

Hamiltonian Systems

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1 Hamiltonian Equations

1.1 Notation and first definitions

1.2 Poisson Bracket

Let $F, G : U \subset \mathbb{R}^{n+1} \rightarrow \mathbb{R}$ be \mathcal{C}^r ($r \geq 1$) functions such that $(q, p, t) \mapsto F(q, p, t), G(q, p, t)$.

We define the **Poisson Bracket (PB)** as a \mathcal{C}^{r-1} function $\{F, G\} : U \rightarrow \mathbb{R}$

$$\begin{aligned} \{F, G\} &= (\nabla_z F)^T J (\nabla_z G) \\ &= (\nabla_q F)^T (\nabla_p G) - (\nabla_p F)^T (\nabla_q G) \\ &= \sum_{i=1}^n \left(\frac{\partial F}{\partial q_i} \frac{\partial G}{\partial p_i} - \frac{\partial F}{\partial p_i} \frac{\partial G}{\partial q_i} \right) \end{aligned}$$

Properties:

1. **Skew-symmetric:** $\{F, G\} = -\{G, F\}$

In particular: $\{F, F\}$

proof:

$$\begin{aligned} -\{F, G\} &= -\left((\nabla_q F)^T (\nabla_p G) - (\nabla_p F)^T (\nabla_q G) \right) \\ &= (\nabla_p F)^T (\nabla_q G) - (\nabla_q F)^T (\nabla_p G) \\ &= (\nabla_q G)^T (\nabla_p F) - (\nabla_p G)^T (\nabla_q F) \\ &= \{G, F\} \end{aligned}$$



2. **Bilinear:** $\{\alpha F_1 + \beta F_2, G\} = \alpha \{F_1, G\} + \beta \{F_2, G\}, \quad \alpha, \beta \in \mathbb{R}$

proof:

$$\begin{aligned} \{\alpha F_1 + \beta F_2, G\} &= \left(\nabla_z (\alpha F_1 + \beta F_2) \right)^T J (\nabla_z G) \\ &= \left(\nabla_z (\alpha F_1) \right)^T J (\nabla_z G) + \left(\nabla_z (\beta F_2) \right)^T J (\nabla_z G) \\ &= \alpha \left(\nabla_z (F_1) \right)^T J (\nabla_z G) + \beta \left(\nabla_z (F_2) \right)^T J (\nabla_z G) \\ &= \alpha \{F_1, G\} + \beta \{F_2, G\} \end{aligned}$$



3. **Leibnitz rule:** $\{F_1, F_2, G\} = F_1 \{F_2, G\} + F_2 \{F_1, G\}$

proof:

4. **Jacobi identity:** $\{F_1, \{F_2, F_3\}\} + \{F_3, \{F_1, F_2\}\} + \{F_2, \{F_3, F_1\}\} = 0$

proof:

2 N-Body Problem

3 Tópico sobre el que haré el trabajo

4 Exercises

4.1 Chapter 1: Introduction to Hamiltonian systems

Make the phase portrait of the Hamiltonian system

$$\begin{aligned}\dot{x} &= y \\ \dot{y} &= x - \frac{x^3}{3}\end{aligned}$$

and compute its Hamiltonian

[Solución](#)

Make the phase portrait of the Hamiltonian system

$$\begin{aligned}\dot{x} &= x \\ \dot{y} &= -y + x^2\end{aligned}$$

and compute its Hamiltonian

[Solución](#)

(Meyer-Hall-Offin) Let x, y, z be the usual coordinates in \mathbb{R}^3 , $r = xi + yj + zk$, $X = \dot{x}$, $Y = \dot{y}$, $Z = \dot{z}$, $R = \dot{r} = Xi + Yj + Zk$.

1. Compute the three components of angular momentum $mr \times R$.
2. Compute the Poisson bracket of any two of the components of angular momentum and show that it is $\pm m$ times the third component of angular momentum.
3. Show that if a system admits two components of angular momentum as integrals, then the system admits all three components of angular momentum as integrals.

[1. adea](#)

[2. dsa](#)

[3. dadsa](#)

(Meyer-Hall-Offin) **A Lie algebra** A is a vector space with a product: $A \times A \rightarrow A$ that satisfies:

- **Anticommutative:** $ab \neq ba$
- **Distributive:** $a(b + c) = ab + ac$
- **Scalar associative:** $(\alpha a)b = \alpha(ab)$
- **Jacobis identity:** $a(bc) + b(ca) + c(ab) = 0$, $a, b, c \in A$, $\alpha \in \{\mathbb{R}, \mathbb{C}\}$

1. Show that vectors in \mathbb{R}^3 form a Lie algebra where the product $*$ is the cross product.
2. Show that smooth functions on an open set in \mathbb{R}^{2n} form a Lie algebra, where $fg = \{f, g\}$, the Poisson bracket.
3. Show that the set of all $n \times n$ matrices, $gl(n, \mathbb{R})$, is a Lie algebra, where $AB = AB - BA$, the Lie product.

1. bla

2. bla

3. bla

(Meyer-Hall-Offin) The pendulum equation is $\ddot{\theta} + \sin \theta = 0$.

1. Show that $2I = \frac{1}{2}\dot{\theta}^2 + (1 - \cos \theta) = \frac{1}{2}\dot{\theta}^2 + 2\sin^2(\theta/2)$ is an integral.
2. Sketch the phase portrait.
3. Make the substitution $y = \sin(\theta/2)$ to get $\dot{y}^2 = (1 - y^2)(I - y^2)$. Show that when $0 < I < 1$, $y = \text{sn}(t, k)$ solves this equation when $k^2 = I$ (Look at the definition of elliptic sine function of Section 1.6 of Meyer-Hall-Offin).

1. bla

2. bla

3. bla

(Meyer-Hall-Offin) Let $H : \mathbb{R}^{2n} \rightarrow \mathbb{R}$ be a globally defined conservative Hamiltonian, and assume that $H(z) \rightarrow +\infty$ as $|z| \rightarrow +\infty$. Show that all solutions of $\dot{z} = J\nabla H(z)$ are bounded (Hint: Think like Dirichlet).

Solución

Consider a \mathcal{C}^2 Hamiltonian $H = H(q, p, t) : U \subset \mathbb{R}^{2n+1} \rightarrow \mathbb{R}$ such that $\det(\partial_p^2 H) \neq 0$ on U . Define $v = \partial_p H(q, p, t)$. Prove:

1.

$$\begin{aligned}\partial_{q_i} L(q, v, t) &= -\partial_{q_i} H(q, p, t) \\ \partial_{v_i} L(q, v, t) &= p_i \\ \partial_t L(q, v, t) &= -\partial_t H(q, p, t)\end{aligned}$$

2. The Lagrangian L is \mathcal{C}^2 and $\det(\partial_v^2 L) \neq 0$.

3. The Euler-Lagrange equations associated to L and the Hamiltonian equations $\dot{q}_i = \partial_{p_i} H$, $\dot{p}_i = -\partial_{q_i} H$ are equivalent.

1. bla

2. bla

3. bla

4.2 Chapter 2: The N-body problem

Prove that the linear momentum is a first integral and that the center of mass moves with constant velocity for the 3 body problem.

[Solución](#)

Prove that if (a_1, a_2, \dots, a_N) is a central configuration with value λ :

1. For any $\tau \in \mathbb{R}$ then $(\tau a_1, \tau a_2, \dots, \tau a_N)$ is also a central configuration with value $\frac{\lambda}{\tau^3}$.
2. If A is an orthogonal matrix, then $Aa = (Aa_1, Aa_2, \dots, Aa_N)$ is also a central configuration with the same value λ .

1. bal bla

2. bla bla

(Meyer-Hall-Offin) Draw the complete phase portrait of the collinear Kepler problem. Integrate the collinear Kepler problem.

[Solución](#)

(Meyer-Hall-Offin) Show that $\varpi^2(\epsilon^2 - 1) = 2hc^2$ for the Kepler problem. (Attention: Meyer-Hall-Offin has a typo)

[Solución](#)

(Meyer-Hall-Offin) The area of an ellipse is $\pi a^2(1 - \epsilon^2)^{1/2}$, where a is the semi-major axis. We have seen in Keplers problem that area is swept out at a constant rate of $c/2$. Prove Keplers third law: The period p of a particle in a circular or elliptic orbit ($\epsilon < 1$) of the Kepler problem is $p = (\frac{2\pi}{\sqrt{\mu}})a^{3/2}$.

[Solución](#)

4.3 Chapter 3: Linear Hamiltonian systems

4.4 Chapter 6: Symplectic Transformations

4.5 Chapter 8: Geometric Theory

4.6 Chapter 9: Continuation of solutions

(Meyer-Hall-Offin) Show that the scaling used in Section 9.4 of Meyer-Hall-Offin to obtain Hills orbits for the restricted problem works for Hills lunar problem (see previous problem) also. Why does not the scaling for comets work?

[Solución](#)

Prove Lemma 9.7.1 in Meyer-Hall-Offin. Verify that formula (9.11) is the condition for an orthogonal crossing of the line of syzygy in Delaunay elements.

[Solución](#)

4.7 Chapter 10: Normal forms

4.8 Chapter 13: Stability and KAM Theory

(Meyer-Hall-Offin) Using Poincaré elements show that the continuation of the circular orbits established in Section 6.2 (Poincaré orbits) are of twist type and hence stable.

[Solución](#)

5 Apendix

5.1 Complete examples

5.1.1 Harmonic oscillator

5.1.2 The Pendulum

This is a case of a one DEGREE OF FREEDOM of second order, bla bla bla bla

5.2 Needed resoults and definitions

5.2.1 Linear Algebra

matriz ortogonal

bla bla bla bla

DEVOLVER A SU SITIO LAS FOTOS TAMAÑO CARNET

5.2.2 Calculus

teorema punto fijo de bla bla bla bla

Chain Rule: bla bla bla bla

proof:

reason why:

5.3 Geometry

5.4 Funcional Analysis

5.4.1 Differential forms

5.4.2 Measure Theory

5.4.3 Ordinary Differential Equations