

Cosmic kite: Auto-encoding the cosmic microwave background

<https://arxiv.org/abs/2202.05853>

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Outline

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 - b. Boltzmann-Einstein Equations
 - c. The Cosmic Microwave Background (CMB)
 - d. Our Universe
2. Brief introduction to Machine Learning
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Brief introduction to Cosmology

The Friedman-Lemaître-Robertson-Walker Metric

$$ds^2 = -dt^2 + a^2(t) \left[\frac{dr^2}{1 - kr^2} + r^2 d\Omega^2 \right]$$

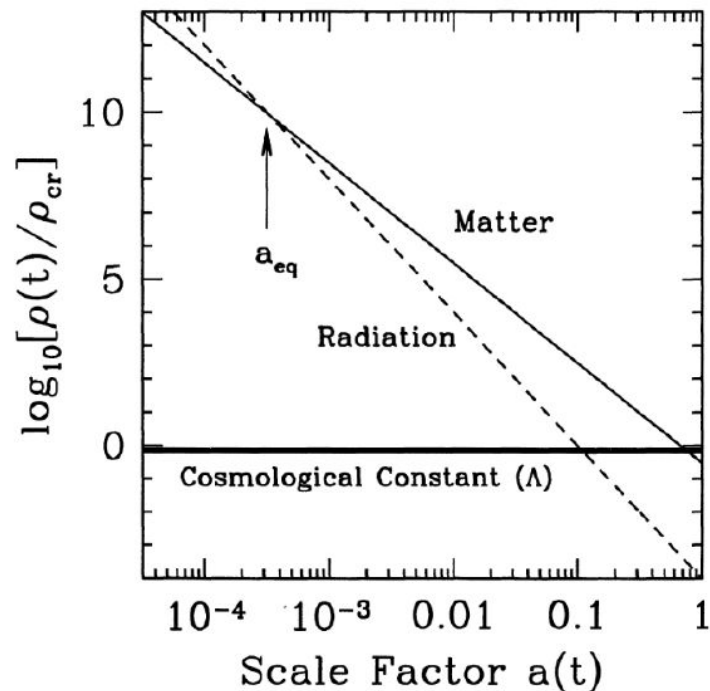
$$\left(\frac{\dot{a}}{a} \right)^2 = \frac{8\pi G}{3} \rho - \frac{k}{a^2}$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3p)$$

$$\frac{\dot{\rho}}{\rho} = -3(1 + \omega) \frac{\dot{a}}{a}$$

$$\rho \propto a^{-3(1+\omega)}$$

$$T = \frac{T_0}{a}$$



Modern Cosmology. Dodelson

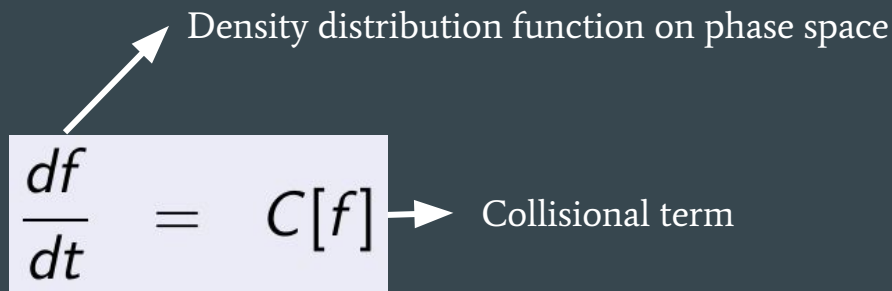
The Friedman-Lemaître-Robertson-Walker Metric + Perturbations (Newtonian Gauge)

$$ds^2 = a^2(\tau)[-(1 + 2\psi)d\tau^2 + (1 - 2\phi)dx_id x^i]$$

The Friedman-Lemaître-Robertson-Walker Metric + Perturbations (Newtonian Gauge)

$$ds^2 = a^2(\tau)[-(1 + 2\psi)d\tau^2 + (1 - 2\phi)dx_i dx^i]$$

The Boltzmann Equations


$$\frac{df}{dt} = C[f]$$

Density distribution function on phase space

Collisional term

The Boltzmann-Einstein Equations

Photons

$$\dot{\Theta} + ik\mu\Theta = -\dot{\phi} - ik\mu\psi - \dot{\tau} \left[\Theta_0 - \Theta + \mu v_b - \frac{1}{2} P_2(\mu) \Pi \right]$$

Baryons

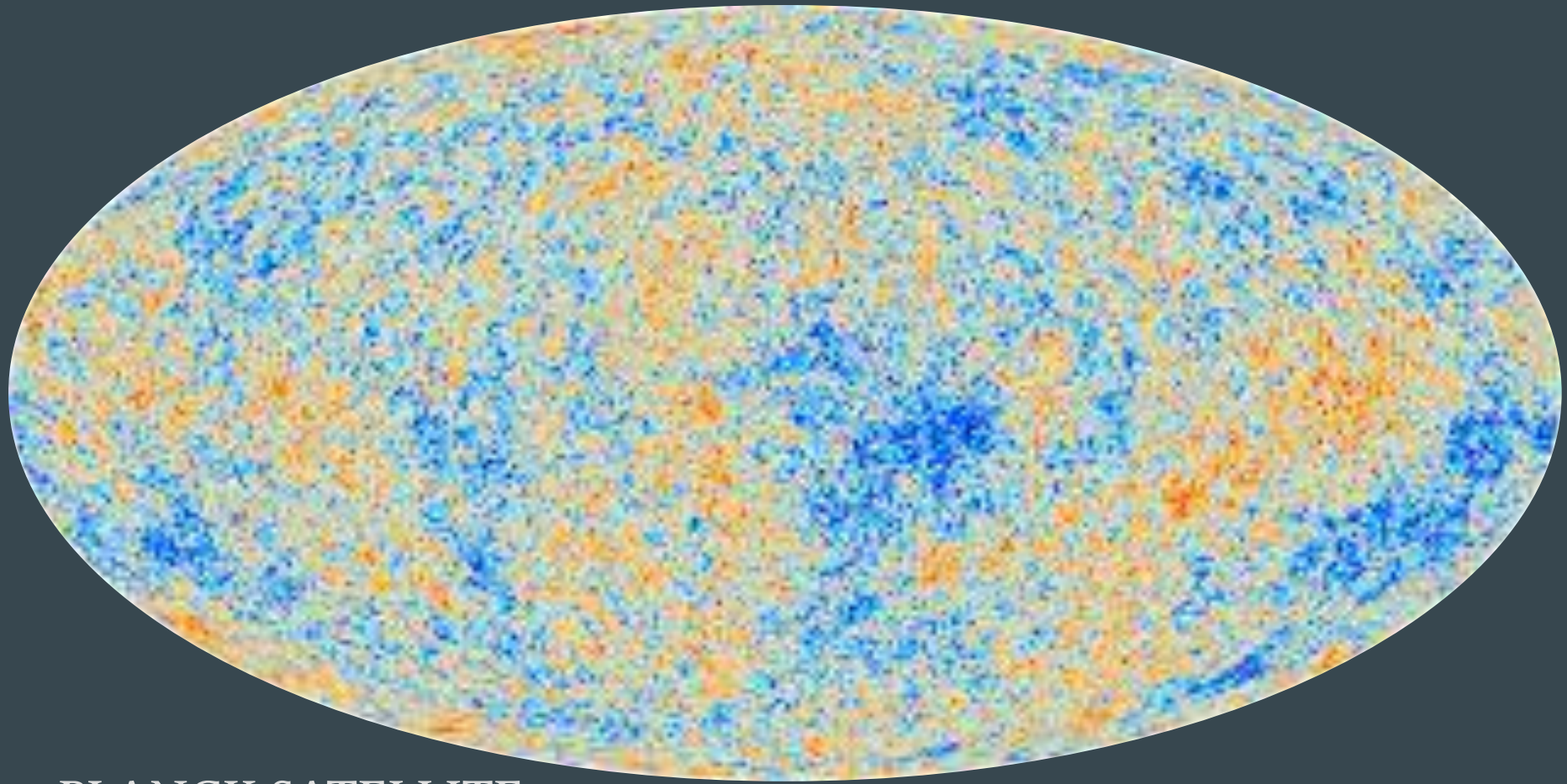
$$\begin{aligned}\dot{\delta}_b + ikv_b &= -3\dot{\phi} \\ \dot{v}_b + \frac{\dot{a}}{a}v_b &= -ik\psi + \frac{\dot{\tau}}{R}[v_b + 3i\Theta_1]\end{aligned}$$

Dark Matter

$$\begin{aligned}\dot{\delta} + ikv &= -3\dot{\phi} \\ \dot{v} + \frac{\dot{a}}{a}v &= -ik\psi\end{aligned}$$

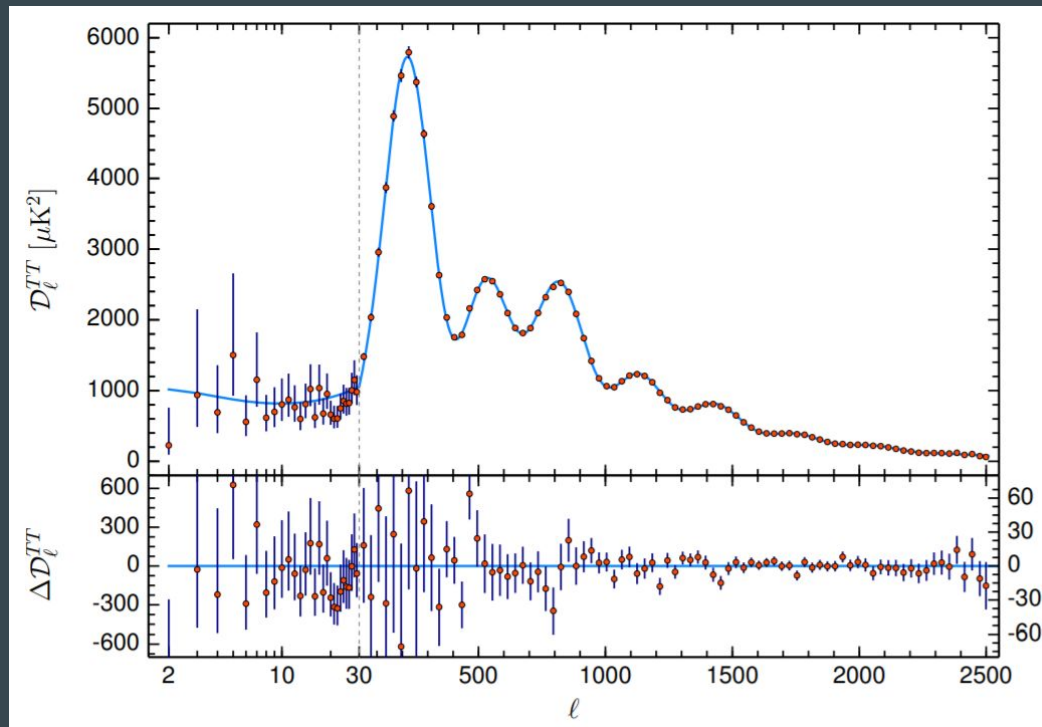
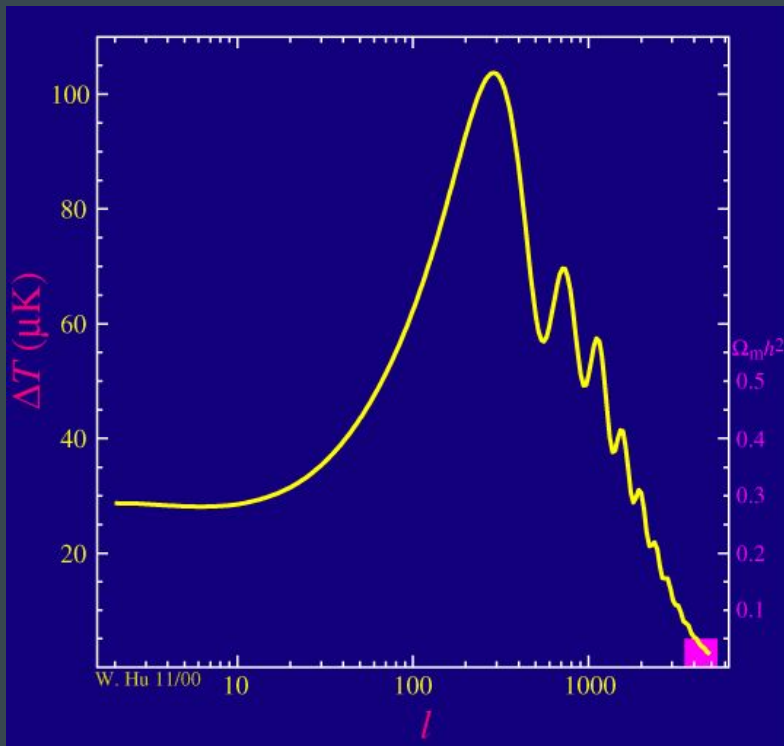
Neutrinos

