visualization_multi_a

March 12, 2023

1 my range [20.5,20.65,10]

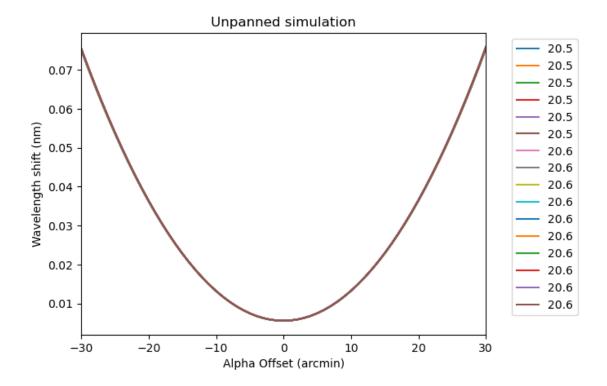
```
[28]: # This file aims to draw conclusion from DN.npz including multi A
      import numpy as np
      import matplotlib.pyplot as plt
      from astropy.modeling import models, fitting
      import math
      from astropy.constants import au, R_sun
      from calculating_DN_2048 import wavelength point num, wavelength list,
       →angle_point_num_alpha
      from calculating_DN_2048 import offaxis_angle_x_alpha, offaxis_angle_y_alpha
      transunit = ((au/R sun).value)**2 / 1000
      DN = np.load("output_DN/DN_my_range_2048.npz")
      DN_alpha_list = DN["DN_alpha_list"] # DN_alpha_list.shape= (10, 61, 15)
      a_list = DN['a_list']
      a_num = len(a_list)
      # Cruciformscan in alpha direction
      offaxis angle x min alpha=offaxis angle x alpha*180*60/math.pi
      # Fit data in DN ??.npz
      wavelength_shift_alpha = np.zeros((a_num, angle_point_num_alpha))
      fit_alpha = [] # List of Gaussian1D
      for i in range(a_num):
          for j in range(angle_point_num_alpha):
              g_init = models.Gaussian1D(amplitude=1E9, mean=0.05, stddev=0.0424)
              # initial value for fitting
             fit_g = fitting.LevMarLSQFitter()
              g = fit_g(g_init, wavelength_list, DN_alpha_list[i][j])
              wavelength_shift_alpha[i][j] = g.mean.value
             fit_alpha.append(g)
```

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[29]: a_list/transunit
```

```
[29]: array([20.5, 20.51, 20.52, 20.53, 20.54, 20.55, 20.56, 20.57, 20.58, 20.59, 20.6, 20.61, 20.62, 20.63, 20.64, 20.65])
```

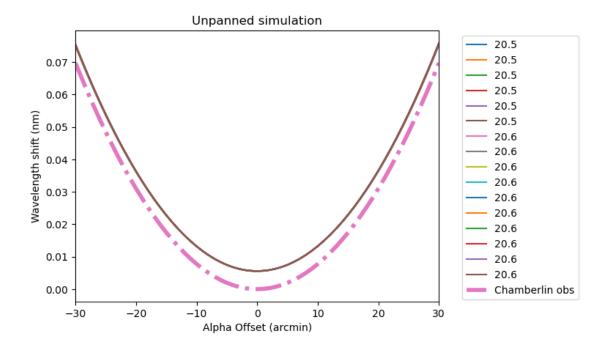
1.1 Multi A

[30]: <matplotlib.legend.Legend at 0x17bb517f280>

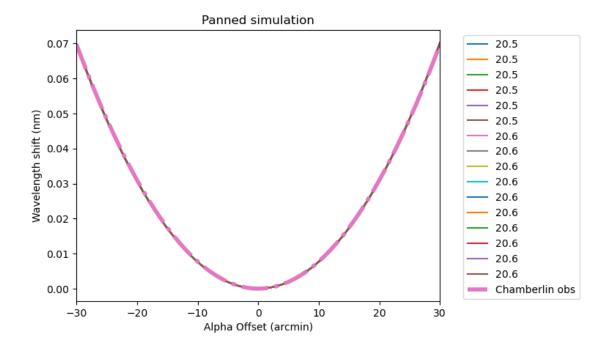


