] = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix} \quad (13)$$

### Task 4: Nadir-Pointing Reference Frame Orientation

In order to point the payload platform axis  $\hat{\mathbf{b}}_1$  towards Mars in the nadir direction, the reference frame  $\mathcal{R}_n$  must be constructed such that  $\hat{\mathbf{r}}_1$  points towards the planet. Additionally, we assume that  $\hat{\mathbf{r}}_2$  is in the direction of the velocity  $\hat{\mathbf{i}}_\theta$ . Therefore we easily can construct a Hill-to-reference DCM which, using our now stated definitions, follows as such:

$$[R_n H] = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{bmatrix} \quad (14)$$

This is the manifestation of a simple  $\pi$  rotation about the second Hill axis, where the reference flips  $\hat{\mathbf{i}}_r$  and  $\hat{\mathbf{i}}_h$ . We can then calculate  $[HN]$  using our procedure from Task 2. We then generate  $[R_n N]$  via the following:

$$[R_n N] = [R_n H][HN] \quad (15)$$

Similarly, given that we are on a circular orbit, and that our reference is an invariant transformation from the Hill frame, we can easily describe  ${}^N\omega_{R_n N}$ . Given that the reference and Hill angular rates are similar, we know that  ${}^H\omega_{R_n N} = [0 \ 0 \ \dot{\theta}]^T$  and can supply the reference angular rate with the following

$${}^N\omega_{R_n N} = [HN]^T {}^H\omega_{R_n N} = [NH][0 \ 0 \ \dot{\theta}]^T \quad (16)$$

### Task 5: GMO-Pointing Reference Frame Orientation

Now we must construct another reference frame  $\mathcal{R}_c$  such that  $-\hat{\mathbf{r}}_1$  = points towards the GMO spacecraft. This is simply done by finding the vector which represents the inertial difference in the position of both spacecraft:  $\Delta\mathbf{r} = \mathbf{r}_{LMO} - \mathbf{r}_{GMO}$ . We can then describe the frame with the following:

$$\hat{\mathbf{r}}_1 = \frac{-\Delta\mathbf{r}}{\|\Delta\mathbf{r}\|} \quad (17)$$

$$\hat{\mathbf{r}}_2 = \frac{\Delta\mathbf{r} \times \hat{\mathbf{n}}_3}{\|\Delta\mathbf{r} \times \hat{\mathbf{n}}_3\|} \quad (18)$$

$$\hat{\mathbf{r}}_3 = \hat{\mathbf{r}}_1 \times \hat{\mathbf{r}}_2 \quad (19)$$

Stacking these unit vectors as such  $[\hat{\mathbf{r}}_1 \ \hat{\mathbf{r}}_2 \ \hat{\mathbf{r}}_3]$  yields a rotation matrix that, when multiplied by, brings vectors from the tracking reference frame to the inertial frame. Therefore, under a transpose operation we get the following:

$$[R_c N] = [\hat{\mathbf{r}}_1 \ \hat{\mathbf{r}}_2 \ \hat{\mathbf{r}}_3]^T \quad (20)$$

Finding  ${}^N\omega_{R_c N}$  is nontrivial and finding an analytical expression for the time derivative of the DCM can be challenging. Instead, we can use a numerical approach to find a usable solution. We know that the derivative of a DCM is that:  $[\dot{C}] = -[\omega^\times][C]$ . Therefore we can find the angular rate with the following:

$$\frac{d[R_c N]}{dt} = -[\omega_{R_c N}^\times][R_c N] \quad (21)$$

$$\frac{[R_c N(t + dt)] - [R_c N(t)]}{dt} [N R_c] = -[\omega_{R_c N}^\times] \quad (22)$$

Because we know have a function that determines this reference DCM at any point in time, this numerical derivative is easy to calculate for a small value  $dt$ . With knowledge of the skew symmetric form, we can de-skew  $[\omega_{R_c N}^\times]$  to find our vector  ${}^{R_c}\omega_{R_c N}$ . To bring this quantity into the inertial frame we perform  ${}^N\omega_{R_c N} = [R_c N(t)]^T {}^{R_c}\omega_{R_c N}$ .