$$\dot{N} + ik\mu N = -\dot{\phi} - ik\mu\psi$$



PLANCK SATELLITE

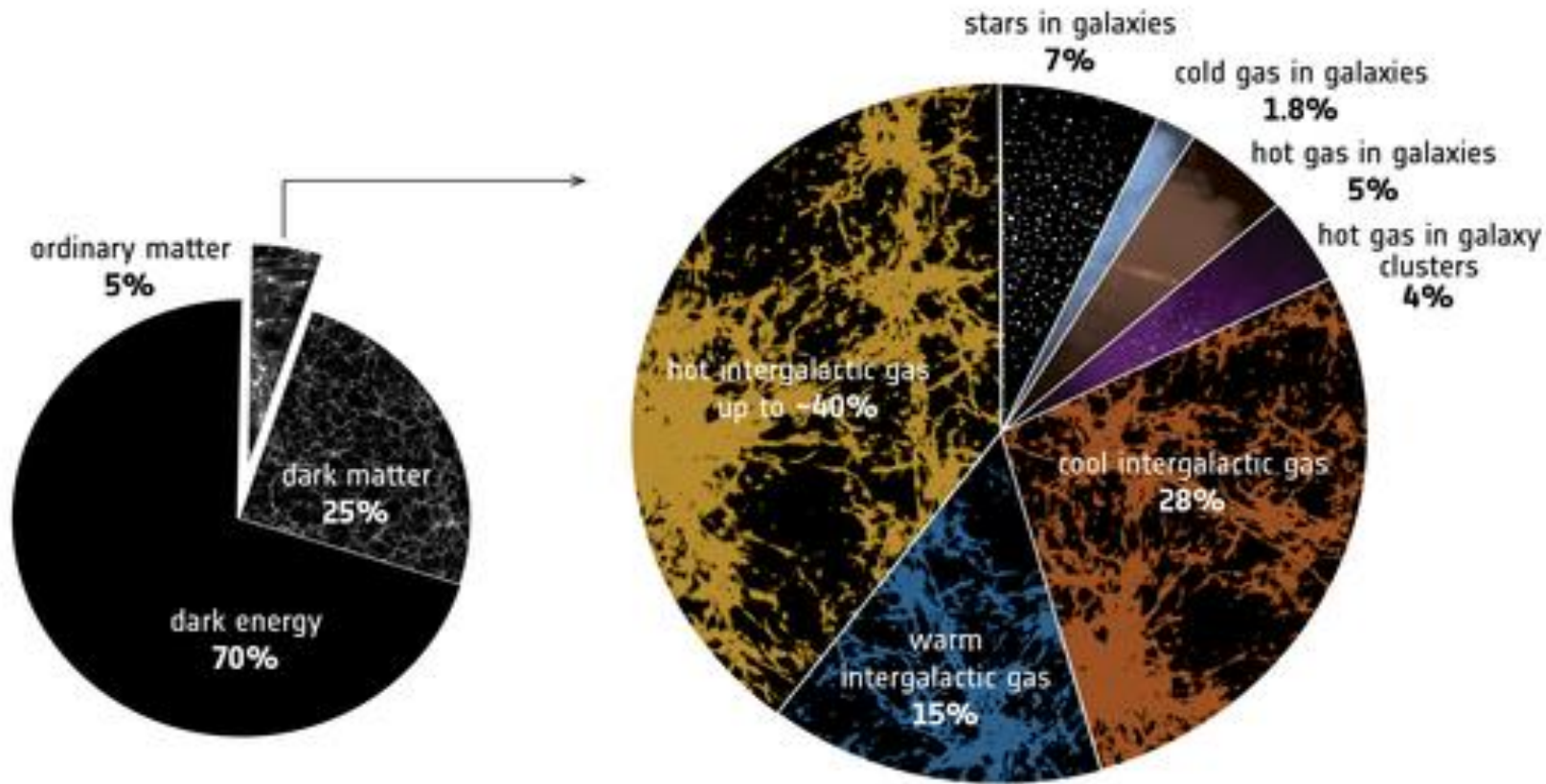
The Cosmic Microwave Background (CMB)



