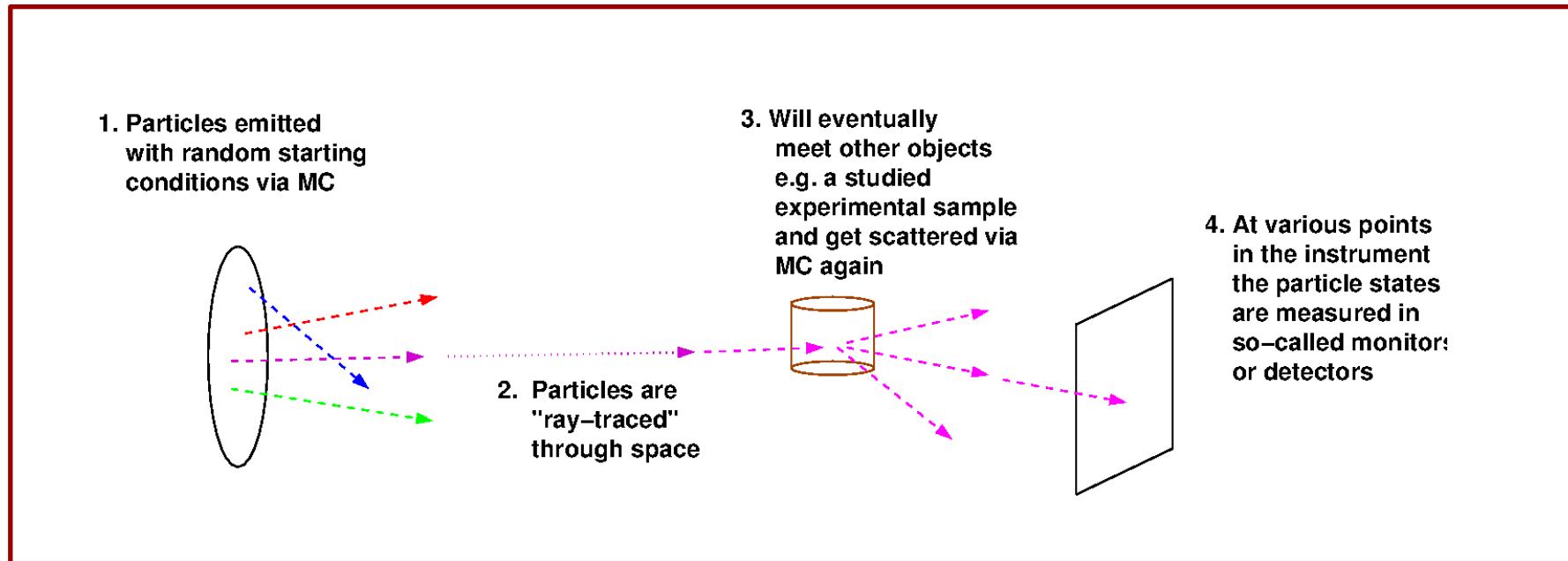


Peter Willendrup

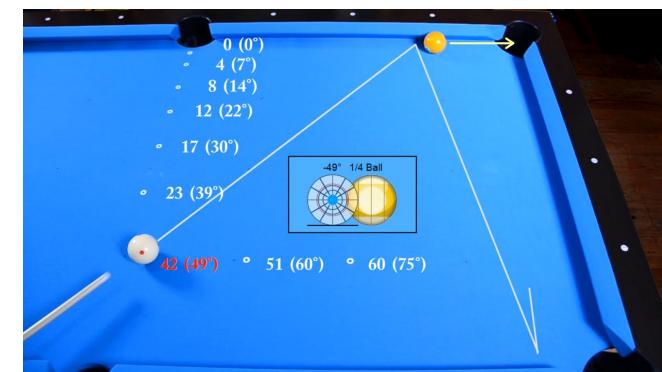
# Introduction to basic concepts of McStas



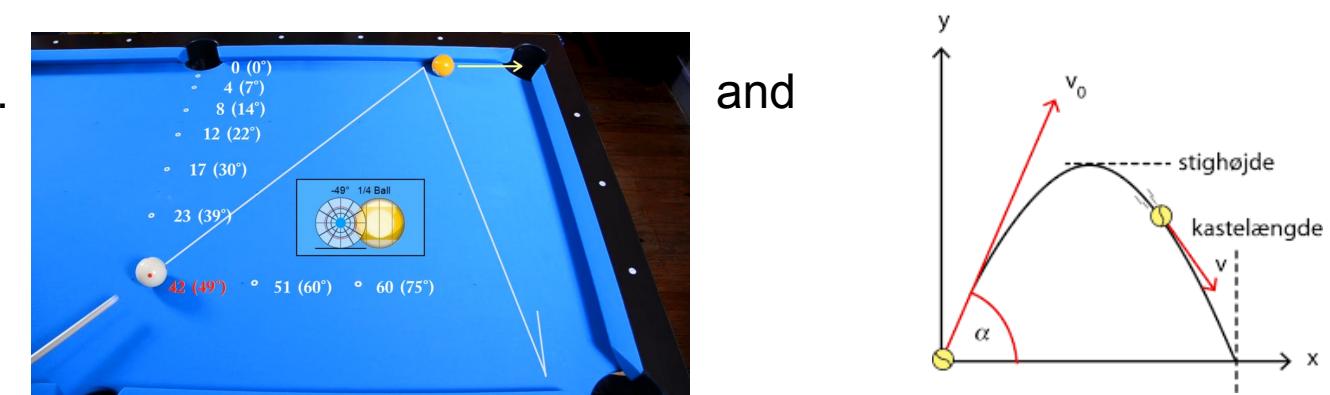
# In the big picture, McStas is this...