## Task 6: Attitude and Angular Rate Error Evaluation

In this section, we must write a function that, given a current attitude state  $\sigma_{BN}$ , angular rate  ${}^B\omega_{BN}$ , and desired reference attitude matrix  $[RN]$ , returns the associate tracking errors  $\sigma_{BR}$  and  ${}^B\omega_{BR}$ . First let us start with the simpler angular velocity error:

$${}^B\omega_{BR} = ({}^B\omega_{BN} - [BN]^N\omega_{RN}) \quad (23)$$

We can find the inertial to body DCM transform by performing  $\text{MRP2C}(\sigma_{BN})$  with the following function:

$$[BN] = \frac{1}{(1 + \sigma^2)^2} \begin{bmatrix} 4(\sigma_1^2 - \sigma_2^2 - \sigma_3^2) + (1 - \sigma^2)^2 & 8\sigma_1\sigma_2 + 4\sigma_3(1 - \sigma^2) & 8\sigma_1\sigma_3 - 4\sigma_2(1 - \sigma^2) \\ 8\sigma_2\sigma_1 - 4\sigma_3(1 - \sigma^2) & 4(-\sigma_1^2 + \sigma_2^2 - \sigma_3^2) + (1 - \sigma^2)^2 & 8\sigma_2\sigma_3 + 4\sigma_1(1 - \sigma^2) \\ 8\sigma_3\sigma_1 + 4\sigma_2(1 - \sigma^2) & 8\sigma_3\sigma_2 - 4\sigma_1(1 - \sigma^2) & 4(-\sigma_1^2 - \sigma_2^2 + \sigma_3^2) + (1 - \sigma^2)^2 \end{bmatrix} \quad (24)$$

Now that we have the tracking error for the angular velocity, we must find the relative error in the modified Rodrigues parameter attitude formalism. We could use the relative MRP formula, but it can be understood more easily by converting to DCMs and using fundamental properties of the SO(3) group:

$$\sigma_{BR} = \text{C2MRP}([BN][RN]^T) \quad (25)$$

The DCM to MRP transform is more complicated and is done by first converting the DCM to a quaternion via Sheppard's method. The first step is to find the maximum of these values, as truth, to constrain that value for the second step:

$$\begin{aligned} \beta_0^2 &= \frac{1}{4}(1 + \text{tr}([BR])) & \beta_2^2 &= \frac{1}{4}(1 + 2[BR]_{22} - \text{tr}([BR])) \\ \beta_1^2 &= \frac{1}{4}(1 + 2[BR]_{11} - \text{tr}([BR])) & \beta_3^2 &= \frac{1}{4}(1 + 2[BR]_{33} - \text{tr}([BR])) \end{aligned} \quad (26)$$

The second step is done by computing the rest of the quaternion entries, using our constrained entry, with the following:

$$\begin{aligned} \beta_0\beta_1 &= ([BR]_{23} - [BR]_{32})/4 & \beta_1\beta_2 &= ([BR]_{12} + [BR]_{21})/4 \\ \beta_0\beta_2 &= ([BR]_{31} - [BR]_{13})/4 & \beta_3\beta_1 &= ([BR]_{31} + [BR]_{13})/4 \\ \beta_0\beta_3 &= ([BR]_{12} - [BR]_{21})/4 & \beta_2\beta_3 &= ([BR]_{23} + [BR]_{32})/4 \end{aligned} \quad (27)$$

The final MRP is calculated with from our quaternion entries using the definition:

$$\sigma_{BR} = \begin{bmatrix} \frac{\beta_1}{1+\beta_0} \\ \frac{\beta_2}{1+\beta_0} \\ \frac{\beta_3}{1+\beta_0} \end{bmatrix} \quad (28)$$

## Task 7: Numerical Attitude Simulator

Now we must numerically integrate our differential equations of motion to simulate the dynamics of our system for both the LMO and GMO spacecraft. Let us define our state vector as the following:

$$\mathbf{X} = \begin{bmatrix} \sigma_{BN} \\ {}^B\omega_{BR} \end{bmatrix} \quad (29)$$