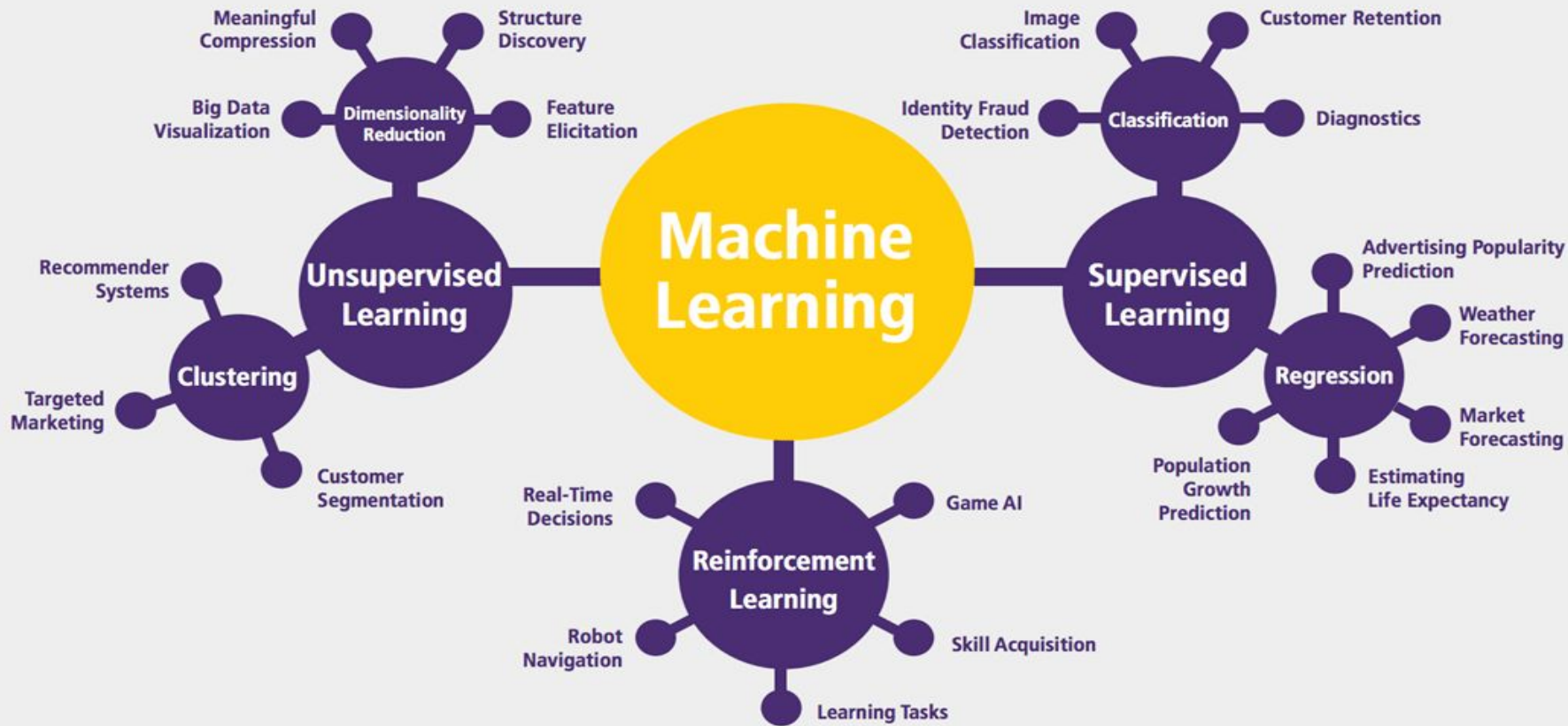
Planck Collaboration 1807.06209

<http://background.uchicago.edu/~whu/animbut/anim2.html>

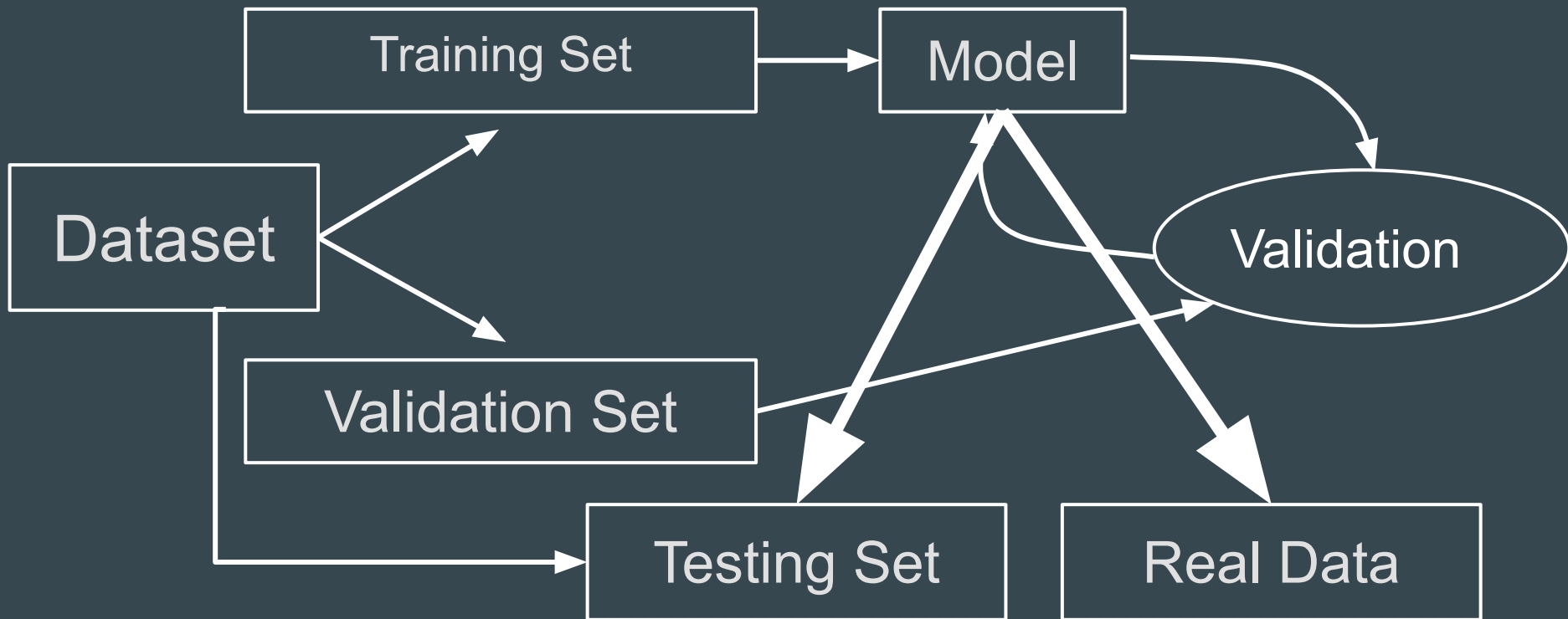
Our Universe



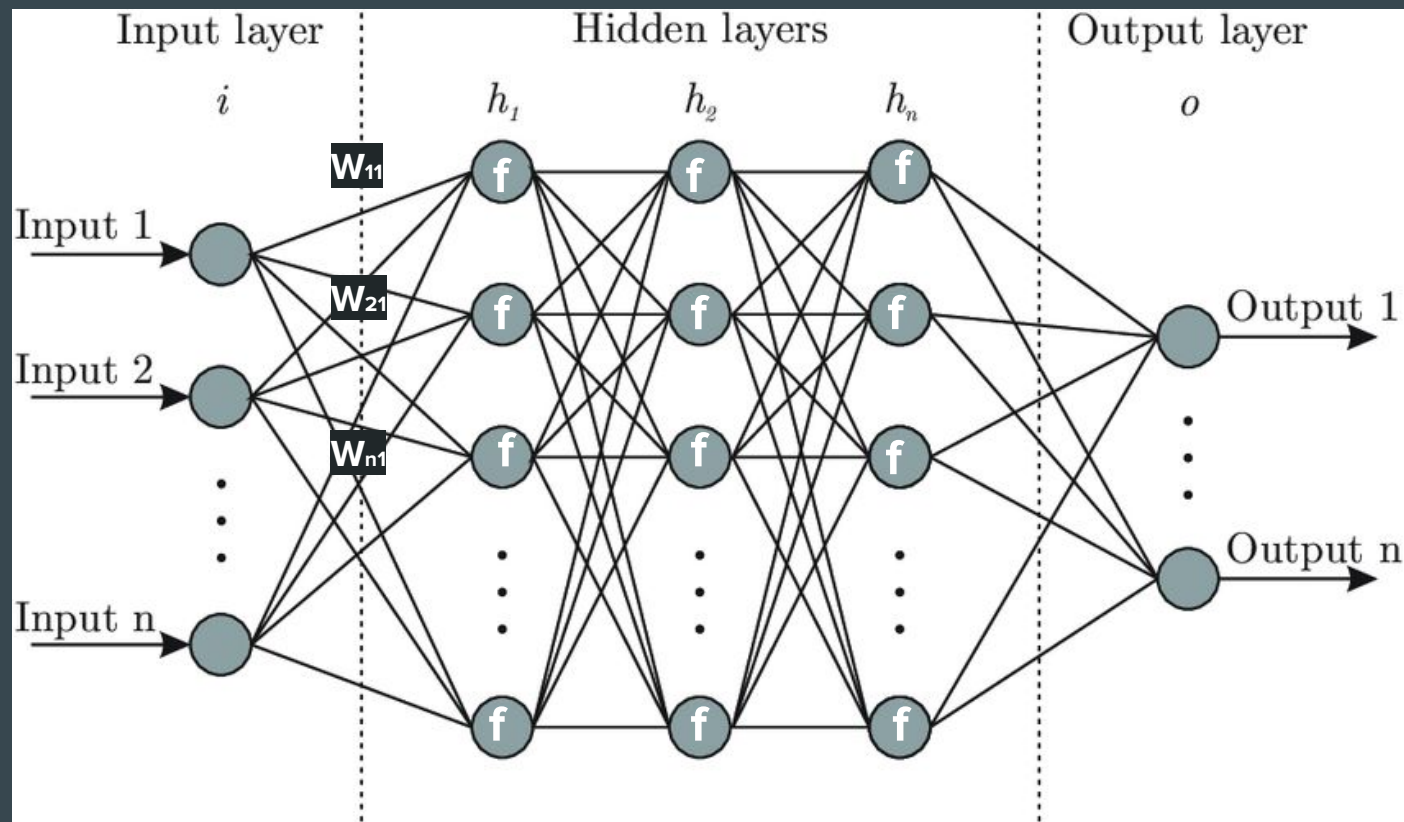
Brief introduction to Machine Learning



Supervised Learning

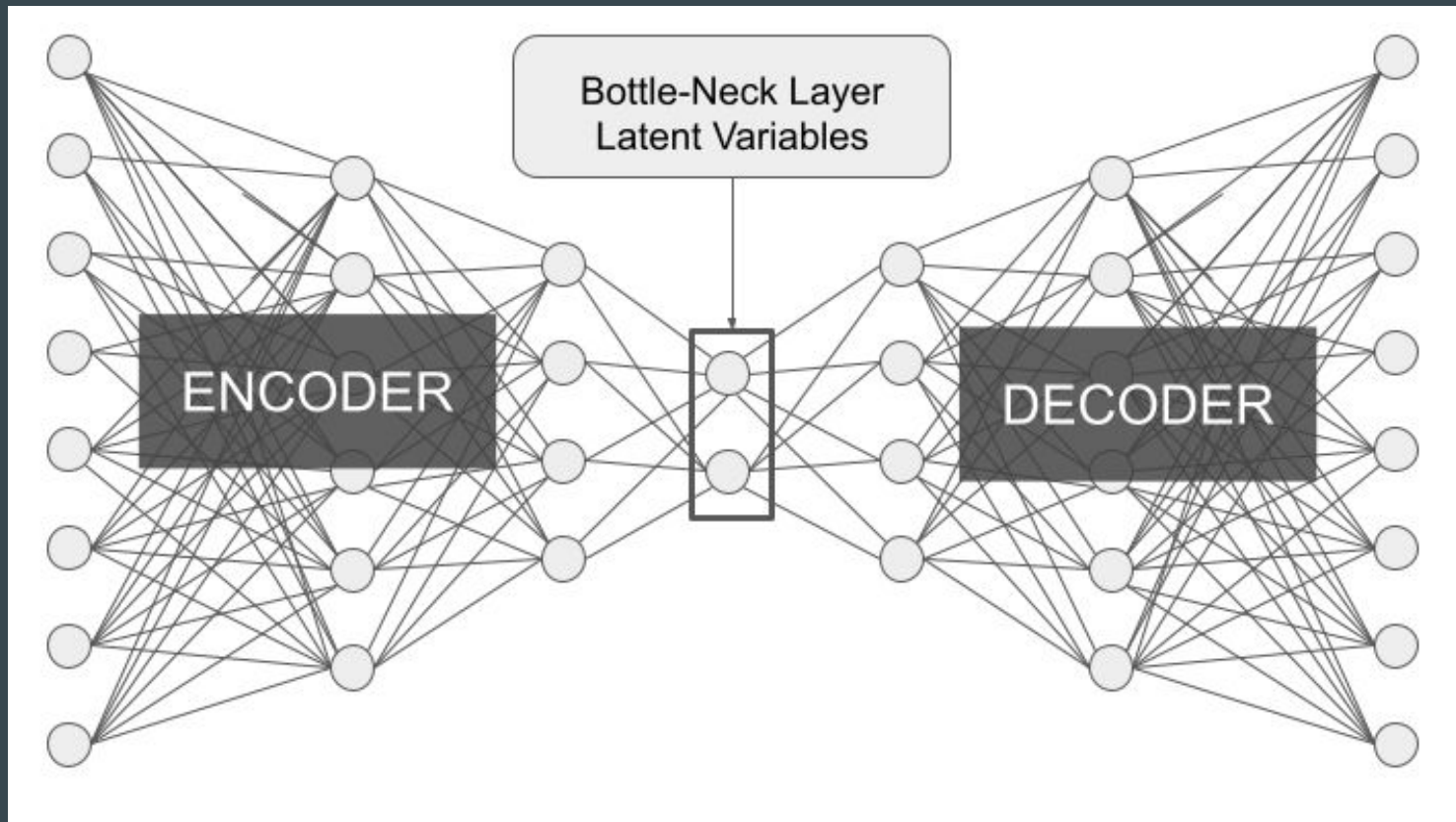


Neural Networks

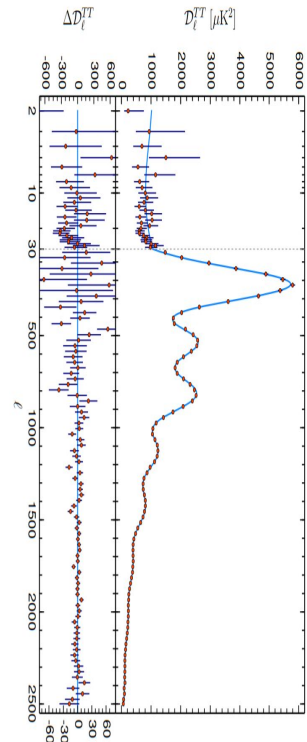
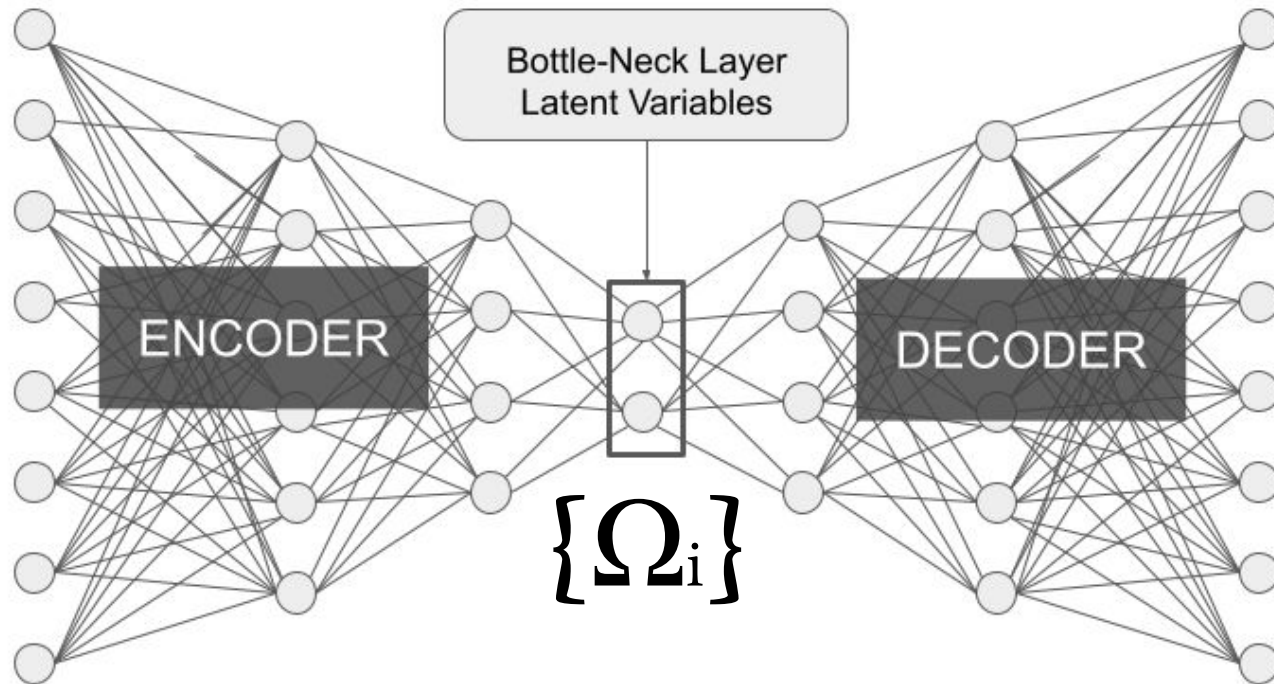
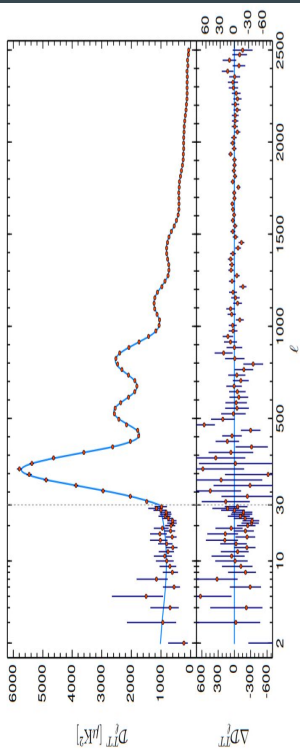


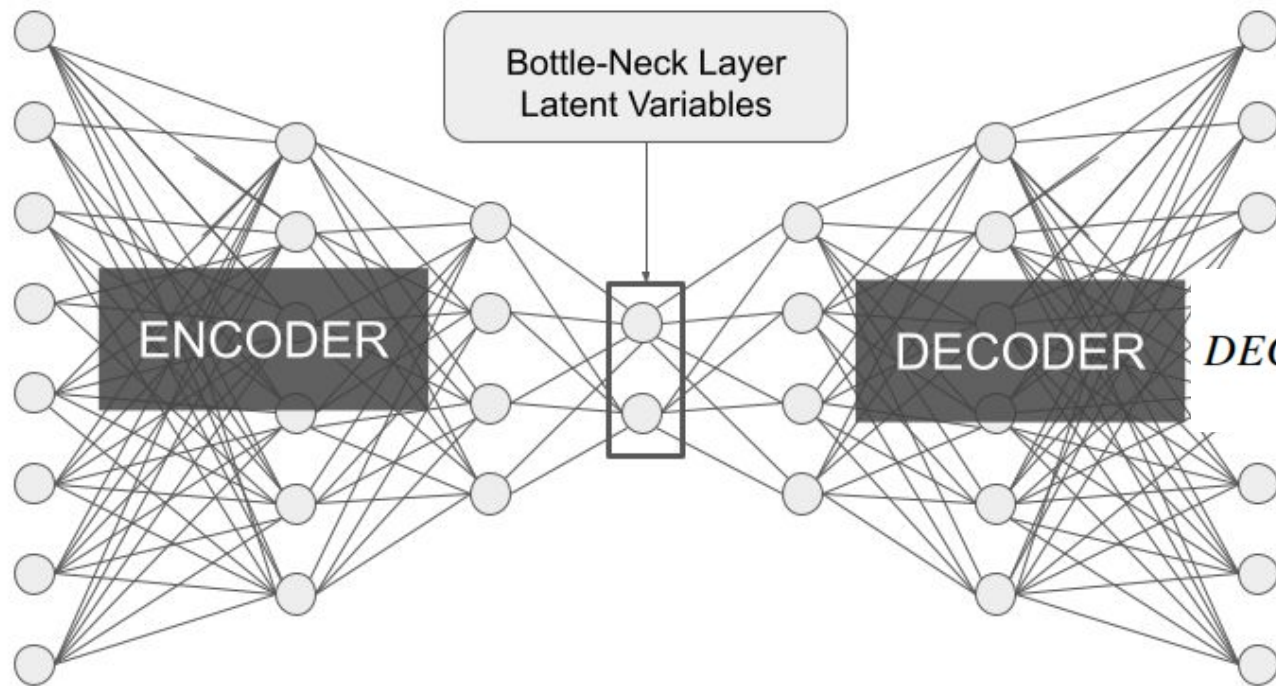
$\text{LOSS}(\text{Pr}, \text{Re})$

Auto-Encoders

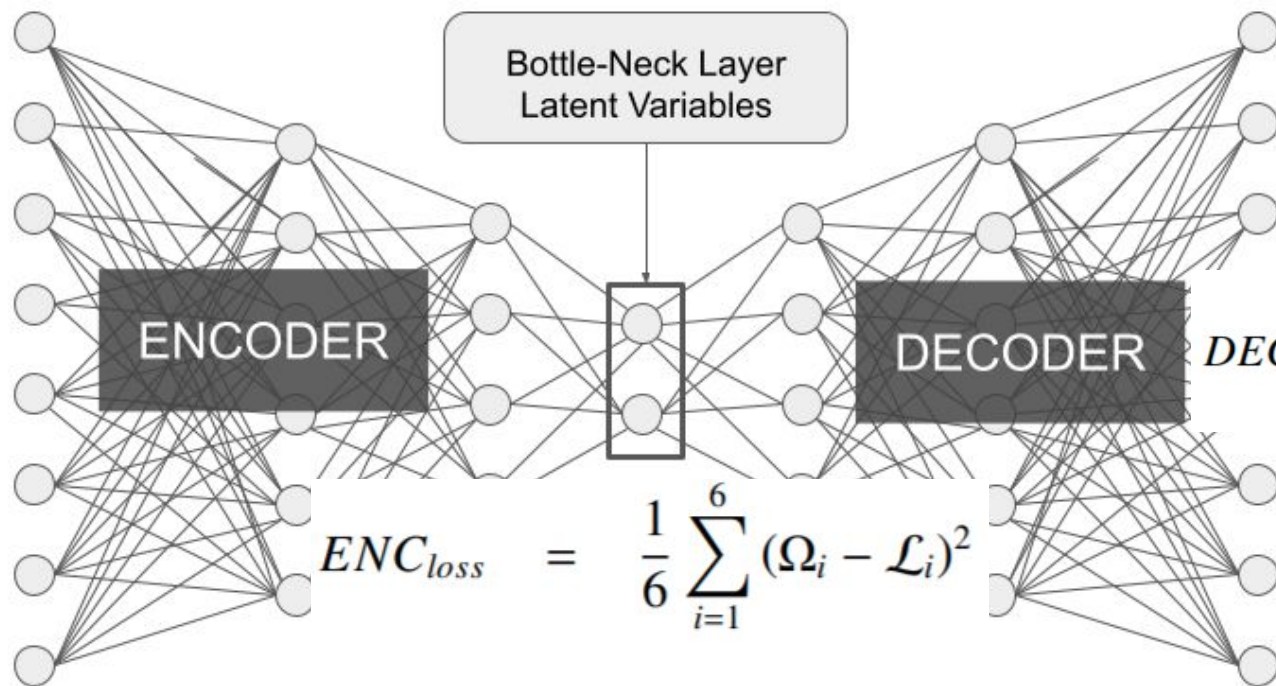


Cosmic-Kite





$$DEC_{loss} = \frac{1}{2648} \sum_{i=2}^{2650} (C_i^R - C_i^P)^2$$



$$LOSS = \frac{1}{N} \sum_{j=1}^N [\omega_{enc} ENC_{loss,j} + \omega_{dec} DEC_{loss,j}]$$

