

Spacecraft Guidance

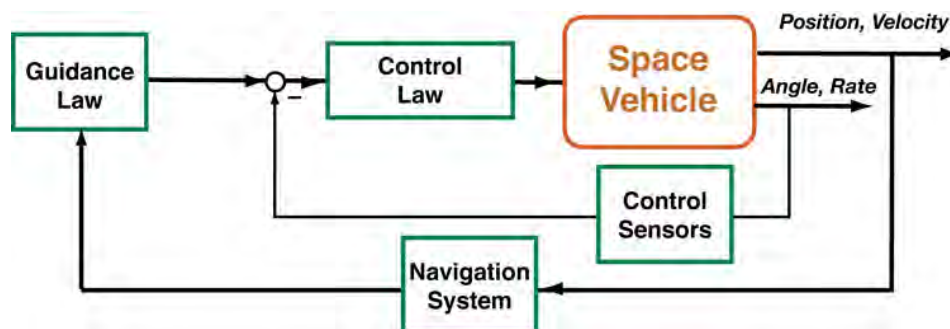
Space System Design, MAE 342, Princeton University
Robert Stengel

- Oberth's "Synergy Curve"
- Explicit ascent guidance
- Impulsive ΔV maneuvers
- Hohmann transfer between circular orbits
- Sphere of gravitational influence
- Synodic periods and launch windows
- Hyperbolic orbits and escape trajectories
- Battin's universal formulas
- Lambert's time-of-flight theorem (hyperbolic orbit)
- Fly-by (swingby) trajectories for gravity assist

Copyright 2016 by Robert Stengel. All rights reserved. For educational use only.
<http://www.princeton.edu/~stengel/MAE342.html>

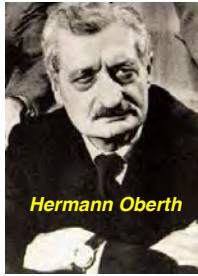
1

Guidance, Navigation, and Control



- **Navigation:** Where are we?
- **Guidance:** How do we get to our destination?
- **Control:** What do we tell our vehicle to do?

2



Hermann Oberth

Energy Gained from Propellant

Specific energy = energy per unit weight

$$\mathbb{E} = h + \frac{V^2}{2g}$$

h : height; V : velocity

Rate of change of specific energy per unit of expended propellant mass

$$\begin{aligned} \frac{d\mathbb{E}}{dm} &= \frac{dh}{dm} + \frac{V}{g} \frac{dV}{dm} = \frac{1}{(dm/dt)} \left(\frac{dh}{dt} + \frac{V}{g} \frac{dV}{dt} \right) \\ &= \frac{1}{(dm/dt)} \left(\frac{dh}{dt} + \frac{1}{g} \mathbf{v}^T \frac{d\mathbf{v}}{dt} \right) = \frac{1}{(dm/dt)} \left(V \sin \gamma + \frac{1}{g} \mathbf{v}^T (\mathbf{T} - m\mathbf{g}) \right) \\ &= \frac{1}{(dm/dt)} \left(V \sin \gamma + \frac{VT}{mg} \cos \alpha - V \sin \gamma \right) \end{aligned}$$

3

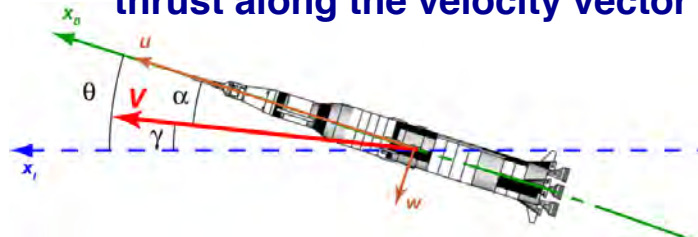
Oberth's Synergy Curve

γ : Flight Path Angle

θ : Pitch Angle

α : Angle of Attack

$d\mathbb{E}/dm$ maximized when $\alpha = 0$, or $\theta = \gamma$, i.e., thrust along the velocity vector



Approximate round-earth equations of motion

$$\begin{aligned} \frac{dV}{dt} &= \frac{T}{m} \cos \alpha - \frac{Drag}{m} - g \sin \gamma \\ \frac{d\gamma}{dt} &= \frac{T}{mV} \sin \alpha + \left(\frac{V}{r} - \frac{g}{V} \right) \cos \gamma \end{aligned}$$

4

Gravity-Turn Pitch Program

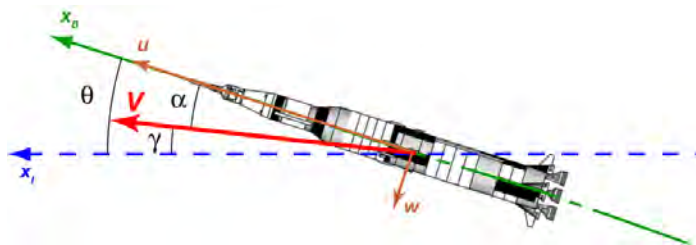
With angle of attack, $\alpha = 0$

$$\frac{d\gamma}{dt} = \frac{d\theta}{dt} = \left(\frac{V}{r} - \frac{g}{V} \right) \cos \gamma$$

Locally optimal flight path

Minimizes aerodynamic loads

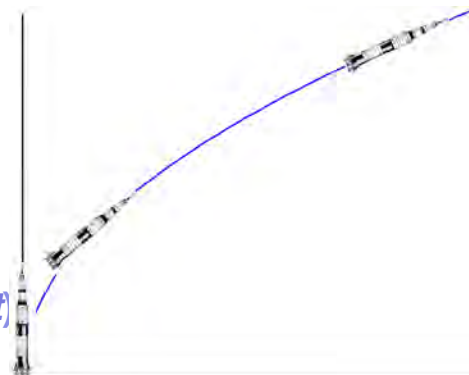
Feedback controller minimizes α or load factor



5

Gravity-Turn Flight Path

- Gravity-turn flight path is function of 3 variables
 - Initial pitchover angle (from vertical launch)
 - Velocity at pitchover
 - Acceleration profile, $T(t)/m(t)$

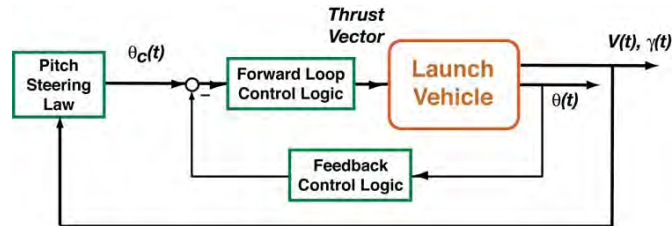


Gravity-turn program closely approximated by tangent steering laws (*see Supplemental Material*)

6

Feedback Control Law

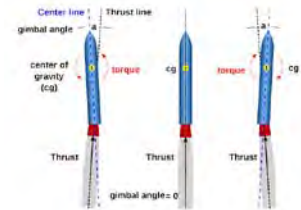
Errors due to disturbances and modeling errors corrected by feedback control



Motor Gimbal Angle $(t) \triangleq \delta_G(t) = c_\theta [\theta_{des}(t) - \theta(t)] - c_q q(t)$

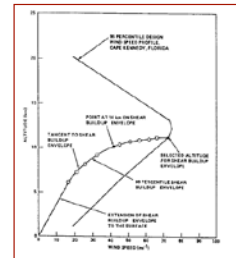
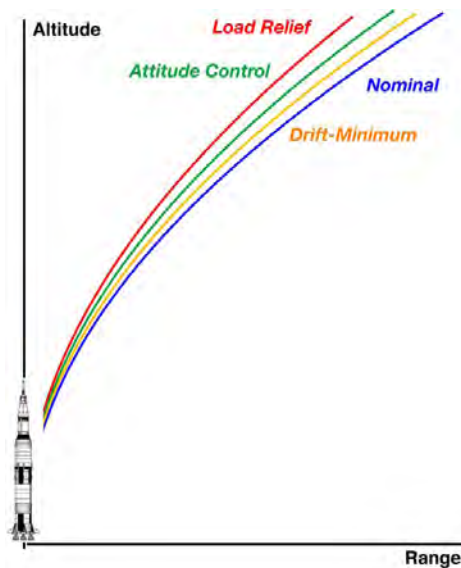
θ_{des} = Desired pitch angle; $q = \frac{d\theta}{dt}$ = pitch rate

c_θ, c_q : Feedback control law gains



7

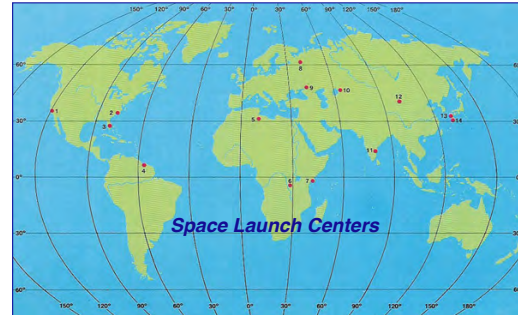
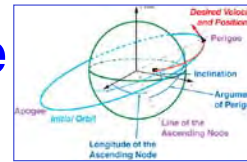
Thrust Vector Control During Launch Through Wind Profile



- **Attitude control**
 - Attitude and rate feedback
- **Drift-minimum control**
 - Attitude and accelerometer feedback
 - Increased loads
- **Load relief control**
 - Rate and accelerometer feedback
 - Increased drift

8

Effect of Launch Latitude on Orbital Parameters



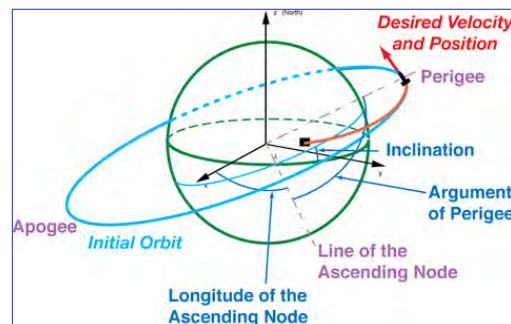
- Launch latitude establishes minimum orbital inclination (without “dogleg” maneuver)
- Time of launch establishes line of nodes
- Argument of perigee established by
 - Launch trajectory
 - On-orbit adjustment

9

Guidance Law for Launch to Orbit

(Brand, Brown, Higgins, and Pu, CSDL, 1972)

- Initial conditions
 - End of pitch program, outside atmosphere
- Final condition
 - Insertion in desired orbit
- Initial inputs
 - Desired radius
 - Desired velocity magnitude
 - Desired flight path angle
 - Desired inclination angle
 - Desired longitude of the ascending/ descending node
- Continuing outputs
 - Unit vector describing desired thrust direction
 - Throttle setting, % of maximum thrust

