



Restoration of images corrupted by Gaussian and uniform impulsive noise

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ABSTRACT

Many approaches to image restoration are aimed at removing either Gaussian or uniform impulsive noise. This is because both types of degradation processes are distinct in nature, and hence they are easier to manage when considered separately. Nevertheless, it is possible to find them operating on the same image, which produces a hard damage. This happens when an image, already contaminated by Gaussian noise in the image acquisition procedure, undergoes impulsive corruption during its digital transmission. Here we propose a principled method to remove both types of noise. It is based on a Bayesian classification of the input pixels, which is combined with the kernel regression framework.

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1. Introduction

Digital image manipulation involves a series of procedures which includes the acquisition and codification of the image in a digital file (image registration) and the transmission of the digital file over some communication channel (image transmission). There is a wide range of processes that affect both procedures negatively. Some of them, such as motion blur [1–3], do not imply information loss. Hence it is conceivable to retrieve the original data by means of finding the inverse of the relevant transformation [4]. Others introduce noise, so that some of the original information is lost. Consequently, it can only be expected to produce a good approximation to the original [5]. In the second case, most techniques rely upon the particular properties of visual data [6].

Even if we restrict our attention to noise removal, there are some different types commonly found in practise. Perhaps the most commonly occurring is additive Gaussian noise [7,8]. It is used to model thermal noise, and under certain conditions it is also the limit of other noises, such as photon counting noise and film grain noise [9]. Its mathematical tractability has led to the proposal of a number of different approaches for its removal. State of the art strategies include wavelets [10,11] and kernel regression [12,13]. Wavelet denoising techniques usually lead to the thresholding (shrinkage) of the corrupted wavelet coefficients [14–17]. Here the difficulty lies on developing an adaptive threshold which takes into account the characteristics of the input data, because excessive shrinking of the coefficients could lead to loss of details. Kernel regression estimates an underlying function corresponding to the

original image from the observed image data. This is done with the help of an adequate weighting of the input data from the pixels which are closest to the particular pixel to be estimated. The output pixels are then weighted averages of the original pixels which lie in their vicinity, and this way the noise is averaged out.

On the other hand, impulse noise arises in digital image transmission over noisy channels, and with faulty equipment [18]. Two kinds of impulse noise are distinguished in literature [19,20]:

- (a) Salt-and-pepper noise, where the impulse noise pixels can only have extreme values. Its detection is relatively easy, as the corrupted pixels differ significantly from their neighbours. Its removal is typically carried out by median filters [21] and other robust statistics approaches [22,23]. There is also a need to preserve the small details of the image, which could be lost if they are mistaken as impulses.
- (b) Uniform noise, where the impulse pixels can have any valid pixel value. In this case, these less outlying values are harder to spot. Consequently, it has received continued attention [18,24,25]. In this paper we are interested in this second class of impulse noise.

There is a fundamental difference among mainstream approaches to Gaussian and impulse noise removal. In the Gaussian case, it is commonly assumed that every pixel can be corrected by subtracting the random additive Gaussian error [26]. In contrast to this, impulse corrupted pixels are conceived as unrecoverable, and the task is to locate and remove the extraneous information they convey [27]. Both assumptions work well for their fields of application, but the situation changes dramatically if Gaussian and uniform noise are present in the same image, a problem which is

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commonly found in practise [28,29]. Typically, the image registration procedure introduces Gaussian noise, and then digital transmission errors produce uniform impulsive noise. In this situation, it is hard to distinguish a Gaussian corrupted pixel which happens to have a high error from an impulse corrupted pixel. Total variation regularization [30,31] has a completely different rationale, since it minimizes a functional that takes into account both the fidelity to the input data and the smoothness of the solution. In principle this makes the total variation strategy less dependent on the type of noise. If applied to an image affected by both Gaussian and uniform noise, the smoothness requirement would take care of the large local changes produced by uniform noise. On the other hand, the Gaussian noise would be averaged out by the combined effects of the fidelity and smoothness requirements. Other approaches include combinations of averaging and robust order statistics [32,20,33]. These strategies improve the results achieved by order statistics (aimed to the removal of impulse corrupted data) by averaging techniques.

Our proposal addresses the problem by estimating the probability that a certain pixel is affected by any of the two kinds of noise. This informs us about the reliability of that pixel, and allows developing an adequate weighting of the input pixels. As we will see, it is suited for heavy noise conditions. Typical scenarios where we find such noise levels include radio frequency interferences which affect wireless transmission of images [34], exposures taken at very high ISO settings [35,36] and photodiode leakage currents in CMOS image sensors [37].

The outline of the paper is as follows. Section 2 presents a probabilistic noise model and a method to learn their parameters from the input image. In Section 3 we obtain a kernel regressor for image restoration which is derived from that noise model. A discussion of the differences among known methods and our proposal is carried out in Section 4. Finally, computational results are shown in Section 5.

2. Noise modelling

2.1. Model definition

Let \mathbf{x}_i be the 2D coordinates of the i -th pixel of an image of size $A \times B$, and let $[0, \nu]$ be the interval of valid pixel values. Our data measurement model assumes that the observed values can be corrupted by Gaussian and uniform impulsive noise. As discussed in [38] and [39], this happens in the real world when the following sequence of events occurs:

- (1) First, a Gaussian noise process affects all pixels. Hence, all the original (uncorrupted) pixel values $z(\mathbf{x}_i) \in [0, \nu]$ are changed to $z(\mathbf{x}_i) + \sigma^2 \varepsilon_i$, where ε_i are independent Gaussian random variables with zero mean and unit variance and $\sigma^2 > 0$ is the common variance of the Gaussian noise. This process will typically result from the physical limitations of the image acquisition procedure: thermal noise, photon counting noise and film grain noise [9], as explained before.
- (2) After that, a uniform impulse noise process operates on the already Gaussian corrupted image. Then the observed value of

the pixel at \mathbf{x}_i is given by

$$y_i = \begin{cases} z(\mathbf{x}_i) + \sigma^2 \varepsilon_i & \text{with probability } 1 - P_{lm} \\ u_i & \text{with probability } P_{lm} \end{cases} \quad (1)$$

where u_i are independent uniform random variables in the interval $[0, \nu]$, and $P_{lm} \in [0, 1]$ is the probability that a given pixel is corrupted by impulse noise. Note that the impulse corrupted pixels do not carry any information about the original image or the previous Gaussian noise. This impulse noise process comes from errors in the subsequent digital processing of the image: transmission errors, faulty storage equipment, and so on.

The overall process is depicted in Fig. 1. It is mathematically expressed by (1), which is equivalently rewritten as

$$y_i = \delta_i(z(\mathbf{x}_i) + \sigma^2 \varepsilon_i) + (1 - \delta_i)u_i \quad (2)$$

where δ_i is a binary random variable which takes the value 0 with probability P_{lm} , and 1 with probability $1 - P_{lm}$.

