

# Quantum Mechanics II

*lecture notes*

Mr. Adrian Signer

**Notes**

Marc Maetz

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# INTRODUCTION

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These are my lecture notes of the lecture. You are welcome to tell any mistakes to: `mmaetz AT student.ethz.ch`. Unfortunately, some lectures are missing because lenovo/IBM is incompetent to give me a working laptop (thinkpad) after two months (even after about 30 e-mails and 10 phone calls).

**Missing lectures**     • 9th April

- 23th Mai



# APPROXIMATION METHODS FOR STATIONARY PROBLEMS

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- standard QM problem:

- given  $|\psi(t_0)\rangle$
- wanted

$$|\psi(t)\rangle = U(t, t_0) |\psi(t_0)\rangle$$

$$U = e^{-\frac{i}{\hbar} H(t-t_0)}$$

- for time independent  $H$

$$U = e^{-\frac{i}{\hbar} H(t-t_0)} \tag{1.1}$$

find eigenvalue and eigenstates (diagonalize  $H$ )

but: most problems cannot be solved exactly  $\rightarrow$  find approximate solution.

---

## 1.1 Time-independent perturbation theory, non-degenerate case

Assume:

$$H = H_0 + \lambda V \tag{1.2}$$

with  $H_0$  the Hamiltonian that I can solve (“free” Hamiltonian) and the perturbation  $V$  “small” and  $\lambda$  a dimensionless bookkeeping part.

$$\lambda \rightarrow 0, \quad H \rightarrow H_0$$

$$\lambda \rightarrow 1, \quad \text{full } H$$

We know

$$\left| \psi_n^{(0)} \right\rangle, E_n^{(0)}$$

with

$$H_0 \left| \psi_n^{(0)} \right\rangle = E_n^{(0)} \left| \psi_n^{(0)} \right\rangle$$

with

$$\left\langle \psi_n^{(0)} \right| \psi_m^{(0)} \rangle = \delta_{mn}$$

(continuous spectrum also understood.)

We want  $|\psi_n\rangle$  and  $E_n$  with

$$(H_0 + \lambda V) |\psi_n\rangle = E_n |\psi_n\rangle$$

let

$$\begin{aligned} E_n &= E_n^{(0)} + \lambda E_n^{(1)} + \lambda^2 E_n^{(2)} + \dots \\ |\psi_n\rangle &= \left| \psi_n^{(0)} \right\rangle + \lambda \left| \psi_n^{(1)} \right\rangle + \lambda^2 \left| \psi_n^{(2)} \right\rangle + \dots \end{aligned}$$

seems obvious, but assumption. (convergence?)

$$\begin{aligned} &\left( H_0 - E_n^{(0)} \right) \left| \psi_n^{(0)} \right\rangle + \lambda \left( \left( H_0 - E_n^{(0)} \right) \left| \psi_n^{(1)} \right\rangle \right) \\ &+ \lambda^2 \left( \left( H_0 - E_n^{(0)} \right) \left| \psi_n^{(2)} \right\rangle + \left( V - E_n^{(1)} \right) \left| \psi_n^{(1)} \right\rangle - E_n^{(2)} \right) + \mathcal{O}(\lambda^3) = 0 \end{aligned}$$

with  $\mathcal{O}(1)$  “step 0”,  $\mathcal{O}(\lambda)$  “step 1”,  $\mathcal{O}(\lambda^2)$  “step 2”.

**Step 0:** nothing to do

**Step 1** multiply by  $\left\langle \psi_m^{(0)} \right|$

$$\begin{aligned} &\left\langle \psi_m^{(0)} \right| H_0 - E_n^{(0)} \left| \psi_m^{(0)} \right\rangle + \left\langle \psi_m^{(0)} \right| V - E_n^{(1)} \left| \psi_m^{(0)} \right\rangle = 0 \\ &= \left( E_m^{(0)} - E_n^{(0)} \right) \left\langle \psi_m^{(0)} \right| \psi_n^{(1)} \rangle + \left\langle \psi_m^{(0)} \right| V \left| \psi_m^{(0)} \right\rangle - E_n^{(1)} \delta_{mn} = 0 \end{aligned}$$

to get  $\left| \psi_n^{(1)} \right\rangle$

$$\begin{aligned} \left| \psi_n^{(1)} \right\rangle &= \sum_m \underbrace{\left\langle \psi_m^{(0)} \right| \psi_n^{(1)} \rangle}_{\left\langle \psi_m^{(0)} \right| V \left| \psi_m^{(0)} \right\rangle - E_n^{(1)} \delta_{mn}} \left| \psi_m^{(0)} \right\rangle \\ &= \sum_m \frac{\left\langle \psi_m^{(0)} \right| V \left| \psi_m^{(0)} \right\rangle}{E_n^{(0)} - E_m^{(0)}} \left| \psi_m^{(0)} \right\rangle + \left| \psi_n^{(0)} \right\rangle \left\langle \psi_n^{(0)} \right| \psi_n^{(1)} \rangle \end{aligned}$$



from normalization

$$\langle \psi_n | \psi_n \rangle \stackrel{!}{=} 1 = \underbrace{\langle \psi_n^{(0)} | \psi_n^{(0)} \rangle}_1 + \lambda \underbrace{\langle \psi_n^{(0)} | \psi_n^{(1)} \rangle}_0 + \lambda \langle \psi_n^{(1)} | \psi_n^{(0)} \rangle + \mathcal{O}(\lambda^2) \quad (1.3)$$

has to be small. If  $E_n^{(0)} = E_m^{(0)}$ ?? degeneracy!  $\rightarrow$  sec 1.2 if  $E_n^{(0)} \simeq E_m^{(0)}$  quasi degenerate

**step 2** take  $\mathcal{O}(\lambda^2)$  terms  $\langle \psi_k^{(0)} |$

$$(E_k^{(0)} - E_n^{(0)}) \langle \psi_k^{(0)} | \psi_n^{(2)} \rangle + \langle \psi_k^{(0)} | V | \psi_k^{(0)} \rangle - E_n^{(1)} \langle \psi_k^{(0)} | \psi_n^{(1)} \rangle = E_n^{(2)} \delta_{kn} \quad (1.4)$$

for  $k = n$

$$\begin{aligned} E_n^{(2)} &= \langle \psi_n^{(0)} | V | \psi_n^{(0)} \rangle \\ &= \sum_{m \neq n} \frac{\langle \psi_n^{(0)} | V | \psi_n^{(0)} \rangle \langle \psi_m^{(0)} | V | \psi_m^{(0)} \rangle}{E_n^{(0)} - E_m^{(0)}} \\ &= \sum_{m \neq n} \frac{\|V_{nm}^2\|}{E_n^{(0)} - E_m^{(0)}} \end{aligned} \quad (1.5)$$

Note  $E_n^{(2)} < 0$  for ground state.

Next compute  $|\psi_n^{(2)}\rangle$ : initially fix normalization such that

$$\langle \psi_n^{(0)} | \psi_n^{(i)} \rangle = \delta_{i0} \quad (1.6)$$

this is in conflict

$$\langle \psi_n | \psi_n \rangle = 1 \quad (1.7)$$

$\rightarrow$  sort out at the end.

$$\begin{aligned} |\psi_n^{(2)}\rangle &= \sum_{k \neq n} |\psi_k^{(0)}\rangle \langle \psi_k^{(0)} | \psi_n^{(2)} \rangle + 0 \\ &= \sum_{k \neq n} |\psi_k^{(0)}\rangle \left( \frac{\langle \psi_k^{(0)} | V | \psi_k^{(0)} \rangle}{E_n^{(0)} - E_k^{(0)}} - \frac{E_n^{(1)} \langle \psi_k^{(0)} | \psi_n^{(1)} \rangle}{E_n^{(0)} - E_k^{(0)}} \right) \end{aligned} \quad (1.8)$$

plug in  $|\psi_n^{(1)}\rangle$  and sort out normalization

$$|\psi_n\rangle_N = Z_n^{1/2} |\psi_n\rangle \quad (1.9)$$

fix such that

$${}_N \langle \psi_n | \psi_n \rangle_M = 1 \quad (1.10)$$

$$\begin{aligned} {}_N \langle \psi_n | \psi_n \rangle_N &= 1 \\ &= Z_n \langle \psi_n | \psi_n \rangle \\ &= Z_n \left( \langle \psi_n^{(0)} | + \lambda \langle \psi_n^{(1)} | + \lambda^2 \langle \psi_n^{(2)} | \right) \\ &\quad \times \left( | \psi_n^{(0)} \rangle + \lambda | \psi_n^{(1)} \rangle + \dots \right) \\ &= E_n^{(2)} \delta_{kn} \\ &= Z_n \left( 1 + \lambda^2 \langle \psi_n^{(1)} | \psi_n^{(1)} \rangle + \mathcal{O}(\lambda^3) \right) \end{aligned} \quad (1.11)$$

$$Z_n^{1/2} = 1 - \frac{\lambda^2}{2} \langle \psi_n^{(1)} | \psi_n^{(1)} \rangle \mathcal{O}(\lambda^3) \quad (1.12)$$

$$\begin{aligned} &\Rightarrow | \psi_n^{(2)} \rangle \\ &= \sum_{k \neq n} \sum_{m \neq n} | \psi_k^{(0)} \rangle \left[ \frac{V_{km} - V_{mn}}{(E_n^{(0)} - E_k^{(0)}) (E_m^{(0)} - E_n^{(0)})} - \frac{V_{kn} V_{nn}}{(E_n^{(0)} - E_n^{(0)})} \right] \\ &\quad - \frac{1}{2} \sum_{k \neq n} | \psi_n^{(0)} \rangle \frac{\| V_{kn}^2 \|}{(E_n^{(0)} - E_k^{(0)})} \end{aligned} \quad (1.13)$$

## 1.2 Time-independent perturbation theory: degenerate case

Assume  $E_n^{(0)}$  is  $\alpha$ -fold degenerate i.e.

$$H_0 | \psi_{n_i}^{(0)} \rangle = E_n^{(0)} | \psi_{n_i}^{(0)} \rangle, \quad 1 \leq i \leq \alpha \quad (1.14)$$

fix

$$\langle \psi_{n_i}^{(0)} | \psi_{n_j}^{(0)} \rangle = \delta_{ij} \quad (1.15)$$

Any linear combination

$$| \chi_n^{(0)} \rangle = \sum_{i=1}^{\alpha} c_{n_i} | \psi_{n_i}^{(0)} \rangle \quad (1.16)$$

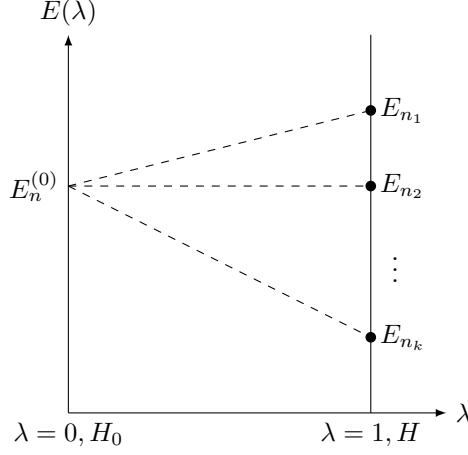


Figure 1.1:

is an eigenstate of  $H_0$  with evaluation  $E_n^{(0)}$ . Typically  $V$  “lifts” degeneracy at least partially since often

$$[H_0, V] \neq 0 \quad (1.17)$$

Pick one of the evals  $E_{n_k}$  with

$$H |\psi_{n_k}\rangle = E_{n_k} |\psi_{n_k}\rangle \quad (1.18)$$

for  $\lambda \rightarrow 0$ :  $E_{n_k} \rightarrow E_n^{(0)}$  and

$$\begin{aligned} |\psi_{n_k}\rangle &\rightarrow |\chi_{n_k}(0)\rangle \\ &= \sum_{i=1}^{\alpha} c_{n_k i} \underbrace{|\psi_{n_i}^{(0)}\rangle}_{\text{some lin comb}} \end{aligned} \quad (1.19)$$

have to find “good” linear combination, i.e. coeff  $c_{n_k i}$ . Main idea as before:

$$|\psi_{n_k}\rangle = |\chi_{n_k}^{(0)}\rangle + \lambda |\psi_{n_k}^{(1)}\rangle \quad (1.20)$$

$$0 = (H_0 - E_n^{(0)}) |\psi_{n_k}^{(1)}\rangle + (V - E_{n_k}^{(1)}) |\chi_{n_k}^{(1)}\rangle \quad (1.21)$$

with

$$|\psi_{n_k}^{(1)}\rangle = \sum_{\ell=1}^{\dim(H_0)} a_{n_\ell} |\psi_\ell^{(0)}\rangle \quad (1.22)$$

and

$$\sum_{i=1}^{\alpha} c_{n_k i} \left| \psi_{n_i}^{(0)} \right\rangle \quad (1.23)$$

multiply by  $\left\langle \psi_{n_j}^{(0)} \right|$ .

$$\begin{aligned} \sum_{\ell=1}^{\dim H_0} \underbrace{(E_{\ell}^{(0)} - E_n^{(0)})}_{=0 \text{ for } n=\ell} a_{n\ell} \underbrace{\left\langle \psi_{n_j}^{(0)} \left| \psi_{\ell}^{(0)} \right\rangle\right.}_{=0 \text{ for } n \neq \ell} + \sum_{i=1}^{\alpha} c_{n_k i} \left( \left\langle \psi_{n_j}^{(0)} \left| V \right| \psi_{n_j}^{(0)} \right\rangle \right. \\ \left. - E_{n_k}^{(1)} \underbrace{\left\langle \psi_{n_j}^{(0)} \left| \psi_{n_i}^{(0)} \right\rangle\right.}_{V_{ji}} - E_{n_k}^{(1)} \underbrace{\left\langle \psi_{n_j}^{(0)} \left| \psi_{n_i}^{(0)} \right\rangle\right.}_{\delta_{ij}} \right) \end{aligned} \quad (1.24)$$

→ solve

$$\det \begin{pmatrix} V_{11} - E_{n_k}^{(1)} & V_{12} & \dots & V_{1\alpha} \\ V_{21} & V_{22} - E_{n_k}^{(1)} & \dots & \vdots \\ \vdots & \vdots & \ddots & \vdots \\ V_{\alpha 1} & \dots & \dots & V_{\alpha\alpha} - E_{n_k}^{(1)} \end{pmatrix} = 0 \quad (1.25)$$

→ eq. of order  $\propto$  in  $E_{n_k}^{(1)}$

→  $\alpha$  solutions

### 1.2.1 easy way out (sometimes)

if  $V_{ij} = 0$  for  $i \neq j$  problem already solved

→  $\alpha$  solutions are

$$E_{n_i}^{(1)} = \left\langle \psi_i^{(0)} \left| V \right| \psi_i^{(0)} \right\rangle \quad (1.26)$$

Note: if  $\exists$  operator  $A$  with

$$[A, V] = 0 \quad (1.27a)$$

and

$$A \left| \psi_{n_i}^{(0)} \right\rangle = a_{n_i} \left| \psi_{n_i}^{(0)} \right\rangle, \quad (1.27b)$$

with

$$a_{n_i} \neq a_{n_k}, \quad \text{for } k \neq i \quad (1.27c)$$

then these  $|\psi_{n_i}^{(0)}\rangle$  are “good” eigenstates

**Proof:**

$$\langle \psi_{n_i}^{(0)} | [A, V] | \psi_{n_i}^{(0)} \rangle = 0 \quad (1.28a)$$

$$= \underbrace{(a_{n_i} - a_{n_k})}_{\neq 0} \underbrace{\langle \psi_{n_i}^{(0)} | V | \psi_{n_i}^{(0)} \rangle}_{V_{ik}} \Rightarrow V_{ik} = 0 \quad (1.28b)$$

### 1.3 The variational principle

Useful to get good estimate of ground-state energy  $E_0$  of complicated systems.

Claim

$$E_0 \leq \frac{\langle \psi | H | \psi \rangle}{\langle \psi | \psi \rangle} = \langle \psi | H | \psi \rangle \quad (1.29)$$

if  $|\psi\rangle$  normalized.

**Proof:** Let

$$|\psi\rangle = \sum c_n |\psi_n\rangle, \quad (1.30a)$$

with

$$H |\psi_n\rangle = E_n |\psi_n\rangle \quad (1.30b)$$

and

$$\langle \psi | \psi \rangle = 1 \quad (1.30c)$$

$$\Rightarrow \sum \|c_n\|^2 = 1 \quad (1.30d)$$

then

$$\begin{aligned} \langle \psi | H | \psi \rangle &= \sum_{m,n} c_m^* c_n \langle \psi_m | H | \psi_n \rangle = \sum_{m,n} c_m^* c_n E_n \underbrace{\langle \psi_m | \psi_n \rangle}_{\delta_{mn}} \\ &= \sum_n \|c_n\|^2 E_n \geq E_0 \sum_n \|c_n\|^2 = E_0 \end{aligned} \quad (1.30e)$$

**Example 1.3.1 (Harmonic oscillator):**

$$H = -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} - \frac{m}{2} \omega^2 x^2 \quad (1.31)$$

(of course we know  $E_0 = \frac{\hbar}{2}\omega$ ). Let

$$\psi(x) = Ae^{-bx^2} \quad (1.32a)$$

since

$$\begin{aligned} \langle \psi | \psi \rangle &\stackrel{!}{=} 1 = \int dx \|A\|^2 e^{-2bx^2} \\ &= \|A\|^2 \sqrt{\frac{\pi}{2b}} \end{aligned} \quad (1.32b)$$

compute

$$\begin{aligned} \langle \psi | H | \psi \rangle &= \|A\|^2 \int dx e^{-bx^2} \left( -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} - \frac{m}{2} \omega^2 x^2 \right) e^{-bx^2} \\ &= \dots \\ &= \frac{\hbar^2 b}{2m} + \frac{m\omega^2}{8b} \\ &= \langle \psi | H | \psi \rangle \\ &\geq E_0 \end{aligned} \quad (1.32c)$$

Minimize with respect to  $b$

$$\begin{aligned} \frac{d}{db} \langle \psi | H | \psi \rangle &= \frac{\hbar^2}{2m} - \frac{m\omega^2}{8b^2} \\ &= 0 \end{aligned} \quad (1.33a)$$

$$b_{\min} = \frac{m\omega}{2\hbar} \quad (1.33b)$$

$$\begin{aligned} E_0 &\leq \langle \psi | H | \psi \rangle_{\min} \\ &= \frac{\hbar\omega}{2} \end{aligned} \quad (1.33c)$$

in this case we get  $E_0$  exactly is a coincidence, since Ansatz=true wave function.

## 1.4 WKB approximation, semiclassical approximation

WKB for Wentzel, Kramers, Brillouin (see QMI Ch. 8.3.) useful for 1-dim problems with “smooth” potential. Schrödinger:

$$\left( -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} V(x) \right) \psi(x) = E\psi(x) \quad (1.34a)$$

if

$$V(x) \equiv V_0 \text{ const.} \quad (1.34b)$$

$$\psi(x) = e^{\pm \frac{i}{\hbar} \sqrt{2m(E-V_0)}x} \quad (1.34c)$$

if  $V(x)$  is slowly varying. Ansatz

$$\psi(x) = e^{\frac{i}{\hbar} S(x)} \quad (1.34d)$$

Ansatz into Schrödinger:

$$\frac{-i\hbar}{2m} S'' + \frac{1}{2m} (S')^2 + V(x) - E = 0 \quad (1.34e)$$

equivalent to but more complicated than Schrödinger. Note for

$$V(x) \equiv V_0$$

$$S = \pm \sqrt{2m(E-V_0)} \cdot x$$

and

$$S'' = 0$$

first term  $\sim \hbar$  vanishes for

$$V(x) \equiv V_0, \quad (\text{classical limit}),$$

Let

$$S(x) = S_0(x) + \hbar S_1(x) + \mathcal{O}(\hbar^2) \quad (1.35a)$$

plug in into differential equation for  $S$

$$\frac{1}{2m} (S'_0)^2 + V(x) - E = 0 \quad (1.35b)$$

$$\begin{aligned} \Rightarrow S'_0 &= \pm \sqrt{2m(E-V(x))} \\ &\equiv \pm p(x) \end{aligned} \quad (1.35c)$$

$$S'_0 S'_1 - \frac{1}{2} S''_0 = 0 \quad (1.35d)$$

$$\Rightarrow S'_1 = \frac{i}{2} \frac{S''_0}{S'_0} = \frac{i}{2} \frac{p'(x)}{p(x)} \quad (1.35e)$$

solve these differential equation

$$S_0 = \pm \int^x dx' p(x') \quad (1.35f)$$

$$S_1 = \frac{i}{2} \ln p(x) \quad (1.35g)$$

$$\begin{aligned} \Rightarrow \psi(x) &= A e^{\frac{i}{\hbar} (S_0 + \hbar S_1)} \\ &= \frac{A_+}{\sqrt{p(x)}} e^{\frac{i}{\hbar} \int dx' p(x')} + \frac{A_-}{\sqrt{p(x)}} e^{-\frac{i}{\hbar} \int dx' p(x')} \end{aligned} \quad (1.35h)$$



# THE HYDROGEN ATOM

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## 2.1 Basics

Two body problem proton (1)-electron (2)

$$H = -\frac{\hbar^2}{2m_2}\nabla_1^2 - \frac{\hbar^2}{2m_2}\nabla_2^2 + V(\mathbf{r}_1 - \mathbf{r}_2) \quad (2.1)$$

new variables

$$R = \frac{m_2 r_1 + m_1 r_2}{m_1 + m_2} \quad (2.2)$$

$$r = r_1 - r_2 \quad (2.3)$$

$$M = m_1 + m_2 \quad (2.4)$$

$$m = \frac{m_1 m_2}{m_1 + m_2} \quad (2.5)$$

$$H = -\frac{\hbar^2}{2M} \quad (2.6)$$


---

## 2.2 Spin-orbit term

naive “derivation”

(1) Electron with spin  $\rightarrow$  magnetic dipole moment

$$\boldsymbol{\mu} = \frac{e}{m} \frac{g}{2} \mathbf{s}, \quad (2.7)$$

$$\mu = \frac{e}{T} \pi r^2, \quad s = \frac{2\pi}{T} m r^2, \quad g \simeq \text{(from Dirac)} \quad (2.8)$$

(2) Electron feeds magnetic field due to the proton

$$\mathbf{E} \sim \frac{e}{r^3} \mathbf{r} \quad (2.9)$$

$$\begin{aligned} \rightarrow \mathbf{B} &= -\frac{1}{c^2} \nabla \times \mathbf{E} \\ &= -\frac{1}{mc^2 r^3} \mathbf{p} \times \mathbf{r} \\ &= \frac{-\mathbf{L}}{mc^2 r^3} \end{aligned} \quad (2.10)$$

wrong by factor 2 (Thomas precession)

... correct result

$$H_{\text{SO}} = \frac{Ze^2}{2mc^2} \frac{1}{r^3} \mathbf{L} \cdot \mathbf{S}, \quad (\sim -\boldsymbol{\mu} \cdot \mathbf{B}) \quad (2.11)$$

To describe spin

$$\begin{aligned} \left| n, \ell, \left( s = \frac{1}{2}, m_s \right) \right\rangle &= \psi_{n\ell m_\ell m_s} \\ &= \psi_{n\ell m_\ell}(r, \theta, \varphi) \chi_{m_s} \end{aligned} \quad (2.12a)$$

with  $\chi_{m_s}$  spin-orbit

$$\begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (2.12b)$$

Note  $H_{\text{SO}}$  “mixes” states with same  $\ell$ , but different  $m_\ell, m'_\ell$

