

Basic quantum field theory

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ABSTRACT: Lecture notes for the course (Version: 2024/10/4)

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0 About this course

- Self-introduction: Kazuya Yonekura (high energy theory)
- Google classroom (zu7iym4)
- Evaluation: by exam at the end of the semester 2025/1/28 Tuesday, (January 28th)
See the calendar of the graduate school of science.
- No class on 2024/10/25 (October 25th, University Festival),
2024/11/1 (November 1st)
- Assumed knowledge :
analytical mechanics, quantum mechanics (QM), electromagnetism (EM), special relativity

This course: Quantum Field Theory (QFT)

Abbreviation:

EM = electromagnetism

QM = quantum mechanics

QFT = quantum field theory

What is QFT?

- High energy physics: the fundamental framework for the laws of physics, including the standard model of particle physics
- Condensed matter physics: effective description of various many body physics phenomena (e.g. superconductivity)

Examples of fields

- All elementary particles are described by some fields.
e.g. EM field, EM waves \rightarrow photon
- Collective motions in many body systems
e.g. sound waves \rightarrow phonon

Remark

The language of relativistic QFT is used.

Natural units:

$$c : \text{speed of light} \tag{0.1}$$

$$\hbar : \text{Planck constant divided by } 2\pi \tag{0.2}$$

$$c = 1, \quad \hbar = 1 \tag{0.3}$$

I will use natural units later.

For cond-mat., $c \rightarrow v$ (an appropriate speed of some field)

1 Non-relativistic QM

1.1 Notation, convention

$$t : \text{time coordinate} \quad (1.1)$$

$$\vec{x} = (x^i) \quad (i = 1, 2, 3) : \text{space coordinates} \quad (1.2)$$

$$\vec{p} = (p^i) : \text{momentum} \quad (1.3)$$

$$(1.4)$$

A dot : time derivative

$$\dot{\vec{x}} = \frac{d\vec{x}}{dt} \quad \ddot{\vec{x}} = \frac{d^2\vec{x}}{dt^2} \quad (1.5)$$

Einstein summation notation

$$A_i B_i = \sum_{i=1,2,3} A_i B_i = \vec{A} \cdot \vec{B} \quad (1.6)$$

Upper and lower indices mean the same thing: $A_i = A^i$.

Partial derivative:

$$\partial_i = \frac{\partial}{\partial x^i}, \quad \partial_t = \frac{\partial}{\partial t} \quad (1.7)$$

Potentials for EM fields

$$\phi : \text{electric potential} \quad \vec{A} = (A_i) : \text{vector potential} \quad (1.8)$$

EM fields

$$\vec{E} = -\vec{\partial}\phi - \frac{\partial\vec{A}}{\partial t} = -\partial_i\phi - \partial_t A_i \quad (1.9)$$

$$\vec{B} = \vec{\partial} \times \vec{A} = \frac{1}{2}\epsilon_{ijk}(\partial_i A_j - \partial_j A_i) \quad (1.10)$$

ϵ_{ijk} : totally antisymmetric tensor,

$$\epsilon_{ikj} = -\epsilon_{ijk}, \quad \epsilon_{kji} = -\epsilon_{ijk}, \quad \epsilon_{jik} = -\epsilon_{ijk}, \quad \epsilon_{123} = 1. \quad (1.11)$$

Explicitly

$$\epsilon_{123} = \epsilon_{231} = \epsilon_{312} = 1, \quad \epsilon_{213} = \epsilon_{132} = \epsilon_{321} = -1, \quad \text{others} = 0 \quad (1.12)$$

The exterior product

$$(\vec{A} \times \vec{B})_i = \epsilon_{ijk} A_j B_k. \quad (\text{Einstein's summation notation for } j, k) \quad (1.13)$$

$$= \sum_{j,k} \epsilon_{ijk} A_j B_k \quad (1.14)$$

1.2 Analytical mechanics

The action for a particle interacting with EM field:

$$S = \int dt \left(\frac{1}{2} m \dot{\vec{x}}^2 - e\phi \right) + e \int A_i \frac{dx^i}{dt} dt \quad (1.15)$$

$$m: \text{mass, } e: \text{charge} \quad (1.16)$$

Equations of motion (EOM) is obtained by the principle of least action:

Under $\vec{x} \rightarrow \vec{x} + \delta\vec{x}$ ($\delta\vec{x}$: infinitesimal), the action is stationary, $\delta S = 0$.

$$\delta S = \int dt (m \delta \dot{x}^i \dot{x}^i - e \delta x^i \partial_i \phi) + e \int dt (\delta x^j \partial_j A_i \dot{x}^i + A_i \delta \dot{x}^i) \quad (1.17)$$

$$= \int dt (-m \ddot{x}^i - e \partial_i \phi) \delta x^i + e \int dt \delta x^i (\partial_i A_j \dot{x}^j - \partial_j A_i \dot{x}^j - \partial_t A_i) \quad (1.18)$$

where

$$\frac{d}{dt} A_i = \partial_j A_i \dot{x}^j + \partial_t A_i \quad (1.19)$$

is used.

$$\delta S = \int dt (-m \ddot{x}^i - e(\partial_i \phi + \partial_t A_i) + e(\partial_i A_j - \partial_j A_i) \dot{x}^j) \delta x^i = 0 \quad (1.20)$$

EOM

$$m \ddot{x}^i = -e(\partial_i \phi + \partial_t A_i) + e(\partial_i A_j - \partial_j A_i) \dot{x}^j \quad (1.21)$$

$$= eE^i + e\epsilon_{ijk} B^k \dot{x}^j \quad (\partial_i A_j - \partial_j A_i = \epsilon_{ijk} B^k) \quad (1.22)$$

$$= e(\vec{E} + \dot{\vec{x}} \times \vec{B}) : \text{EOM} \quad (1.23)$$

Lagrangian L ($S = \int dt L$)

$$L = \frac{1}{2} m \dot{\vec{x}}^2 - e\phi + e A_i \dot{x}^i \quad (1.24)$$

Canonical momentum

$$p_i = \frac{\partial L}{\partial \dot{x}^i} = m \dot{x}^i + e A^i \quad (1.25)$$

Hamiltonian

$$H = p_i \dot{x}^i - L \quad (1.26)$$

$$= \frac{1}{2} m \dot{\vec{x}}^2 + e\phi \quad (1.27)$$

$$= \frac{1}{2m} (\vec{p} - e\vec{A})^2 + e\phi. \quad (1.28)$$

1.3 Canonical quantization

So far classical mechanics.

Now quantum.

The rule for quantization:

$$x^i, p_j \quad : \text{ operators on the Hilbert space of physical states} \quad (1.29)$$

Impose canonical commutation relations to x^i and p_j ,

$$[x^i, p_j] = i\hbar\delta_j^i \quad (1.30)$$

δ_j^i : the Kronecker delta

$$\delta_j^i = \begin{cases} 1 & (i = j) \\ 0 & (i \neq j) \end{cases} \quad (1.31)$$

More explicitly, p_j is realized on wavefunctions $\Psi(\vec{x})$ by

$$p_j\Psi(\vec{x}) = -i\hbar\partial_j\Psi(\vec{x}) \quad (1.32)$$

Hamiltonian operator

$$H\Psi = \left[\frac{1}{2m}(\vec{p} - e\vec{A})^2 + e\phi \right] \Psi \quad (1.33)$$

$$= \left[\frac{1}{2m}(-i\hbar\partial_i - eA_i)^2 + e\phi \right] \Psi \quad (1.34)$$

(Time dependent) Schrödinger equation

$$i\hbar\partial_t\Psi = H\Psi \quad (1.35)$$

Here, the meaning of $(-i\hbar\partial_i - eA_i)^2$ is to act $(-i\hbar\partial_i - eA_i)$ twice and take some over i .

More explicitly, definite

$$(-i\hbar\partial_i - eA_i)^2\Psi = \sum_i (-i\hbar\partial_i - eA_i)[(-i\hbar\partial_i - eA_i)\Psi]. \quad (1.36)$$

1.4 Covariant derivatives and gauge transformations

Covariant derivative will be an important concept in QFT.

For Ψ in QM,

$$D_i = \partial_i - i\frac{e}{\hbar}A_i, \quad D_t = \partial_t + i\frac{e}{\hbar}\phi \quad (1.37)$$

Schrödinger equation is now

$$i\hbar D_t\Psi = -\frac{\hbar^2}{2m}\vec{D}^2\Psi \quad (1.38)$$

\vec{E}, \vec{B} are invariant under gauge transformations (exercise)

$$A_i \rightarrow A'_i = A_i + \partial_i \alpha, \quad \phi \rightarrow \phi' = \phi - \partial_t \alpha. \quad (1.39)$$

α : an arbitrary function.

The transformation of Ψ is defined as

$$\Psi \rightarrow \Psi' = \exp(i \frac{e}{\hbar} \alpha) \Psi \quad (1.40)$$

Then

$$D'_i \Psi' = \left(\partial_i - i \frac{e}{\hbar} (A_i + \partial_i \alpha) \right) \left(\exp(i \frac{e}{\hbar} \alpha) \Psi \right) \quad (1.41)$$

$$= \exp(i \frac{e}{\hbar} \alpha) \left(\partial_i - i \frac{e}{\hbar} A_i \right) \Psi \quad (1.42)$$

$$= \exp(i \frac{e}{\hbar} \alpha) D_i \Psi \quad (1.43)$$

Similarly

$$(D'_i)^2 \Psi' = \exp(i \frac{e}{\hbar} \alpha) (D_i)^2 \Psi \quad (1.44)$$

$$D'_t \Psi' = \exp(i \frac{e}{\hbar} \alpha) D_t \Psi \quad (1.45)$$

Then

$$i\hbar D_t \Psi = -\frac{\hbar^2}{2m} \vec{D}^2 \Psi \iff i\hbar D'_t \Psi' = -\frac{\hbar^2}{2m} \vec{D}'^2 \Psi' \quad (1.46)$$

Principle: physics is invariant under gauge transformations.

1.5 Harmonic oscillator

Harmonic oscillators will play very important roles in QFT.

In QM (one-dimension),

$$H = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 x^2. \quad (1.47)$$

ω : parameter.

Define creation and annihilation operators a, a^\dagger :

$$a = \sqrt{\frac{m\omega}{2\hbar}} x + i \sqrt{\frac{1}{2\hbar\omega m}} p, \quad (1.48)$$

$$a^\dagger = \sqrt{\frac{m\omega}{2\hbar}} x - i \sqrt{\frac{1}{2\hbar\omega m}} p \quad (1.49)$$

Commutation relations

$$[a, a^\dagger] = \frac{1}{2\hbar}(-i[x, p] + i[p, x]) \quad (1.50)$$

$$= 1 \quad ([x, p] = i\hbar) \quad (1.51)$$

$$x = \sqrt{\frac{\hbar}{2m\omega}}(a + a^\dagger), \quad p = -i\sqrt{\frac{\hbar\omega m}{2}}(a - a^\dagger). \quad (1.52)$$

$$H = -\frac{\hbar\omega}{4}(a - a^\dagger)^2 + \frac{\hbar\omega}{4}(a + a^\dagger)^2 \quad (1.53)$$

$$= \frac{\hbar\omega}{2}(aa^\dagger + a^\dagger a) \quad (1.54)$$

$$= \hbar\omega(a^\dagger a + \frac{1}{2}). \quad (1.55)$$

Define state vectors $|n\rangle$ ($n = 0, 1, 2, \dots$) by

$$|0\rangle : a|0\rangle = 0, \quad (1.56)$$

$$|n+1\rangle = \frac{1}{\sqrt{n+1}}a^\dagger|n\rangle \quad (n = 0, 1, 2, \dots). \quad (1.57)$$

Number operator

$$N = a^\dagger a \quad (1.58)$$

Remark: for any operators A, B, C

$$[AB, C] = ABC - CAB = A(BC - CB) + (AC - CA)B = A[B, C] + [A, C]B \quad (1.59)$$

Then

$$[N, a^\dagger] = [a^\dagger a, a^\dagger] = a^\dagger[a, a^\dagger] + [a^\dagger, a^\dagger]a = a^\dagger \quad (1.60)$$

$$[N, a] = a^\dagger, \quad [N, a] = -a. \quad (1.61)$$

One can show

$$N|n\rangle = n|n\rangle. \quad (1.62)$$

Proof by induction:

1. For $n = 0$, $a^\dagger a|0\rangle = 0$ by definition of $|0\rangle$.

2. If $N|n\rangle = n|n\rangle$, then

$$N|n+1\rangle = \frac{1}{\sqrt{n+1}}Na^\dagger|n\rangle \quad (1.63)$$

$$= \frac{1}{\sqrt{n+1}}([N, a^\dagger] + a^\dagger N)|n\rangle \quad (1.64)$$

$$= \frac{1}{\sqrt{n+1}}(a^\dagger + a^\dagger N)|n\rangle \quad (1.65)$$

$$= \frac{1}{\sqrt{n+1}}a^\dagger(n+1)|n\rangle \quad (1.66)$$

$$= (n+1)|n+1\rangle. \quad (1.67)$$

Proof completed.

Hamiltonian

$$H = \hbar\omega(N + \frac{1}{2}) \quad (1.68)$$

$$[H, a^\dagger] = \hbar\omega a^\dagger, \quad [H, a] = -\hbar\omega a. \quad (1.69)$$

The energy E_n of the state $|n\rangle$:

$$H|n\rangle = E_n|n\rangle \quad (1.70)$$

$$E_n = \hbar\omega(n + \frac{1}{2}). \quad (1.71)$$

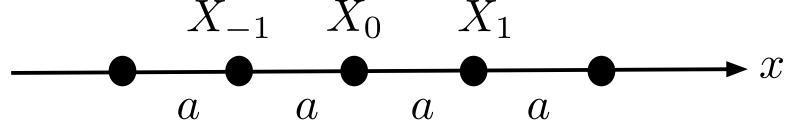
If you are not familiar with QM and analytical mechanics, learn basic things by yourself.

2 A condensed matter example of QFT : phonon

2.1 A lattice model

A simple model: a chain of harmonic oscillators in one-dimension.

$$(X_n, P_n) \quad (n = \cdots, -1, 0, 1, \cdots) \quad (2.1)$$



Canonical commutation relations

$$[X_n, P_m] = i\hbar\delta_{nm} \quad (2.2)$$

$$[X_n, X_m] = 0, \quad [P_n, P_m] = 0 \quad (2.3)$$

Hamiltonian

$$H = \sum_n \left(\frac{P_n^2}{2M} + \frac{1}{2}K(X_{n+1} - X_n - a)^2 \right) \quad (2.4)$$

$$a : \text{lattice spacing}, \quad M : \text{mass of each nucleus}, \quad K : \text{spring constant} \quad (2.5)$$

Classically, the energy is minimized for

$$X_n = na + C \quad (C : \text{constant}) \quad (2.6)$$

We set $C = 0$ for simplicity.

