Attitude Dynamics and Control of a Nano-Satellite Orbiting Mars

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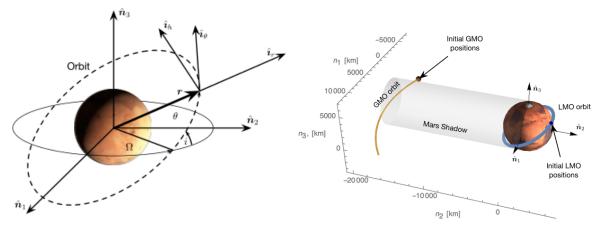
This project for ASEN5010 Spacecraft Dynamics and Control considers a small satellite orbiting Mars at a low altitude. This spacecraft gathers science data and transfers this data to another satellite orbiting at a higher altitude. Periodically, this spacecraft must transition from nadir-pointing, science gathering mode to sunpointing mode to recharge the battery system. The three missions goals are nadir-pointing, communicating with the mother spacecraft, and to sun-point. Both of these spacecraft are in circular orbits.

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Introduction

Problem Statement

Let us begin with defining the orbit of the nano-satellite with the following figure



a) Illustration of the Inertial, Hill, and perifocal geometrical constructions. Taken from ASEN5010 Semester Project sheet.

b) Tracking Error vs Time

Initial Conditions and Problem Setup

Before we begin with our work on the simulation and controller, we must define our initial conditions and operational parameters.

$$\sigma_{B/N}(t_0) = \begin{bmatrix} 0.3\\ -0.4\\ 0.5 \end{bmatrix}$$
 (1)

$$\sigma_{B/N}(t_0) = \begin{bmatrix} 0.3 \\ -0.4 \\ 0.5 \end{bmatrix}$$

$${}^{\mathcal{B}}\omega_{B/N}(t_0) = \begin{bmatrix} 1 \\ 1.75 \\ -2.2 \end{bmatrix} \operatorname{deg/s}$$

$${}^{\mathcal{B}}[I] = \begin{bmatrix} 10 & 0 & 0 \\ 0 & 5 & 0 \\ 0 & 0 & 7.5 \end{bmatrix} \operatorname{kg m}^2$$
(3)

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(3)

We will also use the following Keplerian orbital elements for both spacecraft:

Spacecraft	r	Ω	i	$\theta(t_0)$
LMO	3796.19 km	20°	30°	60°
GMO	20424.2 km	0°	0°	250°

Additionally, we use a gravitational parameter $\mu=42828.3 \mathrm{km}^3/\mathrm{s}^2$. We can also calculate each rotation rate using the circular motion assumption of $\dot{\theta} = \sqrt{\frac{\mu}{r^3}}$. These parameters will be used for all of the tasks and intermediate simulations in this project.

Task 1: Orbit Simulation

Our Hill frame is defined by the basis: $\{\hat{\pmb{i}}_r, \hat{\pmb{i}}_\theta, \hat{\pmb{i}}_h\}$ with the inertial defined as $\{\hat{\pmb{n}}_1, \hat{\pmb{n}}_2, \hat{\pmb{n}}_3\}$. Given the inertial and Hill frame definitions, we know that the position vector of the LMO satellite is \hat{rir} . Additionally we know that since it is a circular orbit, it has a time invariant angular rate $\omega_{H/N} = \dot{\theta} \hat{\mathbf{i}}_h$. Calculating the vectorial inertial derivative:

$$\dot{\mathbf{r}} = \frac{^{N}d}{dt}\mathbf{r} = \frac{^{H}d}{dt}\mathbf{r} + \boldsymbol{\omega}_{H/N} \times \mathbf{r}$$
(4)

$$= \dot{\theta} \hat{i}_h \times r \hat{i}_r \tag{5}$$

$$=r\dot{\theta}\hat{i}_{\theta} \tag{6}$$

Additionally, we can use this information to find the inertial position and velocity vectors by performing transformations using the perifocal frame information. We know that the perifocal frame can be defined by an Euler 3-1-3 rotation defined the set $\{\Omega, i, \theta\}$

$$C_{ECI} = \begin{bmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos i & \sin i \\ 0 & -\sin i & \cos i \end{bmatrix} \begin{bmatrix} \cos \Omega & \sin \Omega & 0 \\ -\sin \Omega & \cos \Omega & 0 \\ 0 & 0 & 1 \end{bmatrix}$$
(7)

Which describes a rotation from Earth Centered Inertial frame. Each portion of the DCM is a single-axis rotation. We can then use this to project scalar values in the Hill frame to inertial vectors with the following:

$${}^{N}\vec{\boldsymbol{r}} = C_{ECI}^{T} \begin{bmatrix} r \\ 0 \\ 0 \end{bmatrix} \tag{8}$$

$${}^{N}\vec{\boldsymbol{v}} = C_{ECI}^{T} \begin{bmatrix} 0 \\ r\dot{\theta} \\ 0 \end{bmatrix} \tag{9}$$

When the ECI direction cosine matrix is calculated, θ must be propagated over time, as the true anomaly is the only perifocal parameter that is time variant. It is calculated as such: $\theta = \theta_0 + t * \dot{\theta}$.

