CHAPTER 4

THE SENSITIVITY TO POLARIZATION IN STRATOSPHERIC AEROSOL RETRIEVALS FROM LIMB SCATTERED MEASUREMENTS

# 4.1 Introduction

Stratospheric aerosols, which are micron-sized spherical liquid droplets of sulfuric acid, cause a cooling effect by scattering the incoming solar irradiance and therefore have an important radiative effect on climate. This effect depends strongly on the aerosol concentration and also the particle size distribution (*Kiehl and Briegleb*, 1993; *Stocker et al*., 2013). Recent studies have proposed a link between the so-called global warming hiatus and an increase in the stratospheric sulfate aerosol layer. (*Solomon et al*., 2011; *Haywood et al*., 2014; *Fyfe et al*., 2013). The increase in stratospheric aerosol over the last decade was primarily caused by a series of somewhat minor, mostly tropical volcanic eruptions (*Vernier et al*., 2011) although the impact of anthropogenic pollution sources continues to be studied (*Neely et al*., 2014). As noted in the recent review paper by *Kremser et al*. (2015) there is a distinct need for continued monitoring with global coverage of aerosol, particularly extending down to tropopause altitudes.

Stratospheric aerosol distributions have been monitored on a global scale since the 1970s with satellite instruments using a variety of remote sensing techniques. The first satellite aerosol extinction profile retrievals were from limb sounding solar occultation measurements, most notably from the NASA SAGE missions (*Russell and McCormick*, 1989; *Thomason and Taha*, 2003). The solar occultation technique has provided a robust and reliable method to retrieve aerosol by directly measuring the atmospheric optical depth. However, the global sampling of occultation measurements is somewhat limited due the necessity of a sunrise or sunset and typically requires months to cover a large range of latitudes. Limb scatter measurements, such as from OSIRIS (*Llewellyn et al*., 2004), SCIAMACHY (*Bovensmann et al.*, 1999), and most recently by OMPS (*Rault and Loughman*, 2013), have better coverage by only requiring the sunlit conditions at the tangent point, but the retrieval of aerosol is more complex requiring computationally heavy forward modelling and inversion compared to occultation (*Bourassa et al.*, 2007; *Bourassa et al.*, 2012b; *Rieger et al.*, 2014). It is worthwhile to note the success of limb scatter aerosol measurements: the combination of the SAGE II and OSIRIS datasets have recently been used to successfully create a single long term merged time series depicting the evolution of the stratospheric aerosol layer (*Rieger et al.*, 2015), and OSIRIS measurements have been used as one of primary extension of the stratospheric aerosol record for the CMIP6 study (*Thomason et al*., in preparation).

OSIRIS, SCIAMACHY, and OMPS measure the spectral radiance of the scattered sunlight from the limb and use non-linear inversion techniques to retrieve aerosol extinction profiles (*Bourassa et al*., 2012b; *Ernst et al*., 2012, *Rault and Loughman*, 2013). For these retrievals, some assumptions regarding particle size distributions and/or composition are always required in the forward model. Most importantly for this study, currently none of these retrievals account for any polarization sensitivity in their respective measurements. However, these instruments have been specifically designed to measure the total radiance by minimizing the instrument sensitivity to polarization. Recently proposed new instruments with the capability to measure aerosol using limb scattering include the Belgium instrument Atmospheric Limb Tracker for the Investigation of the Upcoming Stratosphere (ALTIUS) (*Dekemper et al*., 2012) and the Aerosol Limb Imager (ALI), a Canadian endeavour. Both instruments image the limb and use acousto-optic tunable filters to select the measured wavelength. The use of the acousto-optic filter inherently means that the measured image is of the linearly polarized radiance. Although it has been previously shown that the retrieval of stratospheric aerosol extinction profiles from polarized scattered sunlight measurements are possible (*McLinden et al*., 1999), the full impact of the polarized measurement has not been systematically studied. In this work we perform an analysis with simulated polarized measurements to determine first if there are any clear advantages or disadvantages to making the linearly polarized measurement. Further, we investigate which linear polarization and viewing geometries have the largest sensitivities to aerosol, and how the polarized measurements affect the accuracy and precision of the retrieved aerosol product.

# 4.2 Background and Forward Model

In order to investigate the effect of polarization on the sensitivity to aerosol, an accurate model of the polarized limb radiance must be employed. Additionally, a large number of scenarios, including various atmospheric states and viewing geometries, are required to fully probe the solution space. In this section, the basic background describing the polarization state of the limb signal is developed, and the SASKTRAN-HR model and the various model scenarios used for the analysis are described. Based on the useful spectral range for limb scatter observations of stratospheric aerosol, we have limited our discussion to wavelengths from 500-1500 nm.

## 4.2.1 Polarized Scattered Sunlight and Stratospheric Aerosols

All full description of scattering interactions within the atmosphere can be found in section 2.4.2 to 2.4.4. This section will briefly cover the theory and then use it to analyze the polarization state of earth’s atmosphere in regards to look direction.

The polarization state of electromagnetic waves can be fully defined by the Stokes vector formulation. The Stokes vector is given by a column matrix

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|  | (4.1) |

where the terms of the Stokes vector, defined in a reference frame, are measures of the total radiance, , the difference between horizontal polarization to vertical polarization, , the difference between +45o diagonal polarization to -45o polarization, , and the difference between the counter clockwise circular polarization to clockwise polarization, . Using a reference frame where the x­‑axis is defined to be the horizontal polarization leads to the following definition for the Stokes parameters

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| --- | --- |
|  | (4.2) |

To model the scattering for an incident ray propagating in a given direction the ray undergoes a rotation into the Stokes reference frame and is then is multiplied by the phase matrix for the scattering process. After the multiplication the resulting ray is then rotated back into the ray’s initial coordinate system through the following operation,

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| --- | --- |
|  | (4.3) |

The outgoing, or scattered, and incoming radiances are represented 4 by 1 matrices, *i.e.* Stokes column vectors, given by and , the rotation matrices are given by and rotate the incoming ray and scattered ray by rotation angles and . The phase matrix is a 4 by 4 represented by and is related to the probability that an incoming ray will be scattered at a scattering angle, . It also describes the change in polarization state through the elements of the matrix.

For this work, two primary scattering interactions induce and/or modify the polarization state of the light propagating in the atmosphere. These are scattering by the molecular air density and by stratospheric sulfate aerosols. The molecular atmosphere interaction is referred to as Rayleigh scattering, and has a phase matrix that is determined from the Rayleigh-Gains approximation (*Mishchenko et al*., 2002) given by

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|  | (4.4) |

where is the scattering angle.

