

# Homework 5

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1. First and foremost, we know the formula:

$$f = \frac{1}{e^{\frac{\varepsilon - \mu}{\tau}} + 1}$$

Thus, we can write:

$$\begin{aligned} -\frac{\partial f}{\partial \varepsilon} &= -\frac{\partial}{\partial \varepsilon} \left( \frac{1}{e^{\frac{\varepsilon - \mu}{\tau}} + 1} \right) \\ -\frac{\partial f}{\partial \varepsilon} &= -\left( -\frac{1}{\tau \left( e^{\frac{\varepsilon - \mu}{\tau}} + 1 \right)^2} \cdot e^{\frac{\varepsilon - \mu}{\tau}} \right) \\ -\frac{\partial f}{\partial \varepsilon} &= \frac{e^{\frac{\varepsilon - \mu}{\tau}}}{\tau \left( e^{\frac{\varepsilon - \mu}{\tau}} + 1 \right)^2} \end{aligned}$$

We can then evaluate:

$$f_{\varepsilon}(\varepsilon = \mu) = \frac{e^0}{\tau (e^0 + 1)^2}$$

$$f_{\varepsilon}(\varepsilon = \mu) = \frac{1}{4\tau}$$

3. (a) The partition function is given by the sum of energy values:

$$3 = 1 + \lambda e^{-\frac{\varepsilon}{\tau}} + \lambda^2 e^{-\frac{2\varepsilon}{\tau}}$$

We can then define the ensemble average occupancy:

$$\langle N \rangle = \lambda \frac{\partial}{\partial \lambda} (\ln(3))$$

$$\langle N \rangle = \frac{\lambda}{3} \left( e^{-\frac{\varepsilon}{\tau}} + 2\lambda e^{-\frac{2\varepsilon}{\tau}} \right)$$

$$\langle N \rangle = \frac{1}{1 + \lambda e^{-\frac{\varepsilon}{\tau}} + \lambda^2 e^{-\frac{2\varepsilon}{\tau}}} \left( \lambda e^{-\frac{\varepsilon}{\tau}} + 2\lambda^2 e^{-\frac{2\varepsilon}{\tau}} \right)$$

$$\langle N \rangle = \frac{\left( \lambda e^{-\frac{\varepsilon}{\tau}} + 2\lambda^2 e^{-\frac{2\varepsilon}{\tau}} \right)}{1 + \lambda e^{-\frac{\varepsilon}{\tau}} + \lambda^2 e^{-\frac{2\varepsilon}{\tau}}}$$

(b) In this situation, the partition function becomes:

$$3 = 1 + 2\lambda e^{-\frac{\varepsilon}{\tau}} + \lambda^2 e^{-\frac{2\varepsilon}{\tau}}$$

Just like (a), the ensemble average occupancy can be defined in a similar way:

$$\langle N \rangle = \lambda \frac{\partial}{\partial \lambda} (\ln(3))$$

$$\langle N \rangle = \frac{\lambda}{3} \frac{\partial}{\partial \lambda} (3)$$

$$\langle N \rangle = \frac{\lambda}{3} \left( 2e^{-\frac{\varepsilon}{\tau}} + 2\lambda e^{-\frac{2\varepsilon}{\tau}} \right)$$

$$\langle N \rangle = \frac{2\lambda e^{-\frac{\varepsilon}{\tau}} + 2\lambda^2 e^{-\frac{2\varepsilon}{\tau}}}{1 + 2\lambda e^{-\frac{\varepsilon}{\tau}} + \lambda^2 e^{-\frac{2\varepsilon}{\tau}}}$$

6. Since we know the entropy is a function of volume and temperature, we write:

$$\sigma = \sigma(V, \tau)$$

which gives us:

$$d\sigma = \left( \frac{\partial \sigma}{\partial V} \right)_{\tau} dV + \left( \frac{\partial \sigma}{\partial \tau} \right)_{V} d\tau$$

Furthermore, from a modification of one of Maxwell's equations, we know:

$$\left( \frac{\partial \sigma}{\partial V} \right)_{\tau} = \left( \frac{\partial P}{\partial \tau} \right)_{V}$$

and we also know a relation for the specific heat at constant volume:

$$\tau \left( \frac{\partial \sigma}{\partial \tau} \right)_{V} = C_v$$

Plugging these into our equation, we get:

$$d\sigma = \left( \frac{\partial P}{\partial \tau} \right)_V dV + \frac{C_v}{\tau} d\tau$$

Then, assuming both  $A$  and  $B$  are ideal gasses, we can write:

$$\begin{aligned} PV &= N\tau \\ \left( \frac{\partial P}{\partial \tau} \right)_V &= \frac{N}{V} \end{aligned}$$

