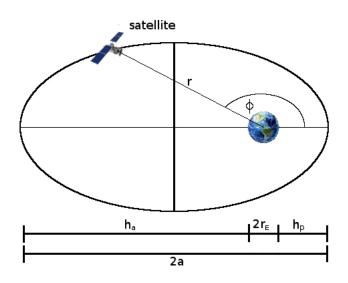
# Formeln SSD

# Bahn- und Orbitdinge

3. Keplersches Gesetz (Quadrate der Umlaufdauern proportional zu den Kuben der großen Halbachsen)

$$T = 2\pi \sqrt{\frac{a^3}{\mu}}$$

Ellipsendinge:



$$2a = h_a + h_p + 2r_E$$
 
$$r_p = h_p + r_E$$
 
$$r_a = h_a + r_E$$
 
$$r_p = a(1 - \varepsilon)$$
 
$$r_a = a(1 + \varepsilon)$$
 
$$\varepsilon = \frac{r_p - r_a}{r_p + r_a}$$

Vis-Viva-Gleichung/Binet-Gleichung:

$$v = \sqrt{\mu \left(\frac{2}{r} - \frac{1}{a}\right)}$$

Vis-Viva-Gleichung für Kreis:

$$r = a \Rightarrow v = \sqrt{\frac{\mu}{r}}$$

für Erde:

$$v = \sqrt{\frac{\mu}{r_E}} = 7.905 \frac{km}{s}$$

(= erste kosmische Geschwindigkeit = minimale Geschwindigkeit, die ein Objekt haben muss, um in einen Erdorbit zu gelangen)

Vis-Viva-Gleichung für Parabel:

$$a\to\infty\Rightarrow v=\sqrt{\frac{2\mu}{r}}$$

1

für Erde:

$$v = \sqrt{\frac{2\mu}{r_E}} = 11.179 \frac{km}{s}$$

(= zweite kosmische Geschwindigkeit = minimale Geschwindigkeit, die ein Objekt haben muss, um das Gravitationsfeld der Erde zu verlassen)

### Geschwindigkeitsdinge für Ellipse:

average angular velocity:

$$\frac{2\pi}{T} \frac{\text{rad}}{\text{min}} \quad \frac{360^{\circ}}{T} \frac{\text{deg}}{\text{min}}$$

maximale Geschwindigkeit (Periapsis):

$$v_{max} = \sqrt{\mu \left(\frac{2}{h_p + r} - \frac{1}{a}\right)}$$

minimale Geschwindigkeit (Apoapsis):

$$v_{min} = \sqrt{\mu \left(\frac{2}{h_a + r} - \frac{1}{a}\right)}$$

wahre Anomalie  $(\Phi, \nu, ...)$ 

$$r = a \frac{1 - \varepsilon^2}{r \cdot \varepsilon \cdot \cos(\nu)} \Rightarrow \nu = \arccos\left(a \frac{1 - \varepsilon^2}{r^2 \cdot \varepsilon}\right)$$

wahre Anomalie ist auch über die exzentrische Anomalie berechenbar:

$$\tan\left(\frac{\nu}{2}\right) = \sqrt{\frac{1+\varepsilon}{1-\varepsilon}} \cdot \tan\left(\frac{E}{2}\right)$$

$$E - \varepsilon \cdot sin(E) = \frac{2\pi}{T}(t - t_{perigee})$$

right ascension of the ascending node:

$$\begin{split} \Delta\Omega_{rev} &= -\frac{3\pi J_2 r_E^2}{a^2 (1-\varepsilon^2)^2} cos(i) & \text{in} \left[\frac{\text{rad}}{\text{rev}}\right] \\ \Delta\Omega_{rev} &= -\frac{3\pi J_2 r_E^2}{a^2 (1-\varepsilon^2)^2} cos(i) \cdot \frac{360^\circ}{2\pi} & \text{in} \left[\frac{\text{deg}}{\text{rev}}\right] \\ \Delta\Omega_{day} &= -\frac{3\pi J_2 r_E^2}{a^2 (1-\varepsilon^2)^2} cos(i) \cdot \frac{\tau_E}{T} & \text{in} \left[\frac{\text{rad}}{\text{day}}\right] \\ \Delta\Omega_{rev} &= -\frac{3\pi J_2 r_E^2}{a^2 (1-\varepsilon^2)^2} cos(i) \cdot \frac{360^\circ}{2\pi} \cdot \frac{\tau_E}{T} & \text{in} \left[\frac{\text{deg}}{\text{day}}\right] \end{split}$$

argument of perigee:

$$\Delta\omega = \frac{3\pi J_2 r_E^2}{2a^2 (1 - \varepsilon^2)^2} (4 - 5sin^2(i)) \qquad \text{in} \left[\frac{\text{rad}}{\text{rev}}\right]$$

angular displacement after n siderial days:

$$\Delta\Omega_n = \frac{n}{T}[\text{rev}] \cdot \Delta\Omega \left[\frac{\text{deg}}{\text{rev}}\right]$$