Task 2: Orbit Frame Orientation

It is simple to generate bases vectors for the Hill frame, under motion, using our new inertial vectors. As stated before, $\mathcal{H} = \{\hat{i}_r, \hat{i}_\theta, \hat{i}_h\}$, which can be constructed with the following:

$$\hat{\boldsymbol{i}_r} = \frac{\boldsymbol{r}_{LM}}{\|\boldsymbol{r}_{LM}\|} \tag{10}$$

$$\hat{i_{\theta}} = \hat{i_h} \times \hat{i_r} \tag{11}$$

$$\hat{\boldsymbol{i}}_h = \frac{\boldsymbol{r}_{LM} \times \dot{\boldsymbol{r}}_{LM}}{\|\boldsymbol{r}_{LM} \times \dot{\boldsymbol{r}}_{LM}\|} \tag{12}$$

If we stack up these vectors into a matrix $[\hat{i_r} \ \hat{i_\theta} \ \hat{i_h}]$, this defines the direction cosine matrix which takes vectors in the Hill frame to the inertial frame: [NH]. We can take the transpose to find the opposite: $[HN] = [\hat{i_r} \ \hat{i_\theta} \ \hat{i_h}]^T$.

Task 3: Sun-Pointing Reference Frame Orientation

The solar panel axis \hat{b}_3 must be pointed at the sun, and a reference frame \mathcal{R}_s must be generated such that \hat{r}_3 points in the sun direction (\hat{n}_2) . Given that the solar reference frame is constant with respect to the inertial frame, the ${}^N\omega_{R_sN}=\mathbf{0}$. And our DCM is easily constructed using our assumptions with the following:

$$[R_s N] = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix} \tag{13}$$

Task 4: Nadir-Pointing Reference Frame Orientation

In order to point the payload platform axis \hat{b}_1 towards Mars in the nadir direction, the reference frame \mathcal{R}_n must be constructed such that \hat{r}_1 points towards the planet. Additionally, we assume that \hat{r}_2 is in the direction of the velocity \hat{i}_{θ} . Therefore we easily can construct a Hill-to-reference DCM which, using our now stated definitions, follows as such:

$$[R_n H] = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{bmatrix}$$
 (14)

This is the manifestation of a simple π rotation about the second Hill axis, where the reference flips \hat{i}_r and \hat{i}_h . We can then calculate [HN] using our procedure from Task 2. We then generate $[R_nN]$ via the following:

$$[R_n N] = [R_n H][HN] \tag{15}$$

Similarly, given that we are on a circular orbit, and that our reference is an invariant transformation from the Hill frame, we can easily describe ${}^{N}\omega_{R_{n}N}$. Given that the reference and Hill angular rates are similar, we know that ${}^{H}\omega_{R_{n}N} = [0\ 0\ \dot{\theta}]^{T}$ and can supply the reference angular rate with the following

$${}^{N}\boldsymbol{\omega}_{R_{n}N} = [HN]^{T} {}^{H}\boldsymbol{\omega}_{R_{n}N} = [NH][0 \ 0 \ \dot{\theta}]^{T}$$
 (16)

Task 5: GMO-Pointing Reference Frame Orientation

Now we must construct another reference frame \mathcal{R}_c such that $-\hat{r}_1$ = points towards the GMO spacecraft. This is simply done by finding the vector which represents the inertial difference in the position of both spacecraft: $\Delta r =$ $r_{LMO} - r_{GMO}$. We can then describe the frame with the following:

$$\hat{\mathbf{r}}_{1} = \frac{-\Delta \mathbf{r}}{\|\Delta \mathbf{r}\|}$$

$$\hat{\mathbf{r}}_{2} = \frac{\Delta \mathbf{r} \times \hat{\mathbf{n}}_{3}}{\|\Delta \mathbf{r} \times \hat{\mathbf{n}}_{3}\|}$$

$$(17)$$

$$\hat{\mathbf{r}}_2 = \frac{\Delta \mathbf{r} \times \hat{\mathbf{n}}_3}{\|\Delta \mathbf{r} \times \hat{\mathbf{n}}_3\|} \tag{18}$$

$$\hat{\boldsymbol{r}}_3 = \hat{\boldsymbol{r}}_1 \times \hat{\boldsymbol{r}}_2 \tag{19}$$

