

## Monte Carlo Theory & Practice (MCNP)

For Radiation Transport Solutions Lecture 1 Dr Dennis Allen

## Course Summary

- Series of five lectures on Monte Carlo theory
  - Introduction the basics radiation transport, codes, applications
  - Random numbers their generation and use
  - Collision modelling
  - Scoring
  - Statistics interpreting the answers
  - Neutron interactions, Gamma interactions, Electrons?
    - Types on interaction, cross-sections, nuclear data
  - Acceleration
- Parallel MCNP workshops
  - Will therefore intersperse theory with MCNP tutorial
  - MCNP its main capabilities
  - MCNP "Building Blocks"
    - Sources, tallies, nuclear data
    - Analogue calculations, "accelerated" calculations
    - Fixed source calculations, neutron multiplication, criticality

## **Boltzmann Transport Equation**

$$\frac{1}{v}\frac{\partial\Phi}{\partial t} + \underline{\Omega}\underline{\nabla}\Phi + \underline{\Sigma}_t.\Phi = \int_{4\pi} \int \underline{\Sigma}_s(E' \to E, \Omega' \to \Omega)\Phi(E', \Omega')dE'd\Omega' + S'$$

Can you identify the terms?

## Radiation Transport

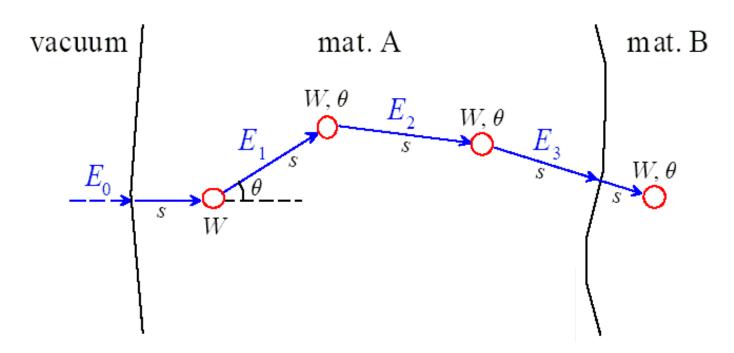
- Cannot usually be solved precisely approximations required
  - Diffusion equation
  - Discreet ordinates
  - Spherical harmonics
- Why would we want to solve it?
  - Examples?

### Monte Carlo - General

- But what if we think we know the physics of individual particle behaviour?
- We can model all the physical processes represented by the transport equation.
- Don't need to introduce approximations to solve it

## Particle Scattering

All interaction events are simulated in chronological succession:



### Monte Carlo - General

- Minimal approximations required
- Can define simple boundary conditions
- Can cope with voids
- But we do need to sample enough particles to
  - Produce good statistical accuracy in the region of interest
  - Make sure that all relevant regions of the problem are adequately sampled (space, energy, angle etc.), so that the simulation is representative of reality

## Law of Large Numbers

- Theoretical basis of Monte Carlo
- The weighted average value of the function, f

$$\bar{f} = \int_{a}^{b} f(x)p(x)dx = \lim_{N \to \infty} \frac{\sum_{i=1}^{N} f(x_i)}{N},$$

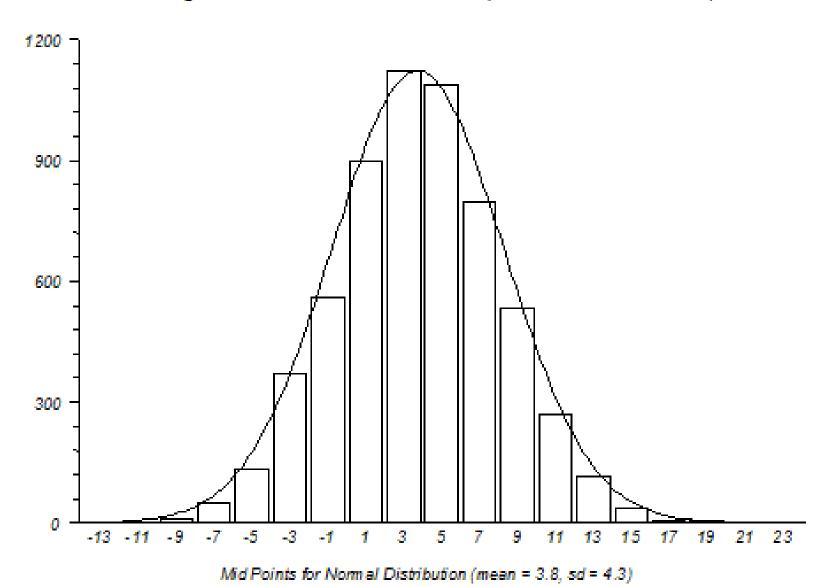
where  $x_i$  are chosen using p(x)

- This relates the result of a <u>continuous</u> integration to the result of a <u>discrete</u> sampling.
- All MC comes down to this

### Central limit theorem

- The sum of a sufficiently large number (N) of independent identically distributed random variables <u>eventually</u> becomes normally distributed as N tends to infinity
- This is useful for us because we can draw useful conclusions from the results from a large number of samples
  - (e.g., 68.7% within one standard deviation, etc.)
- But how large is large enough?

