

Radio luminosity functions with machine learning and Radio Galaxy Zoo

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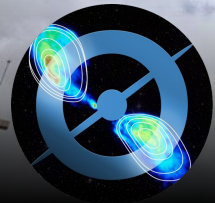
O. Ivy Wong (ICRAR/UWA)

Cheng Soon Ong (Data61/ANU)

Naomi McClure-Griffiths (ANU)

Slides:

<http://www.mso.anu.edu.au/~alger/uhh>



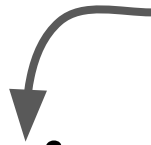
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National
University



Radio luminosity functions

with machine learning and Radio Galaxy Zoo

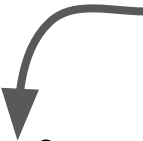
Comoving density of radio sources
as a function of radio luminosity



Radio luminosity functions

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Comoving density of radio sources
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Radio luminosity functions

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Approximating functions
based on existing data

Radio luminosity functions with machine learning and Radio Galaxy Zoo

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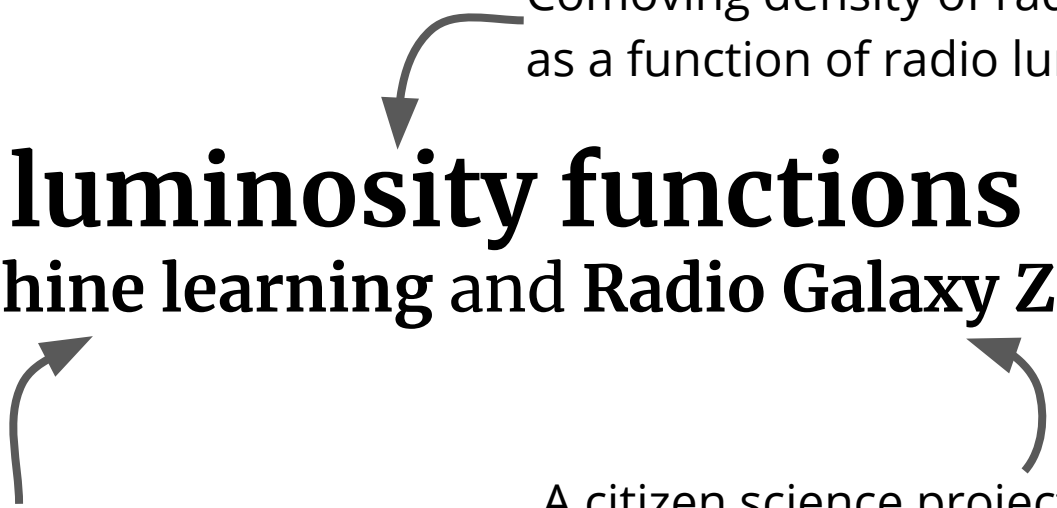


Approximating functions
based on existing data

A citizen science project for
matching radio emission to
infrared galaxies

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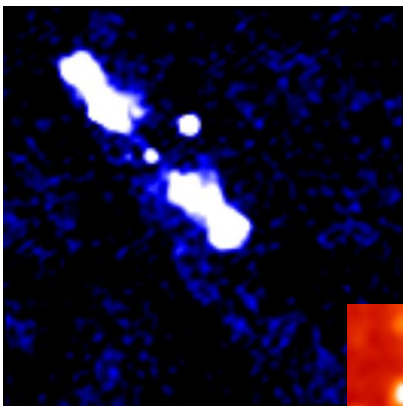
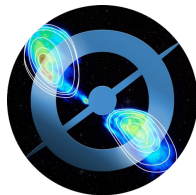


Approximating functions
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A citizen science project for
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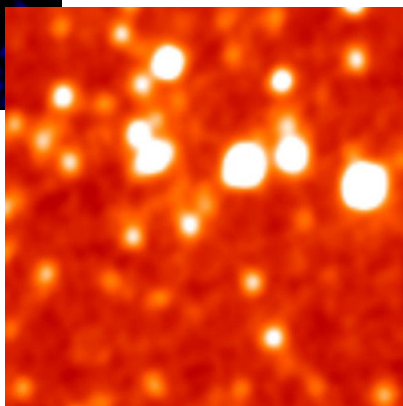
Use machine learning to approximate how citizen scientists cross-identify, then cross-identify everything and make a luminosity function with a huge sample size