Stacking these unit vectors as such $[\hat{r}_1 \ \hat{r}_2 \ \hat{r}_3]$ yields a rotation matrix that, when multiplied by, brings vectors from the tracking reference frame to the inertial frame. Therefore, under a transpose operation we get the following:

$$[R_c N] = [\hat{r}_1 \ \hat{r}_2 \ \hat{r}_3]^T \tag{20}$$

Finding ${}^N\omega_{R_cN}$ is nontrivial and finding an analytical expression for the time derivative of the DCM can be challenging. Instead, we can use a numerical approach to find a usable solution. We know that the derivative of a DCM is that: $[C] = -[\omega^{\times}][C]$ with \times denoting the skew symmetric matrix form of a cross product. Therefore we can find the angular rate with the following:

$$\frac{d[R_c N]}{dt} = -[\boldsymbol{\omega}_{\boldsymbol{R}_c \boldsymbol{N}}^{\times}][R_c N]$$
 (21)

$$\frac{d[R_c N]}{dt} = -[\boldsymbol{\omega}_{\boldsymbol{R}_c \boldsymbol{N}}^{\times}][R_c N]$$

$$\frac{[R_c N(t+dt)] - [R_c N(t)]}{dt}[NR_c] = -[\boldsymbol{\omega}_{\boldsymbol{R}_c \boldsymbol{N}}^{\times}]$$
(21)

Because we know have a function that determines this reference DCM at any point in time, this numerical derivative is easy to calculate for a small value dt. With knowledge of the skew symmetric form, we can de-skew $[\omega_{R,N}^{\times}]$ to find our vector ${}^{R_c}\omega_{R_cN}$. To bring this quantity into the inertial frame we perform ${}^{N}\omega_{R_cN}=[R_cN(t)]^T {}^{R_c}\omega_{R_cN}$.

Task 6: Attitude and Angular Rate Error Evaluation

In this section, we must write a function that, given a current attitude state σ_{BN} , angular rate ${}^B\omega_{BN}$, and desired reference attitude matrix [RN], returns the associate tracking errors σ_{BR} and $^B\omega_{BR}$. First let us start with the simpler angular velocity error:

$${}^{B}\boldsymbol{\omega}_{BR} = ({}^{B}\boldsymbol{\omega}_{BN} - [BN] {}^{N}\boldsymbol{\omega}_{RN}) \tag{23}$$

We can find the inertial to body DCM transform by performing MRP2C(σ_{BN}) with the following function:

$$[BN] = \frac{1}{\left(1 + \sigma^2\right)^2} \begin{bmatrix} 4\left(\sigma_1^2 - \sigma_2^2 - \sigma_3^2\right) + \left(1 - \sigma^2\right)^2 & 8\sigma_1\sigma_2 + 4\sigma_3\left(1 - \sigma^2\right) & 8\sigma_1\sigma_3 - 4\sigma_2\left(1 - \sigma^2\right) \\ 8\sigma_2\sigma_1 - 4\sigma_3\left(1 - \sigma^2\right) & 4\left(-\sigma_1^2 + \sigma_2^2 - \sigma_3^2\right) + \left(1 - \sigma^2\right)^2 & 8\sigma_2\sigma_3 + 4\sigma_1\left(1 - \sigma^2\right) \\ 8\sigma_3\sigma_1 + 4\sigma_2\left(1 - \sigma^2\right) & 8\sigma_3\sigma_2 - 4\sigma_1\left(1 - \sigma^2\right) & 4\left(-\sigma_1^2 - \sigma_2^2 + \sigma_3^2\right) + \left(1 - \sigma^2\right)^2 \end{bmatrix}$$

$$(24)$$

Now that we have the tracking error for the angular velocity, we must find the relative error in the modified rodrigues parameter attitude formalism. We could use the relative MRP formula, but it can be understood more easily by converting to DCMs and using fundamental properties of the SO(3) group:

$$\sigma_{BR} = \text{C2MRP}([BN][RN]^T) \tag{25}$$

The DCM to MRP transform is more complicated and is done by first converting the DCM to a quaternion via Sheppard's method. The first step is to find the maximum of these values, as truth, to constrain that value for the second step:

$$\beta_0^2 = \frac{1}{4}(1 + \operatorname{tr}([BR])) \qquad \beta_2^2 = \frac{1}{4}(1 + 2[BR]_{22} - \operatorname{tr}([BR]))
\beta_1^2 = \frac{1}{4}(1 + 2[BR]_{11} - \operatorname{tr}([BR])) \qquad \beta_3^2 = \frac{1}{4}(1 + 2[BR]_{33} - \operatorname{tr}([BR]))$$
(26)

