ARC - a novel RICH detector for a future e^+e^- collider
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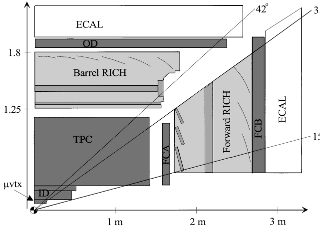
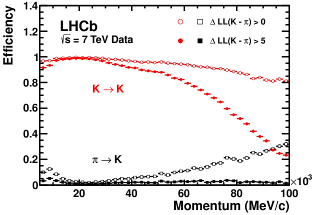
Introduction RICH detectors

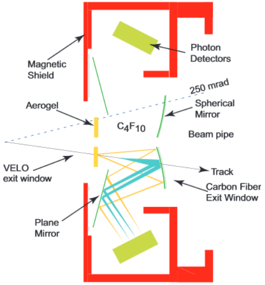
- Excellent hadron PID is crucial in flavour physics
  - Resolve combinatorics and seprate decay modes
- RICH detectors are very powerful for particle ID at high momentum
- At LHCb, π-K separation is excellent up to 100 GeV
- A 4π collider RICH layout was previously used at DELPHI and SLD
  - Challenging because of the space required



(a) DELPHI RICH layout



(b) LHCb RICH performance



(c) LHCb RICH layout

Motivation for RICH at FCC-ee

- FCC-ee will collect 3×10^{12} visible Z boson decays in 4 years
  - Allows for a world-leading flavour physics programme
  - Combined with excellent PID capabilities, FCC-ee will reach an unprecedented precision

- Good PID performance is also required for WW, Higgs and $t\bar{t}$ physics

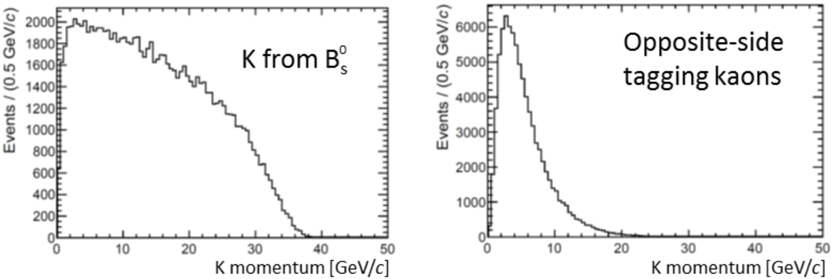


Figure 2: $B_s^0 \to D_s^{\pm} K^{\mp}$

B physics requires pion-kaon separation from low momentum up to 40 GeV

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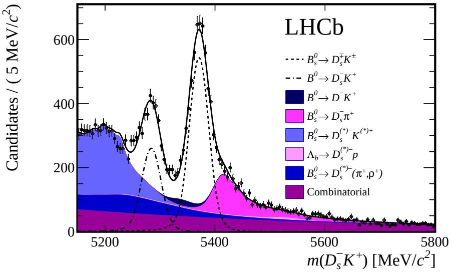


Figure 3: $B_s^0 \to D_s^{\pm} K^{\mp}$

The $B_s^0 \to D_s^{\pm} \pi^{\mp}$ background would be 10 times larger without PID capabilities!

Array of RICH Cells

- **A**rray of **R**I**C**H **C**ells (ARC): A novel RICH detector concept
  - First presented by R. Forty at FCC week 2021
  - Compact, low-mass solution for particle ID for FCC-ee
  - Concept inspired by the compound eyes of an insect
- Adapted to fit into the CLD experiment concept, taking 10% from the tracker volume
  - Radial depth of 20 cm, radius of 2.1 m and a length of 4.4 m
  - Aim to keep material budget below $0.1X_0$
- Aerogel and gas radiators with a spherical mirror
  - Aerogel also acts as thermal insulation between gas and detector

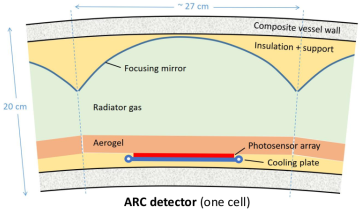


Figure 4: ARC has a cellular structure, similar to an insect's compound eyes

Array of RICH Cells

- All cells are the same size, organised on a hexagonal grid
  - Barrel (endcap) has 945 (384) cells in total, where 18 (21) are unique
  - Hexagonal shape avoids the corners, where performance is worse