→ use degenerate perturbation theory with  $2 \cdot (2\ell + 1) \times \underbrace{2}_{\text{spin}} \underbrace{(2\ell + 1)}_{m_\ell}$  matrix

$$\langle n, \ell, m'_\ell, m'_s | H_{\text{SO}} | n, \ell, m'_\ell, m'_s \rangle \rightarrow \text{diagonalize} \quad (2.13)$$

recall degenerate perturbation theory → find “good” linear combination that diagonalize this matrix by looking for symmetry use total angular momentum

$$J \equiv L + S \quad (2.14)$$

for  $\ell = 0$   $j = \frac{1}{2}$ , for  $\ell \neq 0$   $j = \ell \pm \frac{1}{2}$ . Use states

$$|n, \ell, j, m_j\rangle \quad (2.15)$$

$$|n, \ell, j, m_j\rangle = \sum_{m_\ell, m_s} |n, \ell, m_\ell, m_s\rangle \underbrace{\langle n, \ell, m_\ell, m_s | n, \ell, j, m_j \rangle}_{\text{Clebsch-Gordan}} \quad (2.16)$$

use

$$J^2 = L^2 + 2L \cdot S + S^2 \quad (2.17a)$$

$$L \cdot S = \frac{1}{2} (J^2 - L^2 - S^2) \quad (2.17b)$$

$|n, \ell, j, m_j\rangle$  are eigenstates of

$$H_0, L^2, S^2, J^2, J_z \quad (2.18)$$

with eigenvalues

$$E_n, \hbar^2 \ell(\ell+1), \hbar^2 \frac{3}{4}, \hbar^2 j(j+1), \hbar m_j \quad (2.19)$$

$$\Delta E_{\text{SO}} = \langle n, \ell, j, m_j | H_{\text{SO}} | n, \ell, j, m_j \rangle \quad (2.20)$$

for  $\ell = 0$

$$\Delta E_{\text{SO}} = 0 \quad (2.21a)$$

for  $\ell \neq 0$

$$\begin{aligned} \Delta E_{\text{SO}} &= \frac{Ze^2}{2m^2c^2} \langle n, \ell, j, m_j | \frac{1}{r^3} \frac{1}{2} (J^2 - L^2 - S^2) | n, \ell, j, m_j \rangle \\ &= \frac{Ze^2}{2m^2c^2} \left\langle \frac{1}{r^3} \right\rangle \frac{\hbar^2}{2} \left( j(j+1) - \ell(\ell+1) - \frac{3}{4} \right) \\ &= -E_n \frac{(Z\alpha)^2}{2n(\ell + \frac{1}{2})} \begin{cases} \frac{1}{\ell+1} & j = \ell + \frac{1}{2} \\ -\frac{1}{\ell} & j = \ell - \frac{1}{2} \end{cases} \end{aligned} \quad (2.21b)$$

## 2.3 Darwin term

Sloppy consideration electron position fluctuates by  $\delta r \simeq \lambda_c \simeq \frac{\hbar}{mc}$  electron feels average potential

$$\langle V(r + \delta r) \rangle = \langle V(r) \rangle + \underbrace{\frac{1}{2} \langle \delta \rangle r \cdot \nabla \delta r \cdot \nabla V}_{\text{Darwin term}} \quad (2.22)$$

correct result is

$$\begin{aligned} H_D &= \frac{\hbar^2}{8m^2c^2} \nabla^2 V \\ &= \frac{\pi \hbar^2 Z e^2}{2m^2c^2} \delta(r) \end{aligned} \quad (2.23)$$

only for  $\ell = 0$ !

$$\begin{aligned} \Delta E_D &= \langle n, \ell, j, m_j | H_D | n, \ell, j, m_j \rangle \\ &= \frac{\pi \hbar^2 Z e^2}{2m^2c^2} \|\psi_{n\ell}(0)\|^2 \\ &= -E_n \frac{(Z\alpha)^2}{n} \delta_{\ell 0} \end{aligned} \quad (2.24)$$

## 2.4 Fine structure of hydrogen

Combine  $\Delta E_{\text{rel}}$ ,  $\Delta E_{\text{SO}}$  and  $\Delta E_D$

$$\Delta E_n = -E_n^{(0)} \frac{(Z\alpha)^2}{n^2} \left( \frac{3}{4} - \frac{n}{j + \frac{1}{2}} \right) \quad (2.25)$$

valid for  $\ell = 0$ , i.e.  $j = \frac{1}{2}$  and  $j = \ell \pm \frac{1}{2}$ .

fine structure suppressed  $\sim (Z\alpha)^2$  relative to  $E_n^{(0)}$  only depends on  $j$  (not independently on  $\ell$  and  $s$ ).

Notation for states  $nL_J$

$$n = 1, 2, \dots$$

$$L \equiv S(\ell = 0), P(\ell = 1), D(\ell = 2), F(\ell = 3)$$

### 2.4.1 first few states

$$\begin{array}{cccc} \ell = 0 & \ell = 1 & \ell = 2 & \text{degeneracy } 2n^2 \\ n = 1 & & & \end{array}$$

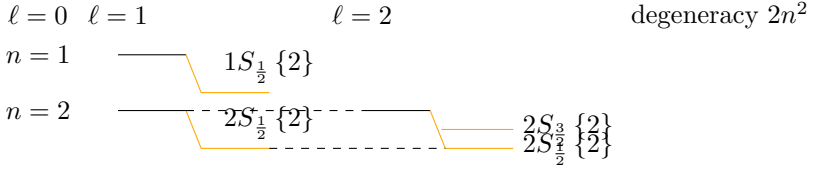


Figure 2.1:

## 2.5 Corrections beyond fine structure

### 2.5.1 Hyperfine structure

Effect of proton spin  $S_p$

$$\boldsymbol{\mu}_p = \frac{e}{2m_p} g_p \mathbf{S}_p \quad (2.26)$$

$\mu_p$  indices  $B$ -field  $\mu$  of electron “feels”  $B$ -field  $\sim \boldsymbol{\mu} \cdot \mathbf{B}$

$$\rightsquigarrow \Delta E_{\text{Hfs}} \sim (Z\alpha)^4 \frac{m_e}{m_p}, \quad (2.27)$$

with  $(Z\alpha)^4$  as always and  $\frac{m_e}{m_p}$  suppression. total spin

$$\begin{aligned} F &= S_e + S_p \\ &= \begin{cases} 1 & \text{triplet} \\ 0 & \text{singlet} \end{cases} \end{aligned} \quad (2.28)$$

### 2.5.2 Lamb shift (needs QED!)

$\rightsquigarrow$  modification of Coulomb potential. splits e.g.  $2S_{\frac{1}{2}}$  and  $2p_{\frac{1}{2}}$

$$\Delta E_{\text{Lamb}} \sim (Z\alpha)^4 \cdot \alpha \left( \frac{1}{\log(Z\alpha)} \right) \dots \quad (2.29)$$



# MANY ELECTRON ATOMS

---

$$H = \sum_{i=1}^N \left( \frac{p^2}{2m} - \frac{Ze^2}{r} \right) + \underbrace{\sum_{i>j} \frac{e^2}{|\mathbf{r}_i - \mathbf{r}_j|}}_1 \quad (3.1)$$

→ complicated! we want:

$$H\psi(1, \dots, N) = E\psi(1, \dots, N) \quad (3.2)$$


---

## 3.1 Identical particles

Consider  $N$  identical particles  $H(1, \dots, N)$  wave function  $\psi(1, \dots, N)$ .

In classical mechanics we can always distinguish these  $n$  particles state. In Quantum Mechanics, we cannot keep track of individual particles if their wave functions overlap. Defin permutation operator  $P_{ij}$  interchanging  $i$  and  $j$

$$P_{ij}\psi(1, \dots, i, \dots, j, \dots, N) = \psi(1, \dots, j, \dots, i, \dots, N) \quad (3.3a)$$

$$P_{ij}^2 = 1 \quad (3.3b)$$

$$\Rightarrow \text{evals } \pm 1 \quad (3.3c)$$

---

<sup>1</sup>interaction between electrons

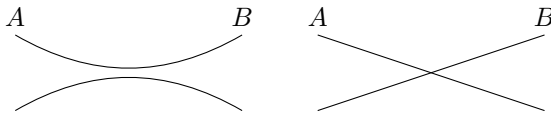


Figure 3.1:

$H$  must be invariant under  $i \leftrightarrow j$

$$\Rightarrow [H, P_{ij}] = 0 \quad (3.4)$$

There are  $N!$  permutations of elements  $1 \cdots N$ , they fall into two classes, even and odd:

$$\text{tr}(P) = \begin{cases} +1 & (-1)^2 \text{ even nr. of Interchanges} \\ -1 & (-1)^2 \text{ odd nr. of Interchanges} \end{cases} \quad (3.5)$$

we have  $[H, P]$ .  $P$  is unitary since

$$\langle \chi | \psi \rangle = \langle P\xi | P\psi \rangle \quad (3.6a)$$

$$= \langle \xi | P^\dagger P | \psi \rangle$$

$$\Rightarrow P^\dagger P = 1 \quad (3.6b)$$

for any observable  $A$  we have

$$[A, P] = 0. \quad (3.6c)$$

(*Identical* part) two combinations are important:

i) totally symmetric  $|\psi\rangle_S$  with

$$P |\psi\rangle_S = |\psi\rangle_S \quad (3.7)$$

with  $|\psi\rangle_S$  completely symmetric linear combination of all  $N!$  Permutations.

ii) totalli antisymmetric  $|\psi\rangle_A$  with

$$P |\psi\rangle_A = (-1)^P |\psi\rangle_A \quad (3.8)$$

...

## 3.2 Thomas-Fermi approximation

$\simeq$  semi classical: assume each electron feels average potential  $\Phi(r)$ , (spherically symmetric)

$$V = -\frac{Ze^2}{r} \xrightarrow{\text{other electron}} -e\Phi(r) \quad (3.9)$$



Poisson equation:

$$\nabla^2 \Phi = -4\pi \tilde{\rho} \stackrel{r \geq 0}{=} 4\pi e \rho(r) \quad (3.10)$$

total charge density (other electron and nucleus)

$$\tilde{\rho} = -e\rho(r) + Ze\delta(r) \quad (3.11)$$

find relation between  $\rho$  and  $\Phi$ . Let  $n$  be nr. states in certain energy range

$$n = \frac{2}{(2\pi\hbar)^3}, \quad \text{if } E = \frac{p^2}{2m} - e\Phi < 0 \quad (3.12a)$$

$$n = 0, \quad \text{if } E = \frac{p^2}{2m} - e\Phi > 0 \quad (3.12b)$$

$$\begin{aligned} \rho &= \int_0^{\sqrt{2me\Phi}} d^3p n \\ &= \frac{(4\pi)2}{(2\pi\hbar)^3} \int_0^{\sqrt{2me\Phi}} \end{aligned} \quad (3.12c)$$

plug this into Poisson  $\rightarrow$  differential equation for  $\Phi$

$$\begin{aligned} \nabla^2 \Phi &= \frac{1}{R} \frac{d^2}{dr^2} \Phi(r) \\ &= \frac{32\pi^2 e}{3(2\pi\hbar)^3} (2me\Phi)^{3/2} \end{aligned} \quad (3.13)$$

solve numerically. boundary condition:

$$\phi(r) \rightarrow \frac{Ze}{r}, \quad \text{for } r \rightarrow 0 \quad (3.14a)$$

normalize

$$4\pi \int dr \rho(r) r^2 = Z \quad (3.14b)$$

$\rightsquigarrow$  “radius” of atom (contains all but one electron)

$$\overline{R} \simeq \text{consanta} Z^{1/3} \quad (3.15)$$

### 3.3 The Hartree approximation

Assume

$$\psi(1 \dots N) = \varphi_1(1) \dots \varphi_N(N) \quad (3.16a)$$

as solution to

$$H\psi = E\psi \quad (3.16b)$$

with

$$H = \sum_i \left( \frac{p_i^2}{2m} - \frac{Ze^2}{r_i} \right) + \sum_{i>j} \frac{e^2}{|\mathbf{r}_i - \mathbf{r}_j|} \quad (3.16c)$$

let  $\varphi_i$  be distinct and orthogonal (partially taking into account Pauli principle) and normalized

$$\int d^3r_i |\varphi_i(r_i)|^2 = 1 \quad (3.17)$$

Want to find stationary state with respect to variation in  $\varphi_i$  taking into account normalization via Lagrange multipliers  $\varepsilon_i$

$$\begin{aligned} \langle H \rangle = & \sum_i \int d^3\mathbf{r} \left( \varphi_i^*(\mathbf{r}) \left( -\frac{\hbar^2}{2m} \nabla^2 - \frac{Ze^2}{r} \right) \varphi_i(\mathbf{r}) \right) \\ & + \sum_{i>j} \int d^3\mathbf{r} \int d^3\mathbf{r}' \varphi_i^*(\mathbf{r}) \varphi_j^*(\mathbf{r}') \frac{e^2}{|\mathbf{r} - \mathbf{r}'|} \varphi_i(\mathbf{r}) \varphi_j(\mathbf{r}') \\ & + \sum_i \varepsilon_i \left( \int d^3r |\varphi_i(r)|^2 - 1 \right) \end{aligned} \quad (3.18)$$

take functional derivative  $\frac{\delta}{\delta \varphi_i^*}$

$$\left( -\frac{\hbar^2}{2m} \nabla^2 - \frac{Ze^2}{r} \right) \varphi_i + V_i(r) \varphi_i(r) = \varepsilon_i \varphi_i(r) \quad (3.19a)$$

$$V_i(r) = \sum_{j \neq i} \int d^3r' \frac{e^2}{|\mathbf{r} - \mathbf{r}'|} |\varphi_j(\mathbf{r}')|^2 \quad (3.19b)$$

interaction of  $i$ -th electron with potential caused by all other ( $j \neq i$ ) electrons  $\varepsilon_i$ : ionization energy of  $i$ -th electron. “solve” numerically with iterative procedure. start with “guess” for  $\varphi_i^{(0)}$ .  $\rightsquigarrow$  into Eq. 3.19b  $\rightsquigarrow V_i^{(0)}$   $\rightsquigarrow$  into Eq. 3.19a solve  $\rightsquigarrow \varphi_i^{(1)}$   $\rightsquigarrow$  etc.

physical interpretation of Lagrange multipliers  $\varepsilon_i \varphi_i^* \cdot 3.19a$

$$\Rightarrow \int d^3r \left( \frac{-\hbar^2}{2m} |\nabla_i \varphi_i|^2 + \left( -\frac{Ze^2}{r_i} + V_i \right) |\varphi_i|^2 \right) = \varepsilon_i \quad (3.20a)$$

with  $\varepsilon_i$  the ionization energy of  $i$ -th electron, assuming others are not affected.

### 3.4 Hartree-Fock approximation

improved ansatz for

$$\psi(1, \dots, N) = \frac{1}{\sqrt{N!}} \begin{vmatrix} \varphi_1(1) & \cdots & \varphi_N(1) \\ \vdots & \ddots & \vdots \\ \varphi_1(N) & \cdots & \varphi_N(N) \end{vmatrix} \quad (3.21a)$$

fully compatible with Pauli  $\rightsquigarrow$  as for Hartree, plug into  $H$  and minimize. Eq. 3.19a stays the same. Eq. 3.19b

$$\begin{aligned} & \frac{1}{2} \sum_{j \neq i} \int d^3\mathbf{r}_i \int d^3\mathbf{r}_j \frac{e^2}{|\mathbf{r}_i - \mathbf{r}_j|} \\ & \times (\varphi_i^*(r_i) \varphi_j^*(r_j) \varphi_i(r_i) \varphi_j(r_j) - \varphi_i^*(r_i) \varphi_j^*(r_j) \varphi_i(r_j) \varphi_j(r_i)) \end{aligned} \quad (3.22a)$$

with  $\varphi_i^*(r_i) \varphi_j^*(r_j) \varphi_i(r_i) \varphi_j(r_j)$  the Hartree term and  $\varphi_i^*(r_i) \varphi_j^*(r_j) \varphi_i(r_j) \varphi_j(r_i)$  the exchange term.

To understand exchange term consider  $N = 2$

$$\psi(1, 2) = \frac{1}{\sqrt{2!}} (\varphi_i(1) \varphi_j(2) - \varphi_j(1) \varphi_i(2)) \quad (3.23a)$$

“new” in H-F write down all terms for

$$\begin{aligned} \langle \psi | \frac{1}{|\mathbf{r}_1 - \mathbf{r}_2|} | \psi \rangle &= \frac{1}{2} \int d^3\mathbf{r}_1 \int d^3\mathbf{r}_2 (\varphi_1^*(r_1) \varphi_2^*(r_2) \varphi_1(r_1) \varphi_2(r_2) + “1 \leftrightarrow 2” \\ &\quad - \varphi_1^*(r_1) \varphi_2^*(r_2) \varphi_1(r_2) \varphi_2(r_1) - “1 \leftrightarrow 2”) \end{aligned}$$

### 3.5 The periodic table and Hund's rules

Electron in atom feels effective potential  $V_{\text{eff}}$  (from nucleus and other electron) which is spherically symmetric.

$$\psi_i = R_{n\ell}(r) Y_{\ell}^m(\theta, \varphi) \chi_{m_s}, \quad (3.24a)$$

with  $\chi_{m_s}$  the spin and  $R_{n\ell}$  different from hydrogen

general rules

$$\begin{cases} n \text{ small} & \text{stronger binding} \\ \ell \text{ small} & \text{electron is closer to nucleus} \end{cases} \tag{3.25}$$

“compete sometimes”. for each  $n$  :

$\ell$	0	1	2	3
name	$S$	$P$	$D$	$F$
$\deg 2(2\ell + 1)$	2	4	6	10

(3.26a)

$K$ -shell	$n = 1$	$\ell = 0$	2 elements	$H, He$	(1s)
$L$ -shell	$n = 2$	$\ell = 0$	2	$Li, Be$	(2s)
		$\ell = 1$	6	$B - Ne$	(2s)
$M$ -shell	$n = 3$	$\ell = 0$	2	$Na - Mg$	(3s)
		$\ell = 1$	6	$Al - Ar$	(3s)
		$\ell = 2$	...	...	(3p)
$N$ -shell	$n = 4$	$\ell = 0$	2		(4s)
	$n = 3$	$\ell = 2$	10	$Al - Ar$	(3d)
	$n = 4$	$\ell = 1$	6	...	(4p)

(3.26b)

configuration of electron  $\rightarrow$  chemical properties of elements. What is the configuration (total spin,  $L, J$ ) of the outer electron.  $\rightarrow$  Hund’s rules (empirical)  
[Notation  $^{2s+1}L_J$ ]

**Example 3.5.1:** Carbon  $(1s)^2(2s)^2(2p)^2$   
for each of the 2  $2p$ -electrons  $2p$ -electrons we can have  $m_\ell = -1, 0, 1$ ,  
 $m_s = -\frac{1}{2}, \frac{1}{2}$ .  $\rightarrow$  6 possibilities  
for both

$$\frac{6 \cdot 5}{2} = 15 \text{ possibilities}$$

(3.27)

$$L = \begin{matrix} 0, 2 & \text{symmetric} \\ 1 & \text{anti-symmetric} \end{matrix}$$

(3.28a)

$$S = \begin{matrix} 1 & \text{symmetric} \\ 0 & \text{anti-symmetric} \end{matrix}$$

(3.28b)

total wave function is antisymmetric

$L$	$S$	$J$	$^{2s+1}L_J$	deg
0	0	0	$^1S_0$	1
		0	$^3P_0$	1
1	1	1	$^3P_1$	3
		2	$^3P_2$	5
2	0	2	$^1D_2$	5
				15

(3.29)

Which one is ground state? → Hund's rules

(1) make spin maximal

$$\begin{aligned}
 &\rightarrow \text{spin part more symmetric} \\
 &\rightarrow \text{orbital part more asymmetric} \\
 &\rightarrow \text{electron further away from each other} \\
 &\rightarrow \text{less repulsion}
 \end{aligned}
 \tag{3.30}$$

for  $C$ :  $s = 1$

(2) make  $L$  maximal

$$\begin{aligned}
 &\rightarrow \text{electron average further away from each other} \\
 &\rightarrow \text{less repulsion}
 \end{aligned}
 \tag{3.31}$$

no impact for  $C$

(3)

$$\Delta E_{\text{SO}} = \text{constant } (j(j+1) - \ell(\ell+1) - s(s+1)) \tag{3.32a}$$

$$\text{constant } \begin{cases} > 0 & \text{if subshell no more than half filled } J = |L - S| \\ < 0 & \text{if subshell more than half filled } J = |L + S| \end{cases} \tag{3.32b}$$

for  $C$  first case and

$$\begin{aligned}
 J &= |L - S| \\
 &= 0
 \end{aligned}
 \tag{3.33}$$

ground state  $^3P_0$



# APPROXIMATION METHODS FOR TIME-DEPENDENT PROBLEMS

---

We now want to know time evolution

$$|\psi(t)\rangle = U(t, t_0) |\psi(t_0)\rangle \quad (4.1)$$

We know:

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = H(t) |\psi(t)\rangle \quad (4.2)$$

with  $H(t)$  now time dependent  $\rightarrow$  more complicated relation between  $H$  and  $U$

---

## 4.1 Time-dependent perturbation theory

Let

$$H(t) = H_0 + \lambda V(t), \quad (4.3a)$$

with  $H_0$  time independent and that can be solved and  $V(t)$  with  $t$  the “only” difference to chapter 1.

$$H_0 |\psi_n^{(0)}\rangle = E_n^{(0)} |\psi_n^{(0)}\rangle \quad (4.3b)$$

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle &= H(t) |\psi(t)\rangle \\ &= (H_0 + \lambda V(t)) |\psi(t)\rangle \end{aligned} \quad (4.3c)$$

for any  $t$ :

$$|\psi(t)\rangle = \sum_n c_n(t) e^{-\frac{i}{\hbar} E_n^{(0)}} |\psi_n^{(0)}\rangle \quad (4.3d)$$

$$\langle \psi(t) | \psi(t) \rangle = 1 \quad (4.3e)$$

$$\Rightarrow \sum_n |c_n|^2 = 1 \quad (4.3f)$$

we can also write

$$V(t) \left| \psi_n^{(0)} \right\rangle = \sum_m \left| \psi_m^{(0)} \right\rangle \left\langle \psi_m^{(0)} \right| V(t) \left| \psi_n^{(0)} \right\rangle \quad (4.3g)$$

with

$$\left\langle \psi_m^{(0)} \right| V(t) \left| \psi_n^{(0)} \right\rangle = V_{mn}(t) \quad (4.3h)$$

→ into Schrödinger

$$\begin{aligned} \sum_n \left( i\hbar \dot{c}_n + E_n^{(0)} c_n \right) \left| \psi_n^{(0)} \right\rangle &= \sum_n c_n e^{-\frac{i}{\hbar} E_n^{(0)} t} \left( E_n^{(0)} \left| \psi_n^{(0)} \right\rangle \right. \\ &\quad \left. + \lambda \sum_m V_{mn}(t) \left| \psi_m^{(0)} \right\rangle \right) \end{aligned} \quad (4.3i)$$

swap labels  $m$  and  $n$  on rhs

$$\sum_n i\hbar \dot{c}_n e^{-\frac{i}{\hbar} E_n^{(0)} t} \left| \psi_n^{(0)} \right\rangle = \sum_{n,m} \lambda c_m e^{-\frac{i}{\hbar} E_m^{(0)} t} V_{nm}(t) \left| \psi_n^{(0)} \right\rangle \quad (4.3j)$$

$$\Rightarrow \dot{c}_n = (i\hbar)^{-1} \lambda \sum_m V_{mn} e^{\frac{i}{\hbar} (E_n^{(0)} - E_m^{(0)}) t} c_m \quad (4.3k)$$

with

$$\omega_{nm} = \frac{E_n^{(0)} - E_m^{(0)}}{\hbar} \quad (4.3l)$$

so

$$\dot{c}_n = (i\hbar)^{-1} \lambda \sum_m V_{nm} e^{i\omega_{nm} t} c_m. \quad (4.3m)$$

Now expand in  $\lambda$  (→ perturbation theory)

$$c_n = c_n^{(0)} + \lambda c_n^{(1)} + \lambda^2 c_n^{(2)} + \dots \quad (4.3n)$$

$$\dot{c}_n^{(0)} = 0, \quad \mathcal{O}(\lambda^0) \quad (4.3o)$$

$$c_n^{(1)} = (i\hbar)^{-1} \sum_m V_{nm} e^{i\omega_{nm} t} c_m^{(0)} \quad (4.3p)$$

...

$$c_n^{(j)} = (i\hbar)^{-1} \sum_m V_{nm} e^{i\omega_{nm} t} c_m^{(j-1)} \quad (4.3q)$$



Let system be in state  $|\psi_i^{(0)}\rangle$  at time to initial condition

$$c_m^{(0)} = \delta_{im} \quad (4.3r)$$

$$\dot{c}_f^{(1)} = (i\hbar)^{-1} V_{fi} e^{i\omega_{fi}t} \quad (4.3s)$$

$$c_f^{(1)}(t) = (i\hbar)^{-1} \int_{t_0}^t dt' V_{fi}(t') e^{i\omega_{fi}t'} \quad (4.3t)$$