The second step is done by computing the rest of the quaternion entries, using our constrained entry, with the following:

$$\beta_{0}\beta_{1} = ([BR]_{23} - [BR]_{32})/4 \quad \beta_{1}\beta_{2} = ([BR]_{12} + [BR]_{21})/4$$

$$\beta_{0}\beta_{2} = ([BR]_{31} - [BR]_{13})/4 \quad \beta_{3}\beta_{1} = ([BR]_{31} + [BR]_{13})/4$$

$$\beta_{0}\beta_{3} = ([BR]_{12} - [BR]_{21})/4 \quad \beta_{2}\beta_{3} = ([BR]_{23} + [BR]_{32})/4$$

$$(27)$$

The final MRP is calculated with from our quaternion entries using the definition:

$$\sigma_{BR} = \begin{bmatrix} \frac{\beta_1}{1+\beta_0} \\ \frac{\beta_2}{1+\beta_0} \\ \frac{\beta_3}{1+\beta_0} \end{bmatrix}$$
 (28)

Task 7: Numerical Attitude Simulator

Now we must numerically integrate our differential equations of motion to simulate the dynamics of our system for both the LMO and GMO spacecraft. Let us define our state vector as the following:

$$X = \begin{bmatrix} \sigma_{BN} \\ B_{\omega_{BR}} \end{bmatrix} \tag{29}$$

For u control torque vector, the rigid body dynamics obey the following:

$$[I]\dot{\boldsymbol{\omega}}_{BN} = -[\boldsymbol{\omega}_{BN}^{\times}][I]\boldsymbol{\omega}_{BN} + \boldsymbol{u} - \boldsymbol{L}$$
(30)

We use a fourth order Runge-Kutte algorithm for integration (RK4). Using the nonlinear dynamics function $\dot{X}=f(t,X)$, the integration is Algorithm 1. Each point $i\in[1:N]$ is 1 integration time step, and therefore the full simulation time dtN. Using this simulation framework, we can study our angular momentum $\boldsymbol{H}=[I]\boldsymbol{\omega}_{BN}$ and kinetic energy $T=\frac{1}{2}\boldsymbol{\omega}_{BN}^T[I]\boldsymbol{\omega}_{BN}$ over time.

Algorithm 1 Fourth Order Runge Kutte Integrator

```
1: for i = 1:N-1 do

2: k_1 = \dot{\boldsymbol{X}}(t(i), \ \boldsymbol{X}(:,i))

3: k_2 = \dot{\boldsymbol{X}}(t(i) + \frac{dt}{2}, \ \boldsymbol{X}(:,i) + \frac{dt}{2}k_1)

4: k_3 = \dot{\boldsymbol{X}}(t(i) + \frac{dt}{2}, \ \boldsymbol{X}(:,i) + \frac{dt}{2}k_2)

5: k_4 = \dot{\boldsymbol{X}}(t(i) + dt, \ \boldsymbol{X}(:,i) + dtk_3)

6: \boldsymbol{X}(:,i+1) = \boldsymbol{X}(:,i) + \frac{dt}{6}(k_1 + 2k_2 + 2k_3 + k_4)

7: end for
```

Task 8: Control for Sun, Nadir, and Communication Pointing

Next, the control architecture must be developed. We shall use the linearized closed loop dynamics to determine the proportional and derivative gains.

Let us consider the PD nonlinear feedback control law:

$$\boldsymbol{u} = -K\boldsymbol{\sigma} - [P]\delta\boldsymbol{\omega} + [I]\left(\dot{\boldsymbol{\omega}}_r - [\boldsymbol{\omega}^{\times}]\boldsymbol{\omega}_r\right) + [\boldsymbol{\omega}^{\times}][I]\boldsymbol{\omega} - \boldsymbol{L}$$
(31)

Let us disregard the external torque modeling error and consider the displacement MRP and deviation angular rate with the following control:

$${}^{\mathcal{B}}\boldsymbol{u} = -K\boldsymbol{\sigma}_{B/R} - P^{\mathcal{B}}\boldsymbol{\omega}_{B/R} \tag{32}$$

Let us look at the closed loop dynamics with this state vector:

$$x = \begin{bmatrix} \sigma \\ \delta \omega \end{bmatrix} \tag{33}$$

Using the differential kinematic equation for the MRPs we arrive at the following nonlinear state space formulation, with the closed loop full-state feedback:

$$\dot{\boldsymbol{x}} = \begin{bmatrix} \dot{\boldsymbol{\sigma}} \\ \delta \dot{\boldsymbol{\omega}} \end{bmatrix} = \begin{bmatrix} 0 & \frac{1}{4}B(\boldsymbol{\sigma}) \\ -K[I]^{-1} & -[I]^{-1}[P] \end{bmatrix} \begin{bmatrix} \boldsymbol{\sigma} \\ \delta \boldsymbol{\omega} \end{bmatrix}$$
(34)

The small angle approximation is employed to linearize this formulation with the following:

$$\begin{bmatrix} \dot{\boldsymbol{\sigma}} \\ \delta \dot{\boldsymbol{\omega}} \end{bmatrix} = \begin{bmatrix} 0 & \frac{1}{4} \mathbb{I}_3 \\ -K[I]^{-1} & -[I]^{-1}[P] \end{bmatrix} \begin{bmatrix} \boldsymbol{\sigma} \\ \delta \boldsymbol{\omega} \end{bmatrix}$$
(35)