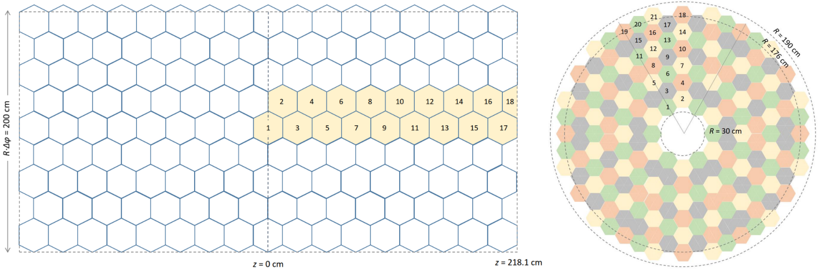
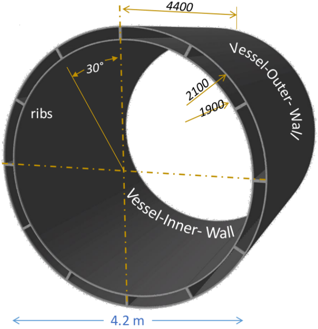


Figure 5: Barrel (left) and endcap (right) cells

Original pressurised ARC

- Original idea by R. Forty was a vessel with pressurised gas
  - Higher photon yield in smaller radial space
- However, a non-pressurised gas was found to have better performance, and also simplify the vessel design



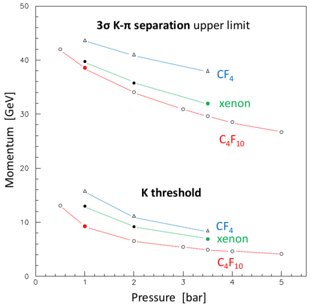
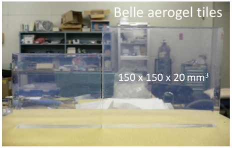


Figure 6: Layout of original carbon fiber vessel (left) and the performance at different pressures (right)

ARC radiators

- C_4F_{10}:
  - Baseline assumption, well known from LHCb RICH1
  - $n = 1.0014 \implies \theta_c = 53\,\text{mrad}$, suitable for high momentum particles
  - C_4F_{10} is a greenhouse gas, substitution with pressurised Ar/Xe possible
- Aerogel:
  - Well known as a RICH radiator, e.g. from ARICH at Belle II
  - $n = 1.01\text{-}1.10 \implies \theta_c = 141\text{-}430\,\text{mrad}$, suitable at low momentum
  - Very low thermal conductivity
    - Suitable to separate gas from detector, which must be cooled
    - Cherenkov photons come for "free" and are focused by the same mirror
  - Drawback: Some loss of photons from scattering



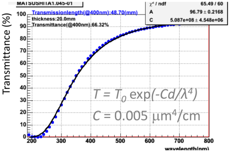


Figure 7: Belle aerogel tiles (left) and aerogel transmission function (right).