Radio Galaxy Zoo

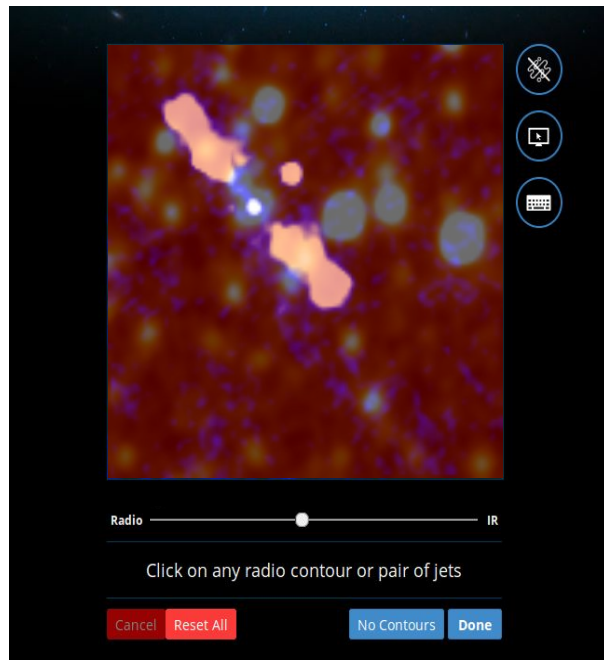


1.4 GHz radio
(FIRST)

3.4 μm infrared
(WISE)

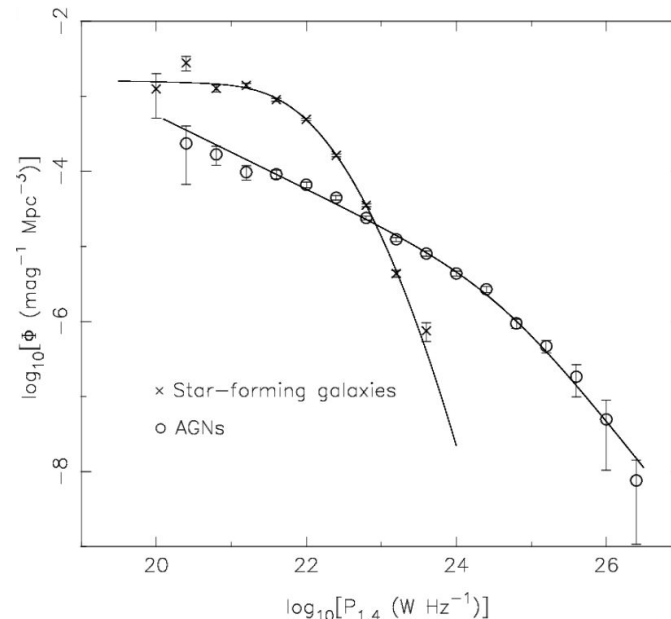


Citizen scientists cross-identify radio emission with infrared host galaxies



Radio luminosity functions

- Comoving density of radio sources as a function of radio luminosity
 - Units of $\text{dex}^{-1} \text{ Mpc}^{-3}$
 - Distribution of radio source luminosities in a *physically meaningful* way
- Fractional radio luminosity functions
 - Luminosity distribution of physically-selected subsets may be different
 - Helps understand evolution and structure of radio galaxies



Radio luminosity function divided into radio due to star formation and radio due to active galactic nuclei.

Image: Mauch & Sadler (2007)

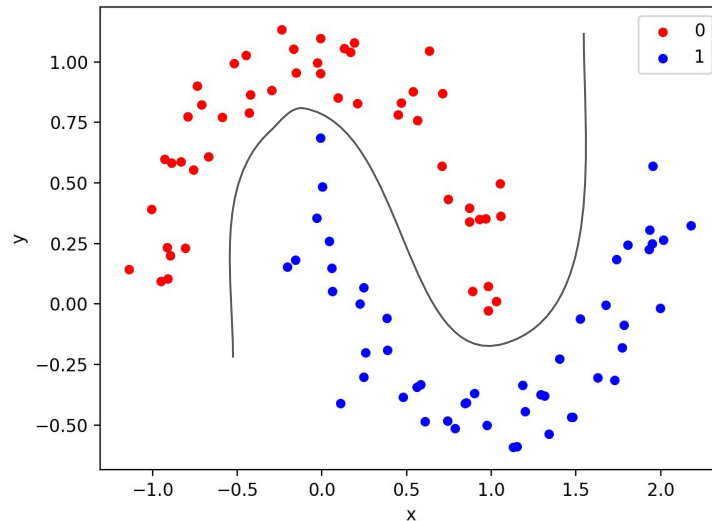
Binary classification

- Find a function that separates objects into two classes
- Well-understood

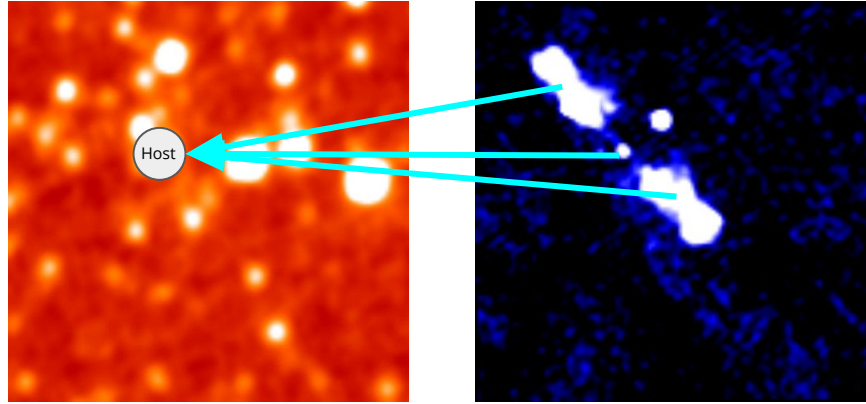
Equivalent:

$$\left\{ \begin{array}{l} f: \mathbb{R}^d \rightarrow \mathbb{R} \\ g: \mathbb{R}^d \rightarrow [0, 1] \\ h: \mathbb{R}^d \rightarrow \{\top, \perp\} \end{array} \right.$$