```
[31]: fig, ax = plt.subplots()
for i in range(a_num):
```

[31]: <matplotlib.legend.Legend at 0x17bb5316f10>



[32]: <matplotlib.legend.Legend at 0x17bb52e5f70>



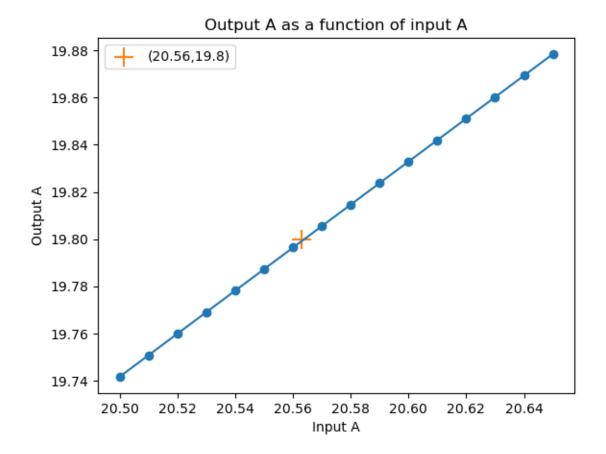
```
[33]: # %% Fit the output A

# Define the polynomial model
model = models.Polynomial1D(degree=2)

# Define the fitter
fitter = fitting.LevMarLSQFitter()

fitted_model = []
for i in range(a_num):
    # Fit the model to the data
    fitted_model.append(
        fitter(model, offaxis_angle_x_alpha, wavelength_shift_alpha[i]))
```

[34]: <matplotlib.legend.Legend at 0x17bb549b910>



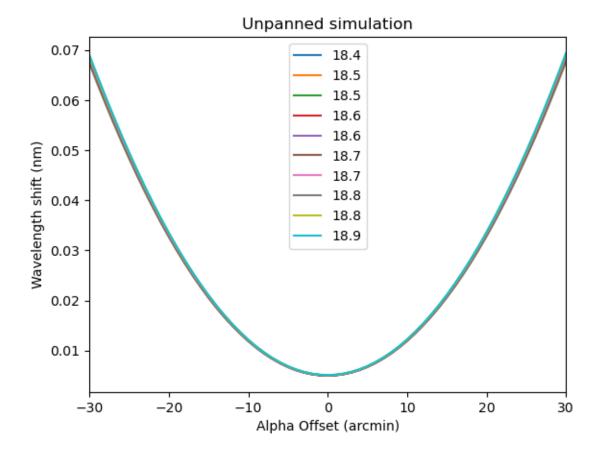
$1.1.1 \quad A=20.72$

2 10a (18.4 18.9 10)

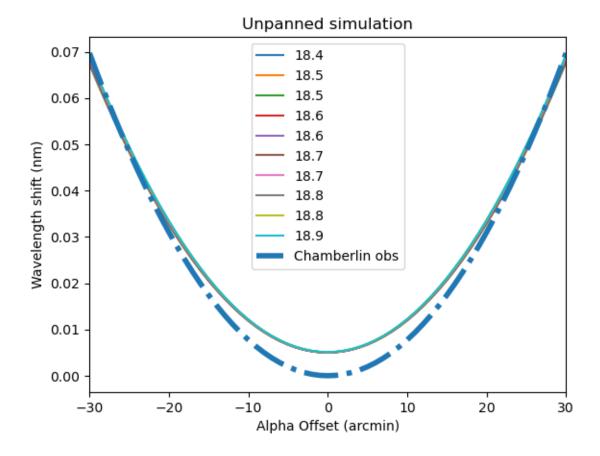
```
[35]: # This file aims to draw conclusion from DN.npz including multi A
      import numpy as np
      import matplotlib.pyplot as plt
      from astropy.modeling import models, fitting
      import math
      from astropy.constants import au, R_sun
      transunit = ((au/R sun).value)**2 / 1000
      DN = np.load("output_DN/DN_10a_2048.npz")
      DN alpha list = DN["DN alpha list"] # DN alpha list.shape= (10, 61, 15)
      a_list = DN['a_list']
      a_num = len(a_list)
      # Initialize
      wavelength point num = 15
      wavelength_list = np.linspace(-0.1, 0.15, wavelength_point_num)
      # Cruciformscan in alpha direction
      angle point num alpha = 61
