Measuring the angular distribution of the cosmological parameters.

Martín de los Rios & Mariano Domínguez

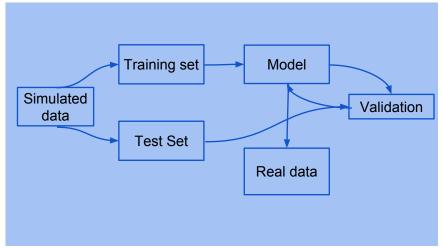
March 31, 2017

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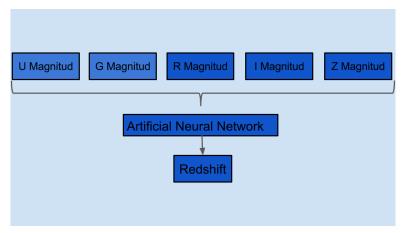
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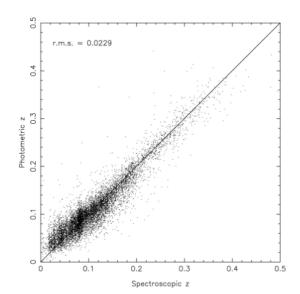
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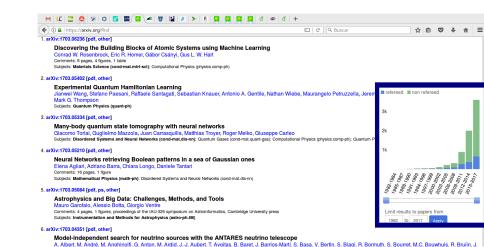
Simple Example: ANN-z

ANNz: Estimating photometric redshift using artificial neural network. Collister & Lahav 2003 (0311058)







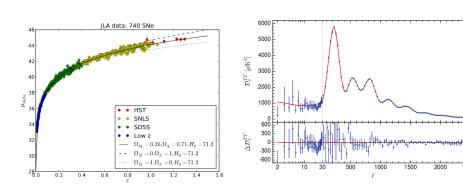


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What are the cosmological parameters?

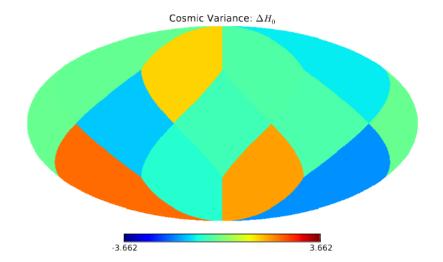
Homogeneous and isotropic Universe → FRW metric $ds^2 = dt^2 - a^2(t) \left[\frac{dr^2}{1 + r^2} + r^2 (d\theta^2 + \sin^2\theta d\phi^2) \right]$ $\left(\frac{H}{H_0}\right)^2 = \Omega_{rad}a^{-4} + \Omega_ma^{-3} + \Omega_{\Lambda} - Kc^2a^{-2}$

How can we measure the cosmological parameters?



Carvalho & Marques 2015 (1512.07869) Planck Collaboration 2015 (1502.01589)





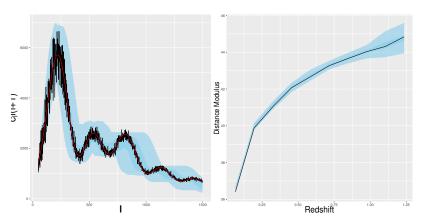
Carvalho & Marques (1512.07869)



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The training sample.

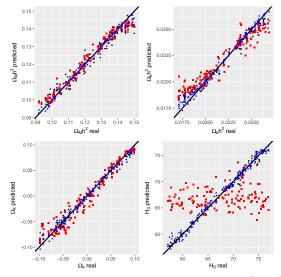
CAMB: Code for Anisotropies in the Cosmic Background (REFERENCIA)



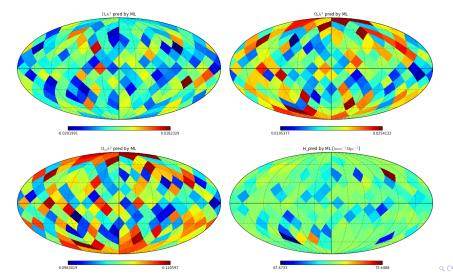


What is Machine Learning.

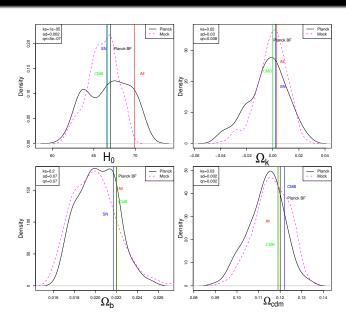
Studying different Machine Learning algorithms.



First Results.



What is Machine Learning.



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Final Remarks

- We developed a machine learning technique that estimate the cosmological parameters in a more efficient way, and allow us to measure the angular distribution of this parameters.
- This technique can be easily extended to use more cosmological information as features (BAO, correlation function, SZ emission, etc.).
- We do not found statistically significant departures from what is expected in an homogeneous and isotropic universe, with the possible exception of a bi-modal H_0 distribution.
- We will extend the parameters space and add polarization information in a forthcoming work.
- We will analyze the correlations between the angular distribution of the cosmological parameters and the large scale structure (voids, filaments, etc.)