$\rightarrow$  transition probability for the system to be found in state  $|\psi_f^{(0)}\rangle$  at time  $t$ .

$$\begin{aligned} P_{i \rightarrow f} &= |c_f^{(1)}|^2 \\ &= \frac{1}{\hbar^2} \left| \int_{t_0}^t dt' V_{fi} e^{i\omega_{fi}t'} \right|^2 + \mathcal{O}(\lambda^2) \end{aligned} \quad (4.3u)$$

approximation only valid if

$$|c_f|^2 \ll 1 \quad (4.3v)$$

Higher orders in  $\lambda$  will be covered section 4.4 and Exercise

## 4.2 Constant perturbation

Let

$$V(t) = \begin{cases} 0 & \text{for } t < t_0 (= 0) \\ V & \text{(constant) for } t > t_0 \end{cases} \quad (4.4a)$$

$$P_{i \rightarrow f} = \frac{1}{\hbar^2} |V_{fi}|^2 \left| \int_{t_0=0}^t dt' e^{i\omega_{fi}t'} \right|^2, \quad (4.4b)$$

using

$$\begin{aligned} \int_{t_0=0}^t dt' e^{i\omega_{fi}t'} &= \frac{2}{\omega^2} (1 - \cos \omega_{fi}t) \\ &= \frac{4}{\omega^2} \sin^2 \left( \frac{\omega_{fi}t}{2} \right), \end{aligned} \quad (4.4c)$$

and

$$\begin{aligned} \delta_t(\alpha) &\equiv \frac{\sin^2(\alpha t)}{\pi \alpha^2 t} \\ &= \begin{cases} \frac{t}{\pi} & \alpha = 0 \\ < \frac{1}{\pi \alpha^2 t} & \alpha \neq 0 \end{cases}, \end{aligned} \quad (4.4d)$$

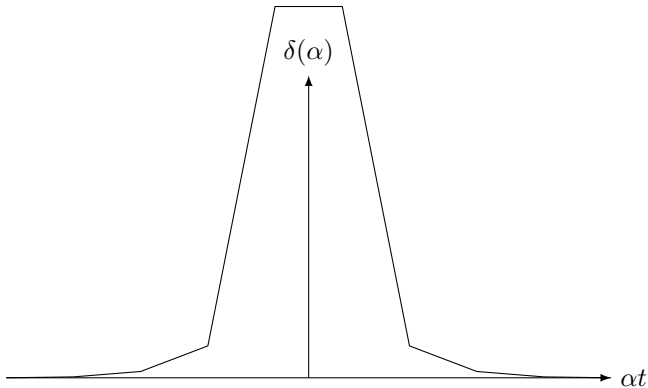


Figure 4.1:

which is plotted in Fig. 4.1

$$\lim_{t \rightarrow \infty} P_{i \rightarrow f} = \frac{\pi t}{\hbar^2} |V_{fi}|^2 \delta \left( \frac{E_f^{(0)} - E_i^{(0)}}{2\hbar} \right) \quad (4.4e)$$

$$P_{i \rightarrow f} = \frac{2\pi t}{\hbar} |V_{fi}|^2 \delta(E_f^{(0)} - E_i^{(0)}) \quad (4.4f)$$

transition rate = probability/time

$$\Gamma_{fi} = \frac{2\pi}{\hbar} |V_{fi}|^2 \delta(E_f - E_i) \quad (4.4g)$$

Consider transitions into continuous spectrum  $\rho(E)$

$$\int_{E_1}^{E_2} dE \rho(E) = \text{number of states in energy range } E_1 - E_2 \quad (4.5a)$$

$$\sum_f \Gamma_{fi} \rightarrow \int dE \rho(E) \Gamma_{fi} = \frac{2\pi}{\hbar} \rho(E_f) |V_{fi}|^2 \quad (4.5b)$$

golden rule!

requires continuum of states and applicability of perturbation theory.

### 4.3 Periodic perturbations

Let

$$V(t) = (V e^{-i\omega t} + V^\dagger e^{+i\omega t}), \quad \text{for } t > t_0 = 0 \quad (4.6)$$

The transition probability  $P_{i \rightarrow f}$  is given by

$$\begin{aligned} P_{i \rightarrow f}(t) &= \frac{1}{\hbar} \left| \int_{t_0}^t dt' (V_{fi} e^{i(\omega_{fi} - \omega)t'} + V_{fi}^\dagger e^{i(\omega_{fi} + \omega)t'}) \right|^2 \\ &= \frac{\pi t}{\hbar^2} \left( |V_{fi}|^2 \sin^2 \left( \frac{t}{2} (\omega_{fi} - \omega) \right) + |V_{fi}^\dagger|^2 \frac{\sin^2 \left( \frac{t}{2} (\omega_{fi} + \omega) \right)}{\pi t \left( \frac{\omega_{fi} + \omega}{2} \right)^2} \right. \\ &\quad \left. + \Re (V_{fi} V_{fi}^\dagger \cdot \mathcal{F}(\omega_{fi}, \omega)) \right) \end{aligned} \quad (4.7a)$$

with  $\mathcal{F}(\omega_{fi}, \omega)$  the interference pattern

the behaviour for large  $t$ ,  $t > \frac{2\pi}{\omega}$  (recall  $\sin^2 \rightarrow \delta$ -function), interference term vanishes, transition rate

$$\Gamma_{i \rightarrow f} = \frac{P_{i \rightarrow f}}{t}, \quad \text{for large } t \quad (4.8a)$$

$$\begin{aligned} \Gamma_{i \rightarrow f} &= \frac{2\pi}{\hbar} \left( |V_{fi}|^2 \delta \left( \underbrace{E_f - E_i - \hbar\omega}_{E_f = E_i + \hbar\omega} \right) \right. \\ &\quad \left. + |V_{fi}^\dagger|^2 \delta \left( \underbrace{E_f - E_i + \hbar\omega}_{E_f = E_i - \hbar\omega} \right) \right) \quad (4.8b) \\ &= \text{absorption of } \hbar\omega + \text{emission of } \hbar\omega \end{aligned}$$

$\rightarrow$  interaction of matter with radiation

### 4.4 The Interaction picture

Consider again evolution operator

$$|\psi(t)\rangle = U(t, t_0) |\psi(t_0)\rangle \quad (4.9)$$

i)  $U(t, t_0) = \mathbf{1}$

ii)  $U(t, t_1)U(t_1, t_0) = U(t, t_0)$

iii)

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle &= H(t) |\psi(t)\rangle \\ &= i\hbar \frac{\partial}{\partial t} U(t, t_0) |\psi(t_0)\rangle \\ &= H(t) U(t, t_0) |\psi(t_0)\rangle \end{aligned} \quad (4.10)$$

$U$  satisfies

$$i\hbar \frac{\partial}{\partial t} U(t, t_0) = H(t) U(t, t_0) \quad (4.11a)$$

Formal solution:

$$U(t, t_0) = 1 + (i\hbar)^{-1} \int_{t_0}^t dt' H(t') U(t', t_0) \quad (4.11b)$$

“solve” by iteration

$$\begin{aligned} U(t, t_0) &= \sum_{n=0}^{\infty} U^{(n)}(t, t_0) \\ &= 1 + U^{(1)} + \dots \end{aligned} \quad (4.11c)$$

$$U^{(1)}(t, t_0) = (i\hbar)^{-1} \int_{t_0}^t dt_1 H(t_1) \quad (4.11d)$$

$$\begin{aligned} U^{(2)}(t, t_0) &= (i\hbar)^{-2} \int_{t_0}^t dt_2 H(t_2) \int_{t_0}^{t_2} dt_1 H(t_1) \\ &= \frac{(i\hbar)^{-2}}{2!} \int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 H(t_2) H(t_1) \end{aligned} \quad (4.11e)$$

(last step see exercise)

$$\vdots = \vdots \quad (4.11f)$$

$$\begin{aligned} U^{(n)}(t, t_0) &= (i\hbar)^{-n} \int_{t_0}^t dt_n H(t_n) \int_{t_0}^{t_n} dt_{n-1} H(t_{n-1}) \dots \int_{t_0}^{t_2} dt_2 H(t_2) \int_{t_0}^{t_1} dt_1 H(t_1) \\ &= \frac{1}{n!} (i\hbar)^{-n} \int_{t_0}^t dt_n \int_{t_0}^{t_n} dt_{n-1} \dots \int_{t_0}^{t_2} dt_2 \int_{t_0}^{t_1} dt_1 H(t_n) \dots H(t_1) \end{aligned} \quad (4.11g)$$

with  $T$  the time ordering operator

$$T(H(t_1) \dots H(t_1)) = H(t_{\tau(1)} \dots H(t_{\tau(n)})) \quad (4.11h)$$

if

$$\begin{aligned} t_{\tau(1)} &> t_{\tau(2)} \\ &> \dots \\ &> t_{\tau(n)} \end{aligned} \quad (4.11i)$$

full solution

$$U(t, t_0) = T \left[ e^{-\frac{i}{\hbar} \int_{t_0}^t dt' H(t')} \right] \quad (4.11j)$$

*cannot* be expanded in a useful way. *All terms are equally important.*

---

## 4.5 Dipole approximation and selection rules

Transitions are governed by  $\alpha^* \langle \psi_0 | j(\mathbf{k}) \cdot \boldsymbol{\epsilon}^*(\mathbf{k}) | \psi_n \rangle$

$$\begin{aligned} j(\mathbf{k}) &= \frac{1}{2m} (\mathbf{p} e^{-i\mathbf{k} \cdot \mathbf{r}} + e^{-i\mathbf{k} \cdot \mathbf{r}} \mathbf{p}) \\ &= \frac{\mathbf{p}}{m} \end{aligned} \quad (4.12a)$$

with Dipole approximation

$$e^{-i\mathbf{k} \cdot \mathbf{r}} = 1 - i\mathbf{k} \cdot \mathbf{r} + \dots \quad (4.12b)$$

$$k \sim \frac{1}{\lambda}, \quad \text{for visible light} \sim (10 \times 10^{-7} \text{ m})^{-1} \quad (4.12c)$$

$$\begin{aligned} r &\sim \text{size of atom} \\ &\sim \text{Bohr radius} \\ &\sim 10 \times 10^{-10} \text{ m} \end{aligned} \quad (4.12d)$$

$$\begin{aligned} \mathbf{j} \cdot \mathbf{A} &\rightarrow \text{Dipole approx.} \\ : \mathbf{j} &= \frac{\mathbf{p}}{m} \end{aligned} \quad (4.12e)$$

(radiation field  $\sim$  constant within atom)

Aside: this is equivalent to a term

$$e\mathbf{r} \cdot \mathbf{E} = -\mathbf{d} \cdot \mathbf{E} \quad (4.12f)$$

with  $\mathbf{d}$  dipole moment

$$\mathbf{d} \equiv -e\mathbf{r} \quad (4.12g)$$

(compare  $\boldsymbol{\mu} \cdot \mathbf{B}$ )

**Proof:**

$$\begin{aligned}
[r_x, H_0] &= \left[ r_x, \frac{\mathbf{p}^2}{2m} \right] \\
&= \frac{1}{2m} ([r_x, \mathbf{p}] \mathbf{p} + \mathbf{p} [r_x, \mathbf{p}])
\end{aligned} \tag{4.13a}$$

with

$$[r_x, \mathbf{p}] = \sum_y i\hbar \delta_{xy} p_y \tag{4.13b}$$

$$[r_x, H_0] = \frac{p_x}{m} i\hbar \tag{4.13c}$$

$$[r, H_0] = i\hbar \frac{\mathbf{p}}{m} \tag{4.13d}$$

emission

$$\begin{aligned}
\langle \psi_0 | e \mathbf{r} \cdot \mathbf{E} | \psi_n \rangle &= -\frac{e}{c} \langle \psi_0 | \mathbf{r} \cdot \dot{\mathbf{A}} | \psi_n \rangle \\
&= -\frac{e}{c} i\omega \langle \psi_0 | \mathbf{r} \cdot \mathbf{A} | \psi_n \rangle \\
&= -\frac{e}{c} \frac{i}{\hbar} (E_n - E_0) \langle \psi_0 | \mathbf{r} \cdot \mathbf{A} | \psi_n \rangle \\
&= -\frac{e}{c} \frac{i}{\hbar} \langle \psi_0 | [\mathbf{r}, H_0] \cdot \mathbf{A} | \psi_n \rangle
\end{aligned} \tag{4.13e}$$

using Eq. 4.13d

$$\begin{aligned}
\langle \psi_0 | e \mathbf{r} \cdot \mathbf{E} | \psi_n \rangle &= -\frac{e}{c} \langle \psi_0 | \frac{\mathbf{p}}{m} \cdot \mathbf{A} | \psi_n \rangle \\
&= \frac{e}{c} \langle \psi_0 | \mathbf{j} \cdot \mathbf{A} | \psi_n \rangle
\end{aligned} \tag{4.13f}$$

**Proof (alternative):** Make gauge trafo with

$$\chi(\mathbf{r}, t) = -\mathbf{A}(t) \cdot \mathbf{r} \tag{4.14a}$$

$$\mathbf{A} \rightarrow \mathbf{A} + \nabla \xi = 0 \tag{4.14b}$$

$$0 = \Phi \rightarrow \Phi - \frac{1}{c} \dot{\chi} = -\frac{1}{c} \dot{\mathbf{A}} \cdot \mathbf{r} = -\mathbf{E} \cdot \mathbf{r} \tag{4.14c}$$

$$H - e\Phi = -e\mathbf{r}\mathbf{E} \tag{4.14d}$$

Atomic transitions are only possible if

$$\langle \psi_{n'\ell'm'_\ell} | \mathbf{r} | \psi_{n\ell m_\ell} \rangle \neq 0 \tag{4.15a}$$

Given  $\ell, m_\ell$  this imposes constraints on  $\ell', m_{\ell'} \rightarrow$  selection rules (in dipole approximation). Can be obtained by (solving at properties of spherical harmonics).

$$\int d\Omega (Y_{\ell'}^{m_{\ell'}}(\theta, \phi))^* \begin{pmatrix} x + iy \\ x - iy \\ z \end{pmatrix} \mathcal{Y}_{\ell}^{m_{\ell}}(\theta, \phi) \quad (4.16a)$$

this is most of the time 0 except if

$$\langle \psi_{n'\ell'm_{\ell'}} | z | \psi_{n\ell m_{\ell}} \rangle \neq 0, \quad \ell' = \ell \pm 1, m_{\ell}' = m_{\ell} \quad (4.16b)$$

$$\langle \psi_{n'\ell'm_{\ell'}} | x \pm iy | \psi_{n\ell m_{\ell}} \rangle \neq 0, \quad \ell' = \ell \pm 1, m_{\ell}' = m_{\ell} \pm 1 \quad (4.16c)$$

(or use Wigner-Ekcart theorem,  $\mathbf{r}$  is a vector operator)

$\Rightarrow$  selection rules for  $E_1$  transitions (dipole approximation)

$$\Delta\ell = \pm 1 \quad (4.16d)$$

$$\Delta m = 0, \pm 1 \quad (4.16e)$$

These rules are violated by “beyond-dipole” transitions (e.g.  $E_2$  quadrupole transitions).

Further selection rules:  $\Delta S = 0$  (spin part of wavefunction not affected by  $\mathbf{r} \cdot \mathbf{E}$ ) always true: *no* transitions between  $j = 0 \rightarrow j = 0$  (total angular momentum conservation)





# POTENTIAL SCATTERING

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We consider

$$H = H_0 + V(r) \quad (5.1a)$$

with

$$\lim_{r \rightarrow \infty} rV(r) = 0, \quad (\text{i.e. not Coloumb}) \quad (5.1b)$$

$V$  is restricted to “small” region.

Want to find stationary solutions

$$\psi(\mathbf{r}, t) = e^{-\frac{i}{\hbar}Et} \psi(\mathbf{r}), \quad \text{for } r \rightarrow \infty \quad (5.1c)$$

steady incoming beam scattered by potential (more general treatment in section 7)

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## 5.1 Elastic scattering and cross sections

we are looking for solutions to

$$(H_0 + V) |\psi\rangle = E |\psi\rangle \quad (5.2a)$$

$$\begin{aligned} E_{\text{in}} &= E_{\text{out}} \\ &= \frac{p^2}{2m} \\ &= \frac{\hbar^2 k^2}{2m} \end{aligned} \quad (5.2b)$$

of the form

$$r \rightarrow \infty \quad (5.2c)$$

$$\begin{aligned} \psi_k(\mathbf{r}) &\rightarrow \psi_{\text{in}}(\mathbf{r}) + \psi_{\text{sc}}(\mathbf{r}) \\ &= e^{i\mathbf{k}\mathbf{r}} + f(\theta, \phi) \frac{e^{ikr}}{r} \end{aligned} \quad (5.2d)$$

where we are not interested in  $\psi(r)$  for small  $r$  (in range of  $V$ ) and  $f(\theta, \phi)$  is the scattering amplitude and  $\frac{e^{ikr}}{r}$  is the outgoing spherical wave

Differential cross section

$$\frac{d\sigma}{d\Omega} = \frac{N}{F} \quad (5.3)$$

$N$ : #particles scattered into  $d\Omega$  per time in  $N d\Omega$

$F$ : flux of incoming particles, number/time/unit area

**Exercise:** For  $|\psi_{\text{in}}|^2 = 1$  (1 part/volume),

$$\begin{aligned} F &= V \\ &= \frac{p}{m} \\ &= \frac{\hbar k}{m} \end{aligned} \quad (5.4a)$$

outgoing:

$$\begin{aligned} \mathbf{j} &= \dots \\ &= \frac{\hbar k}{m} |f(\theta, \phi)|^2 \frac{\mathbf{e}_r}{r^2} \end{aligned} \quad (5.4b)$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = |f(\theta, \phi)|^2 \quad (5.4c)$$

total cross section

$$\sigma = \int d\Omega \frac{d\sigma}{d\Omega} \quad (5.4d)$$

## 5.2 Partial-wave analysis

For central potential  $V(r)$ . No  $\phi$  dependence in  $\psi_k$

$$\psi_k(r, \theta) = \sum_{\ell=0}^{\infty} R_{\ell}(kr) P_{\ell}(\cos \theta) \quad (5.5a)$$

$P_{\ell}(\cos \theta)$  being Legendre polynomials

$$f(\theta) = \sum_{\ell=0}^{\infty} P_{\ell}(k) P_{\ell}(\cos \theta) \quad (5.5b)$$

→ equation for  $R_\ell$ :

$$\frac{d^2}{dr^2} \frac{2}{r} \frac{d}{dr} - \frac{\ell(\ell+1)}{r^2} + 2m(V(r) - E) R_\ell(kr) = 0 \quad (5.5c)$$

Aside: Assume  $V$  is constant ( $\rho = kr$ )

Solutions for  $E > V$ : Spherical Bessel function:

$$j_\ell(\rho) \equiv (-\rho)^\ell \left( \frac{1}{\rho} \frac{d}{d\rho} \right)^\ell \frac{\sin \rho}{\rho} \quad (5.6a)$$

Spherical Neumann function:

$$n_\ell(\rho) = -(-\rho)^\ell \left( \frac{1}{\rho} \frac{d}{d\rho} \right)^\ell \frac{\cos \rho}{\rho} \quad (5.6b)$$

$$j_\ell(\rho) \xrightarrow{\rho \rightarrow 0} \rho^\ell \xrightarrow{\rho \rightarrow \infty} \frac{1}{\rho} \sin\left(\rho - \frac{\pi\ell}{2}\right) \quad (5.7a)$$

$$n_\ell(\rho) \xrightarrow{\rho \rightarrow 0} \frac{1}{\rho^{\ell-1}} \xrightarrow{\rho \rightarrow \infty} -\frac{1}{\rho} \cos\left(\rho - \frac{\pi\ell}{2}\right) \quad (5.7b)$$

$$h_\ell = j_\ell \pm i n_\ell \quad (5.7c)$$

General solution for radial Schrödinger equation for  $r \rightarrow 0$   $\frac{\ell(\ell+1)}{r^2}$  dominant for  $V(r)$  less singular than  $\frac{1}{r^2}$

$$\begin{aligned} R_\ell(r) &\sim j_\ell(\rho) \\ &= j_\ell(kr) \end{aligned} \quad (5.8a)$$

for  $r \rightarrow \infty$   $V(r) \rightarrow 0$  general solution

$$\begin{aligned} R_\ell(r) &= B_\ell(k) j_\ell(kr) + C_\ell n_\ell(kr) \xrightarrow{r \rightarrow \infty} B_\ell(k) \frac{1}{kr} \sin\left(kr - \frac{\pi\ell}{2}\right) \\ &\quad - C_\ell(k) \frac{1}{kr} \cos\left(kr - \frac{\pi\ell}{2}\right) \\ &= \frac{1}{kr} A_\ell(k) \sin\left(kr - \frac{\ell\pi}{2} + \delta_\ell(k)\right), \\ A_\ell &= \sqrt{B_\ell^2 + C_\ell^2}, \text{ phase shift } \tan \delta_\ell = -\frac{C_\ell(k)}{B_\ell(k)} \end{aligned} \quad (5.8b)$$

note: for  $V = 0$ :

$$R_\ell \sim j_\ell(kr) \quad (5.9a)$$

i.e.  $C_\ell$  and  $\delta_\ell = 0$

Solutions for  $r \rightarrow \infty$  are characterized by phase shift  $\delta_\ell$

Next: find relation between  $\delta_\ell \leftrightarrow f(\theta)$

$$\psi_k \rightarrow e^{i\mathbf{k} \cdot \mathbf{r}} + f(\theta) e^{ikr} \frac{1}{r} \quad (5.10)$$

central potential

$$\psi_k = \sum_{\ell=0}^{\infty} R_\ell(kr) P_\ell(\cos \theta) \quad (5.11a)$$

$$f_k = \sum_{\ell=0}^{\infty} f_\ell(k) P_\ell(\cos \theta) \quad (5.11b)$$

Schrödinger equation for  $R_\ell$ , solution  $\stackrel{r \rightarrow \infty}{\sim} j_\ell n_\ell$

$$\begin{aligned} R_\ell(kr) &\stackrel{r \rightarrow \infty}{\sim} \frac{1}{kr} \left( B_\ell \sin \left( kr - \frac{\ell\pi}{2} \right) + C_\ell \cos \left( kr - \frac{\ell\pi}{2} \right) \right) \\ &= \frac{1}{kr} A_\ell \sin \left( kr - \frac{\ell\pi}{2} + \underbrace{\delta_\ell(k)}_{\text{phase shift}} \right) \end{aligned} \quad (5.11c)$$

for  $V = 0$  we have  $C_\ell = 0$ , i.e.  $\rho_\ell = 0$

next: find relation between phase shifts  $\delta_\ell(k)$  and scattering amplitude  $f(\theta)$   
 $(\rightarrow \frac{d\sigma}{d\Omega})$

Aside: free particle eigenfunction in spherical coordinates

$$\psi_{j\ell m_\ell}(r, \phi, \theta) = C_{J\ell}(kr) Y_\ell^{m_\ell}(\theta, \phi), \quad E = \frac{\hbar^2 k^2}{2m} \quad (5.12a)$$

Form a basis, expand

$$e^{i\mathbf{k} \cdot \mathbf{r}} = e^{ikr \cos \theta} \quad (5.12b)$$

in this basis

$$e^{i\mathbf{k} \cdot \mathbf{r}} = \sum_{\ell=0}^{\infty} \sum_{m_\ell}^{\ell} C_{\ell m_\ell} j_\ell(kr) Y_\ell^{m_\ell}(\theta, \phi) \quad (5.12c)$$

having no  $\phi$  dependence

$$\rightsquigarrow e^{i\mathbf{k} \cdot \mathbf{r}} = \sum_{\ell=0}^{\infty} a_\ell j_\ell(kr) P_\ell(\cos \theta) \quad (5.12d)$$

fix coefficient  $a_\ell$  (use orthogonality)

$$e^{i\mathbf{k}\mathbf{r}} = \sum_{\ell=0}^{\infty} (2\ell+1) i^\ell j_\ell(kr) P_\ell(\cos\theta) \quad (5.12e)$$

Put everything into (\*) for  $r \rightarrow \infty$

$$\frac{1}{kr} A_\ell \sin\left(kr - \frac{\pi\ell}{2} + \delta_\ell\right) = (2\ell+1) i^\ell \frac{1}{kr} \sin\left(kr - \frac{\pi\ell}{2}\right) + \frac{A_\ell}{kr} e^{i\delta_\ell} e^{ikr} \quad (5.12f)$$

$$\begin{aligned} \frac{A_\ell}{2i} e^{ikr} e^{-\frac{i\pi\ell}{2}} e^{i\delta_\ell} - \frac{A_\ell}{2i} e^{-ikr} e^{\frac{i\pi\ell}{2}} e^{-i\delta_\ell} &= (2\ell+1) i^\ell \frac{1}{2i} e^{ikr} e^{-\frac{i\ell\pi}{2}} \\ &\quad - (2\ell+1) i^\ell \frac{1}{2i} e^{-ikr} e^{\frac{i\ell\pi}{2}} + k f_\ell e^{ikr} \end{aligned} \quad (5.12g)$$

$$A_\ell = (2\ell+1) i^\ell e^{i\delta_\ell} \quad (5.12h)$$

$$\begin{aligned} f_\ell &= \frac{2\ell+1}{2ik} (e^{2i\delta_\ell} - 1) \\ &= \frac{2\ell+1}{k} e^{i\delta_\ell} \sin(\delta_\ell) \end{aligned} \quad (5.12i)$$

having the full information of  $f(\theta)$ , thus  $\frac{d\sigma}{d\Omega}$

$$\begin{aligned} f(\theta) &= \sum_{\ell=0}^{\infty} \frac{2\ell+1}{k} e^{i\delta_\ell} \sin(\delta_\ell) P_\ell(\cos\theta) \\ &= \sum_{\ell=0}^{\infty} f_\ell P_\ell(\cos\theta) \end{aligned} \quad (5.12j)$$

Total cross section

$$\begin{aligned} \sigma_{\text{tot}} &= \int d\Omega \frac{d\sigma}{d\Omega} \\ &= \int d\Omega |f(\theta)|^2 \\ &= 2\pi \int_{-1}^1 d\cos\theta \sum_{\ell=0}^{\infty} \sum_{\ell'=0}^{\infty} f_\ell f_{\ell'}^* P_\ell(\cos\theta) P_{\ell'}(\cos\theta) \end{aligned} \quad (5.13a)$$

use

$$\int d\cos\theta P_\ell(\cos\theta) P_{\ell'}(\cos\theta) = \frac{2}{2\ell+1} \delta_{\ell\ell'} \quad (5.13b)$$

$$\begin{aligned}
\Rightarrow \sigma_{\text{tot}} &= \sum_{\ell=0}^{\infty} 4\pi \frac{2\ell+1}{k^2} \sin^2(\delta_{\ell}) \\
&\equiv \sum_{\ell=0}^{\infty} \sigma_{\ell}
\end{aligned} \tag{5.13c}$$

### 5.2.1 The optical theorem

$$\begin{aligned}
\text{im}(f(\theta=0)) &= \Im \left( \sum_{\ell=0}^{\infty} f_{\ell} \right) \\
&= \text{im} \left( \sum_{\ell=0}^{\infty} \frac{2\ell+1}{k} e^{i\delta_{\ell}} \sin(\delta_{\ell}) \right) \\
&= \sum_{\ell=0}^{\infty} \frac{2\ell+1}{k} \sin^2(\delta_{\ell})
\end{aligned} \tag{5.14a}$$

$$\begin{aligned}
\Rightarrow \sigma_{\text{tot}} &= \frac{4\pi}{k} \text{im}(f(\theta=0)) \\
&= 0
\end{aligned} \tag{5.14b}$$

often an “easy” way to compute the total cross section by computing  $\text{im}$  of forward scattering amplitude.