We can find the eigenvalues of this matrix to see how the gains affect the stability of the system. The roots act as the following:

$$\lambda_i = \frac{-1}{2I_i} \left(P_i \pm \sqrt{-KI_i + P_i^2} \right) \quad i = 1, 2, 3$$
 (36)

Now, we can choose the gains such that they meet our operation criterion. We want to use single scalar gains for both K and P. We know that we have a decay time of T=120 seconds which constraints out P value with the following:

$$P = \max_{i} \left(\frac{2I_i}{T} \right) \tag{37}$$

Once we have constrained this gain value, we can take a look at the K gain and damping coefficient for each of the modes. We know the relationship $\xi_i = \frac{P}{\sqrt{K I_i}}$. We can find a solution for the critically damped $\xi = 1$ mode, then pick an inertia that results in the other modes being critically or underdamped, where $\xi \leq 1$. We end with the result that $[P \ K] = [0.1666 \ 0.0055]$.

Task 8: Sun Pointing Plots

Let us now simulate the LMO spacecraft performing the sun-point maneuver. We calculate the tracking errors with the function developed earlier and use the reference DCM and angular rate vector developed earlier. These tracking errors are used in our control function with the gains determined in the previous step. With our determined gains, we see how the system tracks the reference for a duration of 400 seconds:

Task 9: Nadir (Science Mode) Pointing Plots

Now we shall simulate the LMO spacecraft performing purely the Mars-nadir pointing mode, where it gathers data by measuring the surface. The only difference between this mode and the sun-pointing mode are our new DCM and angular rate references. Similarly, we simulate for 400 seconds:

Task 10: GMO Spacecraft Communication Pointing Plots

Similarly, we shall apply the control law to the GMO communication-pointing mode. In this mode, the spacecraft tracks the GMO reference attitude and angular rate, in order to uplink information for transmission to Earth. This concludes our individual mode control and we can move onto the full mission simulation with mode mode switching. The 400 second simulation is shown in figure 3(a) with tracking error in 3(b).

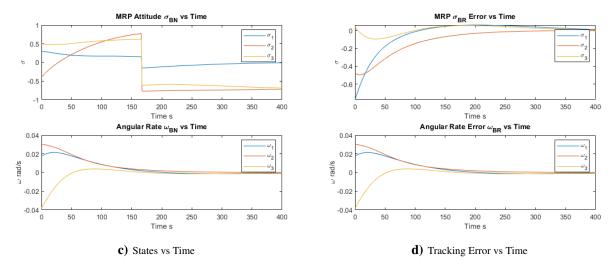


Figure 1: This figures show the states and state tracking errors over times for a 400 second simulation of sunpointing.

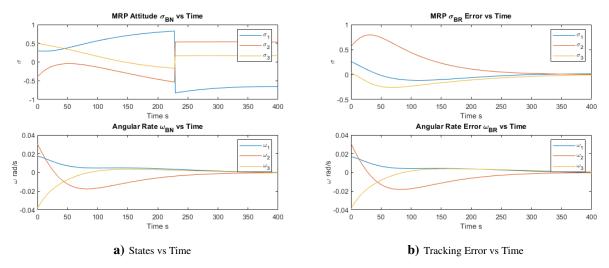


Figure 2: This figures show the states and state tracking errors over times for a 400 second simulation of nadirpointing.

Task 11: Mode Switching Mission Scenario

Next, we must simulate the full mission scenario with mode switching. When the inertial position vector of the LMO spacecraft is positive, or it is in the sun-lit region of the planet, it must point towards the sun to gather power.

I. Conclusion Acknowledgment References

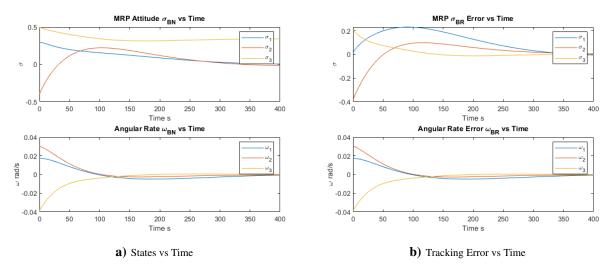


Figure 3: This figures show the states and state tracking errors over times for a 400 second simulation of GMO-pointing.