$h(x) = g(x) > 0$
 $g(x) = \sigma(f(x))$

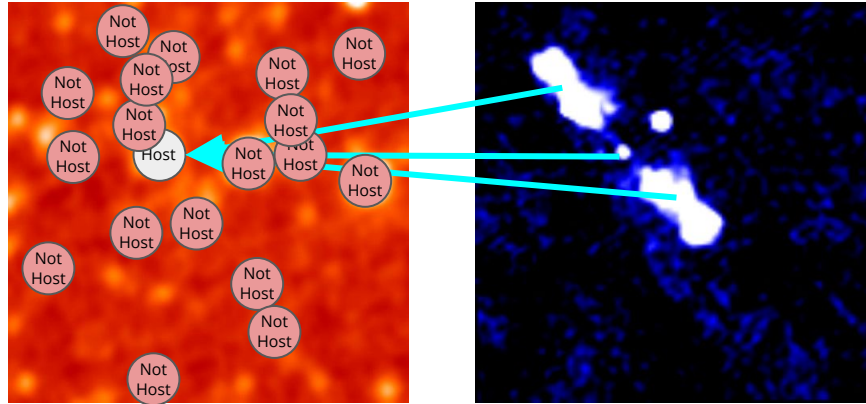


Learning from Radio Galaxy Zoo



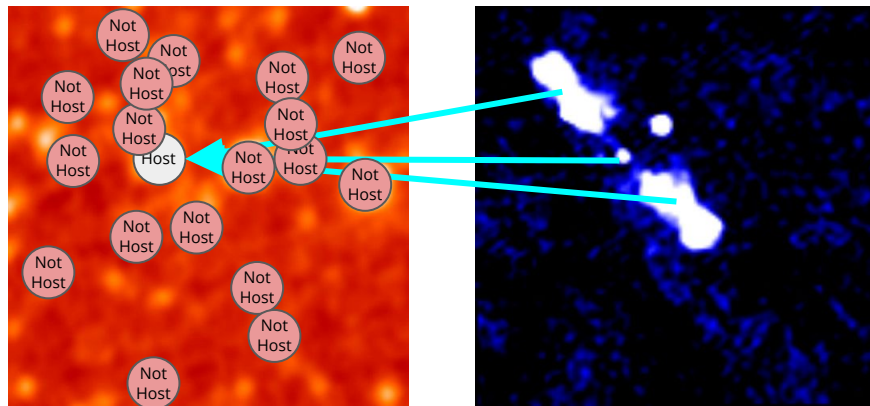
- Assign hosts positive labels

Learning from Radio Galaxy Zoo

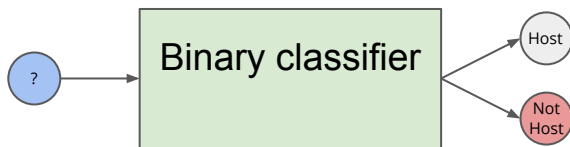


- Assign hosts positive labels
- Assign everything else negative labels

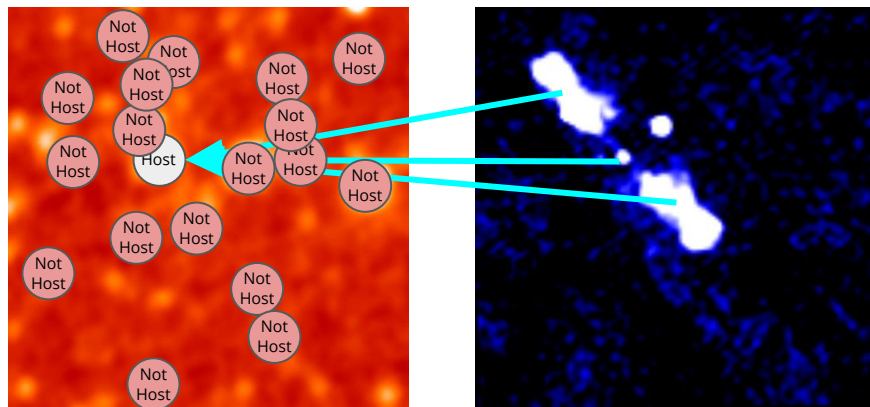
Learning from Radio Galaxy Zoo



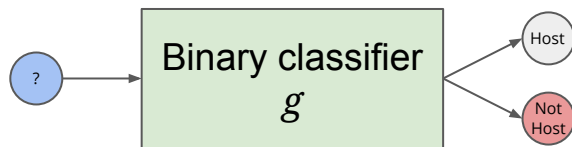
- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify *host* and *not host* classes