Partial-wave useful if not too many  $\sigma_{\ell}$  contribute.

semi-classical: potential of range  $A$ ,  $V(r) = 0$  for  $r > 0$

classical: no scattering if  $b > a$

$$\begin{aligned}
L &\simeq \ell \cdot \hbar \\
&= b \cdot P \\
&= b \cdot \hbar \cdot k
\end{aligned} \tag{5.15}$$

no scattering for

$$\begin{aligned}
b &= \frac{\ell}{k} \\
&> a
\end{aligned} \tag{5.16}$$

### 5.3 Coulomb scattering

So far we have assumed  $\mathbf{r}V(\mathbf{r}) \rightarrow 0$ ,  $|\mathbf{r}| \rightarrow \infty$ , but this is not the case for Coulomb scattering.

However, exact solution is known (recall Hydrogen atom)

$$\left( -\frac{\hbar^2}{2m} \nabla^2 - \frac{Z_1 Z_2 e^2}{r} \right) \psi(\mathbf{r}) = E \psi(\mathbf{r}) \quad (5.17)$$

- hydrogen  $E < 0$  (bound states),

$$\lim_{r \rightarrow \infty} |\psi(\mathbf{r})|^2 = 0 \quad (5.18)$$

- scattering  $E > 0$  with different

$$\nabla^2 = \frac{4}{\xi + \eta} (\partial_\xi \xi \partial_\xi + \partial_\eta \eta \partial_\eta) + \underbrace{\frac{1}{\xi \eta} \frac{\partial^2}{\partial \varphi^2}}_{\spadesuit} \quad (5.19)$$

♠: do not contribute  $\rightarrow$  Confluent hypogeometric equation  $\rightarrow$

2 linearly independent solutions: find the linear combination which is regular at the origin. Result for  $r \rightarrow \infty$

$$\gamma = \frac{m Z_1 Z_2 e^2}{\hbar^2 k} \quad (5.20a)$$

$$\begin{aligned} \psi_\ell(r) \xrightarrow{r \rightarrow \infty} & \underbrace{e^{i(kz + \gamma \log 2k(r-z))}}_{\text{distorted plane wave}} \\ & - \frac{\gamma}{2k \sin^2 \frac{\vartheta}{2}} \frac{\Gamma(1+\gamma)}{\Gamma(1-i\gamma)} e^{-i\gamma \log(\sin \frac{2\vartheta}{2})} \underbrace{\frac{e^{i(kr - \gamma \log 2kr)}}{2}}_{\text{distorted outgoing plane wave}} \end{aligned} \quad (5.20b)$$

This looks the same as

$$\psi(\mathbf{r}) = e^{i\mathbf{k}\mathbf{r}} + f(\vartheta) \frac{e^{ikr}}{r} \quad (5.21)$$

but gets additional phases, because *not*  $rV(r) \xrightarrow{r \rightarrow \infty} \text{Cross-section}$

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= |f_c(\vartheta)| \\ &= \frac{\gamma}{4k^2 \sin^4 \frac{\vartheta}{2}} \\ &= \left( \frac{Z_1 Z_2 e^2}{4E} \right)^2 \frac{1}{\sin^4 \frac{\vartheta}{2}} \end{aligned} \quad (5.22)$$

→ → Rutherford scattering formula (classical!)

→ Phases drop out in this case

Additional phases can have an effect in scattering of 2 identical particles. Here  $2 \rightarrow 2$ ,  $\xrightarrow{A} \xleftarrow{B}$ . Go to the center of mass frame (sec 2.1)

$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + V(\mathbf{r}) \right] \psi(\mathbf{r}) = E\psi(\mathbf{r}) \quad (5.23)$$

•

$$\begin{aligned} \mu &= \frac{m_1 m_2}{m_1 + m_2} \\ &= \frac{m}{2}, \quad \text{reduced mass} \end{aligned} \quad (5.24)$$

$$V \sim \text{interaction potential} \quad (5.25)$$

$$\begin{aligned} \mathbf{r} &= \mathbf{r}_A - \mathbf{r}_B \\ &\sim \text{relative coordinate} \end{aligned} \quad (5.26)$$

In QM the particles are undistinguishable when identical. Two pictures with a particle  $A$  and  $B$  going to each other.

Moreover: Total wave function (spin+space+...) must be either symmetric or antisymmetric under exchange  $A \leftrightarrow B$ ,  $\mathbf{r} \leftrightarrow -\mathbf{r}$ . Spatial wave function

$$\psi_{\text{sym/antysym}} = (e^{i\mathbf{k}\mathbf{r}} \pm e^{-i\mathbf{k}\mathbf{r}}) + (f(\vartheta) \pm f(\pi - \vartheta)) \frac{e^{ikr}}{r} \quad (5.27)$$



**Example:** Coulomb scattering of two protons (spin  $\frac{1}{2}$ , fermions  $\rightarrow$  t.w.f. antisymmetric). Let us look at unpolarized protons and assume that the potential does not depend on spin. Spin wave function:

**prob**  $\frac{1}{4}$  singlet state (anti sym.)

**prob.**  $\frac{3}{4}$  triplet states (sym.)

Spatial wave function

**prob**  $\frac{1}{4}$  symm.

$$\rightarrow \sigma_{\text{sing}} = |f_{\ell}(\vartheta) + f_c(\pi - \vartheta)|^2 \quad (5.28)$$

**prob**  $\frac{3}{4}$  antisym.

$$\rightarrow \sigma_{\text{trp}} = |f_c(\vartheta) - f_c(\pi - \vartheta)|^2 \quad (5.29)$$

Unpolarized cross-section:

$$\begin{aligned} \sigma &= \frac{1}{4} |f_c(\vartheta) + f(\pi - \vartheta)|^2 + \frac{3}{4} |f_c(\vartheta) - f_c(\pi - \vartheta)|^2 \\ &= |f_c(\vartheta)|^2 + |f_c(\pi - \vartheta)|^2 \\ &\quad - \frac{1}{2} (f_c(\vartheta)f_c^*(\pi - \vartheta) + f_c^*(\vartheta)f_c(\pi - \vartheta)) \stackrel{Z_1=Z_2=1}{=} \left( \frac{e}{4E} \right)^2 \left( \underbrace{\frac{1}{\sin \frac{4\theta}{2}} + \frac{1}{\cos^4 \frac{\vartheta}{2}}}_{\text{classical}} \right. \\ &\quad \left. - \frac{\cos(\gamma \log(\tan^2 \frac{\vartheta}{2}))}{\sin^2 \frac{\vartheta}{2} \cos^2 \frac{\vartheta}{2}} \right) \end{aligned} \quad (5.30)$$

*Mott scattering formula*

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## 5.4 Lippman- Schwinger equation & Green's function

Again:

$$E_k = \frac{\hbar^2 k^2}{2m} \quad (5.31a)$$

$$\left( \frac{\hbar^2}{2m} \nabla^2 + E_k \right) \psi_k(\mathbf{r}) = V(\mathbf{r})\psi(\mathbf{r}) \quad (5.31b)$$

If we know the *Green's function* defined by

$$\left( \frac{\hbar^2}{2m} \nabla^2 + E_k \right) g_k(\mathbf{r}) = \delta(\mathbf{r}) \quad (5.32)$$

then we can write a formal solution for  $\psi_k(\mathbf{r})$

$$\psi_k(\mathbf{r}) = e^{i\mathbf{k}\mathbf{r}} + \int d^3\mathbf{r}' g_k(\mathbf{r} - \mathbf{r}') V(\mathbf{r}') \psi_k(\mathbf{r}') \quad (5.33)$$

with  $e^{i\mathbf{k}\mathbf{r}}$  solution to homogeneous equation ( $V = 0$ ) Check it at home.

**Idea:** As in section 4.4 we can turn this formal solution into a series for  $\psi_k$  (in powers of  $V$ )

**First:** Compute  $g_k$ : Go to Fourier space

$$g_k(\mathbf{r}) = \int \frac{d^3\mathbf{q}}{(2\pi)^3} e^{-i\mathbf{q}\mathbf{r}} \tilde{g}_k(\mathbf{q}) \quad (5.34)$$

We get

$$\begin{aligned} \left( \frac{\hbar^2}{2m} \nabla^2 + E_k \right) g_k(\mathbf{r}) &= \int \frac{d^3\mathbf{q}}{(2\pi)^3} \left( -\frac{\hbar^2}{2m} q^2 + \frac{\hbar^2}{2m} k^2 \right) e^{-i\mathbf{q}\mathbf{r}} \tilde{g}_k(\mathbf{q}) \\ &= \delta(\mathbf{r}) \\ &= \int \frac{d^3\mathbf{q}}{(2\pi)^3} e^{-i\mathbf{q}\mathbf{r}} \end{aligned} \quad (5.35a)$$

$$\begin{aligned} \Rightarrow \tilde{g}_k(\mathbf{q}) &= \frac{2m}{\hbar^2} \frac{1}{k^2 - q^2} \\ &= \left( E_k - \frac{\hbar^2 q^2}{2m} \right)^{-1} \end{aligned} \quad (5.35b)$$

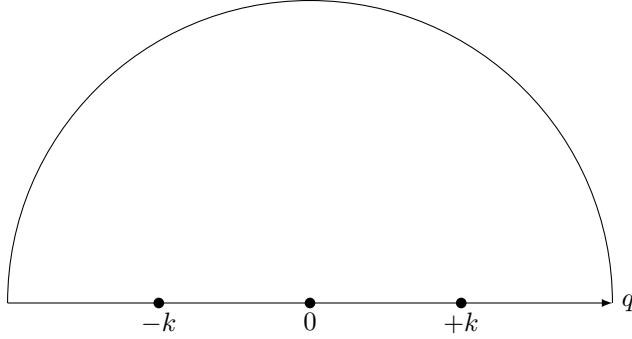


Figure 5.1:

$$\begin{aligned}
 g_k(r) &= \frac{1}{(2\pi)^3} \int_0^\infty dq \int_0^{2\pi} d\varphi \int_{-1}^1 d\cos\vartheta \\
 &\quad \cdot q^2 \frac{2m}{\hbar^2} \frac{1}{k^2 - q^2} e^{-iqr \cos\vartheta} \\
 &= \frac{2m}{\hbar^2} \frac{1}{(2\pi)^2} \\
 &= \frac{2m}{\hbar^2} \frac{1}{(2\pi)^2} \int_0^\infty dq \frac{q^2}{iqr} \frac{1}{k^2 - q^2} \left( e^{iqr} - \underbrace{e^{-iqr}}_{\spadesuit} \right) \\
 &= \frac{m}{2\pi^2 \hbar^2} \frac{1}{ir} \int_{-\infty}^\infty dq \frac{q}{k^2 - q^2} e^{iqr}
 \end{aligned} \tag{5.35c}$$

$\spadesuit$  :

$$\int_0^\infty dq \xrightarrow{q \rightarrow -1} - \int_0^{-\infty} = \int_{-\infty}^0 dq \tag{5.36}$$

Integral is not well defined!  $\rightarrow$  We need a prescription for the poles.

We use contour integral, close it in upper half plane

$$\begin{aligned}
 q &= q_{R\ell} + iq_{\ell m} \\
 &\rightarrow e^{iqr} \\
 &= e^{iq_{R\ell}r} e^{-q_{\ell m}r}
 \end{aligned} \tag{5.37}$$

We deform the contour at  $q = \pm k \rightarrow$  different options, giving *different asympt behaviours* for  $g_k(\mathbf{r})$ ! For negative point outside and positive point inside

curve:

$$\begin{aligned} g_k^+(r) &= 2\pi i \frac{m}{2\pi^2 \hbar^2} \frac{1}{ir} \text{Res}_{q=k} \frac{-q}{(q-k)(q+k)} e^{iqr} \\ &= -\frac{m}{2\pi \hbar^2} \frac{e^{ikr}}{r} \end{aligned} \quad (5.38a)$$

For negative point inside and positive point outside curve:

$$g_k^-(r) = -\frac{m}{2\pi \hbar^2} \frac{e^{-ikr}}{r} \quad (5.38b)$$

for both points inside:

$$g_k^+(r) + g_k^-(r) \quad (5.38c)$$

Both points outside

$$0 \quad (5.38d)$$

$g_k^+ \sim$  outgoing spherical wave

$g_k^- \sim$  incoming spherical wave

What we need is  $g_k^+$ !

$$\psi_k(\mathbf{r}) = e^{i\mathbf{k}\mathbf{r}} - \int d^3\mathbf{r}' \frac{m}{2\pi \hbar^2} \frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} V(\mathbf{r}') \psi(\mathbf{r}') \quad (5.39)$$

Check asymptotic behaviour ( $\mathbf{r}'V(\mathbf{r}') \rightarrow 0$ )

$$\begin{aligned} |\mathbf{r}-\mathbf{r}'| &\xrightarrow{|\mathbf{r}|\rightarrow\infty} r \left( 1 - \frac{r'}{r} \cos\vartheta + \mathcal{O}\left(\left(\frac{r'}{r}\right)\right) \right) \\ &= r - \frac{\mathbf{r}\mathbf{r}'}{r'} \\ &= r - \hat{\mathbf{e}}_r \cdot \mathbf{r}' \end{aligned} \quad (5.40a)$$

We get:

$$\psi_k(\mathbf{r}) = e^{i\mathbf{k}\mathbf{r}} - \frac{m}{2\pi \hbar^2} \int d^3\mathbf{r}' \frac{e^{ikre^{-ik\mathbf{e}_r\mathbf{r}'}}}{r} \times V(\mathbf{r}') \psi_k(\mathbf{r}') \quad (5.40b)$$

...

## 5.5 The Born approximation

Formal solution

$$\psi_k(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}} + \int d^3\mathbf{r}' g_k^+(\mathbf{r} - \mathbf{r}') V(\mathbf{r}') \psi_k(\mathbf{r}') \quad (5.41a)$$

$$g_k^+(\mathbf{r} - \mathbf{r}') = -\frac{m}{2\pi\hbar^2} \frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r} - \mathbf{r}'|} \quad (5.41b)$$

$$f = -\frac{m}{2\pi\hbar^2} \int d^3\mathbf{r}' e^{-ik\mathbf{e}_r\cdot\mathbf{r}'} V(\mathbf{r}') \psi_k(\mathbf{r}') \quad (5.41c)$$

Solve this “pertubatively” by iteration.

$$\psi_k(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}} + \sum_{n=1}^{\infty} \int d^3\mathbf{r}' K(\mathbf{r}, \mathbf{r}') e^{i\mathbf{k}\cdot\mathbf{r}'} \quad (5.42a)$$

$$K_1(\mathbf{r}, \mathbf{r}') = g_k^+(\mathbf{r} - \mathbf{r}') V(\mathbf{r}'), \quad \sim V^1 \quad (5.42b)$$

$$\begin{aligned} K_n(\mathbf{r}, \mathbf{r}') &= \int d^3\mathbf{r}'' K_1(\mathbf{r}, \mathbf{r}'') K_{n-1}(\mathbf{r}'', \mathbf{r}'), \quad \sim V^n \\ &= \int \left( \prod_{i=1}^n d^3\mathbf{r}_i \right) g_k^+(\mathbf{r} - \mathbf{r}) V(\mathbf{r}_n) g_k^+(\mathbf{r}_n - \mathbf{r}_{n-1}) \dots g_k^+(\mathbf{r}_2 \\ &\quad - \mathbf{r}_1) V(\mathbf{r}_1) \end{aligned} \quad (5.42c)$$

Retaining only the 1st term we have (1st) Born approx.

$$\psi_k^{(1)}(\mathbf{r}) = e^{ikz} - \underbrace{\frac{m}{2\pi\hbar^2} \int d^3\mathbf{r}' V(\mathbf{r}') e^{ik(\mathbf{e}_z - \mathbf{e}_r)\cdot\mathbf{r}'} \frac{e^{ikr}}{r}}_{f^{(1)}(\theta, \phi)} \quad (5.43)$$

assuming  $k$  is along the  $z$ -axis.  $f^{(1)}$  is the Fourier transform of potential w.r.t.  $\mathbf{q}$

$$\begin{aligned} \mathbf{q} &\equiv k(\mathbf{e}_z - \mathbf{e}_r) \\ &= \mathbf{k} - \mathbf{k}', \quad \text{momentum transfer} \end{aligned} \quad (5.44a)$$

for central potential  $V(r')$  (not  $V(\mathbf{r}')$ ).

$$\begin{aligned}
 f^{(1)}(\theta, \phi) &= -\frac{m}{2\pi\hbar^2} \int d\phi \int_{-1}^1 d\cos\vartheta \int dr' V(r') e^{iqr' \cos\vartheta} \\
 &= -\frac{m}{\hbar^2} \int r'^2 dr' \frac{2\sin(qr')}{qr'} V(r')
 \end{aligned} \tag{5.45a}$$

$$f^{(1)}(\theta) = -\frac{2m}{\hbar^2 q} \int dr' r' V(r') \sin(qr'), \quad q = 2k \sin \frac{\theta}{2} \tag{5.45b}$$

Note  $f^{(1)}$  is real! Cp optical theorem!?

**Example:** Yukawa potential:

$$V(r) = V_0 \frac{e^{-\mu r}}{r}, \quad \mu \sim \text{range of interaction} \tag{5.46a}$$

$$\begin{aligned}
 f^{(1)} &= -\frac{2m}{\hbar^2 q} V_0 \underbrace{\int_0^\infty dr' e^{-\mu r'} \sin(qr')}_{q/(\mu^2 + q^2)} \\
 &= -\frac{2m}{\hbar^2} \frac{V_0}{\mu^2 + q^2}
 \end{aligned} \tag{5.46b}$$

$$\begin{aligned}
 \left(\frac{d\sigma}{d\Omega}\right)^{(1)} &= |f^{(1)}(\theta)|^2 \\
 &= \left(\frac{2mV_0}{\hbar^2}\right)^2 \frac{1}{(\mu^2 + 2k^2(1 - \cos\theta))^2}
 \end{aligned} \tag{5.46c}$$

total cross section

$$\begin{aligned}
 \sigma_{\text{tot}}^{(1)} &= 2\pi \int_{-1}^1 d\cos\theta \left(\frac{2mV_0}{\hbar^2}\right)^2 \frac{1}{(\mu^2 + 2k^2(1 - \cos\theta))^2} \\
 &= 2\pi \left(\frac{2mV_0}{\hbar^2}\right)^2 \frac{2}{\mu^4 + 4k^2\mu^2}
 \end{aligned} \tag{5.46d}$$

Take limit  $\mu \rightarrow 0$ : “infinite” range of interaction

$$V(\mathbf{r}) = \frac{V_0}{r}, \quad (\text{Coulomb}) \tag{5.47a}$$

$$\left(\frac{d\sigma}{d\Omega}\right)_{\text{Coulomb}}^{(1)} = \left(\frac{2mV_0}{\hbar^2}\right)^2 \frac{1}{(2k^2(1 - \cos\theta))} \tag{5.47b}$$

but total cross section diverges!

Note  $\lim_{r \rightarrow \infty} rV(r) \neq 0$  for Coulomb!

# GENERAL SCATTERING THEORY

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General scattering more complicated than in section 6 e.g. production of new particles. Want to use general representation, so  $|\psi\rangle$  is a general state in Hilbert space (not nec. wave function)

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## 6.1 Dynamics of scattering

Start from time dependent Schrödinger

$$i\hbar\partial_t |\psi, t\rangle = H |\psi, t\rangle \quad (6.1a)$$

$$H = H_0 + V \quad (6.1b)$$

$$\dots = \dots \quad (6.1c)$$

$$\begin{aligned} (i\hbar\partial_t - H_0) |\psi, t\rangle &= V |\psi, t\rangle \\ &= |\chi, t\rangle \end{aligned} \quad (6.1d)$$

**Definition 6.1.1:** Green operation  $G_0(t, t')$  through

$$(i\hbar\partial_t - H_0) G_0(t, t') = \delta(t - t') \cdot \mathbf{1}, \quad (\leftarrow \text{ operators}) \quad (6.2)$$

inhomogeneous differential equation in  $t$

$$G_0^+(t, t') = -\frac{i}{\hbar} \theta(t - t') e^{-\frac{i}{\hbar} H_0(t - t')} \quad (6.3a)$$

$$G_0^-(t, t') = +\frac{i}{\hbar} \theta(t' - t) e^{-\frac{i}{\hbar} H_0(t' - t)} \quad (6.3b)$$

**Proof:**

$$\begin{aligned}
 i\hbar\partial_t G_0^+(t-t') &= i\hbar\left(-\frac{i}{\hbar}\right)\partial_t\theta(t-t')e^{-\frac{i}{\hbar}H_0(t-t')} \\
 &= \delta(t-t')e^{-\frac{i}{\hbar}H_0(t-t')} \\
 &\quad + \theta(t-t')\left(-\frac{i}{\hbar}H_0\right)e^{-\frac{i}{\hbar}H_0(t-t')} \\
 &= \delta(t-t') + H_0G_0^+(t-t')
 \end{aligned} \tag{6.4a}$$

Note: The superscripts  $\pm$  are related to those in section 6

$$G_0^\pm(t, t') = G_0^{pm}(t - t') \tag{6.5a}$$

Write solution to Schrödinger

$$|\psi^\pm, t\rangle = |\psi^0, t\rangle + \int dt' G_0^\pm(t - t') V |\psi^\pm, t'\rangle \tag{6.5b}$$

with  $|\psi^0, t\rangle$  solution to homogeneous problem

$$(i\hbar\partial_t - H_0)|\psi_0, t\rangle = 0 \tag{6.5c}$$

The “physical” solution is given by  $|\psi^+, t\rangle$ , since  $G^+(t - t')$  moves “forward” in time. (retardation)

To make connection with section 6:  $t \rightarrow E$

$$\begin{aligned}
 G_0^+(E) &= \int_{-\infty}^{\infty} dt e^{\frac{i}{\hbar}Et} G_0^+(t) \\
 &= -\frac{i}{\hbar} \int_0^{\infty} dt e^{\frac{i}{\hbar}Et} e^{-\frac{i}{\hbar}H_0t}
 \end{aligned} \tag{6.6a}$$

with

$$G_0^+(t) = G_0^+(t, t' = 0) \tag{6.6b}$$

for  $t \rightarrow \infty$  need  $E \rightarrow E + i0^+, 0^+ > 0$

$$\begin{aligned}
 G_0^\pm &= \frac{1}{E - H_0 \pm i0^+} \\
 &\equiv (E - H_0 \pm i0^+)^{-1}
 \end{aligned} \tag{6.6c}$$



$|\psi^\pm, t\rangle$  evolves with  $H$ , but was equal to free state  $|\psi^0, t\rangle$  @  $t \rightarrow -\infty$

Let  $\alpha$  be a complete set of quantum number of  $H_0$  (including  $E_\alpha$ )

$$\rightarrow |\psi_\alpha^0, t\rangle \leftrightarrow |\psi_\alpha^\pm, t\rangle \quad (6.7a)$$

make Fourier  $t \rightarrow E$  of Eq.  $\pm$

$$\underbrace{\int dt e^{\frac{i}{\hbar}Et} |\psi_\alpha^\pm, t\rangle}_{|\psi_\alpha^\pm(E)\rangle \rightarrow |\psi_\alpha^\pm\rangle} = \underbrace{\int dt e^{\frac{i}{\hbar}Et} |\psi_\alpha^0, t\rangle}_{|\psi_\alpha^0(E)\rangle \sim |\psi_\alpha^0\rangle} + \underbrace{\int dt e^{\frac{i}{\hbar}Et} \int_{t \rightarrow t+t'} dt' G_0^\pm(t-t') V |\psi_\alpha^\pm, t\rangle}_{\quad} \quad (6.7b)$$

$$\int dt e^{\frac{i}{\hbar}Et} G_0^+(t) = G_0^+(E) \quad (6.7c)$$

$$V |\psi_\alpha^\pm(E)\rangle = \int dt' e^{\frac{i}{\hbar}Et'} V (\psi_\alpha^\pm(E)) \quad (6.7d)$$

$$|\alpha^\pm\rangle = |\psi_\alpha^0\rangle + G_0^\pm V |\psi_\alpha^\pm\rangle, \quad (\text{Lippmann-Schwinger}) \quad (6.8a)$$

Solution:

$$\begin{aligned} |\psi_\alpha^\pm\rangle &= (1 - G_0^\pm V)^{-1} |\psi_\alpha^0\rangle \\ &= \frac{1}{(G_0^\pm)^{-1} - V} (G_0^{pm})^{-1} |\psi_\alpha^0\rangle \\ &= \frac{1}{E_\alpha - H_0 \pm i0^+ - V} (E_\alpha - H_0 \pm i0^+ - V + V) |\psi_\alpha^0\rangle \quad (6.9a) \\ &= \frac{1}{\underbrace{E_\alpha - H \pm i0^+}_{G^\pm E_\alpha}} (E_\alpha - H + V \pm i0^+) |\psi_\alpha^0\rangle \\ &= (1 + G^\pm V) |\psi_\alpha^0\rangle \end{aligned}$$

$$\psi_\alpha^\pm = (1 + G^\pm V) |\psi_\alpha^0\rangle \quad (6.9b)$$

can also be obtained from (exercise 2)

$$|\psi^\pm, t\rangle = \lim_{t' \rightarrow \mp\infty} i\hbar G^\pm(t-t') |\psi_0, t'\rangle \quad (6.9c)$$

## 6.2 Møller operators & scattering operator

$\alpha$ : Complete set of quantum numbers

$$H_0 |\psi_\alpha^0\rangle = E_\alpha |\psi_\alpha^0\rangle \quad (6.10a)$$

$$H |\psi_\alpha^\pm\rangle = E_\alpha |\psi_\alpha^\pm\rangle \quad (6.10b)$$

Consider again

$$|\psi_\alpha^\pm\rangle = |\psi_\alpha^0\rangle + (E_\alpha - H_0 + \pm i0^+)^{-1} V |\psi_\alpha^\pm\rangle \quad (6.10c)$$

$$= |\psi_\alpha^0\rangle + \int d\beta \frac{T_{\beta\alpha} |\psi_\beta^0\rangle}{E_\alpha - E_\beta \pm i0^+}$$

$$1 = \int d\beta |\psi_\beta^0\rangle \langle \psi_\beta^0| \quad (6.10d)$$

$$T_{\beta\alpha} \equiv \langle \psi_\beta^0 | V | \psi_\alpha^\pm \rangle \quad (6.10e)$$

Transfer matrix

This state satisfies:

$$\int d\alpha e^{-\frac{i}{\hbar} E_\alpha \tau} f(\alpha) |\psi_\alpha^\pm\rangle \xrightarrow{\tau \rightarrow \mp\infty} \int d\alpha e^{-\frac{i}{\hbar} E_\alpha \tau} f(\alpha) |\psi_\alpha^0\rangle \quad (6.11a)$$

or

$$e^{-\frac{i}{\hbar} H \tau} \int d\alpha f(\alpha) |\psi_\alpha^\pm\rangle \xrightarrow{e^{-\frac{i}{\hbar} H_0 \tau}} \int d\alpha f(\alpha) |\psi_\alpha^0\rangle \quad (6.11b)$$

$$\begin{aligned} \Rightarrow |\psi_\alpha^\pm\rangle &= \lim_{\tau \mp\infty} e^{+\frac{i}{\hbar} H \tau} e^{-\frac{i}{\hbar} H_0 \tau} |\psi_\alpha^0\rangle \\ &= \Omega^\pm |\psi_\alpha^0\rangle \end{aligned} \quad (6.11c)$$

with  $\Omega^\pm$  Møller operators

*Typical* scattering experiment:

At  $t \rightarrow -\infty$  prepare state with quantum number  $\alpha$  of  $H_0$

Q: What is amplitude for this state to end up in (another) eigenstate of  $H_0$  with quantum number  $\beta$ .

A:

$$\begin{aligned}\langle \psi_{\beta}^{-} | \psi_{\alpha}^{+} \rangle &= \langle \psi_{\beta}^0 | (\Omega^{-})^{\dagger} \Omega^{+} | \psi_{\alpha}^0 \rangle \\ &\equiv \langle \psi_{\beta}^0 | S | \psi_{\alpha}^0 \rangle\end{aligned}\quad (6.12)$$

$$\begin{aligned}S &= \lim_{\substack{\tau \rightarrow \infty \\ \tau_0 \rightarrow -\infty}} e^{\frac{i}{\hbar} H_0 \tau} e^{\frac{i}{\hbar} H (\tau_0 - \tau)} e^{-\frac{i}{\hbar} H_0 \tau_0} \\ &= \lim_{\substack{\tau \rightarrow \infty \\ \tau_0 \rightarrow -\infty}} U(\tau, \tau_0) \\ &= U(\infty, -\infty)\end{aligned}\quad (6.13)$$

ex 1: sheet 6:

$$\begin{aligned}i\hbar \frac{d}{d\tau} U(\tau, \tau_0) &= e^{\frac{i}{\hbar} H_0 \tau} (H - H_0) e^{\frac{i}{\hbar} H (\tau_0 - \tau)} e^{-\frac{i}{\hbar} H_0 \tau_0} \\ &= \underbrace{e^{\frac{i}{\hbar} H_0 \tau} V e^{-\frac{i}{\hbar} H_0 \tau}}_{V(\tau)} U(\tau, \tau_0) \\ &= V(\tau) U(\tau, \tau_0)\end{aligned}\quad (6.14a)$$

$V(\tau)$  evolution operator in IA picture section 4.4. Solution

$$U(\tau, \tau_0) = T \left( e^{-\frac{i}{\hbar} \int_{\tau_0}^{\tau} V(\tau) d\tau} \right) \quad (6.15a)$$

and

$$\begin{aligned}S &= U(\infty, -\infty) \\ &= \mathbf{1} - \frac{i}{\hbar} \int_{-\infty}^{\infty} dt_1 V(t_1) + \left( -\frac{i}{\hbar} \right)^2 \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{t_1} dt_2 V(t_1) V(t_2) + \dots\end{aligned}\quad (6.15b)$$

Note:

$$SS^{\dagger} = \mathbf{1}, \quad S \text{ operator is unitary} \quad (6.16)$$

$S$ -matrix

$$\begin{aligned}S_{\beta\alpha} &\equiv \langle \psi_{\beta}^0 | S | \psi_{\alpha}^0 \rangle \\ &= \langle \psi_{\beta}^{-} | \psi_{\alpha}^{+} \rangle\end{aligned}\quad (6.17)$$

**option 1** insert  $S$  operator

**1st term:**

$$\langle \psi_\beta^0 | \mathbf{1} | \psi_\alpha^0 \rangle = \delta(\beta - \alpha) \quad (6.18)$$

**2nd term**

$$\begin{aligned} & -\frac{i}{\hbar} \int_{-\infty}^{\infty} dt_1 \langle \psi_\beta^0 | e^{\frac{i}{\hbar} H_0 t} V e^{-\frac{i}{\hbar} H_0 t} | \psi_\alpha^0 \rangle \\ & = -\frac{i}{\hbar} \int dt e^{-\frac{i}{\hbar} (E_\alpha - E_\beta) t} \langle \psi_\beta^0 | V | \psi_\alpha^0 \rangle \\ & = -2i\pi \delta(E_\alpha - E_\beta) V_{\beta\alpha} \end{aligned} \quad (6.19)$$

**3rd and 4th** → exercise sheet 8

$$\begin{aligned} |\psi_\alpha^\pm\rangle &= |\psi_\alpha^0\rangle + G_0^\pm(E) V |\psi_\alpha^\pm\rangle \\ &= |\psi_\alpha^0\rangle + G^\pm(E) V |\psi_\alpha^0\rangle \end{aligned} \quad (6.20a)$$

$$\begin{aligned} S_{\beta\alpha} &= \langle \psi_\beta^- | \psi_\alpha^+ \rangle \\ &= \langle \psi_\beta^0 | S | \psi_\alpha^0 \rangle \end{aligned} \quad (6.20b)$$

### 6.3 structure of $S$ -matrix

→ exercise 3 sheet 8

$$\begin{aligned} |\psi_\alpha^- \rangle - |\psi_\alpha^+ \rangle &= G^- \\ &= (G^-(E_\alpha) - G^+(E_\alpha)) V |\psi_\alpha^0 \rangle \end{aligned} \quad (6.21a)$$

$$\langle \psi_\alpha^- | - \langle \psi_\alpha^+ | = \langle \psi_\alpha^0 | V (G^+(E_\alpha) - G^-(E_\alpha)) \quad (6.21b)$$

$$\begin{aligned} S_{\beta\alpha} &= \langle \psi_\alpha^+ | \psi_\alpha^+ \rangle \\ &= \left( \langle \psi_\beta^+ | + \langle \psi_\beta^0 | V \left( \frac{1}{E_\beta - H + i0^+} - \frac{1}{E_\beta - H - i0^+} \right) \right) |\psi_\alpha^+ \rangle \\ &= \delta(\beta - \alpha) + \underbrace{\left( \frac{1}{E_\beta - E_\alpha + i0^+} - \frac{1}{E_\beta - E_\alpha - i0^+} \right)}_{\bullet} \end{aligned} \quad (6.21c)$$

$$\spadesuit : \lim_{0+ \searrow 0} \frac{-2i0^+}{(E_\beta - E_\alpha)^2 + (0^+)^2} \quad (6.21d)$$

$$= -2i\pi\delta(E_\beta - E_\alpha)$$

$$\delta(x) = \lim_{\varepsilon \searrow x} \frac{1}{\pi} \frac{\varepsilon}{x^2 + \varepsilon^2} \quad (6.21e)$$

**Remark 6.3.1:**

$$\begin{aligned} \sqrt{H} |\phi\rangle &= \sum \sqrt{E_\alpha} c_\alpha |\psi_\alpha\rangle \\ &= \sum c_\alpha \sqrt{E_\alpha} |\psi_\alpha\rangle \end{aligned} \quad (6.22a)$$

with

$$|\phi\rangle = \sum c_\alpha |\psi_\alpha\rangle, \quad (6.22b)$$

$$H |\psi_\alpha\rangle = E_\alpha |\psi_\alpha\rangle, \quad (6.22c)$$

$$\Rightarrow \sqrt{H} |\psi_\alpha\rangle = \sqrt{E_\alpha} |\psi_\alpha\rangle. \quad (6.22d)$$

$$S_{\beta\alpha} = \delta(\beta - \alpha) - 2i\pi\delta(E_\alpha - E_\beta) \underbrace{\langle \psi_\beta^0 | V | \psi_\alpha^+ \rangle}_{T_{\beta\alpha} \text{ transition matrix}} \quad (6.23a)$$



# QUANTIZATION OF RADIATION FIELD

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## 7.1 Quantization of free radiation field

from section 5.1

$$\begin{aligned} \mathbf{A}(\mathbf{r}, t) = \int \frac{d^3\mathbf{k}}{(2\pi)^3} \sum_{\lambda} (\alpha(k, \lambda) \boldsymbol{\epsilon}(k, \lambda) e^{i\mathbf{k} \cdot \mathbf{r} - i\omega t} \\ + \alpha^*(k, \lambda) \boldsymbol{\epsilon}^*(k, \lambda) e^{-i\mathbf{k} \cdot \mathbf{r} + i\omega_k t}), \quad \omega = c|\mathbf{k}| \end{aligned} \quad (7.1a)$$

with  $\lambda \rightarrow 2$  polarizations,  $\boldsymbol{\epsilon}$  polarization vectors with

$$\mathbf{k} \cdot \boldsymbol{\epsilon}(k, \lambda) = 0 \quad (7.1b)$$

consider time dependence of single mode  $(\mathbf{k}, \lambda)$ :

$$q_{k\lambda} = \alpha(k, \lambda) e^{-i\omega_k t} \quad (7.2a)$$

$$\ddot{q}_{k\lambda} = -\omega_k^2 q_{k\lambda} \quad (7.2b)$$

$\rightarrow$  harmonics oscillator

Recall QMI: harmonic oscillator

$$\hat{H} = \frac{\hbar\omega}{2} (\hat{a}^\dagger \hat{a} + \hat{a} \hat{a}^\dagger) \quad (7.3a)$$

states

$$\hat{a} |0\rangle = 0 \quad (7.3b)$$

$$|n\rangle \equiv \frac{(\hat{a}^\dagger)^n}{\sqrt{n!}} |0\rangle \quad (7.3c)$$

The Hilbert space

$$\hat{a}^\dagger |n\rangle = \sqrt{n+1} |n+1\rangle \dots \quad (7.3d)$$

$$\hat{N} = \hat{a}^\dagger \hat{a} \quad (7.3e)$$

with

$$\hat{N} |n\rangle = n |n\rangle, \quad \text{etc.} \quad (7.3f)$$

outlook  $\mathbf{A} \rightarrow \hat{\mathbf{A}}$  (2nd quantization)

To motivate the interpretation of  $\mathbf{A}$  as collection of independent harmonic oscillator compute

$$H = \frac{1}{8\pi} \int d^3\mathbf{r} (\mathbf{E}^2 + \mathbf{B}^2) \quad (7.4a)$$

$$\begin{aligned} \mathbf{E} &= -\frac{1}{c} \dot{\mathbf{A}} \\ &= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \left( \alpha \boldsymbol{\epsilon} \frac{i\omega}{c} e^{-i\omega t + i\mathbf{k} \cdot \mathbf{r}} - \alpha^* \boldsymbol{\epsilon} \frac{i\omega}{c} e^{i\omega t - i\mathbf{k} \cdot \mathbf{r}} \right) \end{aligned} \quad (7.4b)$$

$$\begin{aligned} \int d^3\mathbf{r} \mathbf{E}^2 &= \int d^3\mathbf{r} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \sum_{\lambda\lambda'} \left( \alpha \boldsymbol{\epsilon} e^{-i\omega t + i\mathbf{k} \cdot \mathbf{r}} - \alpha^* \boldsymbol{\epsilon}^* e^{i\omega t - i\mathbf{k} \cdot \mathbf{r}} \right) \left( \frac{i\omega}{c} \right) \\ &\quad \cdot \left( \alpha' \boldsymbol{\epsilon}' e^{i\omega' t + i\mathbf{k}' \cdot \mathbf{r}} - \alpha'^* \boldsymbol{\epsilon}'^* e^{i\omega' t - i\mathbf{k}' \cdot \mathbf{r}} \right) \left( \frac{i\omega'}{c} \right) \\ &= \int d^3\mathbf{r} \iint \sum_{\lambda\lambda'} \left( \alpha \alpha' \boldsymbol{\epsilon} \times \boldsymbol{\epsilon}'^* \frac{-\omega\omega'}{c^2} e^{-it(\omega-\omega')} e^{i\mathbf{r} \cdot (\mathbf{k}-\mathbf{k}')} \right) \end{aligned} \quad (7.4c)$$

use

$$\int d^3\mathbf{r} e^{i\mathbf{r} \cdot (\mathbf{k}-\mathbf{k}')} = (2\pi)^3 \delta(\mathbf{k}-\mathbf{k}') \quad (7.4d)$$

$$\int \frac{d^3k}{(2\pi)^3} \sum_{\lambda\lambda'} \left( \alpha \alpha' \boldsymbol{\epsilon} \times \boldsymbol{\epsilon}'^* \left( \frac{-\omega^2}{c^2} \right) + \dots \right) \quad (7.4e)$$

$$P = \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \hbar \mathbf{k} a_{k\lambda}^* a_{k,\lambda} \quad (7.5a)$$

2nd quantization: so far  $\mathbf{A}$  classical field,  $a^* = f^{\text{cts}}$

$$A, H, P \rightarrow \hat{A}, \hat{H}, \hat{P} \quad (7.6a)$$



commutation relations

$$\begin{aligned} [\hat{a}_{k\lambda}, \hat{a}_{k'\lambda'}] &= [\hat{a}_{k\lambda}^\dagger, \hat{a}_{k'\lambda'}^\dagger] \\ &= 0 \end{aligned} \quad (7.6b)$$

$$[\hat{a}_{k\lambda}, \hat{a}_{k'\lambda'}^\dagger] = (2\pi)^3 \delta(k - k') \delta_{\lambda\lambda'} \quad (7.6c)$$

$k$  **continuous**  $\rightarrow$  world in a box  $\rightarrow k$  discrete; box  $\rightarrow \infty$ ,  $k \rightarrow \text{const}$

$\lambda$ : discrete

another way to quantise field **A**:

$$\begin{aligned} \mathcal{L} &= \frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} \\ &= \frac{1}{8\pi} (E^2 - B^2) \end{aligned} \quad (7.7a)$$

define conjugate momentum field.

$$\begin{aligned} \pi &= \frac{\partial f}{\partial \mathbf{A}} \\ &= \dots \\ &= -\frac{1}{4\pi c} \mathbf{E}^2 \end{aligned} \quad (7.7b)$$

$$A, \pi \rightarrow \hat{A}, \hat{\pi}$$

$$\begin{aligned} [A_i(x, t), A_j(y, t)] &= [\pi_i(x, t), \pi_j(y, t)] \\ &= 0 \end{aligned} \quad (7.8a)$$

$$[A_i, \pi_j] \approx i\hbar \delta(x - y) \delta_{ij} \quad (7.8b)$$

$$\begin{aligned} \mathbf{A}(\mathbf{r}, t) &= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} (\alpha(k, \lambda) \mathbf{e}(k, \lambda) e^{i\mathbf{k}\cdot\mathbf{r} - i\omega t} \\ &\quad + \alpha^*(\mathbf{k}, \lambda) \mathbf{e}^*(\mathbf{k}, \lambda) e^{-i\mathbf{k}\cdot\mathbf{r} + i\omega t}) \end{aligned} \quad (7.9a)$$

$$H = \frac{1}{8\pi} \int d^3\mathbf{r} (\mathbf{E}^2 + \mathbf{B}^2) \quad (7.9b)$$

$$\begin{aligned}
\int d^3\mathbf{r} \mathbf{E}^2 &= \int d^3\mathbf{r} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \sum_{\lambda\lambda'} \left( \alpha\alpha' \boldsymbol{\epsilon} \right. \\
&\quad \cdot \boldsymbol{\epsilon} \left( -\frac{\omega\omega'}{c^2} \right) e^{-i(\omega+\omega')t} e^{i(\mathbf{k}+\mathbf{k}')\cdot\mathbf{r}} \\
&\quad + \alpha^* \alpha'^* \boldsymbol{\epsilon}^* \boldsymbol{\epsilon}'^* \left( -\frac{\omega\omega'}{c^2} e^{+i(\omega+\omega')t} e^{-i(\mathbf{k}+\mathbf{k}')\cdot\mathbf{r}} \right) \\
&\quad - \alpha\alpha'^* \boldsymbol{\epsilon}^* \boldsymbol{\epsilon}'^* \left( -\frac{\omega\omega'}{c^2} e^{-i(\omega+\omega')t} e^{-i(\mathbf{k}-\mathbf{k}')\cdot\mathbf{r}} \right) \\
&\quad \left. - \alpha^* \alpha' \boldsymbol{\epsilon}^* \boldsymbol{\epsilon}'^* \left( -\frac{\omega\omega'}{c^2} e^{+i(\omega+\omega')t} e^{+i(\mathbf{k}+\mathbf{k}')\cdot\mathbf{r}} \right) \right) \\
&= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda\lambda'} \left( -\frac{\omega^2}{c^2} \right) \left( \alpha_k \alpha_{-k} \boldsymbol{\epsilon}_k \cdot \boldsymbol{\epsilon}_{-k} e^{-2i\omega t} + \alpha_k^* \alpha_{-k}^* \boldsymbol{\epsilon}_k^* \right. \\
&\quad \cdot \boldsymbol{\epsilon}_{-k}^* e^{2i\omega t} - \alpha_k \alpha_{+k}^* \boldsymbol{\epsilon}_k \cdot \boldsymbol{\epsilon}_k^* - \alpha_k^* \alpha_k \boldsymbol{\epsilon}_k^* \cdot \boldsymbol{\epsilon}_k \\
&\quad \left. + \{ \alpha_k \alpha_{-k}, \alpha_k^* \alpha_{-k}^* \text{ terms} \} \right) \\
&= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \left( \frac{\omega^2}{c^2} \right) (\alpha_k \alpha_k^* + \alpha_k^* \alpha_k) \\
&\quad + \{ \alpha_k \alpha_{-k}, \alpha_k^* \alpha_{-k}^* \text{ terms} \} \Big)
\end{aligned}$$

use

$$\int d^3r e^{i(k \pm k')r} = (2\pi)^3 \delta(k \pm k') \quad (7.9c)$$

$\lambda, k'$  integration and  $\omega_{-k} = \omega_k = \omega$

$$\begin{aligned}
\rightsquigarrow \int d^3v \mathbf{E}^2 &= \frac{1}{(2\pi)^3} \int d^3k \left( -\frac{\omega^2}{c^2} \right) (\alpha_k \alpha_{-k} \boldsymbol{\epsilon}_k \boldsymbol{\epsilon}_{-k} e^{-2i\omega t} \\
&\quad + \alpha_k^* \alpha_{-k}^* \boldsymbol{\epsilon}_k^* \boldsymbol{\epsilon}_{-k}^* e^{2i\omega t} - \alpha_k \alpha_k^* \boldsymbol{\epsilon}_k \boldsymbol{\epsilon}_k^* - \alpha_k^* \alpha_k \boldsymbol{\epsilon}_k^* \boldsymbol{\epsilon}_k)
\end{aligned} \quad (7.9d)$$

$$\int d^3\mathbf{r} \mathbf{B}^2 = \dots \quad (7.9e)$$

$$= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \left( \frac{\omega^2}{c^2} \right) (\alpha_k \alpha_k^* + \alpha_k^* \alpha_k - \{\dots\})$$

$$H = \int \frac{d^3\mathbf{k}}{(2\pi)^3} \sum_{\lambda} \frac{\omega^2}{4\pi c^2} (\alpha_k \alpha_k^* + \alpha_k^* \alpha_k) \quad (7.9f)$$

define:

$$\alpha^*(k, \lambda) = \sqrt{\frac{\omega}{2\pi c^2 \hbar}} \alpha^*(k, \lambda) \quad (7.10a)$$

$$\mathbf{A}(\mathbf{r}, t) = \int \frac{d^3k}{(2\pi)^3} \sqrt{\frac{2\pi c^2 \hbar}{\omega}} \sum_{\lambda} (q_{k\lambda} \mathbf{e}_{k\lambda} e^{i\mathbf{k} \cdot \mathbf{r} - i\omega t} + q_{k\lambda}^* \mathbf{e}_{k\lambda}^* e^{-i\mathbf{k} \cdot \mathbf{r} + i\omega t}) \quad (7.10b)$$

$$H = \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \frac{\hbar\omega}{2} (a_{k\lambda}^* a_{k\lambda} + a_{k\lambda} a_{k\lambda}^*) \quad (7.10c)$$

$$\mathbf{P} = \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \hbar \mathbf{k} a_{k\lambda}^* a_{k\lambda} \quad (7.10d)$$