We call $Im(\mathbf{x}_i)$, $G(\mathbf{x}_i)$ the disjoint random events which happen when the pixel at position \mathbf{x}_i is corrupted by Gaussian and impulse noise or only by Gaussian noise, respectively. This means that

$$y_i = \begin{cases} z(\mathbf{x}_i) + \sigma^2 \varepsilon_i & \text{iff } G(\mathbf{x}_i) \\ u_i & \text{iff } Im(\mathbf{x}_i) \end{cases} \quad (3)$$

2.2. Model learning

The noise model defined in the previous subsection depends on two free parameters, namely σ^2 and P_{lm} . In order to learn those parameters, we consider the error e_i at position \mathbf{x}_i :

$$e_i = y_i - z(\mathbf{x}_i) = \begin{cases} \sigma^2 \varepsilon_i & \text{iff } G(\mathbf{x}_i) \\ u_i - z(\mathbf{x}_i) & \text{iff } Im(\mathbf{x}_i) \end{cases} \quad (4)$$

The distribution of the random quantity $u_i - z(\mathbf{x}_i)$ depends on the particular statistics of the original image. Nevertheless, we can simplify the situation by assuming that the pixel values of the original image are uniformly distributed on the interval $[0, \nu]$. As proven in Appendix A, this implies that $u_i - z(\mathbf{x}_i)$ has a triangular distribution with zero mode, minimum value $-\nu$ and maximum value ν , since it is the difference of two uniform random variables in the interval $[0, \nu]$. This allows expressing the probability density of the error e_i as a probabilistic mixture of a Gaussian density and a triangular density:

$$p(e_i) = (1 - P_{lm})N_\sigma(e_i) + P_{lm}Tri_\nu(e_i) \quad (5)$$

where the triangular probability density function (pdf) is

$$Tri_\nu(e_i) = \begin{cases} \frac{\nu + e_i}{\nu^2} & \text{if } -\nu \leq e_i \leq 0 \\ \frac{\nu - e_i}{\nu^2} & \text{if } 0 \leq e_i \leq \nu \\ 0 & \text{otherwise} \end{cases} \quad (6)$$

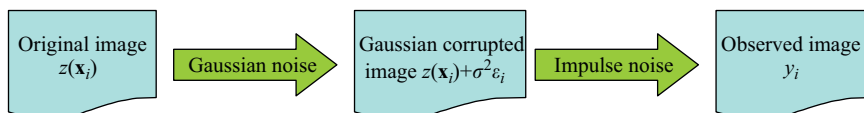


Fig. 1. Data measurement model.

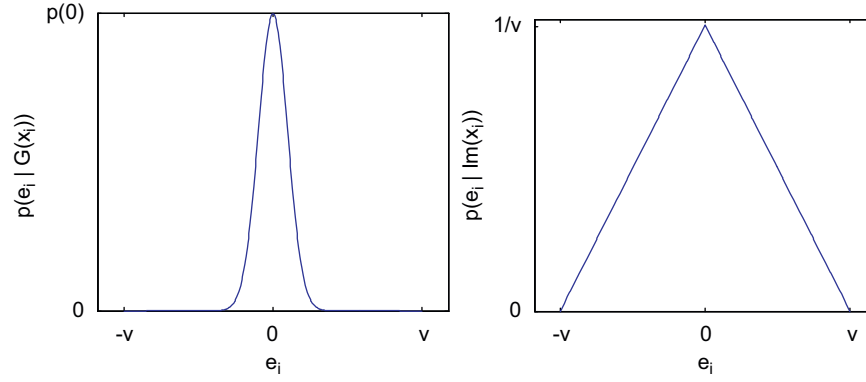


Fig. 2. Probability density function $p(e_i)$ of the error e_i for non impulse corrupted pixels (left) and impulse corrupted pixels (right).

and the Gaussian pdf is given by

$$N_\sigma(e_i) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{e_i^2}{\sigma^2}\right) \quad (7)$$

To get a clearer picture of the situation, please see Fig. 2, where the conditional density functions of the error are shown.

A method is needed to estimate the free parameters σ^2 and P_{lm} from the errors e_i corresponding to the pixels of the image. In particular, we wish to obtain estimators which maximize the data likelihood under the mixture model:

$$L(\sigma, P_{lm}) = \sum_i \log p(e_i | \sigma, P_{lm}) \quad (8)$$

This is accomplished by a specific version of the Expectation-Maximization (EM) algorithm [40,41], which is developed in Appendix B. The update equations for each iteration t of the EM algorithm read:

$$P_{lm}(t+1) = \frac{1}{AB} \sum_i R_{lm,i,t} \quad (9)$$

$$\sigma(t+1) = \sqrt{\frac{\sum_i R_{G,i,t} e_i^2}{AB(1-P_{lm}(t+1))}} \quad (10)$$

where

$$R_{lm,i,t} = \frac{P_{lm}(t) \text{Tri}_v(e_i)}{p(e_i | \theta(t))} \quad (11)$$

$$R_{G,i,t} = \frac{(1-P_{lm}(t)) N_{\sigma(t)}(e_i)}{p(e_i | \theta(t))} \quad (12)$$

and the iteration is continued until convergence.

There is an additional issue. The EM algorithm accepts the errors e_i as inputs, but these values are unknown. We propose to use another image restoration technique to produce predictions of the pixel values $\tilde{z}(\mathbf{x}_i)$, so that the predicted errors can be also computed:

$$\tilde{e}_i = y_i - \tilde{z}(\mathbf{x}_i) \quad (13)$$

Finally, the error predictions \tilde{e}_i are fed to the EM algorithm. The more accurate the pixel predictions $\tilde{z}(\mathbf{x}_i)$, the best estimators for the free parameters σ^2 and P_{lm} we get. We have tested some image restoration techniques, and the iteratively reweighted norm (IRN) approach [30] has been found to work well for this purpose. Hence, we have chosen it for the experiments. The IRN method is based on the minimization of a functional which includes two terms. One of them measures the fidelity of the reconstruction to the input image, and it depends on a weighted L2-norm of the differences among the reconstructed and the original pixel values. The second term is called the regularization

term, and it penalizes reconstructed images with high gradient values. This is aimed to reduce the noise, which is typically accompanied by sharp changes in the pixel values. Then the Jacobian and the Hessian of the functional are obtained, and finally the minimization is carried out by a variation of Newton's method.

3. Kernel regression

3.1. Adaptive image restoration

In the classic 2D kernel regression framework, we would try to estimate $z(\mathbf{x}_i)$ as the mean regression function of the observed data:

$$y_i = z(\mathbf{x}_i) + e_i \Rightarrow z(\mathbf{x}_i) = E[y_i] \quad (14)$$

since $E[e_i] = 0$. Here our alternative approach is to estimate $z(\mathbf{x}_i)$ as the mean regression function of the *non impulse-corrupted observed data*:

$$z(\mathbf{x}_i) = E[y_i | G(\mathbf{x}_i)] \quad (15)$$

Please note that, while the mean regression function to be estimated is the same in both cases,

$$E[y_i] = E[y_i | G(\mathbf{x}_i)] \quad (16)$$

the second option is more convenient, since the Gaussian noise variance σ^2 is expected to be lower than the variance of the combined (Gaussian and impulse) noise:

$$\text{var}[y_i | G(\mathbf{x}_i)] = \text{var}[e_i | G(\mathbf{x}_i)] = \sigma^2 \quad (17)$$

$$\text{var}[y_i] = \text{var}[e_i] = (1-P_{lm})\sigma^2 + P_{lm} \frac{v^2}{6} \quad (18)$$

$$\text{var}[y_i] > \text{var}[y_i | G(\mathbf{x}_i)] \quad (19)$$

where $\text{var}[e_i]$ is obtained in Appendix C, and $\sigma \ll v$. The above equation is the mathematical expression of the following fact: impulse corrupted pixels do not carry any information about the original image.

