

Wind-Mixed Layer Deepening Scaling Law in a Rotating Frame

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Key Points:

- We propose a theory for a scaling law of the deepening of the mixed layer induced by the wind forcing which considers the earth's rotation

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Abstract

The deepening of the mixed layer (ML) is a phenomenon forced by wind stress or heat flux. Scaling laws exist to predict its evolution but they either do not take into account the Earth's rotation or do not explain the turbulent mixing mechanisms involved. Here, we study the deepening of the wind-mixed layer in a linearly stratified and rotating ocean. We describe the turbulent mixing processes in an analytical framework using the $k - \epsilon$ model. We show that mixing is mainly confined to a marginally stable entrainment layer at the bottom of the ML. We propose a scaling law for the time dependence of the deepening rate of the ML $h \sim t^{1/5}$ as a function of the mixing efficiency the potential energy conversion and the dissipation profile.

Keyword: Mixed-layer depth, Mixing-efficiency, TKE, Turbulent convection, Sear-driven turbulence

Plain Language Summary

Understanding the deepening of the mixed layer in the ocean is key to predicting climate and ocean circulation. This process, driven by wind and heat flux, has an impact on the exchange of heat, nutrients and carbon between the ocean surface and deeper ocean. In our study, we investigate how wind-induced mixing affects the depth of the mixed layer in a rotating and stratified ocean. By developing a new scaling law describing the rate at which the mixed layer deepens over time, we gain insight into the complex interactions between turbulent mixing processes and ocean dynamics. This research is leading to a better understanding of the role of the ocean in regulating climate.

1 Introduction

The Mixed layer (ML) deepening is characterised by turbulent convective mixing caused by mechanical stress (momentum flux) or by a buoyancy flux. Its dynamics is also affected by the Earth's rotation. The coriolis parameter plays a stabilising role by reducing turbulent convection mixing (Jones and Marshall (1993); Sander et al. (1995); Mironov et al. (2000)), reducing convective instability (Veronis (1959); Chandrasekhar (1997); Carpenter et al. (2022)) and reducing its velocity (Fernando et al. (1991)).

Pollard et al. (1973) (hereafter P73) studied the evolution of the mixing layer (ML) and formalized a scaling law for its thickness h for a constant surface stress τ with no heat fluxes and linear stratification N . It is a separated expression for time prior and beyond half the inertial time scale $T_f = 2\pi/f$

$$h(t) = (u_*)^{1/2} \left[\frac{4(1 - \cos(ft))}{f^2 N^2} \right]^{1/4} \quad t < \frac{T_f}{2} \quad (1)$$

$$= 1.7 \left(\frac{u_*}{\sqrt{Nf}} \right) \quad t > \frac{T_f}{2} \quad (2)$$

with $u_* = \sqrt{\frac{\tau}{\rho}}$ the friction velocity at the surface and we use the notation

$$L_{P73} = \frac{u_*}{\sqrt{Nf}} \quad (3)$$

as characteristic length from the P73's scaling law.

We also define from the equation (3.1) in the paper of P73 a characteristic velocity of the ML U_{p73}

$$U_{p73} = \frac{u_*^2}{f L_{p73}} = u_* \sqrt{\frac{N}{f}}. \quad (4)$$

This scaling law is used in a large number of models (Zilitinkevich et al. (1979), Lozovatsky et al. (2005)). It is particularly robust with no rotation since at short time,

47 P73 reduced to the expression verify experimentally by Kato and Phillips (1969)

$$h(t) = u_* \left(\frac{1}{2} \right)^{1/4} \sqrt{\frac{t}{N}} \quad (5)$$

48 For longer time, P73 proposed that the ML reaches a maximum and no longer deep-
49 ens. However, it appears that there is a change in the regime after one half of inertial
50 period and the mixing layer continues to deepen (Ushijima and Yoshikawa (2020)) (here-
51 after U20).

52 U20 performed a series of LES experiments in which the friction speed u_* , rota-
53 tion f and stratification N were varied. They showed the dependency of the ML deep-
54 ening on the Rossby number and the Froude number and propose the time-dependant
55 scaling

$$h = 1.5L_{p73} \left(\frac{f}{N} \right)^{-2.2 \times 10^{-2}} \left(\frac{t}{T_f} \right)^{0.18} \quad (6)$$

56 For a range of parameter f/N between $[1.3 \times 10^{-3}$ and $2.0 \times 10^{-4}]$. This scaling is based
57 on an optimal adjustment of the parameters from the LES, However U20 does not de-
58 scribe the mechanisms that explain this scaling and their contribution.

