

# Homework 7

Michael Brodskiy

Professor: Q. Yan

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# 1 Symmetric Quantum Well

1. (a) If  $|\psi(x)|^2$  is symmetric (as specified), then  $\psi(x)$  is symmetric about the origin as well. We know that  $\psi(-x) = \psi(x)$  or  $\psi(-x) = -\psi(x)$ . If neither of these cases is true, then  $|\psi(x)|^2$  can not be symmetric. For example, if  $\psi(-x) \neq \psi(x)$ , then:

$$|\psi(x)| \neq |\psi(-x)|$$

This means:

$$|\psi(x)|^2 \neq |\psi(-x)|^2$$

which contradicts the symmetry of  $|\psi(x)|^2$

Additionally, if  $\psi(-x) \neq -\psi(x)$ , the process becomes similar, as  $-\psi(x)$  becomes positive due to the absolute value sign.

Thus, it is necessary that  $\psi(-x) = \psi(x)$  or  $\psi(-x) = -\psi(x)$ , so that  $|\psi(x)|^2$  maintains its symmetry.

- (b) Assuming the solution to the function is  $A \cos(kx) + B \sin(kx)$ , and knowing that  $U(x)$  is zero:

$$\frac{\partial^2 \psi}{\partial x^2} = -Ak^2 \cos(kx) - Bk^2 \sin(kx)$$

$$\frac{\partial^2 \psi}{\partial x^2} = -k^2 \psi(x)$$

Thus, we get:

$$\frac{\hbar^2 k^2}{2m} \psi(x) = E \psi(x)$$

$$k = \frac{\sqrt{2mE}}{\hbar}$$

Given the above functions, we know that  $k$  must be:

$$k = \frac{n\pi}{L}, \quad n = 1, 2, 3, \dots$$

Using boundary conditions at  $\frac{L}{2}$ :

$$\psi\left(\frac{L}{2}\right) = A \cos\left(\frac{n\pi}{2}\right) + B \sin\left(\frac{n\pi}{2}\right)$$

When  $n$  is odd, this means  $A = 0$ , as  $\cos$  goes to zero, and, due to boundary conditions,  $B = 0$  as well; when  $n$  is even, this means  $B = 0$  as  $\sin$  goes to zero, and, due to boundary conditions,  $A = 0$  as well. Now applying normalization:

When  $n$  is odd:

$$\int_{-\frac{L}{2}}^{\frac{L}{2}} |B \sin\left(\frac{2\pi x}{L}\right)|^2 dx = 1 \Rightarrow b = \sqrt{\frac{L}{2}}$$

When  $n$  is even:

$$\int_{-\frac{L}{2}}^{\frac{L}{2}} |A \cos\left(\frac{\pi x}{L}\right)|^2 dx = 1 \Rightarrow A = \sqrt{\frac{L}{2}}$$

Thus, the equation becomes:

$$\begin{cases} \sqrt{\frac{2}{L}} \cos\left(\frac{n\pi x}{L}\right), & \text{for } n \text{ odd} \\ \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right), & \text{for } n \text{ even} \end{cases}$$

(c) Given the wave functions from (b) and applying them to the wave function:

$$\begin{cases} \sqrt{\frac{2}{L}} \cos\left(\frac{n\pi x}{L}\right) \rightarrow -\frac{n\pi}{L} \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right) \rightarrow -\frac{n^2\pi^2}{L^2} \sqrt{\frac{2}{L}} \cos\left(\frac{n\pi x}{L}\right) \\ \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right) \rightarrow \frac{n\pi}{L} \sqrt{\frac{2}{L}} \cos\left(\frac{n\pi x}{L}\right) \rightarrow -\frac{n^2\pi^2}{L^2} \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right) \end{cases}$$

Putting this into the wave function, we get:

$$-\frac{\hbar^2}{2m} \left( -\frac{n^2\pi^2}{L^2} \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right) \right) = (E - U(x)) \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right)$$