For  $\mathbf{u}$  control torque vector, the rigid body dynamics obey the following:

$$[I]\dot{\omega}_{BN} = -[\omega_{BN}^\times][I]\omega_{BN} + \mathbf{u} - \mathbf{L} \quad (30)$$

We use a fourth order Runge-Kutte algorithm for integration (RK4). Using the nonlinear dynamics function  $\dot{\mathbf{X}} = f(t, \mathbf{X})$ , the integration is Algorithm 1. Each point  $i \in [1 : N]$  is 1 integration time step, and therefore the full simulation time  $dtN$ . Using this simulation framework, we can study our angular momentum  $\mathbf{H} = [I]\omega_{BN}$  and kinetic energy  $T = \frac{1}{2}\omega_{BN}^T[I]\omega_{BN}$  over time.

---

**Algorithm 1** Fourth Order Runge Kutte Integrator

---

```
1: for i = 1:N-1 do
2:    $k_1 = \dot{\mathbf{X}}(t(i), \mathbf{X}(:, i))$ 
3:    $k_2 = \dot{\mathbf{X}}(t(i) + \frac{dt}{2}, \mathbf{X}(:, i) + \frac{dt}{2}k_1)$ 
4:    $k_3 = \dot{\mathbf{X}}(t(i) + \frac{dt}{2}, \mathbf{X}(:, i) + \frac{dt}{2}k_2)$ 
5:    $k_4 = \dot{\mathbf{X}}(t(i) + dt, \mathbf{X}(:, i) + dtk_3)$ 
6:    $\mathbf{X}(:, i+1) = \mathbf{X}(:, i) + \frac{dt}{6}(k_1 + 2k_2 + 2k_3 + k_4)$ 
7: end for
```

---

### Task 8: Control for Sun, Nadir, and Communication Pointing

Next, the control architecture must be developed. We shall use the linearized closed loop dynamics to determine the proportional and derivative gains.

Let us consider the PD nonlinear feedback control law:

$$\mathbf{u} = -K\boldsymbol{\sigma} - [P]\delta\boldsymbol{\omega} + [I](\dot{\boldsymbol{\omega}}_r - [\boldsymbol{\omega}^\times]\boldsymbol{\omega}_r) + [\boldsymbol{\omega}^\times][I]\boldsymbol{\omega} - \mathbf{L} \quad (31)$$

Let us disregard the external torque modeling error and consider the displacement MRP and deviation angular rate with the following control:

$${}^B\mathbf{u} = -K\boldsymbol{\sigma}_{B/R} - P^B\boldsymbol{\omega}_{B/R} \quad (32)$$

Let us look at the closed loop dynamics with this state vector:

$$\mathbf{x} = \begin{bmatrix} \boldsymbol{\sigma} \\ \delta\boldsymbol{\omega} \end{bmatrix} \quad (33)$$

Using the differential kinematic equation for the MRPs we arrive at the following nonlinear state space formulation, with the closed loop full-state feedback:

$$\dot{\mathbf{x}} = \begin{bmatrix} \dot{\boldsymbol{\sigma}} \\ \delta\dot{\boldsymbol{\omega}} \end{bmatrix} = \begin{bmatrix} 0 & \frac{1}{4}B(\boldsymbol{\sigma}) \\ -K[I]^{-1} & -[I]^{-1}[P] \end{bmatrix} \begin{bmatrix} \boldsymbol{\sigma} \\ \delta\boldsymbol{\omega} \end{bmatrix} \quad (34)$$

The small angle approximation is employed to linearize this formulation with the following:

$$\begin{bmatrix} \dot{\boldsymbol{\sigma}} \\ \delta\dot{\boldsymbol{\omega}} \end{bmatrix} = \begin{bmatrix} 0 & \frac{1}{4}\mathbb{I}_3 \\ -K[I]^{-1} & -[I]^{-1}[P] \end{bmatrix} \begin{bmatrix} \boldsymbol{\sigma} \\ \delta\boldsymbol{\omega} \end{bmatrix} \quad (35)$$