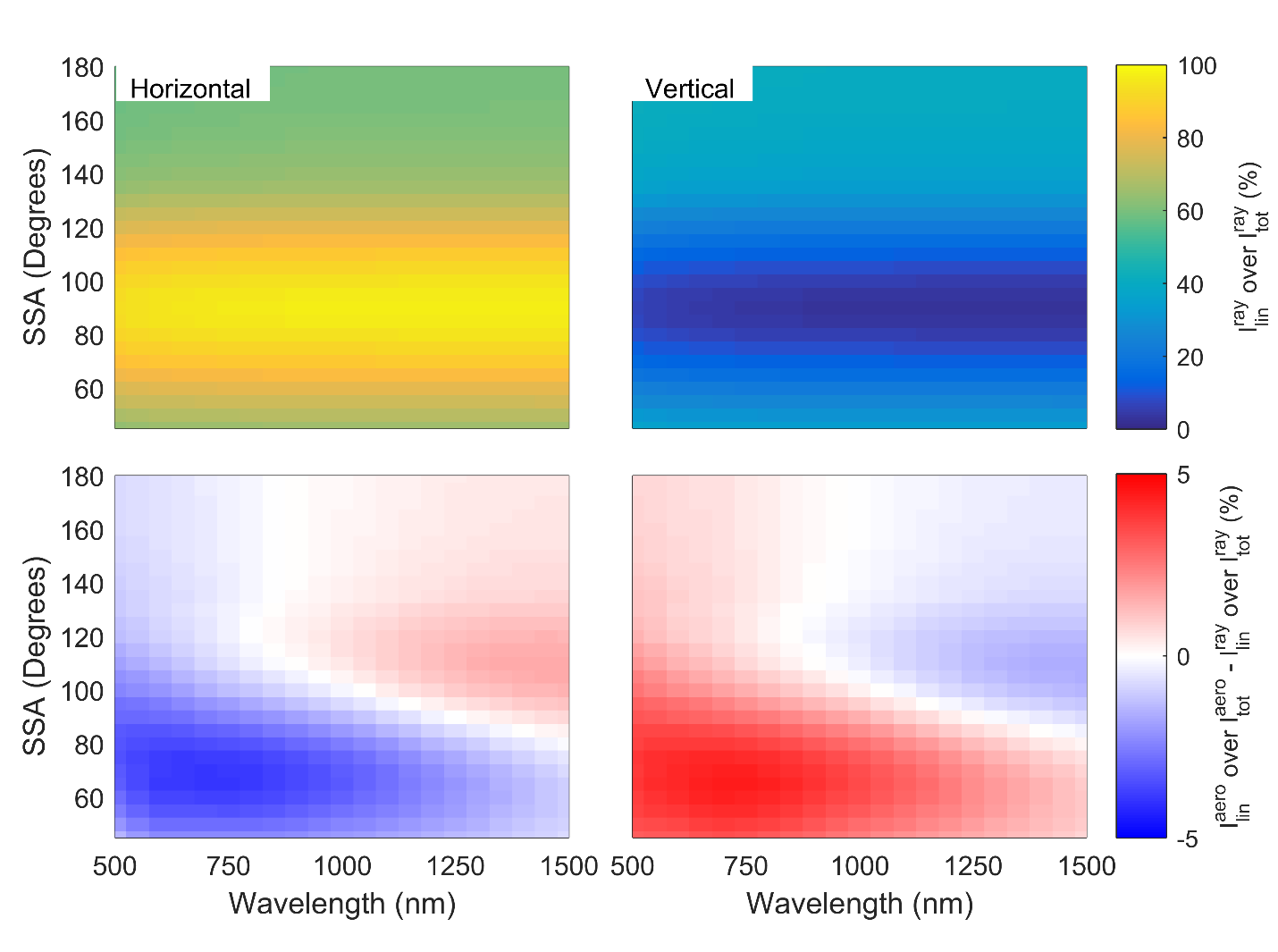
For randomly orientated or spherical particles, such as stratospheric aerosol, only six elements of the phase matrix are required (*van de Hulst*, 1957) which are the following

|  |  |
| --- | --- |
|  | (4.5) |

Additionally, for spherical particles like stratospheric aerosol only four unique terms are required since and . Spherical aerosol scattering at visible and near-infrared wavelengths is fully described by Mie theory (*Mie*, 1908), for which several standard codes have been developed to calculate scattering cross sections and phase matrices based on the particle size distribution and index of refraction (e.g. *Wiscombe*, 1980). A full derivation can be found in *van de Hulst* (1957).

The basic polarization state of the scattered light in the earth’s atmosphere can be understood by first considering a single scattering event of the unpolarized incoming sunlight in a molecular atmosphere. It can be easily seen from the form of the Rayleigh phase matrix (Equation 4.4) that a single scattering event causes the sky to develop a distinct polarization at a solar scattering angle (SSA) of 90 degrees from the incoming solar beam. The scattered sunlight is linearly polarized in the horizontal orientation, which is parallel to the horizon. The degree of polarization gradually decreases at scattering angles greater than or less than 90 degrees (broadly referred to as back-scatter and forward-scatter geometries, respectively). In this single scattering scenario, the radiance is completely unpolarized at solar scattering angles of 0 and 180 degrees. If multiple scattering events are taken into account, the degree of polarization is decreased at 90 degrees SSA, and conversely does not become completely unpolarized at SSAs of 0 and 180 degrees. Simulations with the SASKTRAN-HR radiative transfer model, which is described below, using an atmosphere of molecular air density show that at 90 degrees SSA, the degree of linear polarization of the limb radiance is approximately 95% for a wavelength of 750 nm. This linear polarization effect is strongest at longer wavelengths (*i.e.* approaching 1500 nm) and decreases, on average by 10%, as the wavelength become shorter (*i.e.* down to 500 nm). This is directly related to the greater contribution from multiple scattering at shorter wavelengths. As the SSA increases from 90 degrees, the degree of linear polarization decreases. It is approximately 20% for a back scatter geometry of 180 degrees, and 30% for a scattering angle of 45 degrees. The ratio of the horizontal polarization over the total radiance and the vertical polarization over the total radiance is shown in the top of Figure 4-1. The strong polarized nature can be noted around SSA of 90 degrees where the radiance is almost complete horizontally polarized.

For an atmosphere that contains both the molecular air density as well as a typical background state of stratospheric sulfate aerosol, both Rayleigh and Mie scattering occur in a weighted fraction according to the optical depth of air and aerosol. Compared to the pure Rayleigh scattering case, the addition of aerosol causes a decrease in the degree of linear horizontal polarization for wavelengths shorter than approximately 750 nm. The bottom two panels of Figure **4-1** show the difference in the ratio of the polarized over the total radiance for the atmosphere with aerosol and one without, this effect has a weak dependence on solar scattering angle, with the most depolarization occurring in forward scatter geometries for short wavelengths. Interestingly, for longer wavelengths in back-scatter geometries, the opposite occurs. This is due to the changing fraction of scattering from the molecular air density and aerosol because the Rayleigh scattering cross section falls off much more quickly with wavelength than the aerosol cross section. The magnitude of the observed change in linear polarization from a pure Rayleigh atmosphere to that with typical background aerosol is approximately 5-10%, but it obviously varies depending on aerosol loading and the microphysical parameters of the aerosol.



**Figure 4-1:** (Top) The fraction of a linear polarization (left is horizontal and right is vertical) over the total radiance for molecular air density. (Bottom) The change in linear polarization between an atmosphere that contains aerosol and one with only molecular air density.

## 4.2.2 SASKTRAN-HR Model

The model used for this work is the SASKTRAN-HR radiative transfer model discussed in section 2.4.5 and a brief overview will follow. The High Resolution module of the SASKTRAN radiative transfer framework (*Bourassa et al*., 2007; *Zawada et al*., 2015) was used in this study. The SASKTRAN framework handles built-in and user-specified atmospheric species optical properties and number density profiles, and uses a fully 3D spherical geometry to solve the radiative transfer equation. The High Resolution module uses a successive orders of scattering technique, and solves either the scalar or vector (polarized) radiative transfer equations to arbitrarily many orders of scatter. In this study, the HR module is set to handle the first two orders of scattering in a fully polarized sense, and to handle all scattering into the observer line of sight in a polarized sense. Three orders of scattering into the instrument line of sight are therefore handled in a fully polarized sense. For higher orders of scatter, the higher-than-second orders of scatter are handled in a scalar sense and the polarization state is assumed to be random before the final (polarized) scatter into the observer line of sight. This “pseudo-polarized” approximation has been shown, through comparison against the highly accurate and fully polarized Monte Carlo module of the SASKTRAN framework (*Dueck et al.*, 2016), to be sufficiently accurate for the wavelengths and geometries of interest of this work. All calculations performed with SASKTRAN-HR in this study assume randomly polarized sunlight, and dry air and Mie (H2SO4) scattering events only to model the interaction with the molecular air density and stratospheric aerosol, respectively. Scattering events from the Earth’s surface are assumed to be Lambertian and fully depolarizing.