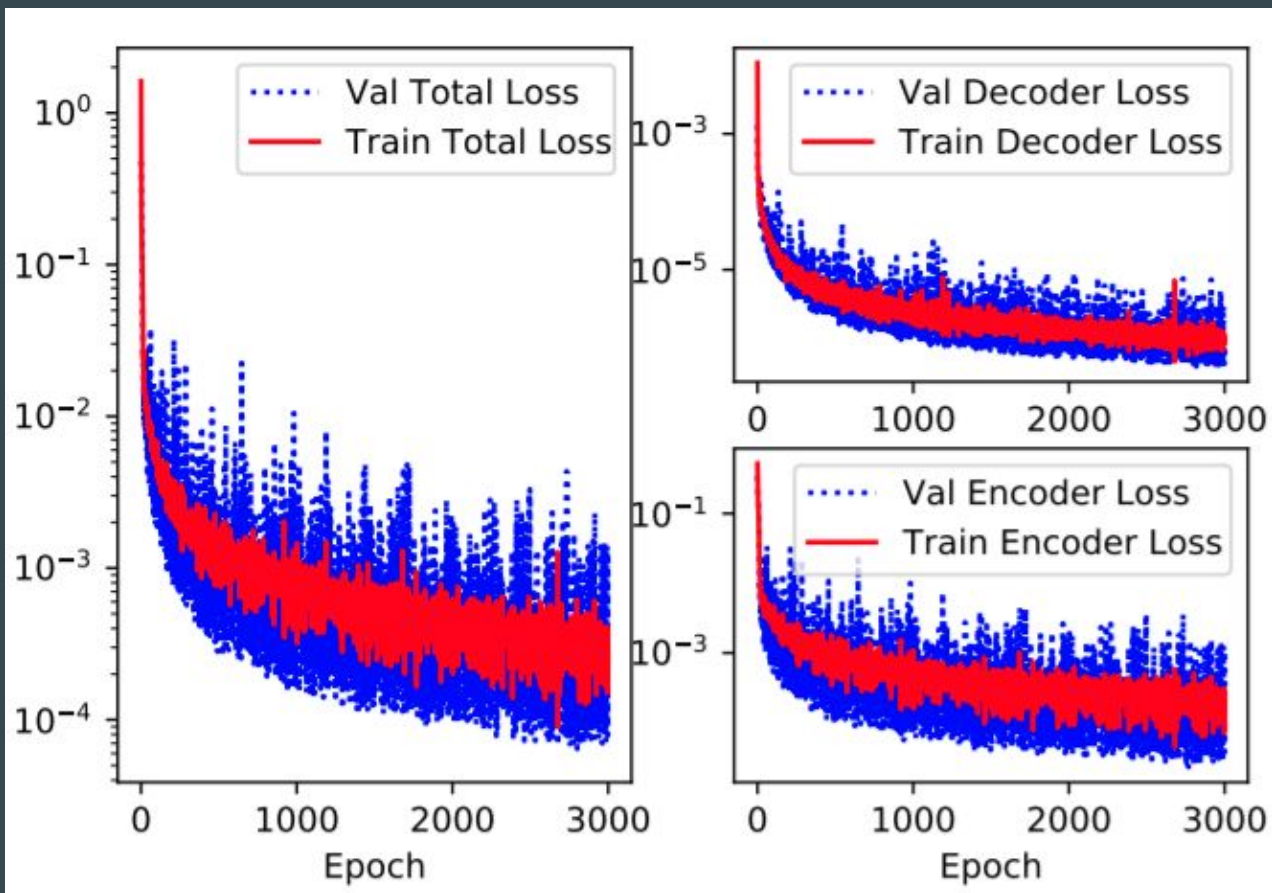
Building of the dataset

- CAMB¹ (Lewis et al.)
- 6 Cosmological Parameters
- 80.000 random cosmologies

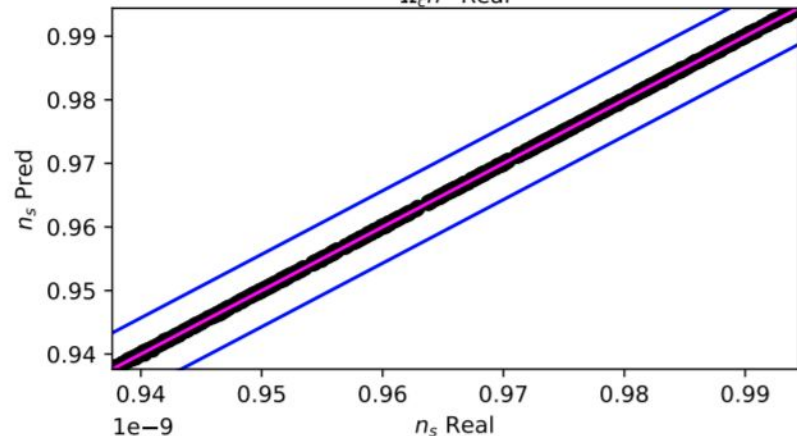
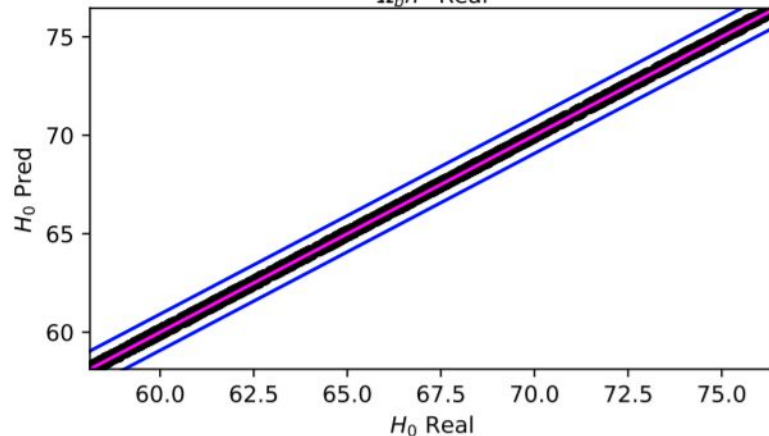
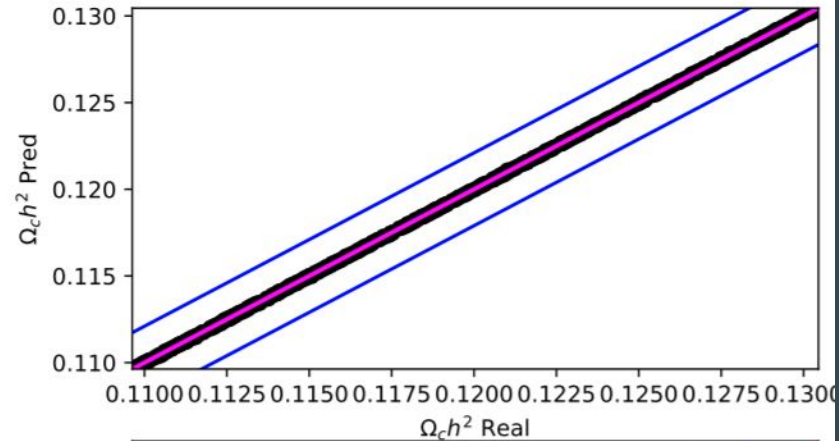
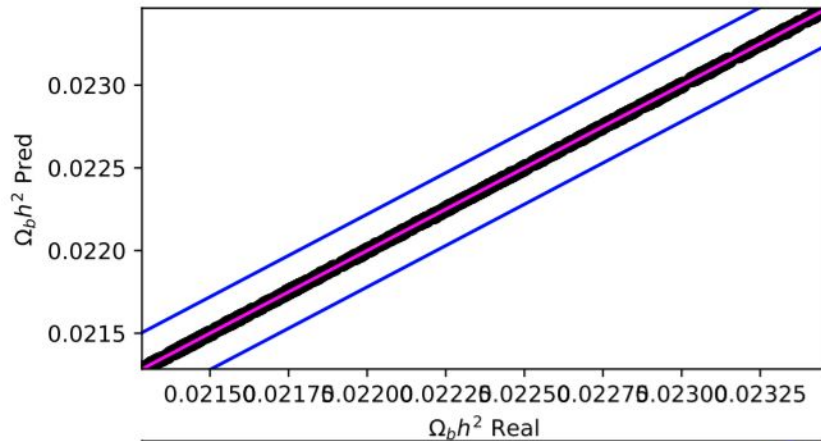
Parameter	Minimum	Maximum	Planck
$\Omega_c h^2$	0.1096	0.130	0.120
$\Omega_b h^2$	0.02128	0.02348	0.02237
H_0	58.12	76.52	67.36
n	0.9375	0.9945	0.9649
A_s	$1.930 * 10^{-9}$	$2.270 * 10^{-9}$	$2.098 * 10^{-9}$
τ	0.014	0.094	0.0544