### 7.1.1 2nd quantization

So far  $\mathbf{A}$  classical field,  $a^{(*)}$  functions

$$a(k, \lambda) \rightarrow \hat{a}(k, \lambda), \quad \text{operator} \quad (7.11a)$$

$$a^*(k, \lambda) \rightarrow \hat{a}^\dagger(k, \lambda), \quad \text{operator} \quad (7.11b)$$

as for harmonic oscillator

$$\mathbf{A}, H, \mathbf{P} \rightarrow \hat{\mathbf{A}}, \hat{H}, \hat{\mathbf{P}} \quad (7.11c)$$

commutation relations:

$$[\hat{a}_{k\lambda}, \hat{a}_{k'\lambda'}] = [\hat{a}_{\dagger}^{k\lambda}, \hat{a}_{\dagger}^{k'\lambda'}] = 0 \quad (7.12a)$$

$$[\hat{a}_{k\lambda}, \hat{a}_{k'\lambda'}^\dagger] = (2\pi)^3 \delta(\mathbf{k}, -\mathbf{k}') \delta_{\lambda\lambda'} \quad (7.12b)$$

$\mathbf{k}$ : continuous label  $\rightarrow$  ften system in a box,  $k$  becomes discrete

$\lambda$ : discrete, l.z

another way to quantize field  $\mathbf{A}$

$$\begin{aligned} \mathcal{L} &= \frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} \\ &= \frac{1}{8\pi} (\mathbf{E}^2 - \mathbf{B}^2) \end{aligned} \quad (7.13a)$$

define conjugate momentum field

$$\begin{aligned}\pi &= \frac{\partial f}{\partial \mathbf{A}} \\ &= \dots \\ &= -\frac{1}{4\pi c} \mathbf{E}^2\end{aligned}\quad (7.13b)$$

$$A, \pi \rightarrow \hat{A}, \hat{\pi}$$

$$\begin{aligned}[A_i(x, t), A_j(y, t)] &= [\pi_i(x, t), \pi_j(y, t)] \\ &= 0\end{aligned}\quad (7.13c)$$

$$[A_i, \pi_j] \approx i\hbar \delta(x - y) \delta_{ij} \quad (7.13d)$$

## 7.2 Fock space

Built up as for single harmonic oscillator with “ladder” operators

$$\hat{a}^\dagger(k, \lambda) = \hat{a}_{k\lambda}^\dagger, \quad \text{creation operator} \quad (7.14a)$$

$$\hat{a}(k, \lambda) = \hat{a}_{k\lambda}, \quad \text{annihilation operator} \quad (7.14b)$$

start with vacuum  $|0\rangle$  definition

$$\hat{a}_{k\lambda} |0\rangle = 0 \quad (7.14c)$$

$$|1(k, \lambda)\rangle = \hat{a}^\dagger |0\rangle \quad (7.14d)$$

state with 1 photon momentum,  $\mathbf{k}$ , polarization  $\lambda$

**General state** Note:

$$\begin{aligned}\hat{a}_{k_j \lambda_j}^\dagger |n_1(k_1, \lambda_1) \dots n_j(k_j, \lambda_j) \dots n_m(k_m, \lambda_m)\rangle \\ = \sqrt{n_j + 1} |n_1 \dots n_{j+1} \dots n_m\rangle\end{aligned}\quad (7.15a)$$

$$\int \frac{d^3 \mathbf{k}}{(2\pi)^3} \sum_\lambda \hat{a}_{k\lambda} |n_1(k_1, \lambda_1)\rangle = \frac{1}{\sqrt{n_1!}} \int \frac{d^3 k}{(2\pi)^3} \sum_\lambda \hat{a}_{k\lambda} \left(\hat{a}_{k_1 \lambda_1}^\dagger\right)^{n_1} |0\rangle \quad (7.15b)$$

with

$$\hat{a}_{k\lambda} \left(\hat{a}_{k_1 \lambda_1}^\dagger\right)^{n_1} |0\rangle = \left(\hat{a}_{k_1 \lambda_1}^\dagger\right)^{n_1-1} n_1 \left[\hat{a}_{k\lambda}, \hat{a}_{k_1 \lambda_1}^\dagger\right] + \left(\hat{a}_{k_1 \lambda_1}^\dagger\right) \hat{a}_{k\lambda} \quad (7.15c)$$

with

$$[\hat{a}_{k\lambda}, \hat{a}_{k_1\lambda_1}^\dagger] = n_1 (2\pi)^3 \delta(\mathbf{k} - \mathbf{k}_1) \delta_{\lambda\lambda_1} \quad (7.15d)$$

$$\begin{aligned} \rightsquigarrow \int \frac{d^3\mathbf{k}}{(2\pi)^3} \sum_{\lambda} \hat{a}_{k\lambda} |n_1(k_1, \lambda_1)\rangle &= \frac{n_1}{\sqrt{n_1!}} \left( \hat{a}_{k_1\lambda_1}^\dagger \right)^{n_1-1} |0\rangle \\ &= \frac{n_1 \sqrt{(n_1-1)!}}{\sqrt{n_1!}} \frac{(\hat{a}_{k_1\lambda_1}^\dagger)^{n_1-1}}{\sqrt{(n_1-1)!}} |0\rangle \\ &= \sqrt{n_1} |(n_1-1)(k_1\lambda_1)\rangle \end{aligned} \quad (7.15e)$$

General:

$$\int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} a_{k\lambda} |n_1 \dots n_m\rangle = \sum_{i=1}^m \sqrt{n_i} |n_1 \dots n_{i-1}, \dots n_m\rangle \quad (7.16a)$$

Compute expectation value of  $\hat{H}$  in state  $|n_1(k_1\lambda_1)\rangle$

$$\langle n_1 | \hat{H} | n_1 \rangle = \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \frac{\hbar\omega}{2} \langle n_1 | \hat{a}_{k\lambda}^\dagger \hat{a}_{k\lambda} + \underbrace{\hat{a}_{k\lambda} \hat{a}_{k\lambda}^\dagger}_{\delta(\mathbf{0})\delta(\mathbf{0})} | n_1 \rangle$$

$$\begin{aligned} \langle n_1 | \hat{H} | n_1 \rangle &= \int \frac{d^3k}{(2\pi)^3} \sum_{\lambda} \hbar\omega \langle n_1 | \hat{a}_{k\lambda}^\dagger \hat{a}_{k\lambda} | n_1 \rangle \\ &= \hbar\omega \langle n_1 | \hat{a}_{k_1\lambda_1}^\dagger | n_1 - 1 \rangle \sqrt{n_1} \\ &= \hbar\omega n_1 \langle n_1 | n_1 \rangle \\ &= n_1 \hbar\omega \end{aligned} \quad (7.16c)$$

$$\rightarrow \langle n_1 \dots n_m | \hat{H} | n_1 \dots n_m \rangle = \sum_{i=1}^m \hbar\omega_i n_i \quad (7.16d)$$

introduce interactions with matter (compare Section 5)

$$V = -\frac{q}{2mc} (\mathbf{p} \cdot \mathbf{A} + \mathbf{A} \cdot \mathbf{p}) + \frac{q^2}{2mc^2} \hat{\mathbf{A}}^2 \quad (7.17a)$$

in Coulomb gauge

$$\nabla \cdot \mathbf{A} = 0 \quad (7.18a)$$

$$\Rightarrow \mathbf{p} \cdot \mathbf{A} = \mathbf{A} \cdot \mathbf{p} \quad (7.18b)$$

$$\hat{V} = -\frac{q}{mc} \mathbf{p} \cdot \hat{\mathbf{A}} + \frac{q^2}{2mc^2} \hat{A}^2 \quad (7.18c)$$

is operator in Fock space and will be used to compute transition matrix elements

$$V_{\beta\alpha} = \langle \psi_\beta^0 | V | \psi_\alpha^0 \rangle \quad (7.18d)$$

### 7.3 Photon emission and absorption

We have considered these processes before in section 5.2

$$\begin{aligned} H &= H_0 + H_{\text{em}} + \hat{V} \\ &= \sum_i \frac{p_i^2}{2m} + \frac{1}{8\pi} \int d^3r \left( \hat{\mathbf{E}}^2 + \hat{\mathbf{B}}^2 \right) \\ &\quad + \text{interaction matter} \leftrightarrow \text{photons} \end{aligned} \quad (7.19a)$$

$$\begin{aligned} \text{eigenstates } |\psi_\alpha^0\rangle &= |\text{matter}\rangle \otimes |\text{photons}\rangle \\ &= \text{wave function} \otimes \text{new, Fockspace} \\ &= |A; n_1(k, \lambda_1) \dots n_m(k_m, \lambda_m)\rangle \end{aligned} \quad (7.19b)$$

#### 7.3.1 Absorption of photon

$$V_{\beta\alpha} = \langle B; (n-1)(k, \lambda) | \hat{V} | A; n(k, \lambda) \rangle \quad (7.20a)$$

with

$B$ : final state of atom

$(n-1)$ : one photon “lost”

$A$ : initial state of atom

$n(k, \lambda)$ :  $n$  photons,  $k, \lambda$

$$\begin{aligned} \rightsquigarrow V_{\beta\alpha} &= \frac{e}{mc} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \sum_{\lambda'} \sqrt{\frac{2\pi\hbar c^2}{\omega'}} \\ &\quad \cdot \langle B; (n-1)(k, \lambda) | \hat{a}(k', \lambda') \mathbf{p} \cdot \boldsymbol{\epsilon}(k', \lambda') e^{i\mathbf{k}' \cdot \mathbf{r}} | A, n(k, \lambda) \rangle \end{aligned} \quad (7.20b)$$

$\hat{a}^\dagger$  term gives no contribution

$$\begin{aligned} \rightsquigarrow V_{\beta\alpha} &= \frac{e}{m} \sqrt{\frac{2\pi\hbar}{\omega}} \sqrt{n} \langle B; | \mathbf{p} \cdot \boldsymbol{\epsilon}(k, \lambda) e^{i\mathbf{k} \cdot \mathbf{r}} | A \rangle \\ &\sim \sqrt{n} \end{aligned} \quad (7.20c)$$

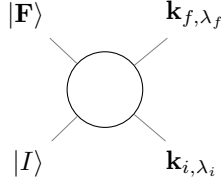


Figure 7.1:

---

### 7.3.2 Emission of photons

$$V_{\beta\alpha} = \langle B; (n+1) (k, \lambda) | \hat{V} | A; n (k, \lambda) \rangle, \quad (7.21a)$$

with  $\hat{V}$  now only  $\hat{a}^\dagger$  part contribution.

$$\begin{aligned} \rightsquigarrow V_{\beta\alpha} &= \frac{e}{m} \sqrt{\frac{2\pi\hbar}{\omega}} \sqrt{n+1} \langle B | \mathbf{p} \cdot \boldsymbol{\epsilon}^* (k, \lambda) e^{-i\mathbf{k} \cdot \mathbf{r}} | A \rangle \\ &\sim \sqrt{n+1} \end{aligned} \quad (7.21b)$$

Note: This is non-zero even for  $n = 0 \rightarrow$  spontaneous emission!  
recall classical:

$$\Gamma_{n0} = \Gamma_{0n} \quad (7.22a)$$

$$\text{absorption} = \text{emission} \quad (7.22b)$$

now

$$\frac{\Gamma_{n0}}{\Gamma_{0n}} = \frac{n_{k\lambda}}{n_{k\lambda} + 1} \quad (7.22c)$$

---

## 7.4 Scattering of photons by atoms

involves creation and annihilation of photon need  $\hat{a}^\dagger(k_f, \lambda_f) \hat{a}(k_i, \lambda_i)$  recall

$$\hat{V} = \frac{e}{mc} \mathbf{p} \cdot \hat{\mathbf{A}} + \frac{e^2}{2mc^2} \hat{\mathbf{A}}^2 \quad (7.23)$$

where  $\hat{\mathbf{A}}$  contains either  $\hat{a}^\dagger$  or  $\hat{a}$  and contributes only at 2nd order,  $\hat{\mathbf{A}}^2$  contains  $\hat{a}^\dagger \hat{a} \rightarrow$  contributes at first order. Both  $\hat{\mathbf{A}}$  and  $\hat{\mathbf{A}}^2$  are  $\sim \frac{e^2}{c^2}$

**first-order contribution**

$$\begin{aligned}
V_{\beta\alpha}^{(1)} &= \langle F, 1(k_f, \lambda_f) | \frac{e^2}{2mc^2} \hat{\mathbf{A}}^2 | I, 1(k_i, f_i) \rangle \\
&= \frac{e^2}{2mc^2} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \sum_{\lambda\lambda'} \frac{2\pi\hbar c^2}{\sqrt{\omega\omega'}} \\
&\quad \cdot \langle F, 1(k_f, \lambda_f) | (\hat{a}\epsilon e^{i\mathbf{k}\cdot\mathbf{r}-i\omega t} + \hat{a}^\dagger(k, \lambda) e^{-i\mathbf{k}\cdot\mathbf{r}+i\omega t}) \\
&\quad \cdot (\hat{a}'(k', \lambda') e^{i\mathbf{k}'\cdot\mathbf{r}-i\omega' t} + \hat{a}'^\dagger \epsilon'^* e^{-i\mathbf{k}'\cdot\mathbf{r}+i'\omega' t}) | I, 1(k_i, \lambda_i) \rangle \\
&= \frac{e^2}{2mc^2} \frac{2\pi\hbar c^2}{\sqrt{\omega_i\omega_f}} 2 \cdot \langle F | \epsilon(k_i, \lambda_i \cdot \epsilon(k_f, \lambda_f) e^{it(\omega_i-\omega_f)}) e^{i\mathbf{r}\cdot(\mathbf{k}_i-\mathbf{k}_f)} | I \rangle
\end{aligned} \tag{7.24}$$

Recall from section 7.2

$$H = H_0 + V \tag{7.25a}$$

$$\begin{aligned}
S_{\beta\alpha} &= \delta(\beta - \alpha) - \frac{i}{\hbar} \int_{\mathbb{R}} dt_1 e^{-\frac{i}{\hbar} t_1 (E_\alpha - E_\beta)} \langle \psi_\beta^0 | V | \psi_\alpha^0 \rangle \\
&\quad - \frac{i}{\hbar} \int_{-\infty}^{\infty} dt_1 e^{-\frac{i}{\hbar} t_1 (E_I + \hbar\omega_i - E_F - \hbar\omega_f)} \left( \frac{2e^2\hbar\pi}{m\sqrt{\omega_i\omega_f}} \epsilon \right. \\
&\quad \left. \cdot \epsilon^* \langle F | \dots | F \rangle \right)
\end{aligned} \tag{7.25b}$$

golden rule (sec 4.2/4.3) Transition rate  $\frac{2\pi}{\hbar} |T|^2 \rho$  cross section  $\frac{d\sigma}{d\Omega}$  (final photon energy  $\hbar\omega_f + a(\hbar\omega_f)$ )

- divide by flux  $v = c$  (section 6.1)
- nr. states

$$\begin{aligned}
\frac{d^3\mathbf{k}}{(2\pi)^3} &= \frac{k^2 dk d\Omega}{(2\pi)^3} \\
&= \frac{\omega_f^2 d\Omega}{(2\pi)^3 c^3 \hbar} d(\hbar\omega_f)
\end{aligned} \tag{7.26a}$$

$$\rho(\omega_f) = \frac{\omega_f^2 d\Omega}{(2\pi)^3 c^3 \hbar} \tag{7.26b}$$



## Feynman diagrams

Figure 7.2:

$$\frac{d\sigma}{d\Omega} = \frac{1}{c} \frac{2\pi}{\hbar} \underbrace{\frac{\omega_f^2}{(2\pi)^3 c^3 \hbar}}_{\rho(\omega_f)} |T|^2 \quad (7.27a)$$

$$= \frac{e^4}{m^2 c^4} \frac{\omega_f}{\omega_i} \left| \mathbf{e}_i \cdot \mathbf{e}_f^* \langle F | e^{i\mathbf{r} \cdot (\mathbf{k}_i - \mathbf{k}_f)} | I \rangle \right|^2$$

$$\left( \frac{\alpha \hbar}{mc} \right)^2 = r_0^2 \quad (7.27b)$$

(classical electron radius)

However, there are further contributions of order  $\alpha^2 \sim e^4$  golden rule (sec 4.2/4.3) Transition rate  $\frac{2\pi}{\hbar} |T|^2 \rho$

cross section  $\frac{d\sigma}{d\Omega}$  (final photon)...

### 7.4.1 2nd order contribution

for  $T_{\beta\alpha}$

$$S_{\beta\alpha} = \dots \left( -\frac{i}{\hbar} \right)^2 \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{t_1} dt_2 \sum_{\gamma} e^{-\frac{i}{\hbar} t_1 (E_{\gamma} - E_{\beta})} e^{-\frac{i}{\hbar} t_2 (E_{\alpha} - E_{\gamma})} V_{\beta\gamma} V_{\gamma\alpha} \quad (7.28a)$$

In our case

$$\begin{aligned} & \left( -\frac{i}{\hbar} \right)^2 \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{t_1} dt_2 \sum_{\lambda\lambda'} dN e^{-\frac{i}{\hbar} (E_N - E_F) t_i} e^{-\frac{i}{\hbar} (E_I - E_N) t_2} \left( \frac{e}{mc} \right)^2 \\ & \cdot \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int \frac{d^3 \mathbf{k}'}{(2\pi)^3} \sum_{\lambda\lambda'} \langle F, P(\omega_f, \lambda_f) | \hat{a}_k \mathbf{p} \cdot \mathbf{e}_k e^{i\mathbf{k} \cdot \mathbf{r} - i\omega t_1 + \hat{a}_k^{\dagger} \mathbf{p} \cdot \mathbf{e}_k^* e^{-i\mathbf{k} \cdot \mathbf{r} + i\omega t_1}} | N \rangle \\ & \cdot \langle N | \hat{a}_{k'} \mathbf{p} \cdot \mathbf{e}_{k'} e^{i\mathbf{k}' \cdot \mathbf{r} - i\omega' t_2} + \hat{a}_{k'}^{\dagger} \mathbf{p} \cdot \mathbf{e}_{k'}^* e^{-i\mathbf{k}' \cdot \mathbf{r} + i\omega' t_2} | I, 1(k_i, \lambda_i) \rangle \end{aligned} \quad (7.29a)$$

need 1  $\hat{a}(k_i, \lambda_i)$  and one  $\hat{a}^\dagger(k_f, \lambda_f)$

$$= \dots \quad (7.29b)$$

$$= \langle F | \hat{a}_{k_f}^\dagger \dots | N \rangle \langle N | \dots \hat{a}_{k_i} | I \rangle + \langle F | \hat{a}_{k_i} \dots | N \rangle \langle N | \dots \hat{a}_{k_f}^\dagger | I \rangle$$

Kramers Heisenberg formula

$$\frac{d\sigma}{d\Omega} = \left( \frac{\alpha \hbar}{mc} \right)^2 \frac{\omega_f}{\omega_i} \left| \boldsymbol{\epsilon}_i \boldsymbol{\epsilon}_f^* \delta_{FI} + \sum_N \frac{\langle F | \mathbf{p} \cdot \boldsymbol{\epsilon}_f^* | N \rangle \langle N | \mathbf{p} \cdot \boldsymbol{\epsilon}_i | I \rangle}{m(E_I - \hbar\omega_f - E_N)} + \frac{\langle F | \mathbf{p} \cdot \boldsymbol{\epsilon}_i | N \rangle \langle N | \mathbf{p} \cdot \boldsymbol{\epsilon}_f^* | I \rangle}{m(E_I - \hbar\omega_f - E_N)} \right| \quad (7.30)$$

→ limiting cases

**Rayleigh scattering** elastic scattering

$$|I\rangle = |F\rangle, \quad (7.31a)$$

$$\omega_i = \omega_f \quad (7.31b)$$

$$\hbar\omega \ll E_I - E_N \quad (7.31c)$$

combine  $\boldsymbol{\epsilon}_i \cdot \boldsymbol{\epsilon}_f$  with other terms

$$\begin{aligned} \langle I | \boldsymbol{\epsilon}_i \boldsymbol{\epsilon}_f^* | I \rangle &= \frac{1}{i\hbar} \langle I | [\mathbf{x} \cdot \boldsymbol{\epsilon}_i, \mathbf{p} \cdot \boldsymbol{\epsilon}_f^*] | I \rangle \\ &= \frac{1}{i\hbar} \sum_N (\langle I | \mathbf{x} \cdot \boldsymbol{\epsilon}_i | N \rangle \langle N | \mathbf{p} \cdot \boldsymbol{\epsilon}_f^* | I \rangle \\ &\quad - \langle I | \mathbf{p} \cdot \boldsymbol{\epsilon}_f^* | N \rangle \langle N | \mathbf{x} \cdot \boldsymbol{\epsilon}_i | I \rangle) \end{aligned} \quad (7.32a)$$

→ put everything together:

$$\frac{1}{E_N - E_I} + \frac{1}{E_I - E_N + \pm \hbar\omega} = \mp \frac{\hbar\omega}{(E_I - E_N)^2} + \frac{(\hbar\omega)^2}{(E_I - E_N)^2} + \dots \quad (7.33a)$$

$$\frac{d\sigma}{d\Omega} = r_0^2 \frac{(\hbar\omega)^4}{m^2} \left| \sum_N \frac{\langle I | \mathbf{p} \cdot \boldsymbol{\epsilon}_f^* | N \rangle \langle N | \mathbf{p} \cdot \boldsymbol{\epsilon}_i | I \rangle}{(E_I - E_N)^2} + \leftrightarrow \right|^2$$

$r_0^2$  classical e-radius,  $\omega^4$  blue sky red sunset

$$\hbar\omega_i \gg E_N - E_I \quad (7.34a)$$

(large compared to binding energy  $\rightsquigarrow$  scattering off “free” electrons)

$$\frac{d\sigma}{d\Omega} = r^2 |\boldsymbol{\epsilon}_i \cdot \boldsymbol{\epsilon}_f|^2 \quad (7.34b)$$

for unpolarized photons

$$\begin{aligned} & \frac{1}{2} \sum_{\lambda_i \lambda_f} \varepsilon_i^a(k_i, \lambda_i) (\varepsilon_f)^a(k_f, \lambda_f) \varepsilon_i^b(\varepsilon_f^*)^b \\ &= \frac{1}{2} \left( \delta_{ab} - \frac{k_i^a k_i^b}{k_i^2} \right) \left( \delta_{ab} - \frac{k_f^a k_f^b}{k_f^2} \right) \\ &= \frac{1}{2} (1 + \cos^2 \theta) \end{aligned} \quad (7.34c)$$

with

$$\theta = \angle(\mathbf{k}_i, \mathbf{k}_f) \quad (7.34d)$$

$$\begin{aligned} \sigma &= \int d\cos\theta \, 2\pi r_0^2 (1 + \cos^2 \theta) \\ &= \frac{8\pi}{3} r_0^2 \end{aligned} \quad (7.34e)$$

**Resonances** What if  $E_N \sim E_I + \hbar\omega_i$

so far : neglected finite lifetime of  $E_N$ ,  $\tau_N = \frac{\hbar}{\Gamma_N}$   
time evolution

$$e^{-\frac{i}{\hbar} E_N t} e^{-\tau_N t} = e^{-\frac{i}{\hbar} t(E_N - i\Gamma_N)} \quad (7.35a)$$

with  $\Gamma_N$  that one cannot neglect

$$\begin{aligned} |T|^2 &\approx \left| \frac{1}{E_I - E_N + i\Gamma_N} \right|^2 \\ &= \frac{1}{(E_I - E_N)^2 + \Gamma_N^2} \end{aligned} \quad (7.35b)$$



# RELATIVISTIC QM

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Will try to find relativistic generalization of Schrödinger as single-particle equation ( $\rightarrow$  we will fail) but will be basis of relativistic (2nd quantized field theory) Rel: Cannot fix number of particles

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## 8.1 Klein-Gordon equation (KGE)

Consider free scalar particle

$$\begin{aligned} X^\mu &\rightarrow X^{\mu'} \\ &= \Lambda^\mu_{\nu'} x^\nu \end{aligned} \tag{8.1a}$$

$$\begin{aligned} \phi(x) &\rightarrow \phi'(x') \\ &= \phi(x) \\ &= \phi(\Lambda^{-1}x') \end{aligned} \tag{8.1b}$$

Now start from

$$E^2 = m^2 c^4 + \mathbf{p}^2 c^2 \tag{8.2a}$$

(not  $E = \frac{p^2}{2m}$ )

$$E \rightarrow i\hbar \frac{\partial}{\partial t} \tag{8.2b}$$

$$\mathbf{p} = -i\hbar \nabla \tag{8.2c}$$

$$\left( -\hbar^2 \frac{\partial^2}{\partial t^2} \right) \phi(\mathbf{x}, t) = (m^2 c^4 - \hbar c^2 \nabla^2) \phi(x, t) \tag{8.2d}$$

in covariant form:

$$\partial_\mu \partial^\mu + \frac{m^2 c^2}{\hbar^2} \phi(x) = 0 \quad (8.2e)$$

$$\begin{aligned} \partial_\mu &= \frac{\partial}{\partial x^\mu} \\ &= \left( \frac{1}{c} \frac{\partial}{\partial t}, \nabla \right) \end{aligned} \quad (8.2f)$$

$$\partial^\mu = \left( \frac{1}{c} \frac{\partial}{\partial t} - \nabla \right) \quad (8.2g)$$

## 8.2 Klein-Gordon equation

KGE:

$$\left( \partial_\mu \partial^\mu + \frac{m^2 c^2}{\hbar^2} \right) \phi = 0 \quad (8.3a)$$

$$\partial_\mu = \left( \frac{1}{c} \frac{\partial}{\partial t}, \nabla \right) \quad (8.3b)$$

Solution

$$\phi(t, \mathbf{x}) = A \cdot e^{i\mathbf{k} \cdot \mathbf{x} - i\omega t} \quad (8.3c)$$

$$\phi(x) = A \cdot e^{-ik_\mu x^\mu} \quad (8.3d)$$

$$k^\mu = \left( \frac{\omega}{c}, \mathbf{k} \right), \quad k^2 = k_\mu k^\mu = \frac{\omega^2}{c^2} - \mathbf{k}^2 = \frac{m^2 c^2}{\hbar^2} \quad (8.3e)$$

or

$$(h\omega) = \pm \sqrt{m^2 c^4 + \mathbf{p}^2 c^2} \quad (8.3f)$$

Negative solutions?!

In analogy to Schrödinger, try to define probability density  $\rho(\mathbf{x}, t)$  and probability current density  $\mathbf{j}(\mathbf{x}, t)$  satisfying

$$\frac{\partial}{\partial t} \rho + \nabla \cdot \mathbf{j} = 0 \quad (8.4a)$$

$$\mathbf{j} = \frac{\hbar}{2mi} (\phi^* (\nabla \phi) - (\nabla \phi^*) \phi) \quad (8.4b)$$

$$\rho = \frac{i\hbar}{2mc^2} \left( \phi^* \frac{\partial \phi}{\partial t} - \frac{\partial \phi^*}{\partial t} \phi \right) \quad (8.4c)$$

covariant form

$$\begin{aligned} j^\mu &= (cp, j) \\ &= \frac{i\hbar}{2m} (\phi^* \partial^\mu \phi - (\partial^\mu \phi^*) \phi) \end{aligned} \quad (8.4d)$$

Note that  $\rho(\mathbf{x}, t)$  is *not* positive definite  $\rightarrow$  cannot be interpreted as probability density

### 8.3 Dirac equation

Try a linear (in  $\frac{\partial}{\partial t}, \nabla$ ) equation. Most general linear equation.

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} \psi &= H\psi \\ &= (-i\hbar c \boldsymbol{\alpha} \cdot \nabla + \beta mc^2) \psi \\ &= (-i\hbar c \alpha_i \nabla_i + \beta mc^2) \psi \end{aligned} \quad (8.5a)$$

With summation convention and

$$\boldsymbol{\alpha} = (\alpha_1, \alpha_2, \alpha_3) \quad (8.5b)$$

and  $\beta$  are 4 non-commuting coefficients

Iterate this equation

$$\begin{aligned} -\hbar^2 \frac{\partial^2}{\partial t^2} \psi &= (-i\hbar c \alpha_i \nabla_i + \beta mc^2) (-i\hbar c \alpha_j \nabla_j + \beta mc^2) \psi \\ &= \left( c^2 \frac{\hbar^2}{2} (\alpha_i \alpha_j + \alpha_j \alpha_i) \nabla_i \nabla_j - i\hbar (\alpha_i \beta + \beta \alpha_i) \nabla_i mc^2 \right. \\ &\quad \left. + \beta m^2 c^4 \right) \psi \end{aligned}$$

Compare to KGE

$$-\hbar^2 \frac{\partial^2}{\partial t^2} \psi = (-\hbar c^2 \nabla_i \nabla_i + m^2 c^4) \psi \quad (8.6a)$$

$$\beta^2 = 1 \quad (8.6b)$$

$$\begin{aligned} (\alpha_i \beta + \beta \alpha_i) &= \{\alpha_i, \beta\} \\ &= 0, \quad (\text{sometimes } [\alpha_i, \beta]_+) \end{aligned} \quad (8.6c)$$

$$\begin{aligned} (\alpha_i \alpha_j + \alpha_j \alpha_i) &= \{\alpha_i, \alpha_j\} \\ &= 2\delta_{ij} \end{aligned} \quad (8.6d)$$

From anticommutation relations we see that coeff. cannot be “numbers”. ( $\rightarrow$  Exercise dim 4 matrices are simplest possibility)

$\rightarrow$  wave function

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix} \quad (8.7)$$

and one possible choice for  $\alpha$  and  $\beta$ .

$$\alpha_i = \begin{pmatrix} 0 & \sigma_i \\ \sigma_i & 0 \end{pmatrix} \quad (8.8a)$$

$$\begin{aligned} \beta &= \begin{pmatrix} \mathbf{1} & 0 \\ 0 & -\mathbf{1} \end{pmatrix} \\ &= \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}, \quad \text{Dirac representation} \end{aligned} \quad (8.8b)$$

Rewrite Dirac equation in terms of  $\gamma$  matrices ( $4 \times 4$  again)

$$\gamma^\mu, \quad \mu \in \{0, 1, 2, 3\}, \gamma^0 = \beta, \gamma^i = \beta\alpha^i \quad (8.9a)$$

in Dirac representation

$$\gamma^0 = \begin{pmatrix} \mathbf{1} & 0 \\ 0 & -\mathbf{1} \end{pmatrix} \quad (8.9b)$$

$$\gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \quad (8.9c)$$

Note can use any other representation:

$$\gamma^\mu \rightarrow U \gamma^\mu U^\dagger \quad (8.10a)$$

Take  $\beta$ : Dirac equation

$$\beta \cdot \left( i\hbar \frac{\partial}{\partial t} \right) \psi = \beta (-i\hbar c \alpha_i \nabla_i + \beta m c^2) \psi \quad (8.10b)$$

$$i\hbar \frac{\partial}{\partial t} \gamma^0 \psi = (-i\hbar c \gamma^i \nabla_i + m c^2) \psi \quad (8.10c)$$

$$(i\hbar \partial_\mu - mc) \psi = 0 \quad (8.10d)$$



Notation any 4-vector  $a^\mu$ :

$$\not{a} \equiv a_\mu \gamma^\mu \quad (8.10e)$$

( $\phi$  and  $\not{\partial}$  have been mistaken...)

$$(i\hbar \not{\partial} - mc) \psi = 0 \quad (8.10f)$$

Properties of  $\gamma$  matrices:

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu} \quad (8.10g)$$

$$\gamma^0 \gamma^\mu \gamma^0 = (\gamma^\mu)^\dagger \quad (8.10h)$$

(There is an old notation in Sakurai.) Dirac equation:

$$(i\hbar \partial_\mu \gamma^\mu - mc) \psi \equiv (i\hbar \not{\partial} - mc) \psi \quad (8.11a)$$

$$= 0$$

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu} \text{Id} \quad (8.11b)$$

The adjoint Dirac equation

$$\begin{aligned} 0 &= \gamma^+ (i\hbar \partial_\mu (\gamma^\mu)^\dagger - mc) \\ &= \psi^\dagger (-i\hbar \overleftarrow{\partial}_\mu \gamma^0 \gamma^\mu \gamma^0 - \gamma^0 \gamma^0 - \gamma^0 \gamma^0 mc) \\ &= \psi^\dagger \gamma^0 (\overleftarrow{\partial}_\mu - mc) \gamma^0 \\ &= 0 \end{aligned} \quad (8.11c)$$

$$\Rightarrow \bar{\psi} (i\hbar \not{\partial} + mc) = 0 \quad (8.11d)$$

$$\bar{\psi} \equiv \psi^\dagger \gamma^0 \quad (8.11e)$$

The current:

$$j^\mu \equiv \bar{\psi} \gamma^\mu \psi \quad (8.12a)$$

$$\begin{aligned} \partial_\mu j^\mu &= (\not{\partial} \bar{\psi}) \psi + \bar{\psi} \not{\partial} \psi \\ &= 0 \end{aligned} \quad (8.12b)$$

$$\begin{aligned} \rho &\equiv j^0 \\ &= \psi^\dagger \gamma^0 \gamma^0 \psi \\ &= \psi^\dagger \psi, \quad \text{positive definite} \end{aligned} \quad (8.12c)$$

( $\rightarrow$  we still will have problem with  $E < 0$  solutions)

## 8.4 Covariance of Dirac equation

Consider LT

$$\begin{aligned} x^\mu &\rightarrow x'^\mu \\ &= \Lambda^\mu{}_\nu x^\nu \\ &= \frac{\partial x'^\mu}{\partial x^\nu} x^\nu, \quad (x \rightarrow \Lambda x) \end{aligned} \quad (8.13a)$$

Dirac

$$(i\hbar\partial_\mu\gamma^\mu - mc)\psi(x) = 0 \quad (8.13b)$$

$$\rightarrow (i\hbar\partial'_\mu - mc)\psi(x') = 0 \quad (8.13c)$$

require transformation

$$\begin{aligned} \psi(x) &\rightarrow \psi'(x') \\ &= S(\Lambda)\psi(x) \end{aligned} \quad (8.13d)$$

such that the “new” equation holds start from  $S(\Lambda) \times$  Dirac equation

$$\begin{aligned} S(\Lambda) \left( i\hbar \frac{\partial}{\partial x^\mu} \gamma^\mu - mc \right) \psi &= S(\Lambda) (i\hbar \Lambda^\nu{}_\mu \partial'_\nu \gamma^\mu - mc) \psi \\ &= S(\Lambda) (i\hbar \Lambda^\nu{}_\mu \partial'_\nu \gamma^\mu - mc) \psi \\ &= (i\hbar S(\Lambda) (\Lambda^\nu{}_\mu \gamma^\mu) S^{-1}(\Lambda) \partial'_\nu \\ &\quad - mc) \underbrace{S(\Lambda)\psi(x)}_{\psi'(x')} \end{aligned} \quad (8.13e)$$

Compare with Dirac in  $S'$

$$S(\Lambda) \Lambda^\nu{}_\mu \gamma^\mu S^{-1}(\Lambda) = \gamma^\nu \quad (8.13f)$$

or

$$\Lambda^\nu{}_\mu \gamma^\mu = S^{-1}(\Lambda) \gamma^\nu S(\Lambda) \quad (8.13g)$$

A proper LT has 6 parameters (3rot, 3boos par)

$$\omega_{\rho\sigma} = -\omega_{\sigma\rho}, \quad (\text{antisymmetric}) \quad (8.13h)$$

Claim: For infinitesimal propel LT:

$$S(\Lambda) = \mathbf{1} + \frac{i}{2} \omega_{\mu\nu} \sigma^{\mu\nu} \quad (8.14a)$$

with

$$\sigma^{\mu\nu} \equiv \frac{i}{2} [\gamma^\mu, \gamma^\nu] \quad (8.14b)$$

or for finite LT

$$S(\Lambda) = e^{\frac{i}{2}\omega_{\mu\nu}\sigma^{\mu\nu}} \quad (8.14c)$$

(compare to QMI rotations)  $\rightarrow$  exercise sheet 12

Fro pariti  $(t, \mathbf{x}) \rightarrow (t, -\mathbf{x})$  or

$$\Lambda^\mu{}_\nu = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad (8.15a)$$

we get for  $\nu = 0$

$$S\gamma^0 S = \gamma^0 \quad (8.15b)$$

and for  $\nu = i$

$$S\gamma^i S = -\gamma^i \quad (8.15c)$$

$$S(\Lambda_p) = \gamma^0 \times \underbrace{\text{Phase}}_1 \quad (8.15d)$$

Define

$$\gamma_5 \equiv i\gamma_0\psi_1\gamma_2\psi_3 \quad (8.16a)$$

in Dirac representation

$$\gamma_5 = \begin{pmatrix} 0 & \mathbf{1} \\ \mathbf{1} & 0 \end{pmatrix} \quad (8.16b)$$

Note

$$(\gamma_5) = \mathbf{1}. \quad (8.16c)$$

Now we can parametrize any  $4 \times 4$  matrix in terms of the following 16 matrices  $\{\mathbf{1}, \gamma_5, \gamma^4, \gamma_5\gamma^4, \sigma^{\mu\nu}\}$ . We know

$$\psi(x) \xrightarrow{\text{LT}} S(\Lambda)\psi(x) \quad (8.17a)$$

$$\bar{\psi}(x) \xrightarrow{\text{LT}} \bar{\psi}(x)S^{-1}(\Lambda) \quad (8.17b)$$

$\rightarrow$  exercise sheet 12

### 8.4.1 Bilinear covariants

$$\bar{\psi}(x)\psi(x) \xrightarrow{\text{pLT}} \bar{\psi}S^{-1}S\psi = \bar{\psi}(x)\psi(x) \quad (8.18a)$$

$$\xrightarrow{\text{Parity}} \bar{\psi}\gamma^0\gamma^0\psi = \bar{\psi}(x)\psi(x) \quad (8.18b)$$

$$\bar{\psi}(x)\gamma_5(x)\psi(x) \xrightarrow{\text{pLT}} \bar{\psi}S^{-1}\gamma^5S\psi = \bar{\psi}\gamma_5\psi \quad (8.18c)$$

$$\{\gamma^5, \gamma^\mu\} = 0 \quad (8.18d)$$

$$\Rightarrow [S, \gamma_5] = 0 \quad (8.18e)$$

$$\bar{\psi}\gamma^\mu\psi \rightarrow \Lambda^\mu{}_\nu \bar{\psi}\gamma^\nu\psi, \quad \text{vector} \quad (8.19a)$$

$$\bar{\psi}\gamma_5\gamma^\mu\psi \rightarrow \text{Det}(\Lambda)\Lambda^\mu{}_\nu \bar{\psi}\gamma^\nu\gamma_5\psi, \quad \text{axial vector} \quad (8.19b)$$

$$\bar{\psi}\sigma^{\mu\nu}\psi \rightarrow \Lambda^\mu{}_\rho \Lambda^\nu{}_\sigma \bar{\psi}\gamma^{\rho\sigma}\psi, \quad \text{tensor rank 2} \quad (8.19c)$$

## 8.5 Solutions to the dirac equation

$$\psi(x) = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix} \quad (8.20)$$

4 comps: 2 spin×2 positive/negative energy  
momentum space

$$\psi^+(x) = e^{-ikx}u(k), \quad \text{positive } E \quad (8.21a)$$

$$\psi^-(x) = e^{ikx}v(k), \quad \text{negative } E \quad (8.21b)$$

$$\begin{aligned} \mathbf{k} \cdot \mathbf{x} &= k_\mu x^\mu \\ &= \omega t - \mathbf{x} \cdot \mathbf{k} \end{aligned} \quad (8.22a)$$

write as:

$$\psi^+ = e^{-\frac{i}{\hbar}(Et - px)}u(p) \quad (8.22b)$$

$$\psi^- = e^{\frac{i}{\hbar}(-Et - px)}u(p) \quad (8.22c)$$

$$p^\mu = \left( \frac{E}{c}, p \right) \quad (8.22d)$$

Dirac

$$(\not{p} - mc) u(p) = 0 \quad (8.23a)$$

$$-\not{p} - mc v(p) = 0 \quad (8.23b)$$

are matrix equations

$$\begin{aligned} (\not{p} \mp mc) &= \begin{pmatrix} \frac{E}{c} \mp mc & -\mathbf{p} \cdot \boldsymbol{\sigma} \\ \mathbf{p} \cdot \boldsymbol{\sigma} & -\frac{E}{c} \mp mc \end{pmatrix} \begin{pmatrix} \chi \\ \varphi \end{pmatrix} \\ &= 0 \end{aligned} \quad (8.23c)$$

par. of  $u$  and/or  $v$

$$u(p) \equiv \begin{pmatrix} \chi \\ \varphi \end{pmatrix} \quad (8.23d)$$

$$\varphi = \left( \frac{c\mathbf{p} \cdot \boldsymbol{\sigma}}{E + mc^2} \right) \chi \quad (8.23e)$$

pick normalization factor  $\sqrt{E + mc^2}$

$$u(p, r) = \begin{pmatrix} \sqrt{E + mc^2} \chi_r \\ \frac{c\mathbf{p} \cdot \boldsymbol{\sigma}}{\sqrt{E + mc^2}} \chi_r \end{pmatrix} \quad (8.23f)$$

and

$$v(p, r) = \begin{pmatrix} \frac{c\mathbf{p} \cdot \boldsymbol{\sigma}}{\sqrt{E + mc^2}} \chi_r \\ \sqrt{E + mc^2} \chi_r \end{pmatrix} \quad (8.23g)$$

$$\chi_1 = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad (8.23h)$$

$$\chi_2 = \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (8.23i)$$

$r$  related to spin, 4-solutions

some propertes of  $u$  and  $v$ :

$$\bar{u}(p, r_i) u(p, r_j) = u^\dagger(p, r_i) \gamma^0 u(p, R - j) \quad (8.23j)$$

$$\begin{aligned} & \left( \sqrt{E + mc^2} \chi_i \quad \frac{c\mathbf{p} \cdot \boldsymbol{\sigma}}{\sqrt{E + mc^2} \chi_i} \right) \begin{pmatrix} \mathbf{1} & 0 \\ 0 & -\mathbf{1} \end{pmatrix} \begin{pmatrix} \sqrt{E + mc^2} \chi_j \\ \frac{c\boldsymbol{\sigma} \cdot \mathbf{p}}{\sqrt{E + mc^2}} \chi_j \end{pmatrix} \\ &= ((E + mc^2) \chi_i \cdot \chi_j - \frac{c\mathbf{p}^2}{(Emc^2)} \chi_i \chi_j) \\ &= 2mc^2 \chi_i \chi_j \\ &= 2mc^2 \delta_{ij} \end{aligned} \quad (8.23k)$$

similar

$$\bar{v}(p, r_i) v(p, r_j) = -2mc^2 \delta_{ij} \quad (8.23l)$$

and

$$\begin{aligned} \bar{v}u &= \bar{u}v \\ &= 0 \end{aligned} \quad (8.23m)$$

Convention: In some books

$$u/v \rightarrow \frac{1}{\sqrt{2mc^2}} u/v \quad (8.23n)$$

$\rightarrow u, v$  form a basis

We can also show:

$$\sum_{i=1}^3 u(p, r_i) \cdot \bar{u}(p, r_i) = c(\not{p} + mc) \quad (8.24a)$$

$\leftarrow$  projection to positive energy states

$$\sum v(p, r_i) \cdot \bar{v}(p, r_i) = c(-\not{p} + mc) \quad (8.24b)$$

$\leftarrow$  negative

Show equivalence for complete set  $u(p, r_j)$  and  $v(p, r_j)$  e.g.

$$\left( \sum_{i=1}^3 u(p, r_i) \cdot \bar{u}(p, r_i) \right) u(p, r_j) = 2mc^2 u(p, r_j) \dots \quad (8.24c)$$

### 8.5.1 Interpretation of solutions and spin

Now we show that Dirac equation describes spin  $\frac{1}{2}$

$$H = \boldsymbol{\alpha} \cdot \mathbf{p} + \beta mc^2 \quad (8.25a)$$

$$\mathbf{L} = \mathbf{x} \times \mathbf{p} \quad (8.25b)$$

$$\begin{aligned} [L_i, H] &= [\varepsilon_{ijk} x_j p_k, \alpha_\ell p_\ell] \\ &= \varepsilon_{ijk} \alpha_\ell [x_j, p_\ell] p_k \\ &= i\hbar \varepsilon_{ijk} \alpha_j p_k \end{aligned} \quad (8.25c)$$

$$\begin{aligned} [\mathbf{L}, H] &= i\hbar \boldsymbol{\alpha} \times \mathbf{p} \\ &\neq 0 \end{aligned} \quad (8.25d)$$

is not conserved. However  $\mathbf{J} = \mathbf{L} + \mathbf{S}$  must be conserved  $\mathbf{S} \neq 0$

need to find  $\mathbf{S}$  such that

$$[\mathbf{J}, H] = 0, \quad (8.26a)$$

i.e.

$$[\mathbf{S}, H] = -i\hbar \boldsymbol{\alpha} \times \mathbf{p} \quad (8.26b)$$

and of course

$$[S_i, S_j] = i\hbar \varepsilon_{ijk} S_k \quad (8.26c)$$

claim:

$$\begin{aligned} \mathbf{S} &= \frac{\hbar}{2} \sum \\ &= \frac{-i\hbar}{2} \alpha^1 \alpha^2 \alpha^3 \boldsymbol{\alpha} \\ &= \frac{\hbar}{2} \begin{pmatrix} \boldsymbol{\sigma} & 0 \\ 0 & \boldsymbol{\sigma} \end{pmatrix} \end{aligned} \quad (8.26d)$$

**Proof:**

$$[S_i, S_j] = i\hbar \varepsilon_{ijk} S_k \quad (8.27a)$$

$$\begin{aligned} [S_i, H] &= \frac{-i\hbar}{2} [\alpha^1 \alpha^2 \alpha^3 \alpha^i, H] \\ &= \frac{-i\hbar}{2} \frac{1}{2} [\alpha^j \alpha^k, \alpha^n p^n + \beta mc^2] \end{aligned} \quad (8.27b)$$

with

$$[\alpha^j \alpha^k, \beta] = 0 \quad (8.27c)$$

use

Let's look at relation between spin and coordinate transformation more care-  
fulle

