

# Indian Institute of Technology, Kanpur Department of Earth Sciences

ES0213A: Fundamentals of Earth Sciences

### Lecture 07. Earth's Energy

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### Aims of this lecture



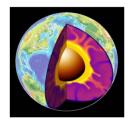
- Earth's resources of energy
- Atmospheric effect on Energy radiation
- Greenhouse effect
- Earth's Energy balance

# Global energy Sources



### **Internal Heat**

- Earth's matter gravity differentiation by density. It results in Earth's stratification into the high-density iron-oxide core, the residual silicate mantle, the light aluminosilicate crust, and the hydrosphere with the atmosphere.
- Decay of radioactive elements causing the release of heat energy



We will have a dedicated class on this...

### **External Heat**

- From Sun
- From impact



### SUN – our star, the furnace

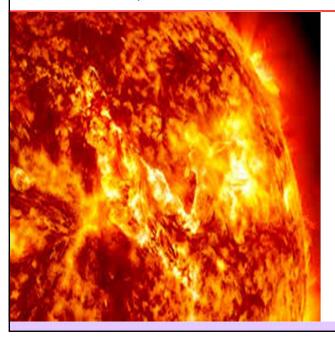




- The sun is a furnace that produces heat energy when H nuclei undergo fusion to He.
- The sun consists of unburned H and He residue
- Sun is considered to be a BLACK BODY as it absorbs all light shined upon it and it emits a spectrum of light (a function of temperature).
- Luminosity (power output): 3.865 x 10<sup>26</sup> Watts.
   Since its birth, it has lost about 30% of its luminosity.
- It still has a life about 3-4 billion years from now, till its H reservoir runs out

# SUN – our star, the furnace





In context to our Earth's energy, the SUN delivers two "products" we are concern about -

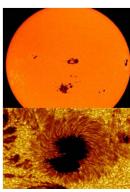
PLASMA

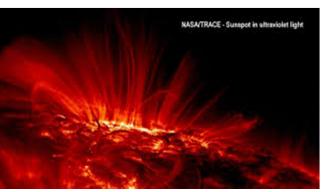
**RADIATION** 

### **PLASMA**

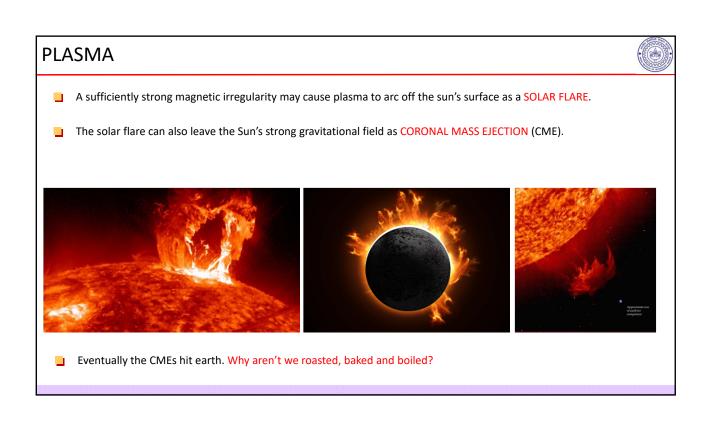


- In the Sun's core, T = 15.6 million K and the pressure is 300 billions times that of Earth's. T on sun's surface is
- These extreme conditions break the Sun's H and He atoms into a swirling mass of charged particles PLASMA
- These plasma particles are always in motion due to sun's strong magnetic field. A burst of such magnetism produces a SUNSPOT. Cooler than average ~ 4500 K.





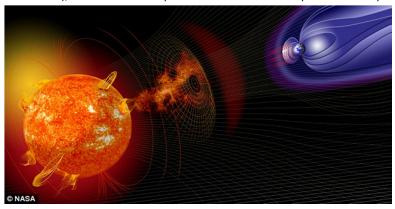
## **PLASMA** On short term, sunspots on 11 year cycle cause a small change in global temperature (~ 0.1 K change) 400 Years of Sunspot Observations Modern 250 Maximum Maunder Minimum 1650 1700 1800 1850 1900 1950 2000 Remember, butterfly effect!



### **PLASMA**



It is our own magnetic field (the protective magnetic shield) that saves us... and 1) produces auroras (borealis and australis), which forms when particles enter earths atmosphere near the poles.



Field Electronics – satellite communications, electrical outages (Quebec 1989).

### **PLASMA**

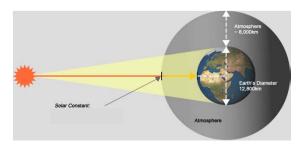


How much time it requires for CMEs (velocity ~ 3 million km/hr) to reach earth?

 $time_{CME}$  = (distance from Sun/ CME velocity) = 50 hours

 $\blacksquare$  How much energy (solar constant,  $S_o$ ) we receive from sun?

[solar constant,  $S_o$  = Luminosity / Area of the coverage sphere]



### Radiation



- The sun continuously emits radiant energy, aka RADIATION
- What is radiation?
  - Transfer of energy that occurs when energized waves travel
  - Does it require a medium? In vacuum?
  - What is the speed of radiation? Constant or variable?
  - Who / what can emit radiation?

### Radiation



Stefan-Boltzmann Law - I: Warmer objects emit radiation more intensely than coo objects do.

 $E = \sigma T^4$ 

: Radiation in Watts/m<sup>2</sup>

: Stefan-Boltsmann constant = 5.67 x 10<sup>-8</sup> Wm<sup>-2</sup> K<sup>-4</sup> σ

: Temperature (in Kelvin)

Wien's law: Warmer objects emit shorter wavelengths; cooler objects emits larger wavelengths

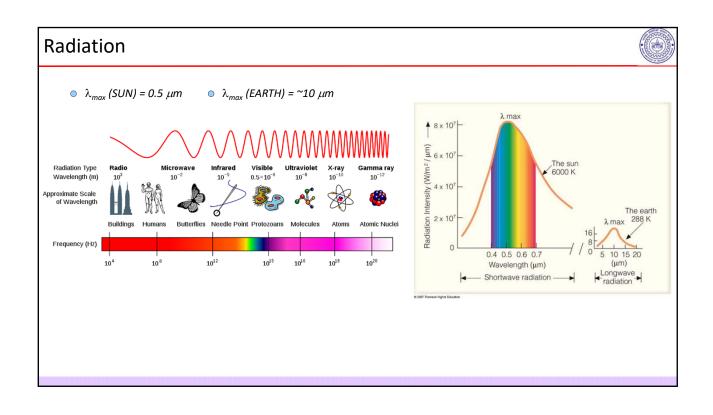
 $\lambda_{max} = \frac{C}{T}$ 

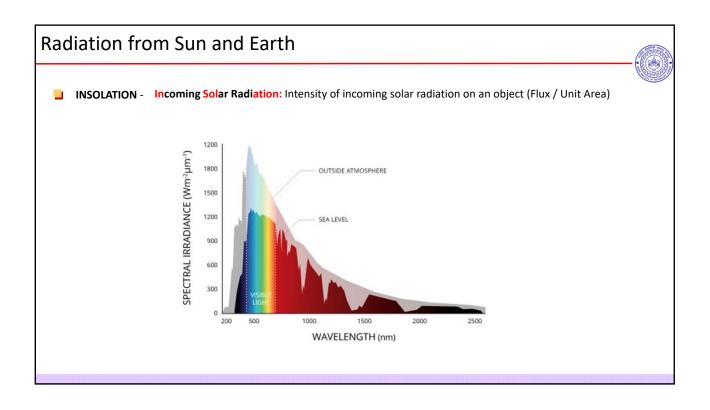
: wavelength (μm)

: constant = 3000  $\mu$ m K

*C* T : Temperature (in Kelvin)

Warmer objects: Large intensity and shorter wavelengths Cooler objects: Less intensity and higher wavelengths





### **Heat Balance**



Stefan-Boltzmann law - II: Each planet must balance net incoming solar radiation with outgoing radiation, determined by its temperature.

$$(1 - \alpha) \pi R^2 S_o = 4\pi R^2 \sigma T^4$$

α : Earth's albedo (30%) R :Radius of Earth (6371 km)

 $S_0$  : Solar Constant for Earth (1367 Wm $^{-2}$ )

σ : Sterfan-Boltzman Constant (5.67 x 10<sup>-8</sup> Wm<sup>-2</sup> K<sup>-4</sup>)

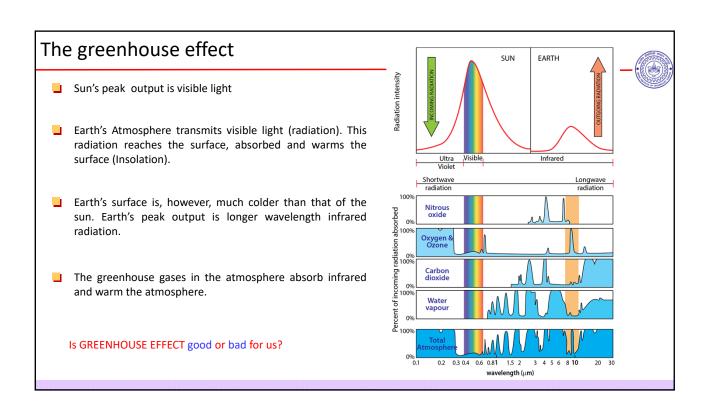
T : Temperature in Kelvin

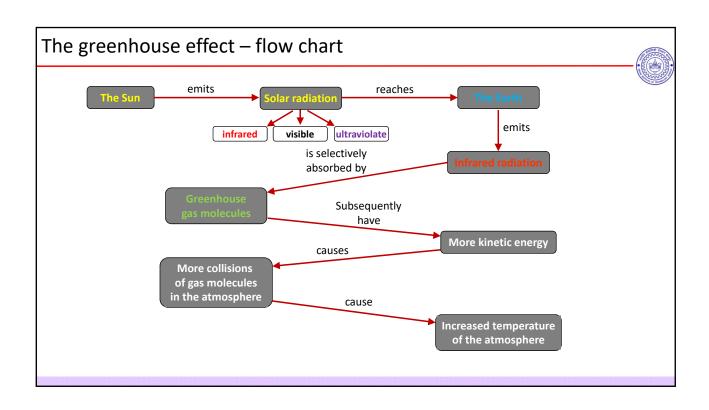
# Temperatures of inner planets

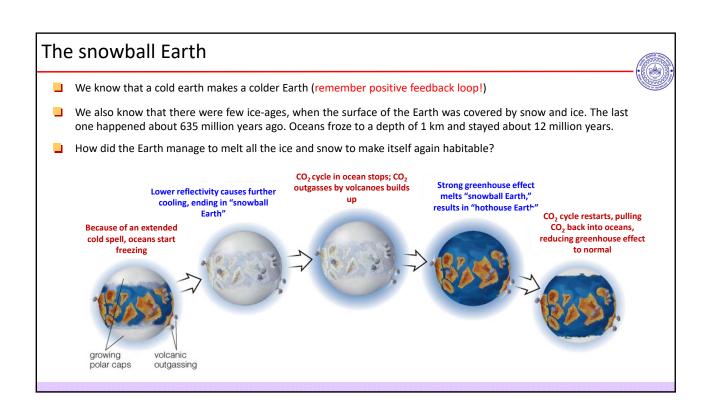


	Albedo (α)	Equilibrium T (°C)	Actual surface T (°C)
Mercury	0.1	162 Just abou	180 ut agrees
Venus	0.59	-10 Disagre	453 es badly
Earth	0.31	-18	15 grees
Mars	0.15	-55 Nearly	-43 agrees

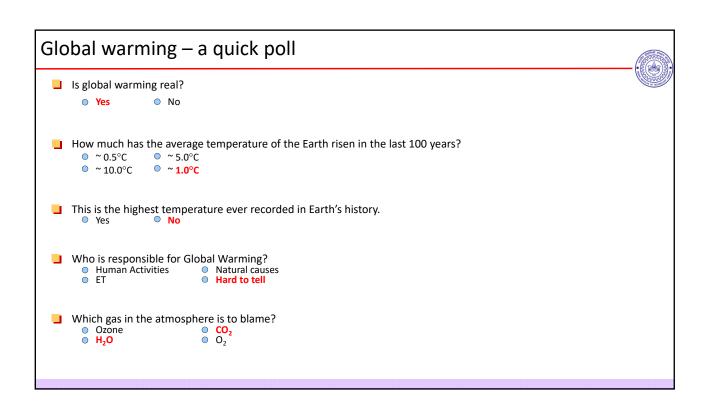
### Temperatures of inner planets Radiative equilibrium works well for Mercury (no **VENUS** atmosphere) and just about for Mars (thin atmosphere).. But not for Earth and Venus. 150 The disagreement for Venus and the Earth is because (km) 100 these two planets have atmospheres containing certain gases which modify their surface temperatures. This is the 'Greenhouse Effect' in action: **MARS** Earth's surface is 34°C warmer than if there were no Troposphere atmosphere 100 200 300 400 500 600 700 800 Temperature (K) Venus has a 'runaway' Greenhouse effect, and is over 400°C warmer Mars atmosphere slightly warms its surface, by about 10°C The existence of the Greenhouse Effect is universally accepted (it is not controversial), and it links the composition of a planet's atmosphere to its surface 0.006 200 Temperature (K) temperature. Copyright © 2005 Pearson Prentice Hall, Inc





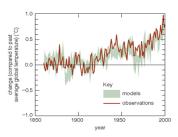


# Global warming — a quick poll Is global warming real? Yes No How much has the average temperature of the Earth risen in the last 100 years? ```O.5°C ```C.0°C ```10.0°C ```10.0°C This is the highest temperature ever recorded in Earth's history. Yes No Who is responsible for Global Warming? Human Activities Natural causes ET Natural causes Hard to tell Which gas in the atmosphere is to blame? Ozone CO2 H20 O20



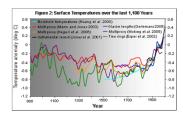
### Global warming - the scenario





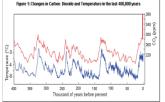
There is a gradual increase in the average temperature of the Earth's atmosphere in the last 100 years...It has risen about 1°C since 1900...

- Are human activities causing global warming?
- What other (non-human) factors can cause global warming?
- How does global warming affect our life?



Reconstructions of surface temperature variations from six research teams. Each curve illustrates a somewhat different history of temperature changes, with a range of uncertainties that tend to increase backward in time (as indicated by the shading).

NRC, 2006. National Academy of Sciences, Courtesy of the National Academies Press. Washington, D.C.



Fluctuations in temperature (blue) and in the atmospheric concentration of carbon dioxide (red) over the past 400,000 years as inferred from Antarctic ice-core records. The vertical red bar is the increase in atmospheric carbon dioxide levels over the past two centuries and before 2006.

A. V. Fedorov et al. Science 312, 1485 (2006).

# What's the big deal if CO<sub>2</sub> causes 1°C temperature increase?



- An increase in atmospheric temperature (human or natural origin) will lead to the increase in the water vapor content of the troposphere.
- Because water vapor is a strong greenhouse gas, the increase in H<sub>2</sub>O vapor in turn causes enhanced greenhouse effect, raising the temperature more.
- Higher atmospheric temperature will cause more evaporation of water
- Which leads to even higher temperature... POSITIVE FEEDBACK LOOP... and effectively culminate to Runaway Greenhouse Effect.