Learning from Radio Galaxy Zoo



- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify *host* and *not host* classes



$\text{xid} : \text{Radio} \rightarrow \text{IR}$

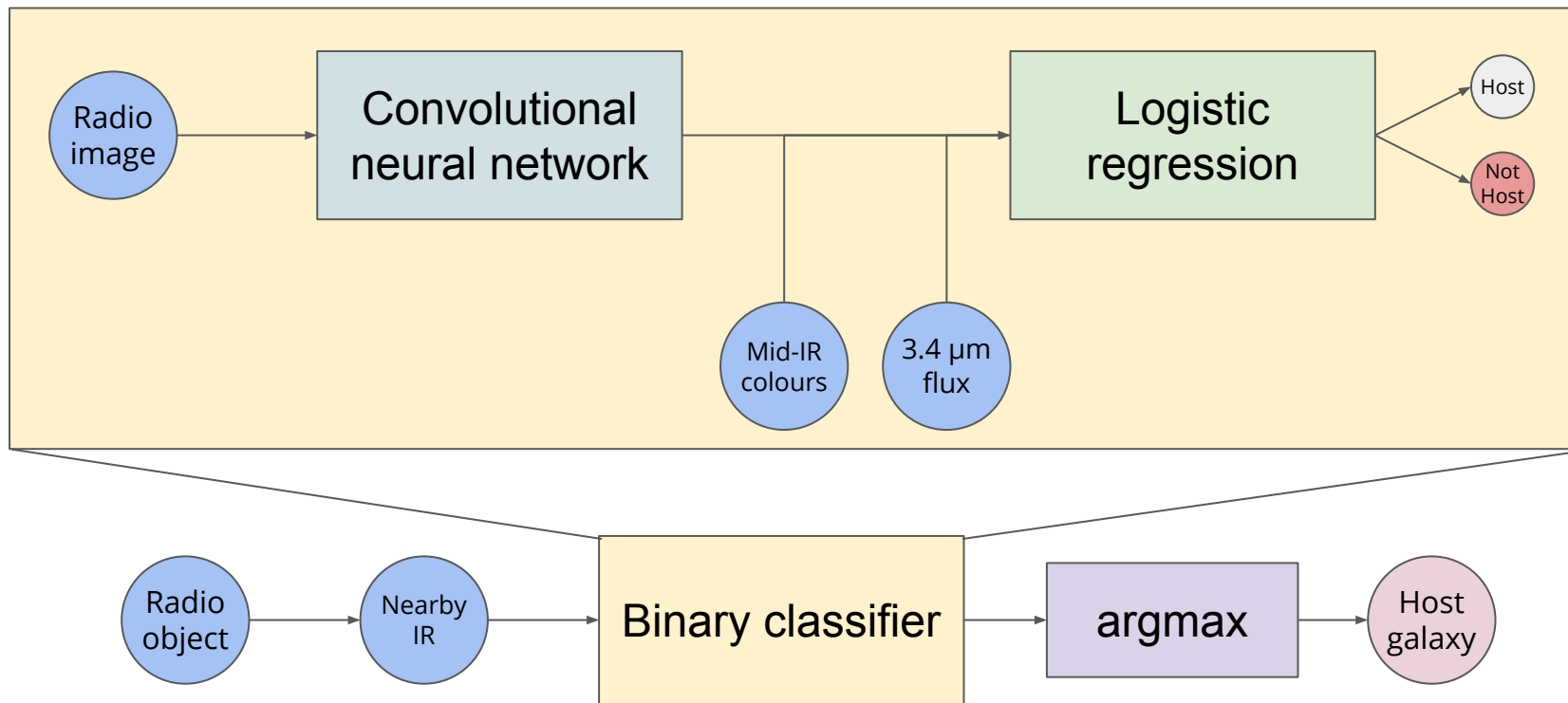
$$\text{xid}(\mathbf{r}) = \underset{i \in \text{IR objects}}{\operatorname{argmax}} g(i) \mathcal{N}(\mathbf{r}, i)$$