Event displays

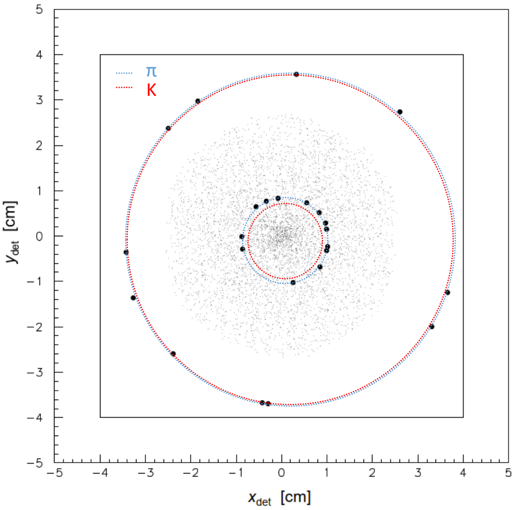


Figure 8: Event display: Photon hits on photodetector

Event displays

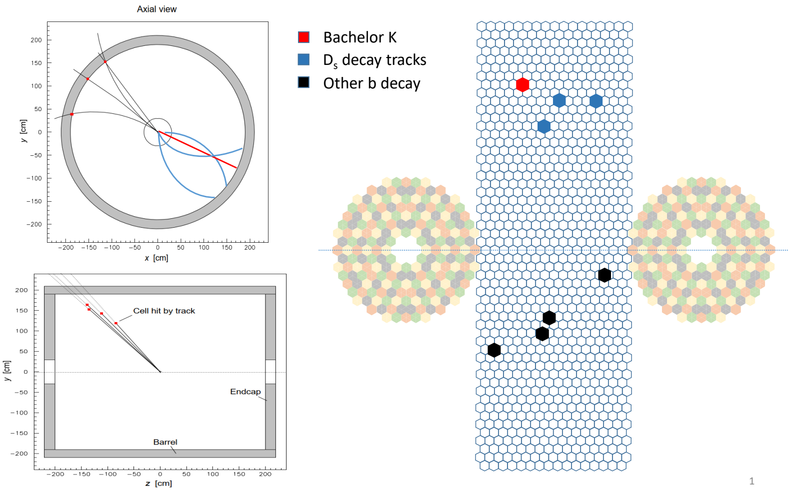


Figure 9: Event display: $B_s \to D_S K$ event

Optimisation of ARC layout

- The following procedure is used to evaluate the ARC performance:
  1. Generate straight particle track from IP and trace it through ARC
  2. Generate Cherenkov photons from gas radiator
  3. Track photons through the optics and to detector
  4. Reconstruct Cherenkov angles and calculate standard deviation
- Two sources of uncertainty are considered:
  1. Emission point uncertainty: Emission point is assumed to be the mid-point of the track inside the gaseous radiator
  2. Chromatic dispersion uncertainty: Spread in Cherenkov angle due to wavelength dependence in refractive index

Minimise the Cherenkov angle uncertainty:

$$\Delta\theta = \frac{1}{\sqrt{N}} \times \frac{1}{1-N} \times \sum_{i=0}^{N-1}(\theta - \bar{\theta})^2$$

Examples of photon tracking through optimised layout

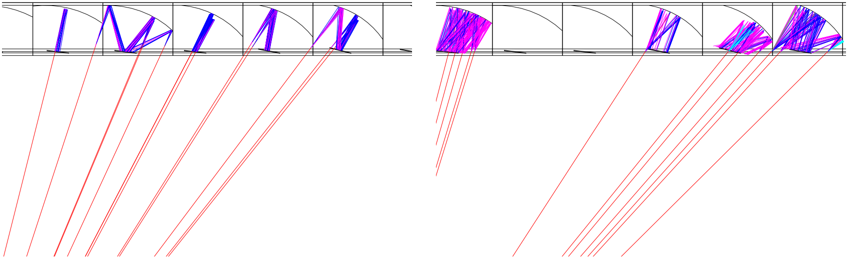
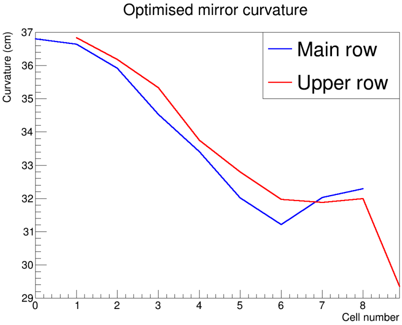


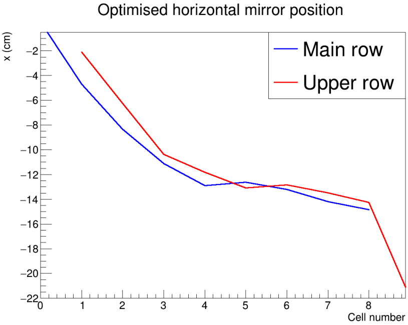
Figure 10: Tracking of photons from gas radiator (left) and aerogel radiator (right) through the ARC optics

- Parameters that are optimised:
    - Mirror curvature
    - Mirror vertical and horizontal position
    - Detector horizontal position and tilt

Optimised mirror curvature and mirror position



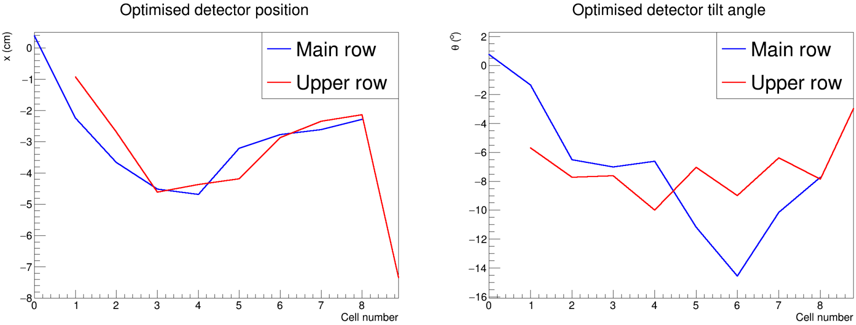
(a) Mirror curvature



(b) Mirror position

Figure 11: Optimised mirror parameters

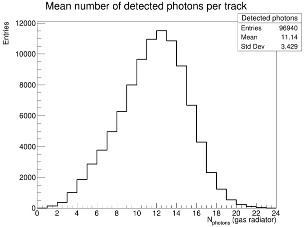
Optimised detector position and detector tilt angle