For later convenience, define new variables (ϕ_n, π_n)

$$X_n = na + \sqrt{\frac{a}{M}}\phi_n, \quad P_n = \sqrt{aM}\pi_n \quad (2.7)$$

Then,

$$[\phi_n, \pi_m] = \frac{1}{a}[X_n, P_m] = \frac{i\hbar\delta_{mn}}{a} \quad (2.8)$$

Also

$$\frac{P_n^2}{2M} = \frac{1}{2}a\pi_n^2 \quad (2.9)$$

$$\frac{1}{2}K(X_{n+1} - X_n - a)^2 = \frac{aK}{2M}(\phi_{n+1} - \phi_n)^2 = \frac{a}{2}v^2 \left(\frac{\phi_{n+1} - \phi_n}{a} \right)^2 \quad (2.10)$$

where

$$v^2 = \frac{a^2 K}{M}. \quad (2.11)$$

Then

$$H = a \sum_n \left(\frac{1}{2} \pi_n^2 + \frac{1}{2} v^2 \left(\frac{\phi_{n+1} - \phi_n}{a} \right)^2 \right) \quad (2.12)$$

$$[\phi_n, \pi_m] = \frac{i \hbar \delta_{mn}}{a} \quad (2.13)$$

2.2 Continuum limit

Suppose we are interested in collective motions whose wavelength λ is large enough,

$$\lambda \gg a, \quad \text{or} \quad ka \ll 1, \quad k = \frac{2\pi}{\lambda} : \text{wavenumber}. \quad (2.14)$$

So, we take the limit

$$a \rightarrow 0. \quad (2.15)$$

Define

$$x = na : \text{the positions of nuclei (at the static positions)} \quad (2.16)$$

(ϕ_n, π_n) are functions of x ,

$$\phi_n \rightarrow \phi(x), \quad \pi_n \rightarrow \pi(x). \quad (2.17)$$

From the definition of derivatives and integrals,

$$\frac{\phi_{n+1} - \phi_n}{a} \rightarrow \frac{\partial \phi}{\partial x}(x) = \partial_x \phi(x), \quad (2.18)$$

$$a \sum_n \rightarrow \int dx. \quad (2.19)$$

Hamiltonian

$$H = a \sum_n \left(\frac{1}{2} \pi_n^2 + \frac{1}{2} v^2 \left(\frac{\phi_{n+1} - \phi_n}{a} \right)^2 \right) \quad (2.20)$$

$$\rightarrow \int dx \left(\frac{1}{2} \pi(x)^2 + \frac{1}{2} v^2 (\partial_x \phi(x))^2 \right) \quad (2.21)$$

Define Dirac delta function:

$$\delta(x) = \begin{cases} \infty & x = 0 \\ 0 & x \neq 0 \end{cases} \quad (2.22)$$

such that

$$\int_{-\infty}^{\infty} dx \delta(x) = 1 \quad (2.23)$$

Then

$$\frac{\delta_{nm}}{a} \rightarrow \delta(x-y), \quad x = na, \quad y = ma \quad (2.24)$$

because

$$\frac{\delta_{nm}}{a} = \begin{cases} \frac{1}{a} \rightarrow \infty & n - m = 0 \\ 0 & n - m \neq 0 \end{cases} \quad (2.25)$$

and

$$a \sum_n \frac{\delta_{nm}}{a} = 1. \quad (2.26)$$

Then

$$[\phi_n, \pi_m] = \frac{i\hbar\delta_{mn}}{a} \quad (2.27)$$

$$\rightarrow [\phi(x), \pi(y)] = i\hbar\delta(x-y). \quad (2.28)$$

Summary:

$$H = \int dx \left(\frac{1}{2}\pi^2 + \frac{1}{2}v^2 (\partial_x \phi)^2 \right) \quad (2.29)$$

$$[\phi(x), \pi(y)] = i\hbar\delta(x-y). \quad (2.30)$$

An example of QFT, called phonon.

2.3 Lagrangian formulation

Let's return to classical theory.

The original Hamiltonian

$$H = \sum_n \left(\frac{P_n^2}{2M} + \frac{1}{2}K(X_{n+1} - X_n - a)^2 \right) \quad (2.31)$$

Lagrangian

$$L = \sum_n \left(\frac{M}{2}(\dot{X}_n)^2 - \frac{1}{2}K(X_{n+1} - X_n - a)^2 \right) \quad (2.32)$$

$$X_n = na + \sqrt{\frac{a}{M}}\phi_n \quad (2.33)$$

$$L = a \sum_n \left(\frac{1}{2}(\dot{\phi}_n)^2 - \frac{1}{2}v^2 \left(\frac{\phi_{n+1} - \phi_n}{a} \right)^2 \right) \quad (2.34)$$

$$\rightarrow \int dx \left(\frac{1}{2}(\partial_t \phi)^2 - \frac{1}{2}v^2 (\partial_x \phi)^2 \right) \quad (2.35)$$

$$= \int dx \mathcal{L} \quad (2.36)$$

Here

$$\mathcal{L} = \frac{1}{2}(\partial_t \phi)^2 - \frac{1}{2}v^2 (\partial_x \phi)^2 \quad : \text{Lagrangian density} \quad (2.37)$$

The canonical momentum

$$\pi_n = \frac{P_n}{\sqrt{aM}} = \sqrt{\frac{M}{a}} \dot{X}_n = \dot{\phi}_n. \quad (2.38)$$

hence

$$\pi(x) = \partial_t \phi(x). \quad (2.39)$$

This can be written as

$$\pi = \frac{\partial \mathcal{L}}{\partial(\partial_t \phi)} \quad (2.40)$$

The action

$$S = \int dt dx \mathcal{L} \quad (2.41)$$

2.4 Classical dynamics

The principles of least action:

$$S = \int dt dx \left(\frac{1}{2}(\partial_t \phi)^2 - \frac{1}{2}v^2 (\partial_x \phi)^2 \right) \quad (2.42)$$

$$\phi \rightarrow \phi + \delta \phi \quad (2.43)$$

$$0 = \delta S = \int dt dx (\partial_t \phi \partial_t \delta \phi - v^2 \partial_x \phi \partial_x \delta \phi) \quad (2.44)$$

$$= \int dt dx (-\partial_t^2 \phi + v^2 \partial_x^2 \phi) \delta \phi \quad (2.45)$$

$$(-\partial_t^2 + v^2 \partial_x^2) \phi = 0 \quad (2.46)$$

If we consider

$$\phi = \cos(kx - \omega t), \quad (2.47)$$

$$\omega : \text{angular frequency}, \quad k : \text{wavenumber} \quad (2.48)$$

then

$$\omega^2 = v^2 k^2 \implies \phi = \cos k(x \mp vt) \quad (2.49)$$

$$v : \text{the speed of sound} \quad (2.50)$$

3 Free scalar field theory

Aim: systematic quantization of a scalar field.

From now on,

- Consider relativistic QFT. (Sometimes good approximation for condensed matter systems if the speed of light c is replaced by an appropriate velocity v .)
- Use natural units

$$\hbar = 1, \quad c = 1. \quad (3.1)$$

They can be recovered by dimensional analysis.

- Consider three spatial dimensions. (The previous example was one dimension. It is straightforward to generalize to other dimensions.)

3.1 Special relativity

Coordinates of spacetime

$$x^\mu = (ct, x^i) = (ct, x^1, x^2, x^3) = (x^0, \vec{x}) \quad (3.2)$$

In natural units $c = 1$, $x^0 = t$.

space indices : Latin letters $i, j, \dots = 1, 2, 3$,

spacetime indices : Greek letters $\mu, \nu, \dots = 0, 1, 2, 3$.

Spacetime metric tensor

$$g_{\mu\nu} = \begin{pmatrix} - & & & \\ & + & & \\ & & + & \\ & & & + \end{pmatrix} \quad (3.3)$$

$$ds^2 = -d\tau^2 = g_{\mu\nu} dx^\mu dx^\nu = -(dx^0)^2 + (d\vec{x})^2. \quad (3.4)$$

It is the distance between two spacetime points x^μ and $x^\mu + dx^\mu$:

- When $ds^2 > 0$, spatial distance ds
- When $ds^2 < 0$, proper time difference $d\tau$

4-vector

$$A^\mu, \quad A_\mu = g_{\mu\nu} A^\nu \quad (3.5)$$

Einstein summation notation

$$A^\mu B_\mu = \sum_{\mu=0}^3 A^\mu B_\mu \quad (3.6)$$

$$A^2 = A^\mu A_\mu = -(A^0)^2 + (\vec{A})^2 \quad (3.7)$$

4-momentum of a particle

$$p^\mu = m \frac{dx^\mu}{d\tau} = m \left(\frac{1}{\sqrt{1 - \vec{v}^2}}, \frac{\vec{v}}{\sqrt{1 - \vec{v}^2}} \right) \quad (3.8)$$

$$\vec{v} = \frac{dx^i}{dt} \quad (3.9)$$

$$p^2 = p^\mu p_\mu = -(p^0)^2 + \vec{p}^2 = -m^2 \quad (3.10)$$

$$\implies p^0 = \sqrt{\vec{p}^2 + m^2} := E_p \quad (3.11)$$

Partial derivative

$$\partial_\mu = \frac{\partial}{\partial x^\mu} \quad (3.12)$$

EM potential

$$A^\mu = (+\phi, \vec{A}), \quad A_\mu = (-\phi, \vec{A}). \quad (3.13)$$

EM tensor

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad (3.14)$$

$$= \begin{pmatrix} 0 & -E^1 & -E^2 & -E^3 \\ E^1 & 0 & B^3 & -B^2 \\ E^2 & -B^3 & 0 & B^1 \\ E^3 & B^2 & -B^1 & 0 \end{pmatrix} \quad (3.15)$$

$$F_{i0} = -F^{i0} = E_i \quad (3.16)$$

$$F_{ij} = \epsilon_{ijk} B^k \quad (3.17)$$

If you are not familiar with special relativity, learn basic things by yourself.

3.2 Real scalar field

Real scalar field ϕ

Lagrangian

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 \quad (3.18)$$

$$S = \int d^4x \mathcal{L} \quad (3.19)$$

Equations of motion: $\phi \rightarrow \phi + \delta\phi$,

$$0 = \delta S = \int d^4x \delta\phi (\partial^2 \phi - m^2). \quad (3.20)$$

$$(\partial^2 - m^2)\phi = 0 : \text{Klein-Gordon eq.} \quad (3.21)$$

If we set

$$\phi = \cos(p \cdot x), \quad p = (p^0, \vec{p}) : \text{constant 4-vector} \quad (3.22)$$

then

$$(p^0)^2 - \vec{p}^2 = m^2 \quad (3.23)$$

$$\implies p^0 = \pm E_p, \quad E_p = \sqrt{\vec{p}^2 + m^2}. \quad (3.24)$$

\vec{p} is later interpreted as momentum of a particle.

3.3 Quantization

Lagrangian

$$\mathcal{L} = \frac{1}{2} \left((\dot{\phi})^2 - (\vec{\partial}\phi)^2 - m^2 \phi^2 \right), \quad \dot{\phi} = \partial_t \phi. \quad (3.25)$$

The canonical momentum

$$\pi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \dot{\phi} \quad (3.26)$$

Hamiltonian

$$H = \int d^3x (\pi \dot{\phi} - \mathcal{L}) \quad (3.27)$$

$$= \int d^3x \frac{1}{2} \left(\pi^2 + (\vec{\partial}\phi)^2 + m^2 \phi^2 \right) \quad (3.28)$$

Equal time commutation relations

$$[\phi(t, \vec{x}), \pi(t, \vec{y})] = i\delta^3(\vec{x} - \vec{y}). \quad (3.29)$$

For technical concreteness, impose periodic boundary condition:

$$x^i \sim x^i + L \quad (i = 1, 2, 3) \quad (3.30)$$

This means

$$\phi(x^1, x^2, x^3) = \phi(x^1 + L, x^2, x^3) = \phi(x^1, x^2 + L, x^3) = \phi(x^1, x^2, x^3 + L) \quad (3.31)$$

Fourier modes

$$\exp(i\vec{p} \cdot \vec{x}) \quad (3.32)$$

$$\vec{p} = \frac{2\pi}{L} \vec{n}, \quad \vec{n} = (n_1, n_2, n_3) ; \text{ integers} \quad (3.33)$$

Fourier mode expansion

$$\phi = \frac{1}{L^{3/2}} \sum_{\vec{p}} Q(\vec{p}) \exp(i\vec{p} \cdot \vec{x}) \quad (3.34)$$

$$\pi = \frac{1}{L^{3/2}} \sum_{\vec{p}} P(\vec{p}) \exp(i\vec{p} \cdot \vec{x}) \quad (3.35)$$

The volume of space

$$\int d^3x = L^3 \quad (3.36)$$

Convenient formula

$$\frac{1}{L^3} \int d^3x \exp(i\vec{p} \cdot \vec{x}) = \begin{cases} 1 & \vec{p} = 0 \\ 0 & \vec{p} = \frac{2\pi}{L} \vec{n} \neq 0 \end{cases} \quad (3.37)$$

Inverse Fourier transformation

$$Q(\vec{p}) = \frac{1}{L^{3/2}} \int d^3x \phi(\vec{x}) \exp(-i\vec{p} \cdot \vec{x}) \quad (3.38)$$

$$P(\vec{p}) = \frac{1}{L^{3/2}} \int d^3x \pi(\vec{x}) \exp(-i\vec{p} \cdot \vec{x}) \quad (3.39)$$

For a real scalar, $\phi = \phi^\dagger, \pi = \pi^\dagger$

$$Q(\vec{p})^\dagger = Q(-\vec{p}), \quad P(\vec{p})^\dagger = P(-\vec{p}) \quad (3.40)$$