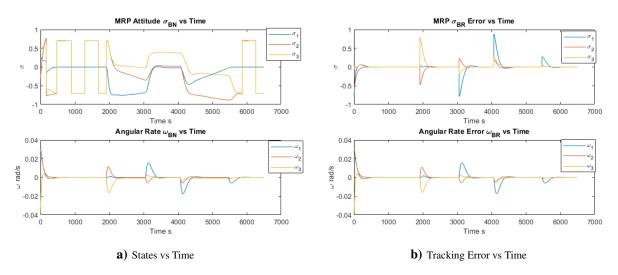


Figure 4: This figures show the states and state tracking errors over times for the full 6500 second simulation of operation.

Listing 1: Main Script (Simulation Framework)

```
%% Padraig Lysandrou April 4th 2019 -- ASEN5010 Final Project
   clc; close all; clear all;
2
   % Problem specified dynamic initial conditions
   sigma_BN_0 = [0.3 - 0.4 0.5].';
5
   omega_BN_B_0 = deg2rad([1 1.75 -2.2]).';
   Ic = diag([10 5 7.5]); p.Ic = Ic;
   % Problem specified orbital parameters
   mu_M = 42828.3;
                                 % km3/s2
10
  h_LMO = 400;
                                 % km
  RM = 3396.19;
                                 % km
12
13 \text{ r_LMO} = \text{h_LMO} + \text{R_M};
14 \text{ r\_GMO} = 20424.2;
   Omega_GMO = 0; inc_GMO = 0; theta_GMO_0 = deg2rad(250);
16 Omega_LMO = deg2rad(20); inc_LMO = deg2rad(30); theta_LMO_0 = deg2rad(60);
in theta_dot_LMO = sqrt(mu_M / (r_LMO^3)); % rad/s, orbital angular rate(circ)
```

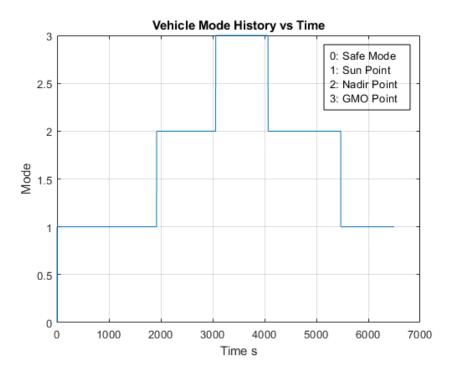


Figure 5: This figure shows the modes of the spacecraft over the duration of the mission.