- Classical Newtonian mechanics, i.e.
- (independent, particles though...)



and

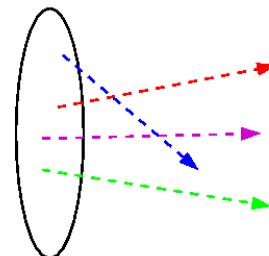




# In the big picture, McStas is this...

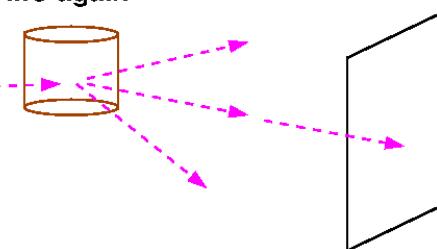
Instrument

1. Particles emitted with random starting conditions via MC



2. Particles are "ray-traced" through space

3. Will eventually meet other objects e.g. a studied experimental sample and get scattered via MC again

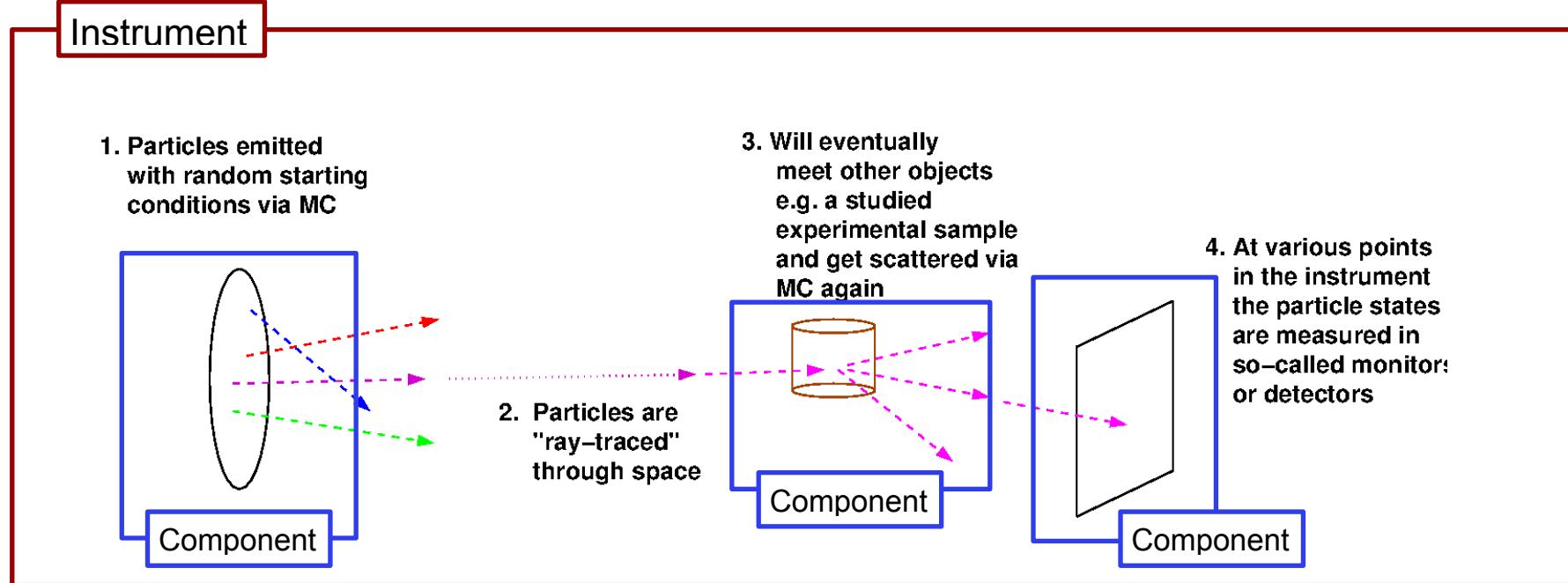


4. At various points in the instrument the particle states are measured in so-called monitors or detectors

The instrument defines our “lab coordinate system”



# In the big picture, McStas is this...

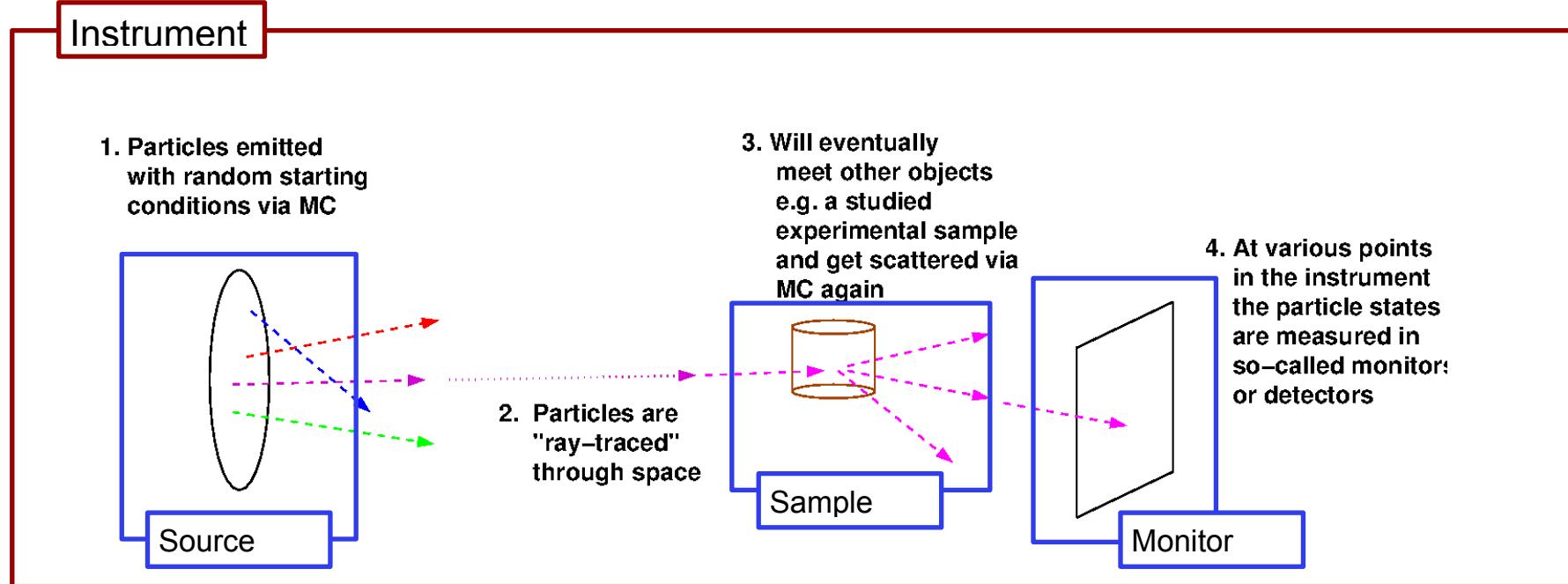


The instrument defines our “lab coordinate system”

