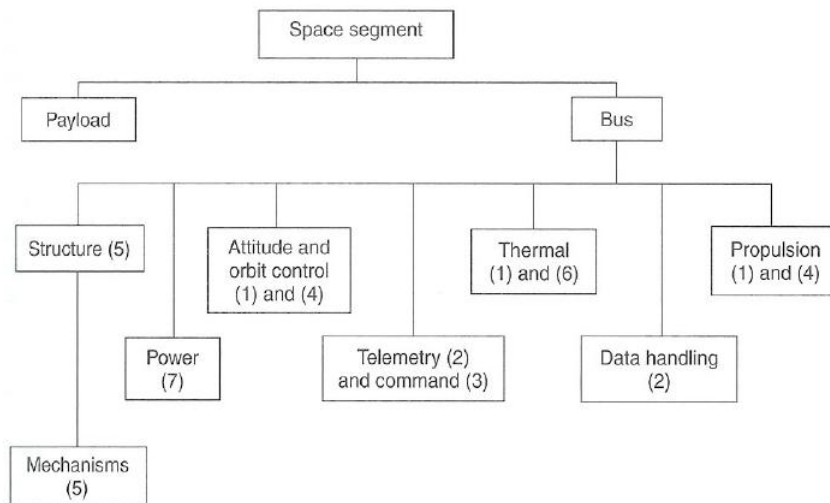


SSD

1 Spacecraft System Design

Mission concept:

- subject (what for)
- orbit and constellation
- payload, bus
- launch element
- ground element
- mission operations
- command, communication, control



2 Space Dynamics/Kepler Orbits

Typical coordinate systems:

- spacecraft-fixed
 - Mittelpunkt des Satelliten = Ursprung
 - nadir = z-Achse, nominale Geschwindigkeit = x-Achse
 - gut, um Position und Orientierung der Satelliteninstrumente festzustellen
- earth-fixed
 - Mittelpunkt der Erde = Ursprung
 - durch greenwich meridian = x-Achse
 - Geolocation, Satellitenbewegung
- roll, pitch and yaw-coordinates
- celestial coordinates
 - Mittelpunkt der Erde = Ursprung
 - Richtung Frühlingspunkt = x-Achse
 - Orbitanalyse, Astronomie

Keplergesetze

1. der Orbit eines jeden Planeten ist eine Ellipse, wobei die Sonne in einem der Fixpunkte liegt
2. die Verbindungslinie zwischen Sonne und Planet überstreicht in gleichen Zeiten gleiche Flächen
3. die Quadrate der Umlaufzeiten sind proportional zu den Kuben der großen Halbachsen

Ellipsendinge

- a ... große Halbachse
- ε , e ... Exzentrizität, "Abplattung" der Ellipse ($\varepsilon=0$: Kreis, $\varepsilon=1$: Parabel, $0 < \varepsilon < 1$: Ellipse)

Begriffe:

- Periapsis: Punkt der Ellipse, der am nächsten an dem Zentralkörper liegt (bei Sonne: Perihel, bei Erde: Perigäum)
- Apoapsis: Punkt der Ellipse, der am weitesten entfernt vom Zentralkörper liegt (bei Sonne: Apohel, bei Erde: Apogäum)
- Distanz zu Periapsis $r_p = a(1 - \varepsilon)$, Distanz zu Apoapsis $r_a = a(1 + \varepsilon)$

6 Bahnelemente:

Lieblingsformel

$$T = 2\pi\sqrt{\frac{a^3}{\mu}}$$

Change of the right ascension of the ascending node

$$\Delta\Omega = -\frac{3\pi J_2 R_E^2}{a^2(1 - \varepsilon^2)^2} \cos(i)$$

Change of the argument of perigee

$$\Delta\omega = \frac{3\pi J_2 R_E^2}{2a^2(1 - \varepsilon^2)^2} (4 - 5\sin^2(i))$$

Orbits

1. Highly Elliptical Orbit HEO

- [illegible]