59 In this paper we are interested in the physical process involved in the ML forma-
60 tion after the half of an inertial period. We propose to diagnose the ML deepening mech-
61 anism in presence of rotation by using idealised 1D numerical simulations. Then to build
62 a theoretical frame for a scaling law. We begin in section 2 by describing the $k-\epsilon$ tur-
63 bulent closure model. In section 3 we quantify the energy exchanges in the water col-
64 umn. A diagnostic of mixing and dissipation is given in section 4. In section 5 we for-
65 malise the scaling law and explore the limits of the proposed assumption. We conclude
66 in section 6.

67 2 Model and Methods

68 2.1 $k - \epsilon$ Model

69 Consider the x, y -momentum equation in a 1D rotating frame and the buoyancy
70 equation without heat fluxes

$$\frac{\partial \bar{u}}{\partial t} + (\bar{\mathbf{u}} \cdot \nabla) \bar{u} - f \bar{v} = \frac{1}{\rho} \frac{\partial \bar{p}}{\partial x} + \nu \nabla^2 \bar{u} - \frac{\partial}{\partial z} \overline{u'w'} \quad (7)$$

$$\frac{\partial \bar{v}}{\partial t} + (\bar{\mathbf{u}} \cdot \nabla) \bar{v} - f \bar{u} = \frac{1}{\rho} \frac{\partial \bar{p}}{\partial y} + \nu \nabla^2 \bar{v} - \frac{\partial}{\partial z} \overline{v'w'} \quad (8)$$

$$\frac{\partial \bar{b}}{\partial t} + (\bar{\mathbf{u}} \cdot \nabla) \bar{b} = \frac{\partial}{\partial z} \left(\nu' \frac{\partial \bar{b}}{\partial z} - \overline{w'b'} \right) \quad (9)$$

71 The $k - \epsilon$ model is a turbulent closure of order 2. It closes the Reynolds's aver-
72 age equation by adding two prognostic transport equations, one for the turbulent kinetic
73 energy $k = u'^2 + v'^2$ and the other for the dissipation ϵ .
74 Neglecting advecting terms.

$$\frac{\partial k}{\partial t} - \frac{\partial}{\partial z} \left(\left(\nu + \frac{\nu_t}{\sigma_k} \right) \frac{\partial k}{\partial z} \right) = P - B - \epsilon \quad (10)$$

$$\frac{\partial \epsilon}{\partial t} - \frac{\partial}{\partial z} \left(\left(\nu + \frac{\nu_t}{\sigma_\epsilon} \right) \frac{\partial \epsilon}{\partial z} \right) = \frac{\epsilon}{k} (c_{1\epsilon} P - c_{3\epsilon} B - c_{2\epsilon} \epsilon) \quad (11)$$

Where σ_k the turbulent Schmidt number, $c_{1\epsilon}, c_{2\epsilon}, c_{3\epsilon}$ empirical parameters (Canuto et al. (2001)).
The turbulent fluxes are computed as follows (Rodi (1987)Burchard (2002b))

$$P = -\overline{u'w'}\partial_z\bar{u} = \nu_t(\partial_z\bar{u})^2 \quad (12)$$

$$B = \overline{w'b'} = -\nu'_t(\partial_z\bar{b}) \quad (13)$$

with $b = -g\frac{\Delta\rho}{\rho}$ and the diffusivity ν_t and ν'_t respectively of momentum and heat are parameterised using stability functions c_μ and c'_μ (Canuto et al. (2001); Burchard and Bolding (2001)) such that :

$$\nu_t = c_\mu \frac{k^2}{\epsilon} \quad \nu'_t = c'_\mu \frac{k^2}{\epsilon} \quad (14)$$

In this model, the buoyancy production is directly linked to the potential energy (Burchard, 2002a)

$$\frac{\partial E_p}{\partial t} = \int_{-H}^0 \nu'_t \frac{\partial b}{\partial z} dz = \int_{-H}^0 -B dz. \quad (15)$$

Here the capital E_p refer to the integrated potential energy $e_p = -zb$.

This model has been shown to be robust in convective processes (Burchard and Bolding (2001); Umlauf et al. (2003)) since it correctly reproduces the basic interaction mechanisms between shear, stratification and turbulence. A review of turbulence closure models, and in particular the $k - \epsilon$ model, is presented in (Burchard, 2002a).

2.2 Numerical Model

We use the open-source numerical model GOTM (General Ocean Turbulence Model) <https://gotm.net/>. This Model solves the momentum, salt and heat transport equations over a 1D water column. All horizontal gradients are either parameterised, neglected, or obtained prognostically where possible. Our configuration solves equation (11) for a Coriolis parameter $f \in [5.910^{-5}; 1.0210^{-4}] \text{ s}^{-1}$, a stratification $N^2 \in [2.510^{-5}; 210^{-3}] \text{ s}^{-1}$ (same range of parameter f/N as U20). A velocity friction $u_* = 1.10^{-2} \text{ ms}^{-1}$, a simulation span of 10 days for a time step of 60 s. The vertical resolution is 50 cm for a domain of 500 m in the z -axis.