      DN_alpha = np.zeros((angle_point_num_alpha, wavelength_point_num))
      offaxis_angle_x_alpha = np.linspace(-math.pi /
                                          360, math.pi/360, angle_point_num_alpha)
      offaxis_angle_x_min_alpha=offaxis_angle_x_alpha*180*60/math.pi
      offaxis_angle_y_alpha = np.zeros(angle_point_num_alpha)
      # Cruciformscan in beta direction
      angle_point_num_beta = 61
      DN_beta = np.zeros((angle_point_num_beta, wavelength_point_num))
      offaxis_angle_x_beta = np.zeros(angle_point_num_beta)
      offaxis angle y beta = np.linspace(-math.pi /
                                         360, math.pi/360, angle_point_num_beta)
      offaxis_angle_y_min_beta=offaxis_angle_y_beta*180*60/math.pi
      # Fit data in DN ??.npz
      wavelength_shift_alpha = np.zeros((a_num, angle_point_num_alpha))
      fit_alpha = [] # List of Gaussian1D
      for i in range(a num):
          for j in range(angle_point_num_alpha):
              g_init = models.Gaussian1D(amplitude=1E9, mean=0.05, stddev=0.0424)
              # initial value for fitting
              fit_g = fitting.LevMarLSQFitter()
              g = fit_g(g_init, wavelength_list, DN_alpha_list[i][j])
              wavelength_shift_alpha[i][j] = g.mean.value
```

2.1 Multi A

[37]: <matplotlib.legend.Legend at 0x17bbacf31c0>

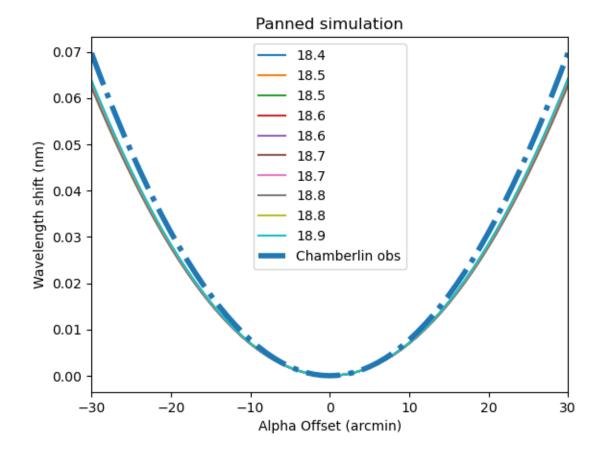


[38]: <matplotlib.legend.Legend at 0x17bbad8bf40>