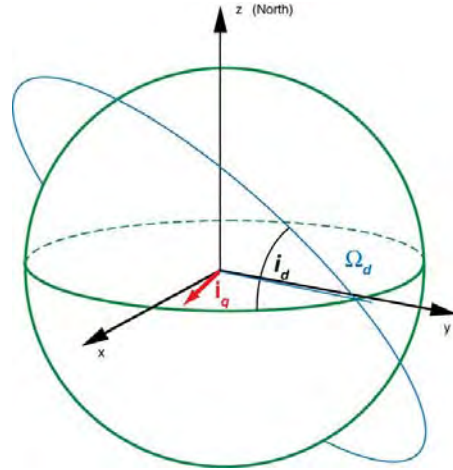


10

Guidance Program Initialization

- Thrust acceleration estimate
- Mass/mass flow rate
- Acceleration limit ($\sim 3g$)
- Effective exhaust velocity
- Various coefficients
- **Unit vector normal to desired orbital plane, \mathbf{i}_q**

$$\mathbf{i}_q = \begin{bmatrix} \sin i_d \sin \Omega_d \\ \sin i_d \cos \Omega_d \\ \cos i_d \end{bmatrix}$$



i_d : Desired inclination angle of final orbit
 Ω_d : Desired longitude of descending node

11

Guidance Program Operation: Position and Velocity

- Obtain thrust acceleration estimate, \mathbf{a}_T , from guidance system
- Compute corresponding mass, mass flow rate, and throttle setting, δT

$\mathbf{i}_r = \frac{\mathbf{r}}{r}$: Unit vector aligned with local vertical
 $\mathbf{i}_z = \mathbf{i}_r \times \mathbf{i}_q$: Downrange direction

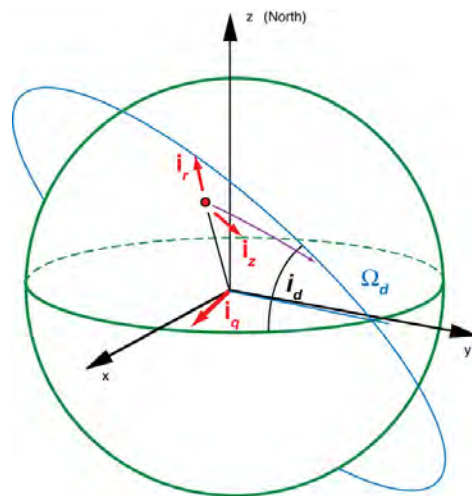
Position

$$\begin{bmatrix} r \\ y \\ z \end{bmatrix} = \begin{bmatrix} |\mathbf{r}| \\ r \sin^{-1}(\mathbf{i}_r \bullet \mathbf{i}_q) \\ open \end{bmatrix}$$

Velocity

$$\begin{bmatrix} \dot{r} \\ \dot{y} \\ \dot{z} \end{bmatrix} = \begin{bmatrix} \mathbf{v}_{IMU} \bullet \mathbf{i}_r \\ \mathbf{v}_{IMU} \bullet \mathbf{i}_q \\ \mathbf{v}_{IMU} \bullet \mathbf{i}_z \end{bmatrix}$$

\mathbf{v}_{IMU} : Velocity estimate in IMU frame



12

Guidance Program: Velocity and Time to Go

Effective gravitational acceleration

$$g_{eff} = -\frac{\mu}{r^2} + \frac{|\mathbf{r} \times \mathbf{v}|^2}{r^3}$$

Time to go (to motor burnout)

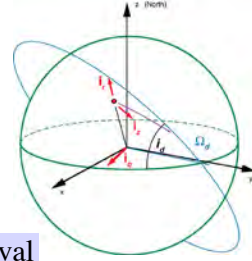
$$t_{go_{new}} = t_{go_{old}} - \Delta t \quad \Delta t : \text{Guidance command interval}$$

Velocity to be gained

$$\mathbf{v}_{go} = \begin{bmatrix} (\dot{r}_d - \dot{r}) - g_{eff} t_{go} / 2 \\ -\dot{y} \\ \dot{z}_d - \dot{z} \end{bmatrix}$$

Time to go prediction (prior to acceleration limiting)

$$t_{go} = \frac{m}{\dot{m}} \left(1 - e^{-v_{go}/c_{eff}} \right) \quad c_{eff} : \text{Effective exhaust velocity}$$



13

Guidance Program Commands

Guidance law: required radial and cross-range accelerations

$$\begin{aligned} a_{T_r} &= a_T [A + B(t - t_o)] - g_{eff} \\ a_{T_y} &= a_T [C + D(t - t_o)] \end{aligned} \quad a_T = \text{Net available acceleration}$$

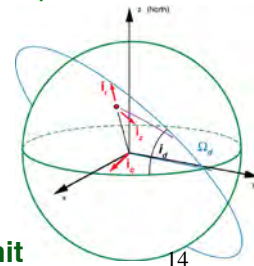
Guidance coefficients, **A**, **B**, **C**, and **D** are functions of

$$\begin{aligned} &\left(r_d, r, \dot{r}, t_{go} \right) \\ &\left(y, \dot{y}, t_{go} \right) \end{aligned} \quad \text{plus } c_{eff}, m/\dot{m}, \text{ Acceleration limit}$$

Required thrust direction, **i_T** (i.e., vehicle orientation in (**i_r**, **i_q**, **i_z**) frame

$$\mathbf{a}_T = \begin{bmatrix} a_{T_r} \\ a_{T_y} \\ \text{what's left over} \end{bmatrix}; \quad i_T = \frac{a_T}{|\mathbf{a}_T|}$$

Throttle command is a function of **a_T** (i.e., acceleration magnitude) and acceleration limit

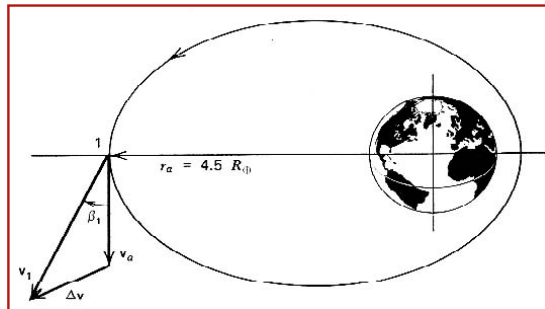


14

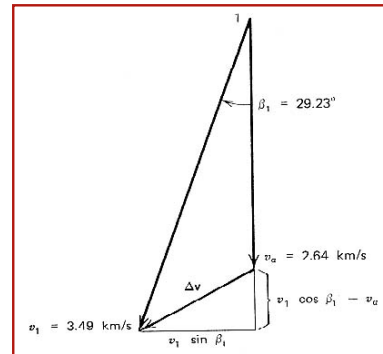
Impulsive ΔV Orbital Maneuver

- If rocket burn time is short compared to orbital period (e.g., seconds compared to hours), **impulsive ΔV approximation** can be made
 - Change in position during burn is \sim zero
 - Change in velocity is \sim instantaneous

Velocity impulse at apogee



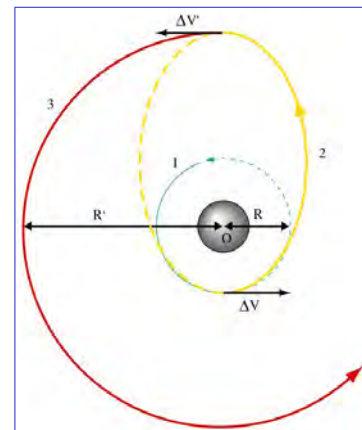
Vector diagram of velocity change



15

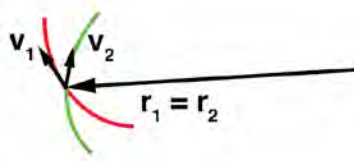
Orbit Change due to Impulsive ΔV

- Maximum energy change accrues when ΔV is aligned with the instantaneous orbital velocity vector
 - Energy change \rightarrow Semi-major axis change
 - Maneuver at perigee raises or lowers apogee
 - Maneuver at apogee raises or lowers perigee
- Optimal transfer from one circular orbit to another involves two impulses [*Hohmann transfer*]
- Other maneuvers
 - In-plane parameter change
 - Orbital plane change



16

Assumptions for Impulsive Maneuver



Instantaneous change in velocity vector

$$\mathbf{v}_2 = \mathbf{v}_1 + \Delta \mathbf{v}_{rocket}$$

Negligible change in radius vector

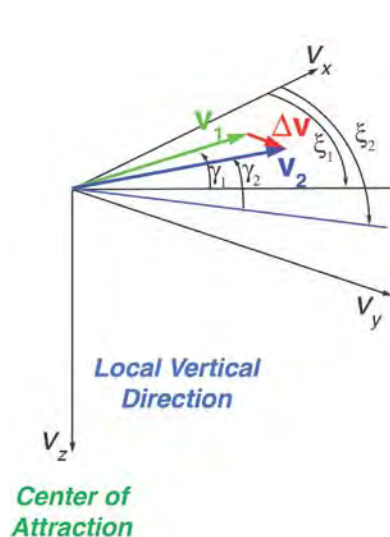
$$\mathbf{r}_2 = \mathbf{r}_1$$

Therefore, new orbit intersects old orbit
Velocities different at the intersection

17

Geometry of Impulsive Maneuver

Change in velocity magnitude, $|\mathbf{v}|$, vertical flight path angle, γ , and horizontal flight path angle, ξ



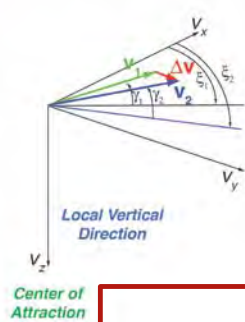
$$\mathbf{v}_1 = \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix}_1 = \begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix}_1 = \begin{bmatrix} V \cos \gamma \cos \xi \\ V \cos \gamma \sin \xi \\ -V \sin \gamma \end{bmatrix}_1$$

$$\mathbf{v}_2 = \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix}_2 = \begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix}_2 = \begin{bmatrix} V \cos \gamma \cos \xi \\ V \cos \gamma \sin \xi \\ -V \sin \gamma \end{bmatrix}_2$$

$$\Delta \mathbf{v} = \begin{bmatrix} \Delta v_x \\ \Delta v_y \\ \Delta v_z \end{bmatrix} = \begin{bmatrix} (v_{x_2} - v_{x_1}) \\ (v_{y_2} - v_{y_1}) \\ (v_{z_2} - v_{z_1}) \end{bmatrix}$$

18

Required $\Delta \mathbf{v}$ for Impulsive Maneuver



$$\Delta \mathbf{v} = \begin{bmatrix} \Delta v_x \\ \Delta v_y \\ \Delta v_z \end{bmatrix} = \begin{bmatrix} v_{x_2} - v_{x_1} \\ v_{y_2} - v_{y_1} \\ v_{z_2} - v_{z_1} \end{bmatrix}$$

$$\Delta \mathbf{v}_{rocket} = \begin{bmatrix} \Delta V_{rocket} \\ \xi_{rocket} \\ \gamma_{rocket} \end{bmatrix} = \begin{bmatrix} \left(\Delta v_x^2 + \Delta v_y^2 + \Delta v_z^2 \right)^{1/2}_{rocket} \\ \sin^{-1} \left(\frac{\Delta v_y}{\left(\Delta v_x^2 + \Delta v_y^2 \right)^{1/2}_{rocket}} \right) \\ \sin^{-1} \left(\frac{\Delta v_z}{\Delta V}_{rocket} \right) \end{bmatrix}$$

19

Single Impulse Orbit Adjustment Coplanar (i.e., in-plane) maneuvers

- Change energy
- Change angular momentum
- Change eccentricity

$$\begin{aligned} \mathbb{E} &= \frac{1}{2} v^2 - \mu/r = (e^2 - 1) \mu^2 / h^2 \\ h &= \sqrt{\frac{\mu^2 (e^2 - 1)}{\mathbb{E}}} = \sqrt{\frac{\mu^2 (e^2 - 1)}{v^2/2 - \mu/r}} \\ e &= \sqrt{1 + 2 \mathbb{E} h^2 / \mu^2} \end{aligned}$$

- Required velocity increment

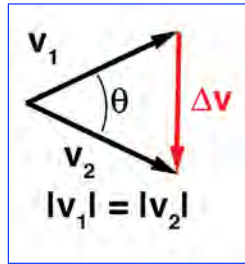
$$\begin{aligned} v_{new} &\triangleq v_{old} + \Delta \mathbf{v}_{rocket} = \sqrt{2(\mathbb{E}_{new} + \mu/r)} \\ &= \sqrt{2[(e_{new}^2 - 1) \mu^2 / h_{new}^2 + \mu/r]} \\ \Delta \mathbf{v}_{rocket} &= v_{new} - v_{old} \end{aligned}$$

20

Single Impulse Orbit Adjustment

Coplanar (i.e., in-plane) maneuvers

- Change semi-major axis
 - magnitude
 - orientation (i.e., argument of perigee); in-plane isosceles triangle



$$a_{new} = \frac{h_{new}^2 / \mu}{1 - e_{new}^2}$$

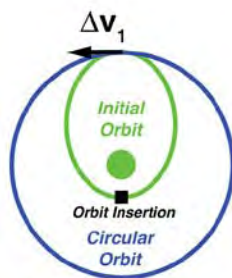
$$\begin{aligned} r_{perigee} &= a(1 - e) \\ r_{apogee} &= a(1 + e) \\ v_{perigee} &= \sqrt{\frac{\mu}{a} \left(\frac{1 + e}{1 - e} \right)} \\ v_{apogee} &= \sqrt{\frac{\mu}{a} \left(\frac{1 - e}{1 + e} \right)} \end{aligned}$$

- Change apogee or perigee
 - radius
 - velocity

21

In-Plane Orbit Circularization

Initial orbit is elliptical, with apogee radius equal to desired circular orbit radius



Initial Orbit

$$\begin{aligned} a &= (r_{cir(target)} + r_{insertion}) / 2 \\ e &= (r_{cir(target)} - r_{insertion}) / 2a \\ v_{apogee} &= \sqrt{\frac{\mu}{a} \left(\frac{1 - e}{1 + e} \right)} \end{aligned}$$

Velocity in circular orbit is a function of the radius

“Vis viva” equation:

$$v_{cir} = \sqrt{\mu \left(\frac{2}{r_{cir}} - \frac{1}{a_{cir}} \right)} = \sqrt{\mu \left(\frac{2}{r_{cir}} - \frac{1}{r_{cir}} \right)} = \sqrt{\frac{\mu}{r_{cir}}}$$

Rocket must provide the difference

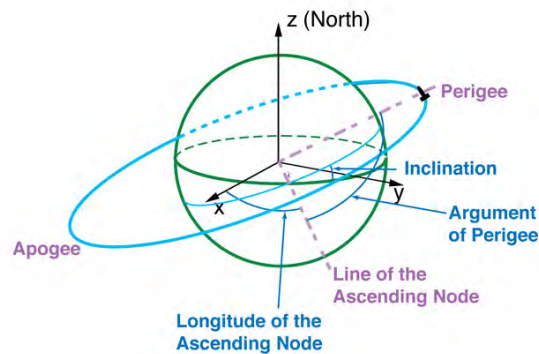
$$\Delta v_{rocket} = v_{cir} - v_{apogee}$$

22

Single Impulse Orbit Adjustment

Out-of-plane maneuvers

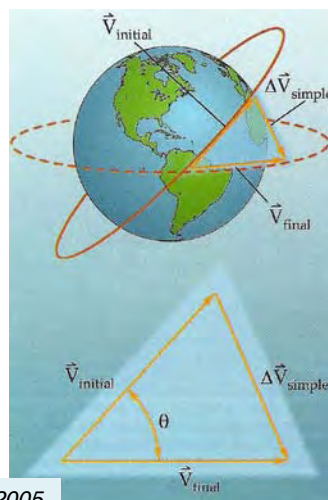
- Change orbital inclination
- Change longitude of the ascending node
- v_1 , Δv , and v_2 form **isosceles triangle** perpendicular to the orbital plane to leave in-plane parameters unchanged



23

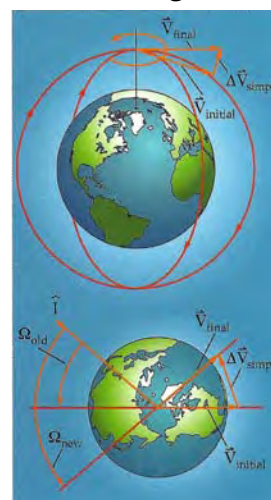
Change in Inclination and Longitude of Ascending Node

Inclination



Sellers, 2005

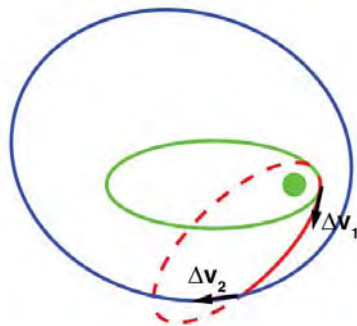
Longitude of Ascending Node



24

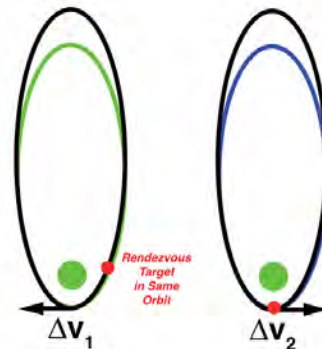
Two Impulse Maneuvers

Transfer to Non-Intersecting Orbit



- 1st Δv produces target orbit intersection
- 2nd Δv matches target orbit
- Minimize $(|\Delta v_1| + |\Delta v_2|)$ to minimize propellant use

Phasing Orbit

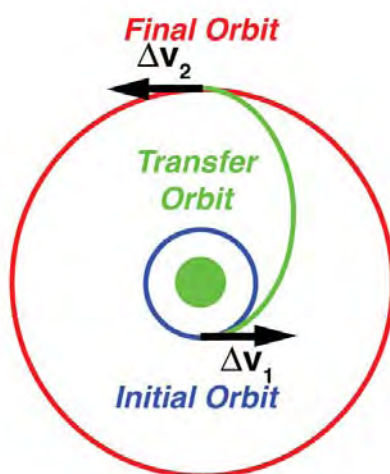


- Rendezvous with trailing spacecraft in same orbit
- At perigee, increase speed to increase orbital period
- At future perigee, decrease speed to resume original orbit

25

Hohmann Transfer between Coplanar Circular Orbits (Outward transfer example)

Thrust in direction of motion at transfer perigee and apogee



$$v_{cir_1} = \sqrt{\frac{\mu}{r_{cir_1}}}$$

$$v_{cir_2} = \sqrt{\frac{\mu}{r_{cir_2}}}$$

Transfer Orbit

$$a = (r_{cir_1} + r_{cir_2})/2$$

$$e = (r_{cir_2} - r_{cir_1})/2a$$

$$v_{p_{transfer}} = \sqrt{\frac{\mu}{a} \left(\frac{1+e}{1-e} \right)}$$

$$v_{a_{transfer}} = \sqrt{\frac{\mu}{a} \left(\frac{1-e}{1+e} \right)}$$

26

Outward Transfer Orbit Velocity Requirements

Δv at 1st Burn

$$\Delta v_1 = v_{p_{transfer}} - v_{cir_1}$$

$$= v_{cir_1} \left(\sqrt{\frac{2r_{cir_2}}{r_{cir_1} + r_{cir_2}}} - 1 \right)$$

Δv at 2nd Burn

$$\Delta v_2 = v_{cir_1} - v_{a_{transfer}}$$

$$= v_{cir_2} \left(1 - \sqrt{\frac{2r_{cir_1}}{r_{cir_1} + r_{cir_2}}} \right)$$

$$v_{cir_2} = v_{cir_1} \sqrt{\frac{r_{cir_1}}{r_{cir_2}}}$$

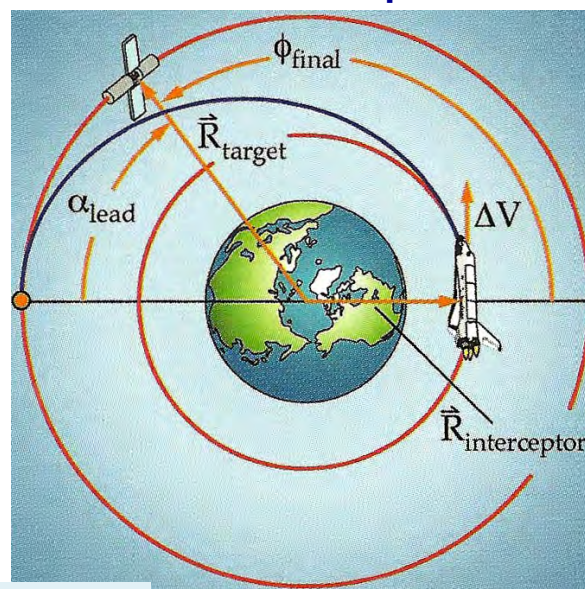
Hohmann Transfer is **energy-optimal** for 2-impulse transfer between circular orbits and $r_2/r_1 < 11.94$

$$\Delta v_{total} = v_{cir_1} \left[\sqrt{\frac{2r_{cir_2}}{r_{cir_1} + r_{cir_2}}} \left(1 - \frac{r_{cir_1}}{r_{cir_2}} \right) + \sqrt{\frac{r_{cir_1}}{r_{cir_2}}} - 1 \right]$$

27

Rendezvous Requires Phasing of the Maneuver

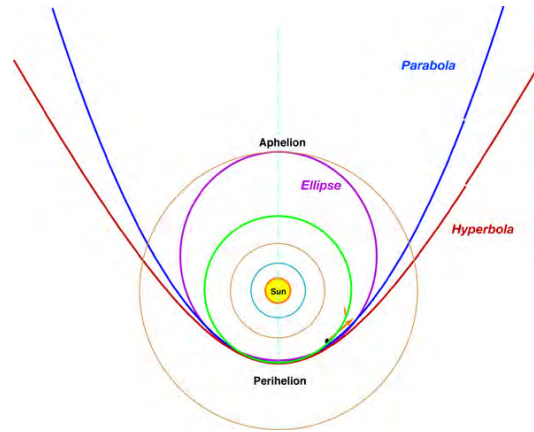
Transfer orbit time equals target's time to reach rendezvous point



Sellers, 2005

28

Solar Orbits



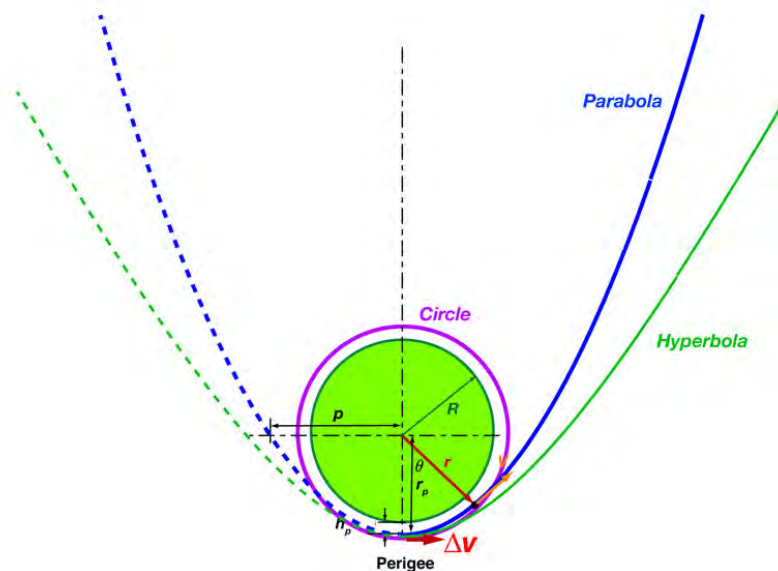
- Same equations used for Earth-referenced orbits
 - Dimensions of the orbit
 - Position and velocity of the spacecraft
 - Period of elliptical orbits
 - Different gravitational constant

$$\mu_{Sun} = 1.3327 \times 10^{11} \text{ km}^3/\text{s}^2$$

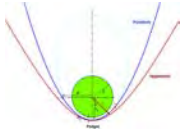
29

Escape from a Circular Orbit

Minimum escape trajectory shape is a **parabola**



30



In-plane Parameters of Earth Escape Trajectories

Dimensions of the orbit

$$p = \frac{h^2}{\mu} = \text{"The parameter" or semi-latus rectum}$$

h = Angular momentum about center of mass

$$e = \sqrt{1 + 2 \frac{\mathbb{E} p}{\mu}} = \text{Eccentricity} \geq 1$$

\mathbb{E} = Specific energy, ≥ 0

$$a = \frac{p}{1 - e^2} = \text{Semi-major axis, } < 0$$

$$r_{\text{perigee}} = a(1 - e) = \text{Perigee radius}$$

31

In-plane Parameters of Earth Escape Trajectories

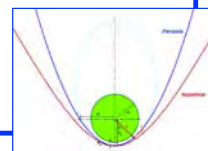
Position and velocity of the spacecraft

$$r = \frac{p}{1 + e \cos \theta} = \text{Radius of the spacecraft}$$

θ = True anomaly

$$V = \sqrt{2\mu \left(\frac{1}{r} - \frac{1}{2a} \right)} = \text{Velocity of the spacecraft}$$

$$V_{\text{perigee}} \geq \sqrt{2\mu / r_{\text{perigee}}}$$



32

Escape from Circular Orbit

Velocity in circular orbit

$$V_c = \sqrt{\mu \left(\frac{2}{r_c} - \frac{1}{r_c} \right)} = \sqrt{\frac{\mu}{r_c}}$$

Velocity at perigee of parabolic orbit

$$V_{perigee} = \sqrt{\mu \left(\frac{2}{r_c} - \frac{1}{(a \rightarrow \infty)} \right)} = \sqrt{\frac{2\mu}{r_c}}$$

Velocity increment required for escape

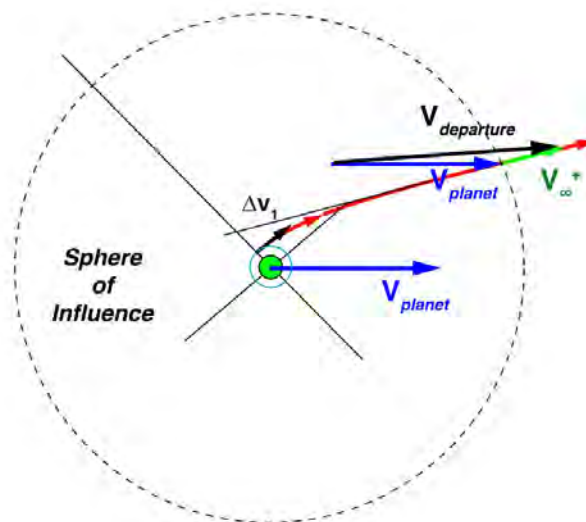
$$\Delta V_{escape} = V_{perigee_{parabola}} - V_c = \sqrt{\frac{2\mu}{r_c}} - \sqrt{\frac{\mu}{r_c}} \approx 0.414 V_c$$

33

Earth Escape Trajectory

Δv_1 to increase speed to escape velocity

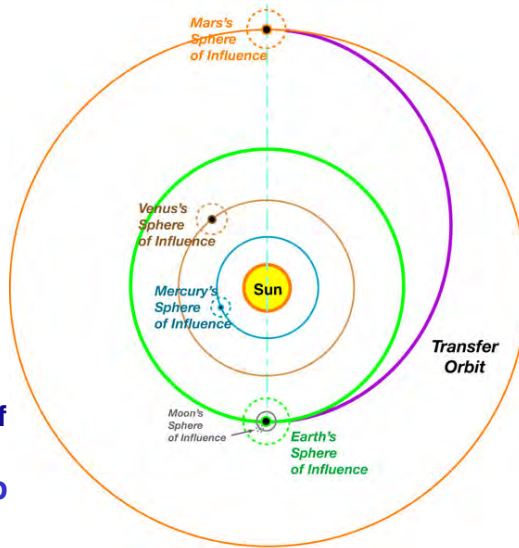
Velocity required for transfer at sphere of influence



34

Transfer Orbits and Spheres of Influence

- **Sphere of Influence (Laplace):**
 - Radius within which gravitational effects of planet are more significant than those of the Sun
- **Patched-conic section approximation**
 - Sequence of 2-body orbits
 - Outside of planet's sphere of influence, Sun is the center of attraction
 - Within planet's sphere of influence, planet is the center of attraction
- **Fly-by (swingby) trajectories dip into intermediate object's sphere of influence for gravity assist**



35

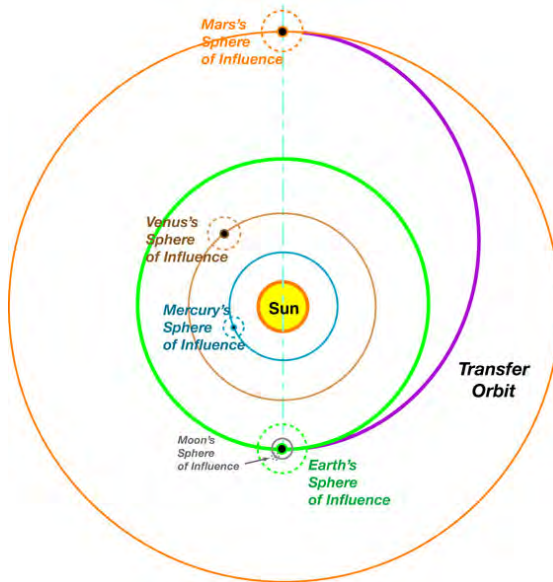
Solar System Spheres of Influence

$$\text{for } \frac{m_{\text{Planet}}}{m_{\text{Sun}}} \ll 1, \quad r_{SI} \simeq r_{\text{Planet-Sun}} \left(\frac{m_{\text{Planet}}}{m_{\text{Sun}}} \right)^{2/5}$$

Planet	Sphere of Influence, km
Mercury	112,000
Venus	616,000
Earth	929,000
Mars	578,000
Jupiter	48,200,000
Saturn	54,500,000
Uranus	51,800,000
Neptune	86,800,000
Pluto	27,000,000-45,000,000

36

Interplanetary Mission Planning



- **Example: Direct Hohmann Transfer from Earth Orbit to Mars Orbit (No fly-bys)**

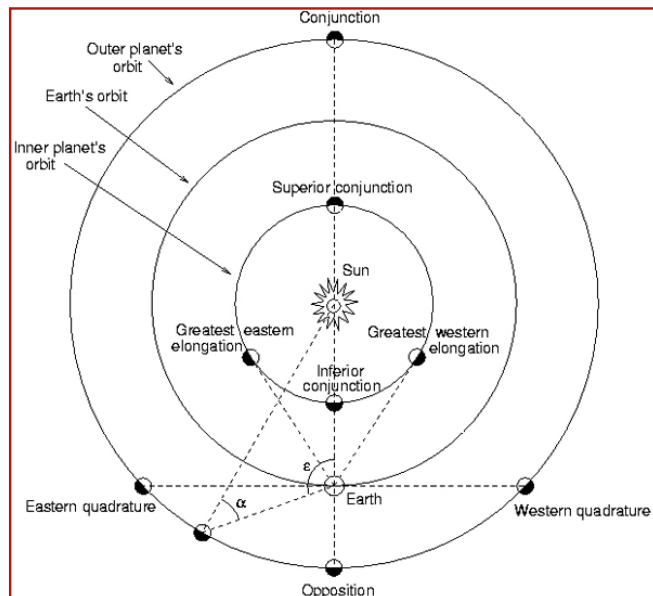
- 1) Calculate required perigee velocity for transfer orbit - Sun as center of attraction: **Elliptical orbit**
- 2) Calculate Δv required to reach Earth's sphere of influence with velocity required for transfer – Earth as center of attraction: **Hyperbolic orbit**
- 3) Calculate Δv required to enter circular orbit about Mars, given transfer apogee velocity – Mars as center of attraction: **Hyperbolic orbit**

37

Launch Opportunities for Fixed Transit Time: The Synodic Period

- **Synodic Period, S_n :**
The time between conjunctions
 - P_A : Period of Planet A
 - P_B : Period of Planet B
- **Conjunction:** Two planets, A and B, in a line or at some fixed angle

$$S_n = \frac{P_A P_B}{P_A - P_B}$$



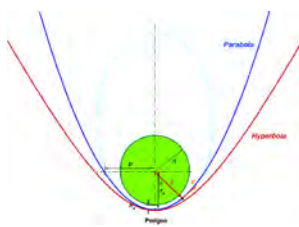
38

Launch Opportunities for Fixed Transit Time: The Synodic Period

Synodic Period with respect to Earth, days		
Planet		Period
Mercury	116	88 days
Venus	584	225 days
Earth	-	365 days
Mars	780	687 days
Jupiter	399	11.9 yr
Saturn	378	29.5 yr
Uranus	370	84 yr
Neptune	367	165 yr
Pluto	367	248 yr

39

Hyperbolic Orbits



Orbit Shape	Eccentricity, e	Energy, E
Circle	0	<0
Ellipse	0 < e < 1	<0
Parabola	1	0
Hyperbola	>1	>0

$$\mathcal{E} = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}, \quad \therefore a < 0$$

Velocity remains positive as radius approaches ∞

$$v \xrightarrow{r \rightarrow \infty} v_{\infty}$$

$$\therefore \mathcal{E}_{\infty} = \frac{v_{\infty}^2}{2}, \quad \text{and } v_{\infty} = \sqrt{-\frac{\mu}{a}} \quad \text{or} \quad a = -\frac{\mu}{v_{\infty}^2}$$

40

The diagram illustrates a hyperbolic trajectory around a central body (red dot labeled 'Center of Attraction'). The trajectory consists of an 'Approach Trajectory' and a 'Departure Trajectory', both asymptotically approaching lines labeled 'Asymptote'. Key parameters shown include:

- V_{∞}^+ and V_{∞}^- : Velocities at infinity for the departure and approach trajectories, respectively.
- Δ : The angle between the asymptotes.
- r : The radial distance from the center of attraction to a point on the trajectory.
- θ : The true anomaly, the angle between the radial line and the reference direction.
- δ : The deflection angle, the angle between the initial and final velocity vectors.
- $\left(\frac{\pi}{2} - \frac{\delta}{2}\right)$: The angle between the asymptote and the line perpendicular to the line of sight.
- p : The semi-latus rectum, the perpendicular distance from the center of attraction to the asymptote.
- r_p : The periapsis distance, the minimum distance from the center of attraction to the trajectory.
- $-a$: The negative semi-major axis.
- B : A point on the trajectory where the radial distance is r .
- Miss Distance : The distance from the center of attraction to the asymptote.

- Δ : Miss Distance, km

δ : Deflection Angle, deg or rad

Hyperbolic Orbits

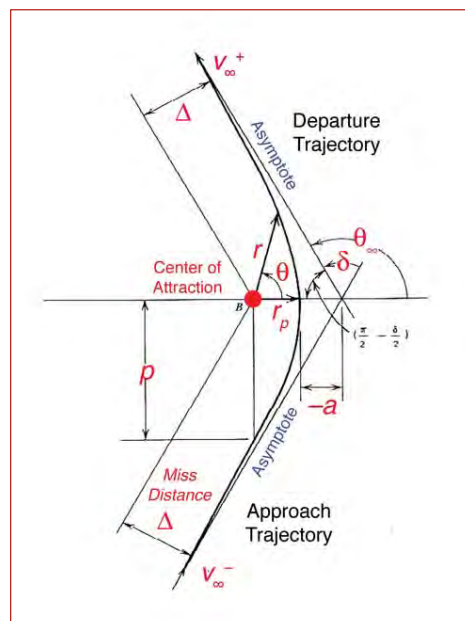
Polar Equation for a Conic Section

$$r = \frac{p}{1 + e \cos \theta} = \frac{a(1 - e^2)}{1 + e \cos \theta}$$

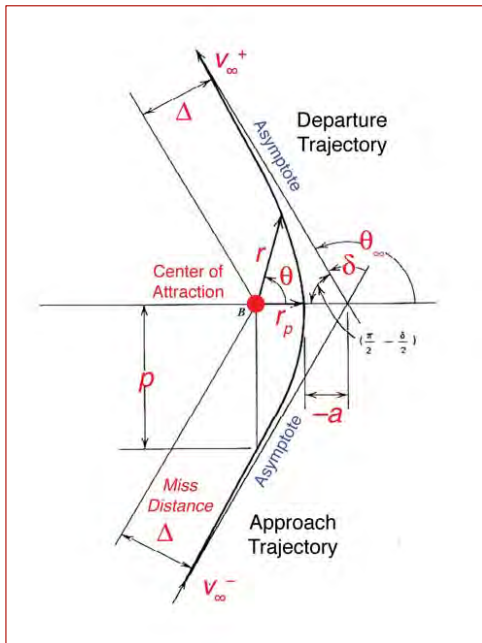
$$\cos \theta = \frac{1}{e} \left[\frac{a(1 - e^2)}{r} - 1 \right]$$

$$\theta \xrightarrow{r \leftrightarrow \infty} \theta_\infty$$

$$\theta_\infty = \cos^{-1}\left(-\frac{1}{e}\right)$$



Hyperbolic Orbits



Angular Momentum

$$h = \text{Constant} = v_{\infty} \Delta$$

$$= \sqrt{\mu p} = \sqrt{\mu a (1 - e^2)} = \sqrt{\frac{\mu^2 (e^2 - 1)}{v_{\infty}^2}}$$

Eccentricity

$$e = \sqrt{1 + \frac{2h^2 \mathcal{E}}{\mu^2}} = \sqrt{1 + \frac{v_{\infty}^4 \Delta^2}{\mu^2}}$$

Perigee Radius

$$r_p = a(1 - e) = \frac{\mu}{v_{\infty}^2} (e - 1)$$

Eccentricity

$$e = \left[1 + \frac{r_p v_{\infty}^2}{\mu} \right]$$

43

Hyperbolic Mean and Eccentric Anomalies

H : Hyperbolic Eccentric Anomaly

$$M = e \sinh H - H$$

Newton's method of successive approximation
to find H from M , similar to solution for E (Lecture 2)

$$\theta(t) = 2 \tan^{-1} \left[\sqrt{\frac{e+1}{e-1}} \tanh \frac{H(t)}{2} \right]$$

$$r = a(1 - e \cosh H)$$

see Ch. 7, Kaplan, 1976

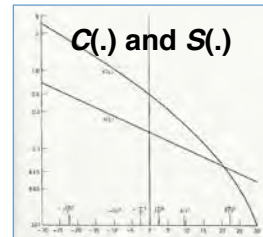
44

Battin's Universal Formulas for Conic Section Position and Velocity as Functions of Time

$$\mathbf{r}(t_2) = \left[1 - \frac{\chi^2}{r(t_1)} C\left(\frac{\chi^2}{a}\right) \right] \mathbf{r}(t_1) + \left[t_2 - \frac{\chi^3}{\sqrt{\mu}} S\left(\frac{\chi^2}{a}\right) \right] \mathbf{v}(t_1)$$

$$\mathbf{v}(t_2) = \frac{\sqrt{\mu}}{r(t_1)r(t_2)} \left[\frac{\chi^3}{a} S\left(\frac{\chi^2}{a}\right) - \chi \right] \mathbf{r}(t_1) + \left[1 - \frac{\chi^2}{r(t_2)} C\left(\frac{\chi^2}{a}\right) \right] \mathbf{v}(t_1)$$

$$\chi = \begin{cases} \sqrt{a} [E(t_2) - E(t_1)], & \text{Ellipse} \\ \sqrt{-a} [H(t_2) - H(t_1)], & \text{Hyperbola} \\ \sqrt{p} \left[\tan \frac{\theta(t_2)}{2} - \tan \frac{\theta(t_1)}{2} \right], & \text{Parabola} \end{cases}$$



see Ch. 7, Kaplan, 1976; also Battin, 1964

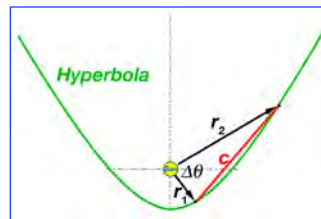
45

Lambert's Time-of-Flight Theorem (Hyperbolic Orbit)

$$(t_2 - t_1) = \sqrt{\frac{-a^3}{\mu}} [(\sinh \gamma - \gamma) + (\sinh \delta - \delta)]$$

where

$$\gamma \triangleq 2 \sinh^{-1} \sqrt{\frac{r_1 + r_2 + c}{-4a}}; \quad \delta \triangleq 2 \sinh^{-1} \sqrt{\frac{r_1 + r_2 - c}{-4a}}$$



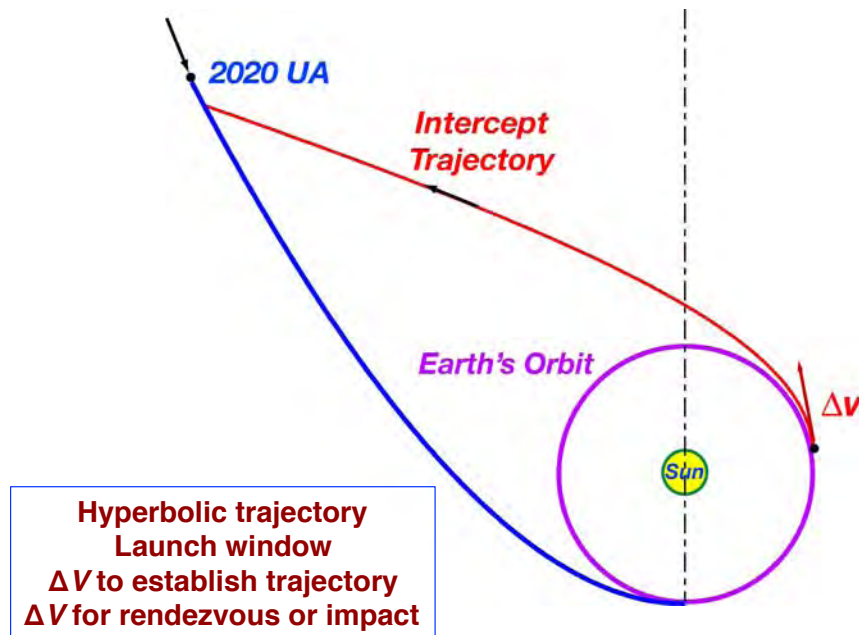
see Ch. 7, Kaplan, 1976; also Battin, 1964

<http://www.mathworks.com/matlabcentral/fileexchange/39530-lambert-s-problem>

<http://www.mathworks.com/matlabcentral/fileexchange/26348-robust-solver-for-lambert-s-orbital-boundary-value-problem>

46

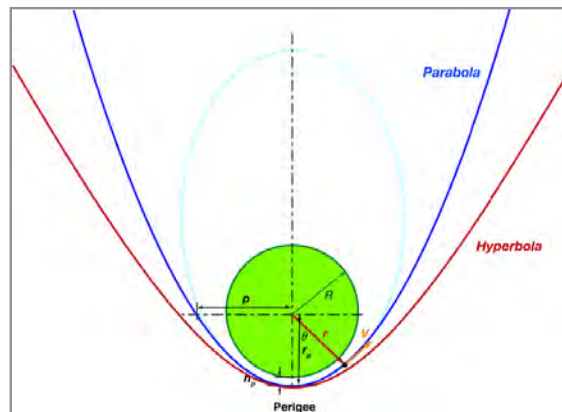
Asteroid Encounter



47

Swing-By/Fly-By Trajectories

- Hyperbolic encounters with planets and the moon provide gravity assist
 - Shape, energy, and duration of transfer orbit altered
 - Potentially large reduction in rocket ΔV required to accomplish mission

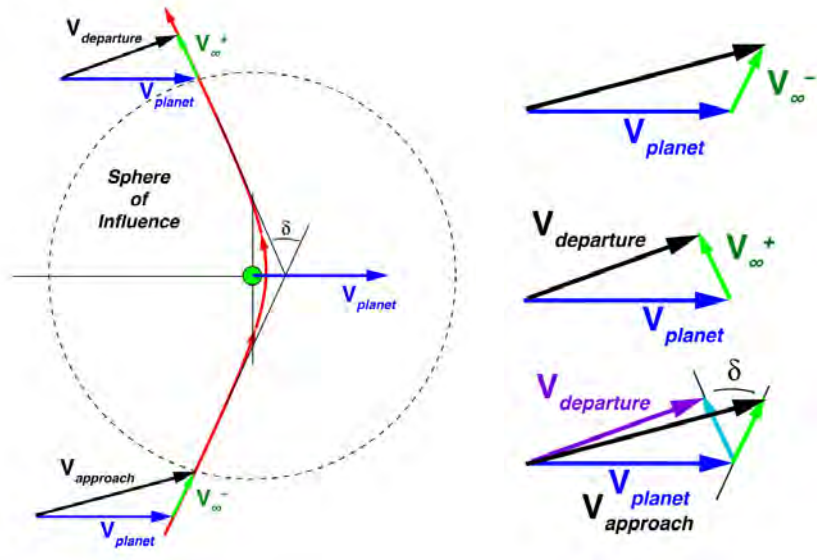


Why does gravity assist work?

48

Effect of Target Planet's Gravity on Probe's Sun-Relative Velocity

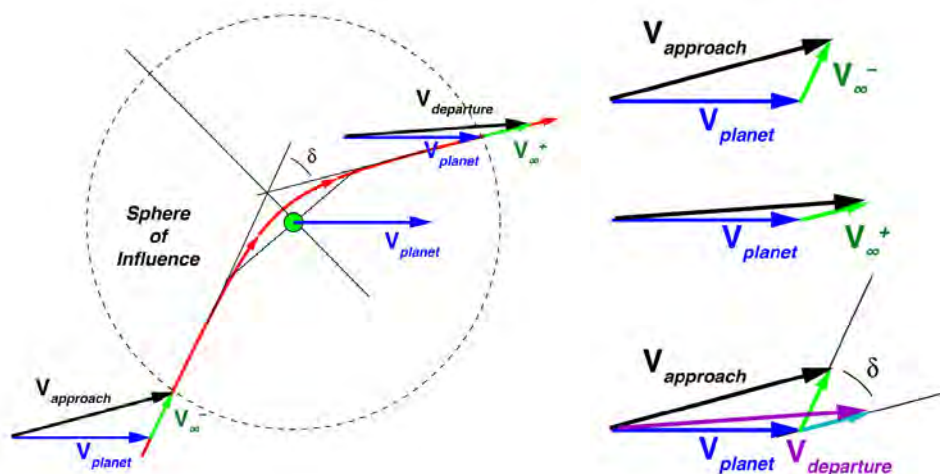
Deflection – Velocity Reduction



49

Effect of Target Planet's Gravity on Probe's Sun-Relative Velocity

Deflection – Velocity Addition

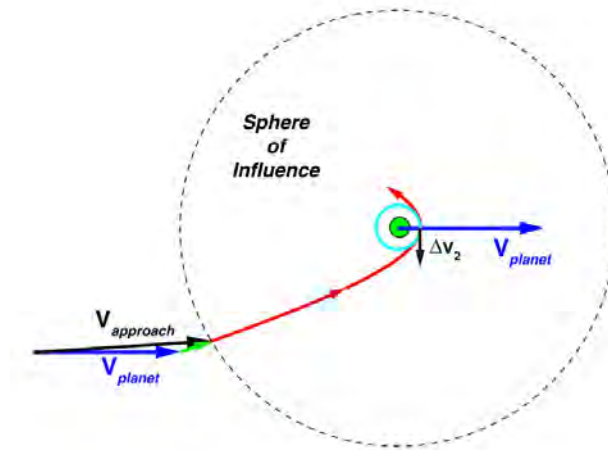


50

Planet Capture Trajectory

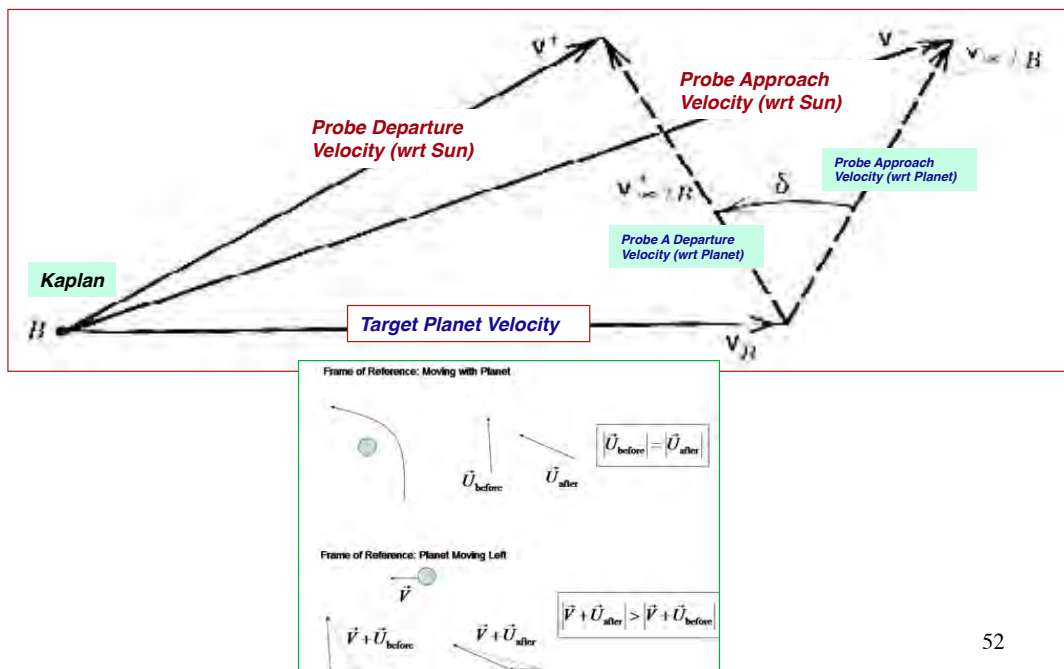
Hyperbolic approach to planet's sphere of influence

Δv to decrease speed to circular velocity



51

Effect of Target Planet's Gravity on Probe's Velocity



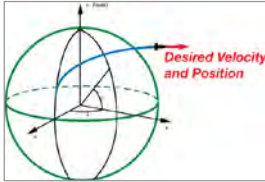
52

***Next Time:
Spacecraft Environment***

53

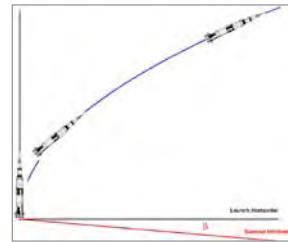
Supplemental Material

54



Phases of Ascent Guidance

- Vertical liftoff
- Roll to launch azimuth
- Pitch program to atmospheric “exit”
 - Jet stream penetration
 - Booster cutoff and staging
- Explicit guidance to desired orbit
 - Booster separation
 - Acceleration limiting
 - Orbital insertion

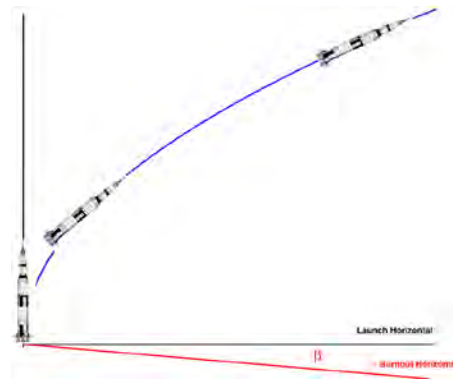


55

Tangent Steering Laws

- Neglecting surface curvature

$$\tan \theta(t) = \tan \theta_o \left(1 - \frac{t}{t_{BO}} \right)$$
- “Open-loop” command, i.e., no feedback of vehicle state



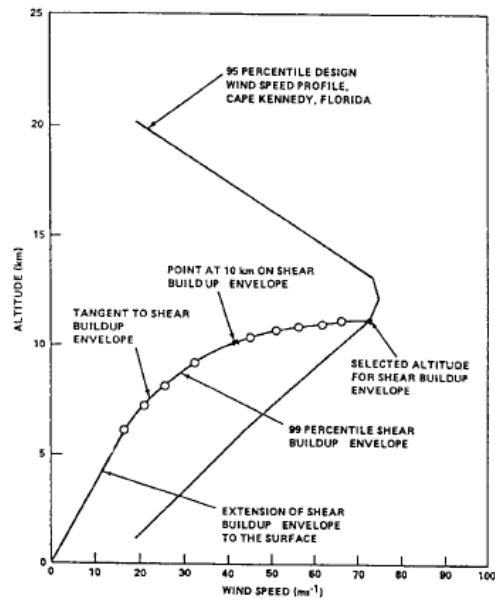
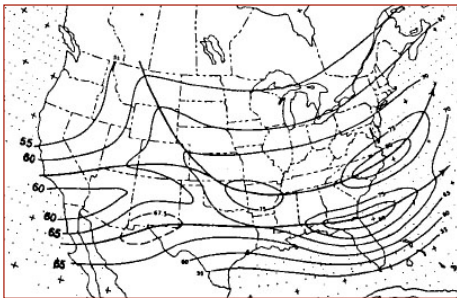
- Accounting for effect of Earth surface curvature on burnout flight path angle

$$\tan \theta(t) = \tan \theta_o \left[1 - \frac{t}{t_{BO}} - \tan \beta \left(\frac{t}{t_{BO}} \right) \right]$$

56

Jet Stream Profiles

- Launch vehicle must be able to fly through strong wind profiles
- Design profiles assume 95th-99th-percentile worst winds and wind shear



57

Longitudinal (2-D) Equations of Motion for Re-Entry



Differential equations for velocity ($x_1 = V$), flight path angle ($x_2 = \gamma$), altitude ($x_3 = h$), and range (x_4)

Angle of attack (α) is optimization control variable

$$\dot{x}_1 = -D(x_1, x_3, \alpha)/m - g \cos x_2$$

$$\dot{x}_2 = [g/x_1 - x_1/(R + x_3)] \sin x_2 - L(x_1, x_3, \alpha)/mx_1$$

$$\dot{x}_3 = x_1 \cos x_2$$

$$\dot{x}_4 = x_1 \sin x_2 / (1 + x_3/R)$$

D = drag
 L = lift
 M = pitch moment
 R = Earth radius

Equations of motion define the **dynamic constraint**

$$\dot{\mathbf{x}}(t) = \mathbf{f}[\mathbf{x}(t), \alpha(t)]$$

58

A Different Approach to Guidance: Optimizing a Cost Function

- Minimize a scalar function, J , of terminal and integral costs

$$J = \phi[\mathbf{x}(t_f)] + \int_{t_o}^{t_f} L[\mathbf{x}(t), \mathbf{u}(t)] dt$$

$L[\cdot]$: Lagrangian

with respect to the control, $\mathbf{u}(t)$, in (t_o, t_f) ,
subject to a dynamic constraint

$$\dot{\mathbf{x}}(t) = \mathbf{f}[\mathbf{x}(t), \mathbf{u}(t)], \quad \mathbf{x}(t_o) \text{ given}$$

$\dim(\mathbf{x}) = n \times 1$
 $\dim(\mathbf{f}) = n \times 1$
 $\dim(\mathbf{u}) = m \times 1$

59

Guidance Cost Function

$\begin{aligned} &\phi[\mathbf{x}(t_f)] \\ &L[\mathbf{x}(t), \mathbf{u}(t)] \end{aligned}$	<ul style="list-style-type: none"> Terminal cost, e.g., in final position and velocity Integral cost, e.g., tradeoff between control usage and trajectory error
--	---

- Minimization of cost function determines the optimal state and control, \mathbf{x}^* and \mathbf{u}^* , over the flight path duration

$$\begin{aligned} \min_u J &= \min_u \left\{ \phi[\mathbf{x}(t_f)] + \int_{t_o}^{t_f} L[\mathbf{x}(t), \mathbf{u}(t)] dt \right\} \\ &= \phi[\mathbf{x}^*(t_f)] + \int_{t_o}^{t_f} L[\mathbf{x}^*(t), \mathbf{u}^*(t)] dt \rightarrow [\mathbf{x}^*(t), \mathbf{u}^*(t)] \end{aligned}$$

60



Example of Re-Entry Flight Path Cost Function

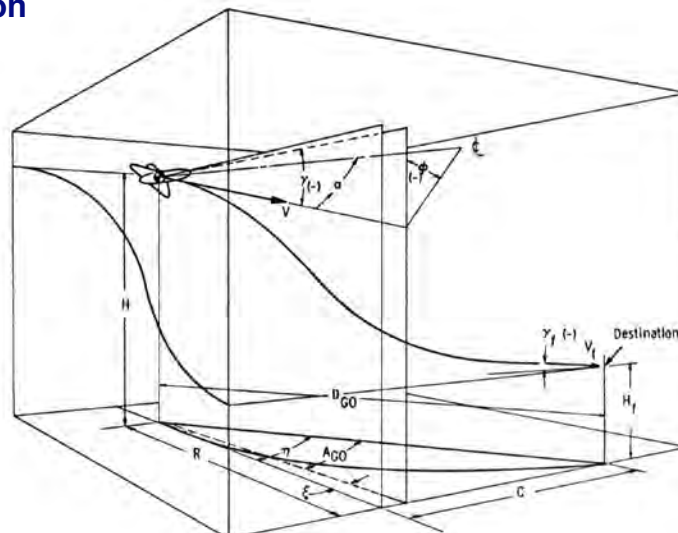
$$J = a[V(t_f) - V_d]^2 + b[r(t_f) - r_d]^2 + \int_{t_o}^{t_f} c[\alpha(t)]^2 dt$$

- **Cost function includes**
 - Terminal range and velocity
 - Penalty on control use
 - **a**, **b**, and **c** tradeoff importance of each factor
- **Minimization of this cost function**
 - Defines the optimal path, **x*(t)**, from t_o to t_f
 - Defines the optimal control, **α*(t)**, from t_o to t_f

61

Extension to Three Dimensions

- Add roll angle as a control; add crossrange as a state
- For the guidance law, replace range and crossrange from the starting point by distance to go and azimuth to go to the destination

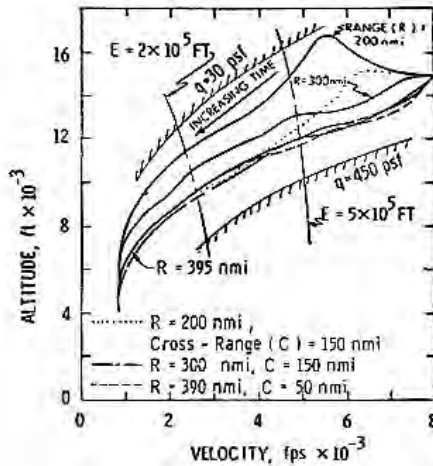


62

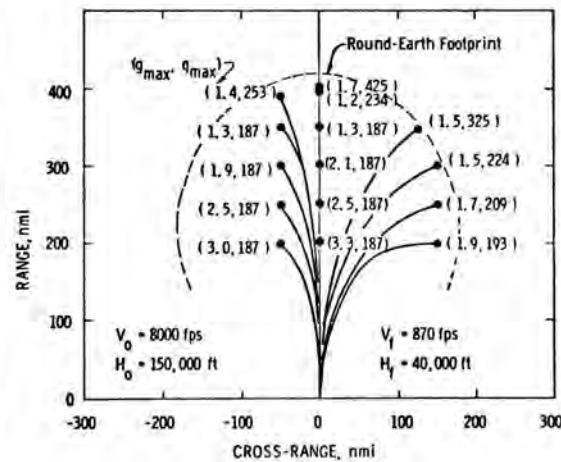


Optimal Trajectories for Space Shuttle Transition

Altitude vs. Velocity



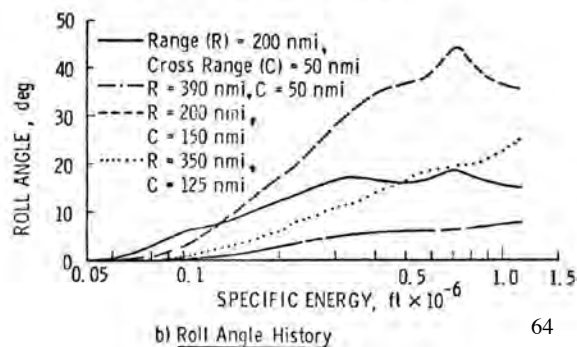
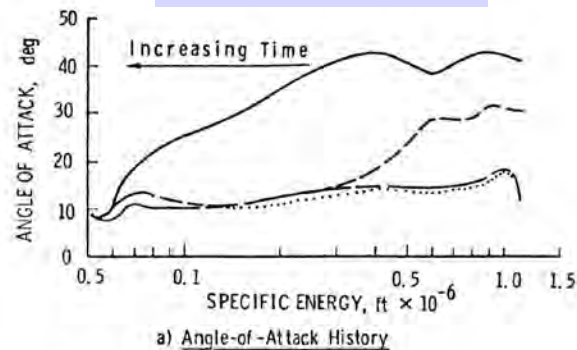
Range vs. Cross-Range ("footprint")



63

Optimal Controls for Space Shuttle Transition

Angle of Attack and Roll Angle vs. Specific Energy



64

Optimal Guidance System Derived from Optimal Trajectories

Angle of Attack and Roll Angle vs. Specific Energy

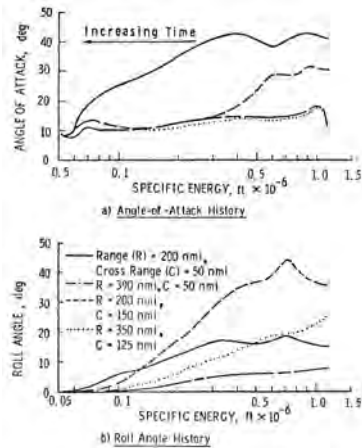
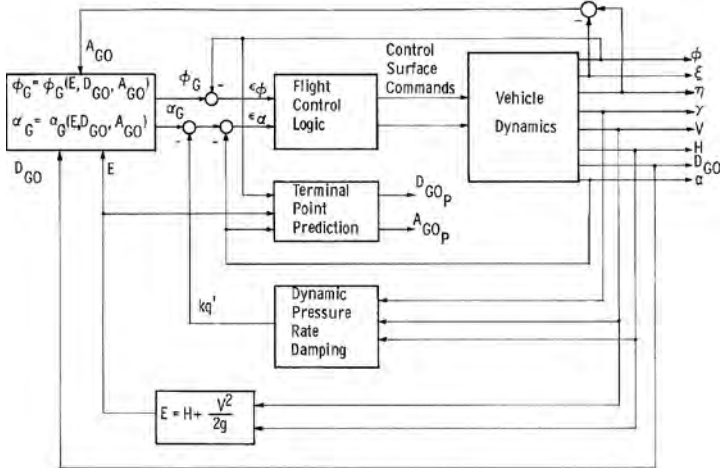


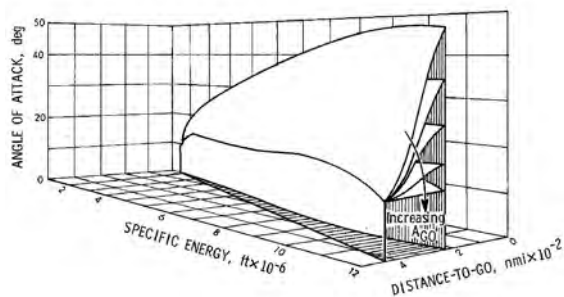
Diagram of Energy-Guidance Law



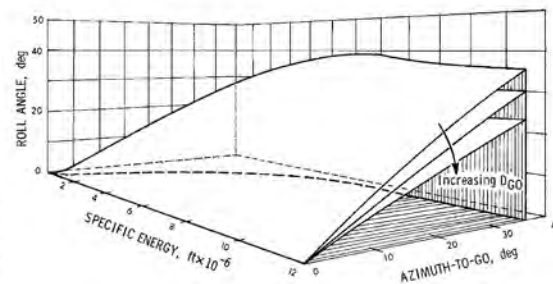
65

Guidance Functions for Space Shuttle Transition

Angle of Attack Guidance Function

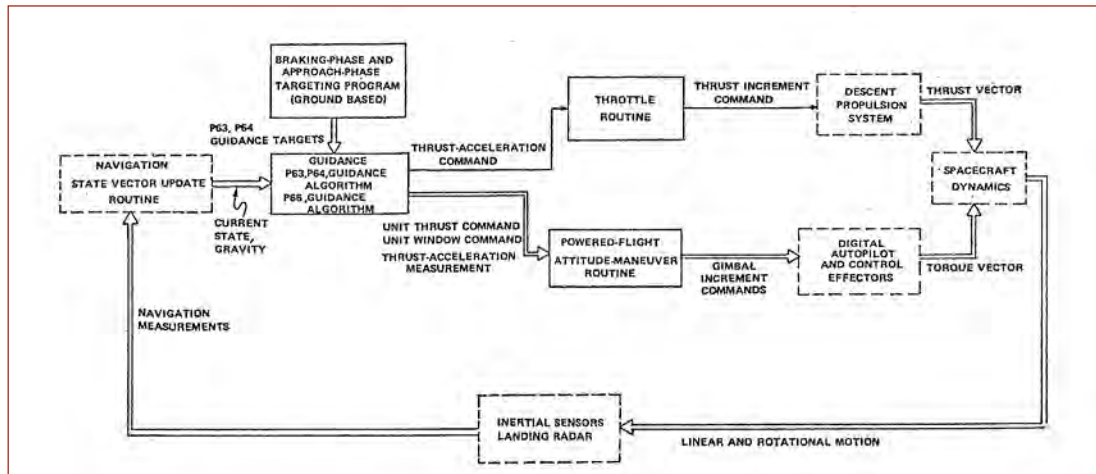


Roll Angle Guidance Function



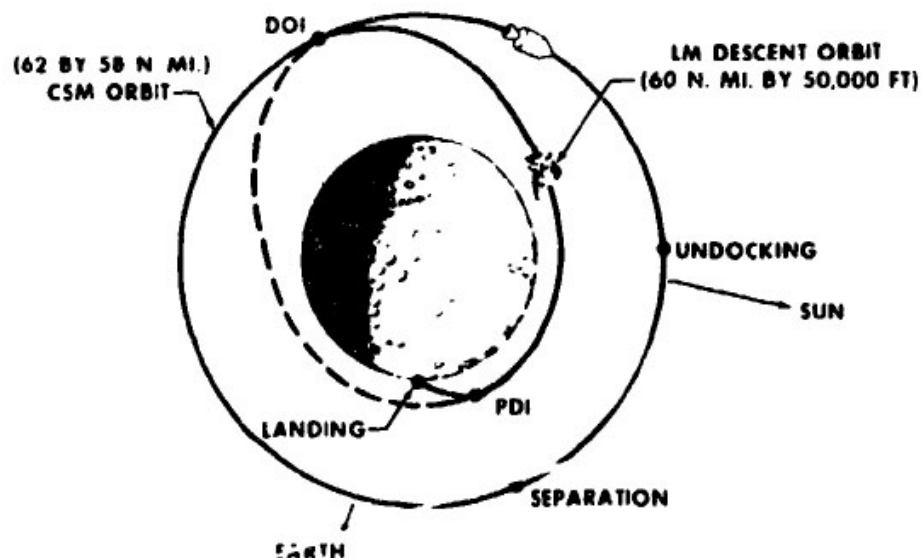
66

Lunar Module Navigation, Guidance, and Control Configuration



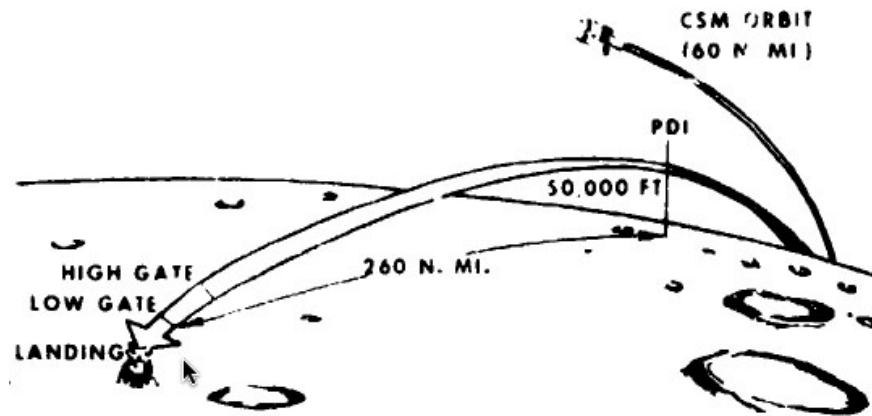
67

Lunar Module Transfer Ellipse to Powered Descent Initiation



68

Lunar Module Powered Descent



PHASE	INITIAL EVENT	DESIGN CRITERIA
BRAKING	PDI	MINIMIZE PROPELLANT USAGE
APPROACH	HIGH GATE	CREW VISIBILITY
LANDING	LOW GATE	MANUAL CONTROL

69

Lunar Module Descent Events

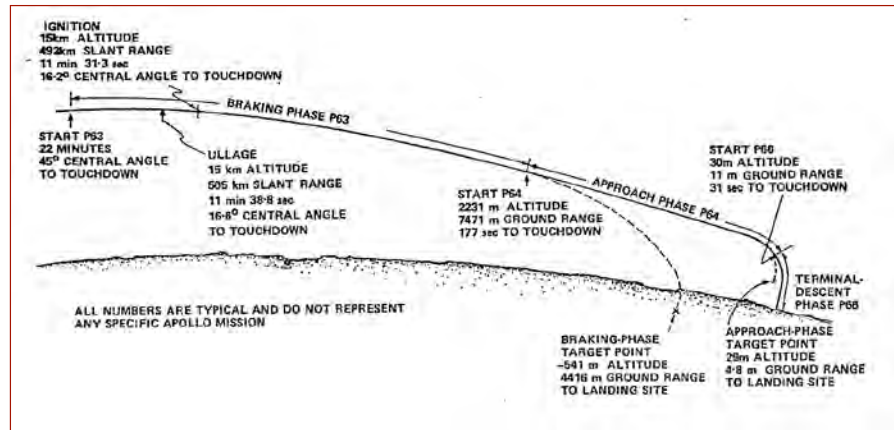
Event	TFL, ^a min:sec	Inertial velocity, fps	Altitude rate, fps	Altitude, ft	ΔV, fps
A Ullage	-00:07				
B Powered descent initiation	00:00	5560	-4	48 814	0
C Throttle to maxi- mum thrust	00:26	5529	-3	48 725	31
D Rotate to windows- up position	02:56	4000	-50	44 934	1572
E LR altitude update	04:18	3065	-89	39 201	2536
F Throttle recovery	06:24	1456	-106	24 635	4239
G LR velocity update	06:42	1315	-127	22 644	4390
H High gate	08:26	506	-145	7 515	5375
I Low gate	10:06	55 (^b 68)	-16	512	6176
J Touchdown (probe contact)	11:54	-15 (^b 0)	-3	12	6775

^aTime from ignition of the DPS.

^bHorizontal velocity relative to surface.

70

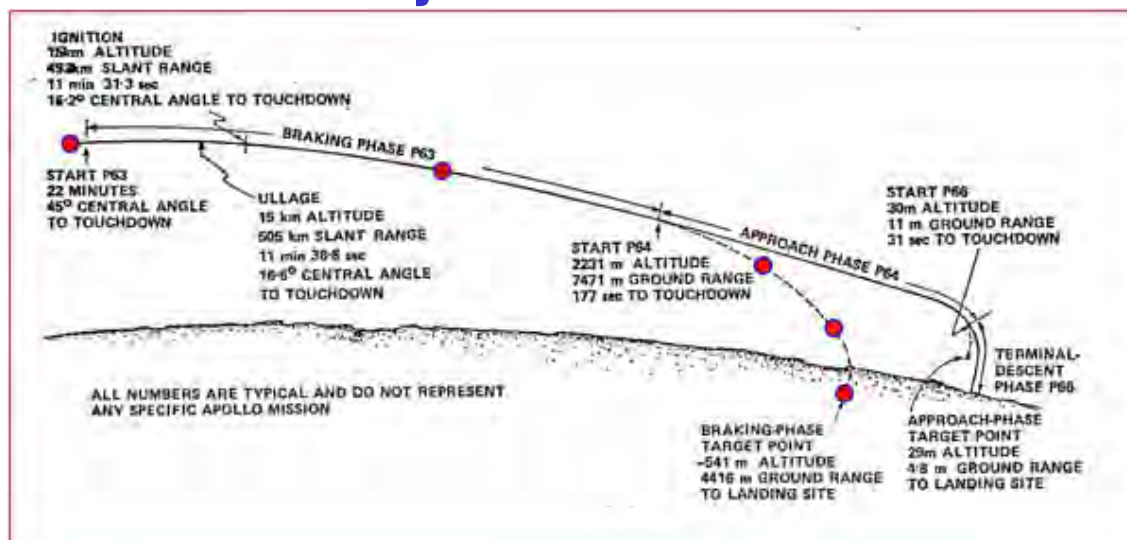
Lunar Module Descent Targeting Sequence



Braking Phase (P63)
Approach Phase (P64)
Terminal Descent Phase (P66)

71

Characterize Braking Phase By Five Points



72

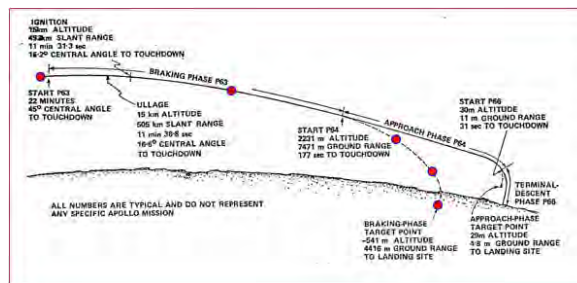
Lunar Module Descent Guidance Logic

(Klumpp, *Automatica*, 1974)

- Reference (nominal) trajectory, $\mathbf{r}_r(t)$, from target position **back** to starting point (Braking Phase example)
 - Three 4th-degree polynomials in time
 - 5 points needed to specify each polynomial

$$\mathbf{r}_r(t) = \mathbf{r}_t + \mathbf{v}_t t + \mathbf{a}_t \frac{t^2}{2} + \mathbf{j}_t \frac{t^3}{6} + \mathbf{s}_t \frac{t^4}{24}$$

$$\mathbf{r}(t) = \begin{bmatrix} x(t) \\ y(t) \\ z(t) \end{bmatrix}$$



73

Coefficients of the Polynomials

$$\mathbf{r}_r(t) = \mathbf{r}_t + \mathbf{v}_t t + \mathbf{a}_t \frac{t^2}{2} + \mathbf{j}_t \frac{t^3}{6} + \mathbf{s}_t \frac{t^4}{24}$$

- \mathbf{r} = position vector
- \mathbf{v} = velocity vector
- \mathbf{a} = acceleration vector
- \mathbf{j} = jerk vector (time derivative of acceleration)
- \mathbf{s} = snap vector (time derivative of jerk)

$$\mathbf{r} = \begin{bmatrix} x \\ y \\ z \end{bmatrix}$$

$$\mathbf{v} = \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix} = \begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix}$$

$$\mathbf{a} = \begin{bmatrix} a_x \\ a_y \\ a_z \end{bmatrix}$$

$$\mathbf{j} = \begin{bmatrix} j_x \\ j_y \\ j_z \end{bmatrix}$$

$$\mathbf{s} = \begin{bmatrix} s_x \\ s_y \\ s_z \end{bmatrix}$$

74

Corresponding Reference Velocity and Acceleration Vectors

$$\mathbf{v}_r(t) = \mathbf{v}_t + \mathbf{a}_t t + \mathbf{j}_t \frac{t^2}{2} + \mathbf{s}_t \frac{t^3}{6}$$

$$\mathbf{a}_r(t) = \mathbf{a}_t + \mathbf{j}_t t + \mathbf{s}_t \frac{t^2}{2}$$

- $\mathbf{a}_r(t)$ is the reference control vector
 - Descent engine thrust / mass = total acceleration
 - Vector components controlled by orienting yaw and pitch angles of the Lunar Module



Guidance Logic Defines Desired Acceleration Vector

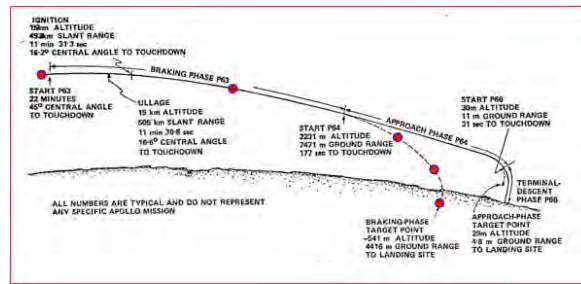
- If initial conditions, dynamic model, and thrust control were perfect, $\mathbf{a}_r(t)$ would produce $\mathbf{r}_r(t)$

$$\mathbf{a}_r(t) = \mathbf{a}_t + \mathbf{j}_t t + \mathbf{s}_t \frac{t^2}{2} \Rightarrow \mathbf{r}_r(t) = \mathbf{r}_t + \mathbf{v}_t t + \mathbf{a}_t \frac{t^2}{2} + \mathbf{j}_t \frac{t^3}{6} + \mathbf{s}_t \frac{t^4}{24}$$

- ... but they are not
- Therefore, **feedback control** is required to follow the reference trajectory



Guidance Law for the Lunar Module Descent



Linear feedback guidance law

$$\mathbf{a}_{command}(t) = \mathbf{a}_r(t) + \mathbf{K}_V [\mathbf{v}_{measured}(t) - \mathbf{v}_r(t)] + \mathbf{K}_R [\mathbf{r}_{measured}(t) - \mathbf{r}_r(t)]$$

\mathbf{K}_V : velocity error gain

\mathbf{K}_R : position error gain

Nominal acceleration profile is corrected for measured differences between actual and reference flight paths

Considerable modifications made in actual LM implementation (see Klumpp's original paper on *Blackboard*)