The components define devices or features available in our instrument



# In the big picture, McStas is this...

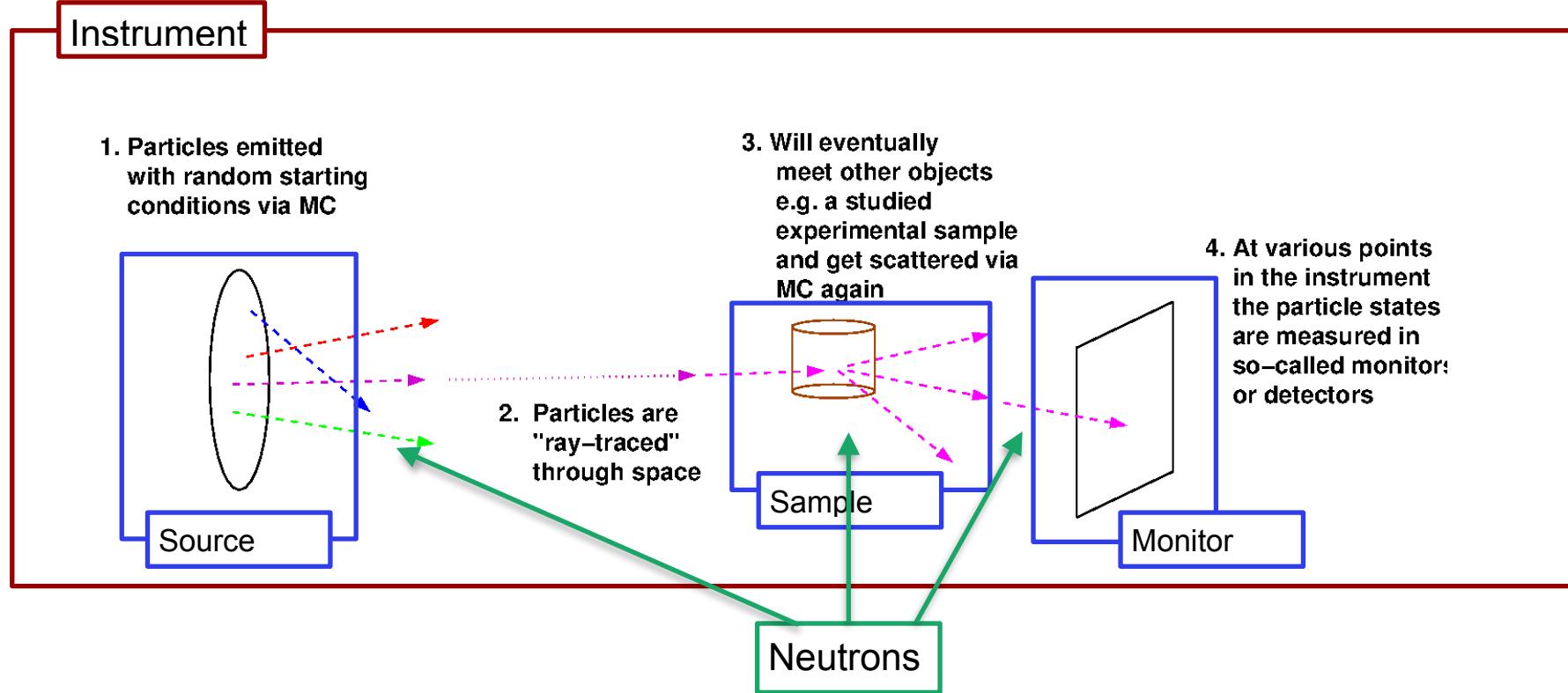


The instrument defines our “lab coordinate system”

The components define devices or features available in our instrument - they have different function



# In the big picture, McStas is this...



The instrument defines our “lab coordinate system”

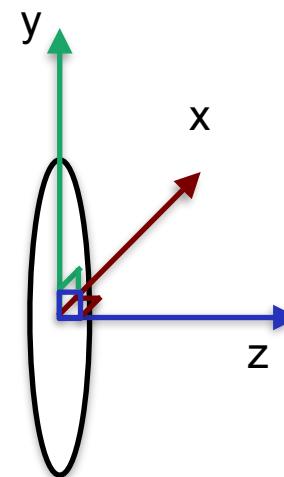
The components define devices or features available in our instrument - they have different function

Neutron particles are passed on from one component to the next, changing state under way

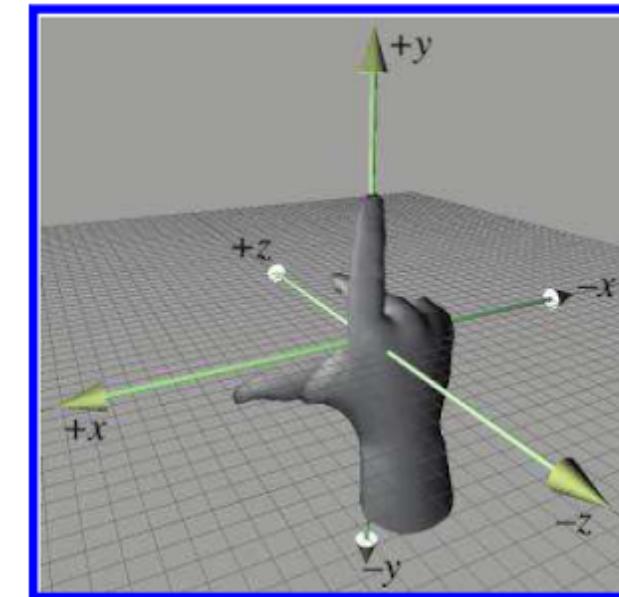


# Placing components - source

- One of the first components in your instrument is typically a source, which has a coordinate system like this....



- z is along neutron beam direction
- y is vertical
- x at an angle of  $90^\circ$  wrt. z,y

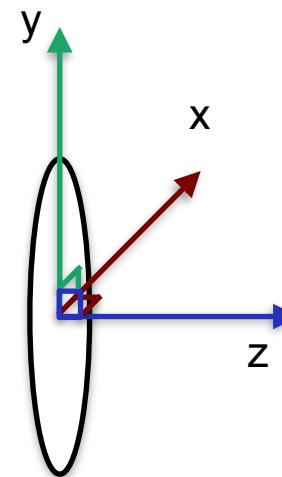


Right-handed  
coordinate system



# Placing components - source

- Often the source coordinate system coincides with the “lab” coordinate system, denoted ABSOLUTE in McStas language, i.e.