Returning to our equation, we obtain:

$$d\sigma = \frac{N}{V} dV + \frac{C_v}{\tau} d\tau$$

Integrating both sides, we obtain:

$$\sigma = N \ln(V) + C_v \ln(\tau)$$

We can defined these separately for each species:

$$\begin{aligned} \sigma_A &= N \ln(V) + C_v^A \ln(\tau) \\ \sigma_B &= N \ln(V) + C_v^B \ln(\tau) \end{aligned}$$

The difference in entropy may be written as:

$$\Delta\sigma = \sigma_{A+B} - (\sigma_A + \sigma_B)$$

We know:

$$\sigma_{A+B} = 2N \ln(2V) + (C_v^A + C_v^B) \ln(\tau)$$

Which then gives:

$$\begin{aligned} \Delta\sigma &= 2N \ln(2V) + \cancel{(C_v^A + C_v^B) \ln(\tau)} - 2N \ln(V) - \cancel{(C_v^A + C_v^B) \ln(\tau)} \\ \Delta\sigma &= 2N \ln(2V) - 2N \ln(V) \\ \boxed{\Delta\sigma} &= \boxed{2N \ln(2)} \end{aligned}$$

Assuming the two are the same species, the volume occupied would become  $2V$ , since each is indistinguishable. Bringing us to the last step, we find:

$$\Delta\sigma = 2N \ln(2V) - 2N \ln(2V)$$

And, thus:

$$\boxed{\Delta\sigma_{A \equiv B} = 0}$$

12. (a) We can begin by defining the partition function as:

$$z_1 = \sum e^{-\frac{\varepsilon}{\tau}} = \int D(\varepsilon) e^{-\frac{\varepsilon}{\tau}} d\varepsilon$$

In two dimensions, the energy density is:

$$D(\varepsilon) d\varepsilon = \frac{Ap}{2\pi\hbar^2} dp = \frac{A}{\pi\hbar^2} d\left(\frac{p^2}{2}\right) = \frac{AM}{\pi\hbar^2} d\left(\frac{p^2}{2M}\right) = \frac{AM}{\pi\hbar^2} d(\varepsilon)$$

Thus, the partition function becomes:

$$z_1 = \frac{L^2 M}{\pi\hbar^2} \int_0^\infty e^{-\frac{\varepsilon}{\tau}} d\varepsilon = \frac{L^2 M \tau}{\pi\hbar^2}$$

We can then write:

$$\begin{aligned} N &= z_1 \lambda \\ N &= \frac{L^2 M \tau}{\pi\hbar^2} e^{\frac{\mu}{\tau}} \\ \frac{\mu}{\tau} &= \ln\left(\frac{N\pi\hbar^2}{L^2 M \tau}\right) \\ \mu &= \tau \ln\left(\frac{N\pi\hbar^2}{L^2 M \tau}\right) \end{aligned}$$

(b) We can sum all of the unidimensional components to get:

$$U = \frac{1}{2} N \tau + \frac{1}{2} N \tau$$

$$\boxed{U = N \tau}$$

(c) We can begin by finding the free energy:

$$F = -\tau \ln(z)$$

We know the partition function may be written as:

$$z = \frac{z_1^N}{N!}$$

Substituting  $z$  into the free energy formula, we get:

$$F = -N \tau \ln\left(\frac{N\pi\hbar^2}{L^2 M \tau}\right) + N \tau \ln(N) - N \tau = -N \tau \left(\ln\left(\frac{N\pi\hbar^2}{L^2 M \tau}\right) + 1\right)$$

We know:

$$\sigma = - \left( \frac{\partial F}{\partial \tau} \right)_{V,N}$$

Which gives us:

$$\begin{aligned}\sigma &= \frac{\partial}{\partial \tau} \left( N \tau \left[ \ln \left( \frac{\pi N \hbar^2}{L^2 M \tau} \right) + 1 \right] \right) \\ \sigma &= N \ln \left( \frac{\pi N \hbar^2}{L^2 M \tau} \right) + N - \tau \frac{\partial}{\partial \tau} \left( \ln \left( \frac{\pi N \hbar^2}{L^2 M \tau} \right) \right) \\ \sigma &= N \ln \left( \frac{\pi N \hbar^2}{L^2 M \tau} \right) + N + \underbrace{\frac{\tau^2 M L^2}{2 \pi \hbar^2} \frac{\partial}{\partial \tau} \left( \frac{\pi N \hbar^2}{L^2 M \tau} \right)}_N\end{aligned}$$

We end up with an expression similar to the Sackur-Tetrode equation:

$$\boxed{\sigma = N \left[ \ln \left( \frac{\pi \hbar^2}{L^2 M \tau} \right) + 2 \right]}$$