```
[39]: fig, ax = plt.subplots()
      for i in range(a_num):
          ax.plot(offaxis_angle_x_min_alpha,
                  wavelength_shift_alpha[i]\
                      -wavelength_shift_alpha[i][int(angle_point_num_alpha/2)],
                  label="{:.1f}".format(a_list[i]/transunit))
          # ax.plot(offaxis_angle_x_min_alpha, wavelength_shift_alpha[i],
                    label="\{:.2f\}".format(a\_list[i]/transunit))
      ax.set_xlim(-30, 30)
      ax.plot(offaxis_angle_x_min_alpha,19.
       ⇔8*transunit*offaxis_angle_x_alpha**2,label="Chamberlin obs",
              linestyle='dashdot',linewidth=4)
      ax.set xlabel("Alpha Offset (arcmin)")
      ax.set_ylabel('Wavelength shift (nm)')
      ax.set_title("Panned simulation")
      ax.legend()
```

[39]: <matplotlib.legend.Legend at 0x17bbb32df70>



2.1.1 18.66

3 A: [0,50,26]

```
# Cruciformscan in alpha direction
offaxis_angle_x_min_alpha=offaxis_angle_x_alpha*180*60/math.pi
# Fit data in DN ??.npz
wavelength_shift_alpha = np.zeros((a_num, angle_point_num_alpha))
fit alpha = [] # List of Gaussian1D
for i in range(a_num):
   for j in range(angle point num alpha):
       g_init = models.Gaussian1D(amplitude=1E9, mean=0.05, stddev=0.0424)
        # initial value for fitting
       fit_g = fitting.LevMarLSQFitter()
        g = fit_g(g_init, wavelength_list, DN_alpha_list[i][j])
       wavelength_shift_alpha[i][j] = g.mean.value
        fit_alpha.append(g)
```

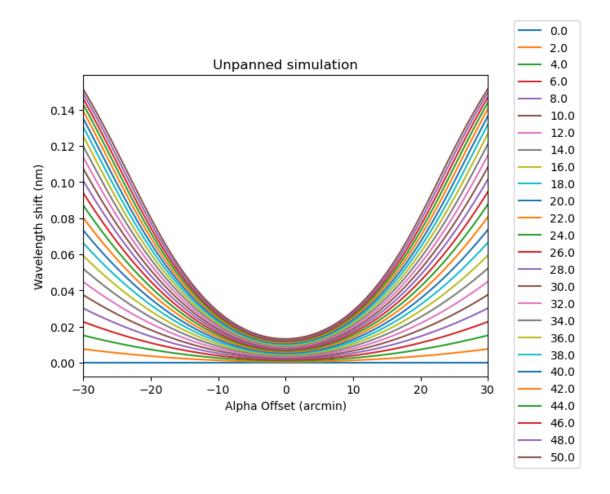
[41]: a_list/transunit

```
[41]: array([ 0., 2., 4., 6., 8., 10., 12., 14., 16., 18., 20., 22., 24.,
            26., 28., 30., 32., 34., 36., 38., 40., 42., 44., 46., 48., 50.])
```

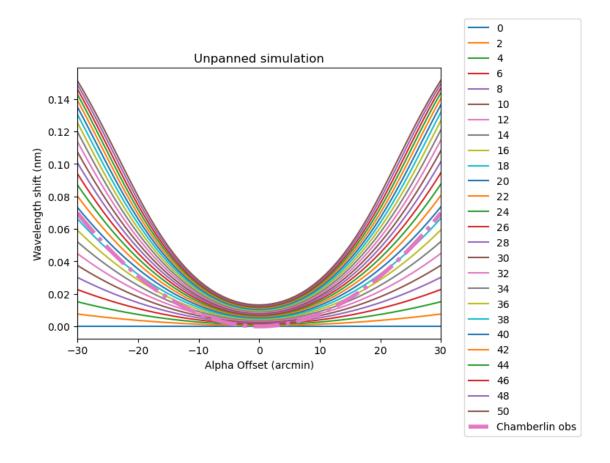
3.1 Multi A

```
[42]: fig, ax = plt.subplots()
      for i in range(a_num):
          # ax.plot(offaxis_angle_x_alpha,
                    wavelength shift alpha[i]\
                        -wavelength_shift_alpha[i][int(angle_point_num_alpha/2)])
          ax.plot(offaxis_angle_x_min_alpha, wavelength_shift_alpha[i],
                  label="{:.1f}".format(a_list[i]/transunit) )
      ax.set xlim(-30,30)
      ax.set_xlabel("Alpha Offset (arcmin)")
      ax.set_ylabel('Wavelength shift (nm)')
      ax.set_title("Unpanned simulation")
      ax.annotate('Label', xy=(3, 6), xytext=(4, 8),
                  arrowprops=dict(arrowstyle='->', lw=2))
      ax.legend(loc='upper left', bbox_to_anchor=(1.05, 1.2))
```

[42]: <matplotlib.legend.Legend at 0x17bbb577760>

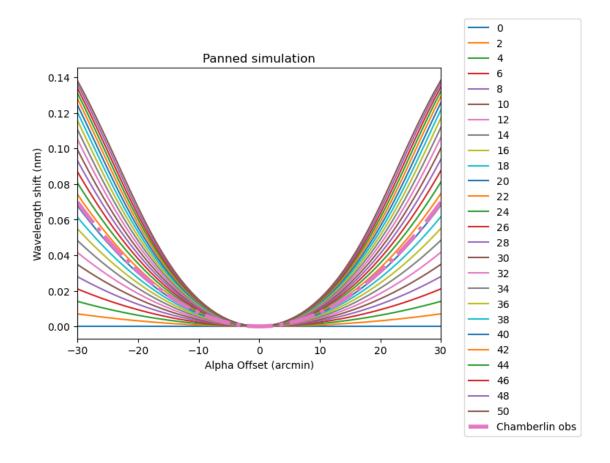


[43]: <matplotlib.legend.Legend at 0x17bb53ed6a0>



