

#### The Magellan Telescopes are twin 6.5m Gregorians

Located at LCO (near La Silla) in Chile (near the GMT site). The Magellan site is excellent for seeing <0.7" median. The MagAO system will go to the Clay (scope on the right).

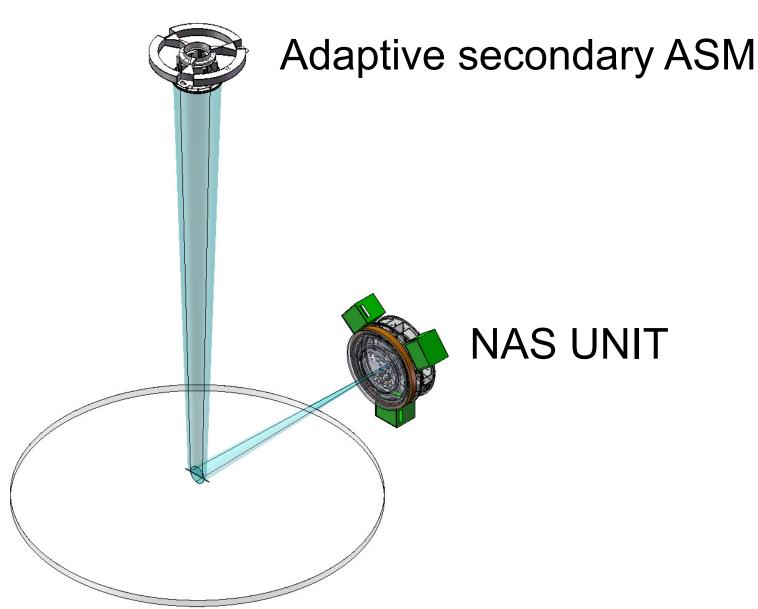
# A Few Magellan AO Key Science Goals (two science cameras)

- 1. The detailed analysis of Exoplanet atmospheres in the thermal IR.
- >> need very high Strehl imaging (hence super-resolution) and coronagraphy 2-5 µm (*Clio2 camera*).
- 2. Extremely high resolution imaging of binaries, faint companions, asteroid surfaces, solar system moons,  $H\alpha/[SII]/Ca$  triplet line imaging etc. (0.65-1.05 µm science)
- 3. Very deep optical imaging of circumstellar and extragalactic environments

>> need broad-band far-red Visible 0.65-1.05 µm wavelength AO (*VisAO*) CCD camera & coronagraph

The P.I. of MagAO is Laird Close, University of Arizona.

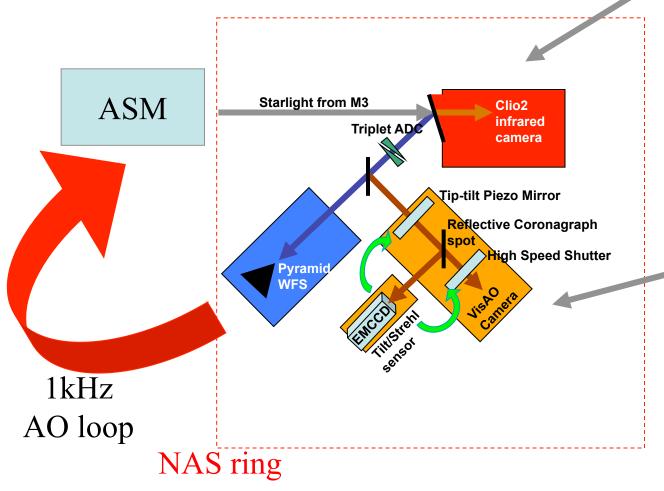
## MagAO



Clio2 (1-5.3 µm)

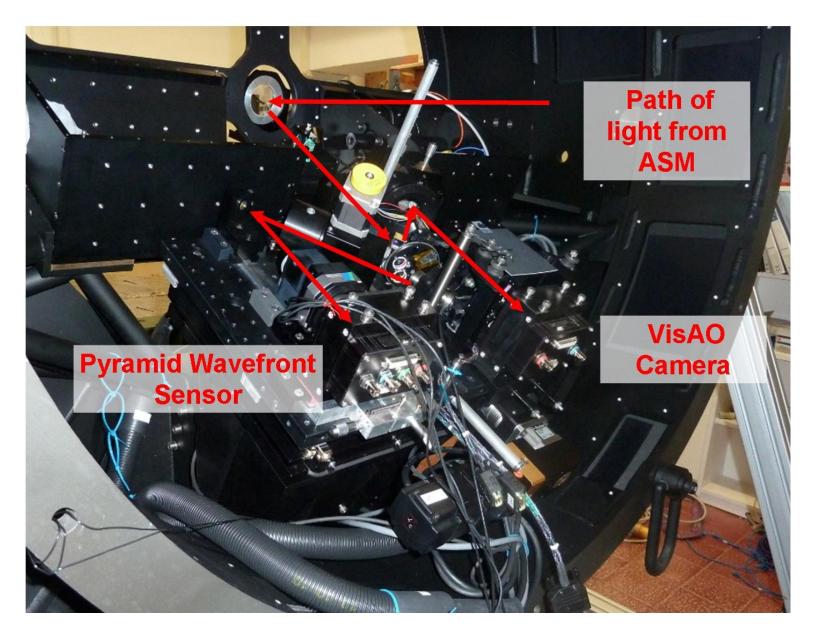
Schematic of MagAO

Camera, coronagraph

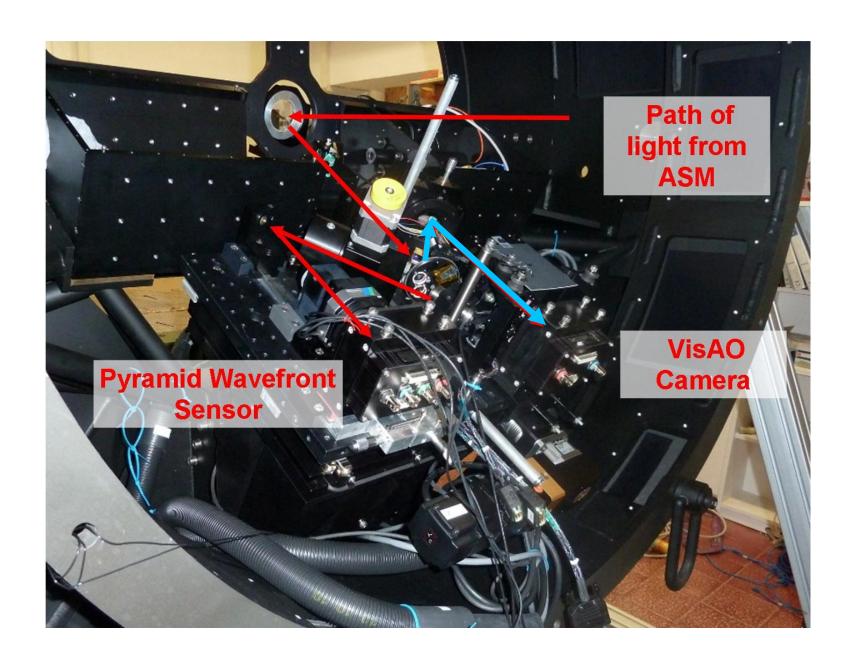


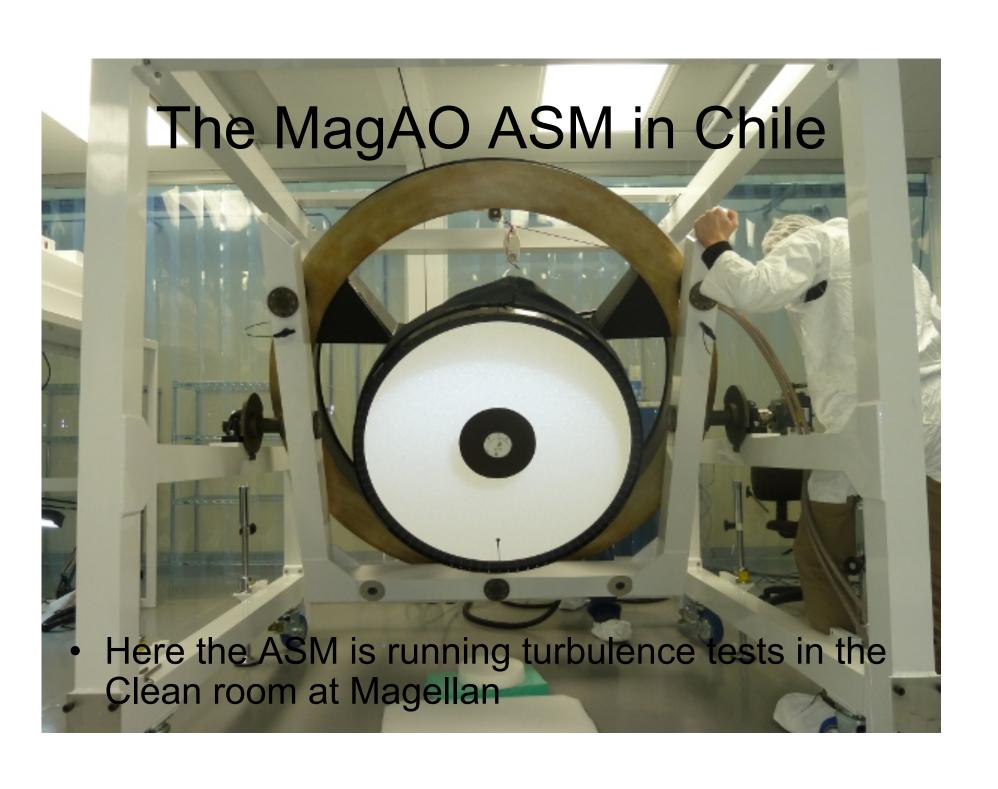
VisAO CCD 20 mas resolution CCD/coronagraph

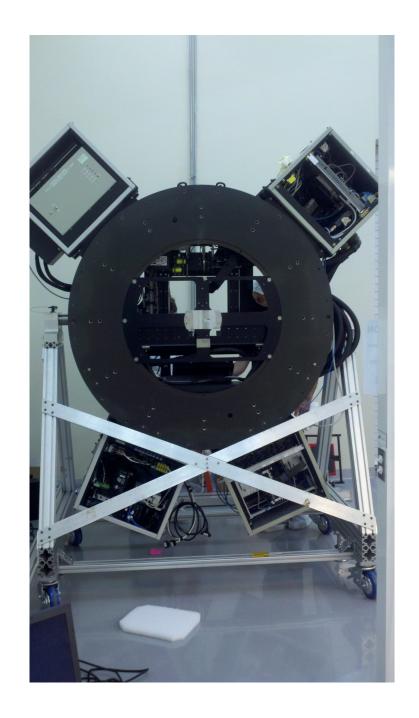
AO imaging at Hα and longer



Close et al. 2012

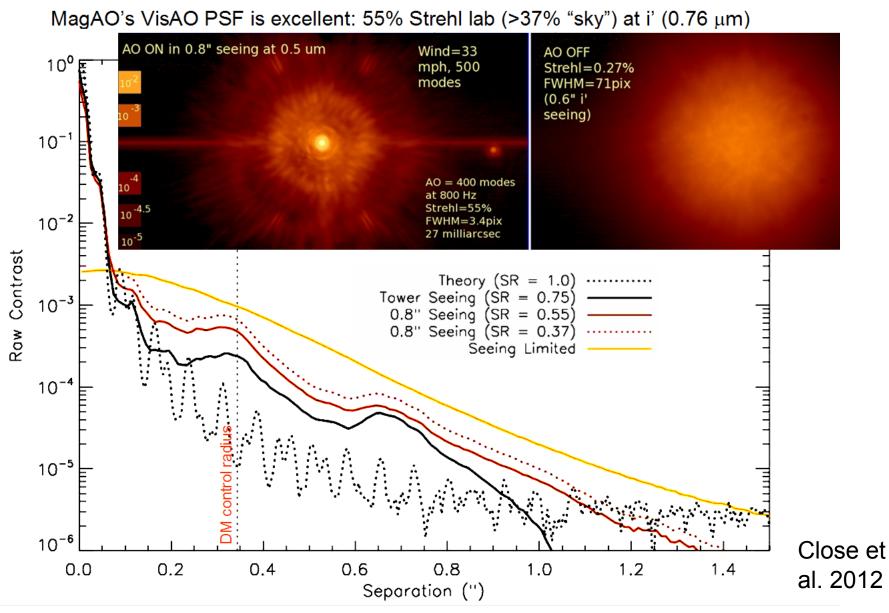




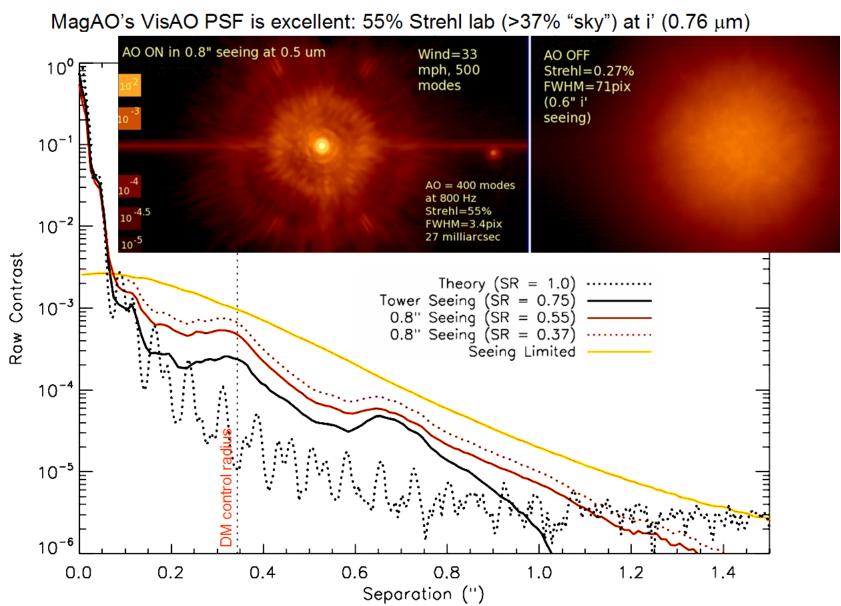


#### The NAS In Chile

 Contains the PWFS, VisAO camera, and Shack Hartmann active optics systems

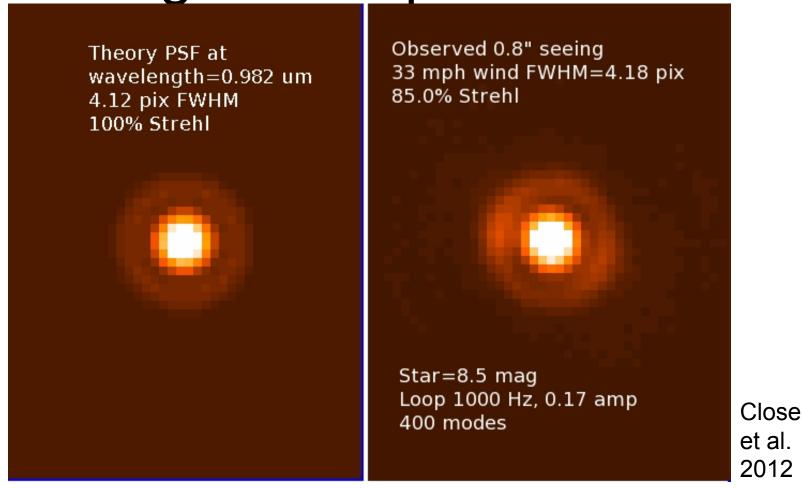


RESULTS OF CLOSED LOOP OPTICAL TESTS IN ARCETRI TEST TOWER



Achieved "on-sky" 122 nm rms wavefront error at 800 Hz --- much better than 190nm goal

Closing the Loop at 1000 Hz



At 1000 Hz closed loop we can achieve 65 nm rms error in tower ~95nm "on-sky". NOTE that *this is a linear stretch*. This is much better than the CDR 190 nm rms top level requirement.

### MagAO Status, June 2012

- The ASM has been successfully delivered and final calibration at Magellan in Chile
- The NAS unit, and VisAO camera, have also had final alignments and tests in Chile
- Clio2 will have its PreShip Review July 4, 2012.
- First Light for MagAO is officially scheduled from Nov 16- Dec 9, 2012.