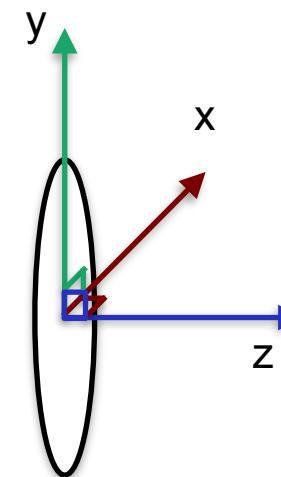
- COMPONENT Source = Source\_simple(...  
AT (0,0,0) ABSOLUTE



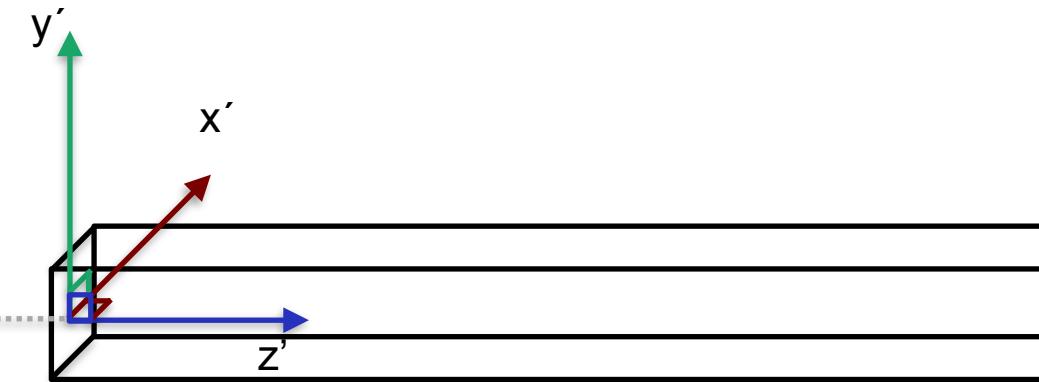
# Placing further components - RELATIVE

Placing further components is done by order of

1. Location, i.e



```
COMPONENT Source = Source_simple(...)  
AT (0,0,0) ABSOLUTE
```

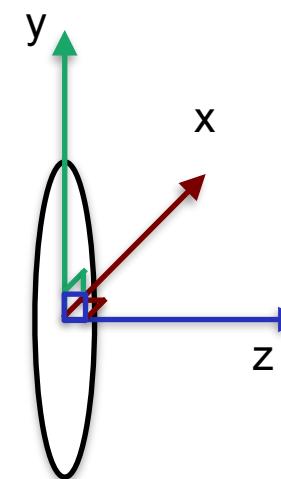


```
COMPONENT Guide = Guide(...)  
AT (0,0,1) RELATIVE Source
```

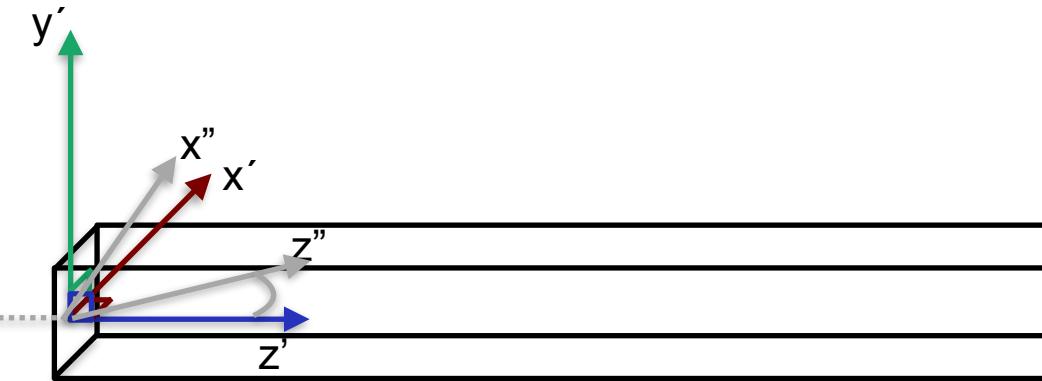


# Placing further components - RELATIVE

Placing further components is done by order of  
2. Rotation, i.e



```
COMPONENT Source = Source_simple(...)  
AT (0,0,0) ABSOLUTE
```

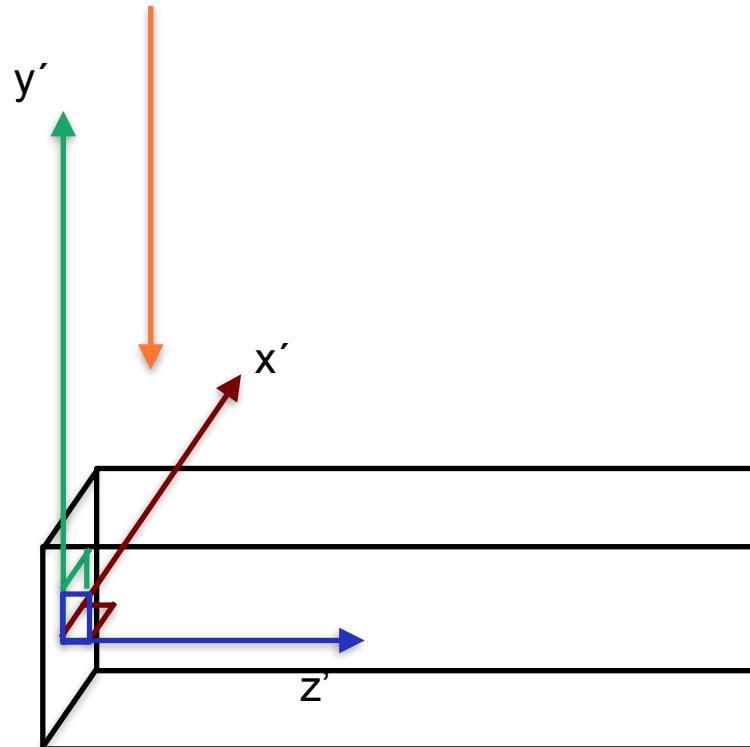
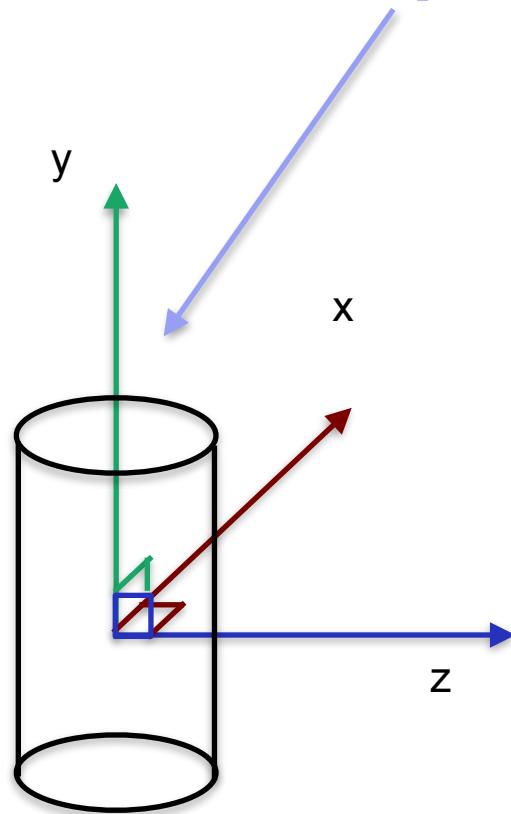


```
COMPONENT Guide = Guide(...)  
AT (0,0,1) RELATIVE Source  
ROTATED (0,30,0) RELATIVE Source
```

(Reference labels can also be PREVIOUS or PREVIOUS+1 etc.)