```
[44]: fig, ax = plt.subplots()
      for i in range(a_num):
          ax.plot(offaxis_angle_x_min_alpha,
                  wavelength_shift_alpha[i]\
                      -wavelength_shift_alpha[i][int(angle_point_num_alpha/2)],
                  label="{:.0f}".format(a_list[i]/transunit))
          # ax.plot(offaxis_angle_x_min_alpha, wavelength_shift_alpha[i],
                    label="\{:.2f\}".format(a\_list[i]/transunit))
      ax.set_xlim(-30, 30)
      ax.plot(offaxis_angle_x_min_alpha,19.
       ⇔8*transunit*offaxis_angle_x_alpha**2,label="Chamberlin obs",
              linestyle='dashdot',linewidth=4)
      ax.set_xlabel("Alpha Offset (arcmin)")
      ax.set_ylabel('Wavelength shift (nm)')
      ax.set_title("Panned simulation")
      ax.legend(loc='upper left', bbox_to_anchor=(1.05, 1.2))
```

[44]: <matplotlib.legend.Legend at 0x17bba187f10>



```
[45]: # %% Fit the output A

# Define the polynomial model
model = models.Polynomial1D(degree=2)

# Define the fitter
fitter = fitting.LevMarLSQFitter()

fitted_model = []
for i in range(a_num):
    # Fit the model to the data
    fitted_model.append(
        fitter(model, offaxis_angle_x_alpha, wavelength_shift_alpha[i]))

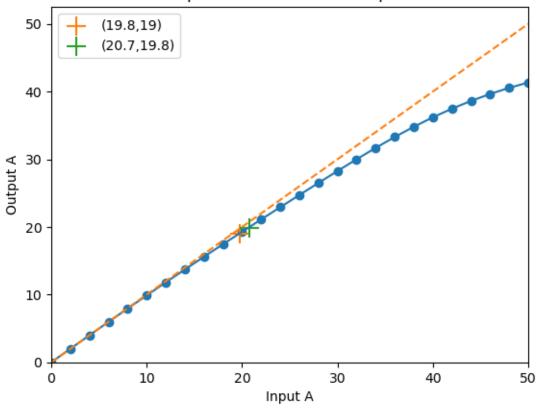
[46]: c2_list = []
for i in range(a_num):
    c2_list.append(fitted_model[i].c2.value)
```

fig, ax = plt.subplots()

ax.plot(a_list/transunit, c2_list/transunit)
ax.scatter(a_list/transunit, c2_list/transunit)

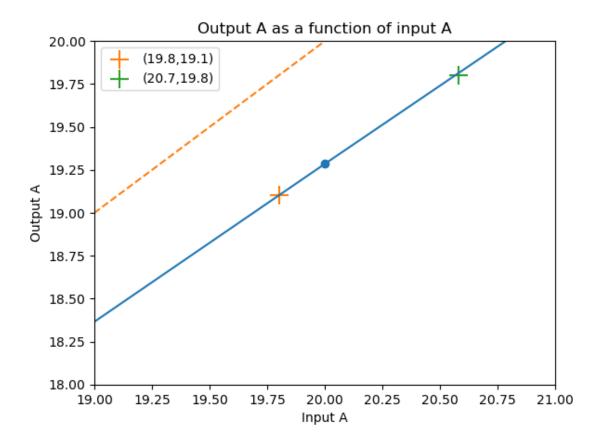
[46]: <matplotlib.legend.Legend at 0x17bba29bfd0>

Output A as a function of input A



```
ax.plot(a_list[8:12]/transunit, c2_list[8:12]/transunit)
ax.scatter(a_list[8:12]/transunit, c2_list[8:12]/transunit)
ax.scatter(19.8,19.1,marker='+',s=200,label="(19.8,19.1)")
ax.scatter(20.58,19.8,marker='+',s=200,label="(20.7,19.8)")
ax.plot(a_list/transunit,a_list/transunit,linestyle="dashed")
# ax.scatter(a_list/transunit,a_list/transunit)
ax.set_xlim(19,21)
# ax.set_ylim(0,)
ax.set_ylim(18,20)
ax.set_ylim(18,20)
ax.set_ylabel("Input A")
# ax.annotate('(19.8,19)', xy=(19.8,19), xytext=(2, 40),
# arrowprops=dict(facecolor='black', shrink=0.05))
ax.set_title("Output A as a function of input A")
ax.legend()
```

[47]: <matplotlib.legend.Legend at 0x17bbabfbdf0>



[]: