

Introductory Computational Mathematics

Semester 2, 2022

Dr Elliot Carr

Tarang Janawalkar

This work is licensed under a Creative Commons
“Attribution-NonCommercial-ShareAlike 4.0 International” license.



Contents

Contents	1
1 Preliminaries	2
1.1 Errors	2
1.2 Floating Point Arithmetic	2
1.2.1 Representing Real Numbers as Floating Point Numbers	3
1.2.2 Converting between Floating Point Number Systems	3
1.2.3 IEEE Floating Point Standard	4
1.3 Catastrophic Cancellation	4
1.4 Taylor Polynomials	5
1.5 Taylor Series	6
2 Ordinary Differential Equations	6
2.1 Initial Value Problems	6
2.2 Time Discretisation	6
2.3 Euler's Method	6
2.4 Local and Global Error	7
2.5 Second Order Taylor Method	7
2.6 Modified Euler Method	8
2.7 Runge-Kutta Method	9
3 Interpolation	9
3.1 Lagrange Form	10
3.1.1 Two Data Points	10
3.1.2 Three Data Points	10
3.1.3 Arbitrary Number of Points	11
3.2 Error in Polynomial Approximation	11
3.3 Newton's Divided Difference Form	12
3.3.1 Derivation	12
3.3.2 Divided Differences	13
3.4 Newton's Forward Difference Form	14

1 Preliminaries

1.1 Errors

Errors in calculations are a common problem in numerical analysis. We can quantify the magnitude of such an error by two measures.

Definition 1.1 (Absolute and relative error). Let \tilde{x} be an approximation of x . Then the **absolute error** is given by

$$\text{absolute error} = |\tilde{x} - x|.$$

The **relative error** is given by

$$\text{relative error} = \frac{|\tilde{x} - x|}{|x|}.$$

It is important to realise that the absolute error can be misleading when comparing different sizes of errors, i.e., it is always small for small values of x and \tilde{x} .

1.2 Floating Point Arithmetic

The set of real numbers \mathbb{R} contains uncountably many elements. Computers have a limited number of bits, and can therefore only represent a small subset of these elements.

The most common approximation of real arithmetic used in computers is known as **floating point arithmetic**.

Definition 1.2 (Floating point number system). A floating point number system $\mathbb{F}(\beta, k, m, M)$ is a *finite subset* of the real number system characterised by the parameters:

- $\beta \in \mathbb{N}$: the base
- $k \in \mathbb{N}$: the number of digits in the significand
- $m \in \mathbb{Z}$: the minimum exponent
- $M \in \mathbb{Z}$: the maximum exponent

Definition 1.3 (Floating point numbers). The floating point numbers $f \in \mathbb{F}(\beta, k, m, M)$ are real numbers expressible in the form

$$f = \pm (d_1.d_2d_3 \dots d_k)_\beta \times \beta^e$$

where $e \in \mathbb{Z}$ is the **exponent** satisfying $m \leq e \leq M$. The quantity $d_1.d_2d_3 \dots d_k$ is known as the **significand**, where d_i are base- β digits, with $d_1 \neq 0$ unless $f = 0$, to ensure a unique representation of f .

Computers primarily use floating point number systems with base $\beta = 2$ (binary), other common bases include $\beta = 10$ (decimal¹) and $\beta = 16$ (hexadecimal).

¹Note that for base-10, we do not need to include the subscript in the significand.

To illustrate the finiteness of the floating point number system, consider the following example:

$$\begin{aligned}\mathbb{F}(10, 3, -1, 1) &= \{0, \\ &\quad \pm 1.00 \times 10^{-1}, \quad \pm 1.01 \times 10^{-1}, \quad \dots, \quad \pm 9.99 \times 10^{-1}, \\ &\quad \pm 1.00 \times 10^0, \quad \pm 1.01 \times 10^0, \quad \dots, \quad \pm 9.99 \times 10^0, \\ &\quad \pm 1.00 \times 10^1, \quad \pm 1.01 \times 10^1, \quad \dots, \quad \pm 9.99 \times 10^1 \} \\ &= \{0, \\ &\quad \pm 0.100, \quad \pm 0.101, \quad \dots, \quad \pm 0.999, \\ &\quad \pm 1.00, \quad \pm 1.01, \quad \dots, \quad \pm 9.99, \\ &\quad \pm 10.0, \quad \pm 10.1, \quad \dots, \quad \pm 99.9 \}\end{aligned}$$

Note that the numbers in this set are not equally spaced, (smaller spacing for smaller exponents).