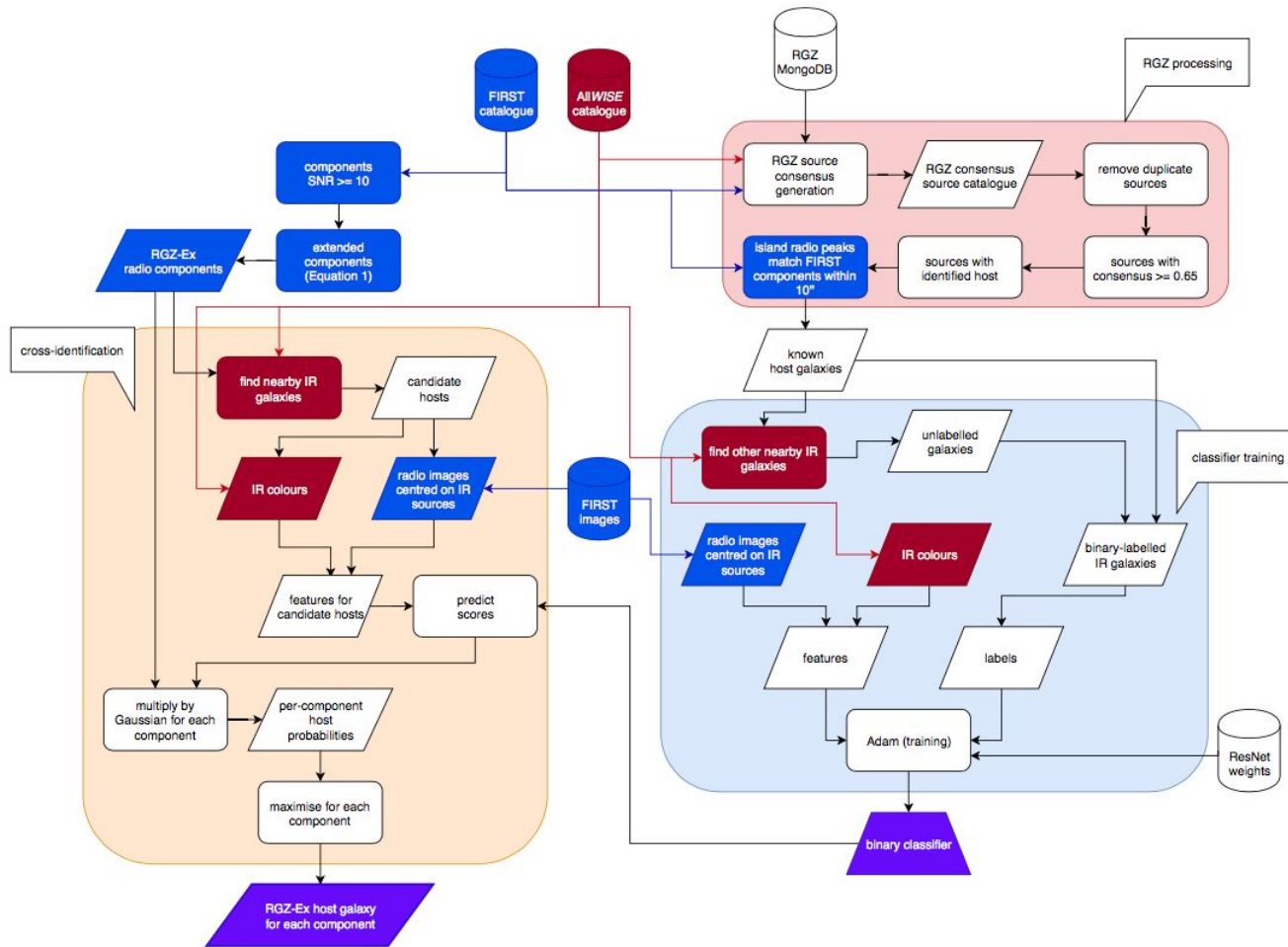
Cross-identification as binary classification