Commutations relations at a fixed time (say $t = 0$)

$$[Q(\vec{p}), P(\vec{q})] = \frac{1}{L^3} \int d^3x d^3y [\phi(\vec{x}), \pi(\vec{y})] \exp(-i\vec{p} \cdot \vec{x} - i\vec{q} \cdot \vec{y}) \quad (3.41)$$

$$= \frac{1}{L^3} \int d^3x d^3y i \delta^3(\vec{x} - \vec{y}) \exp(-i\vec{p} \cdot \vec{x} - i\vec{q} \cdot \vec{y}) \quad (3.42)$$

$$= \frac{i}{L^3} \int d^3x \exp(-i(\vec{p} + \vec{q}) \cdot \vec{x}) \quad (3.43)$$

$$= i \delta_{\vec{p}, -\vec{q}} \quad (\text{Kronecker delta}) \quad (3.44)$$

Hamiltonian

$$H = \frac{1}{2} \int d^3x \left(\pi^2 + (\vec{\partial}\phi)^2 + m^2 \phi^2 \right) \quad (3.45)$$

$$= \frac{1}{2L^3} \int d^3x \sum_{\vec{p}} \sum_{\vec{q}} (P(\vec{p})P(\vec{q}) + (-\vec{p} \cdot \vec{q} + m^2)Q(\vec{p})Q(\vec{q})) \exp(-i(\vec{p} + \vec{q}) \cdot \vec{x}) \quad (3.46)$$

$$= \frac{1}{2} \sum_{\vec{p}} \sum_{\vec{q}} (P(\vec{p})P(\vec{q}) + (-\vec{p} \cdot \vec{q} + m^2)Q(\vec{p})Q(\vec{q})) \delta_{\vec{p}, -\vec{q}} \quad (3.47)$$

$$= \frac{1}{2} \sum_{\vec{p}} (|P(\vec{p})|^2 + (\vec{p}^2 + m^2)|Q(\vec{p})|^2) \quad (3.48)$$

$$= \frac{1}{2} \sum_{\vec{p}} (|P(\vec{p})|^2 + E_p^2 |Q(\vec{p})|^2) \quad (E_p = \sqrt{\vec{p}^2 + m^2}) \quad (3.49)$$

Define creation and annihilation operators

$$A_{\vec{p}} = \frac{1}{\sqrt{2E_p}}(E_p Q(\vec{p}) + iP(\vec{p})) \quad (3.50)$$

$$A_{\vec{p}}^\dagger = \frac{1}{\sqrt{2E_p}}(E_p Q(-\vec{p}) - iP(-\vec{p})) \quad (3.51)$$

Commutation relations

$$[A_{\vec{p}}, A_{\vec{q}}^\dagger] = \frac{1}{2E_p}(-iE_p[Q(\vec{p}), P(-\vec{q})] + iE_p[P(\vec{p}), Q(-\vec{q})]) = \delta_{\vec{p}, \vec{q}} \quad (3.52)$$

In this way,

$$[A_{\vec{p}}, A_{\vec{q}}^\dagger] = \delta_{\vec{p}, \vec{q}}, \quad [A_{\vec{p}}, A_{\vec{q}}] = [A_{\vec{p}}^\dagger, A_{\vec{q}}^\dagger] = 0 \quad (3.53)$$

Hamiltonian

$$H = \frac{1}{2} \sum_{\vec{p}} E_p (A_{\vec{p}}^\dagger A_{\vec{p}} + A_{\vec{p}} A_{\vec{p}}^\dagger) \quad (3.54)$$

$$= \sum_{\vec{p}} E_p \left(A_{\vec{p}}^\dagger A_{\vec{p}} + \frac{1}{2} \right) \quad (3.55)$$

For each Fourier mode \vec{p} , there is one harmonic oscillator with creation/annihilation operators $A_{\vec{p}}, A_{\vec{p}}^\dagger$.

$$[H, A_{\vec{p}}^\dagger] = E_p A_{\vec{p}}^\dagger, \quad [H, A_{\vec{p}}] = -E_p A_{\vec{p}}. \quad (3.56)$$

Define vacuum energy

$$E_{\text{vac}} = \sum_{\vec{p}} \frac{1}{2} E_p \quad (3.57)$$

This is the zero point energy of harmonic oscillators.

$$H = \sum_{\vec{p}} E_p A_{\vec{p}}^\dagger A_{\vec{p}} + E_{\text{vac}}. \quad (3.58)$$

3.4 Hilbert space: Fock space

The Hilbert space of states is the same as a collection of harmonic oscillators.

Define the vacuum state $|\Omega\rangle$ by

$$A_{\vec{p}}|\Omega\rangle = 0 \quad (\text{For all } \vec{p}) \quad (3.59)$$

The ground state of harmonic oscillators.

$$H|\Omega\rangle = E_{\text{vac}}|\Omega\rangle \quad (3.60)$$

E_{vac} is related to cosmological constant. I do not discuss it, and simply set

$$E_{\text{vac}} \rightarrow 0. \quad (3.61)$$

A state

$$A_{\vec{p}}^\dagger |\Omega\rangle \quad (3.62)$$

Its energy

$$H A_{\vec{p}}^\dagger |\Omega\rangle = ([H, A_{\vec{p}}^\dagger] + A_{\vec{p}}^\dagger H) |\Omega\rangle = E_p A_{\vec{p}}^\dagger |\Omega\rangle \quad (E_{\text{vac}} = 0) \quad (3.63)$$

$$E_{\vec{p}} = \sqrt{\vec{p}^2 + m^2} \quad (3.64)$$

Interpretation: A single particle with 4-momentum $p = (E_{\vec{p}}, \vec{p})$.

A particle appears from a field!

More general states:

$$|n\rangle = \left(\prod_{\vec{p}} (A_{\vec{p}}^\dagger)^{n_{\vec{p}}} \right) |\Omega\rangle \quad (3.65)$$

$$n = (n_{\vec{p}}) : \text{A nonnegative integer } n_{\vec{p}} \text{ for each } \vec{p} \quad (3.66)$$

$$(A_{\vec{p}}^\dagger A_{\vec{p}}) |n\rangle = n_{\vec{p}} |n\rangle. \quad (3.67)$$

$$E = \sum_{\vec{p}} n_{\vec{p}} E_p \quad (3.68)$$

Interpretation :

- $n_{\vec{p}}$ particles with $p = (E_{\vec{p}}, \vec{p})$ for each \vec{p} .
- Identical particles are not distinguished.
- Any number $n_{\vec{p}}$ of particles is possible in a single one-particle state \vec{p} : Bose-Einstein statistics.

This kind of Hilbert space : Fock space

3.5 Infinite volume

L was artificial. We are going to take $L \rightarrow \infty$.

$$\phi = \frac{1}{L^{3/2}} \sum_{\vec{p}} Q(\vec{p}) \exp(i\vec{p} \cdot \vec{x}) \quad (3.69)$$

$$A_{\vec{p}} = \frac{1}{\sqrt{2E_p}} (E_p Q(\vec{p}) + iP(\vec{p})) \quad (3.70)$$

$$A_{\vec{p}}^\dagger = \frac{1}{\sqrt{2E_p}} (E_p Q(-\vec{p}) - iP(-\vec{p})) \quad (3.71)$$

$$(3.72)$$

From them

$$Q(\vec{p}) = \frac{1}{\sqrt{2E_p}}(A_{\vec{p}} + A_{-\vec{p}}^\dagger), \quad P(\vec{p}) = -i\sqrt{\frac{E_p}{2}}(A_{\vec{p}} - A_{-\vec{p}}^\dagger) \quad (3.73)$$

$$\phi = \frac{1}{L^{3/2}} \sum_{\vec{p}} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(A_{\vec{p}} \exp(i\vec{p} \cdot \vec{x}) + A_{\vec{p}}^\dagger \exp(-i\vec{p} \cdot \vec{x}) \right) \quad (3.74)$$

Recall

$$\vec{p} = \frac{2\pi}{L} \vec{n}, \quad \vec{n} = (n_1, n_2, n_3). \quad (3.75)$$

When $L \rightarrow \infty$, \vec{p} continuous,

$$\left(\frac{2\pi}{L} \right)^3 \sum_{\vec{n}} \rightarrow \int d^3p \quad (3.76)$$

Delta function

$$\frac{L^3}{(2\pi)^3} \delta_{\vec{p}, \vec{q}} \rightarrow \delta^3(\vec{p}): \text{Dirac delta function.} \quad (3.77)$$

Check:

$$\int d^3p \delta^3(\vec{p} - \vec{q}) = \left(\frac{2\pi}{L} \right)^3 \sum_{\vec{p}} \frac{L^3}{(2\pi)^3} \delta_{\vec{p}, \vec{q}} = 1. \quad (3.78)$$

Define

$$a_{\vec{p}} = L^{3/2} A_{\vec{p}}. \quad (3.79)$$

$$[a_{\vec{p}}, a_{\vec{q}}^\dagger] = L^3 \delta_{\vec{p}, \vec{q}} \quad (3.80)$$

$$\rightarrow (2\pi)^3 \delta^3(\vec{p} - \vec{q}). \quad (3.81)$$

$$\phi = \frac{1}{L^3} \sum_{\vec{p}} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(a_{\vec{p}} \exp(i\vec{p} \cdot \vec{x}) + a_{\vec{p}}^\dagger \exp(-i\vec{p} \cdot \vec{x}) \right) \quad (3.82)$$

$$\rightarrow \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(a_{\vec{p}} \exp(i\vec{p} \cdot \vec{x}) + a_{\vec{p}}^\dagger \exp(-i\vec{p} \cdot \vec{x}) \right). \quad (3.83)$$

$$H = \sum_{\vec{p}} E_p A_{\vec{p}}^\dagger A_{\vec{p}} = \frac{1}{L^3} \sum_{\vec{p}} E_p a_{\vec{p}}^\dagger a_{\vec{p}} \quad (3.84)$$

$$\rightarrow \int \frac{d^3p}{(2\pi)^3} E_p a_{\vec{p}}^\dagger a_{\vec{p}} \quad (3.85)$$

One particle state

$$|\vec{p}\rangle = \sqrt{2E_{\vec{p}}} a_{\vec{p}}^{\dagger} |\Omega\rangle. \quad (3.86)$$

($\sqrt{2E_{\vec{p}}}$ is just convention, often used in high energy physics.)

Normalization

$$\langle \vec{p} | \vec{q} \rangle = \sqrt{2E_{\vec{p}}} \sqrt{2E_{\vec{q}}} \langle \Omega | a_{\vec{p}} a_{\vec{q}}^{\dagger} | \Omega \rangle \quad (3.87)$$

$$= \sqrt{2E_{\vec{p}}} \sqrt{2E_{\vec{q}}} \langle \Omega | \left((2\pi)^3 \delta^3(\vec{p} - \vec{q}) + a_{\vec{q}}^{\dagger} a_{\vec{p}} \right) | \Omega \rangle \quad (3.88)$$

$$= 2E_{\vec{p}} (2\pi)^3 \delta^3(\vec{p} - \vec{q}). \quad (3.89)$$

4 More on a real scalar field

- Heisenberg picture
- Non-relativistic limit
- Second quantization

4.1 Heisenberg picture in QM

One particle quantum mechanics

$$[x, p] = i \quad (\hbar = 1), \quad H = \frac{p^2}{2m} + V(x) \quad (4.1)$$

Schrödinger eq. for time-dependent states $|\Psi, t\rangle$:

$$i \frac{\partial}{\partial t} |\Psi, t\rangle = H |\Psi, t\rangle \quad (4.2)$$

Formal solution

$$|\Psi, t\rangle = e^{-itH} |\Psi\rangle \quad (|\Psi\rangle = |\Psi, t=0\rangle) \quad (4.3)$$

It is called Schrödinger picture.

Here

$$e^A = \sum_{n=0}^{\infty} \frac{1}{n!} A^n \quad \text{for any operator } A. \quad (4.4)$$

Heisenberg picture:

Define time dependent operators

$$x(t) = e^{itH} x e^{-itH}, \quad p(t) = e^{itH} p e^{-itH} \quad (4.5)$$

For any operator O ,

$$O(t) = e^{itH} O e^{-itH} \quad (4.6)$$

The reason for the definition:

Expectation value of O at t is

$$\langle \Psi, t | O | \Psi, t \rangle = \langle \Psi | e^{itH} O e^{-itH} | \Psi \rangle = \langle \Psi | O(t) | \Psi \rangle \quad (4.7)$$

- Schrödinger picture: states depend on t , operators do not.
- Heisenberg picture: operators depend on t , states do not.

They are physically equivalent. Just a matter of interpretation.

Note

$$\frac{\partial}{\partial t} e^{itH} = iH e^{itH} = e^{itH} iH \quad (4.8)$$

Heisenberg equations:

$$\frac{\partial}{\partial t} O(t) = \left(\frac{\partial}{\partial t} e^{itH} \right) O e^{-itH} + e^{itH} O \left(\frac{\partial}{\partial t} e^{-itH} \right) \quad (4.9)$$

$$= i[H, O(t)]. \quad (4.10)$$

Computations for $x(t)$, $p(t)$:

$$[p^2, x] = p[p, x] + [p, x]p = -2ip, \quad (4.11)$$

$$[V(x), p] = [V(x), -i\partial_x] = i\partial_x V(x) \quad (4.12)$$

$$\frac{d}{dt} x(t) = ie^{itH} [H, x] e^{-itH} = ie^{itH} \left[\frac{p^2}{2m}, x \right] e^{-itH} = e^{itH} \frac{p}{m} e^{-itH} = \frac{p(t)}{m} \quad (4.13)$$

$$\frac{d}{dt} p(t) = ie^{itH} [H, p] e^{-itH} = -e^{itH} \partial_x V(x) e^{-itH} = -\partial_x V(x(t)) \quad (4.14)$$

$$p(t) = m\dot{x}(t) \quad (4.15)$$

$$m\ddot{x}(t) = -\partial_x V(x(t)) \quad (4.16)$$

The same form as classical equations, but $x(t), p(t)$ are operators.

4.2 Heisenberg picture in QFT

In QFT, operators $\phi(\vec{x})$ depend on \vec{x} .

Relativity requires to treat (t, \vec{x}) on the same footing.