Let $\mathbf{x} \in [1, A] \times [1, B]$ be a position on the image, which may or may not be coincident with a pixel position, i.e., subpixel accuracy is allowed. The local kernel estimator in the vicinity of \mathbf{x} is given by

$$z(\mathbf{x}_i) = z(\mathbf{x}) + (\nabla z(\mathbf{x}))^T (\mathbf{x}_i - \mathbf{x}) + \frac{1}{2} (\mathbf{x}_i - \mathbf{x})^T (\mathcal{H}z(\mathbf{x})) (\mathbf{x}_i - \mathbf{x}) + \dots \quad (20)$$

where ∇ and \mathcal{H} are the gradient and Hessian operators, respectively, and $G(\mathbf{x}_i)$ verifies, i.e., the input pixel at position \mathbf{x}_i is Gaussian corrupted. If we take into account the symmetry of the

Hessian matrix, we may write

$$z(\mathbf{x}_i) = \beta_0 + \beta_1^T(\mathbf{x}_i - \mathbf{x}) + \beta_2^T \text{svec}((\mathbf{x}_i - \mathbf{x})(\mathbf{x}_i - \mathbf{x})^T) + \dots \quad (21)$$

where $\text{svec}(\cdot)$ is a vectorization of a symmetric matrix,

$$\text{svec}\left(\begin{bmatrix} a & b \\ b & c \end{bmatrix}\right) = (a \ b \ c)^T \quad (22)$$

and the parameters to be determined are:

$$\beta_0 = z(\mathbf{x}) \quad (23)$$

$$\beta_1 = \nabla z(\mathbf{x}) = \left[\frac{\partial z(\mathbf{x})}{\partial x_1}, \frac{\partial z(\mathbf{x})}{\partial x_2} \right]^T \quad (24)$$

$$\beta_2 = \frac{1}{2} \left[\frac{\partial^2 z(\mathbf{x})}{\partial x_1^2}, 2 \frac{\partial^2 z(\mathbf{x})}{\partial x_1 \partial x_2}, \frac{\partial^2 z(\mathbf{x})}{\partial x_2^2} \right]^T \quad (25)$$

We group the parameters for notational convenience:

$$\mathbf{b} = [\beta_0, \beta_1^T, \dots, \beta_N^T]^T \quad (26)$$

with N the order of the estimation.

Note that $\beta_0 = z(\mathbf{x})$ is the estimated image value at \mathbf{x} . These parameters are obtained by solving the following optimization problem, where only the Gaussian corrupted pixels are considered:

$$\hat{\mathbf{b}} = \arg \min_{\mathbf{b}} F(\mathbf{b}) \quad (27)$$

$$F(\mathbf{b}) = \sum_i \delta_i K_i(\mathbf{x}_i - \mathbf{x}) [y_i - \beta_0 - \beta_1^T(\mathbf{x}_i - \mathbf{x}) - \beta_2^T \text{svec}((\mathbf{x}_i - \mathbf{x})(\mathbf{x}_i - \mathbf{x})^T) - \dots]^2 \quad (28)$$

In the above equation K_i is the 2D smoothing kernel function for pixel i , which will be studied in the next subsection.

Since the value of the random variable δ_i for a pixel i is not known, that is, whether i is Gaussian or impulse corrupted, objective function (28) cannot be evaluated directly. Instead of this, we substitute it by its expectation under the observed pixel value y_i , so the optimization problem to be solved in practise is

$$\hat{\mathbf{b}} = \arg \min_{\mathbf{b}} E[F(\mathbf{b})|y_i] \quad (29)$$

$$E[F(\mathbf{b})|y_i] = \sum_i E[\delta_i | y_i] K_i(\mathbf{x}_i - \mathbf{x}) [y_i - \beta_0 - \beta_1^T(\mathbf{x}_i - \mathbf{x}) - \beta_2^T \text{svec}((\mathbf{x}_i - \mathbf{x})(\mathbf{x}_i - \mathbf{x})^T) - \dots]^2 \quad (30)$$

The relevant expectations are computed by Bayes' theorem:

$$E[\delta_i | y_i] = P(G(\mathbf{x}_i) | y_i) = P(G(\mathbf{x}_i) | \tilde{e}_i) = \frac{p(\tilde{e}_i | G(\mathbf{x}_i))}{p(\tilde{e}_i)} \quad (31)$$

where the probability densities p come from the mixture model trained in Section 2, that is, we are using the predicted errors for approximation:

$$p(\tilde{e}_i) \approx p(e_i) \quad (32)$$

$$p(\tilde{e}_i | G(\mathbf{x}_i)) \approx p(e_i | G(\mathbf{x}_i)) \quad (33)$$

Eq. (31) can be rewritten by using (5) to yield a more explicit formulation:

$$E[\delta_i | y_i] = P(G(\mathbf{x}_i) | y_i) = \frac{(1 - P_{lm}) N_\sigma(\tilde{e}_i)}{(1 - P_{lm}) N_\sigma(\tilde{e}_i) + P_{lm} \text{Tri}_v(\tilde{e}_i)} \quad (34)$$

Next we rewrite the problem (29) in matrix form:

$$\hat{\mathbf{b}} = \arg \min_{\mathbf{b}} \|\mathbf{y} - \mathbf{X}_x \mathbf{b}\|_{\mathbf{W}_x}^2 = \arg \min_{\mathbf{b}} (\mathbf{y} - \mathbf{X}_x \mathbf{b})^T \mathbf{W}_x (\mathbf{y} - \mathbf{X}_x \mathbf{b}) \quad (35)$$

where

$$\mathbf{y} = (y_1, \dots, y_M)^T \quad (36)$$

$$\mathbf{W}_x = \text{diag}[E[\delta_1 | y_1] K_1(\mathbf{x}_1 - \mathbf{x}), \dots, E[\delta_M | y_M] K_M(\mathbf{x}_M - \mathbf{x})] \quad (37)$$

