



# Monte Carlo Theory & Practice (MCNP)

For Radiation Transport Solutions  
Lecture 3  
Dr Dennis Allen

# Lecture 3 Summary

- Random numbers
- Probability distribution sampling
- Particle interactions
- More on tallies

# Random Number Generation

- What is Monte Carlo famous for?
- Random number generation
  - The heart of any Monte Carlo method

**DILBERT** By SCOTT ADAMS



*“Random number generation is too important to be left to chance”*

Robert R Coveyou, Research Mathematician, ORNL

Pseudo-random number generation

Pseudo RNG (or deterministic RNG)

Algorithm for generating a sequence of numbers whose distribution appears to be random

Needs to be “seeded” to start the sequence

Can we ever have a true RNG?

# MCNP RNG

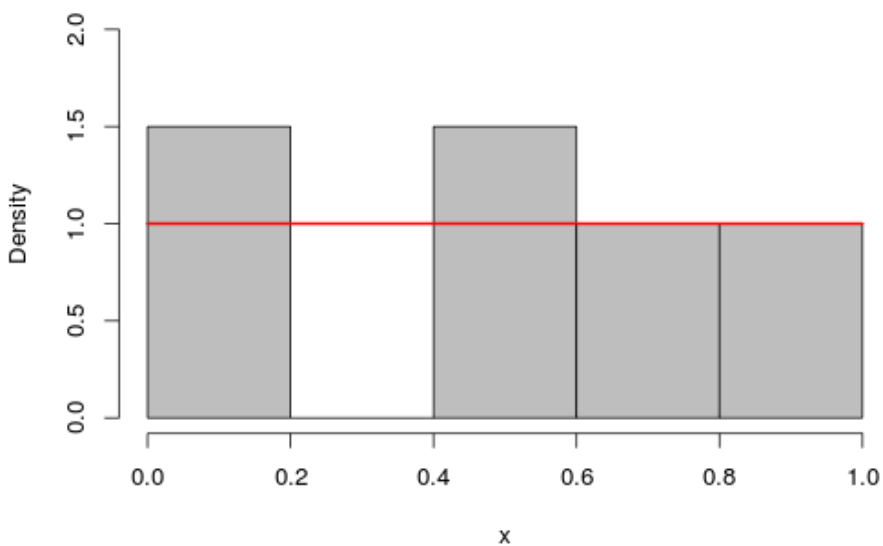
- MCNP uses a pseudo-RNG
- The implementation uses 48-bit integers and has a “period” of  $2^{46}$  values  $\sim 7.04 \times 10^{13}$
- This can be exceeded for some problems!
- Random number “stride”
  - The number of random numbers between those used to select source particles

# Random Number Generation

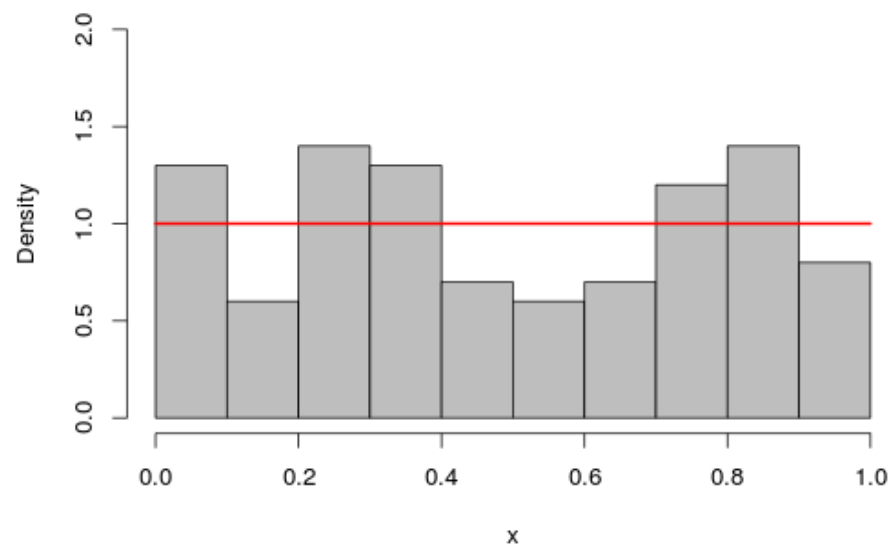
- RNGs aim to generate a sequence of random numbers within a fixed range whose distribution is uniform
- $\xi$  generated “random” number
  - Usually within the unit interval  $0 \leq \xi < 1$
- $P(\xi)$  the probability of selecting  $\xi$
- Clearly  $P(\xi)$  is chosen so that