¹ <https://camb.info/>

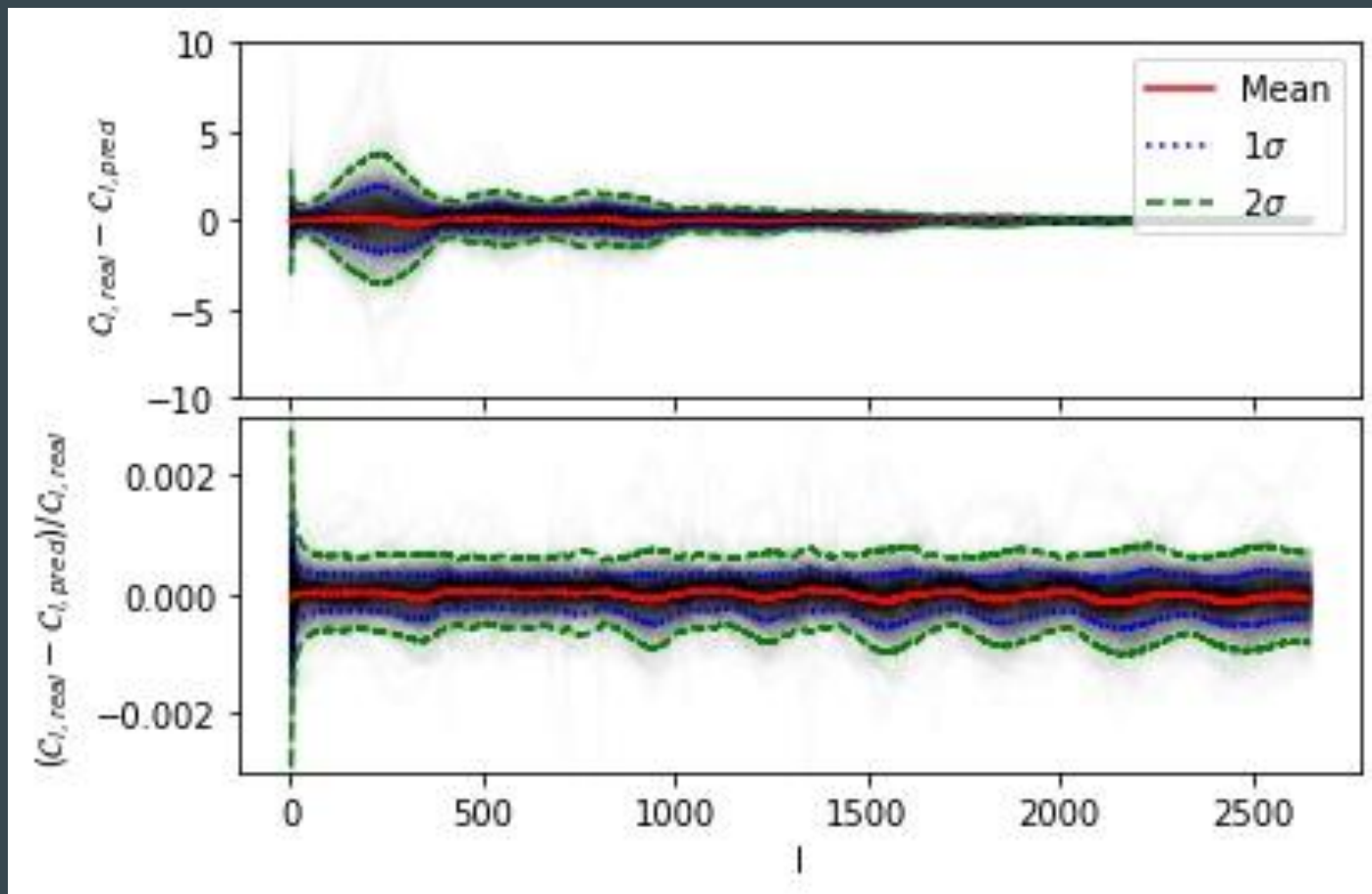
Training

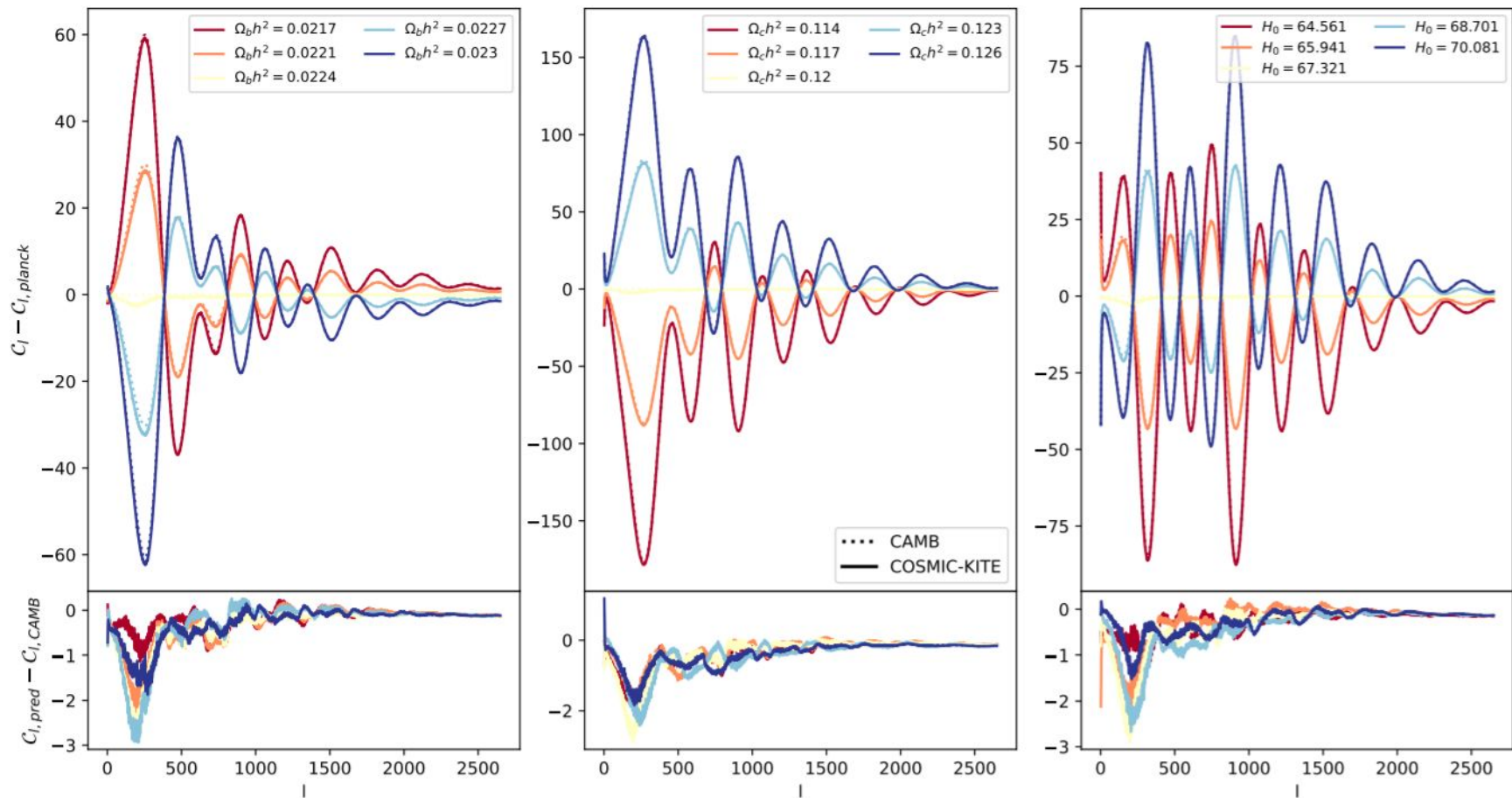


Encoding the CMB

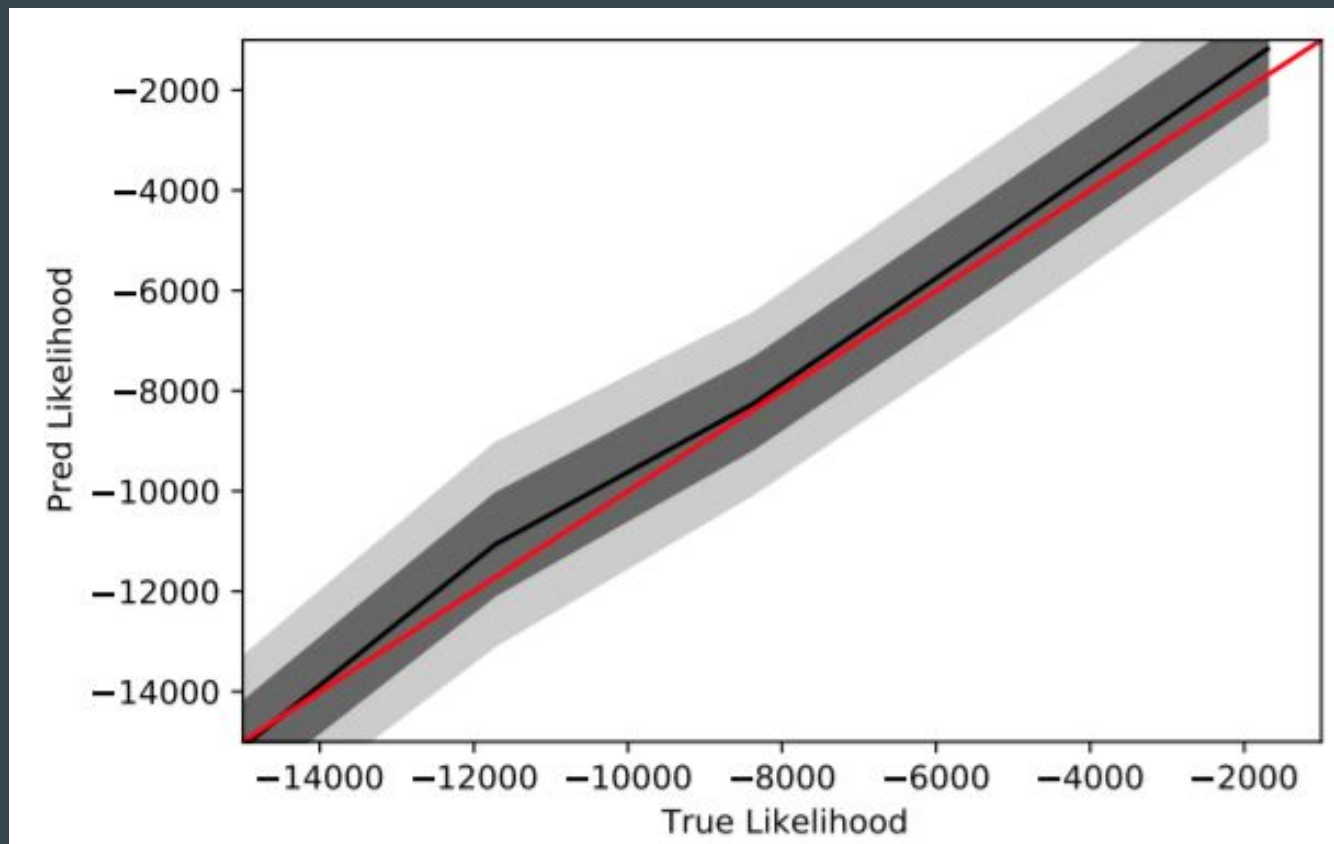


Decoding the CMB

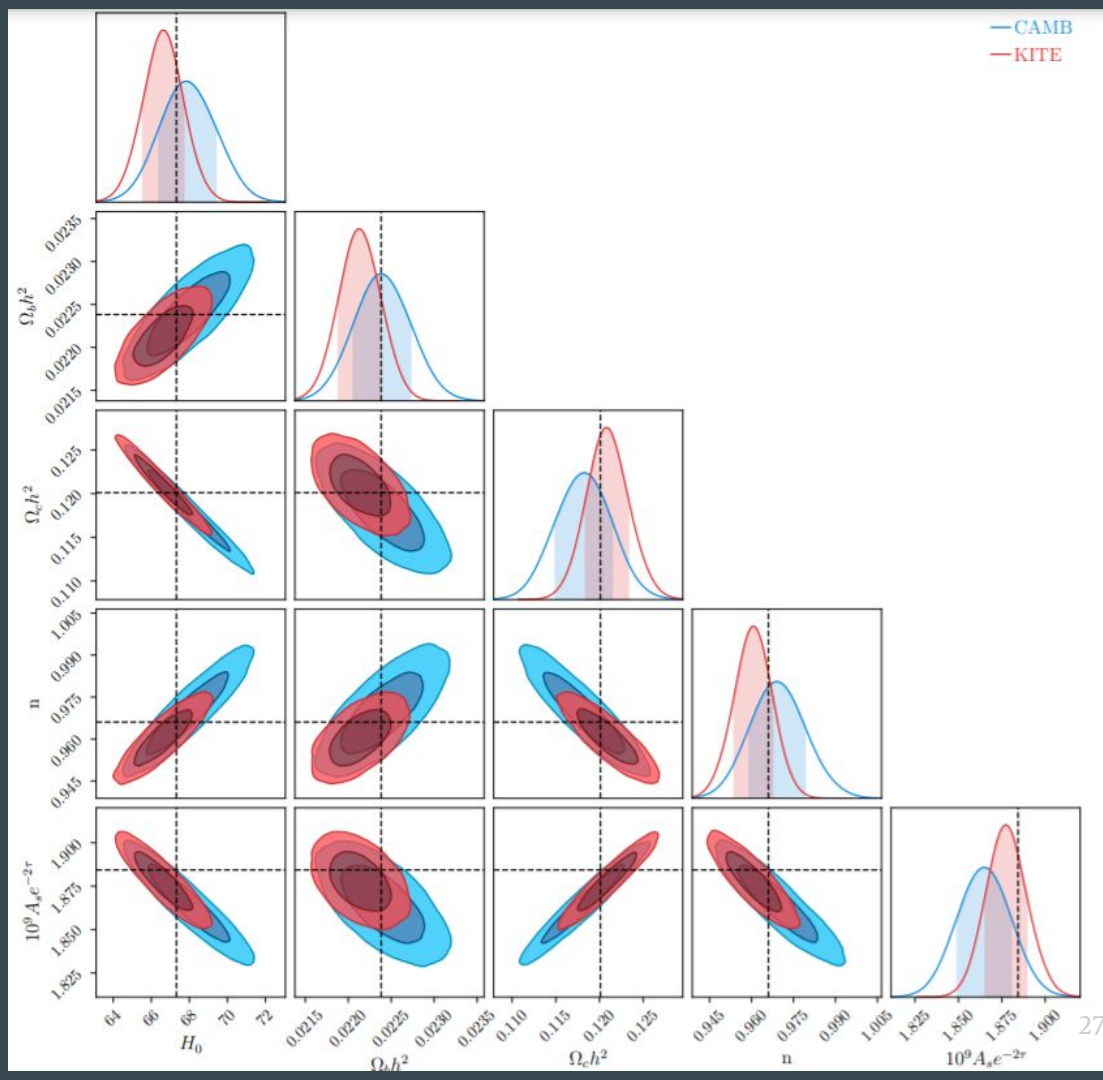




Likelihood Analysis



Bayesian Inference



```
from cosmic_kite import cosmic_kite

H0_true  = 67.32117
omb_true = 0.0223828
omc_true = 0.1201075
n_true   = 0.9660499
tau_true = 0.05430842
As_true  = 2.100549e-9

true_pars = np.array([omb_true, omc_true, H0_true, n_true, tau_true, As_true]).reshape(1, -1)

# The input of the pars2ps function must be an array of shape (n, 6)
#  where n is the number of cosmological models to be computed

ps = cosmic_kite.pars2ps(true_pars)[0]

# The input of the ps2pars function must be an array of shape (n, 2450)
#  where n is the number of cosmological models to be computed
pred_pars = cosmic_kite.ps2pars(ps.reshape(1, -1))[0]
```

Conclusions

Conclusions

- We performed an auto-encoding analysis of the CMB power spectra.
- Using the encoder, we are able to predict the cosmological parameters from the power spectra with an error $\sim 0.2\%$.
- Using the decoder we are able to predict the power spectra from the cosmological parameters with a mean error $\sim 0.0018\%$.
- Although this algorithm does not improve the precision of the measurements compared with the traditional methods, it reduces significantly the computation time.
- Represents the first attempt (to my knowledge) towards forcing the latent variables to have a physical interpretation.
- It can be extended to other signals.

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$



Posterior

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$



Posterior



Likelihood

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$

The diagram illustrates the components of the Bayesian formula. An arrow points from $P(\Omega|X)$ to the word "Posterior". Another arrow points from $P(X|\Omega)$ to the word "Likelihood". A third arrow points from $P(\Omega)$ to the word "Prior".

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$

Diagram illustrating the relationship between the equation and its components:

- $P(\Omega|X)$ is labeled **Posterior**.
- $P(X|\Omega)$ is labeled **Likelihood**.
- $P(\Omega)/P(X)$ is labeled **Evidence**.
- $P(\Omega)$ is labeled **Prior**.

Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$

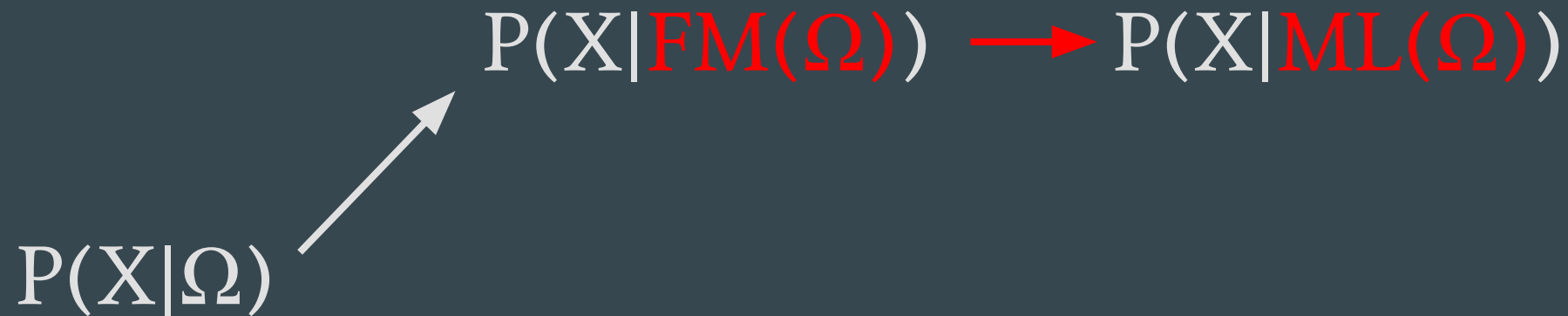
Posterior

Likelihood

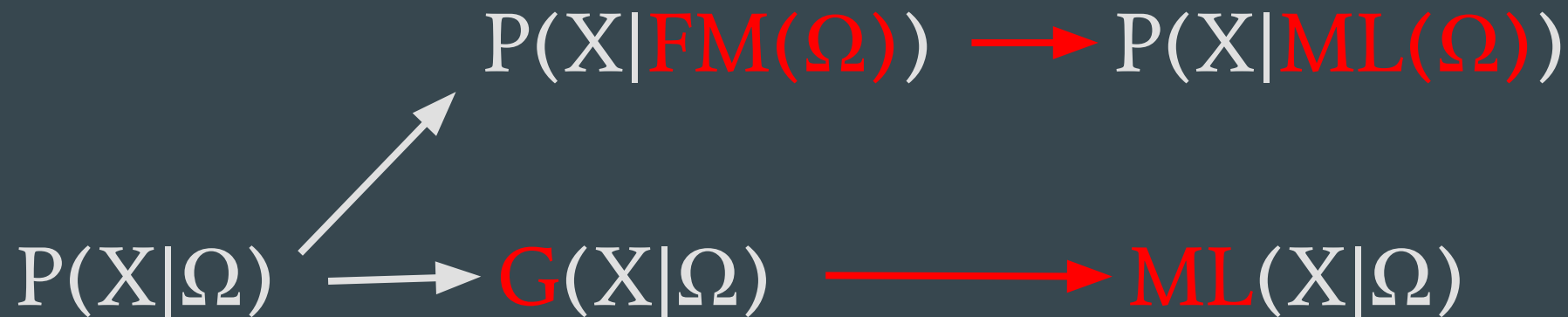
Evidence

Prior

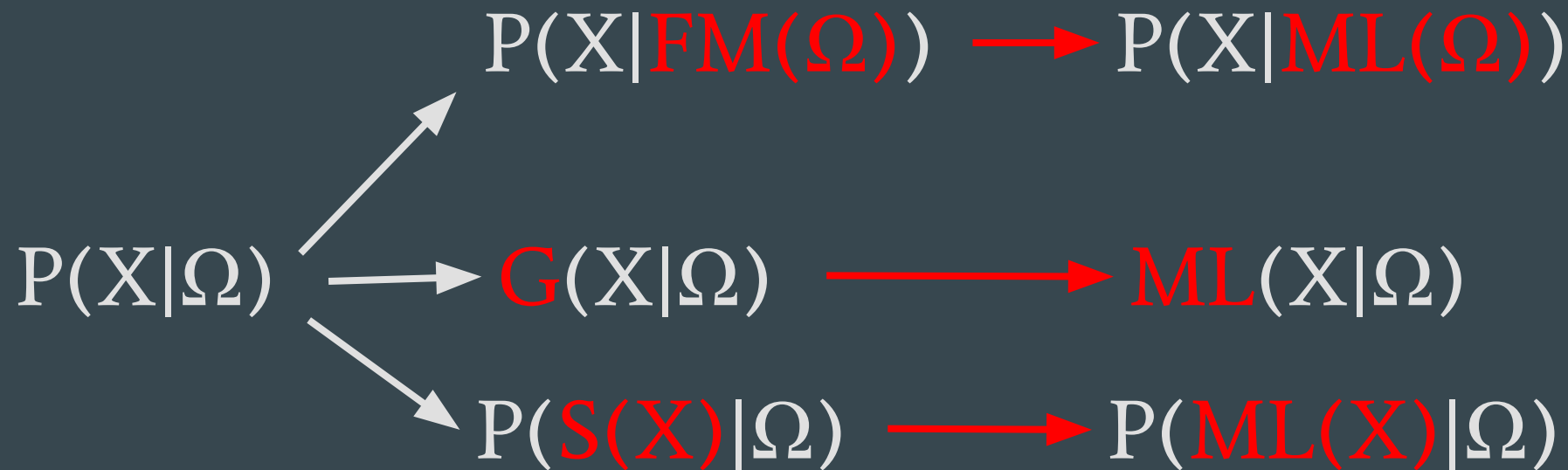
Future Work



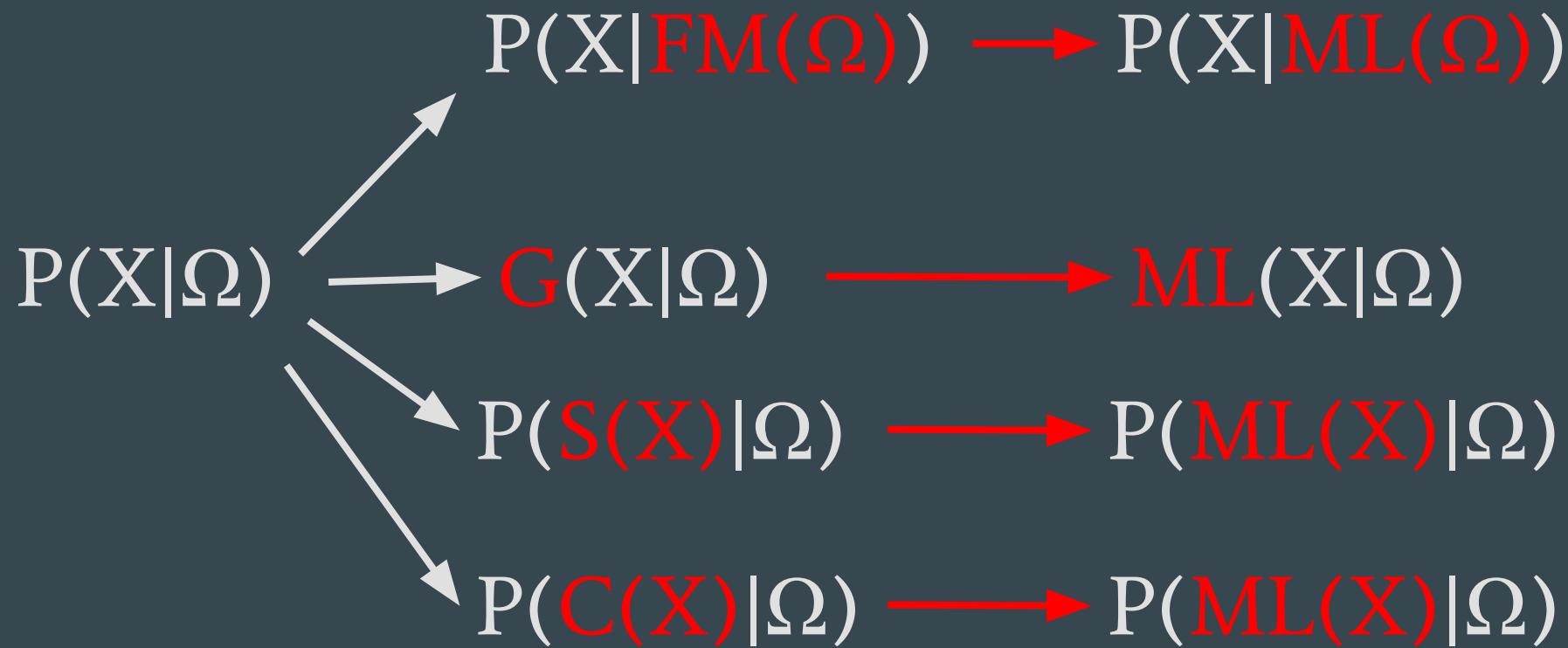
Future Work



Future Work



Future Work



Future Work

$$P(\Omega|X) = P(X|\Omega) P(\Omega)/P(X)$$

Diagram illustrating the relationship between the equation and its components:

- $P(\Omega|X)$ is labeled **Posterior** (indicated by a red arrow).
- $P(X|\Omega)$ is labeled **Likelihood** (indicated by a grey arrow).
- $P(\Omega)/P(X)$ is labeled **Evidence** (indicated by a grey arrow).
- $P(\Omega)$ is labeled **Prior** (indicated by a grey arrow).