# Components often have their origin at the centre of mass, i.e. for samples ... but not for neutron guides



Generally speaking, the component author can choose **the meaningful coordinate system for the given problem!**

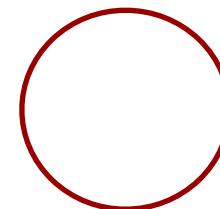
- The McStas system takes care of the transformation between them....



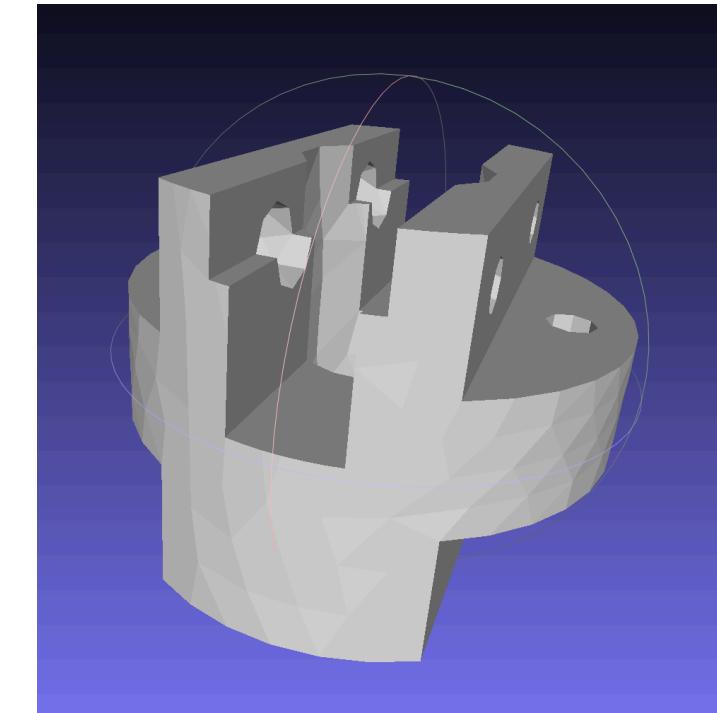
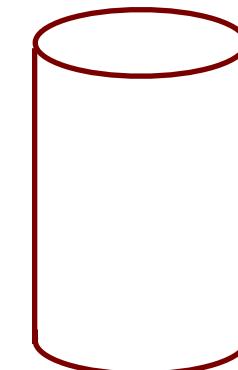
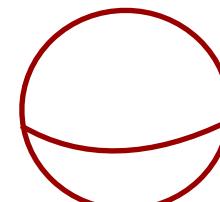
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McStas  
School

# Component geometries are typically simple objects... But some have polygon-description of the surface

2D



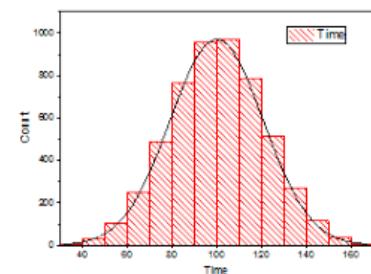
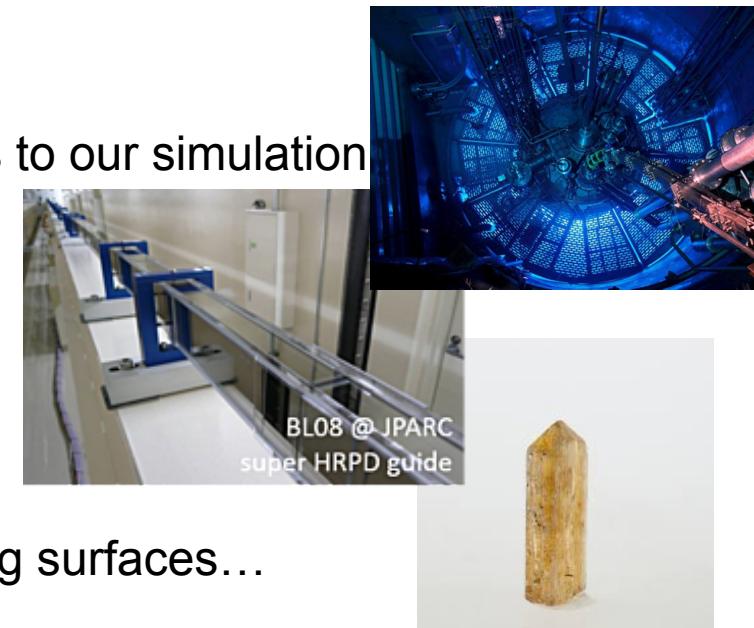
3D





# Component classes

- Sources - these define MC starting conditions / “inject” neutrons to our simulation
- Optics - used to tailor properties of the neutron beam
  - Examples are mirrors, guides, choppers, collimators, slits, ...
- Samples - “matter” of some form
  - Powders, single crystals, liquids, micelles in solution, reflecting surfaces...
- Monitors - may probe the state of the neutron beam and store histograms / event lists
- Misc, obsolete
  - “Other stuff” and “Old stuff”





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INSTITUTE  
FOR  
NUCLEAR  
SCIENCES

