

$$\rightarrow dn = \frac{d^3 p}{(2\pi)^3} \cdot V$$

number of states increases with Volume ...

Number of states in phase space volume

$$= 4\pi p^2 dp \cdot \frac{V}{(2\pi)^3}$$

But: larger volume also needs different normalization of wavefunctions to guarantee limitarity of probability density ...

$$\frac{dn}{dp} = \frac{4\pi p^2}{(2\pi)^3} \cdot V$$

Wavefunction normalization:

$$\psi(x,t) = A e^{i(\vec{p}\vec{x} - Et)} \quad \int \psi^* \psi d\vec{x} = \infty$$

Normalization within volume V:

$$\int_V \psi^* \psi d\vec{x} = \mathcal{S} = 1 \rightarrow A^2 = \frac{1}{V}$$

about Siegel & P. Sillars / Fortgould.

Normalization compensates factor V in phase space as it appears in $|\psi_i|^2 \sim (1/V)^n$; [n: # particles]

One V per final state particle ... in p-space ... in $|\psi_i|^2$... one extra $1/V$ remains from initial state ... for several ... conserved through flux ...

Thus we can choose $V=1$:

$$\frac{dn}{dp} = \frac{4\pi p^2}{(2\pi)^3} \quad \text{and} \quad \mathcal{S}(E) = \left| \frac{dn}{dE} \right|_E = \frac{dn}{dp} \cdot \left| \frac{dp}{dE} \right|_{E_i} = \frac{1}{R}$$

with $E^2 = p^2 + m^2$
 $2E dE = 2p dp$
 $dE = \frac{p}{E} dp$

We use $dn \dots$

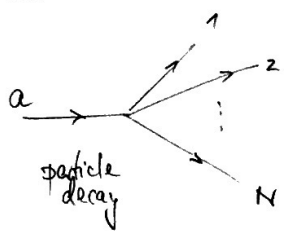
For N-particle final state; N-1 indep. momenta ...

Remember:

$$\Gamma_{fi} = 2\pi \int |\mathcal{M}_{fi}|^2 \delta(E_f - E_i) d\Omega$$

$$dn = \prod_{i=1}^{N-1} dn_i = \prod_{i=1}^{N-1} \frac{d^3 p_i}{(2\pi)^3}$$

$$= (2\pi)^3 \cdot \prod_{i=1}^N \frac{d^3 p_i}{(2\pi)^3} \cdot \delta^3(\vec{p}_a - \sum_{i=1}^N \vec{p}_i) \quad (*)$$



Lorentz-invariant phase-space:

More expression is not Lorentz-invariant as volume changes by factor $1/\gamma = 1/E$; even though normalization of wavefunction would cancel V. More convenient to choose Lorentz-invariant formulation ...

need to use $1/E$...

Choose: $\int_V \psi^* \psi d^3 x = 2E$

which means that extra factors of $2E$ enter Matrix element ... $T_{fi} \rightarrow M_{fi}$

Lorentz invariant normalization of wavefunction $A^2 = 1/E$

$$\rightarrow dLIPS = \prod_{i=1}^N \frac{d^3 p_i}{(2\pi)^3 \cdot 2E_i} = \frac{1}{(2\pi)^3} \dots$$

Lorentz-inv.

$$\Gamma_{if} = \frac{(2\pi)^4}{2E_a} \int |\mathcal{M}_{fi}|^2 \delta(E_a - E_1 - E_2) \delta^3(\vec{p}_a - \vec{p}_1 - \vec{p}_2) \frac{d^3 p_1}{(2\pi)^3 2E_1} \frac{d^3 p_2}{(2\pi)^3 2E_2}$$



From: $M_{fi} = \langle u_1 u_2 | \hat{H} | u_a \rangle \sim \sqrt{2E_1} \sqrt{2E_2} \sqrt{2E_a}$

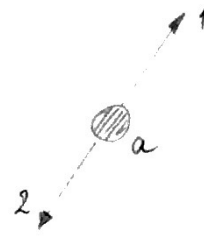
from M_{fi}^2 normalization of initial state Lorentz invariant

With $|\mathcal{M}_{fi}|^2 = 2E_a 2E_1 2E_2 \cdot |T_{fi}|^2$

Two-body decay:

[phase space!]

$$E_a = m_a, \vec{p}_a = 0; \vec{p}_1 = \vec{p}^* = -\vec{p}_2$$



Process:
 $a \rightarrow 1 + 2$
CM-Frame

$$\Gamma_{fi} = \frac{1}{8\pi^2 m_a} \int |M_{fi}|^2 \delta(m_a - E_1 - E_2) \delta^3(\vec{p}_1 + \vec{p}_2) \frac{d^3\vec{p}_1}{2E_1} \frac{d^3\vec{p}_2}{2E_2}$$