#### Subsatellitenpunkt

latitude of SSP:

$$\Theta = \arcsin(\sin(i) \cdot \sin(\omega + \nu))$$

longitude of SSP:

$$\lambda = \arccos\left(\frac{\cos(\omega + \nu)}{\cos(\Theta)}\right) + \Omega + \lambda_{greenwich}$$

#### Hohmann-Transfer:

große Halbachse der Transferellipse:

$$a_{transfer} = \frac{r_A + r_B}{2}$$

Transferdauer:

$$T_{transfer} = \pi \sqrt{\frac{a_{transfer}^3}{\mu}}$$

Geschwindigkeitsänderung, um ersten Orbit zu verlassen (beschleunigen):

$$\Delta v_1 = \sqrt{\frac{\mu}{r_A}} \left( \sqrt{\frac{2r_B}{r_A + r_B}} - 1 \right)$$

Geschwindigkeitsänderung, um auf zweiten Orbit zu kommen (abbremsen):

$$\Delta v_2 = \sqrt{\frac{\mu}{r_B}} \left( 1 - \sqrt{\frac{2r_A}{r_A + r_B}} \right)$$

### noch mehr Bahnparameterdinge:

earth angular radius:

$$\rho = \arcsin\left(\frac{r_E}{r_E + h}\right)$$

maximum nadir angle:

$$\eta_{max} = arcsin(sin(\rho) \cdot cos(\varepsilon_{min}))$$

maximum earth central angle:

$$\lambda_{max} = 90^{\circ} - \varepsilon_{min} - \eta_{max}$$

maximum distance:

$$D_{max} = r_E \left( \frac{sin(\lambda_{max})}{sin(\eta_{max})} \right)$$

minimum earth central angle:

$$\lambda_{min} = arcsin(sin(90^{\circ} - i) \cdot sin(lat_{gs}) + cos(90^{\circ} - i) \cdot cos(lat_{gs}) \cdot cos(long_{gs} - L_{node} + 90^{\circ}))$$

minimum nadir angle:

$$\eta_{min} = \arctan\left(\frac{\sin(\rho) \cdot \sin(\lambda_{min})}{1 - \sin(\rho) \cdot \cos(\lambda_{min})}\right)$$

maximum elevation angle:

$$\varepsilon_{max} = 90^{\circ} - \lambda_{min} - \eta_{min}$$

minimum distance:

$$D_{min} = r_E \left( \frac{sin(\lambda_{min})}{sin(\eta_{min})} \right)$$

maximum angular rate:

$$\dot{\Theta}_{max} = \frac{2\pi(r_E + h)}{T \cdot D_{min}}$$

azimuth range:

$$\Delta \Phi = 2 \cdot arccos\left(\frac{tan(\lambda_{min})}{tan(\lambda_{max})}\right)$$

time in view:

$$T_{view} = \frac{T}{180^{\circ}} \cdot arccos\left(\frac{cos(\lambda_{max})}{cos(\lambda_{min})}\right)$$

Eclipsedinge:

$$\beta = \arcsin\left(\frac{r_E}{r_E + h}\right)$$

eclipsed fraction of a circular orbit:

$$F_e = \frac{1}{\pi} \left( \frac{\sqrt{h^2 + 2r_E \cdot h}}{(r_E + h) \cdot \cos(\beta)} \right)$$

Dezibeldinge:

$$P[W] = 1W \cdot 10^{\frac{P[dBW]}{10}}$$
$$P[dBW] = 10 \cdot log_{10} \left(\frac{P[W]}{1W}\right)$$

gain:

$$G = 10 \cdot log_{10} \left(\frac{P_1}{P_2}\right)$$

EIRP:

$$EIRP = G + P$$

Thermaldinge:

$$Q = A_r \cdot \varepsilon_{IR} \cdot \sigma \cdot T^4$$
 
$$\dot{Q} = \lambda \cdot a \cdot b \cdot \frac{T_1 - T_2}{c} \qquad \text{oder auch} \qquad \dot{Q} = K \cdot A \cdot \frac{T_1 - T_2}{L}$$

series conduction:

$$\sigma_{tot} = \frac{1}{\frac{1}{\sigma_1} + \frac{1}{\sigma_2} + \frac{1}{\sigma_3} + \dots}$$

$$\sigma = \frac{L}{K \cdot A} \quad \text{oder} \quad \sigma = \frac{1}{\alpha \cdot A}$$

