

ISMS conference

The pure rotational spectrum of hydroxymethyl radical reinvestigated to enable its interstellar detection

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V. Lattanzi³, H. Müller⁴ and O. Pirali¹

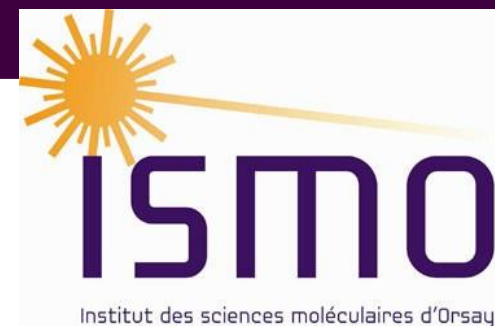


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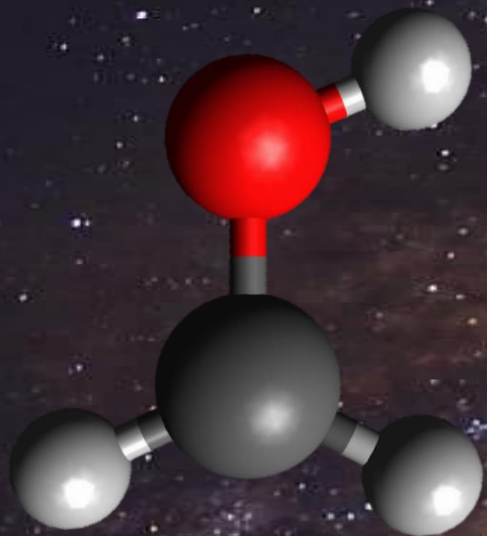
Olivia Chitarra | June 23, 2021

CH_2OH : a missing link in interstellar chemistry?

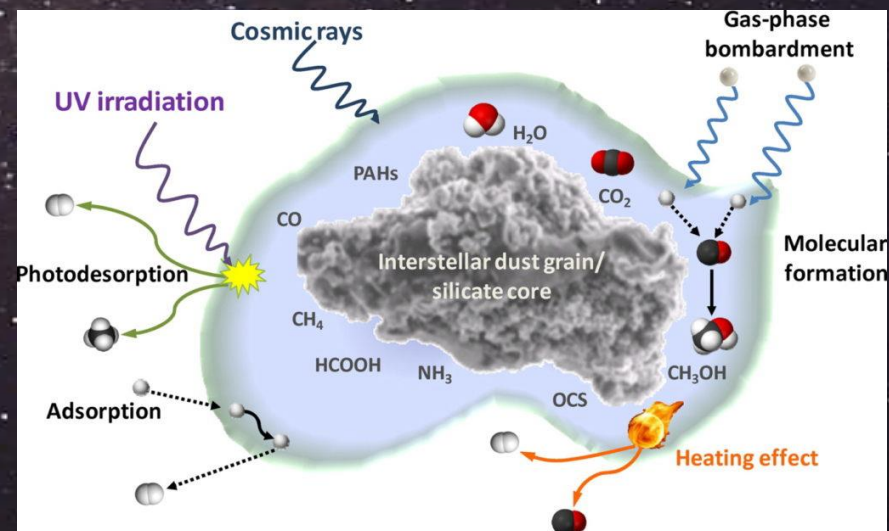
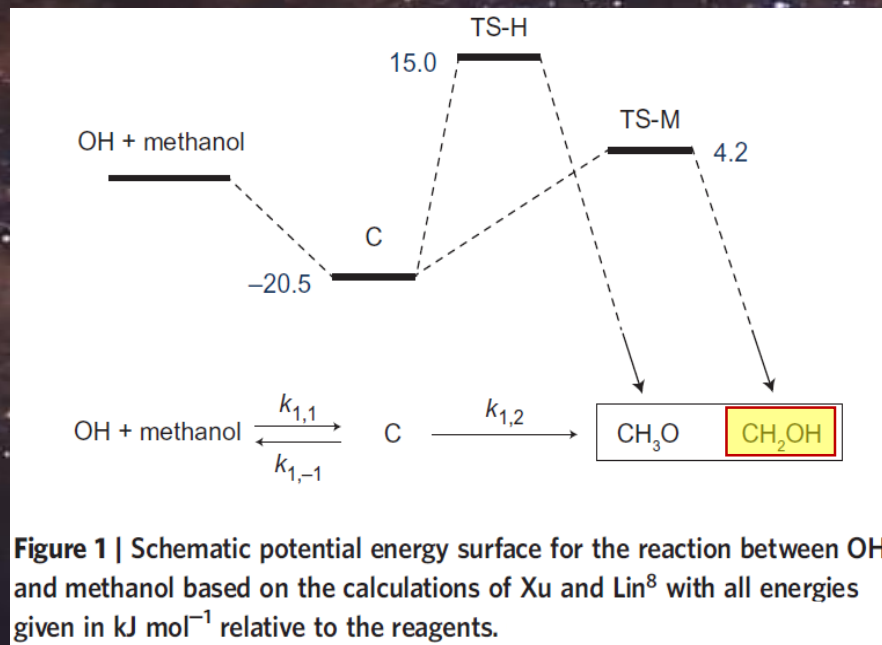


□ Hydrogen, Carbon and Oxygen containing species

CH₂OH: a missing link in interstellar chemistry?



- Hydrogen, Carbon and Oxygen containing species
- Potentially produced from methanol (ice and/ or gas phase)



Shannon R. J. et al., *Nature Chemistry*, **5**, 745-749, 2013

Öberg K. I. et al., *Astronomy & Astrophysics*, **504**, 891-913, 2009

Lucas M. et al., *Chemical Physics Letters*, **619**, 18-22, 2015

Acharyya K. et al., *Molecular Physics*, **113**, 2243-2254, 2015

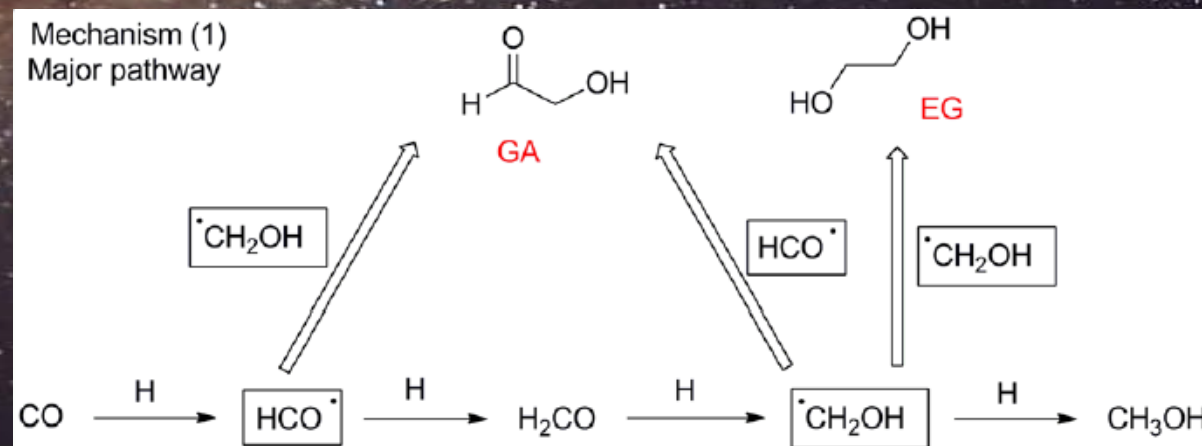
Meinert C. et al., *Physics of Life Reviews*, **8**, 307-330, 2011

Bertin M. et al., *The Astro. J. Letters*, **817**, 2016

CH₂OH: a missing link in interstellar chemistry?



- ❑ Hydrogen, Carbon and Oxygen containing species
- ❑ Potentially produced from methanol (ice and/ or gas phase)
- ❑ Precursor of Complex Organic Molecules (COMs)



Balucani N. et al., *MNRAS*, **449**, L16-L20, 2015

Öberg K. I. et al., *Astronomy & Astrophysics*, **504**, 891-913, 2009

Maity S. et al., *Phys. Chem. Chem. Phys.*, **17**, 3081-3114, 2015

Butscher T. et al., *MNRAS*, **453**, 1587-1596, 2015

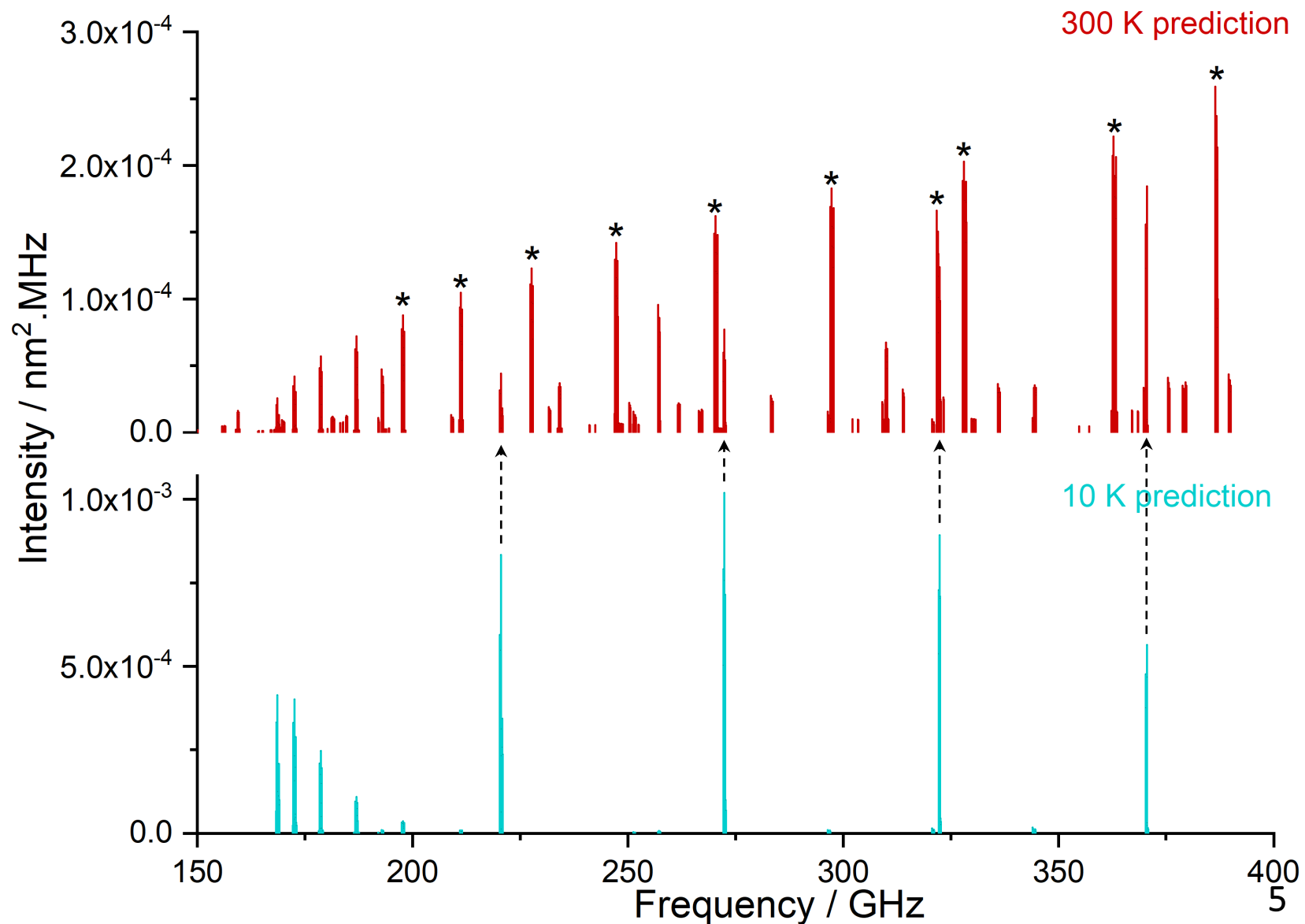
CH₂OH: laboratory challenge

Ro-vibration studies

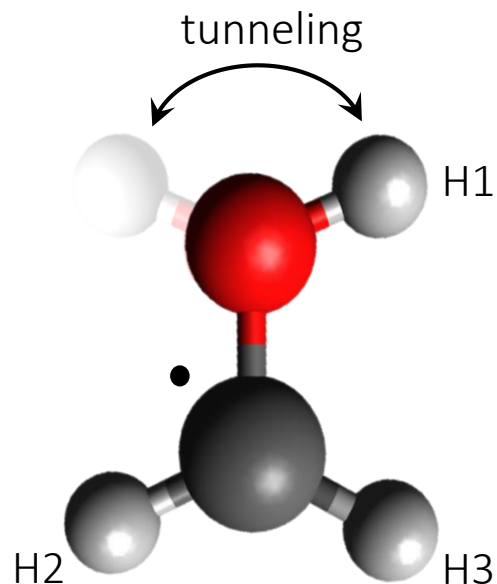
- Infrared spectral region
- Supersonic jet
- Roberts M. A. *et al.*, *The Journal of Physical Chemistry*, **117**, 7042-7049, 2013
- Schuder M. D. *et al.*, *The Journal of Chemical Physics*, **146**, 194307, 2017

Rotational study

- Millimeter-wave spectral region
- First observation of the rotational-tunneling spectrum
- Bermudez C. *et al.*, *Astronomy & Astrophysics*, **598**, A9, 2017



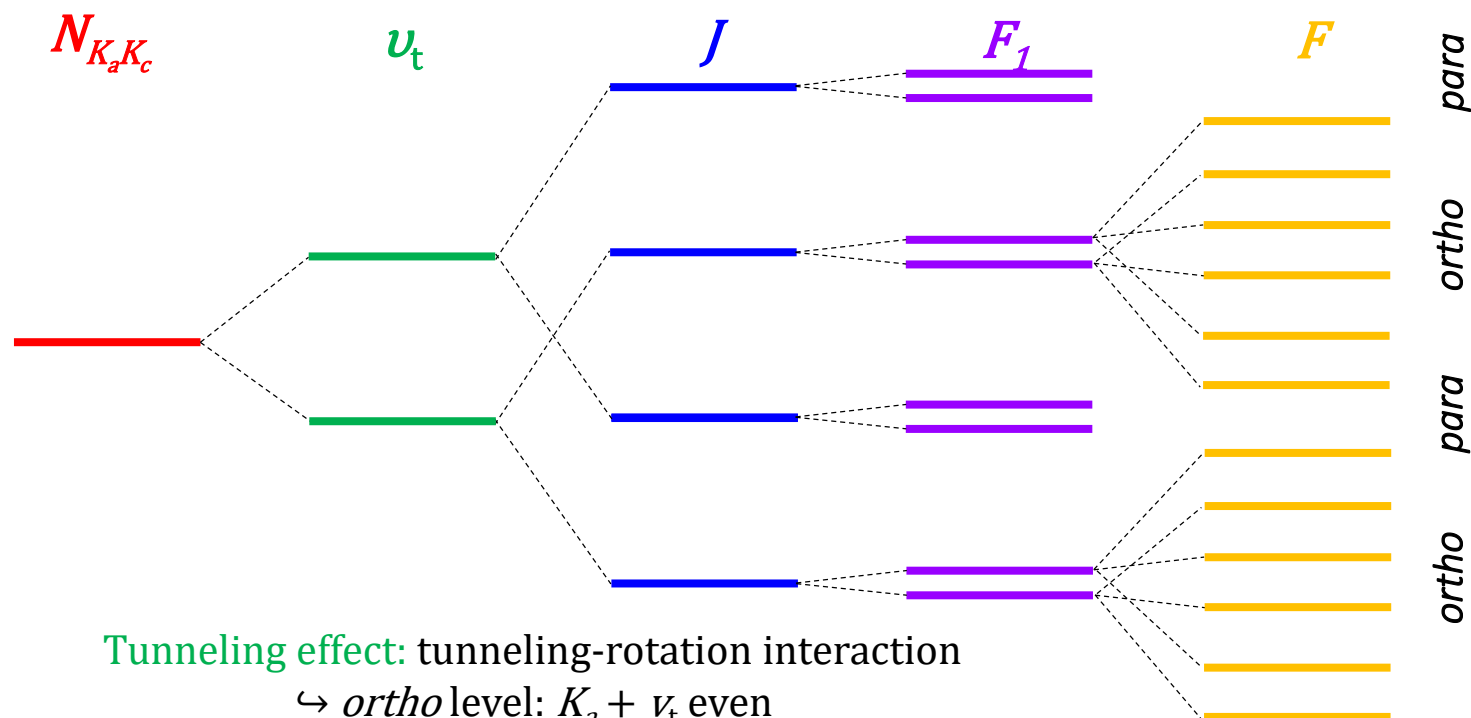
Spectroscopy



❑ Asymmetric top molecule

❑ Planar in $v = 0$

- One unpaired electron
- 3 H (of non-zero nuclear spin)
- Tunneling of H1



Tunneling effect: tunneling-rotation interaction

↪ *ortho* level: $K_a + v_t$ even

↪ *para* level: $K_a + v_t$ odd

Fine structure: spin-rotation interaction

↪ $J = N + S$

Hyperfine Structure: hydroxyl hydrogen nucleus interaction

↪ $F_1 = J + I_{H_1}$

Hyperfine Structure: equivalent methylenic hydrogen nucleus interaction

↪ $F = J + I_{H_2} + I_{H_3}$

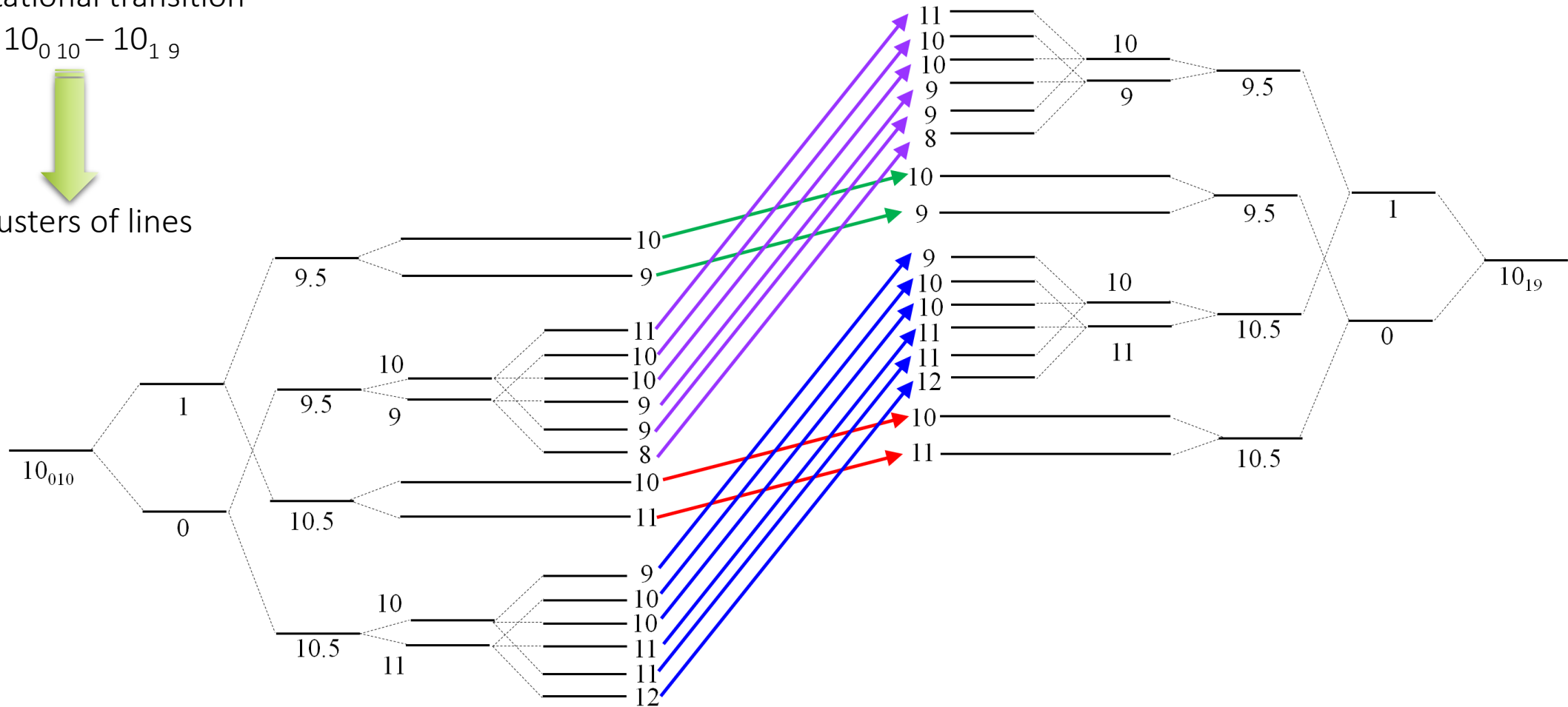
Spectroscopy

One rotational transition

ex: $10_{010} - 10_{19}$



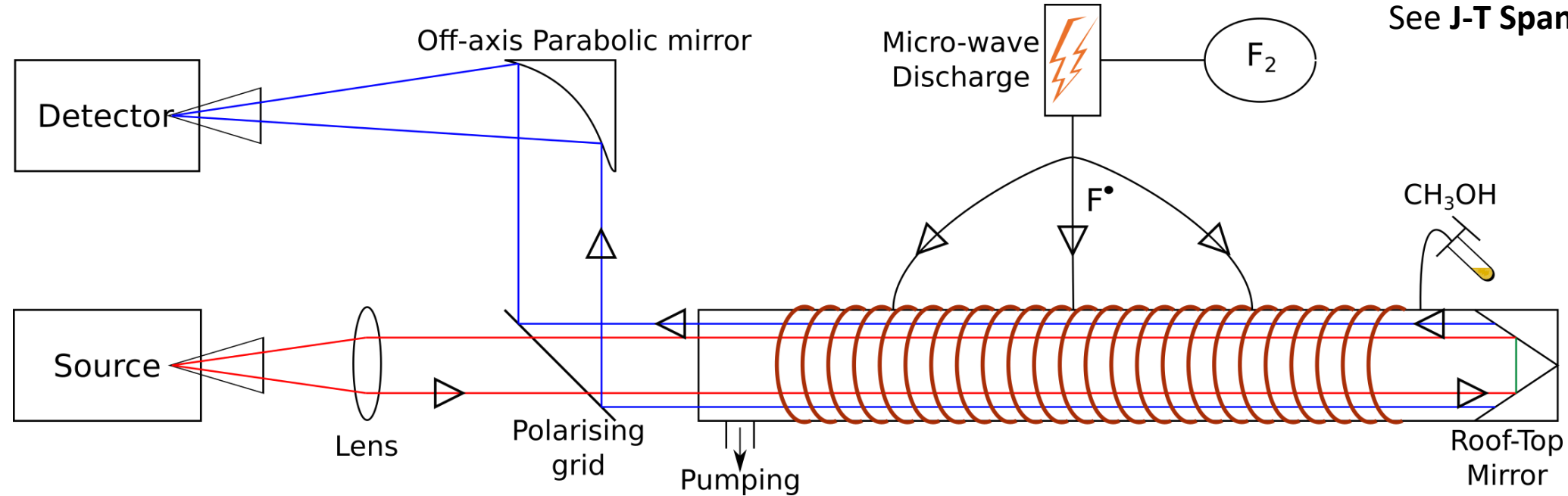
4 clusters of lines



Systematic measurements **compulsory**

Experimental methods

See J-T Spaniol's talk (TI05)



Production

Hydrogen abstraction from methanol by fluorine atom

Acquisition

Frequency multiplication chain spectrometer:
140-900 GHz

Magnetic Field Modulation to distinguish paramagnetic lines

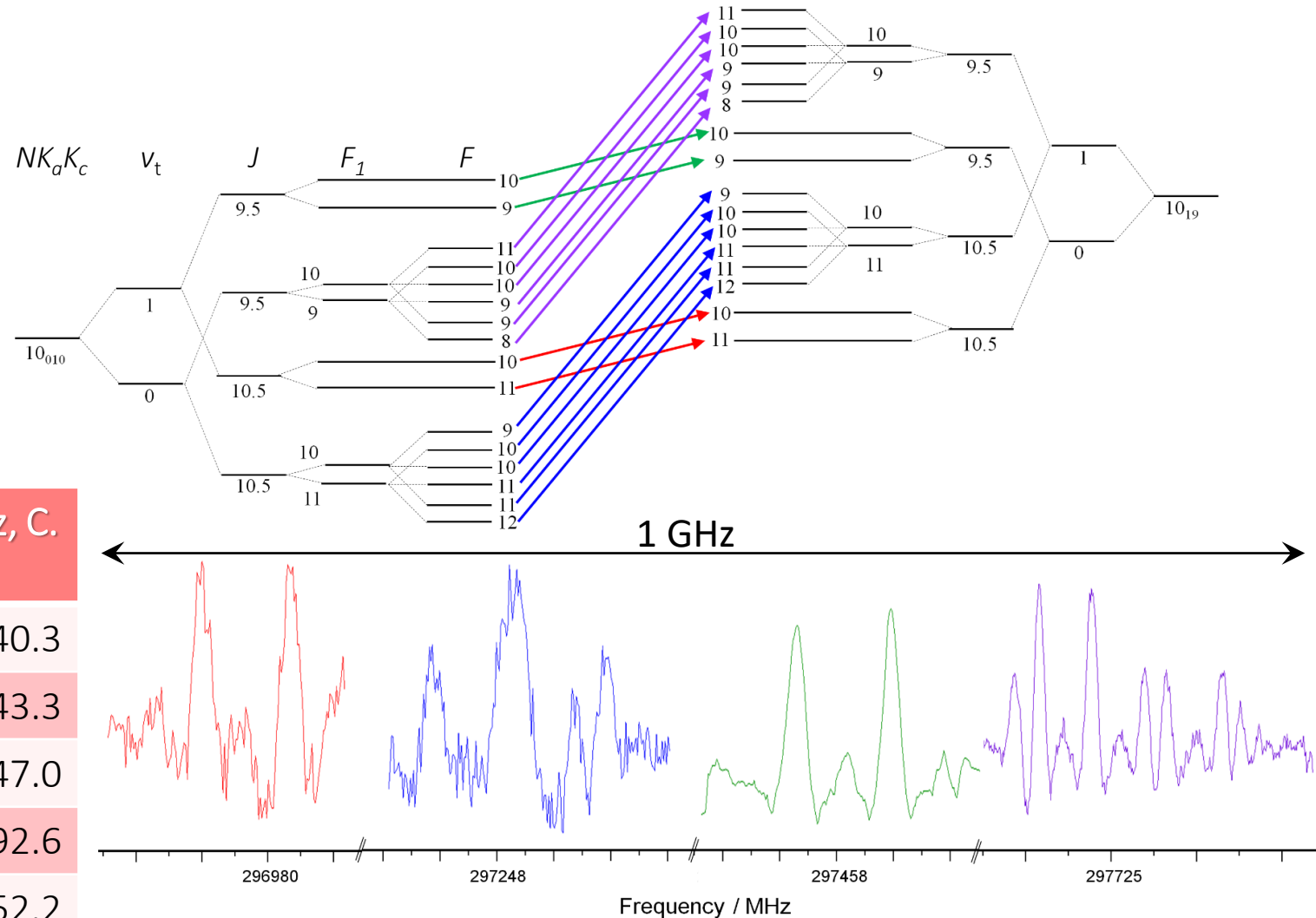
Results and discussion

Experimental measurements

- 676 mmw lines recorded + 96 from Bermudez et al. (2017)
- Freq. accuracies: from 50 to 800 kHz
- $^{\text{r}}\text{R}_{1-2}(1-15)$, $^{\text{r}}\text{Q}_{1-2}(1-16)$ and $^{\text{p}}\text{R}_{0-1}(6-17)$ transitions



Reliable and constrained fit



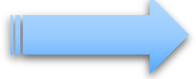
See L Coudert's talk (FC12)

Parameters (MHz)	Present	Chitarra, O. et al.	Bermudez, C. et al.
A	194536.5	194535.4	194540.3
B	28869.5	29843.4	29843.3
C	26921.5	25947.5	25947.0
$D_{NK} \times 10^3$	593.7	584.7	792.6
$a_F(\text{H})$	-54.0	-54.4	-52.2
RMS	0.409	0.081	0.077

Results and discussion

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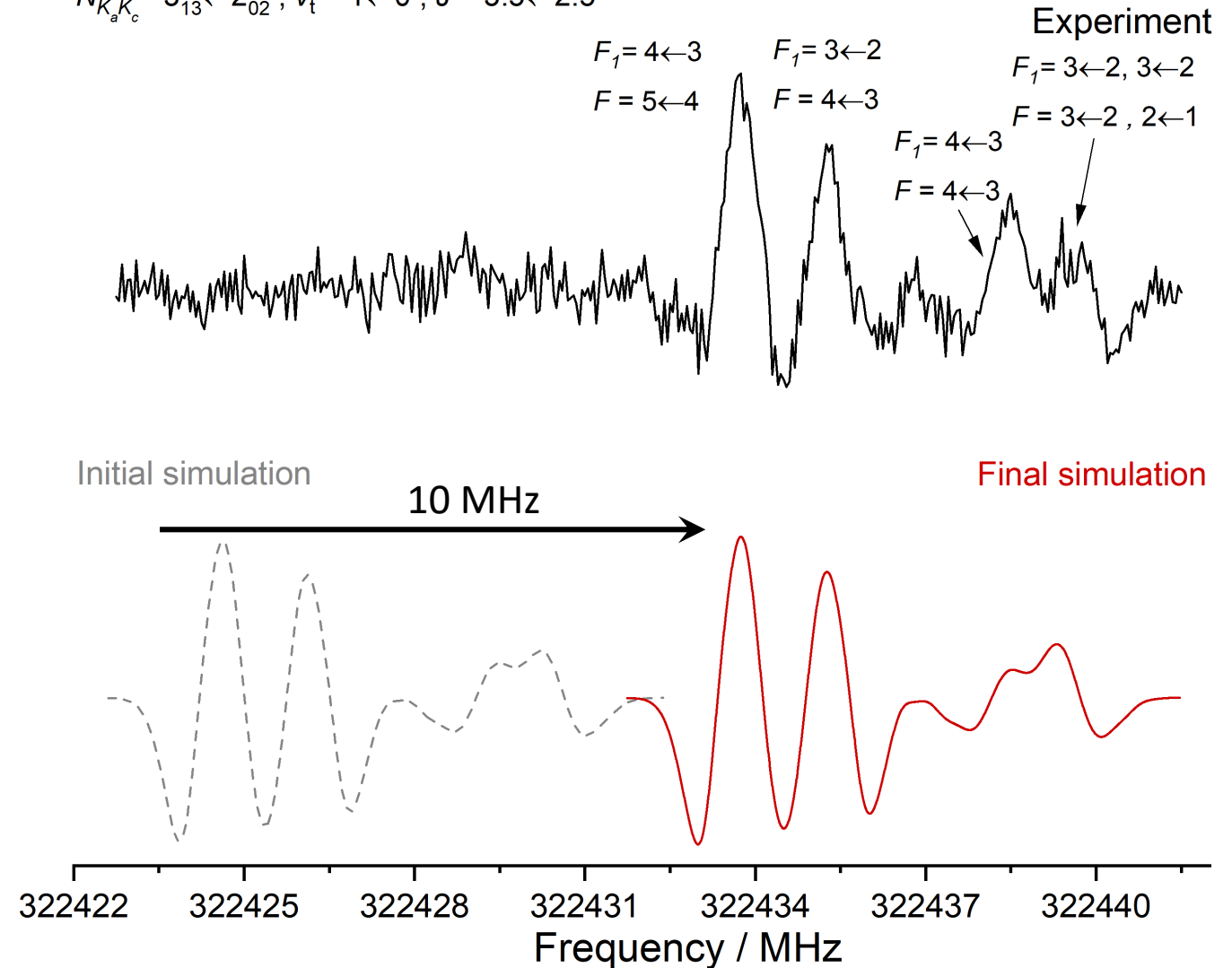
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$$N_{K_a K_c} = 3_{13} \leftarrow 2_{02}, v_t = 1 \leftarrow 0, J = 3.5 \leftarrow 2.5$$



See L Coudert's talk (FC12)

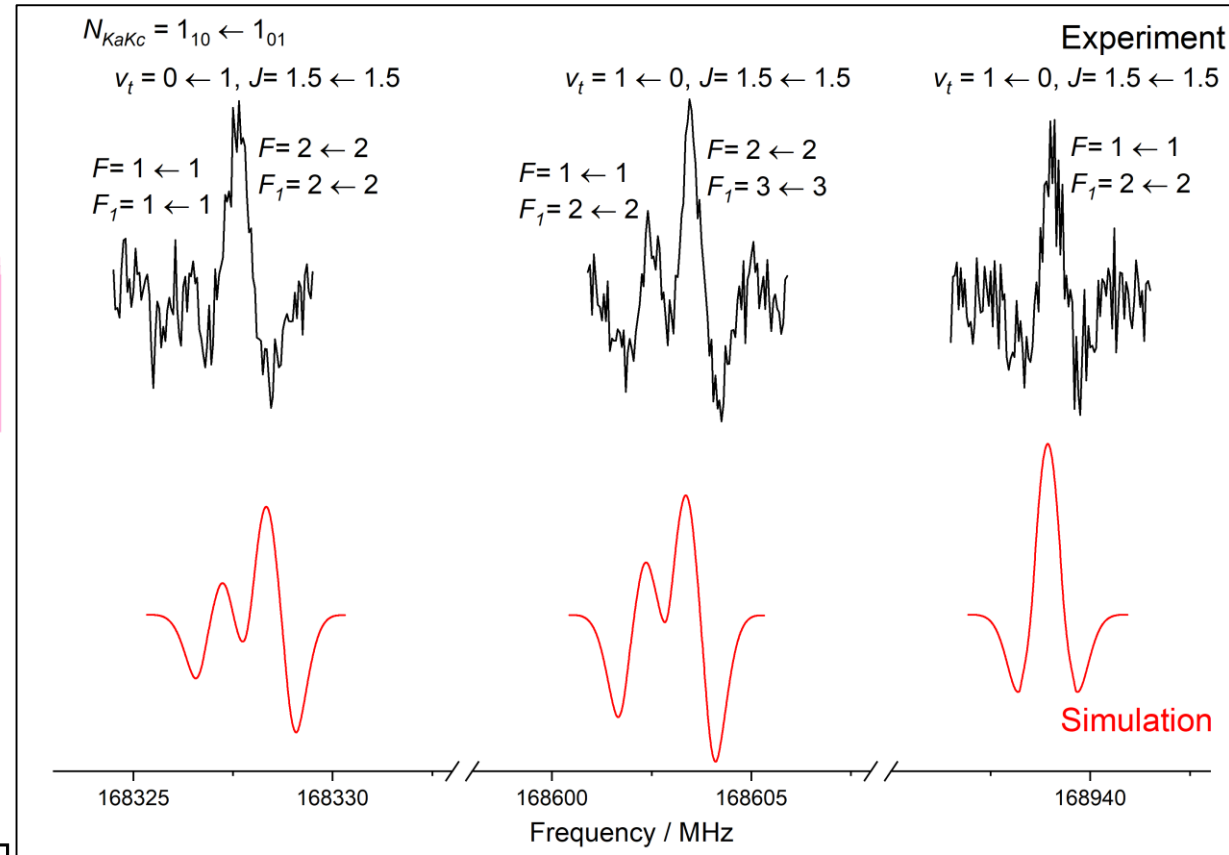
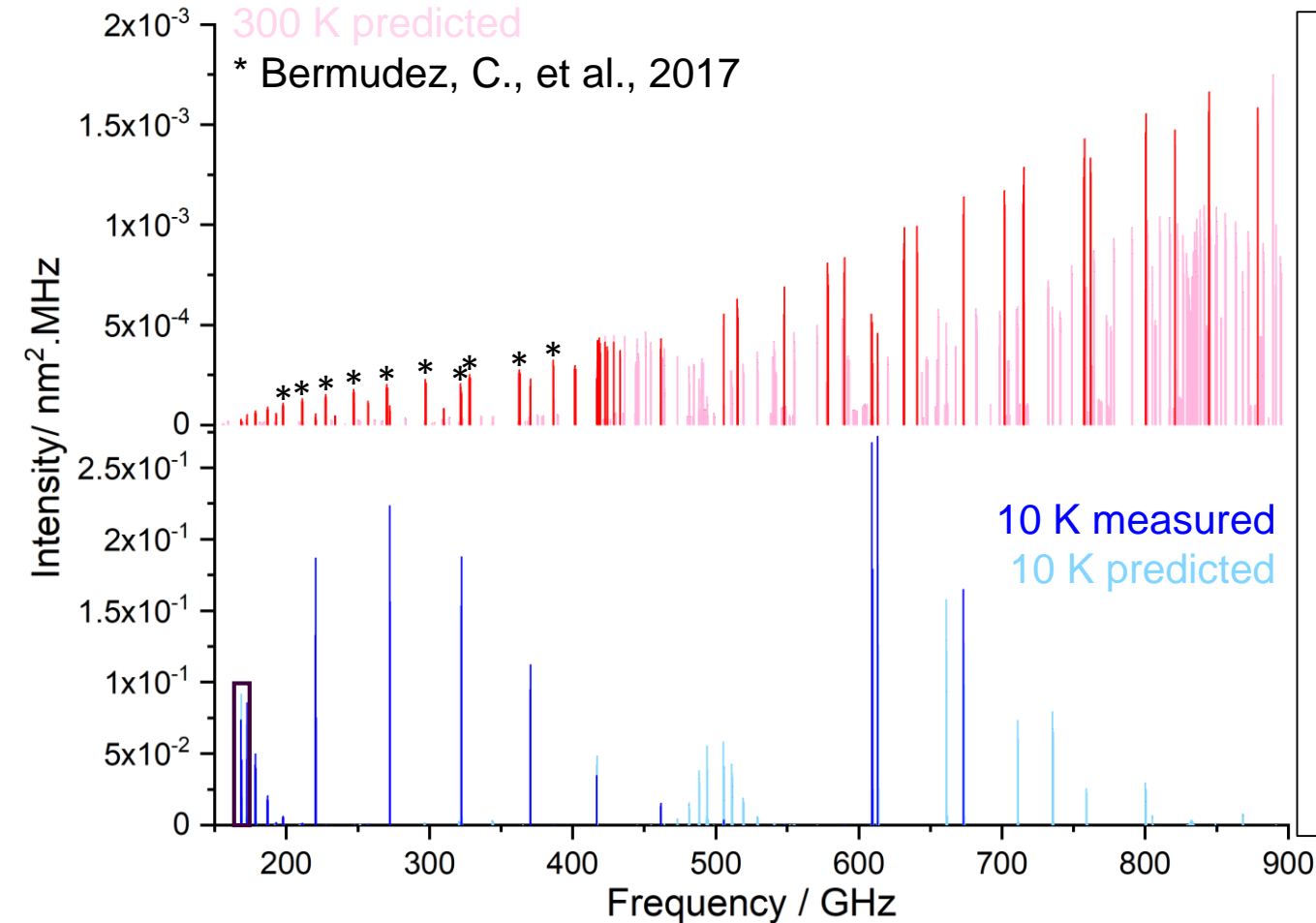
Results and discussion

Astrophysical relevant transitions

300 K measured (Chitarra, O., et al., 2020 + new lines)

300 K predicted

* Bermudez, C., et al., 2017



Perspectives

Our work

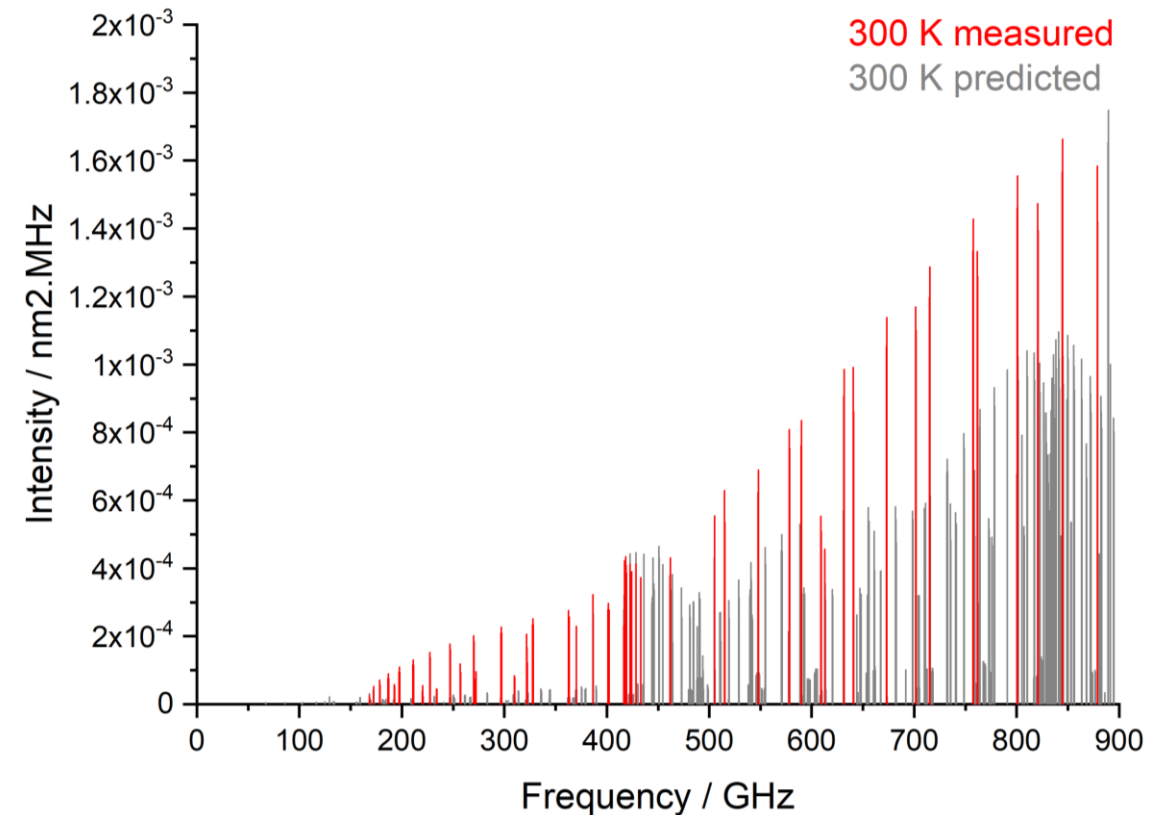
- ❑ More constrained set of rotational parameters

Consequences

- ❑ Better predictions of line positions

Perspectives

- ❑ Toward a more suitable model



Perspectives

Our work

- ❑ More constrained set of rotational parameters
- ❑ Measurement of astrophysical relevant transitions

Consequences

- ❑ Better predictions of line positions
 - ❑ Confident search of the CH_2OH radical in the ISM
- Ongoing observations
(30m IRAM telescope)

Perspectives

- ❑ Toward a more suitable model
- ❑ Confirmation or modification of astrophysical models



Acknowledgments

Collaborators

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Olivier Pirali

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Valerio Lattanzi



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CNRS