13. (a) We know the canonical partition function is:

$$Z_N = \frac{(n_Q V)^N}{N!}$$

We know from the Gibbs sum that:

$$\begin{aligned}\mathfrak{Z} &= \sum_{N=0}^{\infty} \lambda^N Z_N \\ \mathfrak{Z} &= \sum_{N=0}^{\infty} \frac{\lambda^N (n_Q V)^N}{N!} \\ \mathfrak{Z} &= \sum_{N=0}^{\infty} \frac{(\lambda n_Q V)^N}{N!}\end{aligned}$$

Since this sum is of a known form, we can rewrite it as:

$$\boxed{\mathfrak{Z} = e^{\lambda n_Q V}}$$

- (b) We can write the probability as:

$$P(N) = \frac{\lambda^N Z_N}{\mathfrak{Z}}$$

Which becomes:

$$P(N) = \frac{(\lambda n_Q V)^N}{N! e^{\lambda n_Q V}}$$

We then need to find the average concentration:

$$\begin{aligned}
\langle N \rangle &= \frac{1}{3} \sum_{N=0}^{\infty} N \lambda^N Z_N \\
\langle N \rangle &= \frac{1}{3} \sum_{N=0}^{\infty} \frac{N (\lambda n_Q V)^N}{N!} \\
\langle N \rangle &= \frac{1}{3} \sum_{N=0}^{\infty} \frac{(\lambda n_Q V)^N}{(N-1)!} \\
\langle N \rangle &= \frac{\lambda n_Q V}{3} \underbrace{\sum_{N=0}^{\infty} \frac{(\lambda n_Q V)^{N-1}}{(N-1)!}}_3 \\
\langle N \rangle &= \lambda n_Q V
\end{aligned}$$

Now returning to our expression for probability, we can write it as:

$$P(N) = \frac{\langle N \rangle^N}{N! e^{\langle N \rangle}}$$

Bringing the exponential up, we get:

$$P(N) = \frac{\langle N \rangle^N e^{-\langle N \rangle}}{N!}$$

(c) We can write the first summation as:

$$\sum_{N=0}^{\infty} \frac{\langle N \rangle^N e^{-\langle N \rangle}}{N!}$$

This can also be written as:

$$e^{-\langle N \rangle} \underbrace{\sum_{N=0}^{\infty} \frac{\langle N \rangle^N}{N!}}_{e^{\langle N \rangle}}$$

Which becomes:

$$\sum_N P(N) = e^{-\langle N \rangle} e^{\langle N \rangle} = 1$$

We can write the second summation as:

$$\sum_{N=0}^{\infty} \frac{N \langle N \rangle^N e^{-\langle N \rangle}}{N!}$$

This can be written as:

$$\sum_{N=0}^{\infty} \frac{\langle N \rangle^N e^{-\langle N \rangle}}{(N-1)!}$$

$$\langle N \rangle e^{-\langle N \rangle} \sum_{N=0}^{\infty} \frac{\langle N \rangle^{N-1}}{(N-1)!}$$

This then becomes:

$$\boxed{\sum_N NP(N) = \langle N \rangle e^{-\langle N \rangle} e^{\langle N \rangle} = \langle N \rangle}$$

14. (a) For this problem, we can simply use a formula:

$$Q = N\tau \ln \left( \frac{V_2}{V_1} \right)$$

This becomes:

$$Q = Nk_B T \ln \left( \frac{V_2}{V_1} \right)$$

$$Q = (1)(8.314)(300) \ln(2)$$

Which is:

$$\boxed{Q = 1.729[\text{kJ}]}$$

For the second process, since the gas expands isentropically, there is no heat transfer between the gas and its surroundings.

- (b) We can write the ratio of temperatures as:

$$\frac{T_2}{T_1} = \left( \frac{V_1}{V_2} \right)^{\gamma-1}$$

With volumes:

$$V_1 = 2V_o$$

$$V_2 = 4V_o$$

$\gamma$  is the ratio of specific heats. We assume  $\gamma = 1.6\bar{6}$ , as we are dealing with a monatomic gas, which yields:

$$T_2 = (300) \left( \frac{1}{2} \right)^{\frac{2}{3}}$$

$$\boxed{T_2 = 189[\text{K}]}$$

(c) We know that the change in entropy may be written as:

$$\Delta\sigma = N \ln\left(\frac{V_2}{V_1}\right)$$

$$\Delta\sigma = N \ln(2)$$

$$\Delta\sigma = (6.022 \cdot 10^{23}) \ln(2)$$

This gives us:

$$\boxed{\Delta\sigma = 4.17 \cdot 10^{23}}$$