Definition 1.4 (Overflow and underflow). Consider the value $x \in \mathbb{R}$, if x is too small in magnitude to be represented in \mathbb{F} , an **underflow** occurs which typically causes the number to be replaced by zero.

Similarly, if x is too large in magnitude to be represented in \mathbb{F} , an **overflow** occurs which typically causes the number to be replaced by infinity.

Corollary 1.2.0.1. *The smallest and largest values (in magnitude) of \mathbb{F} are given by*

$$\begin{aligned}\min_{f \in \mathbb{F}} |f| &= \beta^m \\ \max_{f \in \mathbb{F}} |f| &= (1 - \beta^{-k}) \beta^{M+1}.\end{aligned}$$

The cardinality of the positive elements in \mathbb{F} , is given by

$$|\{f \in \mathbb{F} : f > 0\}| = (M - m + 1) (\beta - 1) \beta^{k-1}$$

so that by including negative numbers and zero, the cardinality of \mathbb{F} is given by

$$|\mathbb{F}| = 2|\{f \in \mathbb{F} : f > 0\}| + 1.$$

1.2.1 Representing Real Numbers as Floating Point Numbers

If we wish to represent a real number² x that is not exactly representable in \mathbb{F} , we can **round** the number to the nearest *representable* number.

The error committed by this process is known as the **roundoff error**.

1.2.2 Converting between Floating Point Number Systems

Consider $fl : \mathbb{R} \rightarrow \mathbb{F}(\beta, k, m, M)$, defined as function which maps real numbers x to the nearest element in \mathbb{F} . To determine $fl(x)$:

1. Express x in base β .
2. Express x in scientific form.

² x must satisfy $\min(\mathbb{F}) \leq x \leq \max(\mathbb{F})$.

3. Verify that $m \leq e \leq M$:

- If $e > M$, then $x = \infty$.
- If $e < m$, then $x = 0$.
- Otherwise, round the significand to k digits.

The relative error produced by rounding x to $fl(x)$ is bounded according to

$$\frac{|x - fl(x)|}{|x|} \leq \frac{1}{2}\beta^{1-k}.$$

Definition 1.5 (Unit roundoff). The **unit roundoff** or **machine precision** u of a floating point number system $\mathbb{F}(\beta, k, m, M)$ is given by

$$u = \frac{1}{2}\beta^{1-k}.$$

1.2.3 IEEE Floating Point Standard

IEEE 754 is the standard for floating point arithmetic used by most modern computers.

It is a binary format, with several variants. The most common variant is **IEEE double precision**, which is based on $\mathbb{F}(2, 53, -1022, 1023)$.

The basic properties of this format are summarised in the following table.

Unit roundoff	$u = 1.11 \times 10^{-16}$
Largest representable positive number	1.80×10^{308}
Smallest representable positive number	2.23×10^{-308}
Special values	$\pm 0, \pm \infty, \text{NaN}$

1.3 Catastrophic Cancellation

When working with floating point arithmetic, roundoff is a common source of error. Certain operations may bring roundoff errors that are too large to be easily corrected.

Catastrophic cancellation or **cancellation error** is the error that occurs in the floating point subtraction of two numbers that are very close to each other, where at least one of them is not exactly representable.