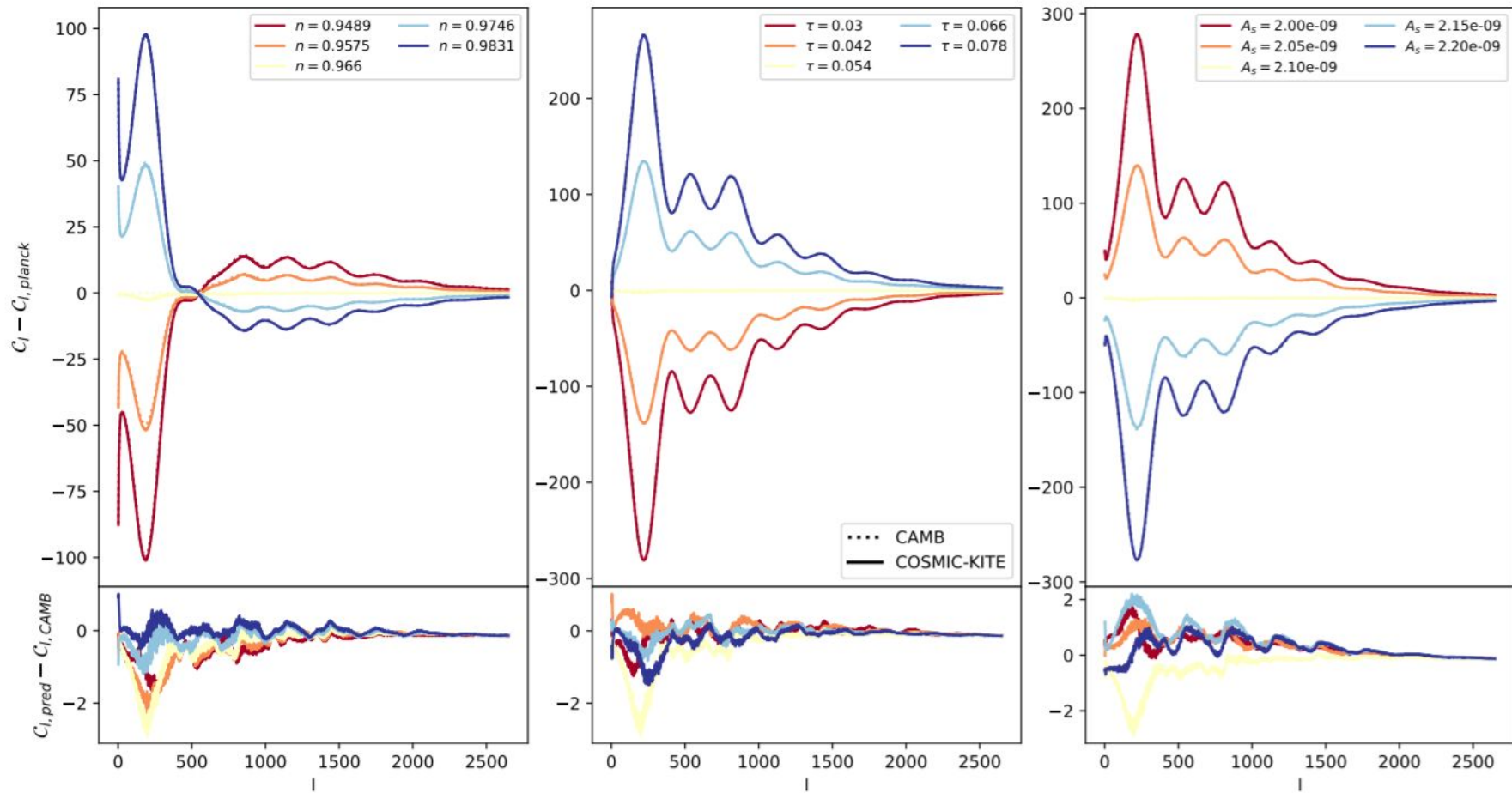
Future Work

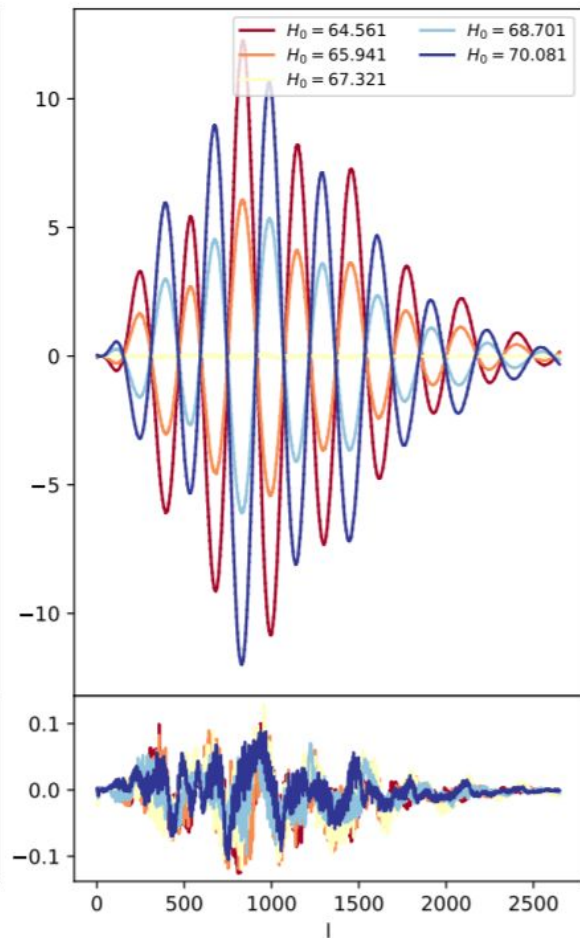
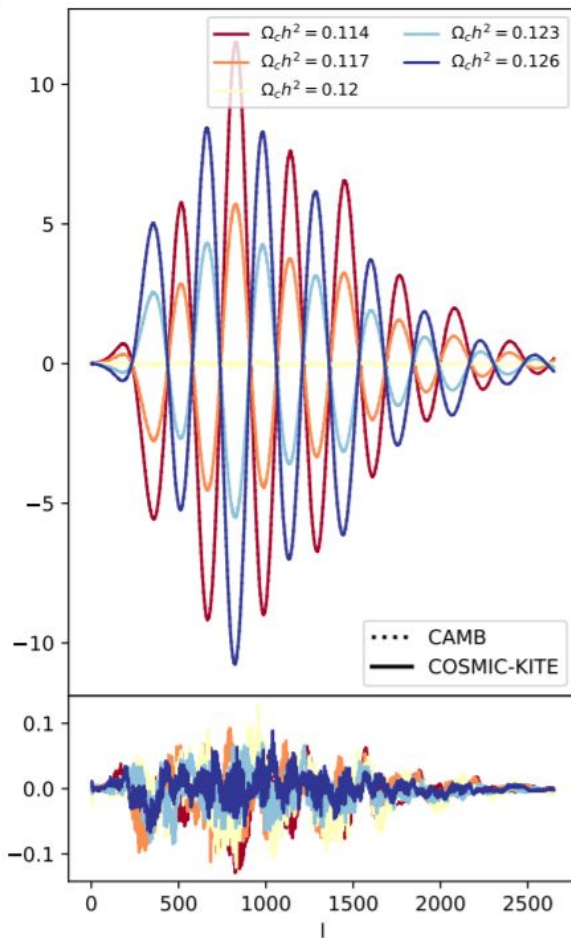
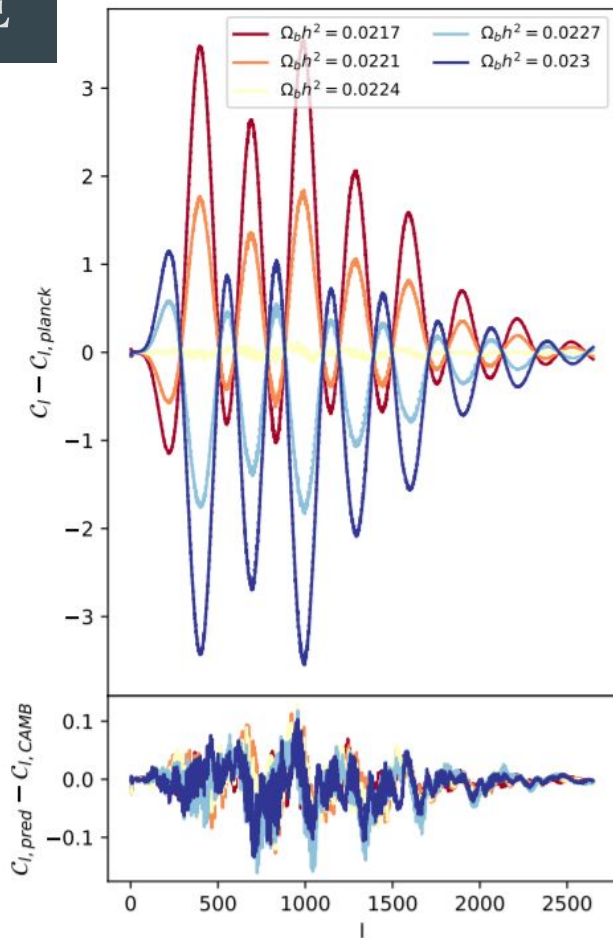
$$P(\Omega|X) \longrightarrow ML(X|\Omega)$$