Heisenberg picture is more convenient.

$$\phi = \phi(t, \vec{x}) = \phi(x), \quad x = (x^\mu) = (t, \vec{x}). \quad (4.17)$$

$$\phi(t=0, \vec{x}) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(a_{\vec{p}} \exp(i\vec{p} \cdot \vec{x}) + a_{\vec{p}}^\dagger \exp(-i\vec{p} \cdot \vec{x}) \right). \quad (4.18)$$

In Heisenberg picture,

$$\phi(t, \vec{x}) = e^{itH} \phi(0, \vec{x}) e^{-itH} \quad (4.19)$$

Recall

$$[H, a_{\vec{p}}] = -E_p a_{\vec{p}}, \quad [H, a_{\vec{p}}^\dagger] = +E_p a_{\vec{p}}^\dagger \quad (4.20)$$

Then

$$e^{itH}a_{\vec{p}}e^{-itH} = e^{-itE_p}a_{\vec{p}}, \quad e^{itH}a_{\vec{p}}^\dagger e^{-itH} = e^{itE_p}a_{\vec{p}}^\dagger \quad (4.21)$$

Proof: If $H|\Psi\rangle = E|\Psi\rangle$, then

$$Ha_{\vec{p}}^\dagger|\Psi\rangle = (E + E_p)a_{\vec{p}}^\dagger|\Psi\rangle \quad (4.22)$$

Then

$$e^{itH}a_{\vec{p}}^\dagger e^{-itH}|\Psi\rangle = e^{itH}a_{\vec{p}}^\dagger e^{-itE}|\Psi\rangle = e^{it(E+E_p)}e^{-itE}a_{\vec{p}}^\dagger|\Psi\rangle \quad (4.23)$$

$$= e^{itE_p}a_{\vec{p}}^\dagger|\Psi\rangle \quad (4.24)$$

$|\Psi\rangle$ was an arbitrary energy eigenstate, so

$$e^{itH}a_{\vec{p}}^\dagger e^{-itH} = e^{itE_p}a_{\vec{p}}^\dagger. \quad (4.25)$$

$e^{itH}a_{\vec{p}}e^{-itH} = e^{-itE_p}a_{\vec{p}}$ is the same.

$$\phi(t, \vec{x}) = e^{itH}\phi(0, \vec{x})e^{-itH} \quad (4.26)$$

$$= \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(e^{itH}a_{\vec{p}}e^{-itH} \exp(i\vec{p} \cdot \vec{x}) + e^{itH}a_{\vec{p}}^\dagger e^{-itH} \exp(-i\vec{p} \cdot \vec{x}) \right) \quad (4.27)$$

$$= \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \left(a_{\vec{p}} e^{ip \cdot x} + a_{\vec{p}}^\dagger e^{-ip \cdot x} \right) \quad (4.28)$$

where

$$p \cdot x = -p^0 t + \vec{p} \cdot \vec{x}, \quad p^0 = E_p. \quad (4.29)$$

Klein-Gordon equation

$$(\partial^2 - m^2)\phi = 0 \quad (4.30)$$

because

$$(\partial^2 - m^2) \exp(\pm ip \cdot x) = (-p^2 - m^2) \exp(\pm ip \cdot x) = 0. \quad (4.31)$$

4.3 Non-relativistic limit

Non-relativistic limit

$$|\vec{p}| \ll m \quad (4.32)$$

In this limit,

$$E_p \simeq m + \frac{\vec{p}^2}{2m} \quad (4.33)$$

m : Einstein's mass energy

Define

$$K_p = \frac{\vec{p}^2}{2m}. \quad (4.34)$$

$$\phi(x) = \frac{1}{\sqrt{2m}} e^{-imt} \varphi + \frac{1}{\sqrt{2m}} (e^{-imt} \varphi)^* \quad (4.35)$$

$$\varphi(x) = \int \frac{d^3 p}{(2\pi)^3} a_{\vec{p}} e^{-iK_p t + i\vec{p} \cdot \vec{x}}. \quad (4.36)$$

From

$$(i \frac{\partial}{\partial t} + \frac{1}{2m} \vec{\partial}^2) e^{-iK_p t + i\vec{p} \cdot \vec{x}} = (K_p - \frac{\vec{p}^2}{2m}) e^{-iK_p t + i\vec{p} \cdot \vec{x}} = 0. \quad (4.37)$$

Equations of motion

$$i \frac{\partial}{\partial t} \varphi = - \frac{1}{2m} \vec{\partial}^2 \varphi \quad (4.38)$$

The same form as Schrödinger eq. with zero potential (and $\hbar = 1$).

Suppose: $|\Psi\rangle$ is a one-particle state: superposition of $a_{\vec{p}}^\dagger |\Omega\rangle$.

Define

$$\Psi(t, \vec{x}) = \langle \Omega | \varphi | \Psi \rangle. \quad (4.39)$$

Then

$$i \frac{\partial}{\partial t} \Psi(t, \vec{x}) = \langle \Omega | i \frac{\partial}{\partial t} \varphi | \Psi \rangle \quad (4.40)$$

$$= \langle \Omega | - \frac{1}{2m} \vec{\partial}^2 \varphi | \Psi \rangle \quad (4.41)$$

$$= - \frac{1}{2m} \vec{\partial}^2 \Psi(t, \vec{x}). \quad (4.42)$$

Schrödinger eq.

Example:

$$|\Psi\rangle = a_{\vec{p}}^\dagger |\Omega\rangle. \quad (4.43)$$

Then

$$\Psi(t, \vec{x}) = \langle \Omega | \left(\int \frac{d^3 q}{(2\pi)^3} a_{\vec{q}} e^{-iK_q t + i\vec{q} \cdot \vec{x}} \right) a_{\vec{p}}^\dagger |\Omega\rangle \quad (4.44)$$

$$= \langle \Omega | \int \frac{d^3 q}{(2\pi)^3} e^{-iK_q t + i\vec{q} \cdot \vec{x}} (2\pi)^3 \delta^3(\vec{p} - \vec{q}) |\Omega\rangle \quad (\because [a_{\vec{p}}, a_{\vec{q}}^\dagger] = (2\pi)^3 \delta^3(\vec{p} - \vec{q})) \quad (4.45)$$

$$= e^{-iK_p t + i\vec{p} \cdot \vec{x}} \quad (4.46)$$

This is the wavefunction for the state with momentum \vec{p} .

4.4 Second quantization

Consider a QFT with action

$$S = \int dt d^3x \left(\frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} (\vec{\partial} \phi)^2 - \frac{1}{2} m^2 \phi^2 - mV(\vec{x}) \phi^2 \right) \quad (4.47)$$

ϕ : field

$V(\vec{x})$: some function.

EOM

$$-\partial_t^2 \phi = \left(-\vec{\partial}^2 + m^2 + 2mV(\vec{x}) \right) \phi. \quad (4.48)$$

Take

$$\phi(x) = \frac{1}{\sqrt{2m}} e^{-imt} \varphi + \frac{1}{\sqrt{2m}} (e^{-imt} \varphi)^* \quad (4.49)$$

In the limit $m \rightarrow \infty$, EOM becomes

$$i\partial_t \phi = \left(-\frac{1}{2m} \vec{\partial}^2 + V(\vec{x}) \right) \phi. \quad (4.50)$$

Terms oscillating rapidly as e^{2imt} are neglected.

The same form as Schrödinger eq.

Find eigenfunctions $\Psi_n(\vec{x})$,

$$H\Psi_n(\vec{x}) = E_n\Psi_n(\vec{x}) \quad H = -\frac{1}{2m} \vec{\partial}^2 + V(\vec{x}). \quad (4.51)$$

(For notational simplicity, assume discrete states.)

Expand the field φ ,

$$\varphi(t, \vec{x}) = \sum_n \Psi_n(\vec{x}) e^{-iE_n t} A_n, \quad (4.52)$$

A_n : operators.

Quantization of the field ϕ turns out to give (exercise)

$$[\varphi(\vec{x}), \varphi(\vec{y})^\dagger] = \delta^3(\vec{x} - \vec{y}), \quad [\varphi(\vec{x}), \varphi(\vec{y})] = 0, \quad [\varphi(\vec{x})^\dagger, \varphi(\vec{y})^\dagger] = 0 \quad (4.53)$$

This gives (exercise)

$$[A_n, A_m^\dagger] = \delta_{nm}, \quad [A_n, A_m] = 0, \quad [A_n^\dagger, A_m^\dagger] = 0. \quad (4.54)$$

Interpretation:

A_n, A_n^\dagger are creation and annihilation operators for the state corresponding to Ψ_n .

Schrödinger eq. is satisfied not by states but by the field φ :

Sometimes called second quantization.

5 Spinors and Dirac equations

5.1 Spins in QM

Spin degrees of freedom $\vec{S} = (S_i)$:

Commutation relations

$$[S_i, S_j] = i\epsilon_{ijk}S_k \quad (\hbar = 1). \quad (5.1)$$

Electron (or quark) : spin $\frac{1}{2}$, two states

$$|\uparrow\rangle : \text{spin up} \quad S_3|\uparrow\rangle = +\frac{1}{2}|\uparrow\rangle \quad (5.2)$$

$$|\downarrow\rangle : \text{spin down} \quad S_3|\downarrow\rangle = -\frac{1}{2}|\downarrow\rangle. \quad (5.3)$$

State vectors (neglecting positions)

$$c_1|\uparrow\rangle + c_2|\downarrow\rangle \quad (5.4)$$

With basis vectors, $(|\uparrow\rangle, |\downarrow\rangle)$, state vectors are two-dimensional vector space

$$\begin{pmatrix} c_1 \\ c_2 \end{pmatrix} \quad (2 \text{ dim. vector}) \quad (5.5)$$

Pauli matrices

$$\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \quad (5.6)$$

Properties:

$$\sigma^i \sigma^j = \delta^{ij} + i\epsilon^{ijk} \sigma^k \quad (5.7)$$

More explicitly,

$$(\sigma^1)^2 = 1, \quad \sigma^1 \sigma^2 = i\sigma^3, \quad \text{etc..} \quad (5.8)$$

Spin angular momentum for an electron is

$$S^i = \frac{1}{2} \sigma^i \quad (5.9)$$

Eigenstates for S_3

$$S^3 \begin{pmatrix} 1 \\ 0 \end{pmatrix} = +\frac{1}{2} \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad S^3 \begin{pmatrix} 0 \\ 1 \end{pmatrix} = -\frac{1}{2} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (5.10)$$

So far we have neglected position \vec{x} . Let's include \vec{x} .

Wave functions

$$\Psi(\vec{x}, t) = \begin{pmatrix} \Psi_1(\vec{x}, t) \\ \Psi_2(\vec{x}, t) \end{pmatrix} \quad (5.11)$$

Interpretation:

$|\Psi_1(\vec{x})|^2$ is the probability density for spin up states. Similarly for $|\Psi_2(\vec{x})|^2$.

Total probability =1

$$\int d^3x |\Psi(\vec{x})|^2 = 1, \quad |\Psi(\vec{x})|^2 = |\Psi_1(\vec{x})|^2 + |\Psi_2(\vec{x})|^2 \quad (5.12)$$

Experimental fact :

Spin \vec{S} and magnetic field \vec{B} interact, with the Hamiltonian

$$H = H_0 + H_1 \quad (5.13)$$

H_0 does not involve spin,

$$H_0 = \frac{1}{2m}(-i\partial_i - eA_i)^2 + e\phi \quad (5.14)$$

H_1 includes spin,

$$H_1 = -g \frac{e}{2m} \vec{S} \cdot \vec{B} \quad (5.15)$$

g : dimensionless parameter

$$\frac{g-2}{2} = \begin{cases} 0.001\,159\,652\,181\,64(76) & \text{(theoretical calculation by QFT)} \\ 0.001\,159\,652\,180\,73(28) & \text{(experimental value)} \end{cases} \quad (5.16)$$

Extremely good agreement.

5.2 Gamma matrices and Clifford algebra

Dirac equation

$$(\gamma^\mu \partial_\mu + m)\psi = 0. \quad (5.17)$$

γ^μ ($\mu = 0, 1, 2, 3$) : gamma matrices, to be discussed.

Historical motivation: relativistic version of Schrödinger equation.

This motivation is not valid today. It is an EOM for fields.

- Klein-Gordon eq. : spin 0 fields
- Dirac eq. : spin $\frac{1}{2}$ fields
- Maxwell (or Yang-Mills) eq. : spin 1 fields

Properties of γ^μ : For plane waves $\Psi \propto \exp(ip_\mu x^\mu)$

$$(\gamma^\mu \partial_\mu + m)\Psi = (i\gamma_\mu p^\mu + m)\Psi = 0 \quad (5.18)$$

$$m\Psi = (-i\gamma_\mu p^\mu)\Psi \quad (5.19)$$

Using it twice,

$$m^2 \Psi = -(\gamma_\mu p^\mu)^2 \Psi \quad (5.20)$$

Requirement: there is a solution with $m^2 = -p^2$.

$$p^2 = (\gamma_\mu p^\mu)^2 = \gamma_\mu \gamma_\nu p^\mu p^\nu = \frac{1}{2}(\gamma_\mu \gamma_\nu + \gamma_\nu \gamma_\mu) p^\mu p^\nu. \quad (5.21)$$

In general, define anticommutator

$$\{A, B\} = AB + BA \quad (5.22)$$

If

$$\{\gamma_\mu, \gamma_\nu\} = \gamma_\mu \gamma_\nu + \gamma_\nu \gamma_\mu = 2g_{\mu\nu} \quad (5.23)$$

Then $p^2 = -m^2$. This anticommutation relation is called Clifford algebra.

γ_μ need to be matrices.

In three (not four) dimensions, Pauli matrices σ_i satisfy the desired relations:

$$\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \quad (5.24)$$

$$\{\sigma^i, \sigma^j\} = (\sigma^i \sigma^j + \sigma^j \sigma^i) = 2\delta_{ij} \quad (5.25)$$

Pauli matrices: 2×2 . In four dimensions, $\{\gamma_\mu, \gamma_\nu\} = 2g_{\mu\nu}$ requires 4×4 matrices.

Explicit examples of γ^μ :

$$\gamma^i = \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} \quad (5.26)$$

$$\gamma^0 = \begin{pmatrix} 0 & -i \\ -i & 0 \end{pmatrix} \quad (5.27)$$

Each block is 2×2 . In total, 4×4

Check of Clifford algebra:

$$\gamma^i \gamma^j = \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 0 & -i\sigma^j \\ i\sigma^j & 0 \end{pmatrix} = \begin{pmatrix} \sigma^i \sigma^j & 0 \\ 0 & \sigma^i \sigma^j \end{pmatrix} \quad (5.28)$$

$$\gamma^i \gamma^j + \gamma^j \gamma^i = \begin{pmatrix} 2\delta^{ij} & 0 \\ 0 & 2\delta^{ij} \end{pmatrix} = 2\delta^{ij} \quad (5.29)$$

$$\gamma^i \gamma^0 = \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 0 & -i \\ -i & 0 \end{pmatrix} = \begin{pmatrix} -\sigma^j & 0 \\ 0 & \sigma^j \end{pmatrix} \quad (5.30)$$

$$\gamma^0 \gamma^i = \begin{pmatrix} 0 & -i \\ -i & 0 \end{pmatrix} \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} = \begin{pmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{pmatrix} \quad (5.31)$$

$$\gamma^i \gamma^0 + \gamma^0 \gamma^i = 0 \quad (5.32)$$

$$\{\gamma^0, \gamma^0\} = 2(\gamma^0)^2 = -2 \quad (5.33)$$

Combining them,

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu} \quad (5.34)$$

By using tensor products in linear algebra,

$$\gamma^i = \sigma^2 \otimes \sigma^i \quad (5.35)$$

$$\gamma^0 = -i\sigma^1 \otimes 1 \quad (5.36)$$

γ_μ need not be the above explicit form.