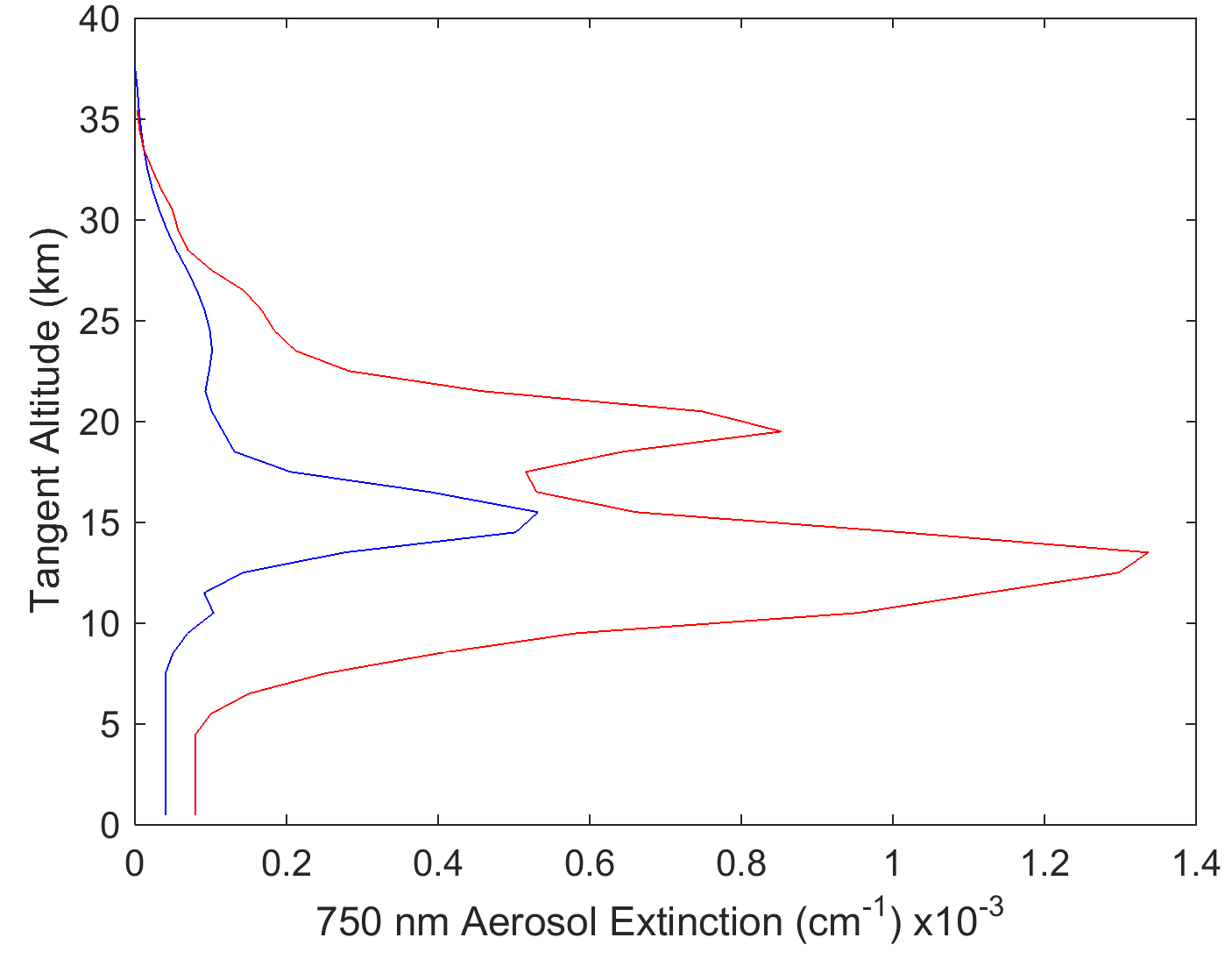
## 4.2.3 Model Scenarios

The impact of using polarized radiance measurements on stratospheric aerosol retrievals is systematically studied with the radiative transfer model by exploring a set of distinct cases that approximately cover the expected range of aerosol parameters, including both particle size and concentration (or extinction) profiles, and viewing geometries. Viewing geometry is an important parameter as even in the case of the total radiance measurements, the geometry can have a substantial effect on the sensitivity of the measurement to aerosol due to asymmetry of the Mie scattering phase function, i.e. element of the phase matrix (*Rieger et al*., 2014). There is strong aerosol scattering in the forward direction and so this results in a weaker relative aerosol signal in the back scatter direction.

To probe the space of possible aerosol measurement scenarios, two aerosol extinction coefficient profiles and four particle size distributions were used. The two extinction profiles, nominally specified at 750 nm, correspond to a background aerosol case, typical of the volcanically quiet period of the early 2000’s (*Deshler et al.*, 2003), and a volcanically enhanced case which was taken from OSIRIS measurements two months after the Nabro eruption in 2012 (*Bourassa et al*., 2012c). Both profiles are shown in Figure 4-2. The four particle size distributions were also chosen to represent typical background and volcanically enhanced cases. The background cases are both single mode lognormal distributions with somewhat different, but still typically observed, size parameters. A bi-modal lognormal particle size distribution was used for the volcanically enhanced cases, with one fine mode and one coarse mode, each comprising an equal fraction of the total extinction. All of the parameters of the size distributions are detailed in Table 4-1. These selected distributions are based on in-situ balloon particle counter measurements from Laramie, Wyoming (*Deshler et al*., 2003). The size distributions were used for translating the extinction profiles, which are specified at 750 nm, to other wavelengths by scaling the extinction by the ratio of the Mie scattering cross sections corresponding to the size distribution at the two wavelengths.

To probe the range of possible viewing geometries from low earth orbit, a range of Solar Zenith Angles (SZAs) and Solar Scattering Angles (SSA) were selected. The ranges give representative selections of the possible geometries of a limb scatter instrument in low earth orbits at a range of local times. The selected values for SZA are 15o, 45o, and 75o and for SSA of 30o, 60o, 90o, 120o, 150o, and 180o.

Simulated measurements were performed at wavelengths of 500, 750, 1000, 1250, 1500 nm, which approximately cover for the spectral range commonly used for aerosol retrievals from limb instruments. For example, OSIRIS and SCHIAMACHY aerosol products use the ratio of 750 nm to 470 nm for the aerosol retrieval (*Bourassa et al*., 2012b; *Ernst et al*., 2012). Additional longer wavelengths have been shown to provide particle size information from limb scatter measurements (*Rieger et al*., 2014) and so the 1000-1500 nm wavelength range was also included in this study. Finally, we also performed simulations for Earth surface albedo values of 0 and 1 in order to cover the full range of potential impact.



**Figure 4-2:** The two aerosol profiles used in this study. The blue is a background aerosol extinction levels, and the red curve is a representative aerosol profile after the Nabro eruption.

**Table 4-1:** Different particle size distributions used to test the sensitivity of the aerosol retrieval.

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| --- | --- | --- | --- | --- | --- |
| Particle size distributions | Fine mode radius (µm) | Fine mode width | Coarse mode radius (µm) | Coarse mode width | Percent extinction coarse mode (%) |
| 1 | 0.04 | 1.8 | -- | -- | 0 |
| 2 | 0.12 | 1.25 | -- | -- | 0 |
| 3 | 0.04 | 1.8 | 0.30 | 1.15 | 50 |
| 4 | 0.12 | 1.25 | 0.30 | 1.15 | 50 |

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# 4.3 Methodology

For the purposes of this study, we have assumed an instrument capable of measuring only the linearly polarized radiance with either a vertical or horizontal orientation. This is representative of newly proposed instruments like ALTIUS (*Dekemper et al*., 2012) and ALI that use an AOTF and by nature only measure one orientation of linearly polarized radiance. We want to answer the question: if the linear polarization is measured, is this an advantage or a disadvantage over a measurement of the total radiance for aerosol retrievals? Further, is there a preferred orientation of linear polarization?

The polarization states used here are defined as the following: the linearly polarized radiance aligned with the horizon is referred to as the horizontal polarization, and the linearly polarized radiance that is perpendicular to the horizon is referred to as the vertical polarization. We also use the total radiance, or alternatively the scalar radiance, as the reference case. Note that the scalar radiance is not precisely equal to the total radiance. For the work presented here the term “total radiance” refers to the first term in the Stokes vector, which is calculated by the SASKTRAN model when solving the vector radiative transfer equation. The term “scalar radiance” refers to the radiance calculated by the SASKTRAN model when solving the scalar transfer equation. Using the Stokes parameter formulation, the horizontal polarization is given by and the vertical polarization is given by , which can be easily shown from the definitions given in Equation 4.2.

Our study further breaks down this problem into three questions. First, how does the fraction of the limb scatter signal that is due to aerosol vary with aerosol load and viewing geometry for both scalar and polarized measurements? Secondly, does the polarized measurement increase sensitivity to assumptions in the retrieval algorithm and therefore increase potential for biased results? And finally, how does the polarized measurement effect the uncertainty estimate of the retrieved profile?

To explore the first question, simulated measurements were calculated with SASKTRAN using the scenarios described in section 4.2.3, including various wavelengths, geometries, aerosol loading and particle size distributions. These simulated measurements are then used to determine the approximate fraction of the limb signal that is due to aerosol. In each case the model is run with a nominal atmosphere that consists of molecular air density, and climatological ozone and nitrogen dioxide profiles.

The fraction of the limb signal due to aerosol is determined by calculating the radiance without aerosol in the model atmosphere, *i.e*. that due to Rayleigh scattering only, , and the radiance including aerosol, . To find the fraction, , in percentage, of the signal that is attributed to aerosol, the following formulation is used:

|  |  |
| --- | --- |
|  | (4.6) |

Due to non-linearities from multiple scattering, it is not strictly true that this is the fraction of the signal due to aerosol; however, at most stratospheric tangent altitudes, the wavelengths under study are quite optically thin and this simple percent difference provides an intuitive approximation of the fraction of the signal due to aerosol.

We explore the second question about the effect of the polarized measurement on the aerosol retrieval using simulated measurements and a retrieval method that is essentially similar to that developed by *Bourassa et al.* (2012b) for OSIRIS. A minor change to the algorithm is made where the measurement vector for this study is not normalized by a shorter wavelength. We have made this change as the results from *Rieger et al.* (2014) shows this actually decreases sensitivity to particle size distributions. Although it is advantageous in a retrieval scenario to limit sensitivity to particle size, for this study we explore the worst case scenario under possible limitations of future technology, given that not all instruments may cover a wide enough spectral range for short wavelength normalization.

The limb radiance is calculated using SASKTRAN, again with climatological ozone and NO2 profiles, for each of the scenarios listed in section 4.2.3. This is taken as a simulated measurement and is then used to retrieve aerosol extinction profiles using the *Bourassa et al.* (2012b) technique. This is done similarly for the total radiance and for each orientation of the linearly polarized radiance. Additionally, a retrieval is performed with the scalar SASKTRAN-HR model to see if there is any substantial difference between using the scalar radiance and the total radiance from the vector model. For each aerosol retrieval, the ozone, NO2, and albedo are fixed to the values used in the simulation of the measurement. All four particle size distributions from Table 4-1 are used in the simulations, but following *Bourassa et al*. (2012b), the aerosol particle size is fixed in the retrieval to a single mode log-normal with 0.08 µm mode radius and mode width of 1.6. The assumption of a fixed particle size distribution is common in limb scatter retrieval algorithms and this is used to explore sensitivity of the polarized measurements to particle size distributions, and test if the uncertainty in this assumption greatly effects the retrieved extinction.

Lastly, to answer the third question, an uncertainty estimate is performed on these retrievals in order to check the precision of the retrieved aerosol profile. The precision is determined by mapping the covariance of the measurement vector, through the gain matrix, , which describes the sensitivity of the retrieval to the measurement and the respective noise through the following (*Rodgers*, 2000)

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| --- | --- |
|  | (4.7) |

where is the co-variance matrix for the retrieved aerosol profile. However, the direct calculation of the gain matrix is computationally intensive and numerically requires a retrieval for each measured altitude. A method presented by *Bourassa et al.* (2012a) uses the Jacobian, , to approximate the gain matrix by assuming the problem is linear near the solution state, which is largely a good assumption for limb scatter aerosol retrievals. Using these assumptions, the gain matrix can be determined simply through the inverse of the Jacobian,

|  |  |
| --- | --- |
|  | (4.8) |

Rather than specifying an assumed measurement co-variance to study the behavior of the retrieval precision, we simply replace the measurement co-variance in Equation 4.7 with the identity matrix. Thus the resulting terms of are not absolute quantities but are related to the amplification of the measurement noise when mapped to the retrieved state (*i.e*. the larger the values of the larger the uncertainty for the retrieval). The square root of the elements of the diagonal of the aerosol covariance, typically used to represent the error bars on the retrieved profile, are taken as the amplification of the measurement noise.

This method assumes that the radiance measurements regardless of polarization state have exactly the same signal to noise performance, *i.e*. all measurements have the same co-variance. We also consider the case where the instrument is not compensated such that the magnitude of the various polarization states directly scales the signal to noise performance, *i.e*. the instrument is not compensated to equalize the measurement co-variance when the signal drops due to the measured polarization state. In this scenario the above method must be modified by replacing the identity matrix with the matrix, , to represent the change in signal strength for the various polarizations relative to the scalar case. This matrix is defined as

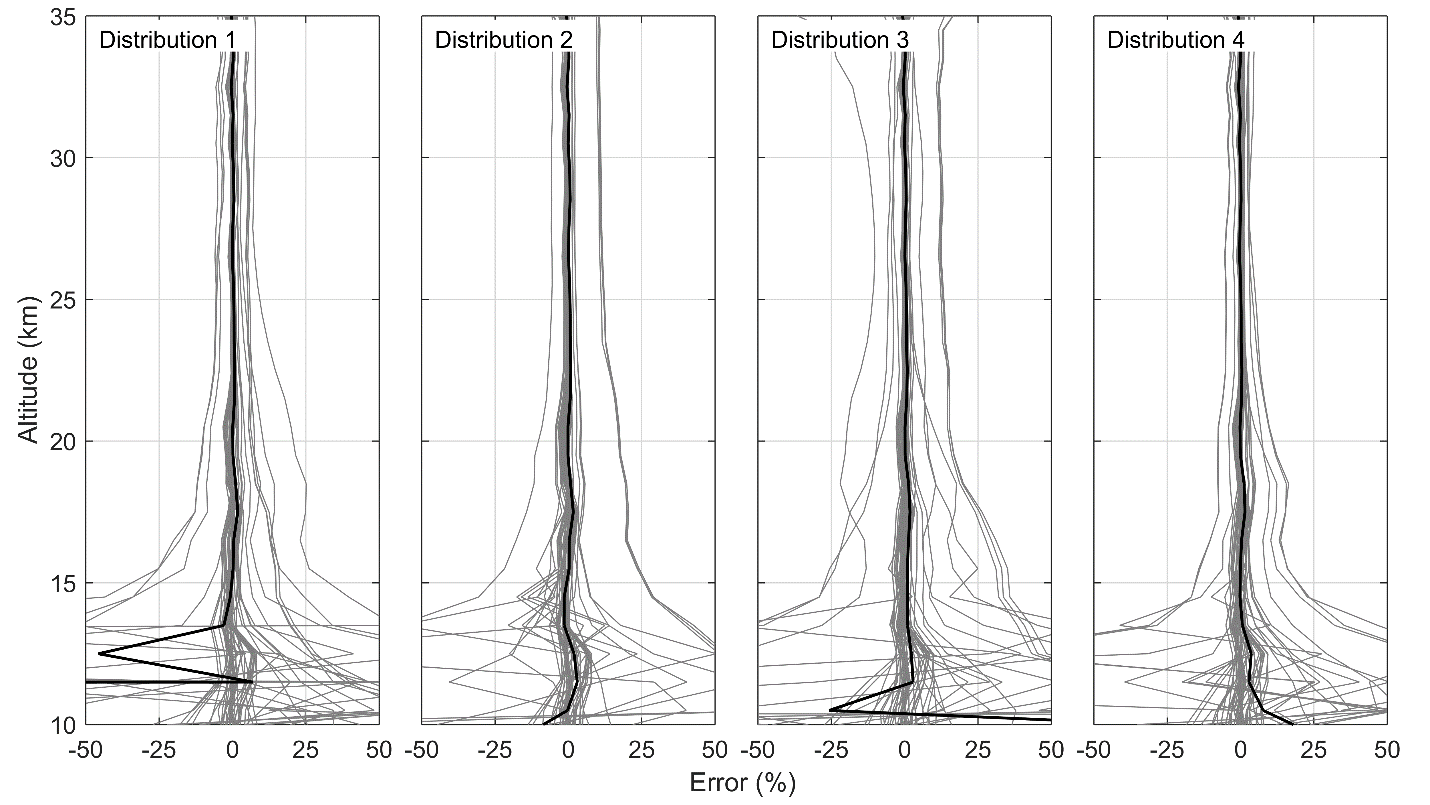
|  |  |
| --- | --- |
|  | (4.9) |

The diagonal of the R matrix is effectively scaled by the inverse of the magnitude of ratio of the polarized radiance, , to the reference scalar case, , for the measurement altitude, .

# 4.4 Analysis

## 4.4.1 Difference in Scalar Retrievals using a Scalar or Vector Model

First, we investigate if there is any significant difference between the use of the scalar radiance and the total radiance for retrievals on measurements of the total radiance. As mentioned above, retrieval algorithms for current limb scatter data sets such as OSIRIS and SCIAMACHY use a scalar radiative transfer model with general success; however, as the total radiance is not strictly equal to the scalar radiance, this may lead to biases in the retrieved extinction profile under certain scenarios. Accounting for the vector component in the model alters the overall total radiance from the scalar solution due to multiple scattering interactions between the various polarization states of each successive order of scattering.

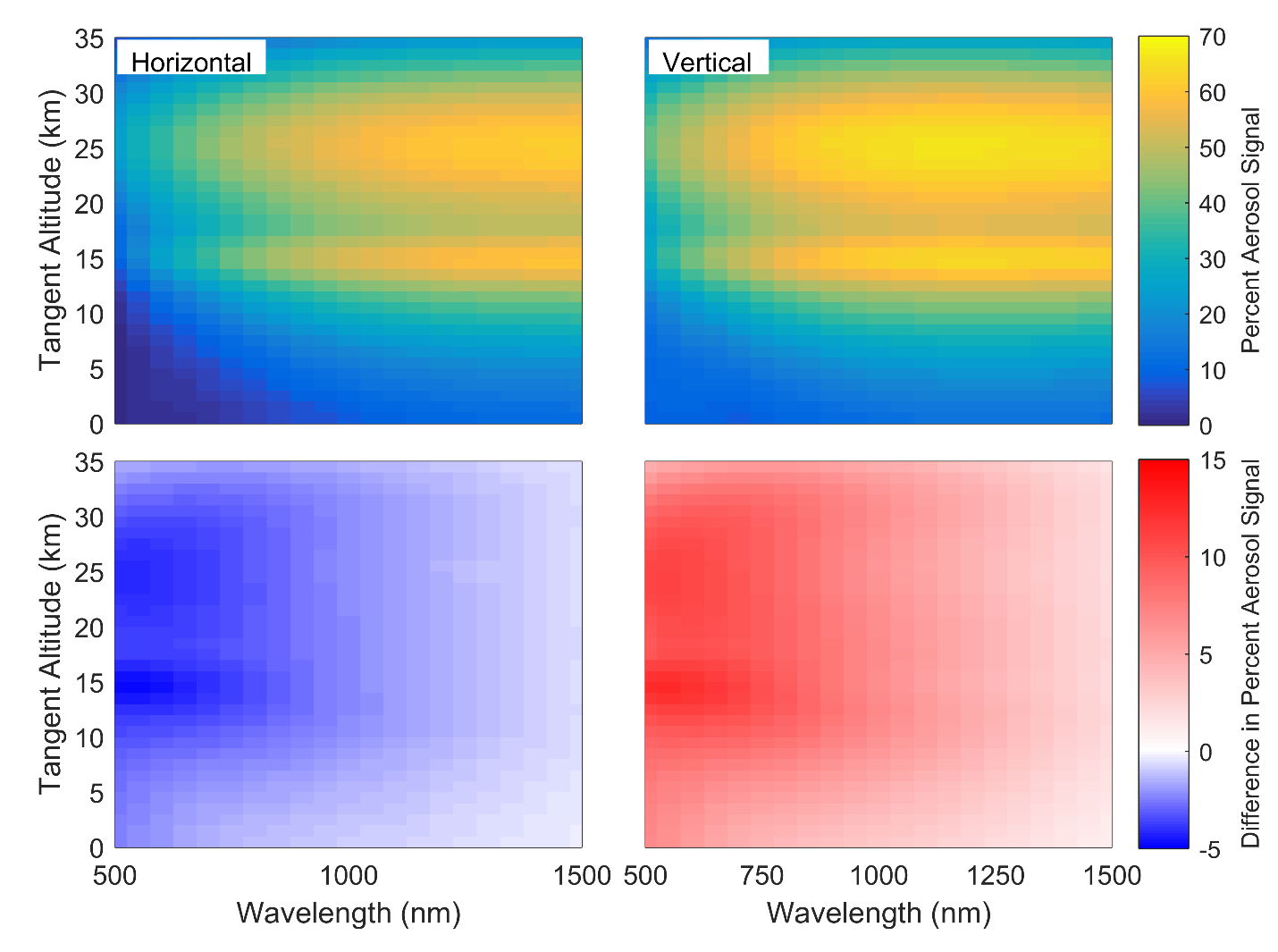


**Figure 4-3**: Percent differences between the vector retrieved aerosol extinction profiles and the scalar retrieval from simulated total radiance measurements. Each column represents a different particle size distribution (see Table 4-1).

The total radiance was simulated with SASKTRAN in vector mode for the full set of wavelengths and viewing geometries, and for the range of aerosol loading scenarios. These were used as input measurements to the retrieval algorithm, which was then performed using both the scalar and vector models. A case-by-case comparison between the retrieved extinctions for the scalar and vector models was performed using a simple percentage difference at each retrieved altitude a can be seen as the grey lines in Figure 4-3. Furthermore, the mean of the bias for each particle size distribution is shown in solid black. These results, given in Figure 4-3, show that across all wavelengths, the mean percent difference is less than 2% from 15 to 37 km. A small number of outlier cases occur where the difference between the retrievals is greater than 7%. All of these cases occur for back scatter geometries and short wavelengths. The reason for this discrepancy is not well understood, although it certainly arises from the differences between the scalar and total radiance due to polarization interactions from the relatively larger contribution of multiply scattered light at shorter wavelengths. Generally, however, any differences between the use of the scalar and vector model for the retrieval are negligible. In fact, any form of discrepancy essentially vanishes for wavelengths past 1000 nm. Since the use of the vector model can increase calculation times by a factor of at least two, it is certainly justifiable to use the scalar model for the overwhelming majority of scenarios. For the rest of the work presented, any reference to the radiance will only refer to the total radiance, , from the vector model.

## 4.4.2 Fraction of Limb Signal due to Aerosol

For a typical background aerosol state, the fractional contribution to the total limb radiance from aerosol was calculated from modelled radiances over a series of stratospheric tangent altitudes using the background aerosol profile and particle size distribution #1, given in Table 4-1. Figure 4-4 shows the percent change in this fraction for horizontally and vertically polarized measurements compared to the total radiance. The viewing geometry, which is a typical low earth orbit scenario, is SZA=45o and SSA=60o, and the albedo is 0. The top of Figure 4-4 shows the percentage of signal that is contributed to aerosol for both horizontal and vertical linear polarizations. As can be seen from the bottom of Figure 4-4, most of the change in the aerosol fraction of the polarized limb signal compared to the total radiance occurs for wavelengths between 500-1000 nm. At these wavelengths the horizontal polarization has a smaller fraction of signal due to aerosol and the vertical polarization has a larger fraction due to aerosol. Overall the change is small and essentially limited to less than 10%.



**Figure 4-4**: (Top) For a horizontal (left) or vertical (right) linear polarization the percent of the signal that is attributed to aerosol. (Bottom) The change in the fraction of the limb signal due to aerosol when compared to the total radiance for the horizontal (left) and vertical (right) polarization. The simulation uses a geometry of SZA=45o and SSA=60o, with the albedo being 0 and the aerosol state the background profile with particle size distribution #1.



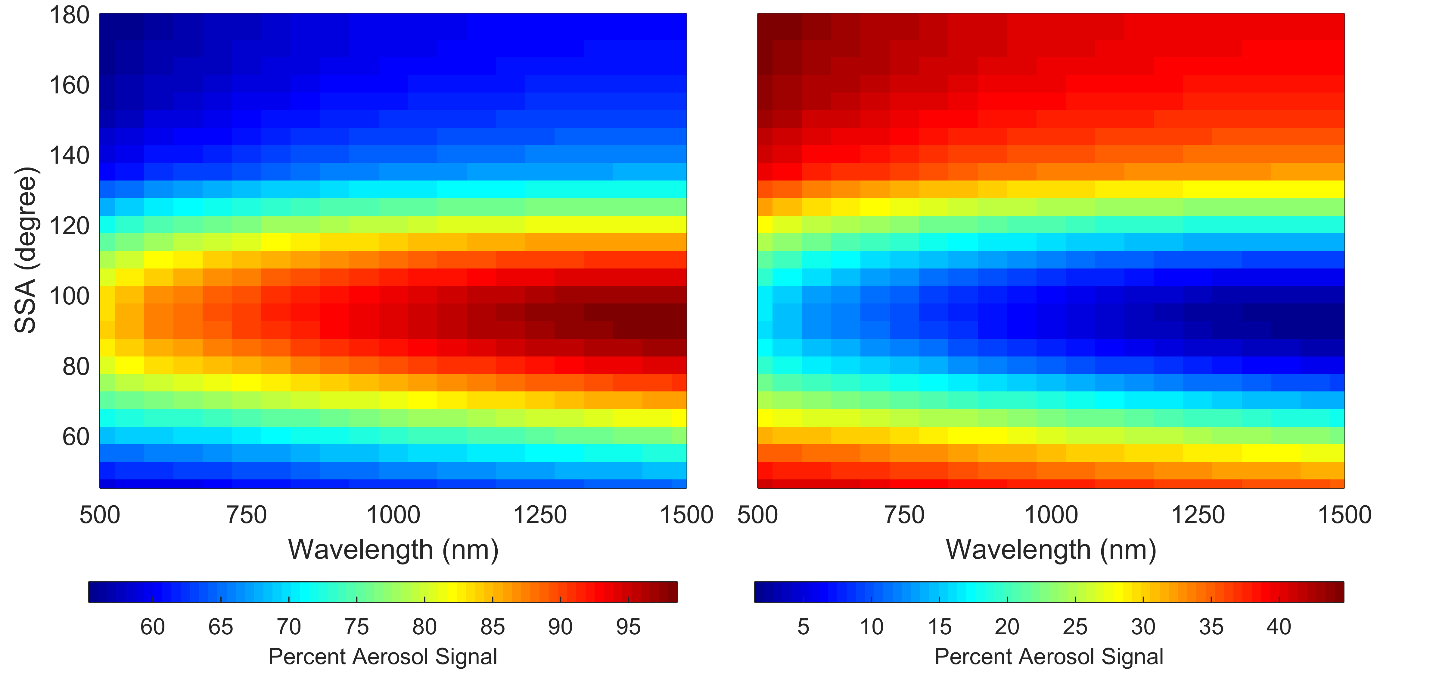
**Figure 4-5**: Dependence of the fraction of the limb spectra due to aerosol on solar scattering angle (left panels) for total radiance (top), horizontal polarization (middle) and vertical polarization (bottom), and the magnitude of the radiance for each case (right panels). Note the low signal near SSA of 90 degrees for the vertical polarization which would be problematic for terminator orbits.

Similar calculations were performed for the range of viewing geometries using the same atmospheric state. Figure 4-5 shows the fraction of limb signal due to aerosol for the total radiance, and both orientations of the linearly polarized radiance. This calculation was performed for 15 km tangent altitude, and other stratospheric tangent altitudes show very similar patterns. An important difference is noted between the forward and back scattering geometries. Remembering the horizontal polarization is given by , the total and horizontal polarization cases have a similar dependence on viewing geometry, with the strongest aerosol signal from long wavelengths in the forward scatter direction. The vertical polarization, given by , has a strong aerosol signal contribution for forward scattering directions, especially at visible wavelengths, in comparison to the total and horizontal polarization cases. For back scattering geometries, somewhat less aerosol signal is observed. For reference, the magnitude of the limb radiance in each case is shown in the right hand column of Figure 4-5 taking note that the high end of the scale is saturated to emphasize the smaller values. It is important to note that the vertical polarization has a very low magnitude at scattering angles near 90 degrees. This makes this combination of viewing geometry and polarization very difficult to use reliably.

We performed these same calculations for the full range of SZAs and found that the SZA only effects the fraction of the signal due to aerosol by less than 1%. Also, when the albedo is changed from 0 to 1, the aerosol signal decreases for all polarizations and wavelengths thus reducing overall sensitivity to aerosol as albedo increases in all cases. Note, however, that the SASKTRAN-HR model assumes that all ground reflection is randomly polarized; the addition of a BRDF model may change the sensitivity to aerosol with higher albedo.

This same analysis was also performed for two other additional polarization orientations, the +45 degree and -45 degree linear polarizations (*i.e* and ) to investigate sensitivity to aerosol. It was found that these two polarization orientations had similar aerosol contribution to the total radiance case with approximately a loss of a third of the overall signal.

In general, the contribution to the limb radiance from aerosol for the horizontally polarized and total radiance cases is approximately the same. The vertical polarization has more asymmetry in aerosol signal between forward and back scattering geometries with very low signal magnitude near 90 degrees scattering angle. Given that essentially all low earth orbit scenarios will cover forward and backward scattering angles, including 90 degrees scattering angle, it is clear that the horizontal orientation overall shows a more favorable response to aerosol. This is particularly true for a terminator orbit such as that for OSIRIS.



**Figure 4-6**: The ratio of the linearly polarized radiance to the total radiance for horizontal (left) and vertical (right) orientations. Note that the scale for each plot is different. The simulation was performed with a SSA of 60 degrees with volcanic aerosol loading for a tangent altitude of 20 km.

A distinct disadvantage of measuring a linear polarization rather than the total radiance is the loss of overall signal magnitude. In Figure 4-6, the ratio of the polarized radiance to the total radiance is shown for a series of SSAs for a tangent altitude of 20 km, but in this case using the volcanic aerosol extinction profile, which serves to enhance the fraction of signal due to aerosol. Measuring the horizontal polarization results in observing signal levels approximately one half to two thirds of the total radiance, with the greatest effect at the shorter wavelengths. The other forward scatter geometries are similarly affected. For back scatter geometries, the signal levels are also approximately half of the total radiance, but with less spectral dependence. For solar scattering angles near 90o, the horizontal polarization encompasses a large fraction of the total radiance resulting in signals of 80-95% of the total. Across the full parameter space of viewing geometries, wavelengths, and aerosol loading scenarios, the magnitude of the horizontal polarization is on average 60-70% of the total radiance.

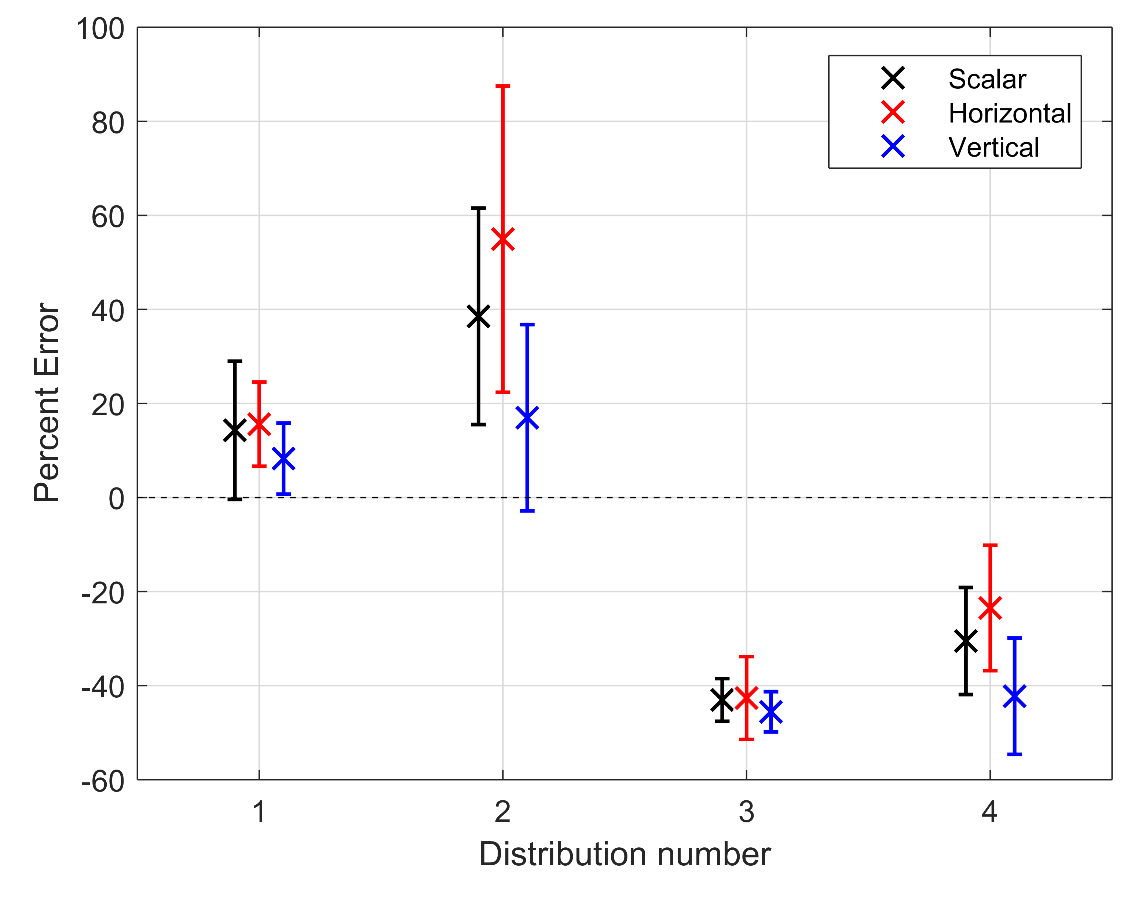
Although the vertical polarization shows a relatively larger fraction of the signal due to aerosol particularly in forward scatter geometries, the overall signal levels are substantially lower. For forward scatter geometries, the magnitude of the vertically polarized radiance is only approximately one third of the total radiance across the spectral range. Back scatter geometries are only slightly better with slightly less than half of the total radiance. Again, near solar scattering angles of 90o the limb radiance is almost fully horizontally polarized and the vertically polarized signal is only 5-20% of the total. On average across the entire parameter space, the vertical polarization is on 30-40% of the total signal.

It is clear from this relatively simplistic analysis of the aerosol signal in polarized limb radiance that there are trade-offs between viewing geometries and polarization orientation, and changing sensitivity across the spectral range. While there is not an overwhelming case to be made for one particular option over the wide range of scenarios that can be considered, the overall response of the horizontally polarized radiance is essentially similar to the total radiance, but with somewhat reduced magnitude that can most likely be mitigated through instrument design considerations. The vertical polarization has much more widely varying sensitivity to aerosol with very low signal levels near 90 degrees scattering angle, and is a much more challenging choice in terms of instrument performance for aerosol measurements.

## 4.4.3 Potential for Retrieval Bias

In this section we directly explore the effect of the polarized measurement on the results of a typical retrieval algorithm through application of the algorithm to simulated measurements across the full parameter space.

We explore the potential of an effect of polarization on the bias in retrieved extinction caused by uncertainty in the assumed particle size distribution. The set of radiances for all cases across the parameter space were again used as simulated input measurements to the retrieval algorithm. This time, retrievals were performed on the horizontally polarized radiance, the vertically polarized radiance, and the total radiance. The radiance calculations in the iterations of the retrieval were set to match the polarization states of the input radiance, but the total solution was used to approximate the total radiance. In all cases, the retrieval was performed using an assumed particle size distribution, which was log-normal with a mode radius and width of 0.08 µm and 1.6 respectively. Note that this assumed size distribution is different than all four of the size distributions used as the “true” state for the simulated input radiances. For the total radiance case, this uncertainty is well known to cause biases of up to 20-30% in retrieved extinction (*Rieger et al.*, 2014). A summary of the differences between the retrieved and true aerosol extinction for 750 nm and 20 km altitude is shown in Figure 4-7. Errors bars on each point represent one standard deviation of the variability in the results for the range of viewing geometries. These results are representative of the level of agreement also found for other wavelengths and altitudes. There is no substantial difference between the results for the background and volcanic extinction profiles. It should be noted that cases with solar scattering angle of 90o have been removed for the vertical polarization due to the very low values of signal, which manifests as a large dependency on the particle size distribution and a highly biased retrieval. This large bias is very sensitive to scattering angle and is nearly eliminated for even 85o or 95o.



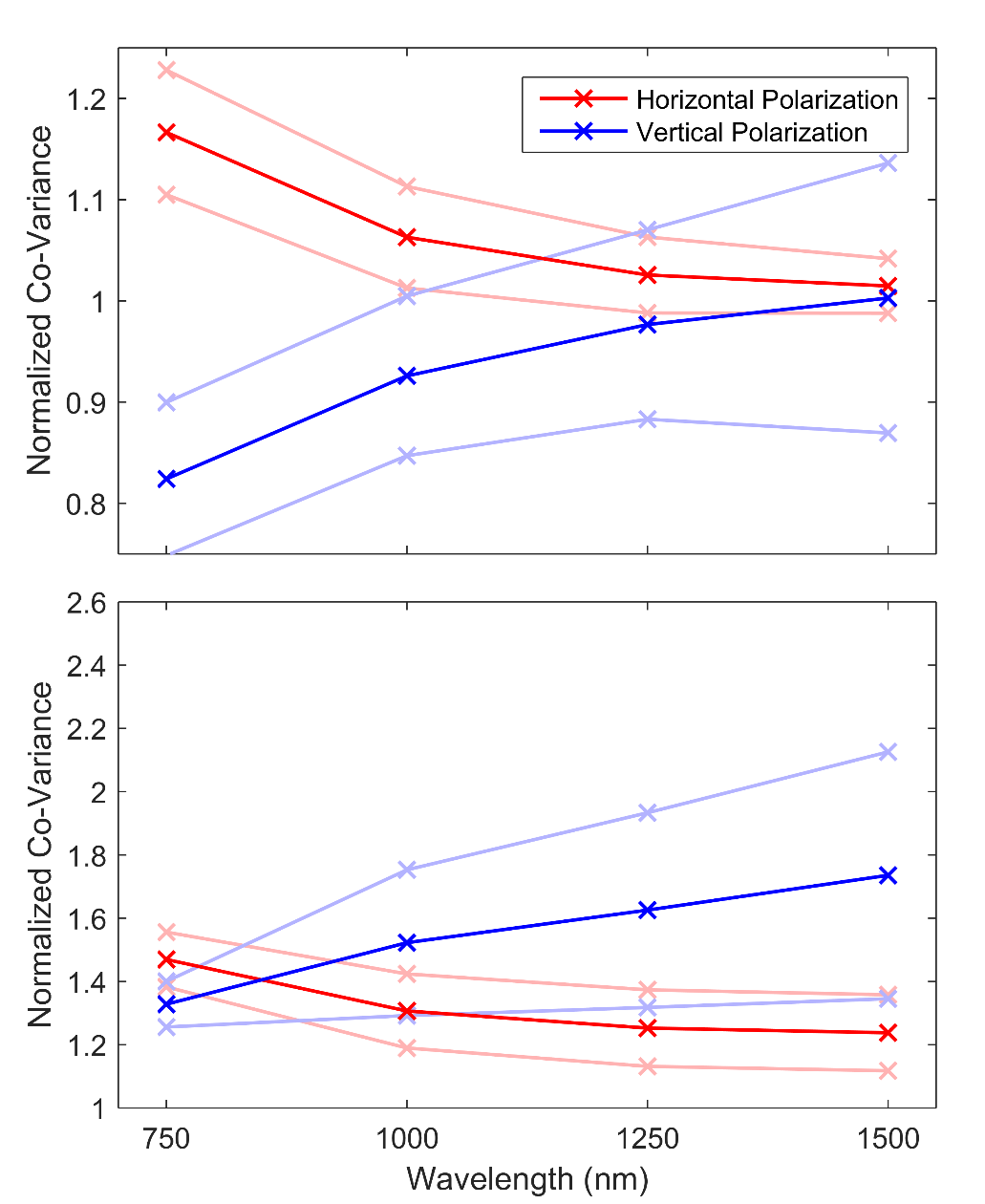
**Figure 4-7**: The mean percent difference between the retrieved aerosol extinction profile with an assumed particle size distribution and the true state corresponding to the indicated particle size distribution (see Table 4-1). Error bars represent one standard deviation of the variability across all viewing geometries. Results shown are for 750 nm and 20 km altitude.

It is clear that the major element of observed bias is simply the difference between the true and assumed particle size distribution. In all four cases there is some difference between the solutions for the various polarization states; however, for each particle size distribution the biases between the polarization states are essentially the same within the variability observed across the various viewing geometries. Neither of the linearly polarized states perform consistently better than the total radiance case; however, they do not perform any worse either (excepting the vertical polarization near 90 degree solar scattering angle), which is an equally important result.