$$\int_0^1 P(\xi) d\xi = 1$$

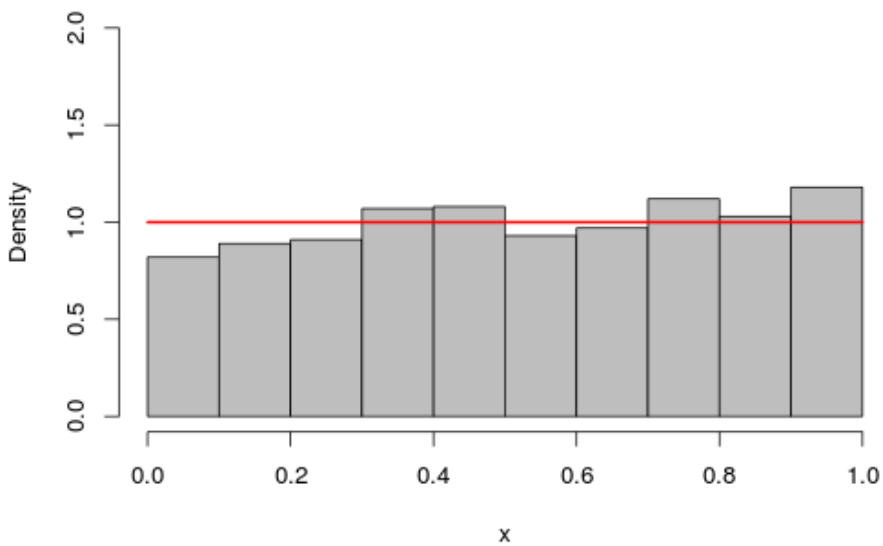
**10 Draws**



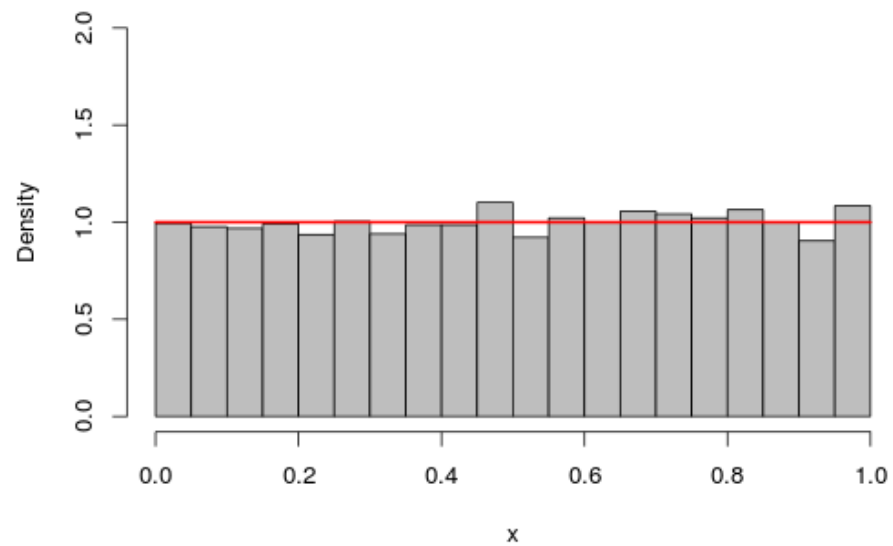
**100 Draws**



**1000 Draws**



**10000 Draws**

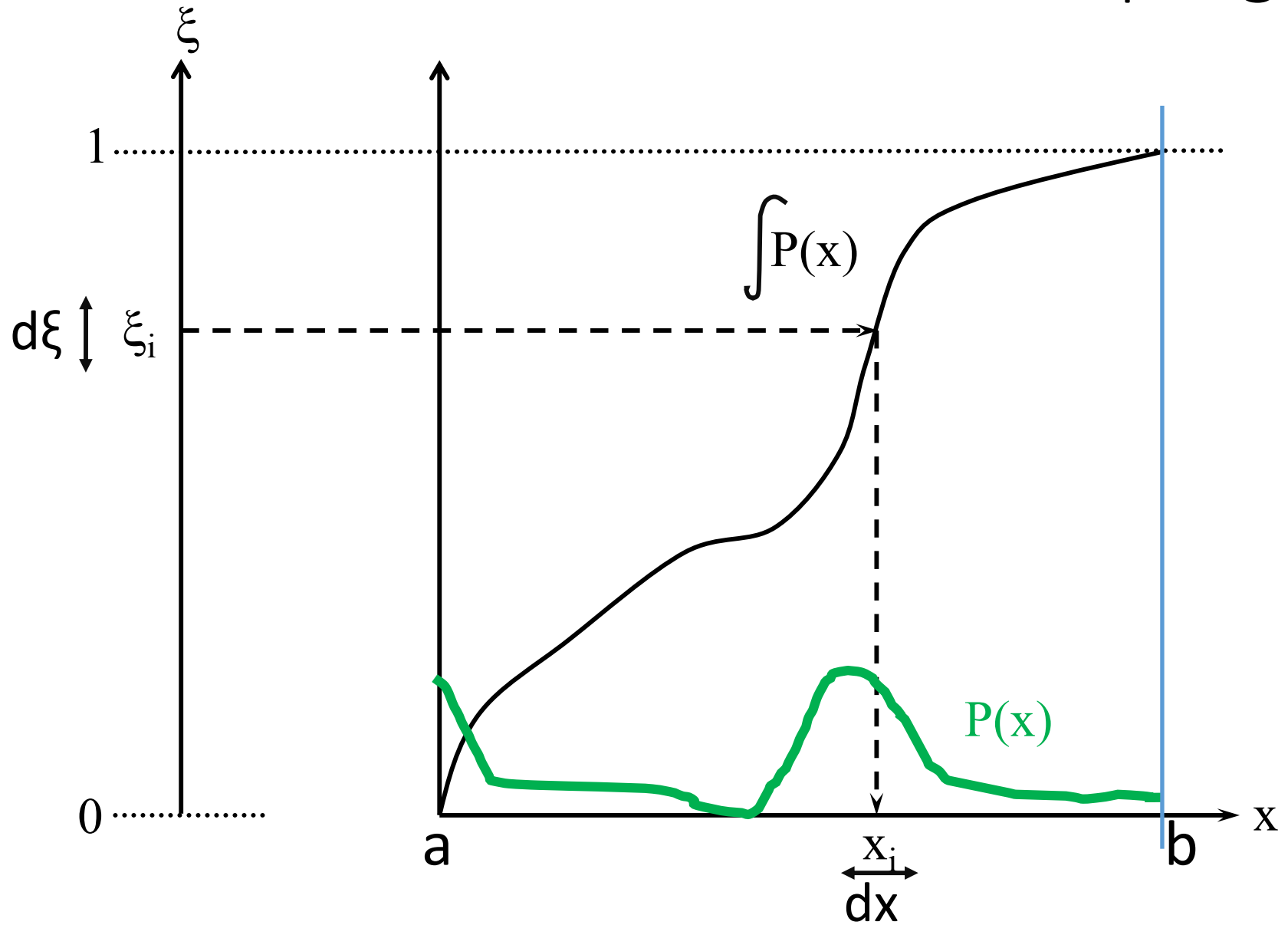


# Non-Uniform Sampling

- Random variable  $x$  defined in the range  $a \leq x \leq b$
- We think we know the probability distribution  $P(x)$
- Wish to generate a series of random values of  $x$  to match this distribution
- We have a pseudo-RNG to generate  $\xi$  within the range  $0 \leq \xi < 1$



# Non-Uniform Sampling



# Non-Uniform Sampling

- Transform from  $\xi$  space to  $x$  space by equating  
probability of obtaining  $x'$  in range  $dx'$  about  $x'$   
= probability of obtaining  $\xi$  in range  $d\xi$  about  $\xi$   
Integrate from lower limit to a value  $x$

$$\int_a^x P(x') dx' = \int_0^{\xi} d\xi' = \xi$$

- Random Sampling Rule

# Applied to Radiation Transport

- The cross-section
- Describes the probability of a particle interaction

$$\Sigma_T = \Sigma_s + \Sigma_a$$

- Macroscopic cross-sections are required

$$\Sigma(\text{cm}^{-1}) = \sigma(\text{cm}^2) n(\text{atoms/cm}^3)$$

$$n(\text{atoms/cm}^3) = \rho(\text{g/cm}^3) N_A \div A(\text{g/mol})$$

# Distance to Next Collision

- Basic Monte Carlo question
  - Where is the next collision?
  - *Diagram of scatter point*
- Probability of collision within  $dr'$  about  $r'$  is  $P(r')dr'$

$$P(r') = \Sigma_T e^{-\Sigma_T r'}$$

- Use random sampling rule

$$\int_0^r P(r') dr' = \int_0^\xi d\xi' = \xi$$

# Distance to Next Collision

$$\int_0^r \Sigma_T e^{-\Sigma_T r'} dr' = \int_0^\xi d\xi' = \xi$$
$$= 1 - e^{-\Sigma_T r}$$

- Rearrange to

$$r = \frac{-\ln(1 - \xi)}{\Sigma_T} \equiv \frac{-\ln(\xi)}{\Sigma_T}$$

# Where is the Collision?

- Distance to cell boundary =  $d$
- If  $r < d$  collision occurs within cell
- If  $r > d$  particle leaves cell and enters next one
  - Need to recalculate new  $d$  and next collision distance ( $r$ )
  - New cell may have different cross-sections

# Then What?

- We know where it collides, but what happens next?
- Choose which atom it collides with
  - For a material composed of several types of atom
  - a, b, c, ... (isotopes considered separately for neutrons)

$$\Sigma_T = \Sigma_{Ta} + \Sigma_{Tb} + \Sigma_{Tc} + \dots$$

Normalise by dividing by  $\Sigma_T$  so sum now = 1

Generate  $\xi$  and select from a, b, c, etc.

# Then What?

- Depends on particle type – photon, neutron, low energy neutron
- Select type of collision in the same way
- E.g. for photon we have to choose from:
  - Compton scatter, photo-electric absorption & pair production

$$\Sigma_T = \Sigma_{cs} + \Sigma_{pe} + \Sigma_{pp}$$

Normalise by dividing by  $\Sigma_T$  so sum now = 1

Generate  $\xi$  and select from cs, pe & pp



# Neutron Interactions

- For neutrons, in general there are many to choose from:
  - Elastic scatter  $(n,n)$
  - Inelastic scatter  $(n,n')$
  - Absorption  $(n,abs) - (n,\gamma), (n,\alpha), (n,d), (n,t), (n,p), \dots$
  - Multiplication  $(n,2n), (n,3n) \dots$
  - Fission  $(n,f)$  - this too implies further neutron production
- The collision isotope determines which ones are available and what the cross-sections are
  - Also depends on neutron energy

# Neutron Interactions - MCNP

- Identify collision nuclide
- Is  $S(\alpha, \beta)$  scattering required?
  - Low energy neutrons only – material dependent
- Are any photons generated?
  - Depends on mode of problem and collision type.
- Is neutron captured?
- Is collision elastic, inelastic or a fission event
  - Determines energy of ongoing neutron
  - Whether more particles (photons & neutrons) are generated
- For neutrons  $E < 4\text{eV}$ , need to take thermal motion of collision atoms into account.
  - Temperature matters
  - Target nuclide velocity sampling required – subtracted from velocity of incoming neutron

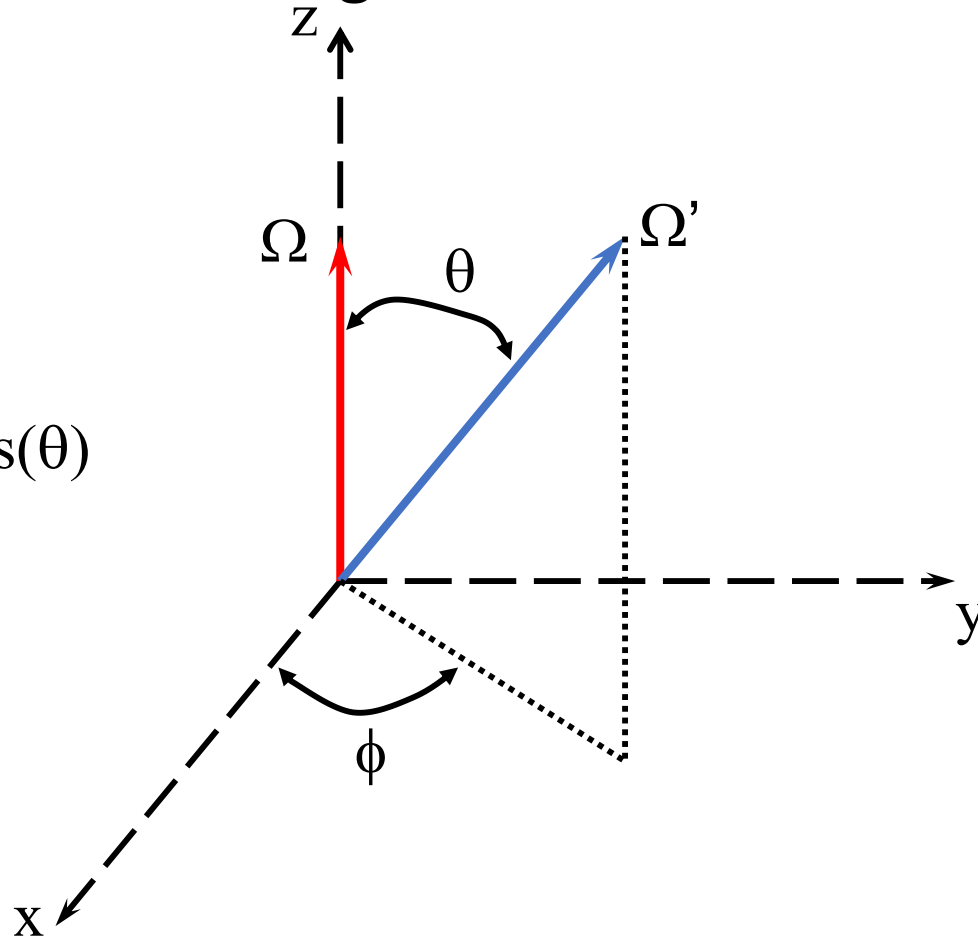
# Absorption

- Can lead to a number of potential events.
  - $(n,\gamma)$ ,  $(n,\alpha)$ ,  $(n,p)$ ,  $(n,d)$ ,  $(n,2n)$ ,  $(n,f)$ , ....
- Gammas are tracked (in MODE N P)
- Fissions lead to more neutrons
  - By default MCNP will track fission-generated neutrons
  - Can be switched off
- All particles generated are “banked” if required to be tracked
  - Record location of production
  - Sample direction and energy later

# Scattering - Selection of Scattering Angle

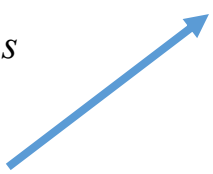
Defining the scatter angles:

Define:  $\mu = \cos(\theta)$



# Angle of Scatter (polar)

Probability of scattering within  $d\mu$  about  $\mu$

$$P(\mu) = \int_{-1}^{\mu} \frac{\Sigma_s(\mu') 2\pi d\mu'}{\Sigma_s} = \int_{-1}^{\mu} \frac{d\mu'}{2} = (\mu + 1) / 2$$


For isotropic scattering

$\Sigma_s(\mu)$  is constant and  $(\mu + 1) / 2 = \xi$

$$\Sigma_s(\mu) / \Sigma_s = 1 / 4\pi \qquad \mu = 2\xi - 1$$

For non-isotropic scatter use random sampling rule

$$\int_a^x P(x') dx' = \int_0^{\xi} d\xi' = \xi$$

# Azimuthal Scattering Angle

$$P(\phi) = \int_0^{\phi} \frac{1}{2\pi} \cdot d\phi' = \frac{\phi}{2\pi}$$

$$\frac{\phi}{2\pi} = \xi$$

$$\phi = 2\pi\xi$$

# Energy After Collision

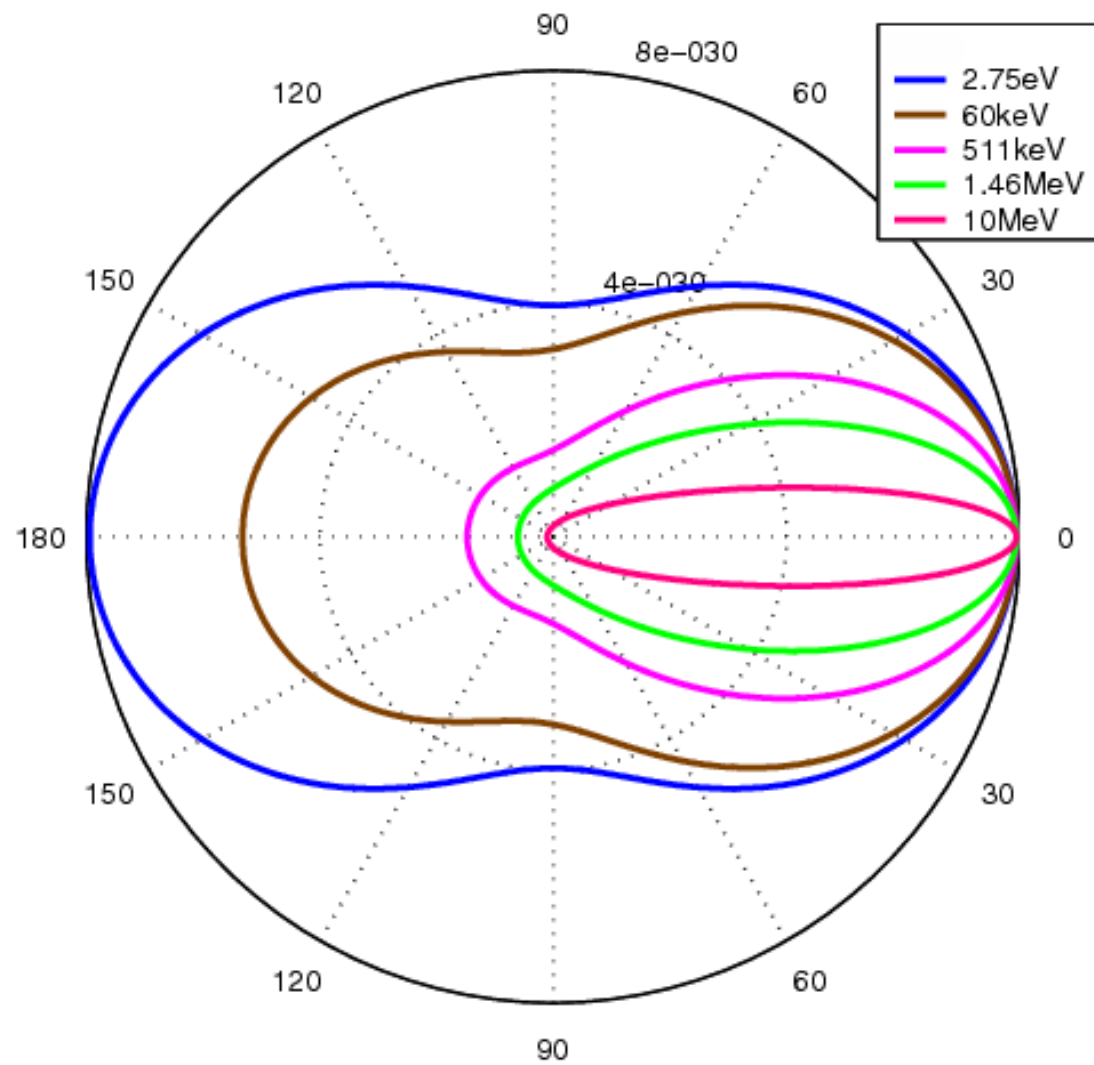
- For neutrons (elastic scatter)

$$E = E_{in} \left[ \frac{1 + A^2 + 2A \cos \theta}{(1 + A)^2} \right]$$

- For photons – Klein Nishina formula

$$\frac{d\sigma}{d\Omega} = \alpha^2 r_c^2 P(E_\gamma, \theta)^2 [P(E_\gamma, \theta) + P(E_\gamma, \theta)^{-1} - 1 + \cos^2(\theta)]/2$$

$$P(E_\gamma, \theta) = \frac{1}{1 + (E_\gamma/m_e c^2)(1 - \cos \theta)}$$





# Basic Flux Tally

$$\text{Flux } \phi = \int_V \int_t \int_E \phi(\vec{r}, E, t) dE dt \frac{dV}{V}$$

- Each particle,  $i$ , is assigned a “weight”  $W_i$
- Track length  $S_i$
- Volume of cell  $V$
- $N$  source particles

$$\phi = \frac{1}{V N} \sum_i S_i w_i$$

# Variations on Tallies

**Fn4 - cell flux (track length)**

**\*Fn4 – cell energy-weighted flux**

$$\text{Flux } \phi = \int_V \int_t \int_E E \phi(\vec{r}, \vec{\epsilon}, t) d\epsilon dt \frac{dV}{V}$$

$$\phi = \frac{1}{V N_i} \sum E_i s_i w_i$$

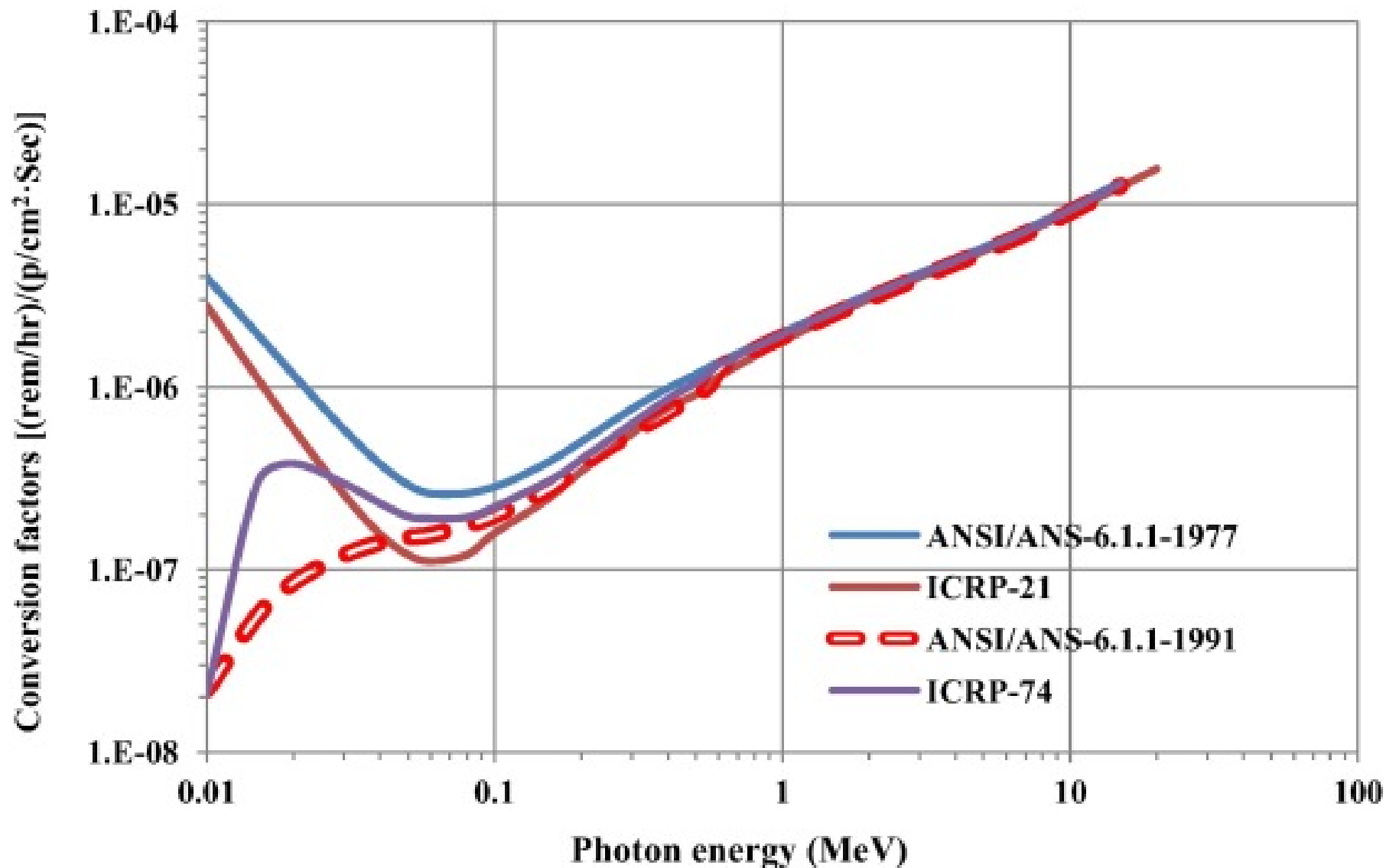
# Tallies – Summary of Units

Type	Particles	Description	Fn Units	*Fn Units
F1	N P E	Surface current	Particles	MeV
F2	N P E	Surface flux	Particles/cm <sup>2</sup>	MeV/cm <sup>2</sup>
F4	N P E	Cell flux	Particles/cm <sup>2</sup>	MeV/cm <sup>2</sup>
F5	N P	Point detector	Particles/cm <sup>2</sup>	MeV/cm <sup>2</sup>
F6	N P E	Energy deposition average over cell	MeV/g	jerks/g
F7	N	Fission energy deposited over a cell	MeV/g	jerks/g
F8	P E or P,E	Energy distribution of pulses in a detector	Pulses	MeV
+F8	E	Charge deposition in a cell	Charge	n/a

1 MeV =  $1.602191 \times 10^{-22}$  jerks

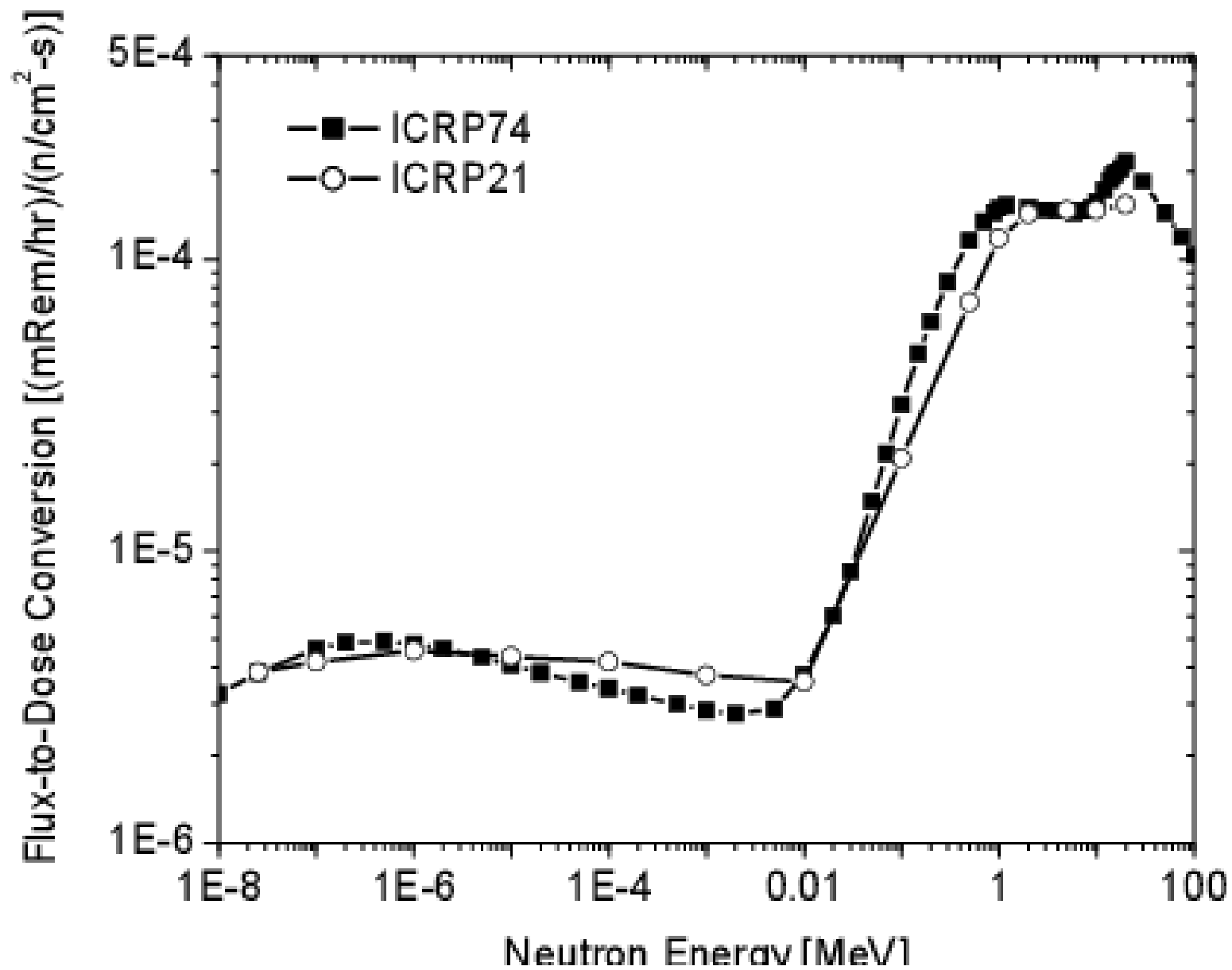
All tally types except 8 can be modified

# Response Functions



**Table H.2**  
**Photon Flux-to-Dose Rate Conversion Factors**

<b>ANSI/ANS-6.1.1-1977</b>		<b>ICRP-21</b>	
<b>Energy, E</b> <b>(MeV)</b>	<b>DF(E)</b> <b>(rem/hr)/(p/cm<sup>2</sup>-s)</b>	<b>Energy, E</b> <b>(MeV)</b>	<b>DF(E)</b> <b>(rem/hr)/(p/cm<sup>2</sup>-s)</b>
0.01	3.96E-06	0.01	2.78E-06
0.03	5.82E-07	0.015	1.11E-06
0.05	2.90E-07	0.02	5.88E-07
0.07	2.58E-07	0.03	2.56E-07
0.1	2.83E-07	0.04	1.56E-07
0.15	3.79E-07	0.05	1.20E-07
0.2	5.01E-07	0.06	1.11E-07
0.25	6.31E-07	0.08	1.20E-07
0.3	7.59E-07	0.1	1.47E-07
0.35	8.78E-07	0.15	2.38E-07
0.4	9.85E-07	0.2	3.45E-07
0.45	1.08E-06	0.3	5.56E-07



# Response Functions

- Tally Multipliers – Response functions
- Enable conversion from fluence to dose or anything

DEn  $E_1 E_2 E_3 \dots E_k$

DFn  $F_1 F_2 F_3 \dots F_k$

The values of  $F$  represent a pointwise function of energy ( $E$ ).  
Logarithmic (default) or linear interpolation is used between these values. Use DEn LIN  $E_1 E_2 E_3 \dots E_k$  for linear.

Can be used to calculate any required response function  
e.g. dose rate, damage rate, reaction rate, etc.

If value of response function at energy  $E$  is  $R(E)$ , then

$$\phi = \frac{1}{V N_i} \sum S_i w_i R(E_i)$$

# General Tally Multiplier FM Card

- Can specify several multipliers for one tally
  - MCNP reports the results for each
  - i.e. the same tracks are weighted in different ways to achieve different results
  - E.g. to tally several reaction rates in the same cell

*FMn (multiplier set 1) (multiplier set 2)*

*n* = tally number

*For each multiplier set*

*C m reaction list*

*C* = multiplicative constant (e.g. atom density)

*m* = material number (defined on a *Mm* card)

*reaction list* = list of reaction numbers required

see Appendix G of manual e.g. 102 = (n, $\gamma$ ), 103 = (n,p), 19 = (n,f)



# FM Example from Manual

Consider the following input cards.

F4:N	10		
FM4	0.04786	999	102
M999	92238.13	1	

$$C \int \phi(E) R_m(E) dE$$

The F4 neutron tally is the track length estimate of the average fluence in cell 10. Material 999 is  $^{238}\text{U}$  with an atomic fraction of 100%.

$C = 0.04786$	normalization factor (such as atom/barn·cm)
$M = 999$	material number for $^{238}\text{U}$ as defined on the material card (with an atom density of 0.04786 atom/barn·cm)
$R_1 = 102$	ENDF reaction number for radiative capture cross-section (microscopic)

The average fluence is multiplied by the microscopic  $(n,\gamma)$  cross section of  $^{238}\text{U}$  (with an atomic fraction of 1.0) and then by the constant 0.04786 (atom/barn·cm). Thus the tally 4 printout will indicate the number of  $^{239}\text{U}$  atoms/cm<sup>3</sup> produced as a result of  $(n,\gamma)$  capture with  $^{238}\text{U}$ .

- Cell 10 need not actually contain material 999
- You can pretend that it does

# Variations on Tallies

## **CFn - Cell flagging**

F12:N 5     \$ Flux of neutrons crossing surface 5

CF12 1 10 \$ Separate tally for those passing through either cell 1 or cell 10

## **SFn - Surface flagging**

F114:P 5     \$ Flux of photons in cell 5

SF114 100 \$ Separate tally for those which have crossed surface 100

## **Tally Segmentation**

Can be used to subdivide cells or surfaces into sub-tallies

F2:N 1 \$ neutron flux over surface 1

FS2 -3 \$ Separate tally for those inside surface 3 and those outside

# Other Tally Multipliers

- Energy multiplier

$EMn \quad M_1 M_2 M_3 \dots M_k$

$n$  = tally number

$M_k$  = multiplier for the  $k^{\text{th}}$  energy bin

- Time bin multiplier  $TMn$
- Cosine bin multiplier  $CMn$  (Type 1 only)

$EM0$ ,  $TM0$  or  $CM0$

- 0 tally bin multipliers apply to all tallies unless specifically overridden by another EM, TM or CM card