Canceling all (2π) 's...
Matrix element (see later)
Phase space

Using: $E_2^2 = m_2^2 + \vec{p}_1^2$, $d^3\vec{p}_1 = p_1^2 dp_1 \sin\theta d\theta d\phi = p_1^2 dp_1 d\Omega$
 $E_1^2 = m_1^2 + \vec{p}_1^2$, $\vec{p}_1 = -\vec{p}_2$...
 ... and properties of the Dirac delta-functions ...

yields:

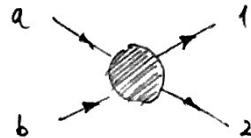
$$\Gamma_{fi} = \frac{p^*}{32\pi^2 m_a^2} \int |M_{fi}|^2 d\Omega$$

Valid for
all two-body decays

Requires:
 $\delta(f(x)) = \left| \frac{df}{dx} \right|^{-1} \delta(x-x_0)$
[see H. Thoenen]
[Hence work!]

Remark: This is the decay rate in the rest frame of particle a.
This is generally quoted as decay width or $\Gamma = 1/\tau$ lifetime.
 Γ_{fi} in its original definition is however not Lorentz invariant
as Lorentz transform changes $m_a \mapsto E_a$...

Cross Section and Lorentz-invariant flux:



$$\sigma = \frac{\Gamma_{fi}}{(v_a + v_b)} = \frac{(2\pi)^4}{4E_a E_b (v_a + v_b)} \int |M_{fi}|^2 \delta^4(p_a + p_b - p_1 - p_2) \frac{d^3\vec{p}_1}{(2\pi)^3 2E_1} \frac{d^3\vec{p}_2}{(2\pi)^3 2E_2}$$

Lorentz-inv. Flux
Lorentz-invariant Matrix element
Lorentz invariant

Γ Lorentz-inv. flux:

$$F = 4E_a E_b (v_a + v_b) = 4E_a E_b \left(\frac{|\vec{p}_a|}{E_a} - \frac{|\vec{p}_b|}{E_b} \right) = 4(E_a |\vec{p}_b| + E_b |\vec{p}_a|) \quad \left[\text{since } \beta = \frac{v}{c} \right]$$

$$F^2 = 16(E_a^2 \vec{p}_b^2 + E_b^2 \vec{p}_a^2 + 2E_a E_b |\vec{p}_a| |\vec{p}_b|)$$

$$4\text{-vectors: } (p_a \cdot p_b)^2 = (E_a E_b + |\vec{p}_a| |\vec{p}_b|)^2 = E_a^2 E_b^2 + \vec{p}_a^2 \vec{p}_b^2 + 2E_a E_b |\vec{p}_a| |\vec{p}_b|$$

$$\rightarrow F^2 = 16[(p_a \cdot p_b)^2 - (E_a^2 - \vec{p}_a^2)(E_b^2 - \vec{p}_b^2)] \rightarrow F = 4[(p_a \cdot p_b)^2 - m_a^2 m_b^2]^{1/2}$$

$$E_a^2 E_b^2 + 2E_a E_b |\vec{p}_a| |\vec{p}_b| + E_a^2 \vec{p}_b^2 + E_b^2 \vec{p}_a^2 - E_a^2 E_b^2 - \vec{p}_a^2 \vec{p}_b^2 + E_a^2 \vec{p}_b^2 + E_b^2 \vec{p}_a^2$$

4-vector product squared!!
masses

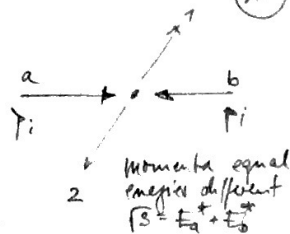
Lorentz invariant.

Scattering in CM-frame:

$$4(E_a \vec{p}_1 + E_b \vec{p}_1)$$

(11)

$$F = 4 E_b^* E_a^* (v_a^* + v_b^*) = 4 p_i^* (E_a^* + E_b^*) = 4 p_i^* \sqrt{s}$$



$$\sigma = \frac{1}{(2\pi)^2} \cdot \frac{1}{4 p_i^* \sqrt{s}} \cdot \int |M_{fi}|^2 \delta(\sqrt{s} - E_1 - E_2) \delta^3(\vec{p}_1 + \vec{p}_2) \frac{d^3 p_1}{2E_1} \frac{d^3 p_2}{2E_2}$$

$$\left(\frac{(2\pi)^4}{(2\pi)^6} \right)$$

$$= \frac{1}{16\pi^2 p_i^* \sqrt{s}} \cdot \frac{p_f^*}{4\sqrt{s}} \int |M_{fi}|^2 d\Omega^*$$

needs again:
 $E_2^2 = m_2^2 + p_1^2$, $d^3 p_1 = p_1^2 dp_1 \sin\theta d\theta d\phi$
 $= p_1^2 dp_1 d\Omega$