As an example, the quadratic formula

$$x_1 = \frac{-b + \sqrt{b^2 - 4ac}}{2a} \qquad x_2 = \frac{-b - \sqrt{b^2 - 4ac}}{2a}$$

experiences catastrophic cancellation for $b^2 \gg 4ac$, as $b^2 - 4ac \approx b^2$ so that $\sqrt{b^2 - 4ac} = \sqrt{b^2} = |b|$:

$$x_1 = \frac{-b + |b|}{2a} \qquad x_2 = \frac{-b - |b|}{2a}$$

When $b > 0$, $|b| = b$, so that

$$x_1 = \frac{-b + b}{2a} = \frac{b - b}{2a}.$$

And when $b < 0$, $|b| = -b$, so that

$$x_2 = \frac{-b - (-b)}{2a} = \frac{b - b}{2a}.$$

This cancellation can be avoided by taking the product of the two roots to determine the exact result of the root that suffers from catastrophic cancellation.

$$x_1 x_2 = \frac{c}{a}.$$

1.4 Taylor Polynomials

$$\cos(x) = \sum_{k=0}^n \frac{(-1)^k}{(2k)!} x^{2k}.$$

Suppose we have a function $f(x)$ that is n differentiable at the point $x = x_0$. This function can be approximated by a sum of polynomials that agrees with its first n derivatives at that point.

Definition 1.6 (Taylor polynomial). The **Taylor polynomial** of degree n of f , centred at x_0 is defined by

$$P_n(x) = \sum_{k=0}^n \frac{f^{(k)}(x_0)}{k!} x^k.$$

The Taylor polynomial can be used to approximate a function f for values of x near x_0 , the following theorem addresses how accurate the approximation is.

Definition 1.7 (Taylor's theorem). Suppose that f is $n+1$ times differentiable on an interval $[a, b]$ containing x_0 , and let P_n be the degree n Taylor polynomial for f , centred on x_0 . Then for all $x \in [a, b]$, there exists a value $x_0 < c < x$ such that

$$f(x) = P_n(x) + \frac{f^{(n+1)}(c)}{(n+1)!} (x - x_0)^{n+1}.$$

The term

$$R_n(x) = \frac{f^{(n+1)}(c)}{(n+1)!} (x - x_0)^{n+1}$$

is called the **error term** or **remainder term** for P_n .

To determine the absolute error from the Taylor series polynomial, consider the error term:

$$|f(x) - P_n(x)| = |R_n(x)|.$$

The maximum value of $|R_n(x)|$ on the interval $[a, b]$ gives the bound on the maximum error incurred when approximating f by P_n on that interval.

1.5 Taylor Series

Given an infinitely differentiable function f , we can take the limit $n \rightarrow \infty$ to find Taylor series representation of f , given by:

$$f(x) = \sum_{k=0}^{\infty} \frac{f^{(k)}(x_0)}{k!} (x - x_0)^k.$$

When we truncate this series at a finite n , the error from the Taylor series is known as **truncation error**. In this case, the remainder term gives us a *bound* on the size of this truncation error.

2 Ordinary Differential Equations

For certain classes of ordinary differential equations (ODEs), we can obtain analytical closed-form solutions using known techniques. However for most cases we will need to use numerical techniques to obtain approximate solutions.

2.1 Initial Value Problems

While solutions with arbitrary constants, such as $y = Ce^t$ are acceptable for theoretical analysis, we cannot have such variables when determining an approximate solution in the real world. Hence we require **initial conditions** to obtain a definitive solution. An ODE combined with an initial condition is called an **initial value problem** (IVP).

2.2 Time Discretisation

Consider the initial value problem,

$$\begin{aligned} y(a) &= \alpha \\ \frac{dy(t)}{dt} &= f(t, y(t)), \quad a \leq t \leq b \end{aligned}$$

Divide the interval $[a, b]$ into n subintervals, each with width $h = (b - a) / n$. Then define $t_i = a + ih$, for $i = 0, 1, \dots, n$, so that $t_0 = a$ and $t_n = b$. If we then compute $y_i = y(t_i)$ denoted as w_i , then $w_i \approx y_i$ for all $i = 0, 1, \dots, n$.

2.3 Euler's Method

Euler's method, or the first order Taylor method, uses a Taylor polynomial approximation of y over each subinterval. Assuming y is twice differentiable,

$$y(t_i + h) = y(t_i) + hy'(t_i) + \mathcal{O}(h^2).$$

The remainder term is not shown in its exact form but rather as $\mathcal{O}(h^2)$, “Big-O of h^2 ”, or “order h^2 ”, meaning that the error is proportional to h^2 . Using the ODE and substituting $t_{i+1} = t_i + h$ gives

$$y_{i+1} = y_i + hf(t_i, y_i) + \mathcal{O}(h^2).$$

By removing the remainder term, we obtain the approximation method known as **Euler's method**:

$$\begin{aligned} w_0 &= \alpha \\ w_{i+1} &= w_i + hf(t_i, w_i) \end{aligned}$$

for $i = 0, 1, \dots, n-1$.

2.4 Local and Global Error

Local error is defined as the error that the method would incur in **one step**, assuming the solution was correct at the previous step. The **global error** is defined as the error in the solution after i steps (at $t = t_i$), and is given by:

$$|y_i - w_i|.$$

It is the accumulation of the local errors from steps 1, 2, ..., i .

The **order** of a method refers to the global error of that method. For Euler's method, the local error is $\mathcal{O}(h^2)$ while the global error is $\mathcal{O}(h)$, hence "*first order* Taylor method".

In general, if the local error is $\mathcal{O}(h^{p+1})$, then the global error is $\mathcal{O}(h^p)$ and the method is said to be a p th order method.

$$\text{Global error} \approx n\mathcal{O}(h^{p+1}) = \frac{b-a}{h}\mathcal{O}(h^{p+1}) = \mathcal{O}(h^p).$$

2.5 Second Order Taylor Method

To improve upon the accuracy of Euler's method, we can use additional Taylor polynomial terms by truncating at a higher order. Assuming y is three times differentiable, we have

$$y(t_i + h) = y(t_i) + hy'(t_i) + \frac{h^2}{2}y''(t_i) + \mathcal{O}(h^3)$$

which can be rewritten as

$$y_{i+1} = y_i + hf(t_i, y_i) + \frac{h^2}{2}f'(t_i, y_i) + \mathcal{O}(h^3).$$

Again by removing the remainder term, we obtain the approximation w_i of y_i , known as the **second order Taylor method**:

$$\begin{aligned} w_0 &= \alpha \\ w_{i+1} &= w_i + hf(t_i, w_i) + \frac{h^2}{2}f'(t_i, w_i) \end{aligned}$$

for $i = 0, 1, \dots, n-1$. This method has a local error of $\mathcal{O}(h^3)$, and therefore a global error of $\mathcal{O}(h^2)$.

2.6 Modified Euler Method

Although the second order Taylor method is the more accurate than Euler's method, we require computing $f'(t, y)$.

Suppose we use a numerical approximation of the derivative of f .

By definition, the derivative of a function is the limiting value of the slope of the line connecting two nearby points on a curve,

$$f'(t_i, y_i) = \lim_{h \rightarrow 0} \frac{f(t_{i+1}, y_{i+1}) - f(t_i, y_i)}{h}.$$

For small values of h , we can approximate the derivative with

$$f'(t_i, y_i) \approx \frac{f(t_{i+1}, y_{i+1}) - f(t_i, y_i)}{h}.$$

By deriving the error term we find that

$$f'(t_i, y_i) = \frac{f(t_{i+1}, y_{i+1}) - f(t_i, y_i)}{h} + \mathcal{O}(h).$$

Hence the second order Taylor polynomial becomes

$$\begin{aligned} y_{i+1} &= y_i + hf(t_i, y_i) + \frac{h^2}{2} f'(t_i, y_i) + \mathcal{O}(h^3) \\ &= y_i + hf(t_i, y_i) + \frac{h^2}{2} \left[\frac{f(t_{i+1}, y_{i+1}) - f(t_i, y_i)}{h} + \mathcal{O}(h) \right] + \mathcal{O}(h^3) \\ &= y_i + hf(t_i, y_i) + \frac{h}{2} f(t_{i+1}, y_{i+1}) - \frac{h}{2} f(t_i, y_i) + \mathcal{O}(h^3) + \mathcal{O}(h^3) \\ &= y_i + \frac{h}{2} f(t_i, y_i) + \frac{h}{2} f(t_{i+1}, y_{i+1}) + \mathcal{O}(h^3) \\ &= y_i + \frac{h}{2} [f(t_i, y_i) + f(t_{i+1}, y_{i+1})] + \mathcal{O}(h^3) \\ &\approx y_i + \frac{h}{2} [f(t_i, y_i) + f(t_{i+1}, y_{i+1})] \\ w_{i+1} &= w_i + \frac{h}{2} [f(t_i, w_i) + f(t_{i+1}, w_{i+1})] \end{aligned}$$

As this formula involves itself, we will use Euler's method on $w_{i+1} = w_i + hf(t_i, w_i)$ on the RHS.

$$\begin{aligned} w_{i+1} &= w_i + \frac{1}{2} [hf(t_i, w_i) + f(t_{i+1}, w_{i+1})] \\ &= w_i + \frac{1}{2} \left[\underbrace{hf(t_i, w_i)}_{k_1} + \underbrace{hf\left(t_{i+1}, w_i + \underbrace{hf(t_i, w_i)}_{k_1}\right)}_{k_2} \right] \\ &= w_i + \frac{1}{2} (k_1 + k_2) \end{aligned}$$

This result leads to what is known as the **modified Euler method**.

$$\begin{aligned}w_0 &= \alpha \\w_{i+1} &= w_i + \frac{1}{2} (k_1 + k_2) \\k_1 &= hf(t_i, w_i) \\k_2 &= hf(t_i + h, w_i + k_1)\end{aligned}$$

This method has a local error of $\mathcal{O}(h^3)$, and therefore a global error of $\mathcal{O}(h^2)$.

2.7 Runge-Kutta Method

Another popular method that agrees with the Taylor method for high orders is the Runge-Kutta method. It is a fourth order method, referred to as RK4. It has the following form:

$$\begin{aligned}w_0 &= \alpha \\w_{i+1} &= w_i + \frac{1}{6} (k_1 + 2k_2 + 2k_3 + k_4) \\k_1 &= hf(t_i, w_i) \\k_2 &= hf\left(t_i + \frac{h}{2}, w_i + \frac{k_1}{2}\right) \\k_3 &= hf\left(t_i + \frac{h}{2}, w_i + \frac{k_2}{2}\right) \\k_4 &= hf(t_i + h, w_i + k_3)\end{aligned}$$

for $i = 0, 1, \dots, n-1$. The local error is $\mathcal{O}(h^4)$, and the global error is $\mathcal{O}(h^3)$.

3 Interpolation

Definition 3.1 (Interpolating function). A function f is said to interpolate the data points

$$\{(x_0, y_0), (x_1, y_1), \dots, (x_n, y_n)\}$$

if it satisfies $f(x_0) = y_0, f(x_1) = y_1, \dots, f(x_n) = y_n$.

There are many kinds of functions that may satisfy an interpolating function, however, the most common choice is a **polynomial**.

The following theorem addresses the existence and uniqueness of interpolating polynomials, given distinct x -values, or **abscissas**.

Theorem 3.0.1. *Let the abscissas x_0, x_1, \dots, x_n be distinct and the function values y_0, y_1, \dots, y_n be given. Then there exists a **unique** interpolating polynomial P_n of degree (at most) n , such that $P_n(x_i) = y_i$ for $i \in [0, n]$.*

The following methods are used to find the interpolating polynomial.

3.1 Lagrange Form

3.1.1 Two Data Points

Consider constructing the interpolating polynomial through the points $\{(x_0, y_0), (x_1, y_1)\}$. We require a polynomial of degree 1 to satisfy the interpolating function. Let $P_1(x) = a_0 + a_1x$ satisfy the following conditions:

$$\begin{aligned} P_1(x_0) &= y_0 && \Leftrightarrow a_0 + a_1x_0 = y_0 \\ P_1(x_1) &= y_1 && \Leftrightarrow a_0 + a_1x_1 = y_1 \end{aligned}$$

solving this linear system of equations, we obtain the following:

$$\begin{aligned} a_0 &= \frac{x_1y_0 - x_0y_1}{x_1 - x_0} \\ a_1 &= \frac{y_1 - y_0}{x_1 - x_0} \end{aligned}$$

so that

$$P_1(x) = \frac{x_1y_0 - x_0y_1}{x_1 - x_0} + \frac{y_1 - y_0}{x_1 - x_0}x.$$

To generalise this process, consider the following rearrangement of the polynomial:

$$\begin{aligned} P_1(x) &= \frac{x_1y_0 - x_0y_1}{x_1 - x_0} + \frac{y_1 - y_0}{x_1 - x_0}x \\ &= \frac{x_1y_0}{x_1 - x_0} - \frac{x_0y_1}{x_1 - x_0} + \frac{y_1x}{x_1 - x_0} - \frac{y_0x}{x_1 - x_0} \\ &= \left(\frac{-x}{x_1 - x_0} + \frac{x_1}{x_1 - x_0} \right) y_0 + \left(\frac{x}{x_1 - x_0} - \frac{x_0}{x_1 - x_0} \right) y_1 \\ &= \frac{x - x_1}{x_0 - x_1} y_0 + \frac{x - x_0}{x_1 - x_0} y_1. \end{aligned}$$

This form clearly shows that each point satisfies the interpolating function.

- The coefficient of y_0 is **1** when $x = x_0$, and **0** when $x = x_1$.
- The coefficient of y_1 is **0** when $x = x_0$, and **1** when $x = x_1$.

we can then write $P_1(x)$ as

$$P_1(x) = L_{1,0}(x)y_0 + L_{1,1}(x)y_1$$

where $L_{1,0}(x) = \frac{x-x_1}{x_0-x_1}$ and $L_{1,1}(x) = \frac{x-x_0}{x_1-x_0}$ are the Lagrange basis functions. The first index corresponds to the degree of the polynomial, and the second index corresponds to the index of the point for which it is the coefficient of.

3.1.2 Three Data Points

For three data points, consider the reverse problem:

- The coefficient of y_0 is **1** when $x = x_0$, **0** when $x = x_1$, and **0** when $x = x_2$.

- The coefficient of y_1 is **0** when $x = x_0$, **1** when $x = x_1$, and **0** when $x = x_2$.
- The coefficient of y_2 is **0** when $x = x_0$, **0** when $x = x_1$, and **1** when $x = x_2$.

then solve for the Lagrange basis functions:

$$\begin{aligned} P_2(x) &= L_{2,0}(x)y_0 + L_{2,1}(x)y_1 + L_{2,2}(x)y_2 \\ &= \frac{(x-x_1)(x-x_2)}{(x_0-x_1)(x_0-x_2)}y_0 + \frac{(x-x_0)(x-x_2)}{(x_1-x_0)(x_1-x_2)}y_1 + \frac{(x-x_0)(x-x_1)}{(x_2-x_0)(x_2-x_1)}y_2. \end{aligned}$$

3.1.3 Arbitrary Number of Points

Given the abscissas $\{x_0, x_1, \dots, x_n\}$ and function values $\{y_0, y_1, \dots, y_n\}$ the polynomial $P_n(x)$ defined by:

$$P_n(x) = \sum_{i=0}^n L_{n,i}(x)y_i$$

where

$$L_{n,i}(x) = \prod_{j=0, j \neq i}^n \frac{x-x_j}{x_i-x_j}$$

is called the **Lagrange form** of the interpolating polynomial.

This arises from the fact that we can construct the Lagrange form using

$$\begin{aligned} P_n(x) &= L_{n,0}(x)y_0 + L_{n,1}(x)y_1 + \dots + L_{n,n}(x)y_n \\ &= \sum_{i=0}^n L_{n,i}(x)y_i \end{aligned}$$

where the Lagrange basis function is as follows:

$$\begin{aligned} L_{n,i}(x) &= \frac{(x-x_0)(x-x_1)\dots(x-x_{i-1})(x-x_{i+1})\dots(x-x_n)}{(x_i-x_0)(x_i-x_1)\dots(x_i-x_{i-1})(x_i-x_{i+1})\dots(x_i-x_n)} \\ &= \prod_{j=0, j \neq i}^n \frac{x-x_j}{x_i-x_j} \end{aligned}$$