But we require

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}, \quad (\gamma_0)^\dagger = -\gamma_0, \quad (\gamma_i)^\dagger = \gamma_i. \quad (5.37)$$

5.3 Lorentz invariance

Lorentz transformation

$$y^\mu = \Lambda^\mu_\nu x^\nu \quad (5.38)$$

ds^2 must be invariant:

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu = g_{\mu\nu} dy^\mu dy^\nu \quad (5.39)$$

$$\implies g_{\rho\sigma} \Lambda^\rho_\mu \Lambda^\sigma_\nu = g_{\mu\nu} \quad (5.40)$$

Transformation of Dirac equation

$$(\gamma^\mu \frac{\partial}{\partial x^\mu} + m)\psi = (\gamma^\mu \frac{\partial y^\nu}{\partial x^\mu} \frac{\partial}{\partial y^\nu} + m)\psi \quad (5.41)$$

$$= (\gamma^\mu \Lambda^\nu_\mu \frac{\partial}{\partial y^\nu} + m)\psi \quad (5.42)$$

Assume ψ also transforms:

$$\psi'(y) = S(\Lambda)\psi(x) \quad (5.43)$$

$S(\Lambda)$: some matrix given in terms of γ^μ , to be determined.

Comparison: a vector field $A^\mu(x)$ transforms as

$$A'^\mu(y) = \Lambda^\mu_\nu A^\nu(x). \quad (5.44)$$

ψ is neither scalar nor vector. It is called spinor.

$$(\gamma^\mu \Lambda^\nu_\mu \frac{\partial}{\partial y^\nu} + m)\psi = \left(\gamma^\mu \Lambda^\nu_\mu \frac{\partial}{\partial y^\nu} + m \right) S(\Lambda)^{-1} \psi' \quad (5.45)$$

$$= S(\Lambda)^{-1} \left(\Lambda^\nu_\mu S(\Lambda) \gamma^\mu S(\Lambda)^{-1} \frac{\partial}{\partial y^\nu} + m \right) \psi' \quad (5.46)$$

If

$$\Lambda^\nu_\mu S(\Lambda) \gamma^\mu S(\Lambda)^{-1} = \gamma^\nu \quad (5.47)$$

$$\Longleftrightarrow S(\Lambda)^{-1} \gamma^\mu S(\Lambda) = \Lambda^\mu_\nu \gamma^\nu \quad (5.48)$$

then

$$\left(\gamma^\mu \frac{\partial}{\partial x^\mu} + m \right) \psi(x) = S(\Lambda)^{-1} \left(\gamma^\mu \frac{\partial}{\partial y^\mu} + m \right) \psi'(y) \quad (5.49)$$

Then Dirac equation is Lorentz invariant.

We need to find $S(\Lambda)$.

Infinitesimal Lorentz transformation

Consider

$$\Lambda^\mu_\nu = \delta^\mu_\nu + \omega^\mu_\nu, \quad \omega : \text{infinitesimal} \quad (5.50)$$

$$g_{\mu\nu} = g_{\rho\sigma} \Lambda^\rho_\mu \Lambda^\sigma_\nu \quad (5.51)$$

$$= g_{\rho\sigma} (\delta^\rho_\mu + \omega^\rho_\mu) (\delta^\sigma_\nu + \omega^\sigma_\nu) \quad (5.52)$$

$$= g_{\mu\nu} + \omega_{\nu\mu} + \omega_{\mu\nu} \quad (5.53)$$

Here indices are raised and lowered by $g_{\mu\nu}$:

$$\omega_{\mu\nu} = g_{\mu\rho} \omega^\rho_\nu \quad (5.54)$$

We get

$$\omega_{\nu\mu} + \omega_{\mu\nu} = 0 \quad : \text{antisymmetric} \quad (5.55)$$

$S(\Lambda)$ for $\Lambda^\mu_\nu = \delta^\mu_\nu + \omega^\mu_\nu$.

The answer that works :

$$S(\Lambda) = 1 + \frac{1}{4}\omega_{\rho\sigma}\gamma^\rho\gamma^\sigma \quad (5.56)$$

Not so many possibilities if it is given in terms of γ^μ and $\omega_{\mu\nu}$.

Check of $S(\Lambda)^{-1}\gamma^\mu S(\Lambda) = \Lambda^\mu_\nu\gamma^\nu$:

For infinitesimal $\omega_{\mu\nu}$,

$$S(\Lambda)^{-1} = 1 - \frac{1}{4}\omega_{\rho\sigma}\gamma^\rho\gamma^\sigma \quad (5.57)$$

Then

$$S(\Lambda)^{-1}\gamma^\mu S(\Lambda) \quad (5.58)$$

$$= (1 - \frac{1}{4}\omega_{\rho\sigma}\gamma^\rho\gamma^\sigma)\gamma^\mu(1 + \frac{1}{4}\omega_{\rho\sigma}\gamma^\rho\gamma^\sigma) \quad (5.59)$$

$$= \gamma^\mu - \frac{1}{4}\omega_{\rho\sigma}(\gamma^\rho\gamma^\sigma\gamma^\mu - \gamma^\mu\gamma^\rho\gamma^\sigma) \quad (5.60)$$

$$= \gamma^\mu - \frac{1}{4}\omega_{\rho\sigma}[\gamma^\rho(\gamma^\sigma\gamma^\mu + \gamma^\mu\gamma^\sigma) - (\gamma^\rho\gamma^\mu + \gamma^\mu\gamma^\rho)\gamma^\sigma] \quad (5.61)$$

$$= \gamma^\mu - \frac{1}{4}\omega_{\rho\sigma}[2\gamma^\rho g^{\mu\sigma} - 2g^{\rho\mu}\gamma^\sigma] \quad (5.62)$$

$$= \gamma^\mu + \omega^\mu_\rho\gamma^\rho \quad (5.63)$$

$$= \Lambda^\mu_\rho\gamma^\rho \quad (5.64)$$

This confirms the desired property of $S(\Lambda)$.

Finite Lorentz transformation

For finite ω , suppose

$$\Lambda = \exp(\omega) \quad (5.65)$$

In general, for a matrix $\omega = (\omega^\mu_\nu)$,

$$\exp(\omega) = \sum_{k=0}^{\infty} \frac{1}{k!}\omega^k. \quad (5.66)$$

Another formula:

$$\exp(\omega) = \lim_{N \rightarrow \infty} \left(1 + \frac{1}{N}\omega\right)^N \quad (5.67)$$

$\exp(\omega)$ is given by doing an infinitesimal $(1 + \frac{1}{N}\omega)$ many times.
Then

$$S(\Lambda) = \exp(\frac{1}{4}\omega_{\rho\sigma}\gamma^\rho\gamma^\sigma) \quad (5.68)$$

$$= \lim_{N \rightarrow \infty} \left(1 + \frac{1}{4} \frac{\omega_{\mu\nu}}{N} \gamma^\mu \gamma^\nu\right)^N \quad (5.69)$$

Transformation law

$$\psi'(y) = S(\Lambda)\psi(x), \quad (y = \Lambda x) \quad (5.70)$$

A field transforming like this: spinor field.

ψ has 4 components

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix} \quad (5.71)$$

$S(\Lambda)$ is a 4×4 matrix.

Summary: Under Lorentz transformation $y^\mu = \Lambda^\mu_\nu x^\nu$,

- Scalar : $\phi'(y) = \phi(x)$
- Vector : $A'^\mu(y) = \Lambda^\mu_\nu A^\nu(x)$
- Spinor : $\psi'(y) = S(\Lambda)\psi(x)$

5.4 More on Lorentz transformation

Summary of previous discussions:

$$y^\mu = \Lambda^\mu_\nu x^\nu \quad (5.72)$$

$$\psi'(y) = S(\Lambda)\psi(x) \quad (5.73)$$

Here

$$\Lambda = \exp(\omega) \quad S(\Lambda) = \exp(\frac{1}{4}\gamma^\mu\gamma^\nu\omega_{\mu\nu}) \quad (5.74)$$

$$S(\Lambda)^{-1}\gamma^\mu S(\Lambda) = \Lambda^\mu_\nu \gamma^\nu. \quad (5.75)$$

$S(\Lambda)$ is not a unitary matrix.

For later purposes, study $S(\Lambda)^\dagger$.

From $\gamma^i\gamma^0 = -\gamma^0\gamma^i$ and $(\gamma^i)^\dagger = \gamma^i$,

$$(\gamma^i)^\dagger = -(\gamma^0)\gamma^i(\gamma^0)^{-1} \quad (5.76)$$

From $\gamma^0 \gamma^0 = \gamma^0 \gamma^0$ and $(\gamma^0)^\dagger = -\gamma^0$,

$$(\gamma^0)^\dagger = -(\gamma^0) \gamma^0 (\gamma^0)^{-1} \quad (5.77)$$

Summarizing,

$$(\gamma^\mu)^\dagger = -(\gamma^0) \gamma^\mu (\gamma^0)^{-1} \quad (5.78)$$

Then

$$(\gamma^\mu \gamma^\nu \omega_{\mu\nu})^\dagger = (\gamma^\nu)^\dagger (\gamma^\mu)^\dagger \omega_{\mu\nu} \quad (5.79)$$

$$= (\gamma^0) \gamma^\nu (\gamma^0)^{-1} (\gamma^0) \gamma^\mu (\gamma^0)^{-1} \omega_{\mu\nu} \quad (5.80)$$

$$= (\gamma^0) (\gamma^\nu \gamma^\mu \omega_{\mu\nu}) (\gamma^0)^{-1} \quad (5.81)$$

$$= -(\gamma^0) (\gamma^\mu \gamma^\nu \omega_{\mu\nu}) (\gamma^0)^{-1} \quad (\because \omega_{\mu\nu} = -\omega_{\nu\mu}). \quad (5.82)$$

Recall the definition of exp of matrices,

$$\exp(A) = \sum_{k=0}^{\infty} \frac{1}{k!} A^k \quad (5.83)$$

From it,

$$\exp(A)^\dagger = \exp(A^\dagger), \quad \exp(BAB^{-1}) = B \exp(A) B^{-1}, \quad \exp(A)^{-1} = \exp(-A) \quad (5.84)$$

Then

$$S(\Lambda)^\dagger = \exp\left(\frac{1}{4}(\gamma^\mu \gamma^\nu \omega_{\mu\nu})^\dagger\right) \quad (5.85)$$

$$= \exp\left(-\frac{1}{4}(\gamma^0)(\gamma^\mu \gamma^\nu \omega_{\mu\nu})(\gamma^0)^{-1}\right) \quad (5.86)$$

$$= (\gamma^0) \exp\left(-\frac{1}{4}(\gamma^\mu \gamma^\nu \omega_{\mu\nu})\right) (\gamma^0)^{-1} \quad (5.87)$$

$$= (\gamma^0) S(\Lambda)^{-1} (\gamma^0)^{-1} \quad (5.88)$$

$S(\Lambda)^{-1}$ is the inverse matrix of $S(\Lambda)$.

Summary:

$$S(\Lambda)^{-1} = (\gamma^0)^{-1} S(\Lambda)^\dagger (\gamma^0). \quad (5.89)$$

Notation :

$$\overline{\psi} = i\psi^\dagger \gamma^0 \quad (5.90)$$

Lorentz transformation

$$\overline{\psi}'(y) = i\psi^\dagger(x) S(\Lambda)^\dagger \gamma^0 \quad (5.91)$$

$$= i\psi^\dagger(x) \gamma^0 (\gamma^0)^{-1} S(\Lambda)^\dagger \gamma^0 \quad (5.92)$$

$$= \overline{\psi} S(\Lambda)^{-1} \quad (5.93)$$

This is useful for constructing Lorentz covariant quantities.

Some examples:

1. A scalar

$$\bar{\psi}'\psi' = \bar{\psi}S(\Lambda)^{-1}S(\Lambda)\psi = \bar{\psi}\psi. \quad (5.94)$$

2. A vector

$$\bar{\psi}'\gamma^\mu\psi' = \bar{\psi}S(\Lambda)^{-1}\gamma^\mu S(\Lambda)\psi = \Lambda^\mu{}_\nu\bar{\psi}\gamma^\nu\psi. \quad (5.95)$$

More general tensors

$$\bar{\psi}'\gamma^{\mu_1}\dots\gamma^{\mu_k}\psi' = \Lambda^{\mu_1}{}_{\nu_1}\dots\Lambda^{\mu_k}{}_{\nu_k}\bar{\psi}\gamma^{\nu_1}\dots\gamma^{\nu_k}\psi \quad (5.96)$$

3. Another scalar

$$\bar{\psi}'\gamma^\mu\frac{\partial}{\partial y^\mu}\psi' = \bar{\psi}\gamma^\mu\frac{\partial}{\partial x^\mu}\psi \quad (5.97)$$

and so on.