[Homework]

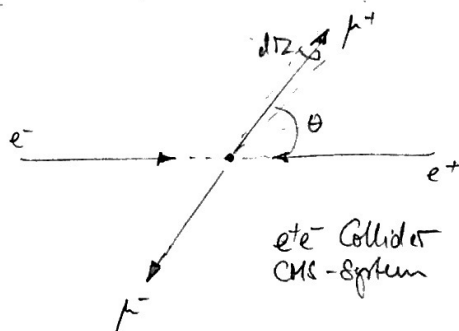
$$\sigma = \frac{1}{64\pi^2 s} \cdot \frac{p_f^*}{p_i^*} \int |M_{fi}|^2 d\Omega^*$$

Total Cross Section for $2 \rightarrow 2$ process.

Differential Cross Section:

$$\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2} \cdot \frac{p_f^*}{p_i^*} |M_{fi}|^2$$

Also: $\frac{d\sigma}{dE}, \frac{d\sigma}{dE d\Omega} \dots$



More complicated if to be calculated for fixed target...
 [if interested see U. Thomson 3.5]

IV. Dirac Equation

Reminder: Description of free particles...

A. Schrödinger Equation.

$$\left[i\frac{\partial}{\partial t} \psi = -\frac{1}{2m} \vec{\nabla}^2 \psi \right] \text{ describes free non-relativistic particles ...}$$

Remember:
 $\hbar = 1$

Solution for energy $E = \vec{p}^2/2m$:

$$\psi(\vec{r}, t) = \frac{1}{V} e^{i(\vec{p}\vec{r} - Et)}$$

ψ^2 = probability density

$$\vec{p} = \hbar \vec{k}$$

$$E = \hbar \omega$$

The Schrödinger equation uses the classical E - p relation $E = \vec{p}^2/2m$
 replacing $E \rightarrow i\frac{\partial}{\partial t}$, $\vec{p} \rightarrow -i\vec{\nabla}$

Continuity equation:

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

with $\rho = \psi^* \psi$, $\vec{j} = -\frac{i}{2m} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*)$

ρ : prob. density
 \vec{j} : current density

Proof: A: $i\hbar \left(\psi + \frac{1}{2m} \nabla^2 \psi \right) \cdot (-i\psi^*) \rightarrow \psi^* \psi - \frac{1}{2m} \psi^* \nabla^2 \psi = 0$
 B: $-i\hbar \left(\psi^* + \frac{1}{2m} \nabla^2 \psi^* \right) \cdot (i\psi) \rightarrow -\psi^* \psi - \frac{1}{2m} \psi^* \nabla^2 \psi^* = 0$
 A - B: $\psi^* \psi + \psi^* \psi - \frac{1}{2m} (\psi^* \nabla^2 \psi - \psi \nabla^2 \psi^*) = 0$
 $\rightarrow \dot{\rho} + \vec{\nabla} \cdot \vec{j} = 0$

B. Klein - Gordon Equation

$$\frac{\partial^2}{\partial t^2} \phi - \nabla^2 \phi + m^2 \phi = 0$$

describes free relativistic spin-0 quantum fields
not free particles as we will see.

The Klein-Gordon equation uses $E^2 = p^2 + m^2$
replacing again $E \rightarrow i\hbar \frac{\partial}{\partial t}$, $\vec{p} \rightarrow -i\hbar \vec{\nabla}$.

Lorentz-invariant form
using co- and contravariant derivatives:

$$(\partial_\mu \partial^\mu + m^2) \phi = 0$$

with $\partial_\mu \partial^\mu = \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2}$

Plane-wave solution:

$\phi = N \cdot e^{i(\vec{p} \cdot \vec{x} - E_\pm t)}$ with $E_\pm = \pm \sqrt{p^2 + m^2}$

$E < 0$ possible

$E < 0$: unphysical?

Would be easy way out; but: negative energy values cannot be ignored as the solution would be incomplete without it.

need complete system of eigenfunctions

More problematic ...

Continuity Equation:

$\nabla^2 \phi - \frac{\partial^2}{\partial t^2} \phi - m^2 \phi = 0$ $\quad | \cdot (-i\phi^*)$ (from left?)
 + $\nabla^2 \phi^* - \frac{\partial^2}{\partial t^2} \phi^* - m^2 \phi^* = 0$ $\quad | \cdot (i\phi)$ (from right?)
 $\frac{\partial}{\partial t} [i\phi^* \dot{\phi} - i\dot{\phi} \phi^*] + \vec{\nabla} \cdot [-i\phi^* \vec{\nabla} \phi - \phi \vec{\nabla} \phi^*] = 0$
 $= \dot{\rho}$ $= \vec{\nabla} \cdot \vec{j}$ see eg. Heisenberg/Thirring