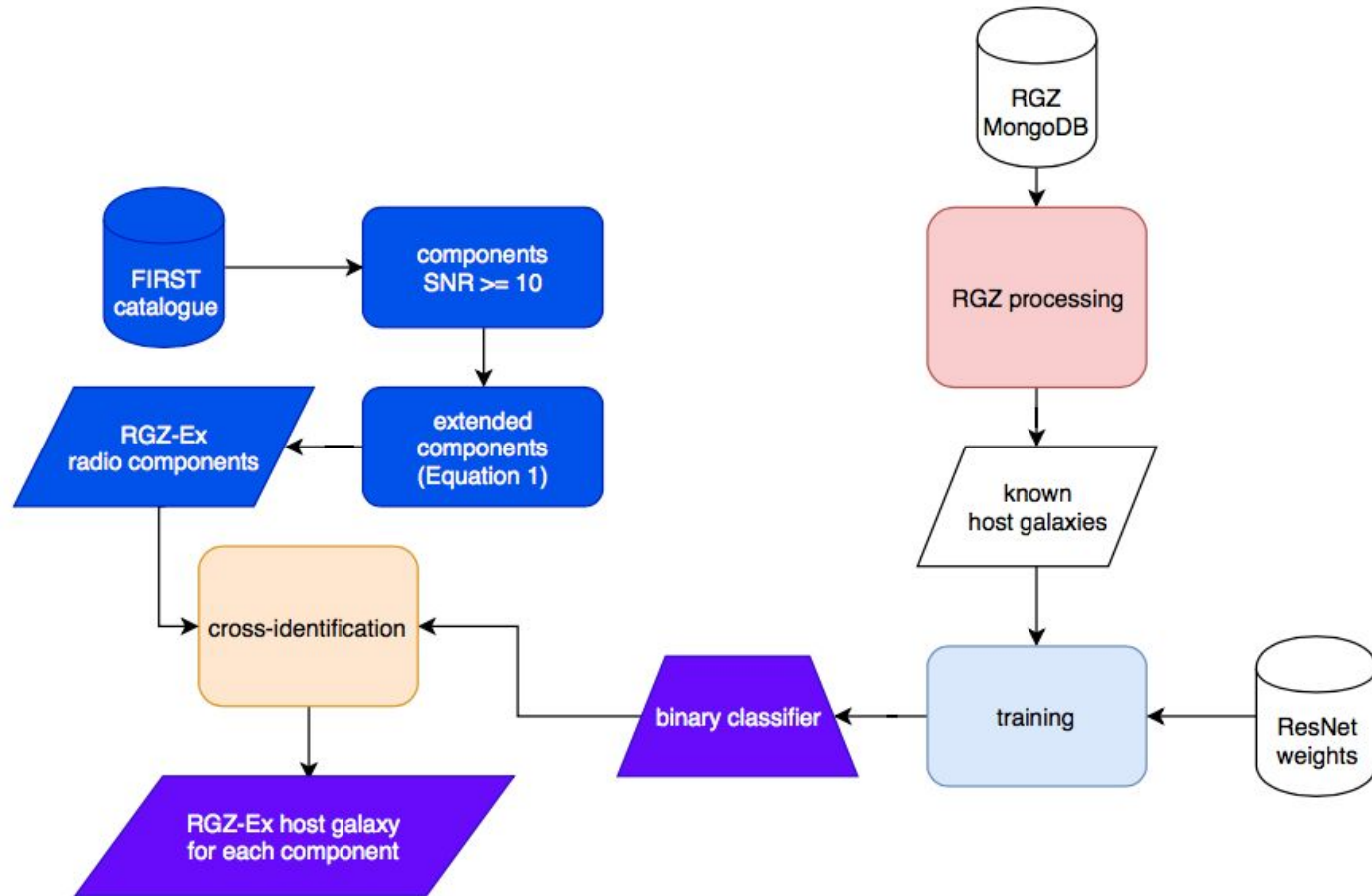
$$\begin{aligned} \text{xid} : \text{Radio} &\rightarrow \text{IR} \\ \text{xid}(r) &= \underset{i \in \text{candidate IR hosts}}{\operatorname{argmax}} f(i; r) \end{aligned}$$

$$\begin{aligned} &\text{where} \\ &f : \mathbb{R}^d \rightarrow \mathbb{R} \\ &f(i) = p(\text{host} \mid i) \\ &\text{is a binary classifier} \end{aligned}$$

Binary classification model







Luminosity function

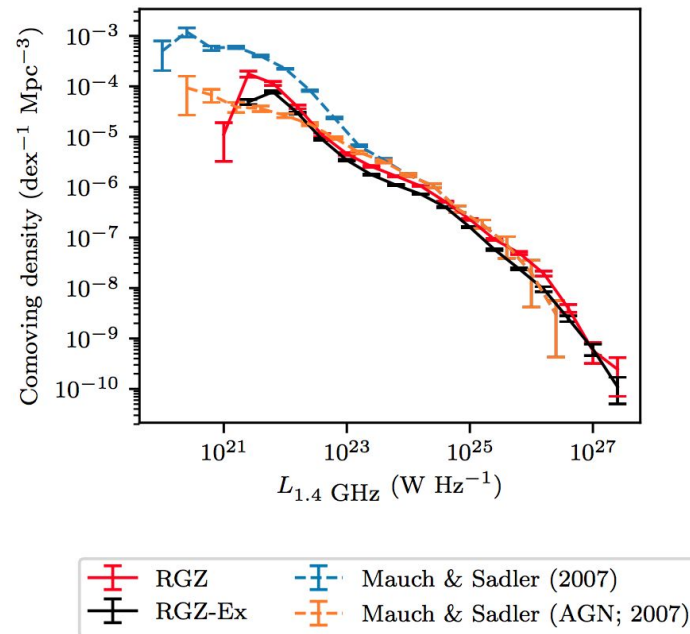
- RGZ-Ex contains 157 007 cross-identified radio sources with 30 743 redshifts
- Large sample allows us to build a radio luminosity function of extended sources
 - Luminosities up to 10^{27} W/Hz
 - Close match to Mauch and Sadler (2007) radio AGN luminosity function

Radio Galaxy Zoo: radio luminosity functions of extended sources

M. J. Alger^{1,2*} et al.

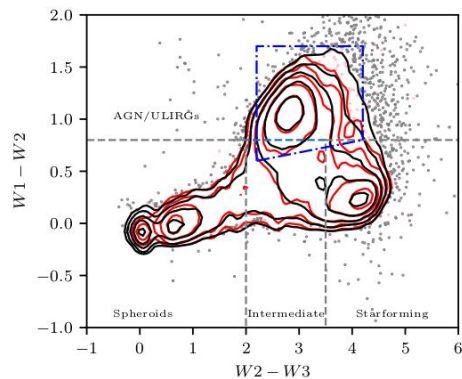
¹Research School of Astronomy and Astrophysics, The Australian National University, Canberra, ACT 2611, Australia