ESS

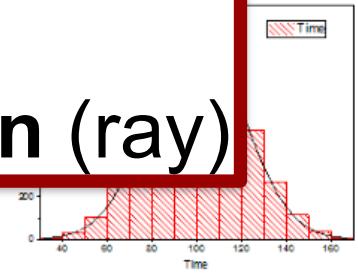
EUROPEAN  
SPALLATION  
SOURCE

# Component classes

- Sources - these define MC starting conditions / “inject” neutrons to our simulation
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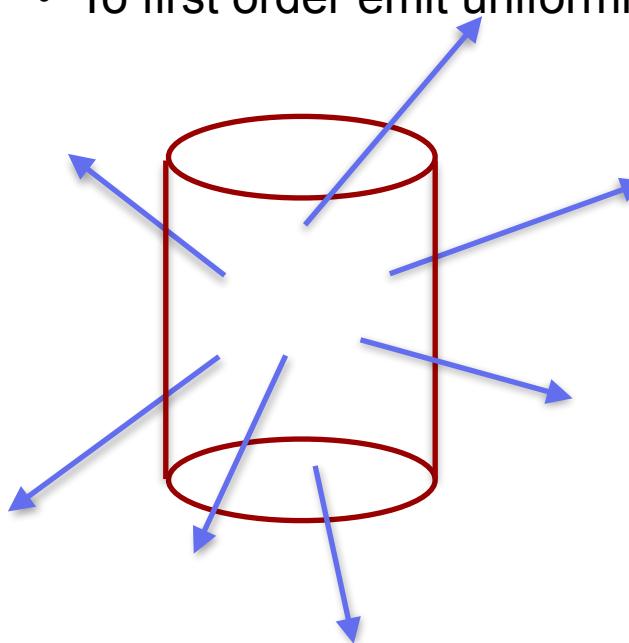
Common to all components:  
They set, manipulate/interact with  
or measure the **state of the neutron (ray)**





# Neutron sources, i.e. moderators

- To first order emit uniformly into  $4\pi$  steradian





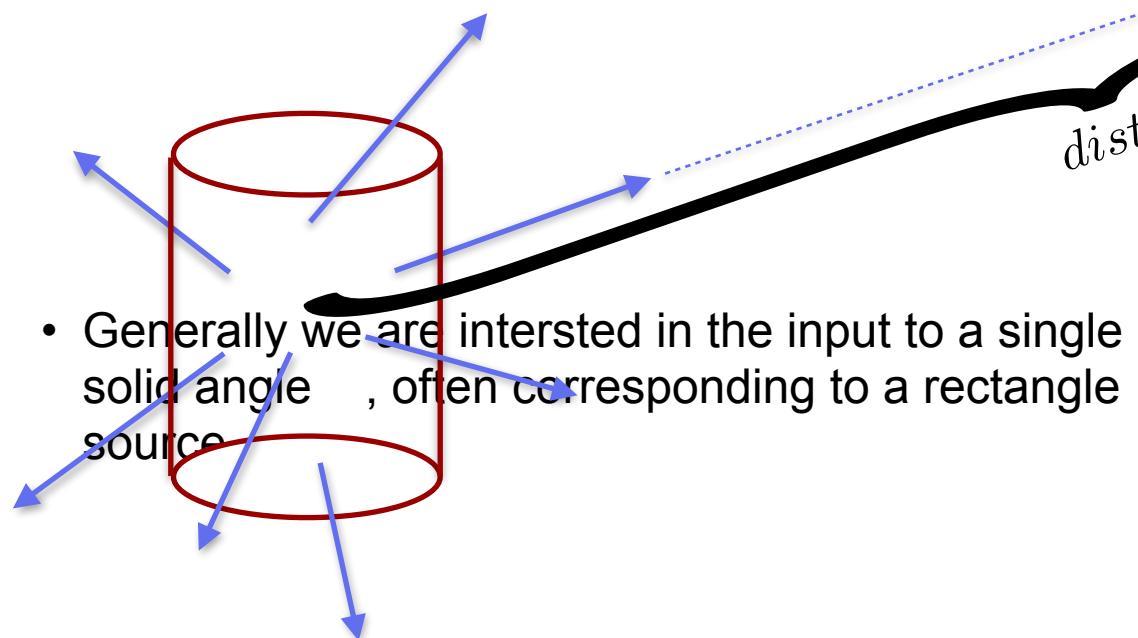
McStas

n

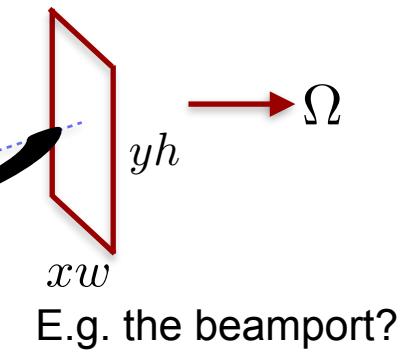
DTU

ESS  
EUROPEAN SPALLATION SOURCE

# Neutron sources, i.e. moderators



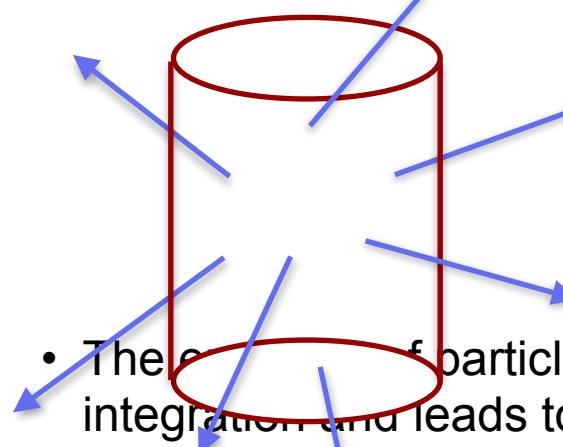
- Generally we are interested in the input to a single instrument, characterised by a certain solid angle  $\Omega$ , often corresponding to a rectangle  $xw \times yh$  at a distance  $dist$  from the source





# Neutron sources, i.e. moderators

- The emission intensity into our chosen solid angle  $\Omega$  can be a function of wavelength, time (pulsed sources) and possibly point of origin on the source surface



- The count of particles into the solid angle  $\Omega$  is in fact an integration and leads to a simulated “intensity”  $I(\lambda, t)$ .
- In McStas, that integrated intensity is partitioned over a given set of particle rays referred to as **ncount**, **-n** or **--ncount**
- The default **ncount** is 1e6 rays

$$I(\lambda, t)$$

$[n/s/str]$

$[n/s/str]$

$[n/s/str]$

$$\Omega$$

$$I_\Omega [n/s]$$

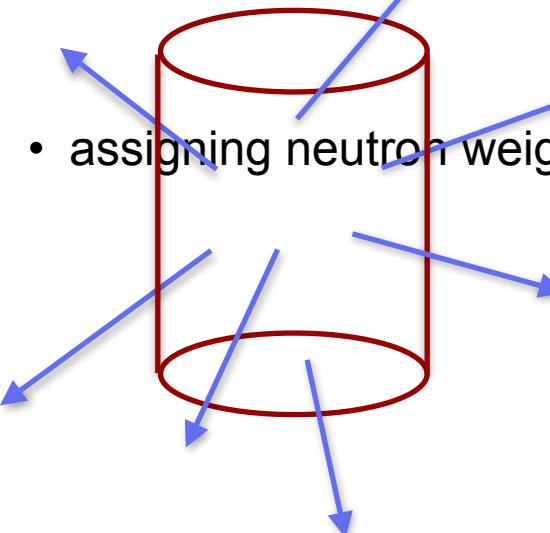
$$\Omega$$

$$\Omega$$



# Neutron sources, i.e. moderators

- Our neutron rays are emitted randomly, sampling “spectrum”, i.e. wavelength, time and area and all variables of the source
- assigning neutron weights such that



$$I_{\Omega}(\lambda, t, \vec{r}) [n/s]$$

$p$

$$\sum_{j=1}^{\text{ncount}} p_j = \int d\lambda, dt, d\vec{r} I_{\Omega}(\lambda, t, \vec{r})$$