$$\mathbf{X}_x = \begin{bmatrix} 1 & (\mathbf{x}_1 - \mathbf{x})^T & \text{svec}^T((\mathbf{x}_1 - \mathbf{x})(\mathbf{x}_1 - \mathbf{x})^T) & \dots \\ 1 & (\mathbf{x}_2 - \mathbf{x})^T & \text{svec}^T((\mathbf{x}_2 - \mathbf{x})(\mathbf{x}_2 - \mathbf{x})^T) & \dots \\ \vdots & \vdots & \vdots & \vdots \\ 1 & (\mathbf{x}_M - \mathbf{x})^T & \text{svec}^T((\mathbf{x}_M - \mathbf{x})(\mathbf{x}_M - \mathbf{x})^T) & \dots \end{bmatrix} \quad (38)$$

with 'diag' producing a diagonal matrix, and M being the number of pixels in a suitable neighbourhood V of position \mathbf{x} . This facilitates to obtain its solution by standard weighted least squares theory [42,43]:

$$\hat{\mathbf{b}} = (\mathbf{X}_x^T \mathbf{W}_x \mathbf{X}_x)^{-1} \mathbf{X}_x^T \mathbf{W}_x \mathbf{y} \quad (39)$$

provided that $\mathbf{X}_x^T \mathbf{W}_x \mathbf{X}_x$ is invertible. In our experiments, V has been chosen to comprise the pixels in a circle with fixed radius r , centred in \mathbf{x} .

3.2. Adaptive kernel estimation

The adaptation to the input data is enhanced if we choose the 2D smoothing kernel K_i to depend on the local gradient covariance matrix \mathbf{C}_i (see [13]):

$$K_i(\mathbf{x}_i - \mathbf{x}) = \frac{\sqrt{\det(\mathbf{C}_i)}}{2\pi h^2} \exp\left(-\frac{1}{2h^2} (\mathbf{x}_i - \mathbf{x})^T \mathbf{C}_i (\mathbf{x}_i - \mathbf{x})\right) \quad (40)$$

where h is a global smoothing parameter and \mathbf{C}_i is given by (see [44]):

$$\mathbf{C}_i = \langle (\nabla z(\mathbf{x})) (\nabla z(\mathbf{x}))^T \rangle_i \quad (41)$$

$$\langle \zeta(\mathbf{x}) \rangle_i = \int \int_{V_i} \zeta(\mathbf{x}) d\mathbf{x} \quad (42)$$

with V_i a local neighbourhood of pixel \mathbf{x}_i . Like before, in practise V_i is a circle with fixed radius r , centred in \mathbf{x}_i .

In order to estimate \mathbf{C}_i , Takeda et al. [13] proposed to compute the truncated singular value decomposition (SVD) of the local gradient matrix \mathbf{G}_i :

$$\mathbf{G}_i = \begin{pmatrix} \dots \\ (\nabla z(\mathbf{x}_j))^T \\ \dots \end{pmatrix} = \mathbf{U}_i \mathbf{S}_i \mathbf{V}_i^T, \text{ with } j \in V_i \quad (43)$$

where \mathbf{S}_i is a 2×2 diagonal matrix representing the energy in the dominant directions, and \mathbf{V}_i is a 2×2 orthogonal matrix whose second column $(v_1, v_2)^T$ defines the dominant orientation angle θ_i :

$$\theta_i = \arctan\left(\frac{v_1}{v_2}\right) \quad (44)$$

The local elongation parameter ρ_i is computed from the diagonal elements s_1, s_2 of \mathbf{S}_i :

$$\rho_i = \frac{s_1 + \lambda'}{s_2 + \lambda'} \quad (45)$$

where $\lambda' \geq 0$ is a regularization parameter. The local scaling parameter γ_i is given by:

$$\gamma_i = \sqrt{\frac{s_1 s_2 + \lambda''}{M}} \quad (46)$$

where $\lambda'' \geq 0$ is another regularization parameter, and M stands for the number of pixels in the local neighbourhood V_i .

Then, the local gradient covariance matrix estimator $\hat{\mathbf{C}}_i$, which we will only use for Gaussian corrupted pixels, is obtained as follows:

$$G(\mathbf{x}_i) \Rightarrow \hat{\mathbf{C}}_i = \gamma_i \Theta_i \begin{pmatrix} \rho_i & 0 \\ 0 & \rho_i^{-1} \end{pmatrix} \Theta_i^T \quad (47)$$

$$\Theta_i = \begin{pmatrix} \cos \theta_i & \sin \theta_i \\ -\sin \theta_i & \cos \theta_i \end{pmatrix} \quad (48)$$

Careful estimation of \mathbf{C}_i is of paramount importance for kernel regression in our context, because impulse corrupted pixels may introduce considerable errors. To take this into account, we consider that the best estimation of \mathbf{C}_i for an impulse corrupted pixel is the null matrix, which corresponds to a locally constant image:

$$Im(\mathbf{x}_i) \Rightarrow \hat{\mathbf{C}}_i = \mathbf{0} \quad (49)$$

Hence, the estimator $\hat{\mathbf{C}}_i$ for an arbitrary pixel is derived from (47) and (49):

$$\hat{\mathbf{C}}_i = \delta_i \gamma_i \Theta_i \begin{pmatrix} \rho_i & 0 \\ 0 & \rho_i^{-1} \end{pmatrix} \Theta_i^T \quad (50)$$

As in the previous subsection, the above equation cannot be implemented directly, since we do not know whether a particular pixel i is Gaussian or impulse corrupted. Hence, in practise we use the expectation of $\hat{\mathbf{C}}_i$ under the observed pixel value y_i :

$$E[\hat{\mathbf{C}}_i | y_i] = E[\delta_i | y_i] \gamma_i \Theta_i \begin{pmatrix} \rho_i & 0 \\ 0 & \rho_i^{-1} \end{pmatrix} \Theta_i^T \quad (51)$$

where $E[\delta_i | y_i]$ is obtained from (34), as before.

3.3. Algorithm

In this subsection we specify how the above presented techniques can be combined in order to develop a restoration algorithm.

The first stage involves obtaining the pixel predictions $\tilde{z}(\mathbf{x}_i)$ from IRN or any other suitable method, so that the noise model can be learnt by the EM algorithm (Section 2.2).

Then we execute a preliminary kernel regression. As the gradient is not known, we take

$$E[\hat{\mathbf{C}}_i | y_i] = E[\delta_i | y_i] \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \quad (52)$$

as a first approach. This produces preliminary estimators of the gradient $\nabla z(\mathbf{x})$, which we use to compute $E[\hat{\mathbf{C}}_i | y_i]$ more accurately, by means of the procedure explained in Section 3.2.

Finally, these more accurate values of $E[\hat{\mathbf{C}}_i | y_i]$ are fed into the kernel regression to yield the definitive restored image.

