

(10) We have a star, it's very far away,  
so shouldn't we have only 1 pixel to  
represent it?

→ We have a smear

→ Image was blurred and noised up

→ Look

What happens when we have Blur/  
Degradation in addition to noise?

$$g(x,y) = h(x,y) * f(x,y) + n(x,y)$$

$$G(u,v) = H(u,v) F(u,v) + N(u,v)$$

↙  
Could come from moving camera or objects  
moving

→ We talked about how to estimate/model  
 $N(u,v)$ . How about estimating  $H(u,v)$ ?

## (Estimating H<sub>UV</sub>)

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- Guessing : e.g. Take a piece of the degraded image and guess what the original image should have looked like

$$H(u,v) \approx \frac{G_s(u,v)}{\hat{F}_s(u,v)}$$

(S = Guess seen)

corrupted image

(We're working with regions of the image)

- Experimentation (if you have access to the imaging device) -

Directly acquiring the impulse response (point spread function)

- Get a camera in a dark room and have a very small very bright light source (~~source~~)

Impulse  $\rightarrow$  2D impulse response

(e.g. Gaussian)

- Estimate / Model  $H(u,v)$

(from simple functions)

e.g. Gaussian

(Blur)

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(contd from last t23)

get psf (in) 51:24Get point spread functionwe have the Degraded Image  $f(x, y)$ we have the estimated Blur,  $H(x, y)$   
so tryInverse Filteringwe've estimated  $H(u, v)$ 

$$\hat{F}(u, v) = \frac{G(u, v)}{H(u, v)}$$

← sounds good  
in theory but not  
practice

$$= F(u, v) + \frac{N(u, v)}{H(u, v)}$$

If  $H(u, v) \approx 0$  for some  $(u, v)$ , or

$\frac{N(u, v)}{H(u, v)}$  is Large  $\rightarrow$  Poor Result

$$J = \text{deconvwnr}(I, PSF, NSR)$$

Supply Image and point spread function, assume

Noise = 0 (NSR)  $\rightarrow$  we have inverse filter

$$iminv = \text{deconvwnr}(Iinv, h);$$

imshow(iminv)

$\rightarrow$  very Bad!

why is this happening?

we see,

$$\hat{F}(u, v) = \frac{F(u, v)H(u, v) + N(u, v)}{H(u, v)}$$

$$= F(u, v) + \frac{N(u, v)}{H(u, v)} \quad \begin{array}{l} \text{we have noise} \\ \rightarrow \text{for small } H, \\ \text{large noise} \end{array}$$

- mitigate?

$\rightarrow$  Apply when the filter has content  
(close to low frequencies)

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One solution: only apply inverse filter at low frequencies (Small  $u, v$ )  
 otherwise, keep  $G(u, v)$ .

$$F(u, v) = \begin{cases} \frac{G(u, v)}{H(u, v)} & \sqrt{u^2 + v^2} \leq r^2 \\ G(u, v) & \text{otherwise} \end{cases}$$

Smooth, r.w.

$\sqrt{u^2 + v^2} \leq r^2$

We want a smooth transition between these two (we don't want a step function)

$H$  has it  
smooth

{ step function -

but we will do it with  $H$  ←  
(smooth and  $H$ )

The Right Thing to do;

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## Wiener FILTER

Minimum-Mean-Square Error filtering

Goal: Minimize

$$E \left( \underset{\text{expected value}}{(f(x,y) - \hat{f}(x,y))^2} \right)$$

Wiener filter

$$\hat{F}(u,v) = \left[ \frac{1}{H(u,v)} \frac{|H(u,v)|^2}{\left( |H(u,v)|^2 + \frac{S_n(u,v)}{S_f(u,v)} \right)} \right] G(u,v)$$

$$S_n(u,v) = |N(u,v)|^2 = \text{Power Spectrum of Noise}$$

$$S_f(u,v) = |F(u,v)|^2 = \text{Power Spectrum of Signal}$$

$$\hat{F}(u,v) = \left[ \frac{H^*(u,v) S_f(u,v)}{S_f(u,v) |H(u,v)|^2 + S_n(u,v)} \right] G(u,v)$$

(b)  $H$  is our psf point spread function

if we didn't have

$$\frac{|H(u,v)|^2}{|H(u,v)|^2 + \frac{S_n}{S_f}}$$

we would be left with the ~~numerical~~ noise filter

if  $H(u,v) \rightarrow 0$ , we don't explode

if  $H(u,v) \rightarrow 0$  - inverse filter  
we don't have the bad numerical properties

of the inverse filter  $\frac{1}{H(u,v)} = |H(u,v)|^{-1} = (u,v)^{-2}$

there are stuff for the wiener filter

we must know but we don't have access to.

$$\left[ \frac{(u,v)^{-2} |H(u,v)|}{(u,v)^{-2} + |H(u,v)|^2} \right] = (u,v)^{-2}$$

$S_n \rightarrow$  Estimate noise, get pdf  
 (from flat region)  
 or constant (white noise)

$S_f$  is hard to estimate as we don't have access to it

tune  $\frac{S_n}{S_f}$  to match up w/ the SNR constant K

We may be able to estimate  $S_n(u,v)$   
 But we don't know  $S_f(u,v)$  [requires original image]  
 Instead we usually use

$$\hat{F}(u,v) = \left[ \begin{array}{c} \frac{1}{H(u,v)} & \frac{|H(u,v)|^2}{|H(u,v)|^2 + K} \\ \frac{|H(u,v)|^2}{|H(u,v)|^2 + K} & \frac{1}{K} \end{array} \right] G(u,v)$$

Improves noisy  
and blurry images

tunable parameters  
 (related to estimate of SNR)

(18) Wiener filter moon image

point spread function,  $h$

$J = \text{deconvnr}(I, PSF, NSR)$

Noise to Signal Power ratio,  $\frac{0.2}{42}$

$imw = \text{deconvnr}(im, h, 0.0001)$

We underestimate the noise so

this is close to inverse F. filter

try 0.001 → better but not great

try 0.01 → much better, but not

Amazing

Star is much brighter

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$K = 0.1$   $\Rightarrow$  some things are better

$\Rightarrow$  star is compact

$\Rightarrow$  Preserved edge and less noisy

needed psf and K

$\rightarrow$  this is used in practice