 $\Omega$



# Neutron rays in McStas - what are they?

- Defining the neutron starting conditions imply setting:
  - The **starting point** on the surface, i.e. (in the code variables  $x, y, z$ )
  - The **direction** into and our /  $\lambda$  ( $E$  in the code variables  $vx, vy, vz$ )
  - The **starting time** (in the code the variable  $t$ )
  - The initial **intensity** / weight of the neutron ray (in the code the variable  $p$ )
  - If needed the initial **polarisation** (in the code the variables  $sx, sy, sz$ )

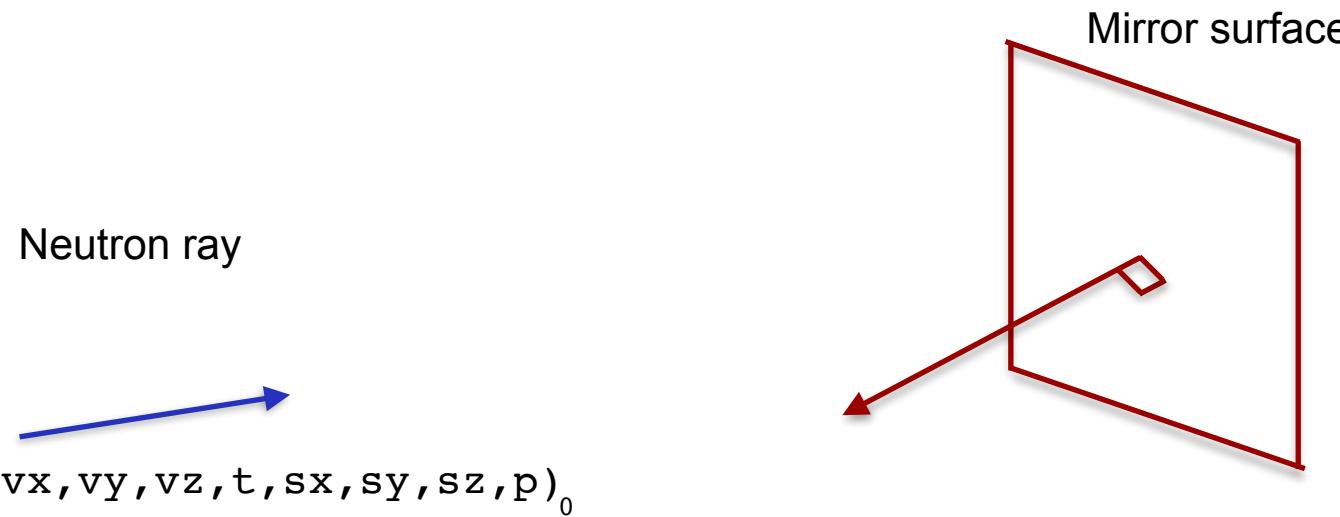
Neutron ray in McStas:	
Location	$x, y, z$
Velocity	$vx, vy, vz$
Time	$t$
Polarisation.	$sx, sy, sz$
Intensity	$p$



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# Neutron (ray)-matter interaction 1: reflecting surface

- 1 starting situation



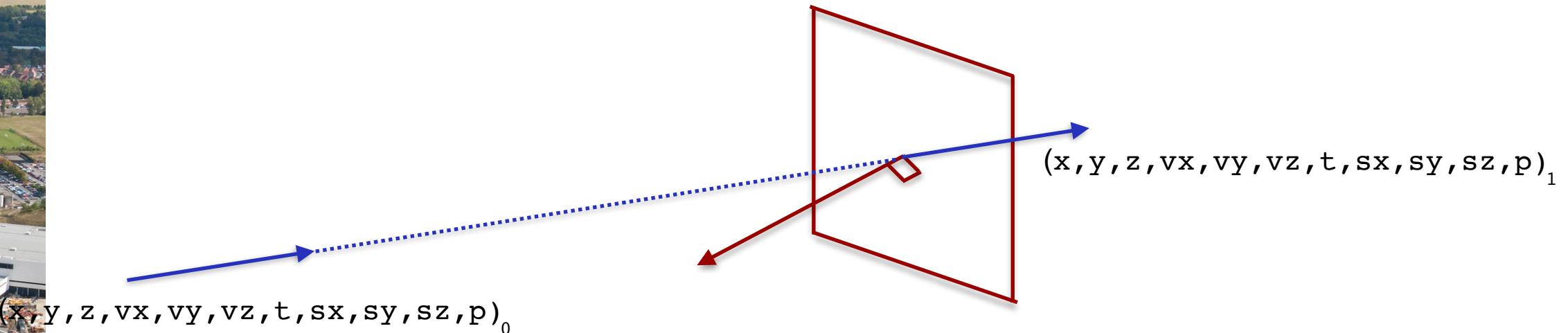
$$(x, y, z, vx, vy, vz, t, sx, sy, sz, p)_0$$



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# Neutron (ray)-matter interaction 1: reflecting surface