We can find the eigenvalues of this matrix to see how the gains affect the stability of the system. The roots act as the following:

$$\lambda_i = \frac{-1}{2I_i} \left( P_i \pm \sqrt{-KI_i + P_i^2} \right) \quad i = 1, 2, 3 \quad (36)$$

Now, we can choose the gains such that they meet our operation criterion. We want to use single scalar gains for both  $K$  and  $P$ . We know that we have a decay time of  $T = 120$  seconds which constraints out  $P$  value with the following:

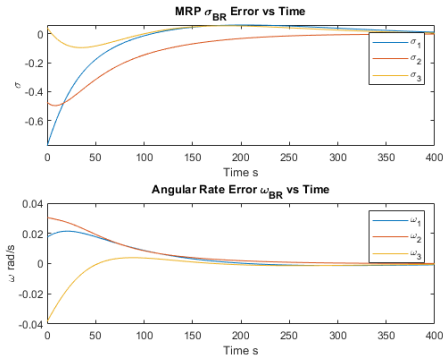
$$P = \max_i \left( \frac{2I_i}{T} \right) \quad (37)$$

Once we have constrained this gain value, we can take a look at the  $K$  gain and damping coefficient for each of the modes. We know the relationship  $\xi_i = \frac{P}{\sqrt{KI_i}}$ . We can find a solution for the critically damped  $\xi = 1$  mode, then pick an inertia that results in the other modes being critically or underdamped, where  $\xi \leq 1$ . We end with the result that  $[P \ K] = [0.1666 \ 0.0055]$ .

### Task 8: Sun Pointing Plots

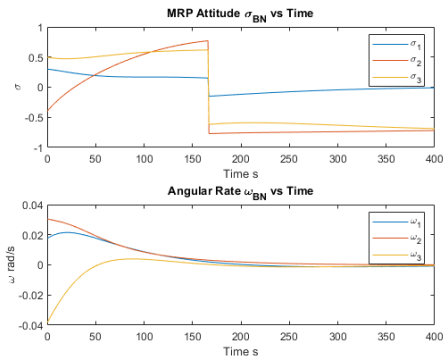
Let us now simulate the LMO spacecraft performing the sun-point maneuver. We calculate the tracking errors with the function developed earlier and use the reference DCM and angular rate vector developed earlier. These tracking errors are used in our control function with the gains determined in the previous step.

With these gains, we see how the system tracks the reference for a duration of 400 seconds:



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**Figure 2: A subfigure**



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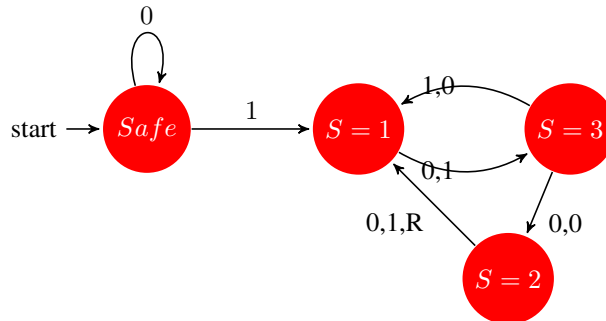
**Figure 3: A subfigure**

**Figure 4: A figure with two subfigures**

**Task 9: Sun Pointing Plots**

**Task 10: Sun Pointing Plots**

**Task 11: Mode Switching Mission Scenario**



**I. Conclusion**

**Acknowledgment**

**References**

### Listing 1: Main Script

```
1 %% Padraig Lysandrou April 4th 2019 -- ASEN5010 Final Project
2 clc; close all; clear all;
3
4 % Problem specified dynamic initial conditions
5 sigma_BN_0 = [0.3 -0.4 0.5].';
6 omega_BN_B_0 = deg2rad([1 1.75 -2.2]).';
7 Ic = diag([10 5 7.5]); p.Ic = Ic;
8
9 % Problem specified orbital parameters
10 mu_M = 42828.3; % km3/s2
11 h_LMO = 400; % km
12 R_M = 3396.19; % km
13 r_LMO = h_LMO + R_M;
14 r_GMO = 20424.2;
15 Omega_GMO = 0; inc_GMO = 0; theta_GMO_0 = deg2rad(250);
16 Omega_LMO = deg2rad(20); inc_LMO = deg2rad(30); theta_LMO_0 = deg2rad(60);
17 theta_dot_LMO = sqrt(mu_M / (r_LMO^3)); % rad/s, orbital angular rate(circ)
18 theta_dot_GMO = sqrt(mu_M / (r_GMO^3)); % rad/s, orbital angular rate(circ)
19
20 % theta_dot_GMO = 0.0000709003;
21
22 % Take care of timing
23 tend = 6500;
24 dt = 1;
25 t = 0:dt:tend;
26 npoints = length(t);
27
28 % Set up the dynamic function as well as state tracking, init conds
29 f_dot = @(t_in,state_in,param) dynamics(t_in,state_in,param);
30 vehicle_state = zeros(6,npoints);
31 vehicle_state(:,1) = [sigma_BN_0; omega_BN_B_0];
32 p.L = [0.0 0 0].';
33 p.u = [0 0 0].';
34
35 % Setup the tracking error stuff, and input history stuff
36 sigma_BR = zeros(3,npoints);
37 omega_BR = zeros(3,npoints);
38 control_input = zeros(3, npoints);
39 sig_int = [0 0 0].';
40
41 % Set the initial conditions
42 H = zeros(3,npoints);
43 H(:,1) = Ic*omega_BN_B_0;
44 T = zeros(1,npoints);
45 T(1) = 0.5*omega_BN_B_0.'*Ic*omega_BN_B_0;
46
47 % PD controller gain initialization
48 P = max(diag(Ic).*(2/120));
49 K = (P^2)./Ic(2,2); % still not sure why this is the gain he wants...
50
51 % xi_1 = sqrt((P^2)./(K.*Ic(1,1))) % keeps <= 1
52 % xi_2 = sqrt((P^2)./(K.*Ic(2,2))) % does not keep <= 1
53 % xi_3 = sqrt((P^2)./(K.*Ic(3,3))) % does not keep <= 1
54
55
56 %% Start the simulation
57 vehicle_mode = 4;
58
59 tic
60 for i = 1:npoints-1
61     % Pull out state values to be used below
62     sigma_BN = vehicle_state(1:3,i);
63     omega_BN = vehicle_state(4:6,i);
64
65     % Determine the inertial small sat and GMO vectors
66     theta_LMO = theta_LMO_0 + t(i)*theta_dot_LMO;
67     [rN_LMO,-] = oe2rv_schaub(r_LMO,mu_M,Omega_LMO,inc_LMO,theta_LMO);
```



```

68 theta_GMO = theta_GMO_0 + t(i)*theta_dot_GMO;
69 [rN_GMO,-] = oe2rv_schaub(r_GMO,mu_M,Omega_GMO,inc_GMO,theta_GMO);
70
71 % Determine the mode of the spacecraft
72 if rN_LMO(2) ≥ 0
73     vehicle_mode = 1;
74 elseif acosd((rN_LMO.'*rN_GMO)/(norm(rN_LMO)*norm(rN_GMO))) < 35
75     % put the system into GMO data transfer mode
76     vehicle_mode = 3;
77 else
78     % put the system into nadir point science mode
79     vehicle_mode = 2;
80 end
81
82
83 % Determine the trackgin DCM/Omega based on vehicle mode
84 switch vehicle_mode
85     case 1 % Sun pointing energy gather mode
86         RN = dcm_RsN(t(i));
87         omega_RN = omega_RsN(t(i));
88         [sigma_BR(:,i), omega_BR(:,i)] = ...
89             track_error(sigma_BN,omega_BN,RN,omega_RN);
90     case 2 % Nadir pointing science mode
91         RN = dcm_RnN(t(i));
92         omega_RN = omega_RnN(t(i),theta_dot_LMO);
93         [sigma_BR(:,i), omega_BR(:,i)] = ...
94             track_error(sigma_BN,omega_BN,RN,omega_RN);
95     case 3 % GMO comm pointing mode
96         RN = dcm_RcN(t(i));
97         omega_RN = omega_RcN(t(i));
98         [sigma_BR(:,i), omega_BR(:,i)] = ...
99             track_error(sigma_BN,omega_BN,RN,omega_RN);
100    case 4 % Safemode hold
101        sigma_BR(:,i) = [0 0 0].';
102        omega_BR(:,i) = [0 0 0].';
103    otherwise % Safemode hold
104        sigma_BR(:,i) = [0 0 0].';
105        omega_BR(:,i) = [0 0 0].';
106 end
107
108 % Calculate the control input from mode error
109 sig_int = sig_int + dt.*(sigma_BR(:,i));
110 p.u = (-K.*sigma_BR(:,i)) + (-P.*omega_BR(:,i)); % + -0.01.*sig_int ;
111 control_input(:,i) = p.u;
112
113
114 % Pull out the angular momentum and the energy for debugging and
115 % verification
116 H(:,i) = Ic*omega_BN;
117 T(i) = 0.5*omega_BN.'*Ic*omega_BN;
118
119 % RK4 step for the spacecraft dynamics
120 k_1 = f_dot(t(i),vehicle_state(:,i),p);
121 k_2 = f_dot(t(i)+0.5*dt, vehicle_state(:,i)+0.5*dt*k_1,p);
122 k_3 = f_dot((t(i)+0.5*dt), (vehicle_state(:,i)+0.5*dt*k_2), p);
123 k_4 = f_dot((t(i)+dt), (vehicle_state(:,i)+k_3*dt), p);
124 vehicle_state(:,i+1) = vehicle_state(:,i) + (1/6)*(k_1+(2*k_2)+(2*k_3)+k_4)*dt;
125
126 % Perform the nonsingular MRP propagation attitude check
127 s = norm(vehicle_state(1:3,i+1));
128 if s > 1
129     vehicle_state(1:3,i+1) = -(vehicle_state(1:3,i+1) ./ (s^2));
130 end
131 end
132 toc
133
134
135 %% Plot data here.

```

```

136
137 % figure; plot(t, H); title('angular momentum');
138 figure; plot(t, vehicle_state(1:3,:)); title('mrps over time')
139 figure; plot(t, vehicle_state(4:6,:)); title('omegas over time')
140 % figure; plot(t, vecnorm(H)); title('Angular Momentum over time')
141
142 % figure; semilogy(t, T); title('system angular energy over time')
143 figure; plot(t, sigma_BR); title('MRP tracking error vs time');
144 figure; plot(t, omega_BR); title('Angular rate tracking error vs time');
145
146
147
148 % Saving things for Checkoff 7
149 % save_to_txt('H500B.txt', H(:,end));
150 % save_to_txt('T500.txt', T(end));
151 % save_to_txt('MRP500.txt', vehicle_state(1:3,end));
152 % save_to_txt('H500N.txt', MRP2C(vehicle_state(1:3,end)).'*H(:,end));
153
154 % close all;
155 % Saving things for Checkoff 8-10:
156 % save_to_txt('gains.txt', [P K]);
157 % save_to_txt('MRP15.txt', vehicle_state(1:3,16));
158 % save_to_txt('MRP100.txt', vehicle_state(1:3,101));
159 % save_to_txt('MRP200.txt', vehicle_state(1:3,201));
160 % save_to_txt('MRP400.txt', vehicle_state(1:3,401));
161
162 % close all;
163 % Saving things for Checkoff 11:
164 % save_to_txt('MRP300.txt', vehicle_state(1:3,301));
165 % save_to_txt('MRP2100.txt', vehicle_state(1:3,2101));
166 % save_to_txt('MRP3400.txt', vehicle_state(1:3,3401));
167 % save_to_txt('MRP4400.txt', vehicle_state(1:3,4401));
168 % save_to_txt('MRP5600.txt', vehicle_state(1:3,5601));

```