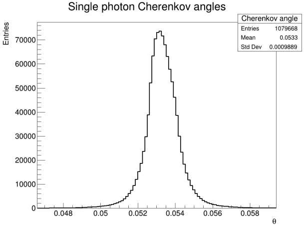
(a) Detector position **(b)** Detector tilt angle

Figure 12: Optimised detector parameters

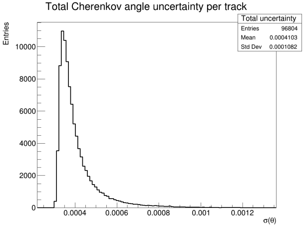
Cherenkov angle uncertainty for gas radiator



(a) Mean number of photons detected

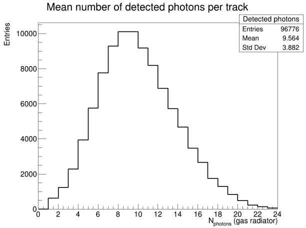


(b) Single photon emission point uncertainty



(c) Total emission point uncertainty

Cherenkov angle uncertainty for aerogel radiator



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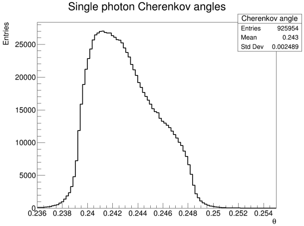

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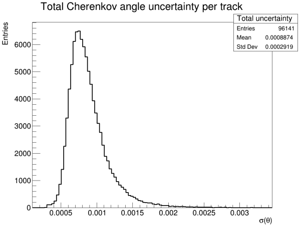

(c) Total emission point uncertainty

Performance of optimised ARC

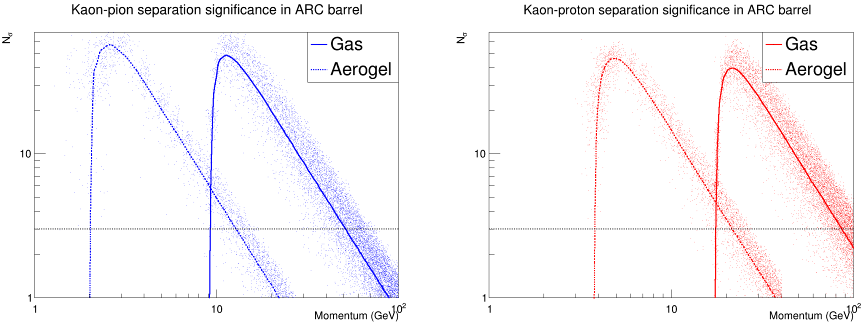

Figure 15: Separation significance per track for π-K (left) and p-K (right)

- Gas (aerogel) provides over 3σ pion-kaon separation in the range 10-50 GeV (2-10 GeV)
  - Improved from earlier studies due to gaining some space for more radiator by no longer pressurising, as well as optimisation of the layout
  - Magnetic field not yet included in these studies
- Combined, the aerogel and gas ensure great PID performance over the whole range of interest to flavour physics

Summary and next steps

- ARC is a low mass and compact cellular PID detector designed to occupy minimum space (e.g. 20 cm in barrel) in a 4π detector at an e^+e^- collider such as FCC-ee
- We have developed an optimised layout that achieves a 3σ kaon-pion separation in the range 2-50 GeV, which is useful for a wide range of flavour physics measurements
    - Our studies focus mainly on flavour physics at the Z-pole
- ARC will allow us to fully exploit the flavour physics potential of future e^+e^- colliders
    - Will enhance the capabilities in WW, Higgs and top physics
- Next steps will include better optimisation, R&D of photodetectors and studies of magnetic field effects

Thanks for your attention!

Backup: Material budget breakdown

Units of radiation length X/X_0

Detector component	Pressurised	Non-pressurised
Vessel walls	5%	1%
Photosensor array/electronics	1%	1%
Cooling plate (3 mm CF)	1%	1%
Aerogel ($n = 1.03$)	5%	0.5%
C_4F_10 gas	5%	0.5%
Focusing mirror	1%	1%
Total	10%	5%

Backup: Technical details about minimisation

- $f(\vec{x})$ is not easily to calculate analytically
- Approximate by simulating a large number of charged tracks
- Finite number of photons $\implies f(\vec{x})$ is not differentiable
  - Cannot be minimised using conventional methods (Minuit, etc)
- I have experimented with a new type of minimisation algorithms: Stochastic optimisation
  - Differential evolution
  - Start with a population of possible solutions, form new solutions by combining (mutating) existing solutions
  - Advantage: Doesn't require initial guess, robust against functions that a not continuous, noisy, change over time, etc
  - Disadvantage: No way to tell if optimal solution has been found, so it requires many iterations