CMBFAST

Floor Terra

June 23, 2010

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The big bang

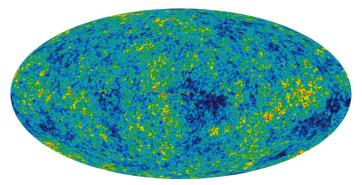
A simple model

- The universe starts small, hot and dense
- The universe expands and cools
- Recombination (z = 1100, T = 4000K)
- Surface of last scattering
- Universe expands while photons travel freely
- CMB is measured by Arno Penzias and Robert Woodrow Wilson ($z=1,\ T=2.725K$)

4□ > 4□ > 4 = > 4 = > = 90

WMAP

What do we see today?

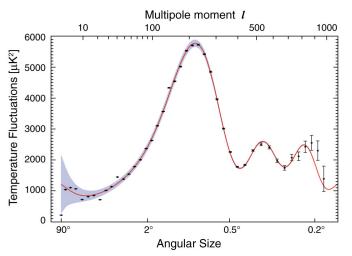


WMAP 7 year data T = 2.725K fluctuations of order 10^{-5}

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WMAP

The data



http://map.gsfc.nasa.gov/media/080999/080999_PowerSpectrumM.jpg

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Perturbations

Where do they come from?

Theory

- Quantum fluctuations
- Inflation
- Density fluctuations

CMBFAST

- Plasma physics
- Gravitation and pressure
- Decoupling
- Last scattering surface

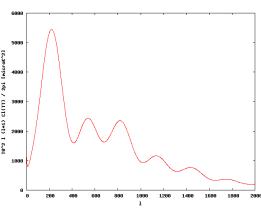
Today

- Gravitation
- Large scale structure

The calculations

======= CMBFAST Processing ========

```
CMB Parameters
ict
            2000
                                   4000.00
1 mo
                    , akmax0
akmaxt
                5.00. nlnkt
                                      5
ntf
z(1)
               0.00
ndyn
                   0, wdyn
                                     -1.00
omegab
               0.05, omegac
                                      0.22
               0.73, omegan
                                      0.00
omegav
h0
              70.00, tcmb
                                      2.72
vhe
               0.24. annunr
                                     3.04
annunr
               0.00, gsnunr
                                      0.00
rcflag
                   0
riflag
                1.00, optdlss
                                      0.09
              50.00, rif
zri
                                      0.20
itflag
                    , nn
itn
                    . irt
an(1)
                     0.96
alphans(1)
                    0.00
ant(1)
                     0.00
alphant(1)
                    0.00
rat(1)
                     0.00
lensflag =
                    , initfl
```



http://lambda.gsfc.nasa.gov/toolbox/tb_cmbfast_form.cfm

Q_rms-ps = 0.5394E+02 micro K

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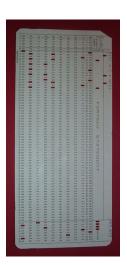
The CMBFAST code

A line-of-sight integration approach to cosmic microwave background anisotropies

- Written by Uros Seljak and Matias Zaldarriaga.
- Article published in 1996
- The first fast CMB code
- Written in the FORTRAN programming language

The problems with CMBFAST

- Sparse documentation
- Designed for interactive use
- FORTRAN



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py-cmbfast

A python wrapper around the CMBFAST code

- Suited for both interactive and scripted use
- Easy to use

```
from libcmb import CMB
cmb = CMB()
# Generate a table with Bessel function values
cmb.jlgen(1500, 3000)
```

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Goals

- Finish wrapping all CMBFAST functionality
- Write a CMBFAST manual
- Write a small (improved) interactive interface



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Questions