Canceling out common terms, we get:

$$E = \frac{\hbar^2 n^2 \pi^2}{2mL^2} + U(x)$$

Because  $E$  is constant and independent of  $x$ , and  $U(x)$  is zero within the well, the possible energies become:

$$E = \frac{\hbar^2 n^2 \pi^2}{2mL^2}$$

## 2 Speed of Electron in Bohr Model

2. (a) We know that  $\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c}$

Due to the principles of angular momentum, we obtain:

$$mvr = n\hbar$$

The radius of the  $n$ -th state is:

$$r_n = \frac{4\pi\epsilon_o\hbar^2 n^2}{me^2}$$

Using the two formulas and substituting, we get:

$$v = \frac{1}{n} \left( \frac{e^2}{4\pi\epsilon_o\hbar} \right)$$

Multiplying by  $c$ , we get:

$$v = \frac{1}{n} \left( \frac{e^2}{4\pi\epsilon_o\hbar} \right) \frac{c}{c}$$

$$v = \frac{c}{n} \left( \frac{e^2}{4\pi\epsilon_o\hbar c} \right)$$

And thus,  $v$  becomes:

$$v = \frac{\alpha c}{n}$$

- (b) Given a hydrogen-like atom with charge  $Ze$ , the coulomb force is  $Z^2$  times its original value. The radius, however, is decreased by a factor of  $Z$ , meaning that the velocity becomes:

$$v_Z = \frac{Z\alpha c}{n}$$

### 3 Ionized Helium

3. Ionized helium energy levels would look as follows:

The major difference between hydrogen and helium is the result of the two protons in the nucleus, which modify the energy levels of each quanta. The energy of each  $n$ -th orbit is given by:

$$E_n = -13.6 \frac{z^2}{n^2}$$

Thus, the energies for the different levels, given  $z = 2$  instead of hydrogen's  $z = 1$ , we obtain:

$$E_{1,2,3,4} = -54.4, -13.6, -6.04, -3.4[\text{eV}]$$

Using the wavelength for emission spectra formula:

$$\frac{1}{\lambda} = Rz^2 \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

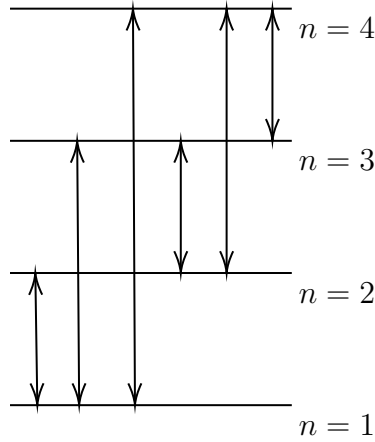


Figure 1: Helium Energy Transition Diagram

Using a calculator<sup>1</sup>, the following wavelengths are obtained:

$$\lambda_{1\leftrightarrow 2} = 30.39[\text{nm}]$$

$$\lambda_{1\leftrightarrow 3} = 25.64[\text{nm}]$$

$$\lambda_{1\leftrightarrow 4} = 24.3[\text{nm}]$$

$$\lambda_{2\leftrightarrow 3} = 164.1[\text{nm}]$$

$$\lambda_{2\leftrightarrow 4} = 121.5[\text{nm}]$$

$$\lambda_{3\leftrightarrow 4} = 468.8[\text{nm}]$$

## 4 Optical Transition and Momentum Conservation

4. The momentum of a photon, given a difference in energy, is:

$$p = \frac{E_1 - E_2}{c}$$

Conservation of linear momentum means that the atom and photon will have equal but opposite momentums. Using the above momentum in the formula for energy, we get:

$$K_R = \frac{p^2}{2m} = \frac{(E_1 - E_2)^2}{2c^2m}$$

Therefore, the provided formula is correct in describing the recoil energy:

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<sup>1</sup>GNU Octave

$$K_R \cong \frac{(E_1 - E_2)^2}{2Mc^2}$$