Figure 6 show the evolution of h in time using this model, we observe an evolution law $h(t) \sim t^{1/5}$ which is close the power found by U20 $h_{U20}(t) \sim t^{0.18}$. This numerical model is in good agreement with U20. Therefore GOTM and thus the $k-\epsilon$ model are adapted to the modelling of the turbulent convection. and represent this process with a reliability which is near of the one offered by LES.

3 Energetic of the Mixed Layer Deepening

We examine the analytically the evolution of the potential energy associated with the depth of the mixing layer. We analyse how this potential energy is affected by changes in the density profile, and identify a structure for the water column based on vertical density and velocity profiles.

3.1 Potential Energy Evolution

The initial density profile is given by

$$\rho_i(z) = \rho_0(1 - \frac{N^2}{g}z) \quad (16)$$

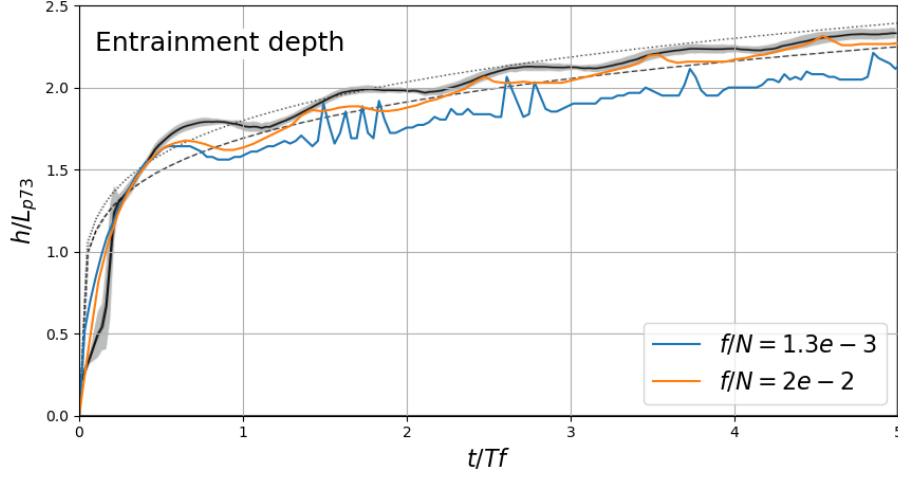


Figure 1. Temporal variation of the MLD normalized by L_{p73} the color curves are plotted on the Fig 3a) of U20. The solid dark line is the averaged value of all the U20 experiment. The Dashed and dotted lines are the estimation of the (Eq 6) with the greatest ($f/N = 2.0 \times 10^{-2}$) and smallest ($f/N = 1.3 \times 10^{-3}$) f/N parameter range of the U20's study. The color lines are GOTM simulation for the two different f/N parameter.

109 The density of the ML is

$$\rho_m = \frac{\rho_i(z=0) + \rho_i(z=-h)}{2} = \rho_0 \left(1 - \frac{N^2}{2g} h\right). \quad (17)$$

110 As h deepens, the density ρ_m of the mixing layer increases because it has been mixed
 111 with the denser water below. Therefore, over time, the centre of mass is shifted upward.
 112 Considering a vertical density profile of a water column composed by a mixed layer and
 113 a stratified one, we can write the expression of the specific potential energy.

$$E_p = \frac{1}{\rho_0} \int_{-H}^0 \rho g z dz \quad (18)$$

114 We plug in the density profile of the ML (eq. 17) and the lower part at rest (eq. 16)

$$E_p = \left(-\frac{gH^2}{2} - \frac{N^2 H^3}{3} + \frac{1}{12} N^2 h^3 \right) \quad (19)$$

115 The first two terms correspond to the initial potential energy and the last term rep-
 116 resent the gain due to the lift of the centre of mass.
 117 Taking the derivative of this equation gives the time evolution of the potential energy,
 118 which is positive. Thus, a deepening of the ML correspond to gain of potential energy.

