

Exercise 4 - Solution

Task 4.1

1.

$$\lambda = \frac{90^\circ}{2} = 45^\circ$$

$$\eta = 90^\circ - \lambda - \varepsilon_{min} = 42^\circ$$

$$\rho = \arcsin\left(\frac{\sin(\eta)}{\cos(\varepsilon_{min})}\right) = 42.07^\circ$$

$$h = r_E \left(\frac{1}{\sin(\rho)} - 1 \right) = 3141 km$$

2.

$$\lambda = \frac{90^\circ}{3} = 30^\circ$$

$$\rho = 57.12^\circ$$

$$h = 1217 km$$

Task 4.2

given: $R_E = 6378 km$, $\mu_E = 398600 \frac{km^3}{s^2}$, $h = 900 km$, $J_2 = 1082.63 \cdot 10^{-6}$, $\tau_e = 86164.10555 + 0.015 Cs$, $\tau_s = 3.155815 \cdot 10^7 s$

1.

$$T = 2\pi \sqrt{\frac{a^3}{\mu_E}} = 2\pi \sqrt{\frac{(h + R_E)^3}{\mu_E}} = 6179.16 s$$

$$v = \sqrt{\frac{\mu_E}{a}} = \sqrt{\frac{\mu_E}{R_E + h}} = 7.40 \frac{km}{s}$$

2.

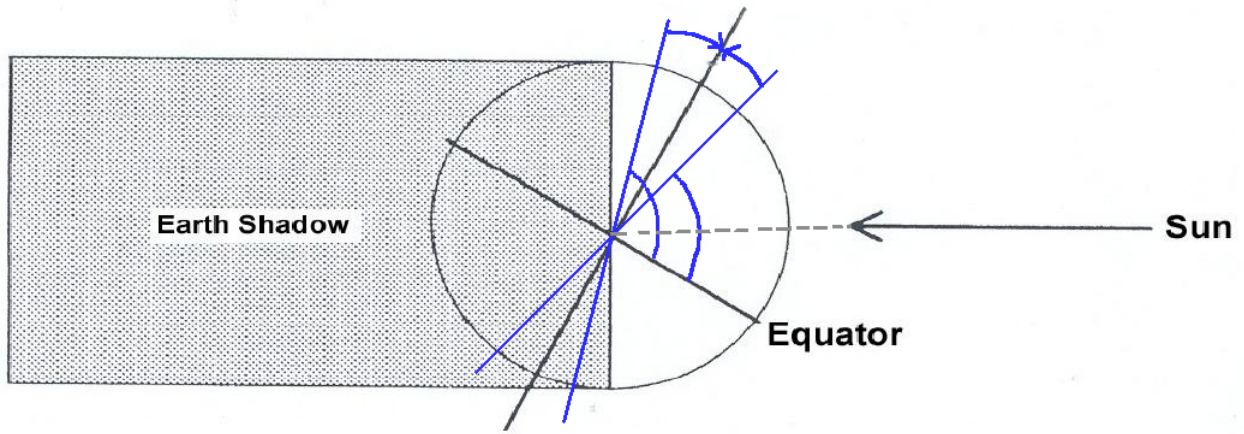
$$\Delta\Omega_{day} = 360^\circ \frac{\tau_e}{\tau_s} = 0.983 \frac{deg}{day}$$

$$\Delta\Omega_{rev} = 360^\circ \frac{\tau_e}{\tau_s} \cdot \frac{T}{\tau_e} = 0.0705 \frac{deg}{rev}$$

$$\Delta\Omega_{rev} = -\frac{3\pi J_2 R_E^2}{a^2 (1 - \epsilon^2)^2} \cdot \cos(i)$$

circular orbit $\Rightarrow \epsilon=0$

$$\Rightarrow i = \arccos\left(-\frac{\Delta\Omega_{rev} a^2}{3\pi J_2 R_E^2}\right) = 99.03^\circ$$



$$\beta^* = \arcsin\left(\frac{R_E}{R_E + h}\right) = 61.20^\circ$$

- orbit 1:

$$\beta_1 = 90^\circ - 23.5^\circ + 9.03^\circ = 75.53^\circ$$

$$\Rightarrow \beta_1 > \beta^* \Rightarrow \text{no eclipse}$$

- orbit 2:

$$\beta_2 = 90^\circ - 23.5^\circ - 9.03^\circ = 57.53^\circ$$

$$\Rightarrow \beta_2 < \beta^* \Rightarrow \text{eclipse}$$

$$F_e = \frac{1}{\pi} \arccos\left(\frac{\sqrt{h^2 + 2R_E h}}{(R_E + h)\cos(\beta_2)}\right) = 0.1456 \approx 15\%$$

$$T_{ecl} = T \cdot F_e = 14.99 \text{ min}$$

Both orbits will have an eclipse since the earth rotation axis stays fixed. The shadow periods of orbit 2 occur half a year later for orbit 1 which was originally eclipse free.

Task 4.3

given: $R_E = 6378 \text{ km}$, $\mu_E = 398600 \frac{\text{km}^3}{\text{s}^2}$, $J_2 = 1082.63 \cdot 10^{-6}$, $i_e = 0^\circ$, $i = 53^\circ$

- 1.

$$T = 2\pi \sqrt{\frac{(R_E + h)^3}{\mu_E}} = 5801.07 \text{ s}$$

$$v = \sqrt{\frac{\mu_E}{(R_E + h)}} = 7.55 \frac{\text{km}}{\text{s}}$$

- 2.

$$\Delta\Omega = -\frac{3\pi J_2 R_E^2}{(R_E + h)^2 (1 - \epsilon^2)^2} \cdot \cos(i)$$

a) circular orbit $\Rightarrow \epsilon=0$

$$\Delta\Omega = -\frac{3\pi J_2 R_E^2}{(R_E + h)^2} \cdot \cos(i) = -0.00513 \frac{rad}{rev} = -0.076197 \frac{rad}{day}$$

b)

$$t_{ecl} = \frac{2 \cdot \beta^*}{360^\circ} = \frac{2 \cdot \arcsin\left(\frac{r_E}{r_E+h}\right)}{360^\circ} = 2129.19s = 35.49min$$

c) given: $r_a = (600 + 6378)km$, $r_b = (900 + 6378)km$

$$\Delta v_1 = \sqrt{\frac{\mu_E}{r_a}} \left(\sqrt{\frac{2r_b}{r_a + r_b}} - 1 \right) = 0.07911 \frac{km}{s}$$

$$\Delta v_2 = \sqrt{\frac{\mu_E}{r_b}} \left(1 - \sqrt{\frac{2r_a}{r_a + r_b}} \right) = 0.07828 \frac{km}{s}$$