- hohe Exzentrizität
- große Halbachsen
- dadurch lange Kontaktdauer zum Satelliten
- Werte für Perigäum: 200 bis 15.000 km
- Werte für Apogäum: 50.000 bis 140.000 km
- für Forschung (z.B. Weltraumteleskope), Telekommunikation, Militär
- Beispiel: Molniya-Orbit (feste Inklination von $63,4^\circ$, Periodendauer von einem halben Sterntag (23h56m4s))

The diagram illustrates Earth's orbit around the Sun. The Sun is represented by a large yellow sphere at the center. Earth is shown at four positions along its elliptical orbit, marked with dates: 21/06 (top), 23/09 (left), 21/03 (right), and 22/12 (bottom). Each Earth is depicted as a blue sphere with a red line representing its axis of rotation and a green line representing its orbital path. The angle between the red and green lines is labeled $\Delta\Omega_{SSO}$. At the top position (21/06), the angle is labeled $\Delta\Omega_{Orbit} = 0^\circ/\text{Year}$. At the left position (23/09), the angle is labeled $\Delta\Omega_{SSO}$. At the bottom position (22/12), the angle is labeled $\Delta\Omega_{SSO} = \Delta\Omega_{Earth} = 360^\circ/\text{Year}$. A curved arrow labeled $\Delta\Omega_{Earth}$ indicates the Earth's rotation. The orbit is shown as a dashed white line.

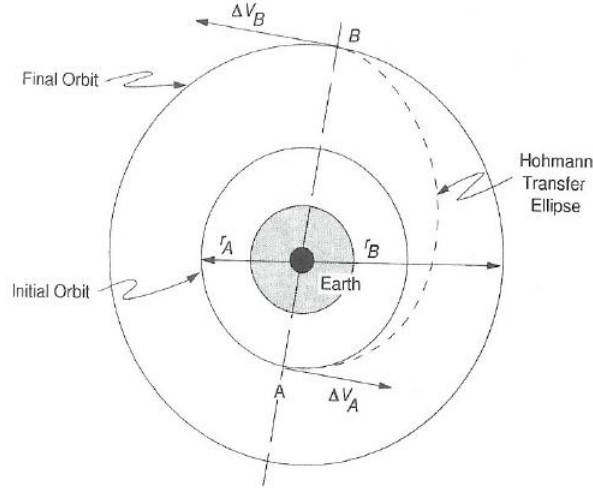
- 4

- Inklination: leicht retrograd ($\approx 98^\circ$)
- Umlaufdauer: 96-100min

3. Geostationary Orbit GEO

- kreisförmiger Orbit
- Höhe: 35.786km
- Umlaufdauer: 24h
- Wettersatelliten, Kommunikationssatelliten, Fernsehsatelliten

Subsatellite Point = intersection of the line between satellite and earth center with the earth's surface
Hohmann Transfer



- Calculate a transfer between two circular orbits with radius r_A to r_B . The velocity at pericenter of the transfer ellipse:

$$v_P^2 = 2\mu \left(\frac{1}{r_A} - \frac{1}{r_A + r_B} \right) = 2\mu \frac{r_B}{r_A(r_A + r_B)}$$

- The required Δv_A to inject from the transfer orbit:

$$\Delta v_A = v_P - v_A = \sqrt{\frac{\mu}{r_A}} \left(\sqrt{\frac{2r_B}{r_A + r_B}} - 1 \right)$$

- The required Δv to inject from the transfer orbit into orbit with r_B :

$$\Delta v_B = v_B - v_{apo} = \sqrt{\frac{\mu}{r_B}} \left(1 - \sqrt{\frac{2r_A}{r_A + r_B}} \right)$$

where v_B is the circular velocity at r_B .