$$\frac{dE_p}{dt} = \frac{N^2}{12} \frac{dh^3}{dt} \quad (20)$$

119 The Fig (2) show the structure of the water column. From the top to the bottom,
 120 three layers can be defined.

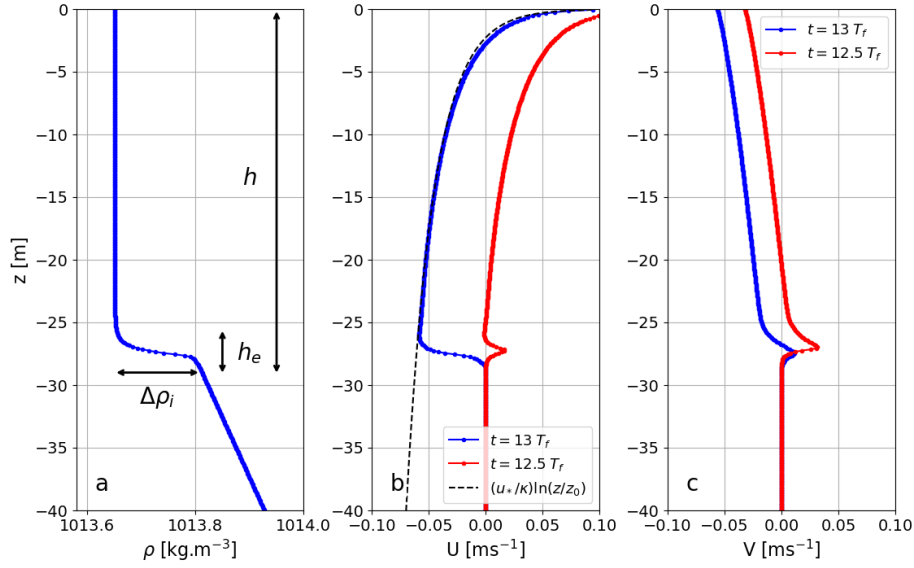


Figure 2. Vertical profile of (a) density, (b) along-wind velocity U and (c) cross-wind velocity V . In blue, the last inertial period $t = 13 T_f$. In red the last half inertial period $t = 12.5 T_f$. In black the logarithmic profile of the velocity from the "Law of the wall" $\frac{u_*}{\kappa} \ln(\frac{z}{z_0})$. The parameters of the simulation are : $f = 10^{-4} \text{ s}^{-1}$, $N^2 = 10^{-4} \text{ s}^{-1}$ and the wind stress $\tau = 0.1027 \text{ m}^2 \text{ s}^{-2}$

- A well mixed layer $z \in [0, -h + h_e]$ in which the density is uniform. The along-wind velocity follow a logarithmic profile

$$\frac{\partial \bar{u}}{\partial z} = \frac{u_*}{\kappa z} \quad (21)$$

with κ the Von Karman constant. This expression come from the *law of the wall* due to the frictional behaviour of the wind stress. This velocity profile is expected due to the $k - \epsilon$ model implementation.

- An entrainment layer $z \in [-h + h_e, -h]$. This layer forms the transition between the other two layers. It contains on the thickness h_e , the density jump $\Delta\rho = N^2 h$ and the velocity jump ΔU .

- A lower part with $z \in [-h, -H]$. This layer is not affected by the surface shear and is therefore at rest with the background stratification

3.2 Local Turbulent Kinetic Energy Balance

The input of energy at the top of the water column is converted into turbulent Shear, which is transported in the water column by turbulent processes. The evolution of the MLD can be retrieved by the buoyancy production. We thus need to consider the turbulent kinetic energy (k) equation (11).

At first, as the near sub surface velocity increase, accelerating the mixed layer. After half of the inertial period we see in Fig 3 (Top) that the velocity drops due to the onset of an inertial oscillation. As a result the layer reaches a periodic equilibrium. However there is still an injection of energy at the surface and the mixing layer continues to deepen as the fig ?? shows but does not accelerate. As a result we observe TKE reaching steady as shown Fig (3) (Bottom)