5.5 Spin

Generally,

$$\text{symmetry} \implies \text{conserved quantity} \quad (5.98)$$

In particular,

$$\text{rotational symmetry} \implies \text{angular momentum} \quad (5.99)$$

The angular momentum is found by infinitesimal rotation

$$y^\mu = \Lambda^\mu{}_\nu x^\nu \quad (5.100)$$

$$\psi'(y) = S(\Lambda)\psi(x) \quad (5.101)$$

Change notation $y \rightarrow x$

$$\psi'(x) = S(\Lambda)\psi(\Lambda^{-1}x). \quad (5.102)$$

Rotation : $\omega_{\mu\nu}$ has only spatial components

$$\omega_{ij} \neq 0, \quad \omega_{ij} = -\omega_{ji} \quad (5.103)$$

$$\Lambda^\mu{}_\nu = \delta^\mu{}_\nu + \omega^\mu{}_\nu.$$

Introduce ω^i

$$\omega_{ij} = \epsilon_{ikj}\omega^k, \quad \omega_{0i} = 0 \quad (5.104)$$

Upper and lower indices are not distinguished for $i = 1, 2, 3$.

$$\Lambda^i{}_j x^j = x^i + \epsilon^{ikj}\omega_k x_j = \vec{x} + \vec{\omega} \times \vec{x} \quad (5.105)$$

By previous results,

$$S(\Lambda) = 1 + \frac{1}{4}\omega_{ij}\gamma^i\gamma^j = 1 - \frac{1}{4}\epsilon_{ijk}\omega^i\gamma^j\gamma^k \quad (5.106)$$

$$\psi'(\vec{x}) = S(\Lambda)\psi(\Lambda^{-1}\vec{x}) \quad (5.107)$$

$$= (1 - \frac{1}{4}\epsilon^{ijk}\omega_i\gamma_j\gamma_k)\psi(\vec{x} - \omega \times \vec{x}) \quad (5.108)$$

$$= \psi(\vec{x}) - \left(\frac{1}{4}\epsilon_{ijk}\omega^i\gamma^j\gamma^k + \epsilon^{ijk}\omega_j x_k \partial_i \right) \psi(\vec{x}) \quad (5.109)$$

$$= \psi(\vec{x}) - i\omega^i J_i \psi(\vec{x}) \quad (5.110)$$

where

$$J_i = -\frac{1}{4}i\epsilon_{ijk}\gamma^j\gamma^k - i\epsilon_{ijk}x^j\partial^k = S_i + L_i \quad (5.111)$$

The generator for infinitesimal rotation.

$$\text{orbital angular momentum : } L_i = -i\epsilon_{ijk}x^j\partial^k = \vec{x} \times \vec{p} \quad (p_i = -i\partial_i) \quad (5.112)$$

$$\text{spin angular momentum : } S_i = -\frac{1}{4}i\epsilon_{ijk}\gamma^j\gamma^k \quad (5.113)$$

$$\text{total angular momentum : } J_i = L_i + S_i \quad (5.114)$$

Computation of S_i : using

$$\gamma^i = \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} \quad (5.115)$$

then

$$S_i = -\frac{1}{4}i\epsilon_{ijk}\gamma^j\gamma^k \quad (5.116)$$

$$= -\frac{1}{4}i\epsilon_{ijk} \begin{pmatrix} \sigma^i\sigma^j & 0 \\ 0 & \sigma^i\sigma^j \end{pmatrix} \quad (5.117)$$

$$= \frac{1}{2} \begin{pmatrix} \sigma^i & 0 \\ 0 & \sigma^i \end{pmatrix} \quad (5.118)$$

$\frac{1}{2}\sigma^i$: spin operator in QM.

In particular,

$$S_z = S_3 = \frac{1}{2} \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & 1 & \\ & & & -1 \end{pmatrix} \quad (5.119)$$

$$\text{eigenvalue : } \pm \frac{1}{2} \quad : \quad \text{spin } \frac{1}{2} \quad (5.120)$$

Remark:

The relation between QM wavefunction Ψ and the operator ψ :

$$\Psi(x) = \langle \Omega | \psi(x) | \Psi \rangle, \quad |\Psi\rangle : \text{a single particle state in QFT} \quad (5.121)$$

S_i is the spin operator for QM wavefunction Ψ .

5.6 Coupling to EM fields

Recall

$$A_\mu = (-\phi, A_i) \quad : \text{EM potential} \quad (5.122)$$

Gauge transformation

$$A_\mu \rightarrow A'_\mu = A_\mu + \partial_\mu \alpha, \quad \alpha : \text{arbitrary function} \quad (5.123)$$

Under it,

$$F'_{\mu\nu} = \partial_\mu A'_\nu - \partial_\nu A'_\mu = \partial_\mu A_\nu - \partial_\nu A_\mu = F_{\mu\nu} \quad (5.124)$$

EM tensor $F_{\mu\nu}$ is invariant.

For wavefunction Ψ ,

$$\Psi \rightarrow \Psi' = \exp(ie\alpha)\Psi. \quad (5.125)$$

Covariant derivative

$$D_\mu \Psi = (\partial_\mu - ieA_\mu)\Psi \quad (5.126)$$

For $D'_\mu = \partial_\mu - ieA'_\mu$

$$D'_\mu \Psi' = \exp(ie\alpha)D_\mu \Psi. \quad (5.127)$$

In the same way, for the operator ψ ,

$$\psi \rightarrow \psi' = \exp(ie\alpha)\psi \quad (5.128)$$

$$D'_\mu \psi' = \exp(ie\alpha)D_\mu \psi. \quad (5.129)$$

More generally, if some $\Phi(x)$ transforms as

$$\Phi(x) \rightarrow \Phi'(x) = \exp(iq\alpha), \quad (5.130)$$

define

$$D_\mu \Phi(x) = (\partial_\mu - iqA_\mu)\Phi. \quad (5.131)$$

q : charge of the operator

For example,

$$\bar{\psi} = i\psi^\dagger \gamma^0 \quad (5.132)$$

$$\bar{\psi} \rightarrow \bar{\psi}' = \exp(-ie\alpha)\bar{\psi} \quad (5.133)$$

Then

$$D_\mu \bar{\psi} = (\partial_\mu + ieA_\mu)\bar{\psi}. \quad (5.134)$$

Integration by parts: for example,

$$\int d^4x \bar{\psi} \gamma^\mu (\partial_\mu - ieA_\mu) \psi = \int d^4x (-\partial_\mu \bar{\psi} - ieA_\mu \bar{\psi}) \gamma^\mu \psi \quad (5.135)$$

hence

$$\int d^4x \bar{\psi} \gamma^\mu D_\mu \psi = - \int d^4x (D_\mu \bar{\psi}) \gamma^\mu \psi. \quad (5.136)$$

Dirac eq. when $A_\mu \neq 0$: replace

$$(\gamma^\mu \partial_\mu + m)\psi = 0 \quad (5.137)$$

by

$$(\gamma^\mu D_\mu + m)\psi = 0 \quad (5.138)$$

Feynman slash notation : For any vector B^μ ,

$$\not{B} = \gamma^\mu B_\mu. \quad (5.139)$$

Then

$$(\not{D} + m)\psi = (\gamma^\mu D_\mu + m)\psi = 0. \quad (5.140)$$

5.7 Magnetic moment

Purpose: to compute the interaction $\propto \vec{B} \cdot \vec{S}$ in QM.

Dirac eq.

$$(\gamma^\mu D_\mu + m)\psi = 0 \quad (5.141)$$

$$D_\mu = \partial_\mu - ieA_\mu$$

$$m^2 \psi = (-\gamma^\mu D_\mu)^2 \psi \quad (5.142)$$

$$= \gamma^\mu D_\mu \gamma^\nu D_\nu \psi \quad (5.143)$$

$$= \left(\frac{1}{2} \{ \gamma^\mu, \gamma^\nu \} + \frac{1}{2} [\gamma^\mu, \gamma^\nu] \right) D_\mu D_\nu \psi \quad (5.144)$$

$$= \left(g^{\mu\nu} + \frac{1}{2} [\gamma^\mu, \gamma^\nu] \right) D_\mu D_\nu \psi \quad (5.145)$$

$$= \left(D^2 + \frac{1}{4} [\gamma^\mu, \gamma^\nu] [D_\mu, D_\nu] \right) \psi \quad (5.146)$$

$$(5.147)$$

Commutation relation

$$[D_\mu, D_\nu] = [\partial_\mu - ieA_\mu, \partial_\nu - ieA_\nu] \quad (5.148)$$

$$= -ie(\partial_\mu A_\nu - \partial_\nu A_\mu) \quad (5.149)$$

$$= -ieF_{\mu\nu} \quad (5.150)$$

$$\left(D^2 - m^2 - ie\frac{1}{2}\gamma^\mu\gamma^\nu F_{\mu\nu} \right) \psi = 0 \quad (5.151)$$

Suppose

$$A_\mu = (0, A_i), \quad \partial_t A_i = 0. \quad (5.152)$$

In particular, $\vec{B} \neq 0$, $\vec{E} = 0$.

$$F_{ij} = \epsilon_{ijk} B^k, \quad F_{0i} = 0. \quad (5.153)$$

Then

$$-ie\frac{1}{2}\gamma^\mu\gamma^\nu F_{\mu\nu} \quad (5.154)$$

$$= -ie\frac{1}{2}(\gamma^i\gamma^j\epsilon_{ijk}B^k) \quad (5.155)$$

$$= 2eS_k B^k \quad (S_k = -\frac{i}{4}\gamma^i\gamma^j\epsilon_{ijk} : \text{spin}) \quad (5.156)$$

$$\left(-\partial_t^2 + \vec{D}^2 - m^2 + 2eS_k B^k \right) \psi = 0 \quad (5.157)$$

Non-relativistic limit $m \rightarrow \infty$: define

$$\psi = e^{-imt}\tilde{\psi} + e^{imt}\tilde{\psi}' \quad (5.158)$$

$$-\partial_t^2(e^{-imt}\tilde{\psi}) = e^{-imt}(m^2 + 2im\partial_t - \partial_t^2)\tilde{\psi} \quad (5.159)$$

EOM

$$0 = e^{imt} \left(-\partial_t^2 + \vec{D}^2 - m^2 + 2eS_k B^k \right) \psi \quad (5.160)$$

$$= \left(-\partial_t^2 + 2im\partial_t + \vec{D}^2 + 2eS_k B^k \right) \tilde{\psi} + e^{2imt}(\dots)\tilde{\psi}' \quad (5.161)$$

In the limit $m \rightarrow \infty$, neglect

(1) ∂_t^2 compared to $m\partial_t$

(2) rapidly oscillating e^{2imt}

Then

$$\left(2im\partial_t + \vec{D}^2 + 2eS_k B^k \right) \tilde{\psi} = 0 \quad (5.162)$$

$$\implies i\partial_t \tilde{\psi} = \left(-\frac{1}{2m}\vec{D}^2 - \frac{e}{m}\vec{S} \cdot \vec{B} \right) \tilde{\psi} \quad (5.163)$$

In QM, it corresponds to the interaction

$$H_1 = -g \frac{e}{2m} \vec{S} \cdot \vec{B}, \quad (5.164)$$

$$g = 2 \quad (5.165)$$

This is a good approximation to the actual value

$$g = 2.002 \dots \quad (5.166)$$

5.8 Solutions of Dirac equation

Purpose: to obtain solutions of Dirac eq. when $A_\mu = 0$:

$$(\not{\partial} + m)\psi = 0. \quad (5.167)$$

Modes with momentum p ,

$$\psi = e^{\pm i p \cdot x} w, \quad p^0 > 0, \quad w = \begin{pmatrix} w_1 \\ w_2 \\ w_3 \\ w_4 \end{pmatrix} : \text{constant}. \quad (5.168)$$

Then

$$(\not{\partial} + m)(e^{\pm i p \cdot x} w) = e^{\pm i p \cdot x} (\pm i \not{p} + m)w = 0. \quad (5.169)$$

$$\implies (\pm i \not{p} + m)w = 0 \quad (5.170)$$

Computation of w :

Recall

$$\gamma^i = \begin{pmatrix} 0 & -i\sigma^i \\ i\sigma^i & 0 \end{pmatrix} \quad \gamma^0 = \begin{pmatrix} 0 & -i \\ -i & 0 \end{pmatrix} \quad (5.171)$$

Write

$$w = \begin{pmatrix} \chi \\ \eta \end{pmatrix} \quad (5.172)$$

χ, η : 2-component

$$\begin{pmatrix} m & \mp(p^0 - \vec{\sigma}\vec{p}) \\ \mp(p^0 + \vec{\sigma}\vec{p}) & m \end{pmatrix} \begin{pmatrix} \chi \\ \eta \end{pmatrix} \quad (5.173)$$

Then

$$\pm(p^0 + \vec{\sigma}\vec{p})\chi = m\eta, \quad \pm(p^0 - \vec{\sigma}\vec{p})\eta = m\chi. \quad (5.174)$$

From them,

$$m^2 \eta = (p^0 + \vec{\sigma} \vec{p})(p^0 - \vec{\sigma} \vec{p}) \eta \quad (5.175)$$

$$= [(p^0)^2 - (\vec{\sigma} \vec{p})^2] \eta \quad (5.176)$$

$$= [(p^0)^2 - \vec{p}^2] \eta \quad (5.177)$$

$$= -p^2 \eta \quad (5.178)$$

Here

$$(\vec{\sigma} \vec{p})^2 = \frac{1}{2} \{\sigma_i, \sigma_j\} p^i p^j = \delta_{ij} p^i p^j. \quad (5.179)$$

$$m^2 = -p^2 \implies p^0 = E_p = \sqrt{\vec{p}^2 + m^2} \quad : \text{ on-shell} \quad (5.180)$$

$p^0 - \vec{\sigma} \vec{p}$ and $p^0 + \vec{\sigma} \vec{p}$ are

- (i) hermitian matrices
- (ii) with positive eigenvalues (exercise)
- (iii) and commute with each other .

For any hermitian A with positive eigenvalues, \sqrt{A} such that $(\sqrt{A})^2 = A$ can be defined.

Proof: If $A = U D U^\dagger$, with U unitary and D diagonal

$$D = \begin{pmatrix} D_1 & & \\ & D_2 & \\ & & \ddots \end{pmatrix} \quad (5.181)$$

Then $\sqrt{A} = U \sqrt{D} U^\dagger$ with

$$D = \begin{pmatrix} \sqrt{D_1} & & \\ & \sqrt{D_2} & \\ & & \ddots \end{pmatrix} \quad (5.182)$$

If $[A, B] = 0$, they are simultaneously diagonalizable and

$$\sqrt{A} \sqrt{B} = \sqrt{AB}. \quad (5.183)$$

We can define

$$\sqrt{p^0 - \vec{\sigma} \vec{p}}, \quad \sqrt{p^0 + \vec{\sigma} \vec{p}} \quad (5.184)$$

and

$$\sqrt{p^0 - \vec{\sigma} \vec{p}} \sqrt{p^0 + \vec{\sigma} \vec{p}} = \sqrt{(p^0 - \vec{\sigma} \vec{p})(p^0 + \vec{\sigma} \vec{p})} = \sqrt{(p^0)^2 - \vec{p}^2} = m. \quad (5.185)$$

Solution for w :

$$w = \begin{pmatrix} \chi \\ \eta \end{pmatrix} = \begin{pmatrix} \sqrt{p^0 - \vec{\sigma} \vec{p}} \xi \\ \pm \sqrt{p^0 + \vec{\sigma} \vec{p}} \xi \end{pmatrix}, \quad \xi : \text{arbitrary} \quad (5.186)$$

In fact,

$$\pm (p^0 + \vec{\sigma}\vec{p})\chi = \pm(p^0 + \vec{\sigma}\vec{p})\sqrt{p^0 - \vec{\sigma}\vec{p}}\xi = \pm m\sqrt{p^0 + \vec{\sigma}\vec{p}}\xi = m\eta. \quad (5.187)$$

$$\pm (p^0 - \vec{\sigma}\vec{p})\eta = (p^0 - \vec{\sigma}\vec{p})\sqrt{p^0 + \vec{\sigma}\vec{p}}\xi = m\sqrt{p^0 - \vec{\sigma}\vec{p}}\xi = m\chi. \quad (5.188)$$

Define

$$u_{\vec{p}}(\xi) = \begin{pmatrix} \sqrt{p^0 - \vec{\sigma}\vec{p}}\xi \\ \sqrt{p^0 + \vec{\sigma}\vec{p}}\xi \end{pmatrix} \quad (5.189)$$

$$v_{\vec{p}}(\xi) = \begin{pmatrix} \sqrt{p^0 - \vec{\sigma}\vec{p}}\xi \\ -\sqrt{p^0 + \vec{\sigma}\vec{p}}\xi \end{pmatrix} \quad (5.190)$$

ξ : arbitrary.

