

# The Rutherford-Bohr Model of the Atom

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# 1 Basic Properties of Atoms

- Around 1900, knowledge about atoms was:
  - Size: Small,  $1[\text{\AA}]/.1[\text{nm}]$
  - Stable: Forces balance
  - Atoms contain electrons ( $e^-$ ) and maintain neutral charge
  - Atoms can emit and absorb electromagnetic radiation

# 2 Scattering Experiments and the Thomson Model

- An early atom model: J.J. Thomson (1904):

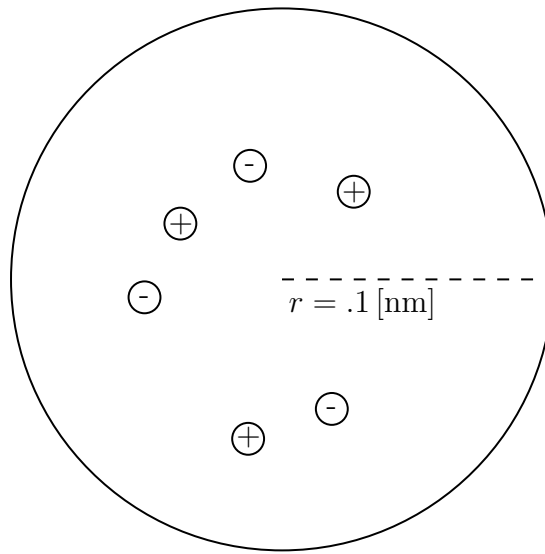


Figure 1: The Jelium Model (“Plum-pudding Model”)

# 3 The Rutherford Nuclear Atom

- Rutherford discovered two rays,  $\alpha$  and  $\beta$ -rays, which he used to experiment with atoms
  - $\alpha$  rays are essentially  $\text{He}^{2+}$  atoms
    - \* The positive charge would mean that the atoms should deflect  $\alpha$  rays
- The Geiger-Marsden Observation

- Their observation found:

$$p(\text{backscattering}) \approx 10^{-4}$$

- This is much larger than expected
- Rutherford proposed that the charge and mass of atoms are concentrated in a region called the nucleus

- The Bohr Model

- Proposed by Niels Bohr (1913)
- Atoms resembled a miniature planetary system

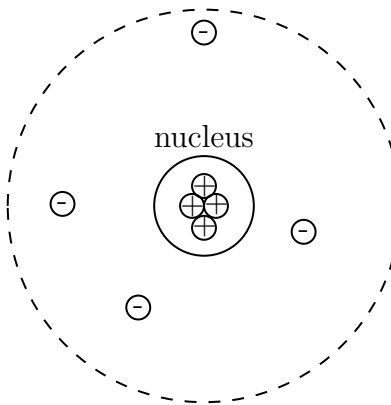


Figure 2: The Bohr Model

- From this, the Coulomb interaction (force) was determined:

$$F = \frac{1}{4\pi\epsilon_o} \frac{|q_1||q_2|}{r^2}$$

- The kinetic energy was determined as:

$$K = \frac{1}{8\pi\epsilon_o} \frac{e^2}{r}$$

- If the electron is radiating, it slows down and moves toward nucleus to collapse?
  - \* Niels Bohr hypothesized that electrons may exist in “stationary states” without radiating electromagnetic energy
- $L = rp = rmv = n\hbar \rightarrow v = \frac{n\hbar}{mr}$ 
  - \* Where  $L$  is the angular momentum,  $r$  is the radius,  $n$  is a quantized number, and  $m$  is the mass

- Substituting this into kinetic energy, we get the permitted radius,  $r$ :

$$r_n = \frac{4\pi\epsilon_o\hbar^2}{me^2}n^2$$

- Using electron information, we get:

$$\frac{4\pi\epsilon_o\hbar^2}{me^2} = .0529[\text{nm}]$$

- This value is known as the Bohr radius

- Hydrogen Atom and Bohr Model:

- Radius:  $r_n = a_on^2$ ,  $n = 1, 2, 3$
- Energy:  $E_n = \frac{-13.6[\text{eV}]}{n^2}$
- $\Delta E = E_n - E_1$  is the excitation energy, where  $E_n$  is the  $n$ th excited state, and  $E_1$  is the ground state
  - \*  $|E_n|$  is the binding energy of  $e^-$  in state  $n$  (ionization energy)
- Optical transitions<sup>1</sup> result in absorption or emission of a photon
  - \* In stationary state, there is no electromagnetic energy radiation
  - \*  $e^-$  can emit radiation when moving from  $n_1$  to  $n_2$

## 4 Line Spectra

- The absorption or emission from atoms may be used to create an emission spectra
- A general equation was generated:

$$\lambda = \lambda_{\text{limit}} \frac{n^2}{n^2 - n_o^2}, \quad n = n_o + 1, n_o + 2, \dots$$

- For the Balmer series,  $n_o = 2$
- For the Lyman series,  $n_o = 1$
- The Ritz combination principle:

$$f_1 + f_2 = f_3$$

- This principle shows that the sum of any two emission frequencies results in a frequency that is also in the spectrum

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<sup>1</sup>Transitions using a photon

- The wavelength of the transition becomes:

$$\lambda = \frac{c}{f} = \frac{64\pi^3 \epsilon_o^2 \hbar^3 c}{m e^4} \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} = \frac{1}{R_\infty} \frac{n_1^2 n_2^2}{n_1^2 - n_2^2}$$

- Where  $R_\infty$  is the Rydberg constant, equal to  $1.097 \cdot 10^7 [\text{m}^{-1}]$
- Lyman series are only observed in the absorption spectrum
- The following summarizes emission spectra
  - Isolated atoms are in the ground state most of the time
  - Excited state lives for a short time (picoseconds to femtoseconds)
  - The absorption spectrum only occurs from the ground state
  - Balmer series are not found in the absorption spectrum

## 5 The Bohr Model

- For an atom with  $z > 1$ 
  - For a nucleus of charge  $ze$ , the Coulomb force is:

$$F = \frac{1}{4\pi\epsilon_o} \frac{|q_1||q_2|}{r^2} = \frac{ze^2}{4\pi\epsilon_o r^2}$$