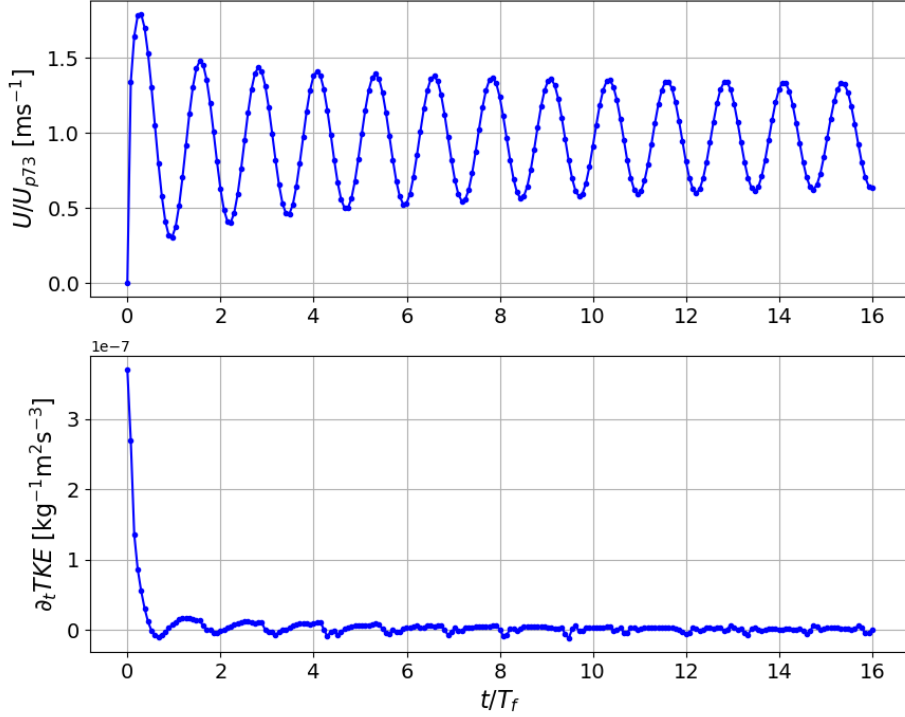


Figure 3. Time evolution of the time derivative of the surface along-wind velocity (Top) and TKE integrated over the water column (Bottom). $f = 10^{-4} \text{ s}^{-1}$, $N^2 = 10^{-4} \text{ s}^{-1}$ and the wind stress $\tau = 0.1027 \text{ m}^2 \text{ s}^{-2}$

Thus there is a balance between the terms on the right hand sides.
Hence

$$P - B = \epsilon. \quad (22)$$

We can use the expression of the Richardson flux $Ri_f = \frac{-B}{P}$ in the equation (22)

$$1 + \frac{P}{B} = -\frac{\epsilon}{B} \quad (23)$$

$$-B = \frac{Ri_f}{1 + Ri_f} \epsilon \quad (24)$$

$$-B = \Gamma \epsilon \quad (25)$$

With the mixing efficiency $\Gamma = \frac{Ri_f}{1 + Ri_f}$ defined empirically ((Osborn, 1980)) with a value of $\Gamma_{t0} = 0.2$. Fig 4 represent its vertical profile. From it we can approximate the vertical profile of Gamma as being a piecewise constant function null everywhere except in the entrainment layer and with a value of $\Gamma_0 = 0.25$. Note that below the entrainment layer, Gamma tends towards 1 for reasons of numerical stability (B takes a threshold value and P is null). The mixing efficiency is zero in this region of the water column. From now we will use the approximated function instead of the real profile. We call it Γ for more simplicity.

This equation is a relationship between the production of buoyancy, dissipation and mixing efficiency.

Although we consider a ML, nothing suggests that the mixing process is uniform in this layer. The subject of the next section is, therefore, the nature of the mixing and dissipation and their vertical distribution.

4 Mixing and Dissipation law

4.1 Interface Mixed/Stratified layer

We have seen above with the density and velocity profiles that the water column has a vertical structure which consists of three distinct layers. Let's perform a heuristic diagnostic to measure the thickness of the entrainment layer by describing the vertical structure of the mixture.

In the upper layer, the velocity has a logarithmic profile and, since the layer is already mixed and therefore uniform (fig 2 a) , we can assume there is very little mixing.

In contrast, the entrainment layer contains a very high density jump and a very high velocity shear on the thickness h_e . It is therefore in this layer that the mixing is greatest. This mixing is caused by a velocity shear which is all the greater the thinner its thickness. This layer is therefore marginally stable.

Then the bottom layer is not affected by the surface shear and is therefore at rest with the background stratification and so zero mixing.

$$Ri = \frac{g \frac{\partial \rho}{\partial z}}{(\frac{\partial U}{\partial z})^2} = Ri_c \quad (26)$$

with the density leap: $\frac{\partial \rho}{\partial z} = \frac{\Delta \rho_i}{h_e} = \frac{N^2 h}{h_e}$

We observe in Fig (3 (b)), that the velocity in the ML follow a logarithmic profile

$$\frac{\partial \bar{u}}{\partial z} = \frac{u_*}{\kappa z} \quad (27)$$

with κ the Von Karman constant. This expression come from the *law of the wall* due to the frictional behaviour of the wind stress. We also see that the velocity profile is oscillating between a peak and a trough of the inertial period. As a result, the shear at the bottom of the mixing layer varies within an inertial period. The first consequence of this is that we need to estimate an average value for this shear between $-h + h_e$ and $-h$.

We use the characteristic velocity U_{p73} (4). Furthermore, in order to obtain a power law, we need to discard a logarithmic correction. We therefore obtain an expression for the velocity shear between the mixed layer and the stratified layer at rest

$$\frac{\partial U}{\partial z} = \frac{u(-h + h_e) - 0}{h_e} = \frac{u_* \sqrt{\frac{N}{f}}}{h_e}. \quad (28)$$

The second consequence is that the marginal stability condition is also periodic since it is only valid for a sufficiently large shear. The precise evaluation of the Richardson number under these conditions is complex, so we will adopt a bulk formulation in the entrainment layer with $Ri_c = O(1)$. This formulation is commonly used in ML deepening models (Pollard et al. (1973); Price et al. (1986)).