Then

$$(i\not{p} + m)u_{\vec{p}}(\xi) = 0, \quad (-i\not{p} + m)v_{\vec{p}}(\xi) = 0. \quad (5.191)$$

Solutions of Dirac equation are

$$e^{ip \cdot x}u_{\vec{p}}(\xi), \quad e^{-ip \cdot x}v_{\vec{p}}(\xi) \quad (p^0 = E_p) \quad (5.192)$$

Orthonormality properties:

$$u_{\vec{p}}(\xi)^\dagger u_{\vec{p}}(\xi') = \left(\xi^\dagger \sqrt{p^0 - \vec{\sigma}\vec{p}} \quad \xi^\dagger \sqrt{p^0 + \vec{\sigma}\vec{p}} \right) \begin{pmatrix} \sqrt{p^0 - \vec{\sigma}\vec{p}}\xi' \\ \sqrt{p^0 + \vec{\sigma}\vec{p}}\xi' \end{pmatrix} \quad (5.193)$$

$$= \xi^\dagger ((p^0 - \vec{\sigma}\vec{p}) + (p^0 + \vec{\sigma}\vec{p})) \xi' \quad (5.194)$$

$$= 2E_p \xi^\dagger \xi' \quad (5.195)$$

In the same way,

$$v_{\vec{p}}(\xi)^\dagger v_{\vec{p}}(\xi') = 2E_p \xi^\dagger \xi'. \quad (5.196)$$

Also,

$$u_{\vec{p}}(\xi)^\dagger v_{-\vec{p}}(\xi') = \left(\xi^\dagger \sqrt{p^0 - \vec{\sigma}\vec{p}} \quad \xi^\dagger \sqrt{p^0 + \vec{\sigma}\vec{p}} \right) \begin{pmatrix} \sqrt{p^0 + \vec{\sigma}\vec{p}}\xi' \\ -\sqrt{p^0 - \vec{\sigma}\vec{p}}\xi' \end{pmatrix} \quad (5.197)$$

$$= \xi^\dagger (m - m) \xi' \quad (5.198)$$

$$= 0. \quad (5.199)$$

Setting

$$\xi_1 = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \xi_2 = \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (5.200)$$

and defining

$$u_{(\vec{p},s)} = u_{\vec{p}}(\xi_s), \quad v_{(\vec{p},s)} = v_{\vec{p}}(\xi_s) \quad (s = 1, 2) \quad (5.201)$$

Summary:

For a given \vec{p} , the following four vectors are orthogonal, with absolute value $2E_p$.

$$u_{(\vec{p},1)}, \quad u_{(\vec{p},2)}, \quad v_{(-\vec{p},1)}, \quad v_{(-\vec{p},2)} \quad (5.202)$$

6 Quantization of Dirac field

Recall: for Lorentz transformation $\Lambda = (\Lambda^\mu_\nu)$, $S(\Lambda)$ is a 4×4 matrix such that

$$S(\Lambda)^{-1} \gamma^\mu S(\Lambda) = \Lambda^\mu_\nu \gamma^\nu \quad (6.1)$$

Dirac field ψ transforms

$$\psi'(x) = S(\Lambda) \psi(\Lambda^{-1}x) \quad (6.2)$$

ψ has 4 components

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix} \quad (6.3)$$

γ^μ and $S(\Lambda)$ are 4×4 matrices.

Feynman slash notation : For any vector B^μ ,

$$\not{B} = \gamma^\mu B_\mu \quad (6.4)$$

In particular,

$$\not{\partial} = \gamma^\mu \partial_\mu \quad (6.5)$$

$$\not{D} = \gamma^\mu D_\mu \quad (6.6)$$

6.1 Analytical mechanics

The Lagrangian

$$\mathcal{L} = -\bar{\psi}(\not{D} + m)\psi \quad (6.7)$$

$$\bar{\psi} = \psi^\dagger (i\gamma^0) \quad (6.8)$$

This is Lorentz invariant.

Action

$$S = \int d^4x \mathcal{L} \quad (6.9)$$

Equations of motion : $\psi \rightarrow \psi + \delta\psi$,

$$\delta S = - \int d^4x \delta\bar{\psi}(\not{D} + m)\psi + \bar{\psi}(\not{D} + m)\delta\psi \quad (6.10)$$

$$= - \int d^4x \delta\bar{\psi}(\not{D} + m)\psi + (-D_\mu \bar{\psi} \gamma^\mu + m\bar{\psi})\delta\psi \quad (6.11)$$

where

$$D_\mu \bar{\psi} = (\partial_\mu + ieA_\mu) \bar{\psi}. \quad (6.12)$$

$$(\not{D} + m)\psi = 0 : \text{ Dirac eq.} \quad (6.13)$$

$$(-D_\mu \bar{\psi} \gamma^\mu + m \bar{\psi}) = 0. \quad (6.14)$$

The second equation follows from the first one:

$$(\gamma^\mu D_\mu + m)\psi = 0 \quad (6.15)$$

$$\implies D_\mu \psi^\dagger (\gamma^\mu)^\dagger + m \psi^\dagger \quad (\because \text{taking } \dagger) = 0. \quad (6.16)$$

Recall

$$(\gamma^\mu)^\dagger = -\gamma^0 \gamma^\mu (\gamma^0)^{-1}, \quad \bar{\psi} = i \psi^\dagger \gamma^0. \quad (6.17)$$

Then

$$0 = D_\mu \psi^\dagger (\gamma^\mu)^\dagger + m \psi^\dagger = (-D_\mu \bar{\psi} \gamma^\mu + m \bar{\psi}) (\gamma^0)^{-1}. \quad (6.18)$$

Canonical momentum : I do not give a systematic discussion.

From

$$\mathcal{L} = -\bar{\psi} \gamma^0 \partial_t \psi \dots = i \psi^\dagger \partial_t \psi + \dots \quad (6.19)$$

define canonical momentum

$$\pi = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} = i \psi^\dagger \quad (6.20)$$

Regard ψ and ψ^\dagger as independent canonical variables.

Hamiltonian

$$H = \int d^3x (\pi \partial_t \psi - \mathcal{L}) \quad (6.21)$$

6.2 Anti-commutation relation

Important point : quantization is different from the usual case.

For a harmonic oscillator,

$$\text{canonical commutation relation} \implies \text{creation/annihilation operators } a, a^\dagger \quad (6.22)$$

$$[a, a^\dagger] = 1, \quad [a, a] = [a^\dagger, a^\dagger] = 0 \quad (6.23)$$

$$a|0\rangle = 0 \quad (6.24)$$

$$|n\rangle \propto (a^\dagger)^n |0\rangle, \quad n = 0, 1, 2, 3, \dots \quad (6.25)$$

$$H = \omega(a^\dagger a + \frac{1}{2}), \quad H|n\rangle = \omega(n + \frac{1}{2})|n\rangle. \quad (6.26)$$

Applied to a scalar field, it has led to Bose-Einstein statistics.

This is not the case for fermions.

Fact: Electrons obey Fermi-Dirac statistics. (6.27)

For each single particle state, there is 0 or 1 particle. (6.28)

A general theorem:

spin-statistics theorem : $\begin{cases} \text{integer spin} = \text{Bose-Einstein statistics: boson} \\ \text{half-integer spin} = \text{Fermi-Dirac statistics: fermion} \end{cases}$ (6.29)

A theoretical consequence of QFT, agreeing with experiment.

Fermi-Dirac requires anticommutation relation:

Creation and annihilation operators a, a^\dagger

$$\{a, a^\dagger\} = aa^\dagger + a^\dagger a = 1, \quad \{a, a\} = \{a^\dagger, a^\dagger\} = 0 \quad (6.30)$$

In particular,

$$(a)^2 = \frac{1}{2}\{a, a\} = 0, \quad (a^\dagger)^2 = \frac{1}{2}\{a^\dagger, a^\dagger\} = 0 \quad (6.31)$$

Only two states:

$$\begin{aligned} |0\rangle, & \quad a|0\rangle = 0 \\ |1\rangle = a^\dagger|0\rangle, & \quad a^\dagger|1\rangle = (a^\dagger)^2|0\rangle = 0 \end{aligned} \quad (6.32)$$

Interpretation: 0 or 1 particle.

Number operator

$$N = a^\dagger a \quad (6.33)$$

In fact,

$$[a^\dagger a, a] = (a^\dagger a)a - a(a^\dagger a) = a^\dagger a^2 - (aa^\dagger)a = 0 - (1 - a^\dagger a)a = -a \quad (6.34)$$

$$\implies [N, a] = -a. \quad (6.35)$$

In the same way,

$$[N, a^\dagger] = a^\dagger \quad (6.36)$$

$$N|0\rangle = a^\dagger a|0\rangle = 0, \quad (a|0\rangle = 0). \quad (6.37)$$

$$N|1\rangle = Na^\dagger|0\rangle = (a^\dagger + a^\dagger N)|0\rangle = a^\dagger|0\rangle = |1\rangle. \quad (6.38)$$

If the Hamiltonian is

$$H = \omega a^\dagger a \quad (6.39)$$

Then

$$H|0\rangle = 0, \quad H|1\rangle = \omega|1\rangle. \quad (6.40)$$

For a Dirac field ψ with canonical momentum $\pi = i\psi^\dagger$:

Define quantization by anti-commutation relations

$$\{\psi(t, \vec{x}), \pi(t, \vec{y})\} = i\delta^3(\vec{x} - \vec{y})I_4 \quad \pi = i\psi^\dagger \quad (6.41)$$

I_4 : 4×4 unit matrix. More explicitly,

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}, \quad \pi = (\pi_1, \pi_2, \pi_3, \pi_4) \quad (6.42)$$

Then

$$\{\psi_a(t, \vec{x}), \pi_b(t, \vec{y})\} = i\delta^3(\vec{x} - \vec{y})\delta_{ab} \quad (a, b = 1, 2, 3, 4) \quad (6.43)$$

We will abbreviate it by omitting even I_4 ,

$$\{\psi(t, \vec{x}), \pi(t, \vec{y})\} = i\delta^3(\vec{x} - \vec{y}), \quad \text{both sides regarded as } 4 \times 4 \text{ matrices} \quad (6.44)$$

Using $\pi = i\psi^\dagger$,

$$\{\psi(t, \vec{x}), \psi^\dagger(t, \vec{y})\} = \delta^3(\vec{x} - \vec{y}) \quad (6.45)$$

6.3 Creation and annihilation operators

Recall solutions of Dirac equations

$$e^{ip \cdot x} u_{(\vec{p}, s)}, \quad e^{-ip \cdot x} v_{(\vec{p}, s)}, \quad s = 1, 2, \quad p^0 = E_p \quad (6.46)$$

Properties :

$$(i\not{p} + m)u_{(\vec{p}, s)} = 0, \quad (-i\not{p} + m)v_{(\vec{p}, s)} = 0, \quad (6.47)$$

Four vectors

$$u_{(\vec{p}, s)}, \quad v_{(-\vec{p}, s')} \quad (s = 1, 2) \quad (6.48)$$

are orthogonal.

Expand

$$\psi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(a_{(\vec{p},s)} u_{(\vec{p},s)} \exp(ip \cdot x) + b_{(\vec{p},s)}^\dagger v_{(\vec{p},s)} \exp(-ip \cdot x) \right) \quad (6.49)$$

$a_{(\vec{p},s)}$, $b_{(\vec{p},s)}$: operators.

The adjoint on $b_{(\vec{p},s)}$ is for later convenience.

Commutation relations turn out to be

$$\{a_{(\vec{p},s)}, a_{(\vec{q},s')}^\dagger\} = \delta_{s,s'} (2\pi)^3 \delta(\vec{p} - \vec{q}) \quad \{b_{(\vec{p},s)}, b_{(\vec{q},s')}^\dagger\} = \delta_{s,s'} (2\pi)^3 \delta(\vec{p} - \vec{q}) \quad (6.50)$$

$$\text{other anticommutators} = 0 \quad (6.51)$$

Derivation:

At $t = 0$,

$$\psi(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(a_{(\vec{p},s)} u_{(\vec{p},s)} + b_{(-\vec{p},s)}^\dagger v_{(-\vec{p},s)} \right) \exp(i\vec{p} \cdot \vec{x}) \quad (6.52)$$

Recall

$$\{\psi(\vec{x}), \psi^\dagger(\vec{y})\} = \delta^3(\vec{x} - \vec{y}) \quad (6.53)$$

Using

$$\int d^3x \exp(i\vec{p} \cdot \vec{x}) = (2\pi)^3 \delta^3(\vec{p}) \quad (6.54)$$

and orthogonality

$$u_{(\vec{p},s)}^\dagger u_{(\vec{p},s')} = 2E_p \delta_{ss'}, \quad v_{(-\vec{p},s)}^\dagger v_{(-\vec{p},s')} = 2E_p \delta_{ss'}, \quad u_{(\vec{p},s)}^\dagger v_{(-\vec{p},s')} = 0. \quad (6.55)$$

then (exercise)

$$a_{(\vec{p},s)} = \frac{1}{\sqrt{2E_p}} \int d^3x u_{(\vec{p},s)}^\dagger \psi(\vec{x}) \exp(-i\vec{p} \cdot \vec{x}) \quad (6.56)$$

$$b_{(-\vec{p},s)}^\dagger = \frac{1}{\sqrt{2E_p}} \int d^3x v_{(-\vec{p},s)}^\dagger \psi(\vec{x}) \exp(-i\vec{p} \cdot \vec{x}) \quad (6.57)$$

From it,

$$\{a_{(\vec{p},s)}, a_{(\vec{q},s')}^\dagger\} \quad (6.58)$$

$$= \frac{1}{\sqrt{2E_p} \sqrt{2E_q}} \int d^3x d^3y u_{(\vec{p},s)}^\dagger \{\psi(\vec{x}), \psi(\vec{y})^\dagger\} u_{(\vec{q},s')} \exp(-i\vec{p} \cdot \vec{x} + i\vec{q} \cdot \vec{y}) \quad (6.59)$$

$$= \frac{1}{\sqrt{2E_p} \sqrt{2E_q}} \int d^3x d^3y u_{(\vec{p},s)}^\dagger \delta^3(\vec{x} - \vec{y}) u_{(\vec{q},s')} \exp(-i\vec{p} \cdot \vec{x} + i\vec{q} \cdot \vec{y}) \quad (6.60)$$

$$= \frac{1}{\sqrt{2E_p} \sqrt{2E_q}} \int d^3x u_{(\vec{p},s)}^\dagger u_{(\vec{q},s')} \exp(-i(\vec{p} - \vec{q}) \cdot \vec{x}) \quad (6.61)$$

$$= \frac{1}{2E_p} u_{(\vec{p},s)}^\dagger u_{(\vec{p},s')} (2\pi)^3 \delta^3(\vec{p} - \vec{q}) \quad (6.62)$$

$$= \delta_{s,s'} (2\pi)^3 \delta^3(\vec{p} - \vec{q}) \quad (6.63)$$

Other relations can be obtained similarly.