- The Hohmann transfer is the most energy-efficient transfer between two circular orbits.

$$\Delta v_{\text{total}} = \Delta v_A + \Delta v_B = \sqrt{\mu} \left[\sqrt{\left(\frac{2}{r_A} - \frac{2}{r_A + r_B} \right)} - \sqrt{\frac{1}{r_A}} + \sqrt{\frac{2}{r_B} - \frac{2}{r_A + r_B}} - \sqrt{\frac{1}{r_B}} \right]$$

3 Mission Analyses

Earth-Synchronous Orbit

- the ground track repeats after a specific period of time
- Earth's rotation rate is the sidereal rotation period = sidereal day τ_E
- τ_E is varying with time $\tau_E = 86164.10555 + 0.15 \cdot C$ [s] where C is the centuries since year 2000
- as the Earth rotates eastward, the satellite is thus moving relative to the surface in westward direction by

$$\Delta\Phi_r = 2\pi \frac{T}{\tau_E} \text{ [rad/rev]}$$

- second effect influencing the shift of the subsatellite point is the rotation of the satellite's orbit plane $\Delta\Omega$
- as $\Delta\Omega$ is positive in eastward direction, these two effects are combined to the total angular shift $\Delta\Phi$ at subsequent equator passages

$$\Delta\Phi = \Delta\Phi_r - \Delta\Omega \text{ [rad/rev]}$$

- to be Earth-Synchronous:

$$n\Delta\Phi = m \cdot 2\pi$$

Sun-Synchronous Orbit

- die Erde braucht $\tau_S = 3.155815 \cdot 10^7$ s, um einmal um die Sonne zu kreisen
- bei einem sonnensynchronen Orbit muss der Winkel zwischen Sonnenrichtung und Orbitebene konstant bleiben
- also muss sich die Ebene pro Tag um einen Winkel θ drehen

$$\theta = 2\pi \frac{\tau_E}{\tau_S} \text{ [rad/day]} = 2\pi \frac{\tau_E}{\tau_S} \frac{T}{\tau_E} \text{ [rad/rev]}$$

Earth- and Sun-Synchronous Orbit

-

$$\Delta\Omega = \theta \Rightarrow T \left(\frac{1}{\tau_E} - \frac{1}{\tau_S} \right) = \frac{m}{n}$$

- angular shift between two subsequent orbits

$$\Delta\Phi = \Delta\Phi_r - \Delta\Omega = 2\pi T \left(\frac{1}{\tau_E} - \frac{1}{\tau_S} \right) \text{ [rad/rev]}$$

worst case between subsequent orbits $\Delta\Phi \cdot R_E$.

Eclipse periods angle between Earth-Sun vector and normal vector to orbit plane: $\sin \beta = \vec{s} \cdot \vec{n}$

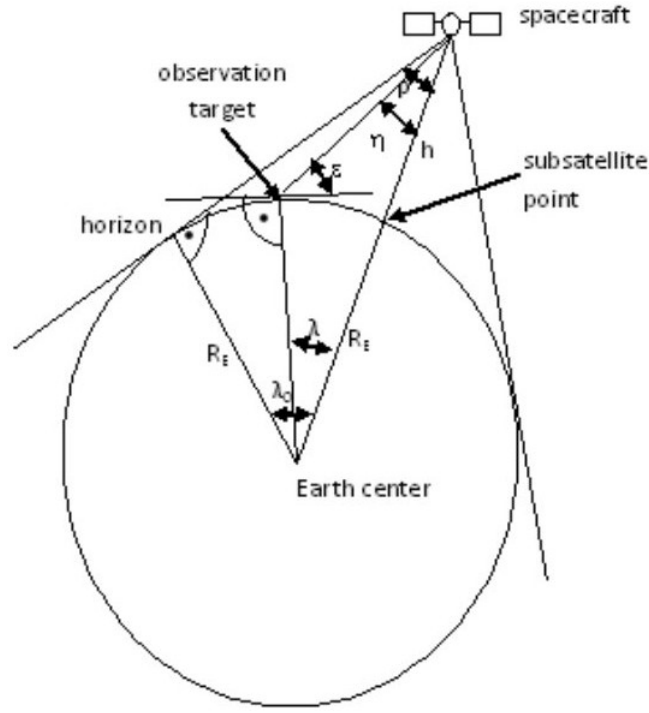
Earth central angular radius at entry into eclipse: $\beta^* = \sin^{-1} \left(\frac{R_E}{h+R_E} \right)$