²Data61, CSIRO, Canberra, ACT 2601, Australia

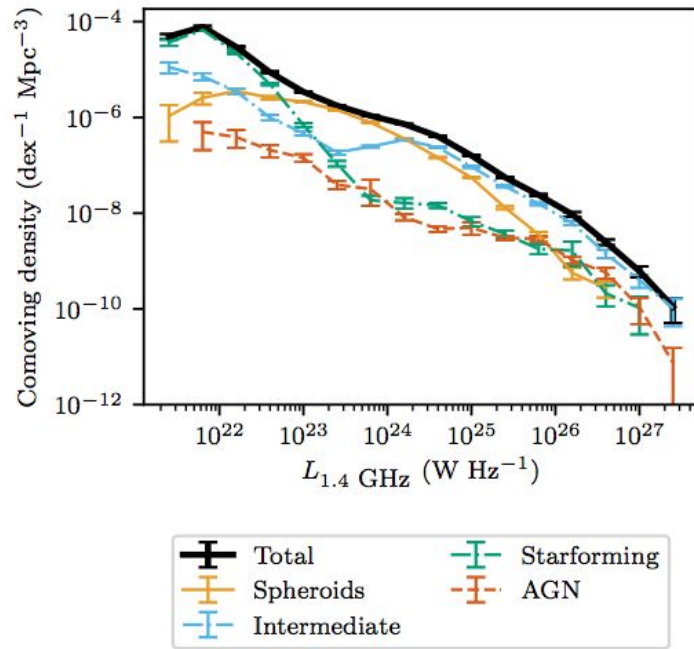


Fractional luminosity function (Mid-IR)

- Divide radio luminosity function based on mid-infrared host colours
 - “Extended” star-forming sources below 10^{23} W/Hz (visually verified)
 - Radio-loud sources dominated by “intermediate” galaxies

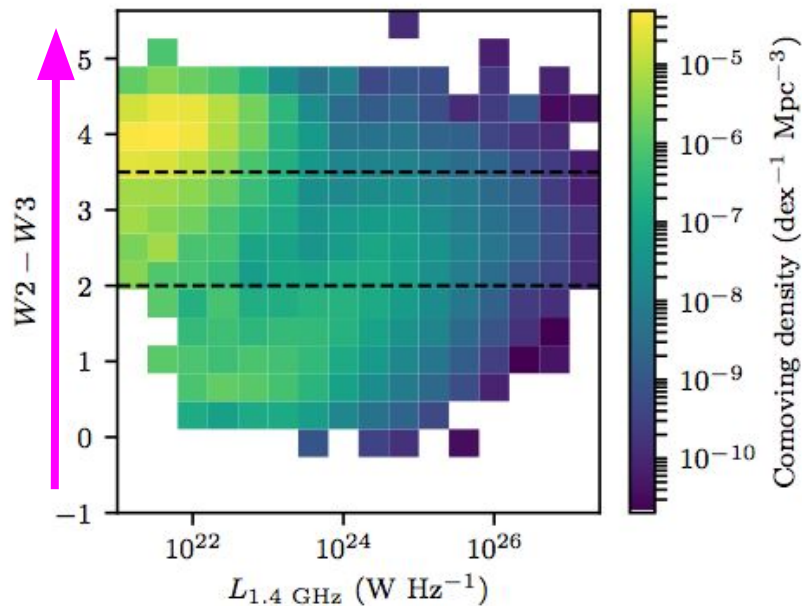
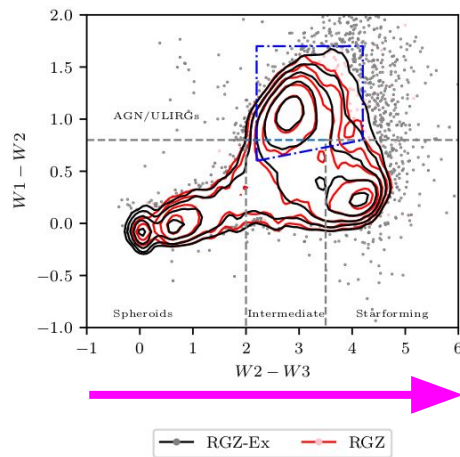