Note that $L_{n,i}(x)$ is equal to **1** when $x = x_i$, and **0** for all other abscissa:

$$L_{n,i}(x_j) = \delta_{ij}$$

3.2 Error in Polynomial Approximation

Theorem 3.2.1 (Lagrange form of remainder). *Suppose that f is $n+1$ times continuously differentiable on $[a, b]$ ³ with distinct abscissas x_0, x_1, \dots, x_n and interpolating polynomial $P_n(x)$. Then for all $x \in [a, b]$ there exists a $c \in [a, b]$ such that*

$$f(x) = P_n(x) + \frac{f^{(n+1)}(c)}{(n+1)!} (x-x_0)(x-x_1)\dots(x-x_n).$$

³ $a = \min(x_i)$ and $b = \max(x_i)$.

where the term

$$R_n(x) = \frac{f^{(n+1)}(c)}{(n+1)!} (x-x_0)(x-x_1)\cdots(x-x_n)$$

is called the Lagrange form of the remainder term for $P_n(x)$.

3.3 Newton's Divided Difference Form

While the Lagrange form is straightforward to analyse, it is not ideal for numerical computation. Suppose that we have found the interpolating polynomial through $n+1$ points, and are now interested in including a new point in the interpolation. This requires us to recompute the Lagrange basis functions for all $n+2$ points.

3.3.1 Derivation

The Newton form of the interpolating polynomial is given by:

$$P_n(x) = a_0 + a_1(x-x_0) + a_2(x-x_0)(x-x_1) + \cdots + a_n(x-x_0)(x-x_1)\cdots(x-x_{n-1})$$

The interpolation conditions:

$$\begin{aligned} P_n(x_0) &= y_0 \\ P_n(x_1) &= y_1 \\ &\vdots \\ P_n(x_n) &= y_n \end{aligned}$$

give the following system of equations for coefficients a_0, a_1, \dots, a_n :

$$\begin{aligned} a_0 &= y_0 \\ a_0 + a_1(x_1 - x_0) &= y_1 \\ a_0 + a_1(x_2 - x_0) + a_2(x_2 - x_0)(x_2 - x_1) &= y_2 \\ &\vdots \\ a_0 + a_1(x_n - x_0) + a_2(x_n - x_0)(x_n - x_1) + \cdots + a_n(x_n - x_0)(x_n - x_1)\cdots(x_n - x_{n-1}) &= y_n \end{aligned}$$

Solving for the first three of these coefficients gives:

$$\begin{aligned} a_0 &= y_0 \\ a_1 &= \frac{y_1 - y_0}{x_1 - x_0} \\ a_2 &= \frac{\frac{y_2 - y_1}{x_2 - x_1} - \frac{y_1 - y_0}{x_1 - x_0}}{x_2 - x_0} \end{aligned}$$

To consisely represent these coefficients, we will need to define divided differences.

3.3.2 Divided Differences

Definition 3.2 (Zeroth divided difference). The zeroth divided difference of f with respect to x_i is denoted $f[x_i]$ and defined by:

$$f[x_i] = y_i$$

Definition 3.3 (k th divided difference). The k th divided difference of f with respect to $x_i, x_{i+1}, \dots, x_{i+k}$ is denoted $f[x_i, x_{i+1}, \dots, x_{i+k}]$ and defined by:

$$f[x_i, x_{i+1}, \dots, x_{i+k}] = \frac{f[x_{i+1}, \dots, x_{i+k}] - f[x_i, \dots, x_{i+k-1}]}{x_{i+k} - x_i}$$

For example,

$$\begin{aligned} f[x_0] &= y_0 \\ f[x_1] &= y_1 \\ f[x_0, x_1] &= \frac{f[x_1] - f[x_0]}{x_1 - x_0} = \frac{y_1 - y_0}{x_1 - x_0} \\ f[x_1, x_2] &= \frac{f[x_2] - f[x_1]}{x_2 - x_1} = \frac{y_2 - y_1}{x_2 - x_1} \\ f[x_0, x_1, x_2] &= \frac{f[x_1, x_2] - f[x_0, x_1]}{x_2 - x_0} = \frac{\frac{y_2 - y_1}{x_2 - x_1} - \frac{y_1 - y_0}{x_1 - x_0}}{x_2 - x_0} \end{aligned}$$