Hence, we obtain a relation for the entrainment layer thickness

$$h_e = Ri_c \frac{u_*^2}{N f} \frac{1}{h} \quad (29)$$

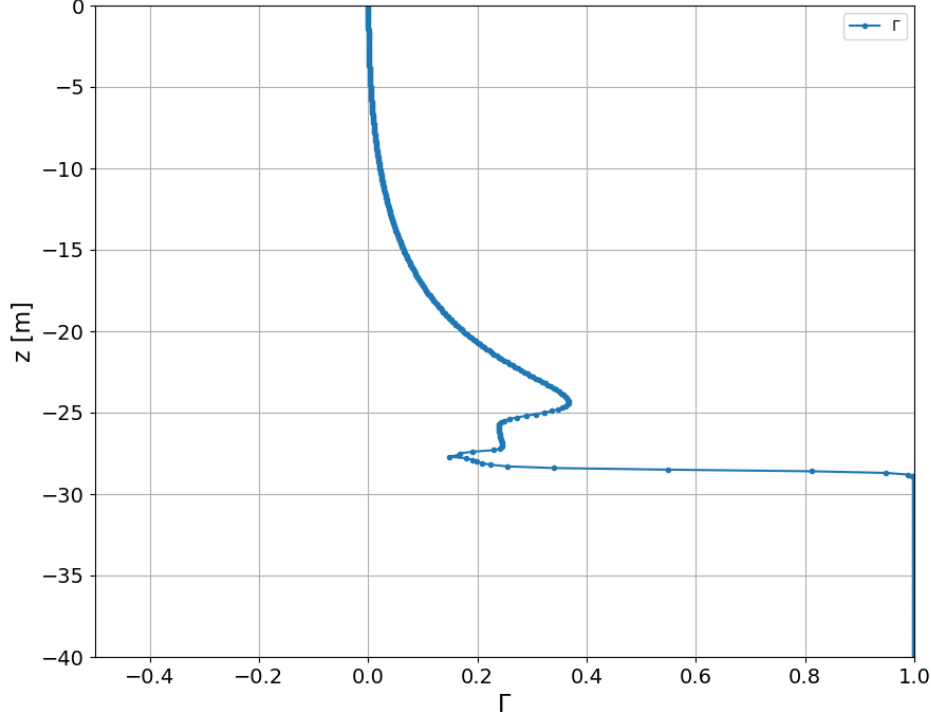


Figure 4. Vertical profile of the mixing efficiency at the last moment of the simulation. The bold line is the profile obtained by the simulations. Implicitly, the contour of the function delimits the three parts: from top to bottom: the Mixed layer, the entrainment layer and the bottom layer. $f = 10^{-4} \text{ s}^{-1}$, $N^2 = 10^{-4} \text{ s}^{-1}$ and the wind stress $\tau = 0.1027 \text{ m}^2 \text{ s}^{-2}$. Below the mixed layer the shear stress is null and thus the Γ tend to 1.

4.2 Mixing efficiency

The integral of the combination of the eq (15) and (25) can be decomposed on the three intervals describe above. Due to the zero value of the mixing efficiency in the mixing layer and the deep layer, we can therefore write

$$\frac{\partial E_p}{\partial t} = \int_{-h}^{-h+h_e} \Gamma \epsilon dz \simeq \Gamma_0 h_e \epsilon(z = -h) \quad (30)$$

Thus

$$\frac{\partial E_{pot}}{\partial t} = Ri_c \Gamma_0 \frac{u_*^2}{Nf} \frac{1}{h} \epsilon(z = -h). \quad (31)$$

4.3 Dissipation Vertical Profile

To determine the nature of the vertical profile of the ϵ dissipation we consider the TKE balance equation. At the surface we have $B = 0$. And as our system is quasi-stationary we have that $\frac{\partial k}{\partial t} = 0$. Hence, near the surface

$$\epsilon = -P.$$

198 The vertical profile of the dissipation is assumed by interpolating its surface expression
 199 throughout the mixed layer. Using the logarithmic profile of the velocity (Fig 27) and
 200 using the constant flux approximation $u'w' = u_*^2$