Free Space Path Loss:

$$L = \left(\frac{4\pi \cdot d}{\lambda}\right)^2 = \left(\frac{4\pi \cdot d \cdot f}{c}\right)^2$$

oder auch (ungenauer)

$$L = 92.4dB + 20 \cdot log_{10}(F[GHz]) + 20 \cdot log_{10}(d[km])$$

thermal noise:

B...bandwidth, R...resistor, T...temperature, k...Boltzmann constant  $(1.38\cdot 10^{-23}\frac{J}{K})$ 

$$V_n = \sqrt{4kTBR}$$

Noise Power:

 $T_e$  = system receiver noise temperature

$$N[dBW] = 10 \cdot log_{10}(kT_eB)$$

Noise Figure:

$$NF[dB] = 10 \cdot log_{10} \left( 1 + \frac{T_e}{290^{\circ}} \right) = 10 \cdot log_{10}(F)$$

Noise Factor:

$$F = \frac{\left(\frac{S}{N}\right)_{in}}{\left(\frac{S}{N}\right)_{out}} = 1 + \left(\frac{T_e}{290^{\circ}}\right)$$

# Fermienergie:

Wahrscheinlichkeit, dass ein Teilchen mit Energie E

$$f(E) = \frac{1}{1 + e^{\frac{E - E_F}{k \cdot T}}}$$

## Power System Dinge:

solar array power during daylight:

$$P_{sa} = \frac{\frac{P_e \cdot T_e}{X_e} + \frac{P_d \cdot T_d}{X_d}}{T_d}$$

 $P_e$ ...power in eclipse,  $T_e$ ...time in eclipse,  $P_d$ ...power in daylight,  $T_d$ ...time in daylight,  $X_e$ ...path efficiency during eclipse,  $X_d$ ...path efficiency during daylight

spacecraft power budget:

$$P_{sg} = \frac{\frac{P_e \cdot T_e}{\eta_L \cdot \eta_b} + \frac{P_s \cdot T_s}{\eta_L}}{T_s}$$

 $P_{sg}$ ...total power from solar array,  $\eta_L$ ...losses between solar array and loads,  $\eta_b$ ...losses of battery,  $P_S$ ...available power in sunlight,  $P_e$ ...power demand in eclipse

# wichtige Parameter für BOL/EOL-Berechnungen:

- $\eta$  ... production efficiency of solar cells (14-22%)
- X...path efficiency from solar array through batteries to loads (eclipse and daytime):  $X_e = 0.65, X_d = 0.85$  (direct energy transfer),  $X_e = 0.60, X_d = 0.80$  (peak power tracking)
- $I_d$  ... inherent degradation ( $\approx 0.77$ , ranges from 0.49 0.88)
- ullet  $\Theta$  ... angle between array normal and sun vector; typically use worst-case sun-angle
- $L_d$  ... life degradation: micrometeorites, radiation, etc. (2-4% per year)

$$L_d = (1 - \text{degradation per year})^{\text{satellite life}}$$

## From Begin of Life to End of Life

$$P_o = \eta \cdot 1358 \frac{W}{m^2}$$
 (output power) 
$$P_{BOL} = P_o \cdot I_d \cdot cos(\Theta)$$
 
$$P_{EOL} = L_d \cdot P_{BOL}$$

solar array size to meet power requirement:

$$A_{sa} = \frac{P_{sa}}{P_{EOL}}$$

mass of solar array ranges from 14 to  $47\frac{W}{kg}$ :

$$M_{sa} = 0.04 \cdot P_{sa} \qquad \text{(for } 25 \frac{W}{kq}\text{)}$$

## **Battery Sizing**

Power need:

$$P_{avg} = V_{bus} \cdot I$$
 
$$Ah_{avg} = \frac{T_e}{1h} \cdot I$$
 
$$Ah_{total} = \frac{Ah_{avg}}{DoD}$$

Capacity:

$$C_r = \frac{P_{avg} \cdot T_e}{DoD \cdot N_{bat} \cdot \eta}$$