Using this notation we can rewrite the coefficients of P_n as

$$\begin{aligned} a_0 &= f[x_0] \\ a_1 &= f[x_0, x_1] \\ a_2 &= f[x_0, x_1, x_2] \\ &\vdots \\ a_n &= f[x_0, x_1, x_2, \dots, x_n] \end{aligned}$$

so that

$$\begin{aligned} P_n(x) &= f[x_0] + f[x_0, x_1](x - x_0) + f[x_0, x_1, x_2](x - x_0)(x - x_1) + \dots \\ &\quad + f[x_0, x_1, x_2, \dots, x_n](x - x_0)(x - x_1) \cdots (x - x_{n-1}). \end{aligned}$$

Definition 3.4 (Newton's divided difference form). Given the abscissas x_0, x_1, \dots, x_n and the function values y_0, y_1, \dots, y_n the polynomial P_n defined by

$$P_n(x) = \sum_{k=0}^n f[x_0, x_1, \dots, x_k] \left(\prod_{i=0}^{k-1} (x - x_i) \right)$$

is called the **Newton divided difference form of the interpolating polynomial**.

3.4 Newton's Forward Difference Form

In the case where abscissas are equally spaced, the Newton divided difference form can be simplified.

Definition 3.5 (Forward difference operator). The **forward difference operator** Δ is defined by:

$$\Delta y_i = y_{i+1} - y_i$$

Higher order forward differences are defined by repeated application:

$$\begin{aligned}\Delta^2 y_i &= \Delta(\Delta y_i) = \Delta(y_{i+1} - y_i) = \Delta y_{i+1} - \Delta y_i = y_{i+2} - y_{i+1} - (y_{i+1} - y_i) \\ &= y_{i+2} - 2y_{i+1} + y_i \\ \Delta^3 y_i &= \Delta(\Delta^2 y_i) = \Delta(y_{i+2} - 2y_{i+1} + y_i) = \Delta y_{i+2} - \Delta y_{i+1} - \Delta y_i \\ &= y_{i+3} - 3y_{i+2} + 3y_{i+1} - y_i\end{aligned}$$

Theorem 3.4.1 (Simplified divided differences). *If x_0, x_1, \dots, x_n are equally spaced, then*

$$f[x_0, x_1, \dots, x_k] = \frac{\Delta^k y_0}{k! h^k}$$

where $h = x_{i+1} - x_i$ is the spacing between the abscissas.

Using the above theorem and making the substitution $x = x_0 + sh$ (so that $x_i = x_0 + ih$) where $s = \frac{x - x_0}{h}$, we can rewrite the Newton divided difference form as

$$\begin{aligned}P_n(x) &= \sum_{k=0}^n f[x_0, x_1, \dots, x_k] \left(\prod_{i=0}^{k-1} (x - x_i) \right) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k! h^k} \left(\prod_{i=0}^{k-1} ((x_0 + sh) - (x_0 + ih)) \right) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k! h^k} \left(\prod_{i=0}^{k-1} (s - i) h \right) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k! h^k} \left(\prod_{i=0}^{k-1} (s - i) \prod_{i=0}^{k-1} h \right) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k! h^k} \left(\prod_{i=0}^{k-1} (s - i) h^k \right) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k!} s(s-1)(s-2) \cdots (s-(k-1)) \\ &= \sum_{k=0}^n \frac{\Delta^k y_0}{k!} s(s-1)(s-2) \cdots (s-k+1) \frac{(s-k)!}{(s-k)!} \\ &= \sum_{k=0}^n \Delta^k y_0 \frac{s!}{k! (s-k)!} \\ &= \sum_{k=0}^n \binom{s}{k} \Delta^k y_0\end{aligned}$$

Definition 3.6 (Newton's forward difference form). Given equally spaced abscissas x_0, x_1, \dots, x_n and the function values y_0, y_1, \dots, y_n the polynomial P_n defined by

$$P_n(x) = \sum_{k=0}^n \binom{\frac{x-x_0}{h}}{k} \Delta^k y_0$$

where $s = \frac{x-x_0}{h}$ with spacing $h = x_{i+1} - x_i$, is called the **Newton forward difference form of the interpolating polynomial**.