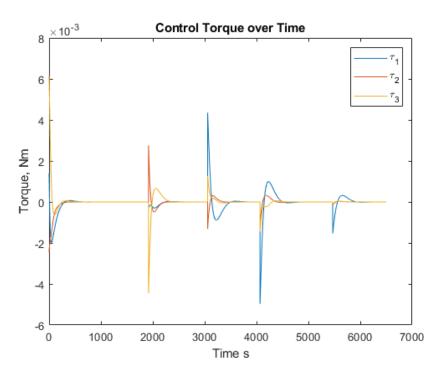


Figure 6: this figure shows the on-axis control torques in Nm.

```
18 theta_dot_GMO = sqrt(mu_M / (r_GMO^3)); % rad/s, orbital angular rate(circ)
20 % theta_dot_GMO = 0.0000709003;
21
22 % Take care of timing
23 tend = 6500:
24 	 dt = 1;
25 t = 0:dt:tend;
26  npoints = length(t);
28 % Set up the dynamic function as well as state tracking, init conds
29 f_dot = @(t_in,state_in,param) dynamics(t_in,state_in,param);
30 vehicle_state = zeros(6, npoints);
vehicle_state(:,1) = [sigma_BN_0; omega_BN_B_0];
p.L = [0.0 \ 0 \ 0].';
33 p.u = [0 0 0].';
34 % vehicle_mode_history = zeros(1, npoints);
35
36 % Setup the tracking error stuff, and input history stuff
37 sigma_BR = zeros(3, npoints);
38 omega_BR = zeros(3, npoints);
39 control_input = zeros(3, npoints);
40 sig_int = [0 0 0].';
42 % Set the initial conditions
43 H = zeros(3, npoints);
44 H(:,1) = Ic*omega_BN_B_0;
45 T = zeros(1, npoints);
46 T(1) = 0.5*omega_BN_B_0.'*Ic*omega_BN_B_0;
48 % PD controller gain initialization
49 P = \max(\text{diag}(Ic) \cdot *(2/120));
50 K = (P^2)./Ic(2,2); % still not sure why this is the gain he wants...
52 \% xi_1 = sqrt((P^2)./(K.*Ic(1,1))) \% keeps <math>\leq 1
si_2 = sqrt((P^2)./(K.*Ic(2,2))) % does not keep < 1
54 % xi_3 = sqrt((P^2)./(K.*Ic(3,3))) % does not keep \leq 1
55
57 %% Start the simulation
vehicle_mode = 0;
s9 vehicle_mode_history = vehicle_mode;
60
61 tic
62 for i = 1:npoints-1
       % Pull out state values to be used below
       sigma_BN = vehicle_state(1:3,i);
64
       omega_BN = vehicle_state(4:6,i);
65
66
       % Determine the inertial small sat and GMO vectors
67
68
       theta_LMO = theta_LMO_0 + t(i)*theta_dot_LMO;
       [rN_LMO,¬] = oe2rv_schaub(r_LMO,mu_M,Omega_LMO,inc_LMO,theta_LMO);
69
       theta_GMO = theta_GMO_0 + t(i)*theta_dot_GMO;
       [rN_GMO,¬] = oe2rv_schaub(r_GMO, mu_M, Omega_GMO, inc_GMO, theta_GMO);
71
73
       % Determine the mode of the spacecraft
74
       if rN_LMO(2) \ge 0
           % put the system into sun pointing mode
75
           vehicle_mode = 1;
76
       elseif acosd((rN_LMO.'*rN_GMO)/(norm(rN_LMO)*norm(rN_GMO))) < 35</pre>
           % put the system into GMO data transfer mode
78
           vehicle_mode = 3;
79
80
       else
           % put the system into nadir point science mode
81
82
           vehicle_mode = 2;
83
84
       vehicle_mode_history = [vehicle_mode_history vehicle_mode];
85
```

```
86
87
       % Determine the trackqin DCM/Omega based on vehicle mode
        switch vehicle_mode
89
            case 1
                        % Sun pointing energy gather mode
                RN = dcm_RsN(t(i));
91
                omega_RN = omega_RsN(t(i));
92
93
                [sigma_BR(:,i), omega_BR(:,i)] = ...
                    track_error(sigma_BN,omega_BN,RN,omega_RN);
94
            case 2
                       % Nadir pointing science mode
                RN = dcm_RnN(t(i));
96
                omega_RN = omega_RnN(t(i),theta_dot_LMO);
                [sigma_BR(:,i), omega_BR(:,i)] = ...
98
99
                    track_error(sigma_BN, omega_BN, RN, omega_RN);
                     % GMO comm pointing mode
100
                RN = dcm_RcN(t(i));
101
                omega_RN = omega_RcN(t(i));
                [sigma_BR(:,i), omega_BR(:,i)] = ...
103
                    track_error(sigma_BN,omega_BN,RN,omega_RN);
104
105
                      % Safemode hold
                sigma_BR(:,i) = [0 \ 0 \ 0].';
106
                omega_BR(:,i) = [0 \ 0 \ 0].';
107
            otherwise % Safemode hold
108
                sigma_BR(:,i) = [0 \ 0 \ 0].';
109
                omega_BR(:,i) = [0 \ 0 \ 0].';
110
111
112
        % Calculate the control input from mode error
113
        sig_int = sig_int + dt.*(sigma_BR(:,i));
114
       p.u = (-K.*sigma_BR(:,i)) + (-P.*omega_BR(:,i)); % + -0.01.*sig_int;
115
       control_input(:,i) = p.u;
116
117
118
119
       % Pull out the angular momentum and the energy for debugging and
       % verification
120
       H(:,i) = Ic*omega_BN;
121
       T(i) = 0.5*omega_BN.'*Ic*omega_BN;
122
123
       % RK4 step for the spacecraft dynamics
124
       k_1 = f_{dot}(t(i), vehicle_state(:,i),p);
125
        k_2 = f_{dot}(t(i)+0.5*dt, vehicle_state(:,i)+0.5*dt*k_1,p);
126
       k_3 = f_{dot}((t(i)+0.5*dt), (vehicle_state(:,i)+0.5*dt*k_2), p);
127
128
       k_4 = f_{dot}((t(i)+dt), (vehicle_state(:,i)+k_3*dt), p);
        \text{vehicle\_state(:,i+1)} = \text{vehicle\_state(:,i)} + (1/6) * (k_1 + (2*k_2) + (2*k_3) + k_4) * dt; 
129
130
       % Perform the nonsingular MRP propagation attitude check
       s = norm(vehicle_state(1:3,i+1));
132
       if s > 1
133
            vehicle\_state(1:3,i+1) = -(vehicle\_state(1:3,i+1) ./(s^2));
134
135
136 end
137
   t.oc
139
140 %% Plot data here.
141 close all;
142
143 % figure; plot(t, H); title('angular momentum');
144 figure; subplot(2,1,1);
145 plot(t, vehicle_state(1:3,:)); title('MRP Attitude \sigma_{BN} vs Time');
146 xlabel('Time s'); ylabel('\sigma'); legend('\sigma_1','\sigma_2','\sigma_3');
147 subplot (2,1,2);
148 plot(t, vehicle_state(4:6,:)); title('Angular Rate \omega_{BN} vs Time')
149 xlabel('Time s'); ylabel('\omega rad/s'); legend('\omega_1','\omega_2','\omega_3');
iso % figure; plot(t, vecnorm(H)); title('Angular Momentum over time')
152 % figure; semilogy(t, T); title('system angular energy over time')
153 figure; subplot(2,1,1);
```

```
plot(t, sigma_BR); title('MRP \sigma_{BR} Error vs Time');
155 xlabel('Time s'); ylabel('\sigma'); legend('\sigma_1','\sigma_2','\sigma_3');
156 subplot (2,1,2);
is plot(t, omega_BR); title('Angular Rate Error \omega_{BR} vs Time')
is xlabel('Time s'); ylabel('\omega rad/s'); legend('\omega_1','\omega_2','\omega_3');
160 figure; plot(t, vehicle_mode_history); title('Vehicle Mode History vs Time')
xlabel('Time s'); ylabel('Mode');
162 grid on; hold on;
dim = [0.7 \ 0.6 \ 0.3 \ 0.3];
164 str = {'0: Safe Mode','1: Sun Point','2: Nadir Point', '3: GMO Point'};
annotation('textbox',dim,'String',str,'FitBoxToText','on');
166 hold off:
168 figure; plot(t, control_input); title('Control Torque over Time');
169 xlabel('Time s'); ylabel('Torque, Nm');
170 legend('\tau_1','\tau_2','\tau_3');
171
172 % Saving things for Checkoff 7
173 % save_to_txt('H500B.txt', H(:,end));
174 % save_to_txt('T500.txt', T(end));
175 % save_to_txt('MRP500.txt',vehicle_state(1:3,end));
% save_to_txt('H500N.txt', MRP2C(vehicle_state(1:3,end)).'*H(:,end));
178 % close all;
179 % Saving things for Checkoff 8-10:
180 % save_to_txt('gains.txt', [P K]);
181 % save_to_txt('MRP15.txt', vehicle_state(1:3,16));
182 % save_to_txt('MRP100.txt', vehicle_state(1:3,101));
184 % save_to_txt('MRP400.txt', vehicle_state(1:3,401));
185
186 % close all;
187 % Saving things for Checkoff 11:
188  % save_to_txt('MRP300.txt', vehicle_state(1:3,301));
189 % save_to_txt('MRP2100.txt', vehicle_state(1:3,2101));
190 % save_to_txt('MRP3400.txt', vehicle_state(1:3,3401));
```