$$\epsilon(z) = \frac{u_*^3}{\kappa z} \quad (32)$$

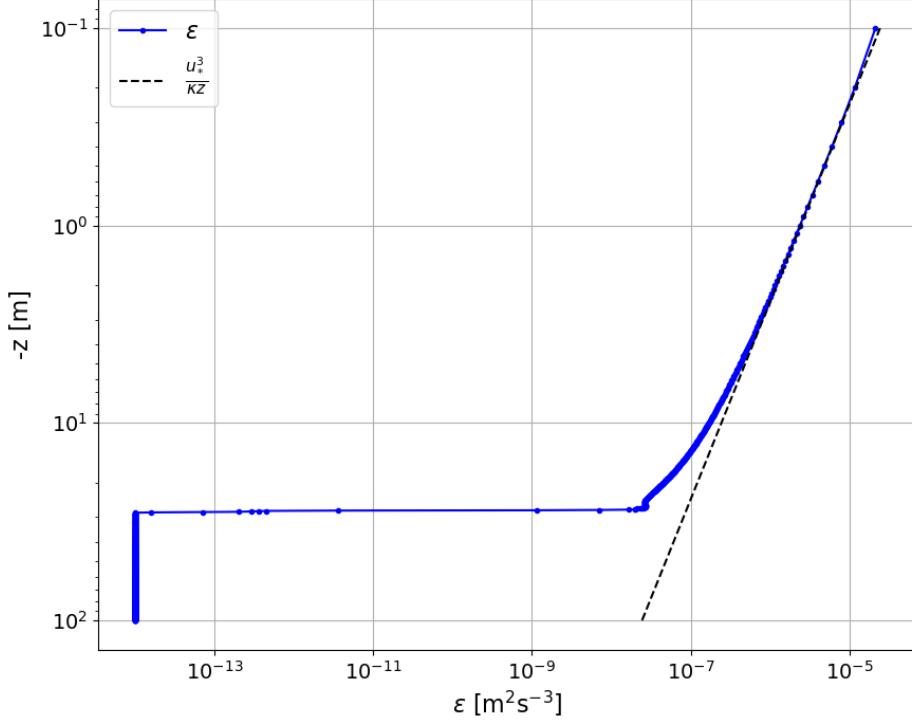


Figure 5. Vertical profile of the dissipation ϵ (blue solid line) the last instant of the simulation. And a law $1/z$ (dashed line). $T. f = 10^{-4} \text{ s}^{-1}$, $N^2 = 10^{-4} \text{ s}^{-1}$ and the wind stress $\tau = 0.1027 \text{ m}^2 \text{ s}^{-2}$

201 The approximation we made by interpolating the surface dissipation expression over
 202 the entire ML is consistent, since we observe in Fig (5) the same order of magnitude at
 203 the top of the entrainment layer, (above the leap).

204 Using the equations (20),(31)

$$\frac{N^2}{12} \frac{\partial h^3}{\partial t} = Ri_c \Gamma_0 \frac{u_*^2}{Nf} \frac{1}{h} \epsilon(z = -h) \quad (33)$$

205 and by replacing the expressions of the dissipation at the depth h using 32

$$\frac{\partial h^5}{\partial t} = \left(20 Ri_c \Gamma_0 \frac{u_*^5}{N^3 f \kappa} \right) \quad (34)$$

206 and by time integration

$$h(t) = \left(20 Ri_c \Gamma_0 \frac{u_*^5}{N^3 f \kappa} \right)^{1/5} t^{1/5}. \quad (35)$$

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We write this equation using the same formulation as Ushijima and Yoshikawa (2020)

$$h(t) = L_p \tau^3 \left(40\pi Ri_c \frac{\Gamma_0}{\kappa} \right)^{1/5} \left(\frac{f}{N} \right)^{1/10} \left(\frac{t}{T_f} \right)^{1/5} \quad (36)$$

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We therefore obtain a law that predicts deepening at a rate $t^{1/5}$. This rate is consistent with the one computed by U20. We also note that the coefficient $\frac{f}{N}$ is quite different from the value of U20 but remain very small and the the sensibility of the scaling law within the range of the $(\frac{f}{N})$ values set by U20 ($\frac{f}{N} = 1.3 \times 10^{-3}$ for the smallest and $\frac{f}{N} = 2.0 \times 10^{-4}$ for the greatest) is negligible.