Hence, the algorithm is as follows:

1. Execute the IRN method (or any other) on the input image to yield predictions $\tilde{z}(\mathbf{x}_i)$ of the original pixel values, and compute the corresponding predicted errors \tilde{e}_i by Eq. (13).
2. Train the noise model by the EM algorithm, Eqs. (9)–(10), until convergence. The input samples e_i for this algorithm are approximated by the predicted errors \tilde{e}_i obtained in Step 1.
3. Perform a preliminary kernel regression, Eq. (39), where the local gradient covariance matrices \mathbf{C}_i are tentatively approximated by the expectations $E[\hat{\mathbf{C}}_i | y_i]$ in Eq. (52). This regression produces preliminary approximations of the gradient $\nabla z(\mathbf{x})$.
4. Find more accurate approximations of the local gradient covariance matrices from Eq. (51), where the necessary gradient approximations come from the output of Step 3.
5. Perform the final kernel regression, Eq. (39), where the local gradient covariance matrices \mathbf{C}_i are approximated by the expectations $E[\hat{\mathbf{C}}_i | y_i]$ obtained in Step 4.

It must be noted that the above algorithm not only produces the restored image values from Eq. (23), but also estimates the

gradient from Eq. (24). In order to apply it to colour images, it must be taken into account that in image coding schemes used in practise there is one colour component which is coded with more spatial resolution, namely that which carries the luminosity information. Hence, we always estimate the local gradient from that component, while the kernel regression is performed separately on each component.

As our approach is aimed to restore very noisy images, we call it *Heavily Damaged Image Restoration (HDIR)*.

4. Discussion

Our method can be conceived as a combination of noise mixture modelling with kernel regression, which is designed to remove two kinds of noise occurring in the same image. Hence it departs from previously known methods, although it shares several properties with some of them. Next we consider those approaches which have something in common with ours.

- (a) Impulse noise removers [45,21,22,46] assume that the erroneous pixels differ significantly from their neighbours. We take this general hypothesis one step further by defining a probabilistic noise model. This allows exploiting the particular statistical properties of its randomness in full, and provides a framework to distinguish impulses from Gaussian corrupted pixels in a principled way.
- (b) Kernel regression methods [12,13] suppose that there is an underlying function whose values are corrupted by some process to yield the observed values. In principle they do not assume any probability distribution for this noise process [42]. Our strategy adds a mixture noise model to this framework, so as to improve its performance by considering each input pixel in light of its likelihood of carrying useful information.
- (c) Many wavelet shrinkage methods use statistical models of either the original image or the corrupted one [14,15]. This is the equivalent in the wavelet domain of our probabilistic noise model. In particular, Luisier et al. [16] use Stein's unbiased risk estimation (SURE) to avoid a statistical model for the wavelet coefficients. The relations between SURE and kernel regression have been studied in [47]. On the other hand, Pizurica and Philips [17] estimate the probability that a certain wavelet coefficient contains a significant noise-free component, much like our estimation of the probability that a given pixel is not impulse corrupted. In spite of these similarities and connections, HDIR is fundamentally different from all of them, since these methods work on the wavelet coefficients, while our proposal works directly on the pixel values.

5. Experimental results

Given the wide range of image restoration methods currently available, we have selected 9 of them which belong to very different approaches to this problem. Next we list them, along with the abbreviation used in the following figures and their most outstanding features:

- (a) Iteratively reweighted norm, IRN [30]. As mentioned before, it is a total variation regularization approach which minimizes a functional that takes into account both the fidelity to the input data and the smoothness of the solution.
- (b) Iterative steering kernel regression, ISKR [13]. This is a state-of-the-art kernel regression method. It performs several

- kernel regressions iteratively, so that the noise is progressively removed. On each iteration the local orientation is estimated, so as not to affect the edges of the original image.
- (c) Progressive switching median filter, PSMF [48]. It has two modules: an impulse detector and a restoring filter, both based on the median as a robust statistic. The filter is applied to a pixel only if an impulse is detected, i.e., it is switched on and off.
- (d) Decision-based algorithm for removal of high-density impulse noises, DBAIN [49]. This algorithm is oriented to removing impulse corrupted pixel with extreme values (salt-and-pepper), i.e., either 0 or v . Hence, it is not designed for the kinds of noise we are considering here, but we include it as an illustration of what happens when it faces non extreme values.
- (e) NeighShrink-SURE, NS [14]. This is a wavelet shrinkage method that determines an optimal threshold and neighbouring window size for every wavelet subband by the Stein's unbiased risk estimate (SURE).

- (f) Interscale orthonormal wavelet thresholding, OWT [16]. It is also a SURE-based wavelet shrinkage approach. It parameterizes the denoising process as a sum of elementary nonlinear processes with unknown weights, and does not hypothesize a statistical model for the original image.
- (g) BiShrink Wavelet-Based Denoising using the Separable Discrete Wavelet Transform, BiS [15]. This method uses a bivariate shrinkage function which models the statistical dependence between a wavelet coefficient and its parent.
- (h) BiShrink Wavelet-Based Denoising using the Dual-Tree Discrete Wavelet Transform, BiS2 [15]. It is analogous to the previous method, but with a different wavelet transform.
- (i) ProbShrink, PS [17]. It estimates the probability that a given wavelet coefficient contains a significant noise-free component, and then the coefficient is multiplied by that probability.

We have selected three benchmark images from the University of Waterloo repertoire [50], which are shown in Fig. 3. One of

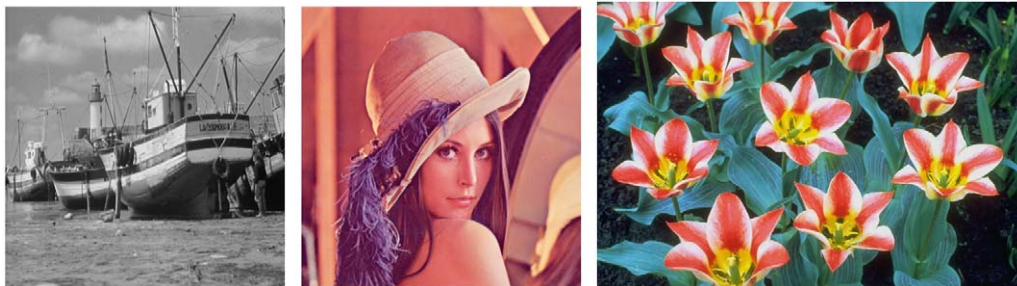


Fig. 3. Original images. From left to right: boat, Lena and tulips.