### Histogram for Normal Distribution (mean = 3.8, sd = 4.3)



## Monte Carlo for Radiation Transport

- What particles can we model with Monte Carlo?
  - Just about anything that we think we understand the physics of
    - Photons (x-rays & gammas)
    - Neutrons (fast, epithermal & thermal)
    - Electrons
    - Positrons
    - Muons
    - Protons
    - lons ...
  - Neutral or charged

### Some Monte Carlo Codes

Code	n	γ	e <sup>-</sup>	e <sup>+</sup>	p <sup>+</sup>	ions <sup>±</sup>
MCNP	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	
MCBEND	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$			
MONK	$\sqrt{}$					
TRIPOLI	$\sqrt{}$	$\sqrt{}$				
EGS5		$\sqrt{}$	$\sqrt{}$	$\sqrt{}$		
GEANT4	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	$\sqrt{}$	V
PENELOPE		$\sqrt{}$	$\sqrt{}$	$\sqrt{}$		
TRIM					$\sqrt{}$	$\sqrt{}$

## Introduction to MCNP

## MCNP – Main Capabilities

- Particles
  - Neutrons 10<sup>-5</sup> eV 20 MeV
  - Photons 1 keV 100 GeV
  - Electrons 1keV 1 GeV
- Full 3D geometry using standard surfaces
  - Repeated geometry features, lattices
- Highly flexible source definition
- Highly flexible scoring methods (tallies)
  - User-defined response functions (e.g. flux to dose)
  - Surface/cell flagging

## MCNP – The Building Blocks

- Cell "Cards"
  - Define 3D regions of space
  - Material within each regions, including density
  - Other attributes can also be defined here (importance)
- Surface "Cards"
  - Define the surface used to form the cells
  - Various possibilities
    - Planes, spheres, cones, cylinders, ellipsoids (hyperboloids & parabolas), general quadratic, tori
    - Special attributed can be defined reflecting or periodic
    - Can be transformed into different coordinate systems
- Data "Cards" everything else
  - Source definitions, acceleration details, tally definitions, tally modifiers, nuclear data ....

# MCNP — The Input File Structure MESSAGE BLOCK (optional)

TITLE CARD – Text to give your model a title CELL CARDS – define the cells

SURFACE CARDS – define the surfaces

DATA CARDS – define everything else C Comment cards – use anywhere

## Card Format (horizontal)

- Limited to 80 characters (columns) per line
  - Continuation lines can be used (end line with "&")
- Alphabetic characters can be upper or lower case
- Cell, surface & data cards begin in 1<sup>st</sup> 5 columns
- "C" in 1<sup>st</sup> column, followed by a blank indicates that whole line is a comment
- "\$" used to indicate the rest of the line is comment
- Entered in "blocks" cell, surface, data cards
- Blank lines used a delimiters between blocks
- Integers must be used where an integer is expected
- Where floating point is expected, any format can be used
  - E.g. 10.0 can be written as "10" or "1.0E+01" or "10.000"

## Surfaces - examples

#### Planes (unbounded)

```
1 PX 1.0 $ plane perpendicular to x-axis at x=1.0
```

```
2 PZ -10.2 $ plane perpendicular to z-axis at z=-10.2
```

```
6 P x1 y1 z1 x2 y2 z2 x3 y3 z3
```

plane passing through the 3 points with coordinates  $(x_1, y_1, z_1)$  etc.

### **Spheres (bounded)**

```
10 S0 100.1 $ sphere centred on origin with radius 100.1
```

```
11 SY 10.0 3.0 $ sphere centred on (0,10,0) of radius 3.0
```

All dimensions are in cm

All parameter need to be separated by at least one space

Only use spaces, DO NOT USE TABS!