Angular arc of orbit in shadow: $2 \cos^{-1} \left(\frac{\cos \beta^*}{\cos \beta} \right)$ **Ground Contact and Coverage Analyses** altitude h , visible horizon characterized by angles ρ and λ_0 : $\rho + \lambda_0 = 90^\circ$

$$\begin{aligned} R_E &= (R_E + h) \cos \lambda_0 \\ &= (R_E + h) \sin \rho \end{aligned}$$

observe Λ_t, Θ_t (long,lat) from known orbit position of satellite, characterized by subsatellite point Λ_s, Θ_s . characteristic paramters:

- nadir angle η
- earth central angle λ
- spacecraft elevation angle ε



calculate nadir angle η :

$$\tan \eta = \frac{\frac{R_E}{R_E+h} \sin \lambda}{1 - \frac{R_E}{R_E+h} \cos \lambda}$$

$$\lambda + \eta + \varepsilon = 90^\circ$$

λ_{\max} : maximum earth central angle \Rightarrow swath width $2\lambda_{\max}$ perpendicular to groundtrack on surface.
Time in view T_{view} for circular orbit with period T :

$$T_{\text{view}} = \frac{T}{180^\circ} \cos^{-1} \left(\frac{\cos \lambda_{\max}}{\cos \lambda} \right)$$

ground station contact periods



- $\sin \eta_{\max} = \cos \varepsilon_{\min} \frac{R_E}{R_E+h}$
- $\lambda_{\max} = 90^\circ - \varepsilon_{\min} - \eta_{\max}$
- max range satellite \leftrightarrow groundstation: $D_{\max} = R_E \frac{\sin \lambda_{\max}}{\sin \eta_{\max}}$
- total time in view: $T_{\text{view}} = \frac{T}{180^\circ} \cos^{-1} \left(\frac{\cos \lambda_{\max}}{\cos \lambda_{\min}} \right)$
- contact only possible, if station-orbit angle $<$ central angle of contact cone

4 Distributed Satellite Systems

- **constellation**: similar trajectories without relative position control.
- **formation**: closed-loop onboard control for topology in the group.
- **swarm**: similar vehicles cooperating without fixed positions, selfdetermined.
- **cluster**: heterogenous system of vehicles for joint objective.

requirements on distributed satellite systems: coordination of

- orbits at different altitudes
- optimal control strategies for position/attitude of components
- activities for heterogenous sensors
- information flow and storage

Walker Delta Pattern Constellation

$$i : t/p/f$$

- i : inclination
- t : total # satellites
- p : # equally shaped orbit planes
- f : relative phase difference between satellites in adjacent planes

Example: Galileo is $56^\circ : 27/3/1$ with circular orbits ($h = 23222km$), nine satellites always in view, one spare satellite in each plane. **earth surface coverage** $s = \frac{t}{p}$ number of satellites equally spaced in plane with angular distance $\Delta v = \frac{360^\circ}{s}$. There are two cases:

- $\Delta v < 2 \cdot \lambda_{\max} \Rightarrow$ area of continuous coverage exists (“street of coverage”)
- $\Delta v > 2 \cdot \lambda_{\max} \Rightarrow$ no street of coverage

Street-width: $\cos \lambda_{\text{street}} = \frac{\cos \lambda_{\max}}{\cos \frac{\Delta v}{2}}$

formation flying architectures and dynamics

- **virtual structure**: treated as single structure
- **behavioral strategies**: distributed control approach, following nature.
- **leader-follower**: divided into leaders and followers. followers track designated leaders with prescribed offset. absolute/relative control architecture.