QMI (nor-rel) [Section 10.4] Translation & rotations: generators  $P^i, J^i$

$$[P^i, P^j] = 0 \quad (8.28a)$$

$$[J^i, J^j] = i\hbar \varepsilon^{ijk} J^k \quad (8.28b)$$

$$[J^i, P^j] = i\hbar \varepsilon^{ijk} P^k \quad (8.28c)$$

coordinate transformation:

$$\begin{aligned}\mathbf{x} &\rightarrow \mathbf{x}' \\ &= R\mathbf{x} + \mathbf{a}\end{aligned}\tag{8.28d}$$

$a$ : state  $|\psi\rangle$  transformation under a certain representation

$$\begin{aligned}|\psi\rangle &\rightarrow |\psi'\rangle \\ &= U(R, a) |\psi\rangle\end{aligned}\tag{8.28e}$$

in relativity add boosts

$$x^\mu \rightarrow \Lambda^\mu{}_\nu x^\nu + a^\mu\tag{8.28f}$$

$$[K_i, K_j] = -i\hbar \varepsilon_{ijk} J_k\tag{8.28g}$$

$$[K_i, J_i] = i\hbar \varepsilon_{ijk} K_k\tag{8.28h}$$

generator

$$J^{\mu\nu} = \begin{cases} J_{0i} = -J_{i0} = K_i \\ J_{ij} = -J_{ji} = i\varepsilon_{ijk} J_k \end{cases}\tag{8.28i}$$

6 generators +4  $P^\mu$

### 8.5.2 Lie algebra of generators

$$[P_\mu, P_\nu] = 0\tag{8.29a}$$

$$[J_{\mu\nu}, J_{\rho\sigma}] = i\hbar (g_{\nu\rho} J_{\mu\sigma} - g_{\mu\rho} J_{\nu\sigma} - g_{\nu\sigma} J_{\mu\rho} + g_{\mu\sigma} J_{\nu\rho})\tag{8.29b}$$

$$[P_\mu, J_{\rho\sigma}] = i\hbar (g_{\mu\rho} P_\sigma - G_{\mu\sigma} P_\rho)\tag{8.29c}$$

under Lorentz transform a state  $|p\rangle$  transforms under a certain representation

$$\begin{aligned}|p\rangle &\rightarrow |\Lambda p\rangle \\ &= U(\Lambda) |p\rangle\end{aligned}\tag{8.29d}$$

$P_\mu P^\mu$  commutes with all generators,  $P^2$  is  $L$ -invariant (Casimir). What else do we need to know (Result: “only” transform under rotations, i.e., “spin”)

Consider any  $P^\mu$ . little group of  $P^\mu$ :

Subgroup of all Poincaré transformations that leave  $P^\mu$  invariant

for  $p^\mu$  in rest frame  $p^\mu = (m, 0, 0, 0)$ : Little group  $\simeq$  rotatinos

Let

$$p^\mu = L^\mu{}_\nu q^\nu\tag{8.30a}$$



i.e. for any  $q^\mu$  with

$$\begin{aligned} q^2 &= m^2 \\ &> 0 \end{aligned} \quad (8.30b)$$

we can find LT  $L(p)$ , s.t.

$$p = L(p)q \quad (8.30c)$$

is in rest frame ... under any LT,  $\Lambda$

$$\begin{aligned} |p\rangle &\rightarrow U(\Lambda) |p\rangle \\ &= U(\Lambda)U(L(p)) |q\rangle \end{aligned} \quad (8.30d)$$

$$U(L(\Lambda p))U^{-1}(L(\Lambda p))U(\Lambda)U(L(p)) |q\rangle \quad (8.30e)$$

$$U(L(\Lambda p)) \cdot U(L^{-1}(\Lambda p))U(\Lambda)U(L(p)) |q\rangle \quad (8.30f)$$

$$U(L(\Lambda p))U(L^{-1}(\Lambda p)\Lambda \cdot L(p)) |q\rangle \quad (8.30g)$$

$\rightarrow$  additional labels in  $|p, s\rangle$  are affected by rotations only

$$\begin{aligned} |p, s\rangle &\rightarrow U(\Lambda) |p, s\rangle \\ &= U(L(\Lambda p)) \underbrace{\sum D_{ss'}}_{\text{def. transformation under rotations}} \\ &= \sum D_{ss'} |\Lambda p, s'\rangle \end{aligned} \quad (8.31a)$$

---

### 8.5.3 2nd casimir operator

$$W_\mu W^\mu = -m^2 \hbar^2 s(s+1) \quad (8.32a)$$

Pauli-Lubanski (axial) vector

$$W_\mu = -\frac{1}{2} \varepsilon_{\mu\nu\rho\sigma} J^{\nu\rho} P^\sigma \quad (8.32b)$$

in rest frame

$$P^\sigma = (m, 0, 0, 0) \quad (8.32c)$$

$$W_\mu = (0, \boldsymbol{\omega}) \quad (8.32d)$$

$$\begin{aligned} W_i &= -\frac{1}{2} \varepsilon_{i\nu\rho\sigma} J^{\nu\rho} P^\sigma \\ &= -\frac{m}{2} \varepsilon_{i\nu\rho\sigma} J^{\nu\rho} \\ &= -m J_i \end{aligned} \quad (8.32e)$$

## 8.6 The non-relativistic limit

Consider expansion in  $\frac{\mathbf{p}}{m} \sim V$  of Dirac equation.

We will derive the Hamiltonian used in section 2 (for fine structure)

Dirac equation in presence of em field:

$$p^\nu \rightarrow p^\nu + \frac{e}{c} A^\nu \quad (8.33a)$$

write

$$\psi = \begin{pmatrix} \chi \\ \eta \end{pmatrix}, \quad \text{2-component-spinors} \quad (8.33b)$$

$$\not{p} + \frac{e}{c} \not{A} - mc \quad (8.33c)$$

$$\begin{pmatrix} \chi \\ \eta \end{pmatrix} = \begin{pmatrix} \frac{E}{c} + \frac{e}{c} \phi - mc & -(\mathbf{p} + \frac{e}{c} \mathbf{A}) \cdot \boldsymbol{\sigma} \\ (\mathbf{p} + \frac{e}{c} \mathbf{A}) \cdot \boldsymbol{\sigma} & -(\frac{E}{c} + \frac{e}{c} \phi + mc) \end{pmatrix} \begin{pmatrix} \chi \\ \eta \end{pmatrix} = 0 \quad (8.33d)$$

$$\Rightarrow \left( \frac{E}{c} + \frac{e}{c} \phi - mc \right) \chi = \left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right) \cdot \boldsymbol{\sigma} \eta \quad (8.33e)$$

$$\left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right) \cdot \boldsymbol{\sigma} \chi = \left( \frac{E}{c} + mc + \frac{e}{c} \phi \right) \eta \quad (8.33f)$$

$$\begin{aligned} \rightsquigarrow \eta &= \frac{(\mathbf{p} + \frac{e}{c} \mathbf{A}) \cdot \boldsymbol{\sigma}}{\frac{E}{c} + mc + \frac{e}{c} \phi} \chi \\ &\approx \frac{(\mathbf{p} + \frac{e}{c} \mathbf{A}) \cdot \boldsymbol{\sigma}}{2mc} \chi \end{aligned} \quad (8.33g)$$

Putting Eq. 8.33g in Eq. 8.33e gives us

$$\begin{aligned} \left( \frac{E}{c} + \frac{e}{c} \phi - mc \right) \chi &= \left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right) \cdot \boldsymbol{\sigma} \left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right) \cdot \boldsymbol{\sigma} \frac{1}{2mc} \chi \\ &= \frac{1}{2mc} \left( \left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right)^2 \right. \\ &\quad \left. + i(-i\hbar) \frac{e}{c} \underbrace{(\nabla \times \mathbf{A} + \mathbf{A} \times \nabla)}_{[\nabla \times \mathbf{A}]} \cdot \boldsymbol{\sigma} \right) \chi \end{aligned} \quad (8.33h)$$

“derivative within  $[\cdot]$  only”

$$\begin{aligned}
 &\Rightarrow \left( \frac{1}{2m} \left( \mathbf{p} \cdot \frac{e}{c} \mathbf{A} \right)^2 + \frac{\hbar e}{2mc} \boldsymbol{\sigma} \cdot \mathbf{B} - e\phi \right) \chi \\
 &= (E - mc^2) \chi \\
 &=: E' \chi
 \end{aligned} \tag{8.33i}$$

where  $\frac{\hbar e}{2mc} \boldsymbol{\sigma} \cdot \mathbf{B}$  has to be compare to section 2.3 with

$$\begin{aligned}
 H &\sim \boldsymbol{\mu} \cdot \mathbf{B} \\
 &= \frac{e}{2m} g \mathbf{S} \cdot \mathbf{B}
 \end{aligned} \tag{8.33j}$$

$$\begin{aligned}
 &= \frac{e}{4m} g \boldsymbol{\sigma} \cdot \mathbf{B} \\
 &\Rightarrow g = 2
 \end{aligned} \tag{8.33k}$$

Let's do this more systematically (expand in  $\frac{|e|}{m}$ )

Find transformation

$$\psi = e^{-iS} \psi', \quad (\text{Foldy-Wonntthuysen transformation}) \tag{8.34a}$$

such thath odd operators are suppressed by  $\left( \frac{|\mathbf{p}|}{m} \right)^n \rightsquigarrow$  mixes  $\chi$  and  $\eta$

$$\sigma := c \boldsymbol{\alpha} \left( \mathbf{p} + \frac{e}{c} \mathbf{A} \right) \tag{8.34b}$$

$$\begin{aligned}
 i\hbar \partial_t \psi' &= i\hbar \partial_t e^{iS} \psi \\
 &= i\hbar \left[ \partial_t e^{iS} \right] \psi + e^{\otimes S} i\hbar \partial_t \psi \\
 &= i\hbar \left[ \partial_t e^{+iS} \right] e^{-iS} \psi' + e^{iS} H e^{-iS} \psi' \\
 &=: H' \psi'
 \end{aligned} \tag{8.34c}$$

$$\begin{aligned}
 H' &= H + i \left[ S, H \right] + \frac{i^2}{2!} \left[ S, \left[ S, H \right] \right] + \frac{i^3}{3!} \dots \\
 &\quad - \hbar \dot{S} - \frac{i\hbar}{2!} \left[ S, \dot{S} \right] - \frac{i^2 \hbar}{3!} \left[ S, \left[ S, \dot{S} \right] \right] \dots
 \end{aligned} \tag{8.34d}$$

Recall

$$H = \beta mc^2 + \sigma + \xi, \quad \text{even: } \xi = -e\phi \tag{8.34e}$$

Set

$$S = \frac{-i\beta}{2mc^2} \sigma \tag{8.34f}$$

Compute

$$\begin{aligned} i[S, H] &= i \left[ \frac{-i\beta}{2mc^2} \sigma, \beta mc^2 + \sigma + \xi \right] \\ &= \frac{\beta}{2mc^2} [\sigma, \xi] + \frac{\beta}{mc^2} \sigma^2 - \sigma \end{aligned} \quad (8.34g)$$

## 8.7 Observables in 2nd quantization

Observables expressed in terms of fields /to operators in Fock space  $\mathcal{F}$ .

Example: particle number density

**QM:** probability density

$$\rho(x) = |\psi(x)|^2 \quad (8.35)$$

**now:**

$$\begin{aligned} \rho(x) &= \hat{\psi}^\dagger(x) \hat{\psi}(x) \\ &= \sum_{ij} \hat{a}_i^\dagger \hat{a}_j \psi_i^*(x) \psi_j(x) \\ &= \sum_{ij} \hat{a}_i^\dagger \hat{a}_j \langle x|i \rangle \langle j|x \rangle \end{aligned} \quad (8.36a)$$

$$\begin{aligned} \int d^3\mathbf{x} \rho(x) &= \sum_{ij} \hat{a}_i^\dagger \hat{a}_j \langle j| \underbrace{\int d^3x |x\rangle \langle x|}_{1} |i\rangle \\ &= \sum_i \hat{a}_i^\dagger \hat{a}_i \\ &= \sum_i n_i \end{aligned} \quad (8.36b)$$

$$\hat{T} = \int d^3\mathbf{x} \hat{\psi}^\dagger \left( -\frac{\hbar^2 \nabla^2}{2m} \right) \hat{\psi}(x) \quad (8.37a)$$

$$\hat{\psi}(x) = \sum_i \hat{a}_i \psi_i(x) \quad (8.38a)$$

$$\hat{\psi}^\dagger(x) = \sum_i \hat{a}_i^\dagger \quad (8.38b)$$

$$\hat{\psi}(x) = \int \frac{d^3k}{(2\pi)^3} \hat{a}_k e^{-i\mathbf{k}\cdot\mathbf{x}} \quad (8.38c)$$

$$\hat{\psi}^\dagger(x) = \int \frac{d^3k}{(2\pi)^3} \hat{a}_k^\dagger e^{i\mathbf{k}\cdot\mathbf{x}} \quad (8.38d)$$

$$\begin{aligned} \hat{T} &= \int d^3\mathbf{x} \int \frac{d^3k'}{(2\pi)^3} \underbrace{e^{i\mathbf{x}(\mathbf{k}-\mathbf{k}')}}_{(2\pi)^3\delta(\mathbf{k}-\mathbf{k}')} \hat{a}_k^\dagger \hat{a}_{k'} \left( -\frac{\hbar^2 (i\mathbf{k}')^2}{2m} \right) \\ &= \int \frac{d^3k}{(2\pi)^3} \frac{\hbar^2 \mathbf{k}^2}{2m} \underbrace{\hat{a}_k^\dagger \hat{a}_k}_{N_k} \end{aligned} \quad (8.38e)$$

$N_k$  counts number of particles of kind  $k$ . Potential

$$U(x) = \int \frac{d^3q}{(2\pi)^3} e^{-i\mathbf{q}\cdot\mathbf{x}} \tilde{U}(\mathbf{q}) \quad (8.38f)$$

$$\begin{aligned} \hat{U} &= \int d^3\mathbf{x} \hat{\psi}^\dagger(x) U(x) \hat{\psi}(x) \\ &= \int \frac{d^3k_1}{(2\pi)^3} \int \frac{d^3k_2}{(2\pi)^3} \int \frac{d^3q}{(2\pi)^3} \int \underbrace{d^3x e^{i\mathbf{k}_1\cdot\mathbf{x}} e^{-i\mathbf{k}_2\cdot\mathbf{x}} e^{-i\mathbf{q}\cdot\mathbf{x}}}_{(2\pi)^3\delta(\mathbf{k}_1-\mathbf{k}_2-\mathbf{q})} \hat{a}_{k_1}^\dagger \hat{a}_{k_2} \tilde{U}(\mathbf{q}) \\ &= \int \frac{d^3\mathbf{k}_1}{(2\pi)^3} \int \frac{d^3\mathbf{k}_2}{(2\pi)^3} \tilde{U}(\mathbf{k}_1\mathbf{k}_2) \end{aligned} \quad (8.38g)$$

$$\begin{aligned} V(x_1, x_2) &= V(x_1 - x_2) \\ &= \int \frac{d^3q}{(2\pi)^3} e^{-i\mathbf{q}(\mathbf{x}_1-\mathbf{x}_2)} \tilde{V}(q) \end{aligned} \quad (8.38h)$$

$$\begin{aligned}
\hat{V} &= \int d^3x_1 \int d^3x_2 \hat{\psi}^\dagger(x_1) \hat{\psi}^\dagger(x_2) V(x_1 - x_2) \hat{\psi}(x_1) \hat{\psi}(x_2) \\
&= \int \prod_{i=1}^4 \frac{dk_i}{(2\pi)^3} \int \frac{d^3q}{(2\pi)^3} \cdot \int d^3x_1 e^{i\mathbf{x}_1(\mathbf{k}_1 - \mathbf{k}_3 - \mathbf{q})} \\
&\quad \cdot \int d^3x_2 e^{i\mathbf{x}_2(\mathbf{k}_2 - \mathbf{k}_4 + \mathbf{q})} \hat{a}_{k_1}^\dagger \hat{a}_{k_2}^\dagger \hat{a}_{k_3} \hat{a}_{k_4} \\
&= \int \frac{d^3\mathbf{k}_1}{(2\pi)^3} \int \frac{d^3\mathbf{k}_2}{(2\pi)^3} \int \frac{d^3q}{(2\pi)^3} \hat{a}_{k_1}^\dagger \hat{a}_{k_2}^\dagger \hat{a}_{\mathbf{k}_1 - \mathbf{q}} \hat{a}_{\mathbf{k}_2 + \mathbf{q}} \tilde{V}(q)
\end{aligned} \tag{8.38i}$$

consider commutation relation between field operators:

$$\begin{aligned}
[\hat{\psi}(x), \hat{\psi}(y)]_{\mp} &= \int \frac{d^3k}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} e^{-i\mathbf{k} \cdot \mathbf{x}} e^{+i\mathbf{k}' \cdot \mathbf{y}} \underbrace{[\hat{a}_k, \hat{a}_{k'}^\dagger]}_{(2\pi)^3 \delta(\mathbf{k} - \mathbf{k}')} \stackrel{!}{=} \delta(\mathbf{x} - \mathbf{y})
\end{aligned} \tag{8.39a}$$

with the  $(2\pi)^3$  of  $(2\pi)^3 \delta(\mathbf{k} - \mathbf{k}')$  being convection and generalization of

$$[\hat{a}_i, \hat{a}_j^\dagger] = \delta_{ij} \tag{8.39b}$$

time dependence of field operator: Schrödinger picture  $\rightarrow$  Heisenberg

$$\underbrace{\hat{\psi}(\mathbf{x}, t)}_{\text{Heisenberg}} = e^{\frac{i}{\hbar} \hat{H} t} \underbrace{\hat{\psi}(\mathbf{x})}_{\text{Schrödinger}} e^{-\frac{i}{\hbar} \hat{H} t} \tag{8.40a}$$

$$\begin{aligned}
i\hbar \frac{\partial}{\partial t} \hat{\psi}(\mathbf{x}, t) &= - [\hat{H}, \hat{\psi}(\mathbf{x}, t)] \\
&= - [\hat{H}, \hat{\psi}(\mathbf{x})] \\
&= -e^{\frac{i}{\hbar} \hat{H} t} [\hat{H}, \hat{\psi}(x)] e^{-\frac{i}{\hbar} \hat{H} t}
\end{aligned} \tag{8.40b}$$

take free case

$$\hat{H} = \hat{T}, \quad \text{bosons} \tag{8.40c}$$

$$[\hat{T}, \hat{\psi}(x)] = \int d^3y \frac{\hbar^2}{2m} [\nabla_y \psi^\dagger(y) \nabla_y \hat{\psi}(y), \hat{\psi}(x)] \tag{8.40d}$$

from now on, don't write  $\nabla_y$  but  $\nabla$

$$\begin{aligned}
 \left[ \nabla_y \psi^\dagger(y) \nabla_y \hat{\psi}(y), \hat{\psi}(x) \right] &= \nabla \hat{\psi}^\dagger(y) \zeta \hat{\psi}(y) \hat{\psi}(x) - \hat{\psi}(x) \nabla \hat{\psi}^\dagger(y) \nabla \hat{\psi}(y) \\
 &= \nabla \hat{\psi}^\dagger(y) \hat{\psi}(x) \nabla \hat{\psi}(y) - \hat{\psi}(x) \nabla \hat{\psi}^\dagger(y) \nabla \hat{\psi}(y) \\
 &= \nabla \left( \left[ \hat{\psi}^\dagger(y), \hat{\psi}(x) \right] \right. \\
 &\quad \left. + \underbrace{\hat{\psi}(x) \hat{\psi}^\dagger(y)}_{=0} \right) \nabla \hat{\psi}(y) - \hat{\psi}(x) \nabla \hat{\psi}^\dagger(y) \nabla \hat{\psi}(y) \\
 &= \int d^3y \frac{-\hbar^2}{2m} \nabla_y \delta(\mathbf{x} - \mathbf{y}) \nabla \hat{\psi}(y) \\
 &= \frac{\hbar^2}{2m} \nabla_x^2 \hat{\psi}(x)
 \end{aligned}$$

in free case

$$i\hbar \frac{\partial}{\partial t} \hat{\psi}(x, t) = -\frac{\hbar^2}{2m} \nabla^2 \hat{\psi}(x, t) + \diamond \quad (8.40e)$$

$\diamond$  with interactions much more complicated  $\rightarrow$  QFT

## 8.8 Quantization of relativistic fields

- Start with Lagrangian density  $\mathcal{L}$  for classical field theory
- Compute conjugate momentum field
- impose equal-time (anti-) commutation relations

**Example 1: free scalar field**  $\Phi(\mathbf{x}, t) = \Phi(x)$

•

$$\mathcal{L} = \frac{1}{2} \partial_\mu \Phi \partial^\mu \Phi - \frac{1}{2} \underbrace{\frac{m^2 c^2}{\hbar^2}}_{\equiv m^2} \Phi^2 - \underbrace{V(\Phi)}_{=0} \quad (8.41a)$$

free field (no interactions). Action:

$$\begin{aligned} S &= \int dt \int d^3x \mathcal{L} \\ &= \int d^3x \mathcal{L}(\Phi, \partial_\mu \Phi) \end{aligned} \quad (8.41b)$$

$$\delta S = 0 \quad (8.41c)$$

$$\rightarrow \frac{\partial \mathcal{L}}{\partial \Phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \Phi)} = 0 \quad (8.41d)$$

for our  $\mathcal{L}$ :

$$-m^2 \Phi - \partial_\mu \partial^\mu \Phi = 0 \quad (8.41e)$$

Klein-Gordan equation

- Conjugate momentum field

$$\begin{aligned} \pi(\mathbf{x}, t) &= \frac{\partial \mathcal{L}}{\partial \dot{\Phi}} \\ &= \dot{\Phi}(\mathbf{x}, t) \end{aligned} \quad (8.42a)$$

Hamiltonian

$$\begin{aligned} H &= \int d^3x \pi \dot{\Phi} - L \\ &= \int d^3x (\pi \dot{\Phi} - \mathcal{L}) \\ &= \frac{1}{2} \int d^3x (\pi^2 + (\nabla \Phi)^2 + m^2 \Phi^2) \end{aligned} \quad (8.42b)$$

with  $\mathcal{L}$  the Lagrangian density

$$\begin{aligned} \int d^3x \mathcal{L} &= L \\ &= \text{Lagrangian} \end{aligned} \quad (8.42c)$$

- now second quantization:

$$\Phi(x) \rightarrow \hat{\Phi}(x) \quad (8.43a)$$



and

$$\pi(x) \rightarrow \hat{\pi}(x) \quad (8.43b)$$

$$\begin{aligned} [\hat{\phi}(\mathbf{x}, t), \hat{\phi}(\mathbf{y}, t)] &= [\hat{\pi}(\mathbf{x}, t), \hat{\pi}(\mathbf{y}, t)] \\ &= 0 \end{aligned} \quad (8.43c)$$

impose

$$[\hat{\Phi}(\mathbf{x}, t), \hat{\pi}(\mathbf{y}, t)] = i\delta(\mathbf{x} - \mathbf{y}) \quad (8.43d)$$

$$\sim [x_i, p_j] = i\delta_{ij}, \quad (\hbar = 1) \quad (8.43e)$$

spin 0, bosonic field use  $[\cdot, \cdot] = [\cdot, \cdot]_-$  and not  $\{\cdot, \cdot\} = [\cdot, \cdot]_+$

now as for radiation field

$$\hat{\Phi}(\mathbf{x}, t) = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k} \left( \hat{a}_k e^{i\mathbf{k}\mathbf{x} - i\omega_k t} + \hat{a}_k^\dagger e^{-i\mathbf{k}\mathbf{x} + i\omega_k t} \right) \quad (8.43f)$$

with  $\omega_k = \sqrt{\mathbf{k}^2 + m^2}$  this satisfies KGE

$$\begin{aligned} \hat{\pi}(\mathbf{x}, t) &= \dot{\hat{\Phi}}(\mathbf{x}, t) \\ &= \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k} \left( -i\omega_k \hat{a}_k e^{i\mathbf{k}\mathbf{x} - i\omega_k t} + i\omega_k \hat{a}_k^\dagger e^{-i\mathbf{k}\mathbf{x} + i\omega_k t} \right) \end{aligned} \quad (8.43g)$$

Note: change in normalization:

$$\int \frac{d^3\mathbf{k}}{(2\pi)^3} \rightarrow \underbrace{\int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{1}{2\omega_k}}_{\text{Lorentz invariant}} \quad (8.43h)$$