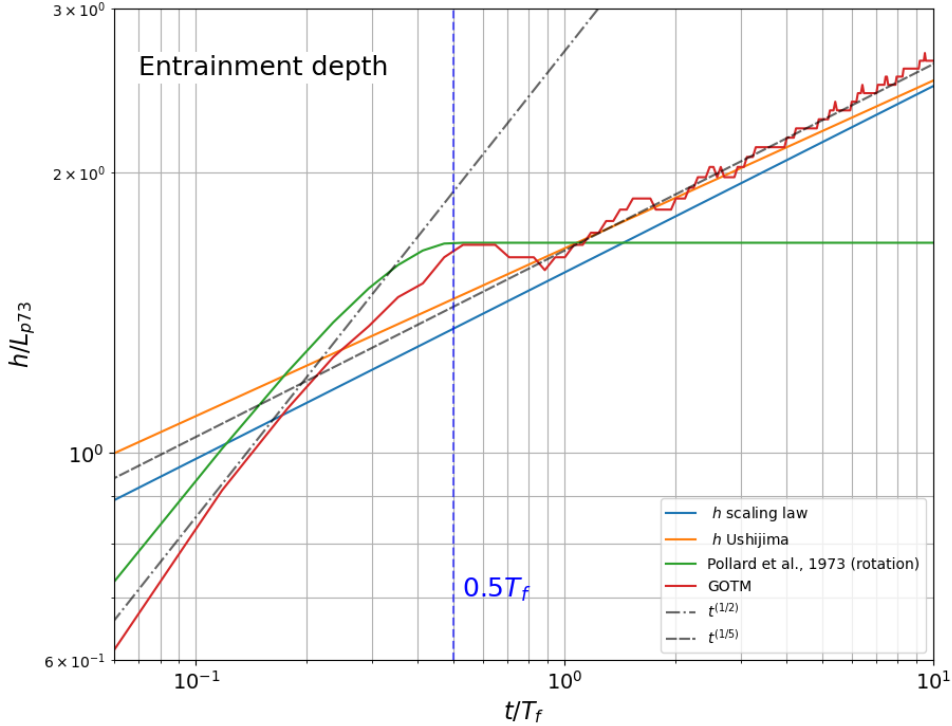


Figure 6. Time evolution of the mixed layer depth (MLD) in solid lines simulated by the numerical model (red), the scaling law predicted by Pollard et al, (green), the scaling law found by U20 (in orange) and the analytical scaling law demonstrated in this paper (in light blue) . Dark dashed line are the temporal evolution power 1/2 and 1/5. The vertical dashed blue line indicates half of the inertial period $0.5T_f$. The Coriolis parameter for the analytical law and the simulation is $f = 10^{-4} \text{ s}^{-1}$, $N^2 = 10^{-4} \text{ s}^{-1}$ and the wind stress $\tau = 0.1027 \text{ m}^2 \text{ s}^{-2}$

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We can compare in Fig 6 the value of our scaling to the value obtain with the scaling of U20 and the numerical simulation. We observe our scaling underestimate the ML deepening. However, because the time dependency is greater in our formulation than the U20 one. The relative error of our MLD estimate to U20 $\frac{h_{U20} - h}{h_{U20}}$ reduce from 4.5 % to 2 % at T_f to 1 % to $12 T_f$. We can therefore conclude to the constancy of our scaling law and show the validity of our theoretical framework.

219 4.4 Limitations

220 We have considered each of these to be valid as long as they respect the same or-
 221 der of magnitude as the simulated values. We summarise them here:

- 222 - We wanted an analytical solution in the form of a power law of the deepening rate of
 223 the ML. Thus we had to consider the stratification to be linear and apply logarithmic
 224 corrections to the velocity at the base of the entrainment layer.
- 225 - We approximated the mixing efficiency profile to be non-null only in the entrainment
 226 layer. In reality, mixing is still occurring in the upper part of the ML since its density
 227 increases as the layer thickens, as shown by its expression eq (17). In the deep layer be-
 228 low the ML, internal waves are generated by the deepening of the ML, creating some mix-
 229 ing.
- 230 - We interpolated the dissipation profile from a valid surface expression.
- 231 - We have used the expression of a critical Richardson number to define the thickness
 232 of the entrainment layer. This approach is more refined than a Bulk Richardson num-
 233 ber over the entire ML (Pollard et al. (1973); Price et al. (1986)). Yet, due to inertial
 234 oscillations, this layer is in a marginal stability state only when the shear at the base of
 235 the ML is maximal. The GOTM model is a faithful representation of the turbulent con-
 236 vection phenomenon, but it appears that for small f/N , the ML deepening is underval-
 237 ued. Hence the dependence of the f/N factor is much greater in our scaling $\left(\frac{f}{N}\right)^{1/10}$
 238 than in that of U20 $\left(\frac{f}{N}\right)^{-2.2 \times 10^{-2}}$.

239 5 Concluding Remarks

240 In this study we have developed a theoretical framework to explain the deepening
 241 of the ML in the case of an ocean forced by surface stress without heat loss in a rotat-
 242 ing frame. The deepening rate was found to be function of the Coriolis parameter and
 243 the stratification with the time dependency $h(t) \sim t^{1/5}$
 244 We have showed that the turbulent kinetic energy reach a steady state after $0.5 T_f$. Thus
 245 the constant input of energy by wind stress on the surface is converted to buoyancy pro-
 246 duction in a layer at the interface between the well-mixed layer above and the stratified
 247 one at rest below. This interface is defined by its marginal stability and concentrate most
 248 of the mixing. Therefore we shown the vertical inhomogeneity of the mixing efficiency
 249 within the mixed layer.

250 All the processes cited here have been studied in the context of wind-driven con-
 251 vection without heat flux. The extension of this work would be to study the case of free
 252 convection subject to rotation, in particular to investigate the creation of inertial waves
 253 in the absence of mean current and wind forcing . Current explanations are based either
 254 on the Taylor Proudman theorem or on the interaction of coherent vortex structures. There
 255 remains much to explore on this question.

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 259 The authors declare that they have no conflict of interest.

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