6.4 Hamiltonian and Fock space

Hamiltonian

$$H = \int d^3x (\pi \partial_t \psi - \mathcal{L}), \quad \pi = i\psi^\dagger \quad (6.64)$$

EOM $(\not{\partial} + m)\psi = 0$ gives

$$\mathcal{L} = -\bar{\psi}(\not{\partial} + m)\psi = 0. \quad (6.65)$$

Thus

$$H = \int d^3x i\psi^\dagger \partial_t \psi \quad (6.66)$$

$$\psi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(a_{(\vec{p},s)} u_{(\vec{p},s)} \exp(ip \cdot x) + b_{(\vec{p},s)}^\dagger v_{(\vec{p},s)} \exp(-ip \cdot x) \right) \quad (6.67)$$

Using orthogonality of $u_{(\vec{p},s)}, v_{(-\vec{p},s)}$, one get (exercise)

$$H = \int \frac{d^3p}{(2\pi)^3} \sum_s E_p \left(a_{(\vec{p},s)}^\dagger a_{(\vec{p},s)} - b_{(\vec{p},s)} b_{(\vec{p},s)}^\dagger \right) \quad (6.68)$$

$$= \int \frac{d^3p}{(2\pi)^3} \sum_s E_p \left(a_{(\vec{p},s)}^\dagger a_{(\vec{p},s)} + b_{(\vec{p},s)}^\dagger b_{(\vec{p},s)} \right) + E_{\text{vac}} \quad (6.69)$$

Here

$$E_{\text{vac}} = \int \frac{d^3p}{(2\pi)^3} \sum_s E_p (-2(2\pi)^3 \delta(0)) \quad (6.70)$$

We neglect it.

$a_{(\vec{p},s)}^\dagger, b_{(\vec{p},s)}^\dagger$ creat energies:

$$[H, a_{(\vec{p},s)}^\dagger] = E_p a_{(\vec{p},s)}^\dagger, \quad [H, b_{(\vec{p},s)}^\dagger] = E_p b_{(\vec{p},s)}^\dagger \quad (6.71)$$

Define vacuum $|\Omega\rangle$

$$a_{(\vec{p},s)} |\Omega\rangle = 0, \quad b_{(\vec{p},s)} |\Omega\rangle = 0 \quad (6.72)$$

One particle states

$$a_{(\vec{p},s)}^\dagger |\Omega\rangle \quad : \text{particle} \quad (6.73)$$

$$b_{(\vec{p},s)}^\dagger |\Omega\rangle \quad : \text{anti-particle} \quad (6.74)$$

Particle and anti-particle has the same energy, opposite charge:

$\psi \sim a + b^\dagger \implies a, b^\dagger$ has the same charge, or a^\dagger and b^\dagger has different charge.

More general states: act each of $a_{(\vec{p},s)}^\dagger, b_{(\vec{p},s)}^\dagger$ on $|\Omega\rangle$ zero or one time.
From

$$(a_{(\vec{p},s)}^\dagger)^2 = 0, \quad (b_{(\vec{p},s)}^\dagger)^2 = 0 \quad (6.75)$$

each of them cannot act twice.

Summary:

- $a_{(\vec{p},s)}^\dagger$ create a particle with momentum \vec{p} and spin s
- $b_{(\vec{p},s)}^\dagger$ create an anti-particle with momentum \vec{p} and spin s
- Each one-particle state has 0 or 1 particle: Fermi-Dirac statistics.

7 Quantization of electromagnetic field

EM field $A_\mu \rightarrow$ photon

Interaction of A_μ and ψ : quantum electrodynamics (QED).

But we only consider free A_μ .

7.1 Analytical mechanics

EM tensor

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad (7.1)$$

Gauge transformation

$$A_\mu \rightarrow A_\mu + \partial_\mu \alpha \quad \alpha : \text{arbitrary} \quad (7.2)$$

Physics invariant under it.

Lagrangian \mathcal{L} and action S

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} = \frac{1}{2}(\vec{E}^2 - \vec{B}^2) \quad (7.3)$$

$$S = \int d^4x \mathcal{L} \quad (7.4)$$

$$F_{ij} = \epsilon_{ijk}B^k, \quad F_{0i} = -E_i$$

Equations of motion: from $\delta S = 0$,

$$\delta S = -\frac{1}{4} \int d^4x 2F^{\mu\nu}(\partial_\mu \delta A_\nu - \partial_\nu \delta A_\mu) \quad (7.5)$$

$$= - \int d^4x F^{\mu\nu} \partial_\mu \delta A_\nu \quad (7.6)$$

$$= \int d^4x (\partial_\mu F^{\mu\nu}) \delta A_\nu \quad (7.7)$$

then

$$\partial_\mu F^{\mu\nu} = 0 : \text{Maxwell eq.} \quad (7.8)$$

Canonical momentum and quantization are subtle. We use ad hoc method.

Terms including time derivatives:

$$\mathcal{L} = +\frac{1}{2}(\partial_0 A_i - \partial_i A_0)^2 + \dots \quad (7.9)$$

Canonical momenta for A_i ($i = 1, 2, 3$) :

$$\pi^i = \frac{\partial \mathcal{L}}{\partial \dot{A}_i} = \partial_0 A_i - \partial_i A_0 = F_{0i} = -E_i \quad (7.10)$$

A_0 does not have time derivative. No canonical momentum for it.
Equation of motion from δA_0

$$\partial_i F^{i0} = \partial_i \pi^i = 0 \quad (7.11)$$

Hamiltonian

$$H = \int d^3x (\pi^i \dot{A}_i - \mathcal{L}) \quad (7.12)$$

$$= \int d^3x (\pi^i (\pi_i + \partial_i A_0) - \mathcal{L}) \quad (7.13)$$

$$= \int d^3x \frac{1}{2} (\vec{\pi}^2 + \vec{B}^2 - A_0 \partial_i \pi^i) \quad (7.14)$$

$$= \int d^3x \frac{1}{2} (\vec{\pi}^2 + \vec{B}^2) \quad (\because \partial_i \pi^i = 0) \quad (7.15)$$

7.2 Gauge invariance of states

Naive canonical commutation relations:

$$[A_i(\vec{x}), \pi^j(\vec{y})] = i \delta_i^j \delta^3(\vec{x} - \vec{y}) \quad (7.16)$$

However,

$$\partial_i \pi^i = 0 \quad (7.17)$$

These two are inconsistent,

$$[A_i(\vec{x}), \partial_j \pi^j(\vec{y})] = i \frac{\partial}{\partial y^i} \delta^3(\vec{x} - \vec{y}) \neq 0 \quad (7.18)$$

Interpretation?

Conceptual (not practical) understanding: wavefunction of \vec{A} ,

$$\Psi(\vec{A}) \quad (7.19)$$

This is a functional of the function \vec{A} . Canonical momentum

$$\pi^i = -i \frac{\delta}{\delta A_i} \quad \text{functional derivative} \quad (7.20)$$

Interpret $\partial_i \pi^i = 0$ as follows:

$\partial_i \pi^i$ is nonzero as an operator.

But impose that physical states satisfy

$$(\partial_i \pi^i) \Psi(\vec{A}) = 0 \quad (7.21)$$

$\partial_i \pi^i = 0$ is not an equation for the operators π^i but a condition for physical states.

More consideration:
For infinitesimal $\alpha(\vec{x})$,

$$0 = \int d^3x \alpha(\vec{x}) \partial_i \pi^i(\vec{x}) \Psi(\vec{A}) \quad (7.22)$$

$$0 = - \int d^3x \partial_i \alpha(\vec{x}) \pi^i(\vec{x}) \Psi(\vec{A}) \quad (7.23)$$

$$= i \int d^3x (\partial_i \alpha) \frac{\delta \Psi(\vec{A})}{\delta A_i} \quad (7.24)$$

$$= i \left[\Psi(\vec{A} + \vec{\partial} \alpha) - \Psi(\vec{A}) \right] \quad (7.25)$$

then

$$\Psi(\vec{A} + \vec{\partial} \alpha) = \Psi(\vec{A}) \quad (7.26)$$

Interpretation: $\Psi(\vec{A})$ is invariant under gauge transformation

$$A_i \rightarrow A_i + \partial_i \alpha \quad (7.27)$$

Physical states are gauge invariant.

7.3 Quantization

Practically, it is convenient to fix the gauge.

Fourier modes

$$A_i(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} \tilde{A}_i(\vec{p}) \exp(i\vec{p} \cdot \vec{x}) \quad (7.28)$$

Gauge transformation in momentum space:

$$\tilde{A}_i(\vec{p}) \rightarrow \tilde{A}_i(\vec{p}) + ip_i \tilde{\alpha}(\vec{p}) \quad (7.29)$$

\tilde{A}_i in the direction $\vec{p}/|\vec{p}|$ is not physical degrees of freedom.

Gauge fixing:

$$p_i \tilde{A}^i = 0 \implies \partial_i A^i = 0 : \text{Coulomb gauge} \quad (7.30)$$

Then

$$\partial_i \pi^i = 0, \quad \partial_i A^i = 0 \quad (7.31)$$

Among 3 components of A_i and π^i , only 2 perpendicular to \vec{p} are physical degrees of freedom.

Mode expansion:

$$\vec{A}_i(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} (\vec{\epsilon}_1 Q_1(\vec{p}) + \vec{\epsilon}_2 Q_2(\vec{p})) \exp(i\vec{p} \cdot \vec{x}) \quad (7.32)$$

$$\pi_i(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} (\vec{\epsilon}_1 P_1(\vec{p}) + \vec{\epsilon}_2 P_2(\vec{p})) \exp(i\vec{p} \cdot \vec{x}) \quad (7.33)$$

$\vec{\epsilon}_1, \vec{\epsilon}_2$: polarization vectors

$$\vec{\epsilon}_k \cdot \vec{\epsilon}_\ell = \delta_{k,\ell}, \quad \vec{p} \cdot \vec{\epsilon}_k = 0 \quad (k, \ell = 1, 2) \quad (7.34)$$

Hamiltonian

$$H = \int d^3x \frac{1}{2} (\vec{\pi}^2 + \vec{B}^2) \quad (7.35)$$

$$= \int d^3x \left(\frac{1}{2} \vec{\pi}^2 + \frac{1}{4} (\partial_i A_j - \partial_j A_i)^2 \right) \quad (7.36)$$

$$= \frac{1}{2} \int d^3x (\vec{\pi}^2 + (\partial_i A_j)^2 - \partial_i A_j \partial_j A_i) \quad (7.37)$$

$$= \frac{1}{2} \int d^3x (\vec{\pi}^2 + (\partial_i A_j)^2) \quad (\text{integration by parts, } \partial_i A^i = 0) \quad (7.38)$$

$$= \dots \text{similar computation to scalar fields} \quad (7.39)$$

$$= \frac{1}{2} \int \frac{d^3p}{(2\pi)^3} (|P_1(\vec{p})|^2 + |P_2(\vec{p})|^2 + \vec{p}^2 (|Q_1(\vec{p})|^2 + |Q_2(\vec{p})|^2)) \quad (7.40)$$

Each pair $(Q_1, P_1), (Q_2, P_2)$ behaves similarly to a real scalar field with $m = 0$.

The rest is similar to real scalar.

Results:

Commutation relations

$$[Q_k(\vec{p}), P_\ell(\vec{q})] = i\delta_{k,\ell} (2\pi)^3 \delta(\vec{p} - \vec{q}) \quad (7.41)$$

Creation, annihilation operators

$$a_{(\vec{p},k)} = \frac{1}{\sqrt{2E_p}} (E_p Q_k(\vec{p}) + iP_k(\vec{p})) \quad (7.42)$$

$$a_{(\vec{p},k)}^\dagger = \frac{1}{\sqrt{2E_p}} (E_p Q_k(-\vec{p}) - iP_k(-\vec{p})) \quad (7.43)$$

$$[a_{(\vec{p},k)}, a_{(\vec{q},\ell)}^\dagger] = \delta_{k,\ell} (2\pi)^3 \delta(\vec{p} - \vec{q}) \quad (7.44)$$

Hamiltonian

$$H = \sum_k \int \frac{d^3p}{(2\pi)^3} E_p \left(a_{(\vec{p},k)}^\dagger a_{(\vec{p},k)} \right) + E_{\text{vac}} \quad (7.45)$$

$$E_p = |\vec{p}|, \quad E_{\text{vac}} = \sum_{\vec{p}} E_p \quad (7.46)$$

Interpretation:

- $a_{(\vec{p},k)}^\dagger$, $(k = 1, 2)$ create two polarization photons with $p^\mu = (|\vec{p}|, \vec{p})$.
- Bose-Einstein statistics

$$\vec{A}(t, \vec{x}) = \sum_k \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \vec{\epsilon}_k \left(a_{(\vec{p},k)} \exp(ip \cdot x) + a_{(\vec{p},k)}^\dagger \exp(-ip \cdot x) \right) \quad (7.47)$$