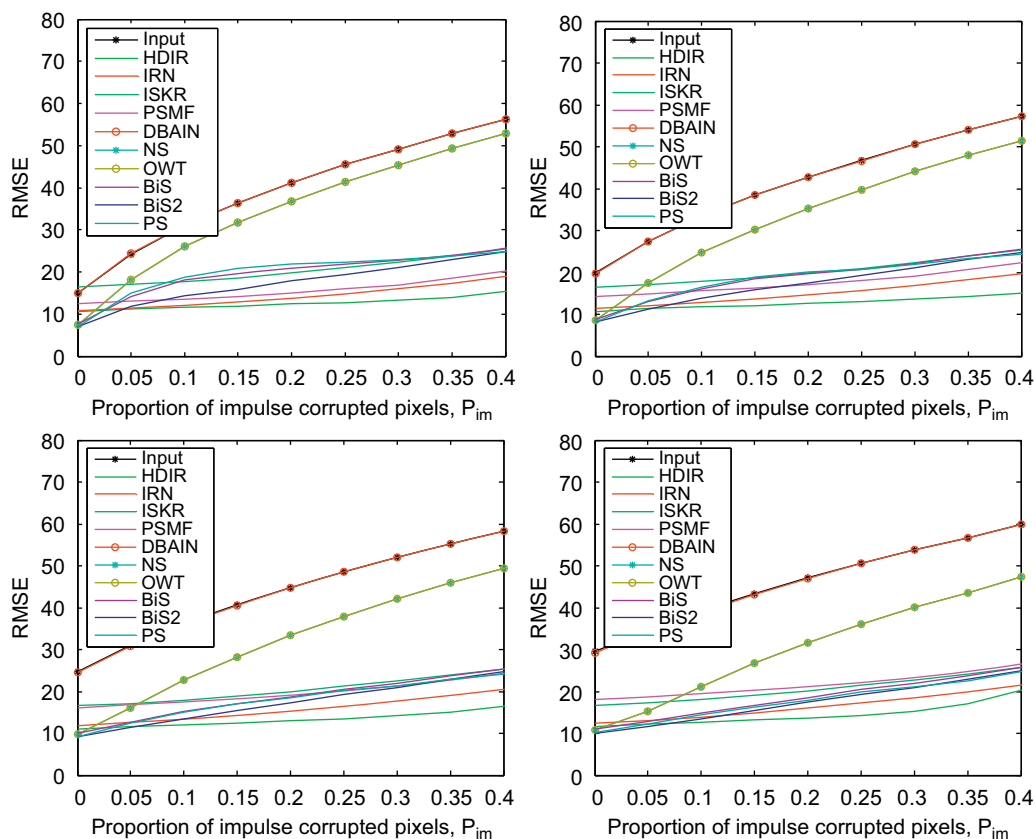


Fig. 4. Results for the boat image. From left to right and from top to bottom: RMSE with Gaussian noise standard deviation $\sigma=15, 20, 25, 30$.

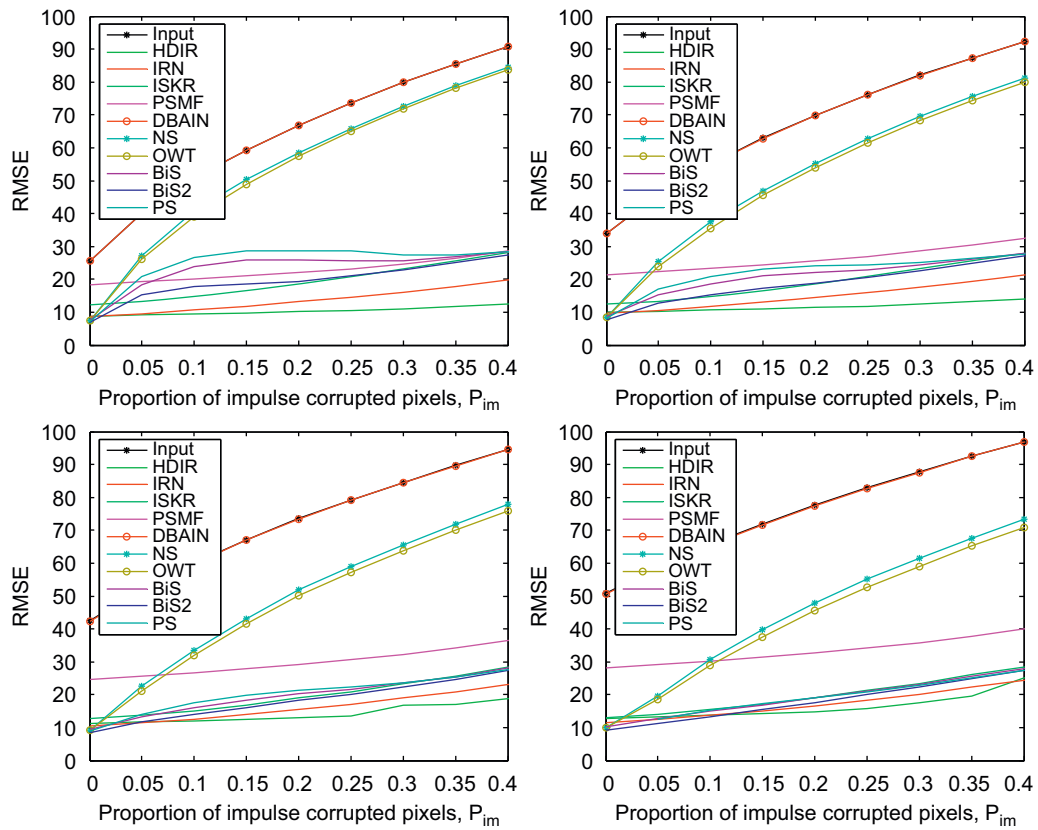


Fig. 5. Results for the Lena image. From left to right and from top to bottom: RMSE with Gaussian noise standard deviation $\sigma=15, 20, 25, 30$.

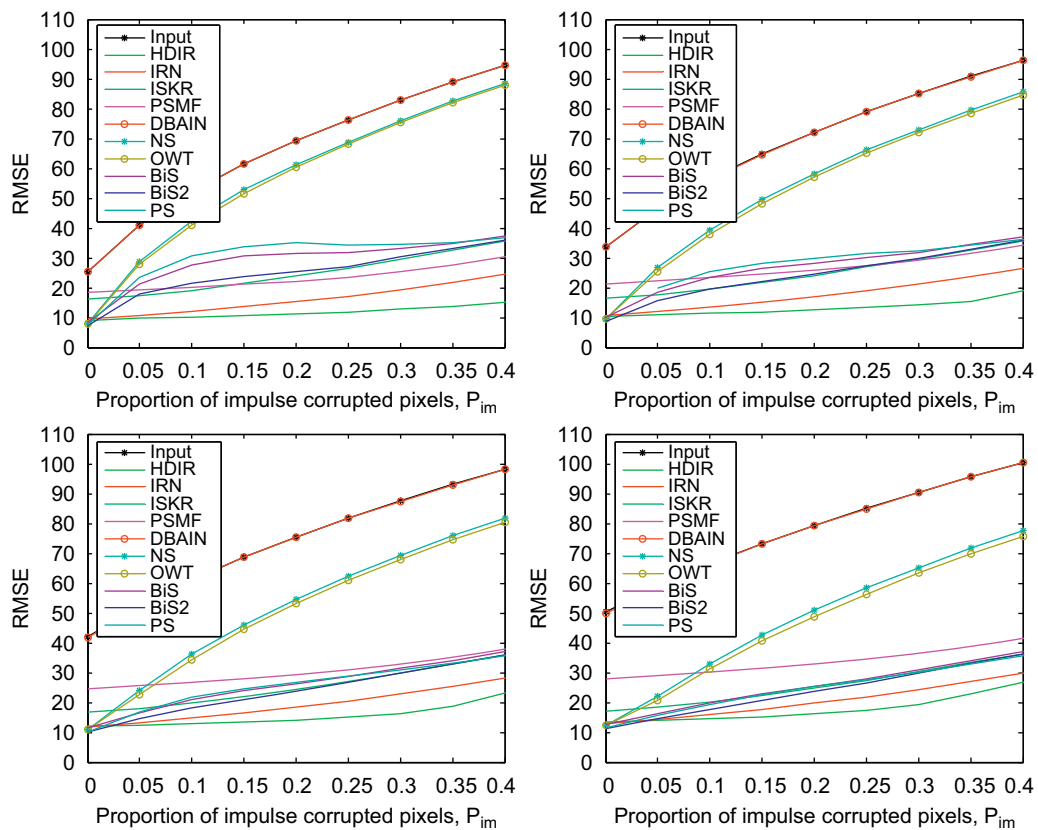


Fig. 6. Results for the tulips image. From left to right and from top to bottom: RMSE with Gaussian noise standard deviation $\sigma=15, 20, 25, 30$.

them (boat) is a 512×512 grayscale image, and the other two (Lena and tulips) are colour images of sizes 512×512 and 768×512 , respectively. For the colour images all the computations have been performed on the YCbCr colour space, since it is widely used in practise to transmit colour information, as done in JPEG image files [51] and MPEG video files [52]. In all cases (grayscale and colour) the pixel values lie in the range $[0,255]$, i.e. we have $v=255$.

We have chosen as a quantitative restoration quality measure the root mean squared error [13] (RMSE, lower is better):

$$RMSE = \sqrt{\frac{1}{AB} \sum_{i=1}^{AB} (y_i - \hat{y}_i)^2} \quad (53)$$

For colour images, we average the squared error over all the spectral channels (see for example [17,53]). The RMSE has the same dimensions as the pixel values, which in our case implies that $RMSE \in [0,255]$. Furthermore, it ranks any compared methods in the same way as the mean squared error (MSE, lower is better)

and the peak signal-to-noise ratio (PSNR, higher is better):

$$MSE = \frac{1}{AB} \sum_{i=1}^{AB} (y_i - \hat{y}_i)^2 \quad (54)$$

$$PSNR = 10 \log_{10} \frac{255^2}{MSE} \quad (55)$$

We have tested four different Gaussian noise levels: $\sigma=15, 20, 25, 30$. For each value of the Gaussian noise standard deviation σ we have tested all the possible proportions of impulse noise P_{im} between 0 and 0.4, in 0.05 increments. The results are shown in Figs. 4–6 for boat, Lena and tulips, respectively. Our HDIR approach yields the best results, followed by IRN, ISKR and some wavelet shrinkage methods (BiS, BiS2 and PS). In particular, IRN minimizes a functional which has a term to represent the fidelity to the input (noisy) image. This term takes into account all input pixels, no matter how noisy they are. Hence, highly noisy pixels can affect the restoration very negatively. On the other hand, HDIR discards the pixels with high noise levels by assigning them a very small weight, i.e. $E[\delta_i|y_i] \approx 0$ in Eq. (30). This means that the restoration is not affected by those pixels.



Fig. 7. Detail of the boat image. From left to right and from top to bottom (RMSE in parentheses): original image, corrupted image (57.1502), HDIR (15.1001), IRN (19.7459), ISKR (25.4810), PSMF (22.3838), DBAIN (57.2077), NS (51.2987), OWT (51.2987), BiS (25.6406), BiS2 (24.7716), PS (24.5345).

The performance of the PSMF method is very dependent on the features of the input image, and degrades quickly as we increase the Gaussian noise level, which is due to its orientation to impulse noise removal. On the other hand, the DBAIN method does not reduce the error (the *RMSE* remains nearly the same as in the input image), as it is designed to remove only extreme impulses.

In order to assess the qualitative performance of the compared approaches, we have selected a set of input conditions for each image which exhibits clearly noticeable differences. In particular we have: for the boat image (Fig. 7), $\sigma=20$ and $P_{im}=0.4$; for the Lena image (Fig. 8), $\sigma=15$ and $P_{im}=0.4$; and for the tulips image (Fig. 9), $\sigma=25$ and $P_{im}=0.3$. In these three figures it can be seen that HDIR preserves many details of the original image, while IRN yields more pixellated reconstructions. On the other hand, ISKR achieves its low *RMSE* error at the expense of excessive smoothing, while the BiS, BiS2 and PS wavelet approaches show very noticeable artifacts in the regions with more impulsive error. The other methods are too conservative and leave a significant portion of the noise. These results agree with the quantitative results stated above, as the comparative quality of each method with respect to the others is the same.

Finally we depict the results of the noise model learning procedure for the Lena image in Fig. 10. As seen, our algorithm is able to obtain a very close approximation to the amount of pixels with a certain error e_i (left) and to the probability that a certain pixel is impulse corrupted given its error e_i (right). This means that our training scheme is able to recover the parameters of the noise process which are relevant for the restoration from the input data, without relying on user tuned parameters.

6. Conclusions

A new image restoration method has been presented, which is aimed to process images corrupted by both Gaussian and uniform impulse noise. A probabilistic theoretical framework has been developed, including a mixture model for the noise to be trained with the input image data. This allows estimating the probability that a certain pixel is impulse corrupted, which in turn provides a principled way to assign weights to the input pixels in the construction of the output image by kernel regression.



Fig. 8. Detail of the Lena image. From left to right and from top to bottom (*RMSE* in parentheses): original image, corrupted image (90.5691), HDIR (12.3469), IRN (19.5763), ISKR (27.8849), PSMF (28.5702), DBAIN (90.5600), NS (84.2008), OWT (83.5118), BiS (28.2878), BiS2 (27.2574), PS (28.4346).



Fig. 9. Detail of the tulips image. From left to right and from top to bottom (*RMSE* in parentheses): original image, corrupted image (87.5721), HDIR (17.1924), IRN (22.8797), ISKR (29.9664), PSMF (32.8684), DBAIN (87.5431), NS (69.4079), OWT (67.9032), BiS (31.4287), BiS2 (29.8594), PS (31.0968).

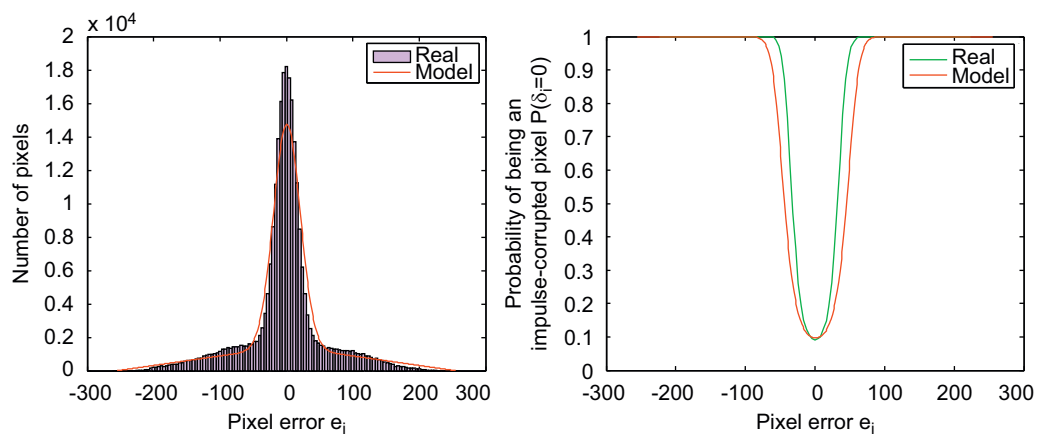


Fig. 10. Noise model learning for the Lena image with $\sigma=15$ and $P_{im}=0.4$. Pixel count versus pixel error e_i (left) and probability of being an impulse corrupted pixel versus e_i (right).

The performance of our proposal has been compared to those of a selection of state of the art techniques that span a wide range of approaches to the image restoration problem. The quantitative and qualitative results with benchmark images show that our method removes the noise while it preserves the fine details of the original image.

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Appendix A. Distribution of the error for impulse corrupted pixels

Let $z(\mathbf{x}_i), y_i$ be the original and observed pixel values at position \mathbf{x}_i , which is affected by impulse noise. Then both random variables are uniform in the interval $[0, v]$. The probability density function of y_i is:

$$f(a) = \begin{cases} \frac{1}{v} & \text{iff } 0 \leq a \leq v \\ 0 & \text{otherwise} \end{cases} \quad (56)$$

The probability density function of $-z(\mathbf{x}_i)$ is:

$$g(a) = \begin{cases} \frac{1}{v} & \text{iff } -v \leq a \leq 0 \\ 0 & \text{otherwise} \end{cases} \quad (57)$$

Now we compute the probability density function of their sum e_i as the convolution of f and g :

$$p(e_i) = (f * g)(e_i) = \int_{-\infty}^{\infty} f(e_i - y_i)g(y_i)dy_i \quad (58)$$

From the definition of g we get:

$$p(e_i) = \frac{1}{v} \int_{-v}^0 f(e_i - y_i) dy_i \quad (59)$$

Now the integrand is zero unless $0 \leq e_i - y_i \leq v$, and then it is $1/v$. Hence,

$$0 \leq e_i \leq v \Rightarrow p(e_i) = \frac{1}{v} \int_{e_i-v}^0 \frac{1}{v} dy_i = \frac{v-e_i}{v^2} \quad (60)$$

On the other hand,

$$-v \leq e_i \leq 0 \Rightarrow p(e_i) = \frac{1}{v} \int_{-v}^{e_i} \frac{1}{v} dy_i = \frac{v+e_i}{v^2} \quad (61)$$

And finally,

$$e_i \notin [-v, v] \Rightarrow p(e_i) = 0 \quad (62)$$

So we arrive at the triangular density with zero mode, minimum value $-v$ and maximum value v :

$$p(e_i) = \text{Tri}_v(e_i) = \begin{cases} \frac{v+e_i}{v^2} & \text{if } -v \leq e_i \leq 0 \\ \frac{v-e_i}{v^2} & \text{if } 0 \leq e_i \leq v \\ 0 & \text{otherwise} \end{cases} \quad (63)$$

Appendix B. Model learning with the Expectation-Maximization algorithm

We must maximize the data likelihood,

$$L(\sigma, P_{lm}) = \sum_i \log p(e_i | \sigma, P_{lm}) \quad (64)$$

where

$$p(e_i | \sigma, P_{lm}) = (1 - P_{lm})N_{\sigma}(e_i) + P_{lm}\text{Tri}_v(e_i) \quad (65)$$

The parameter approximations at time step t are grouped in a parameter vector $\theta(t)$:

$$\theta(t) = (\sigma(t), P_{lm}(t)) \quad (66)$$

For the E step, first we compute the posterior probability of the mixture components for having generated the sample e_i :

$$P(lm(\mathbf{x}_i) | e_i, \theta(t)) = R_{lm,i,t} = \frac{P_{lm}(t)\text{Tri}_v(e_i)}{p(e_i | \theta(t))} \quad (67)$$

$$P(G(\mathbf{x}_i) | e_i, \theta(t)) = R_{G,i,t} = \frac{(1 - P_{lm}(t))N_{\sigma(t)}(e_i)}{p(e_i | \theta(t))} \quad (68)$$

Then we obtain the expectation of the likelihood:

$$E[L(\theta)] = \sum_i (R_{lm,i,t}(\log P_{lm}(t) + \log \text{Tri}_v(e_i)) + R_{G,i,t}(\log(1 - P_{lm}(t)) + \log N_{\sigma}(e_i))) \quad (69)$$

For the M step, both parameters may be optimized independently, since they appear in (69) in separate linear terms.

First we consider P_{lm} :

$$P_{lm}(t+1) = \arg \max_{\alpha} \{ (\sum_i R_{lm,i,t}) \log \alpha + (\sum_i R_{G,i,t}) \log(1 - \alpha) \} \quad (70)$$

This is analogous to the maximum likelihood estimator for the binomial distribution, so we have:

$$P_{lm}(t+1) = \frac{\sum_i R_{lm,i,t}}{\sum_i R_{lm,i,t} + \sum_i R_{G,i,t}} = \frac{1}{AB} \sum_i R_{lm,i,t} \quad (71)$$

On the other hand,

$$\sigma(t+1) = \arg \max_{\alpha} \sum_i R_{G,i,t} \log N_{\mu,\alpha}(e_i) \quad (72)$$

This is analogous to the weighted maximum likelihood estimator of a normal distribution, so we get:

$$\sigma(t+1) = \sqrt{\frac{\sum_i R_{G,i,t} e_i^2}{\sum_i R_{G,i,t}}} = \sqrt{\frac{\sum_i R_{G,i,t} e_i^2}{AB(1 - P_{lm}(t+1))}} \quad (73)$$

Being an EM algorithm, it is guaranteed that these equations converge to a maximum of the likelihood L (see for example [54]).

Appendix C. Noise variance

Here we are interested in the noise variance, $\text{var}[e_i]$. Since the noise has zero mean, we have:

$$\text{var}[e_i] = E[e_i^2] - E[e_i]^2 = E[e_i^2] \quad (74)$$

Since every pixel i is either Gaussian or impulse corrupted,

$$\text{var}[e_i] = P(G(\mathbf{x}_i))E[e_i^2 | G(\mathbf{x}_i)] + P(lm(\mathbf{x}_i))E[e_i^2 | lm(\mathbf{x}_i)] \quad (75)$$

We rewrite in terms of $P_{lm} = P(lm(\mathbf{x}_i))$:

$$\text{var}[e_i] = (1 - P_{lm})E[e_i^2 | G(\mathbf{x}_i)] + P_{lm}E[e_i^2 | lm(\mathbf{x}_i)] \quad (76)$$

Finally, we use the variance of the triangular distribution (see [55,56]) to yield:

$$\text{var}[e_i] = (1 - P_{lm})\sigma^2 + P_{lm} \frac{v^2}{6} \quad (77)$$

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