— RGZ-Ex — RGZ



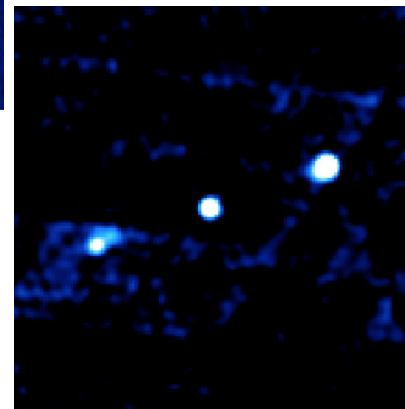
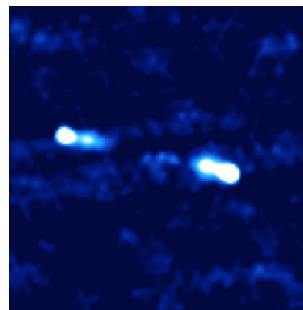
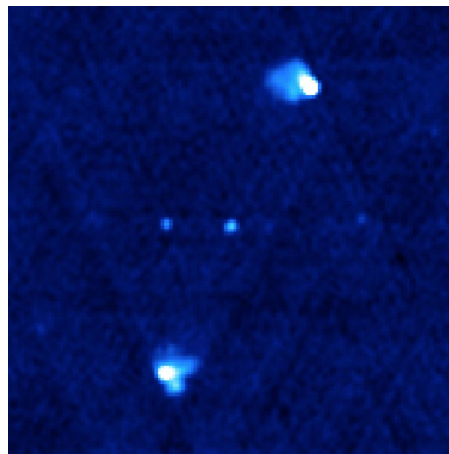
Bivariate luminosity function (Mid-IR)

- Divide radio luminosity function based on 12 μm /4.6 μm colour
 - “Extended” star-forming sources below 10^{23} W/Hz (visually verified)
 - Radio-loud sources dominated by “intermediate” galaxies



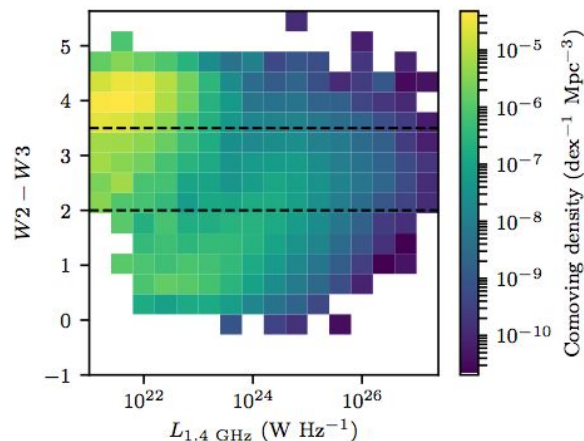
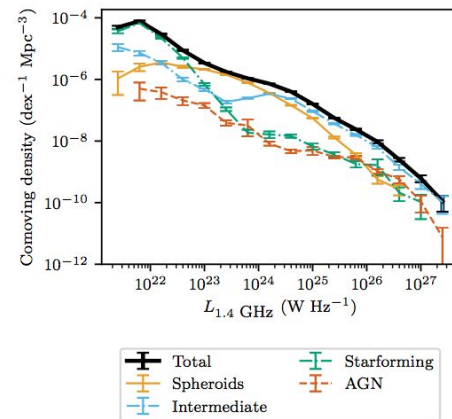
The RGZ-Ex catalogue

- Catalogue of 157 007 candidate radio sources and their hosts
- Large but noisy
- Contains around *fifty* previously unidentified giant radio galaxies (≥ 1 Mpc)



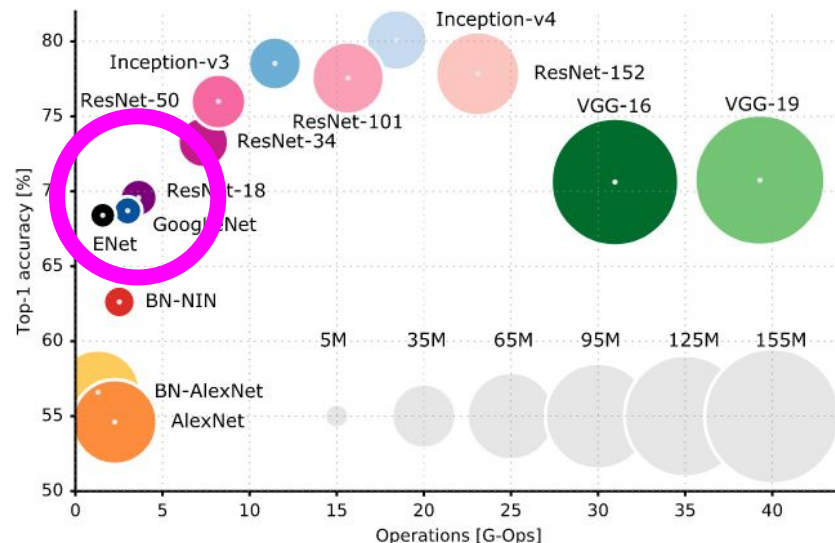
RGZ-Ex and luminosity functions

- We developed an automated, machine learning approach to radio-infrared cross-identification
- We created a huge catalogue of candidate radio sources and their hosts
- We estimated fractional radio luminosity functions of extended radio sources
- We found ~50 new giants



Binary classification model

- ResNet-18 (multiclass)
 - Good accuracy
 - Low complexity
 - Very fast to train and use
- Remove last layer and replace with a binary classifier
- Add non-image features
 - Mid-infrared colours
 - 3.4 μm flux
 - Room for improvement — e.g. add redshifts



Trade-offs between network complexity and accuracy on ImageNet.

Image: Canziani et al. (2016)