Listing 2: System Dynamics RHS

```
function f_x = dynamics(¬, state_in, p)
  %%This function defines all of the system dynamics
       % pull out struct constants
3
       Ic = p.Ic;
       % distribute the states for use! these are row vectors.
       sigma_BN = state_in(1:3);
omega_BN = state_in(4:6);
7
8
       % MRP integration, remember to perform norm check on this badboi: 3X1
10
       sigma_dot = (0.25.*((1 -(sigma_BN.'*sigma_BN))*eye(3) + 2*skew(sigma_BN.') ...
11
12
           + (2*sigma_BN*(sigma_BN.'))))*omega_BN;
13
       % Angular Rate Integration
14
       omega_dot = Ic\((-skew(omega_BN)*(Ic*omega_BN)) + p.L + p.u);
15
16
       % Send out the derivative
17
       f_x = [sigma_dot.' omega_dot.'].';
18
19 end
```

Listing 3: Tracking Error Determination

```
1 function [sigma_BR, omega_BR] = track_error(sigma_BN,omega_BN,RN,omega_RN)
2 %track_error return attitude and angular rate tracking error
3
4 % first tackle the error in attitude...
5 % sigma_RN = C2MRP(RN);
6 % sigma_BR = relMRP(sigma_BN.', sigma_RN.');
7 BN = MRP2C(sigma_BN);
8 sigma_BR = C2MRP(BN*(RN.'));
9 omega_BR = (omega_BN - BN*omega_RN);
10 end
```

Listing 4: Tracking Error Determination

```
function [rN, vN] = oe2rv_schaub(r, mu, Omega, incl, theta)
   %oe2rv JUST FOR THIS PROJECT: NOT ROBUST FOR ANYTHING
       R3\_omega = [cos(theta) sin(theta) 0
                   -sin(theta) cos(theta) 0
4
                           0
                                  1];
5
       R3\_Omega = [cos(Omega) sin(Omega) 0
6
                   -sin(Omega) cos(Omega) 0
                             0
                                     1];
       R1_{inc} = [1]
                          0
                                      0
9
               0 cos(incl) sin(incl)
0 -sin(incl) cos(incl)];
10
11
       % Take perifocal coordinates and transform to ECI coords
12
       perif_to_ECI = (R3_omega*R1_inc*R3_Omega).';
13
       rN = perif_to_ECI*([r 0 0].');
14
       % This is an angular rate that only works for circular orbits:
15
       % basically just mean motion
16
       n = sqrt(mu/r^3);
17
18
       vN = perif_to_ECI* ([0 r*n 0].');
19 end
```