communication in low-earth orbit distributed satellite systems

- comm and tele-operation infrastructure is key element for distributed systems
- transfer position and observation data for formation flying
- amount of data increases with swarm size
- analyse pre-processing procedures, intersatellite links and ground station links

conclusion on distributed satellite systems

- research field due to paradigm shift from one large spacecraft to several smaller crafts
- higher fault tolerance and robustness
- swarms are scaleable
- gun launches into orbit
- combination of big and small spacecrafts
- swarms for surveillance and earth observation
- LEO \rightarrow high spatial resolution
- higher temporal resolution is provided by constellations with several satellites in the same orbit

5 Mechanics

mechanical system engineering

- **mechanical specs:** requirements on satellite, components and equipment
- **verification plan:** “how to prove satellite complies specs?” test, simulations, similarity
- **test plan:** test flow, model philosophy (QM/FM↔PFM)
- **design loads:** simplified load cases for components & equipment

requirements on satellite structures

- external shape
- mass, center of gravity
- resonance frequency
- thermo-elastic distortion
- interfaces
- environment (vacuum, debris, etc.)
- margin of safety

random vibration loads

- white noise: range $20 - 2000Hz$, max levels at $80 - 300Hz$
- mainly acoustic excitation under fairing
- depends on location, orientation, mass
- equivalent design loads: 3 times root mean square (3 sigma value)

shock events

- launcher-induced: stage separation, fairing
- S/C release: clampband, discrete pyro devices
- appendage release: protechnic/deployment shock

structural engineering – fundamentals

- hooke’s law: $\sigma = E \cdot \varepsilon$
- strain def: $\varepsilon = \frac{\Delta L}{L}$
- normal stress in rod: $\sigma = \frac{F}{A}$
- bending stress in beam: $\sigma = \frac{M}{W}$
- thermo-elastic strain: $\varepsilon = CTE \cdot \Delta T$

Margin of safety:

$$MOS = \left(\frac{S_a}{S_e \cdot FOS} - 1 \right) \cdot 100 \stackrel{!}{\geq} 0,0 \quad [\%]$$

- S_a allowable stress
- S_e applied stress
- FOS Factor of safety

safety factors:

- material safety factors on yield
- modelling safety factors covering analysis uncertainties
- specific factors, e.g. for bonded connections

material selection driven by

- ratio stiffness/mass
- ratio strength/mass
- functional aspects
- compability with environment
- thermoelastic behavior
- manufacturing complexity
- sources
- cost

test facilities

- electro-dynamic shaker testing orthogonal axes
- tasks: system identification – model correlation – model adaption

notching – reduction of dynamic loads – prevents exceeding design limit loads, preventing satellite structure damage. primary notching – whole satellite ↔ secondary notching – satellite subsystems

tests

- acoustic noise tests
- separation tests

typical loads:

static	$< 10g$
quasi-static	up to $100g$
sine	up to $100g$
random	sometimes $> 100g$ (3 Sigma)
acoustic noise	$100dB$
shock	$2000g$

critical requirements

- zero grav
- launch loads
- extreme temperatures
- vacuum – lubrication is critical
- no maintainance

6 Thermal Engineering

general task

- engineering: define requirements and design
- analysis: establish thermal mathematical model (TMM), perform distribution calculations
- test: plan, perform, evaluate realistic tests

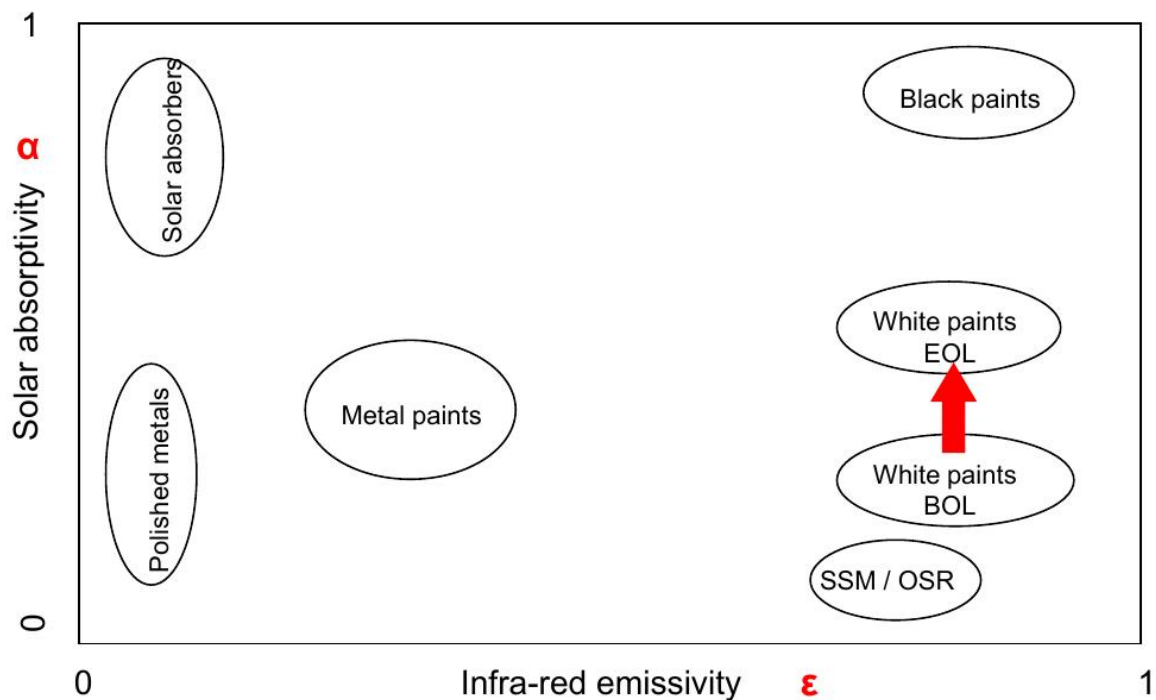
satellite thermal control – requirements

- temperature limits!
- temperature gradients, -stability, -uniformity
- heatflux, -storage
- power and mass-allocation

heat mechanisms:

- radiation: transfer via electromagnetic waves
- conduction: transfer via fluids and solids in absense of fluid motion
- convection: transfer in a flowing fluid and between fluid and wall (mostly irrelevant in satellites)

thermo optical properties of materials



global energy balance

- | | | | |
|--------------------------|---------------------------------------------|----------------------------|---------------------------------------|
| • Solar intensity | $\dot{q}_S = 1316 \dots 1428 \frac{W}{m^2}$ | • Earth albedo | $\dot{q}_A = 0.2 \dots 0.4 \dot{q}_S$ |
| • Earth IR Radiation | $\dot{q}_E = 189 \dots 261 \frac{W}{m^2}$ | • Absorbtivity Coefficient | $\alpha = 0 \dots 1$ |
| • Emissivity Coefficient | $\epsilon = 0 \dots 1$ | • View Factor | $\phi = 0 \dots 1$ |

orbital environment and load cases:

- external loads (sun, earth, moon...)
- LEO, GEO, lagrange, eclipse?
- orientation (earth, sun, deepspace?)
- operating modes

- mission scenario

typical design load cases

	cold case	hot case
environment	cold external, BOL	hot external, EOL
solar intensity	min ($\sim 1320 \frac{W}{m^2}$)	max ($\sim 1420 \frac{W}{m^2}$)
earth albedo	min (~ 0.2)	max (~ 0.4)
earth IR radiation	min ($\sim 200 \frac{W}{m^2}$)	max ($\sim 260 \frac{W}{m^2}$)

thermal design – approach

- insulate against environment
- minimize absorbed heat
- balance internal heat
- define radiator-areas and distribute heat
- install thermal control hardware
- analyze and verify TCS (thermal control system)

critical components

- batteries
- detectors and sensors (instrument & star)
- optical equipment
- mechanisms, tubes, propulsion systems

thermal control hardware

- insulation, surface coating
- thermal interfillers
- thermal doublers, heat straps, heatpipes
- electrical heaters, temperature sensors

7 Rocket Propulsion

8 TT&C

9 Power Generation

10 Power System

11 Thermal Testing

12 Spacecraft Operations