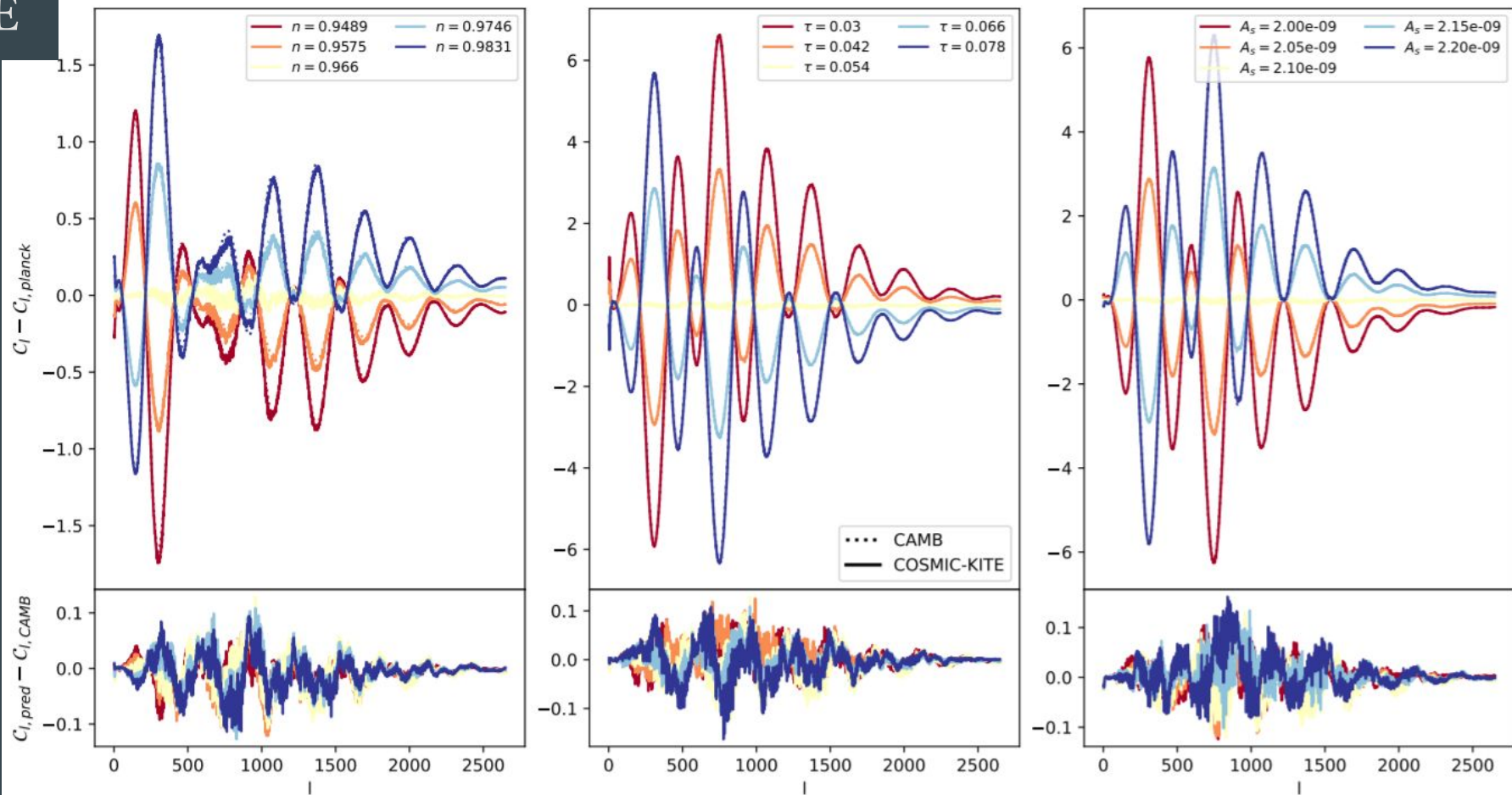


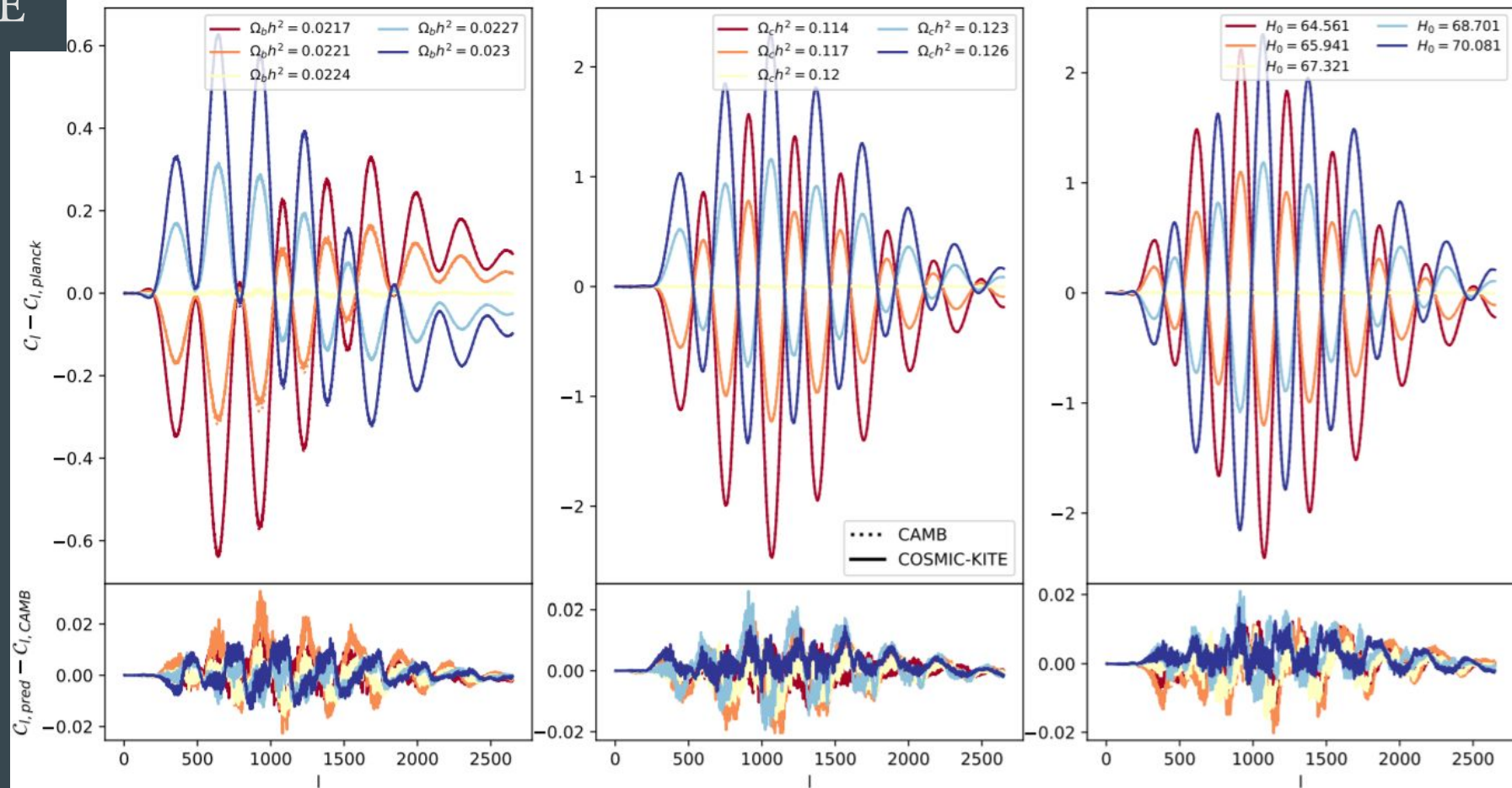
THANK YOU

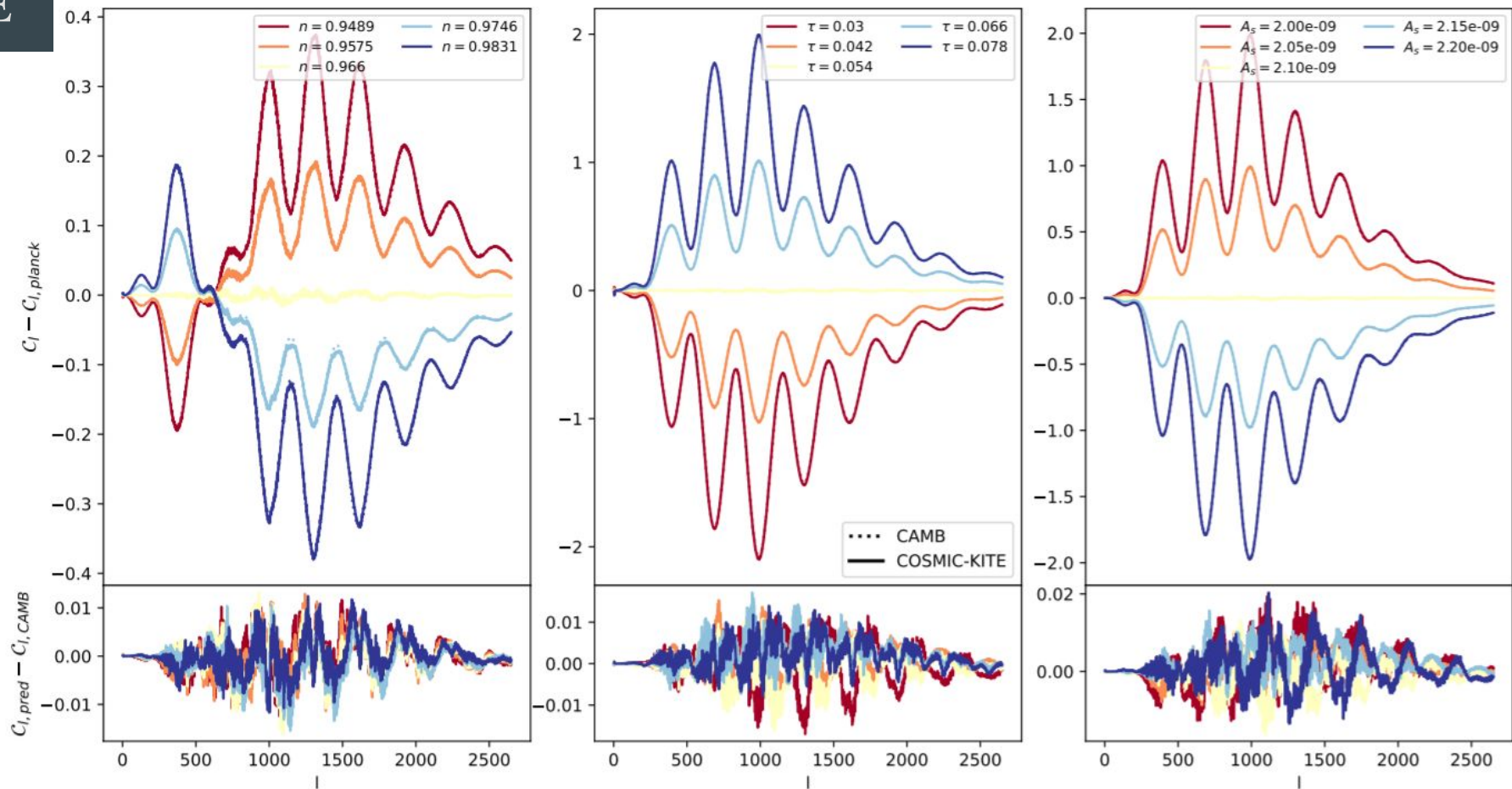
Back-up Slides







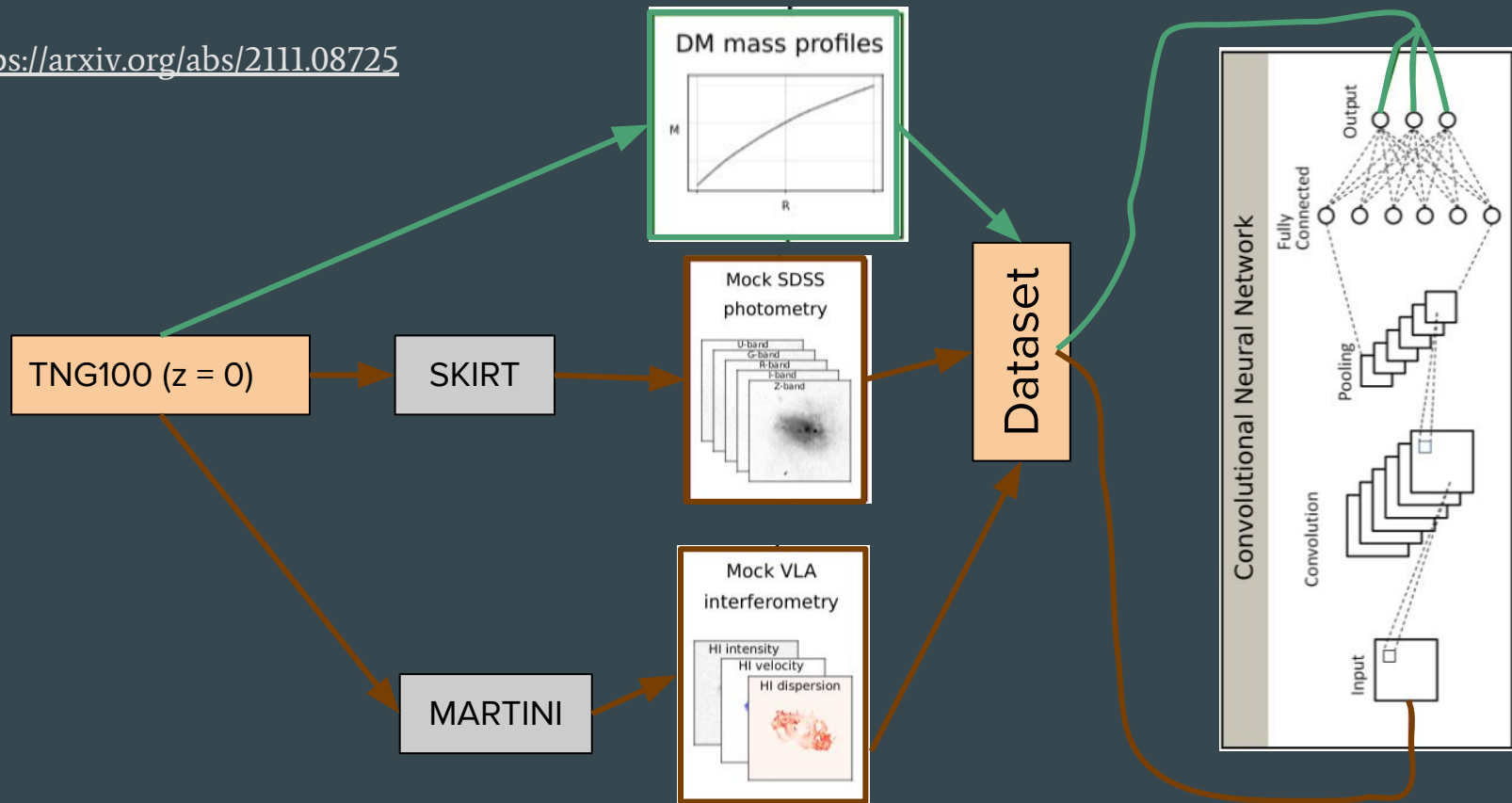




Determining the Dark Matter distribution in galaxies with Deep Learning

Martín de los Ríos, Mihael Petac, Bryan Zaldivar, Nina Bonaventura, Francesca Calore, Fabio Iocco

<https://arxiv.org/abs/2111.08725>

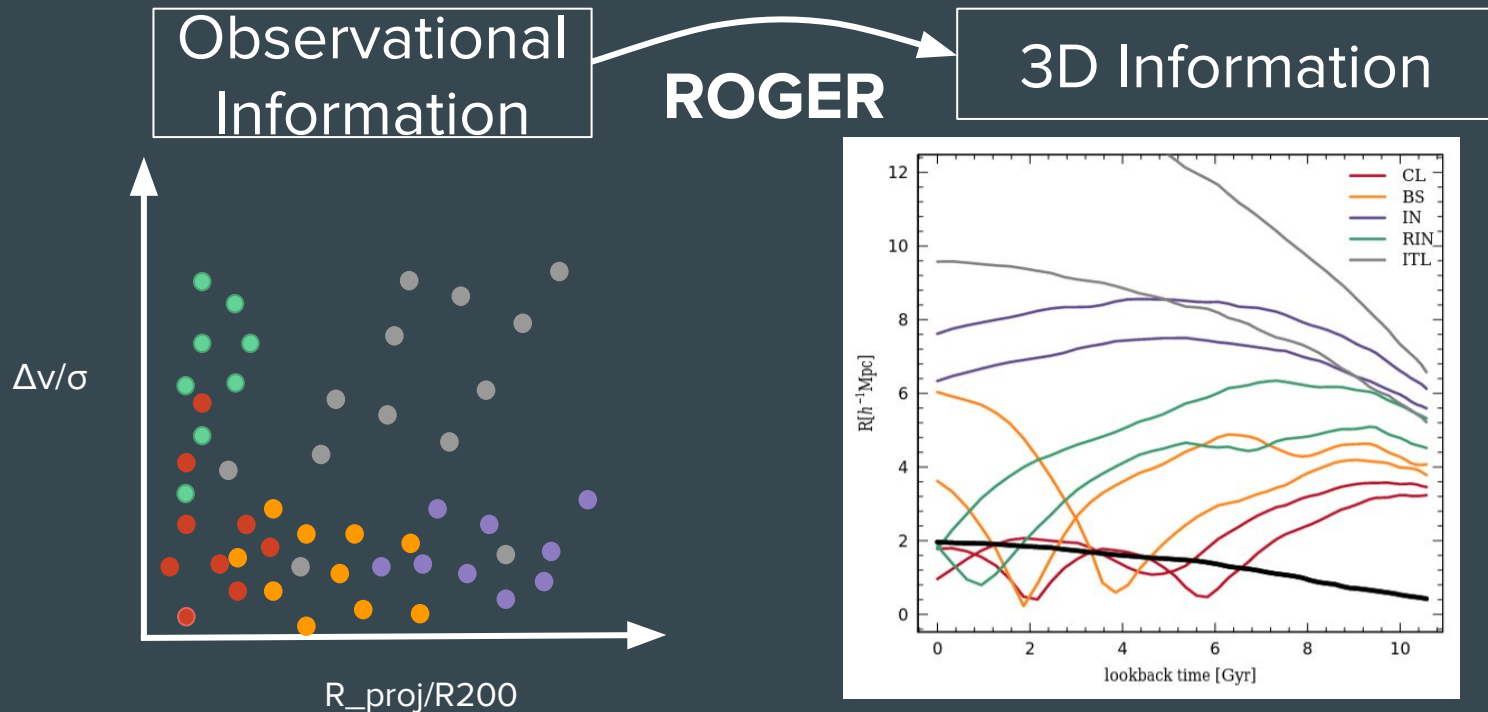


ROGER: Reconstructing Orbits of Galaxies in Extreme Regions using machine learning techniques

Martín de los Ríos, Héctor J. Martínez, Valeria Coenda, Hernán Muriel, Andrés N. Ruiz, Cristian A. Vega-Martínez, Sofía A. Cora

<https://arxiv.org/abs/2010.11959>

<https://arxiv.org/abs/2112.01552>



The MeSsl (Merging Systems Identification) Algorithm

Martín de los Ríos, Mariano J. Domínguez R., Dante Paz, Manuel Merchán

<https://arxiv.org/abs/1509.02524>

<https://arxiv.org/abs/1801.01498>

<https://arxiv.org/abs/1905.10303>