## 4.4.4 Precision analysis

Finally, we study the effect of the polarized measurement on the performance of the retrieval in terms of the precision of the results. We again use simulated measurements across the full range of input parameters as input to the standard retrieval algorithm. Following the methodology outlined in section 4.3, using SASKTRAN-HR the Jacobian matrices were calculated for each retrieved state and used determine the gain matrices, which were then applied as in Equation 4.7 to determine the retrieval precision. It should be noted that not all of the Jacobian matrices could be inverted due to negative sensitivity of the lower tangent altitudes (see discussion in *Bourassa et al.*, 2007) and these were removed from the data set (approximately 9% of total cases). This affected a large fraction of the 500 nm cases, so this wavelength was removed from this section of the analysis. As discussed in section 4.3, we approach this problem from two perspectives: (1) an instrument that is compensated in design and/or operation such that measurements regardless of polarization state or geometries have the same signal to noise ratio, and (2) an uncompensated instrument such that the changing signal level with polarization state and viewing geometry affects the signal to noise ratio of the observation.

For the first case, where the signal to noise ratio is compensated such that it is equal for all cases, the measurement co-variance matrix in Equation 4.7 is replaced with the identity matrix, and the relative size of the square root of the diagonal of the resulting aerosol co-variance represents the amplification of the measurement noise. To compare the performance of the various polarization states, the resulting retrieval co-variances for the linearly polarized measurements were normalized by the retrieval co-variances from the total radiance case. The dependence of the results on the various input parameters, such as wavelength and viewing geometry were examined. Very little altitude dependence was observed and so the results were averaged across the retrieved altitude range.



**Figure 4-8**: The wavelength dependence of the co-variance for the horizontal and vertical polarization retrievals normalized to the total radiance case. The faded line represent one standard deviation of the variability encountered across all input parameters. The top panel is for an instrument design and/or operation that compensates for changing signal levels with polarization and viewing geometry, and the bottom panel is for uncompensated measurements.

In Figure 4-8 the normalized co-variances cases were sorted by wavelength using all geometries and atmospheric states. These bins were then averaged for each wavelength shown with by the red and blue points for the horizontal and vertical polarization respectively. The fainted colours are one standard deviation from the mean. Each of the means in Figure 4-8 contains between 186 to 229 unique data points and values less than one represent co-variance better than the total radiance case and the opposite for values larger than one.

The resulting normalized co-variances have a substantial dependence on wavelength. The situation where the signal to noise ratio is compensated such that it is equal for all cases is shown in the top panel of Figure 4-8, where the vertical polarization has a smaller co-variance, i.e. better precision, by approximately 20% at the shorter wavelengths than the total radiance retrieval. As wavelength increases to 1500 nm, the precision of the vertical polarization case is approximately equal to that of the total radiance case. The horizontal polarization essentially mirrors the vertical case with higher co-variances than total radiance at short wavelengths and approximately equal at 1500 nm. Recall, however, that the vertical polarization has significantly lower magnitude signal levels and in order for the measurement to be compensated to obtain equal signal to noise levels, an increase in instrument sensitivity or exposure time would be required.

**Table 4-2**: The SSA dependence of the normalized co-variance for the horizontal and vertical polarization retrievals. The given numbers are the mean with the standard deviation for each geometry across all wavelengths. Note that the SSA of 90o for the vertical polarization has been removed due to the poor signal in this region.

|  |  |  |  |  |  |
| --- | --- | --- | --- | --- | --- |
| Polarization | 60o | 90o | 120o | 150o | 180o |
| Horizontal (Compensated) | 1.072±0.051 | 1.090±0.082 | 1.047±0.078 | 1.027±0.048 | 1.023±0.042 |
| Vertical (Compensated) | 0.861±0.084 | -- | 0.968±0.157 | 0.977±0.063 | 0.980±0.051 |
| Horizontal (Uncompensated) | 1.289±0.075 | 1.225±0.158 | 1.261±0.165 | 1.341±0.098 | 1.360±0.085 |
| Vertical (Uncompensated) | 1.576±0.167 | -- | 1.852±0.559 | 1.527±0.170 | 1.490±0.130 |

Across the range of solar scattering angles, the vertical polarization has slightly lower co-variance than the other two cases, except at 90 degrees, which is due to the lack of sensitivity in this region noted previously. Table 4-2 shows the calculated means and standard deviations across SSA for the horizontal and vertical polarizations, note the SSA of 90o is missing due to the poor signal, and retrieval quality noted in previous sections which results in poor co-variances for this geometry. Furthermore, the variability of the result across all of the other input parameters increases dramatically as the scattering angle approaches 90 degrees. The precision of the retrieval shows very little dependence on the other input parameters such as solar zenith angle, albedo, particle size distribution, and extinction level. On average across all parameters, the retrieved co-variance from the vertical polarization is approximately 15% smaller than the horizontally polarized retrieval.

In the case of an uncompensated instrument, for example where a linear polarizer is added to the optical chain with no other changes in observation, the scaling of the diagonal of the measurement co-variance is used as outlined in section 4.3. Due to the larger magnitude of the signal in the horizontal polarization compared to the vertical polarization, the horizontal cases generally have lower retrieval co-variance, and this effect increases with longer wavelength as seen in the lower panel of Figure 4-8. Note that in this uncompensated case, since the linear polarizations are always some fraction of the total radiance, the co-variance is always larger than the total radiance case (i.e. the normalized co-variances are always greater than 1). Once again very little dependence on solar zenith angle, albedo, size distribution or extinction level was observed. There was also little dependence on solar scattering angle, except for vertical polarization at 90 degrees. On average across all parameters, vertical and horizontal polarizations have approximately a 60% and 30% larger uncertainty than the total radiance case, respectively.

This analysis shows that the main driver of retrieval precision is the signal to noise level of the observation, as would be expected. Again, this leads to instrument design and/or operational considerations in order to maintain retrieval precision at the same level as the total radiance measurement. A main scientific goal of both the ALI and ALTIUS instruments is obtaining high spatial resolution observations, both vertically and horizontally along, and across, the satellite track. This generally means that images must be collected rapidly and long exposure times are not an affordable luxury. Once more, the relatively higher magnitude signal levels of the horizontal polarization point to this as the more appealing choice of orientation; however, compared to the total radiance case the decreased precision is exaggerated at shorter wavelengths.

# 4.5 Conclusions

We have attempted to address the question of whether or not the measurement of linearly polarized radiance rather than total limb radiance is an advantage or disadvantage with respect to retrievals of stratospheric aerosol. The sensitivity of the polarized limb radiance to aerosol is complex with respect to many parameters, and there are trade-offs in the orientation of the polarization and the orbital viewing geometry. One important point is the very low magnitude of signal observed in the vertical polarization for scattering angles near 90 degrees, which are encountered for a large fraction of observations in low earth orbit, particularly sun-synchronous near-terminator orbits such as for OSIRIS. More generally, it is important to consider the overall lower magnitude of the linearly polarized radiance, which by definition is a fraction of the total radiance. The horizontal polarization has, on average, higher signal levels than the vertical polarization. It also has a weaker dependence on solar scattering angle that is more similar to the total radiance than the vertical polarization, which is more strongly skewed in the forward scatter direction.

One critical bias in limb scatter retrievals of stratospheric aerosol is that due to uncertainty in particle size parameters. We tested four different particle size distribution scenarios, representing background and volcanically perturbed conditions, over a large range of other parameters such as wavelength, viewing geometry, and extinction level and found that there is no significant change in the observed bias for polarized or total radiance measurements. So, with respect to this bias, the polarized measurement is neither an advantage or a disadvantage.

The polarized measurement can have an effect on the precision of the retrieval, where again the main driver is the magnitude of the signal. The polarization can either increase or decrease the precision and there is an apparent trade-off between signal levels and performance of the vertical or horizontal polarization. This is mostly important at shorter wavelengths since at longer wavelengths the differences are negligible.

In conclusion, we have found no clear advantage to the linearly polarized measurement over the total radiance for aerosol retrievals; however, there are also no clear disadvantages assuming the somewhat lower overall signal levels can be handled in the instrument design or operation. With careful choice of the orientation of the measured polarization with respect to the orbital geometry and desired coverage, an instrument such as ALI or ALTIUS is fully capable of obtaining retrieved aerosol products of very similar quality to an equivalent instrument that measures the total radiance.