

Assignment 02, Computational Astrophysics

This assignment is in two parts. The first part involves using some available library functions for astronomical calculations. For these assignments, you need to read the given files and understand (1) what is the purpose of the given function and (2) How to use check out the given function with given test data. It is like a comprehension test; in an examination *unseen* functions would be provided.

The second part involves developing some familiarity in elementary astronomical concepts (like stellar magnitudes, distance modulus etc) in a computational setting. Common strategies involving the Matlab functions `arrayfunc`, `table`, `array2table`, `sprintf` etc would be demonstrated and applied to these problems.

In order to get started with the assignments, unzip the contents of the folder and start working on the bundled live script `assignment2.mlx`. Part of the code is already provided for you! Feel free not to adhere to them and write your own code. You can also write a little script every now and then to check out your ideas before including your code in the live-editor) You can also start by going over to the pdf version of the same assignment. Run the full script and Export to pdf after completing the assignment.

```
% Uncomment and Edit this section after completing the enire assignment and  
% before generating the output pdf.
```

```
%s=date;  
%Name='Hing Ting Chat (Roll 000)';  
%disp( ['Saved on: ' s 'by ' Name ])
```

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Part I

Conversion of Time formats

Question 1 Degrees to DMS

Convert 10°, 15'', 36''' to decimals. Use Ddd.m (Run help Ddd to know about the usage of the function Ddd. Or you can say Edit Ddd to load up Ddd.m in the matlab editor and examine its source code.)

```
% Write your code here
```

Convert Back into DMS Notation using the function.

```
% Write your code here
```

Question 2 DMS to Degrees

Now convert the values: 15.50000, -8.15278, 0.01667, -0.08334 to DMS notation using the DMS function. Put these values in a vector and then make use of Matlab's arrayfun to apply the function to all elements of the time-vector. The syntax would be something like [tD,tM,tS]=arrayfun(@DMS,tvec); Form a matrix out of the columns of t', tD', tM' and tS'. Display it in tabular format using the table2array function. The output should look like the following image. Hints: The table2array function can be called as in array2table(m,'VariableNames',CNames), where CNames is a cell array containing the column headers of the table.

ans = 4x4 table

	Time	Deg	Min	Sec
1	15.5000	15	30	0
2	-8.1528	-8	9	10.0080
3	0.0167	0	1	0.0120
4	-0.0833	0	-5	0.0240

```
clear ;  
%Write your code here
```

Question 3 Julian Day

The function Mjday returns the number of (modified) Julian days. Read the function definition and find the number of Julian days elapsed at the following points of time: Dec 12, 1971 02:31:17 PM, Feb 09, 2023, 08:30:11 AM. How many days passed between these two space time points? You may find it useful to change to format bank while working out the answer (revert back afterwards to the normal output format).

```
% Uncomment and write your code here (replace the "---" by something appropriate  
% d1=Mjday(---);  
% d2=Mjday(---);  
% disp(['d1=',num2str(d1)]); disp(['d2=',num2str(d2)]);disp(['ndays=',num2str(d2-d1)]); format
```

Part II

Common Strategies

We discuss and demonstrate a set of common strategies for the Matlab live-script interface which *also* works with Matlab scripts, i.e these strategies are not interface specific but involve the use of named and unnamed (anonymous) @ function handles and some functions like arrayfun for generalising a function originally devised for scalar inputs to accept list inputs. Some aspects of producing console output in the form of tables using table, array2table, num2str, disp and sprintf (a function that copies a formatted text string to a string variable) have also been touched upon.

According to [Matlab documentation](#) an anonymous function is a function that is *not* stored in a program file, but is associated with a variable whose data type is `function_handle`. Anonymous functions can accept multiple inputs and return one output. They can contain only a single executable statement. We have seen (in the above problems) how a function written as an external file can also be used as function handle). One straightforward example might be that of a function that computes the area of a rectangle:

```
areaRect = @(x,y) x*y;  
a=2.6; b= 9.5;  
xArea= areaRect(a,b);  
disp(['The area is ' num2str(xArea) ' m^2'])
```

It is possible to calculate the area of several rectangles by threading arrayfun over two lists containing the two sides of the rectangles.

```
aa=[1.7 2.9 3.6 8.7];  
bb=[11.2 88.7 123.4 881.8];  
xArea = arrayfun(areaRect,aa,bb);  
sxArea=strtrim( sprintf('%g ', xArea) );  
disp(['The areas are ', sxArea, ' sq m'])
```

Anonymous functions are so called as they do not really need a name, e.g the following code has the same effect although no names have been assigned to the algorithm that calculates the area.

```
disp( ['The areas are ', ...  
      strtrim( sprintf('%g ', arrayfun(@(x,y) x*y,aa,bb) ) ), ...  
      ' sq m'] )
```

Magnitudes and such

Question 4 Magnitudes

For a revision of the concept of stellar magnitudes see this [ksu](#) page (local pdf copy archived [here](#)). Make sure you understand all the connections. Now work out the following problems.

Relative Magnitude: Define an anonymous function (of the brightnesses of two objects) named `relmag` in your workspace (Make it a function of the brightnesses of the two objects) which would calculate the relative magnitude of objects whose brightness ratios vary between 1 to 500 in steps of 5.

Display only 5 representative values of the Relative Magnitudes vs their Brightness ratios including the first and the last values in a table.

```
relmag = @(x,y) -2.50*log10(y/x);
B2=1:5:100;
B1=ones(size(B2));
xMags=arrayfun(relmag,B2,B1);
xMagsDisp=xMags(1:4:end);
B2Disp=B2(1:4:end);
m=[B2Disp' xMagsDisp'];
CNames={'Brt. Ratios', 'Magnitudes'};
array2table(m, 'VariableNames', CNames)
```

Absolute Magnitude: Now Define an anonymous function (of the relative magnitude m and distance d (in pc)) named `absmag` in your workspace which would calculate the absolute magnitude of objects of relative magnitude $m = -2$ and situated between 10 to 1000 pc. This time create spaced arrays of spacing 8 and wrap the tabular output within a `disp`. (Note the difference in output)

```
% Write your code here
```

Question 5 Distance Modulus

The Distance modulus μ is a measure of distance to a celestial object. It is defined as

$$\mu = m - M$$

An object with a distance modulus of 0 is exactly 10 parsecs away. If the distance modulus is negative, the object is closer than 10 parsecs, and its apparent magnitude is brighter than its absolute magnitude. If the distance modulus is positive, the object is farther than 10 parsecs and its apparent magnitude is less bright than its absolute magnitude. Verify that $\mu(d) = 5 \log_{10} d - 5$ and find the distances d (in pc) for $\mu \in (-4, -3, \dots, 5) \cup \{10, 20\}$. Display them in a table.

Procedure: Define functions $\mu(d)$ for calculating the distance modulus as a function of distances in pc and $\lg_{10}d(x)$ for calculating the distance from the distance modulus x first. Then construct a list `muvec` for holding the distance moduli and calculate the distances in another list `dmodvec` using `arrayfun`.

```
% Write your code her
```

Suppose you were viewing the Sun from a planet orbiting another star 40 pc away. What would the Sun's apparent magnitude be? (The Sun has absolute magnitude $M = +4.8$). Ans: 7.8

Procedure: Calculate the distance modulus first. Then find the apparent magnitude from the distance modulus and absolute magnitude.

```
%Write your code here
```

References

- (1) For a revision of the concept of stellar magnitudes see this [ksu](#) page (local pdf copy archived [here](#))
- (2) Las Cumbres Observatory [page](#) on distance modulus
- (3) Wikipedia [entry](#) on distance modulus