

Astrometry forecast

In progress...

November 13, 2024

1 Forecast

From the equation (2.2) at $t = t'$ we have

$$\langle \delta n^i(\mathbf{n}, t) \delta n^j(\mathbf{q}, t') \rangle = \frac{3H_0^2}{32\pi^3} \int df \frac{\Omega_{\text{GW}}}{f^3} \left[H_{ij}^{(0)}(\mathbf{n}, \mathbf{q}) + \beta (4 - n_\Omega) H_{ij}^{(1)}(\mathbf{n}, \mathbf{q}, \mathbf{v}) \right], \quad (1)$$

We identify the monopole and dipole coefficients:

$$M = \frac{3H_0^2}{32\pi^3} \int df \frac{\Omega_{\text{GW}}}{f^3}, \quad (2)$$

$$D_\beta = \frac{3H_0^2}{32\pi^3} \int df \frac{\Omega_{\text{GW}}}{f^3} \beta (4 - n_\Omega). \quad (3)$$

Besides this, we know

$$\Omega_{\text{GW}}(f) = \frac{4\pi^2 f^3}{3H_0^2} I(f) = \frac{4\pi^2 f^3}{3H_0^2} I_0 \left(\frac{f}{f_\star} \right)^{n_I}, \quad (4)$$

Then

$$M = \frac{I_0}{8\pi} \int df \left(\frac{f}{f_\star} \right)^{n_I}, \quad (5)$$

$$D_\beta = \frac{I_0}{8\pi} \int df \left(\frac{f}{f_\star} \right)^{n_I} \beta (4 - n_\Omega). \quad (6)$$

On the other side, from equation (4.1)

$$\langle \delta n_a^i \delta n_b^j \rangle \equiv \mathbf{C} = \Sigma_n p_n H_{ab,n}^{ij} + N_{ab}^{ij} = \sum_n p_n \mathbf{H}_n + \mathbf{N}. \quad (7)$$

with $p_0 = M$ and $p_1 = D_\beta$. From the equations (5) and (6) we see

$$\Delta p_0 \simeq \frac{\Delta I_0}{8\pi} \int df \left(\frac{f}{f_\star} \right)^{n_I}, \quad (8)$$

$$\Delta p_1 \simeq \frac{\Delta I_0}{8\pi} \int df \left(\frac{f}{f_\star} \right)^{n_I} \beta (4 - n_\Omega) \quad (9)$$

There is a relation between I_0 and $\Omega_{\text{GW},0}$ given by

$$\Omega_{\text{GW},0} = \frac{2\pi^2}{3H_0^2} f_\star^2 A^2 = \frac{4\pi^2}{3H_0^2} f_\star^3 I_0 \quad (10)$$

Section 4.1 from the overleaf has a result for $\Delta p_0 = \sqrt{[\mathcal{F}^{-1}]_{00}}$ and $\Delta p_1 = \sqrt{[\mathcal{F}^{-1}]_{11}}$, where \mathcal{F} is given by the equation (4.13) in that section, times the total number of observations n_{obs} . With the relations above, the errors on the parameters p_n can be expressed as errors in $\Delta\Omega_{\text{GW},0}$ as follows,

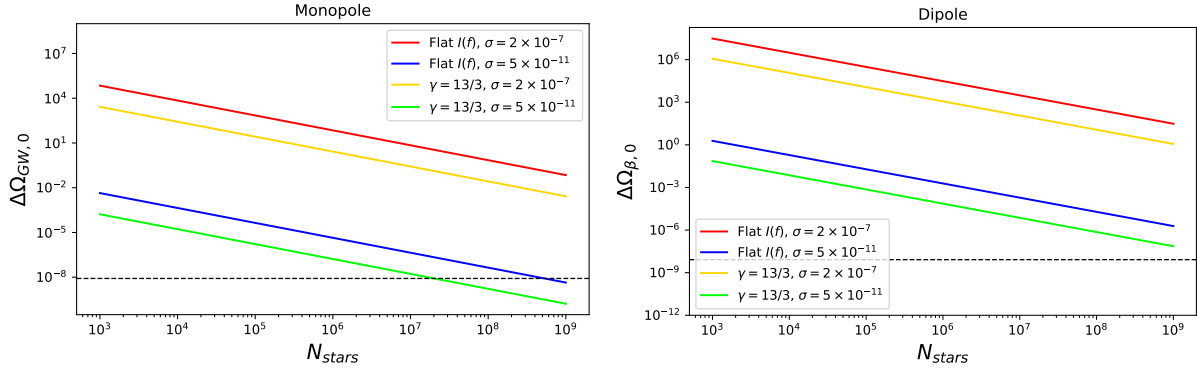


Figure 1: *Forecast for the estimators p_0 (left panel) and p_1 (right panel) in terms of the spectral energy density parameter $\Omega_{\text{GW},0}$. In this case $\Omega_{\beta,0}$ is just a way to make it evident that $\Omega_{\text{GW},0}$ is calculated this time with the dipole estimator.*

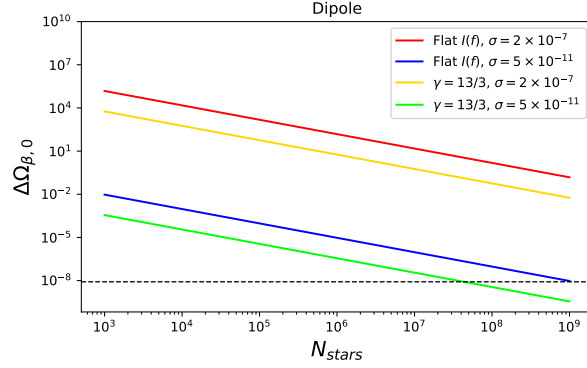


Figure 2: *Forecast for the estimator and p_1 in terms of the spectral energy density parameter $\Omega_{\text{GW},0}$. In this case $\Omega_{\beta,0} = \beta(4 - n_\Omega)\Omega_{\text{GW},0}$.*

$$\Delta\Omega_{\text{GW},0} \simeq \frac{4\pi^2}{3H_0^2} f_\star^3 \Delta I_0 = \frac{4\pi^2 f_\star^3}{3H_0^2} \left(\frac{8\pi \Delta p_0}{\int df \left(\frac{f}{f_\star} \right)^{n_I}} \right), \quad (11)$$

$$\Delta\Omega_{\text{GW},0} \simeq \frac{4\pi^2}{3H_0^2} f_\star^3 \Delta I_0 = \frac{4\pi^2 f_\star^3}{3H_0^2} \left(\frac{8\pi \Delta p_1}{\beta(4 - n_\Omega) \int df \left(\frac{f}{f_\star} \right)^{n_I}} \right) \quad (12)$$

I can also plot equation (12) times the factor $\beta(4 - n_\Omega)$, the result for this case is in Figure 2. Nevertheless, in this way there is no much difference between the Dipole and the Monopole.

References