Beware of other invisible characters

## Surfaces – more examples

### **Cylinders (unbounded)**

```
1 CY 1.5 $ cylinder along y axis, radius=1.5
```

```
10 C/Z 3.0 5.5 2.4 $ cylinder parallel to z axis passing through (3.0,5.5, ...), radius=2.4
```

### Cones – on x axis (unbounded)

```
100 KX x t2 \pm 1 or K/X x y z t2 \pm 1
```

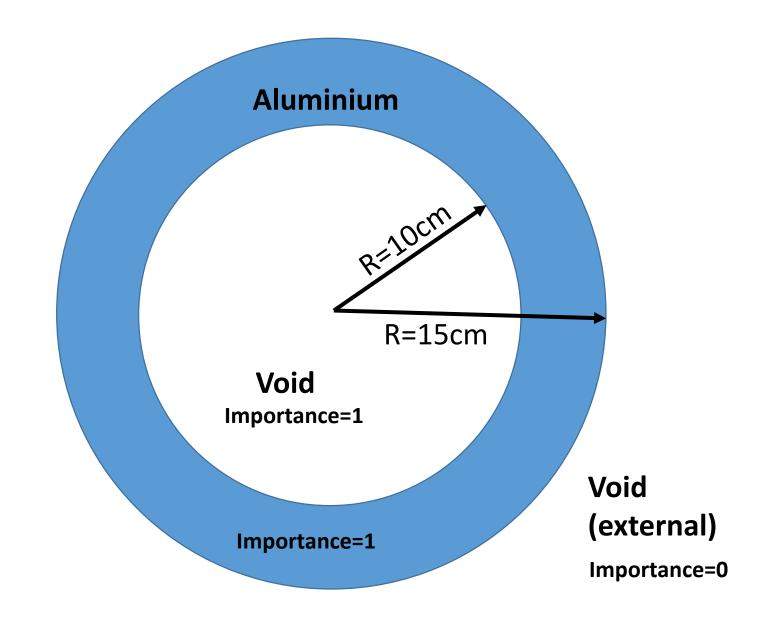
x defines vertex,  $t=tan(\theta)$ ,  $\pm 1$  (optional) defines direction of cone

```
101 KX 3. 5 0. 25 +1 $cone with vertex at (3.5,0,0), \theta=tan<sup>-1</sup>(0.5)
```

C double cone with vertex at (1,2,1.5),  $\theta$ =tan<sup>-1</sup>(0.3)

## MCNP Geometry

- All space must be defined as something
  - Up to the boundary of your problem
- All cells must:
  - be either void or allocated to a defined material with a chosen density
  - be defined using pre-defined surfaces
    - Using surface numbers and a combination of union (:), intersection () and complement (#) operators
  - have an "importance" for each particle type in the problem



A Simple Model of an Aluminium Shell Containing a Void C Model created 6/2/16 by DA Allen, last modified 19/11/16 C The Cell Cards

```
1 0 -1 $ Void cell inside surface 1
2 1 -2. 7 1 -2 $ Shell of m1 outside 1, inside 2, \rho=2.7g/cc
3 0 $ Void cell outside surface 2
```

### C The Surface Cards

```
1 S0 10.0 $ Sphere centred on origin, rad=10.0 2 S0 15.0 $ Sphere centred on origin, rad=15.0
```

C Data Cards – just a single material card in this case

```
MODE P
IMP:P 1 1 0
M1 13000 1.0 $ Aluminium, 100% pure
```

A simple model of an iron cylinder

C Model created 19/11/16 by DA Allen, last modified 19/11/16

### C The Cell Cards

- 1 1 -7.6 -1 2 -3 \$ Cylinder of m1,  $\rho$ =7.5g/cc
- 2 0 1:-2:3 \$ Void cell outside surface 2

### C The Surface Cards

- 1 CX 5.0 \$ Cylinder along x axis, radius=5.0
- 2 PX 0.0 \$ Plane perpendicular to x axis at x=0.0
- 3 PX 20.0 \$ Plane perpendicular to x axis at x=20.0

C Data Cards – just a single material card in this case

MODE F

IMP:P 1 0

M1 26000 1.0 \$ Iron, 100% pure

### Data Cards - Materials

Mm ZAID<sub>1</sub> fraction<sub>1</sub> ZAID<sub>2</sub> fraction<sub>2</sub> .... m is a unique material number ZAID is a code to represent the atoms of the material ZAID = zaaa.nnXz= atomic number (1 or 2 digits) aaa=3 digit atomic weight (can use 000 for photons) nn = a 2 digit number representing the nuclear data library X = C for continuous energy data (usually recommended), D for discreet energy library, m for multigroup cross-sections, y for "dosimetry" data

 $fraction_k$  is the atom fraction for the  $k^{th}$  listed atom If fraction is negative, then it is interpreted as weight fraction

## Mode – what is MCNP simulating?

MODE N neutrons only

MODE P photons only

MODE E electrons only

MODE N P neutrons and photons [can have neutron-induced photons,  $(n, \gamma)$ ]

MODE P E photons and electrons [photon-induced electrons, e.g. Compton scatter, pair production]

MODE N P E all three, including induced production