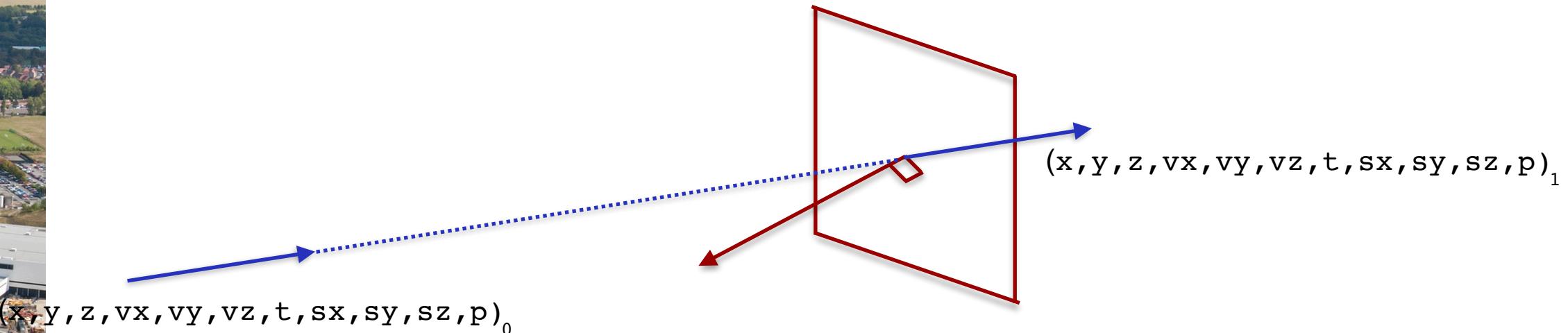
- 2. Propagate to the mirror surface





# Neutron (ray)-matter interaction 1: reflecting surface

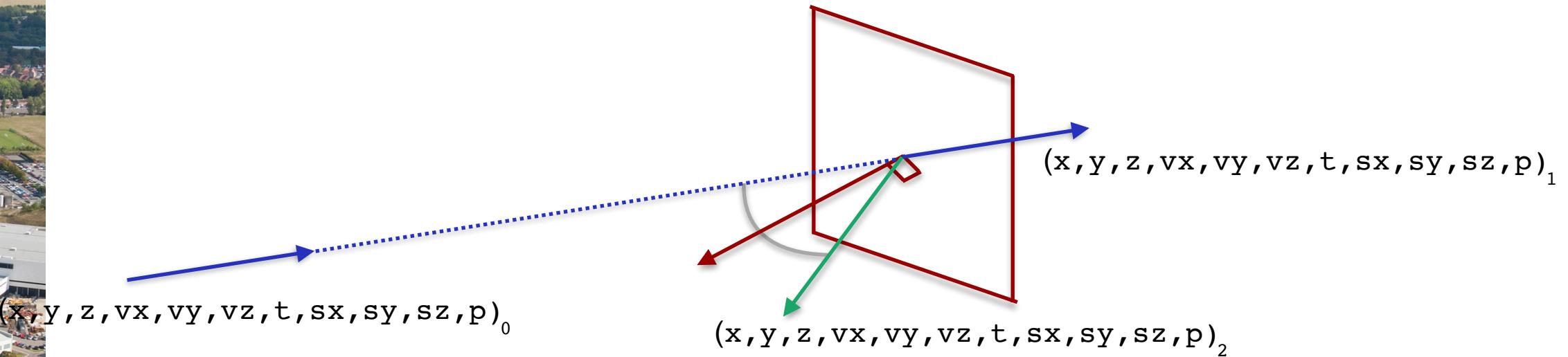
- 3. Checks (are we on surface, what is probability of reflection etc.)





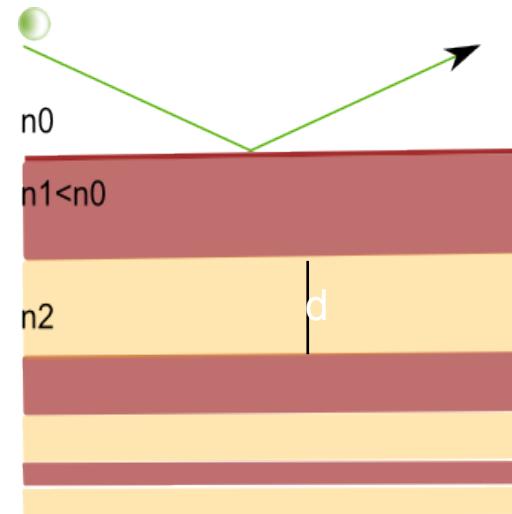
# Neutron (ray)-matter interaction 1: reflecting surface

- 4. Reflect





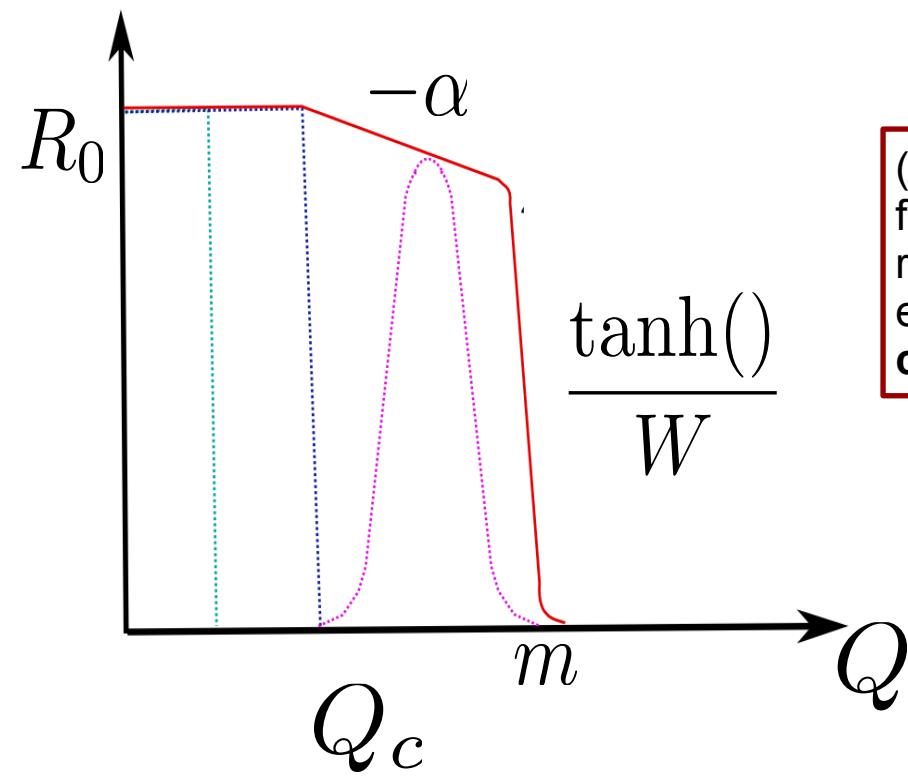
# Parametrisation of reflectivity on mirrors etc.



$$V = \frac{2\pi\hbar^2}{m} bN \quad \sin\theta < \sqrt{\frac{mV}{2\pi^2\hbar^2}}\lambda$$

$$m = \frac{\theta_{mirror}}{\theta_{Ni}}$$

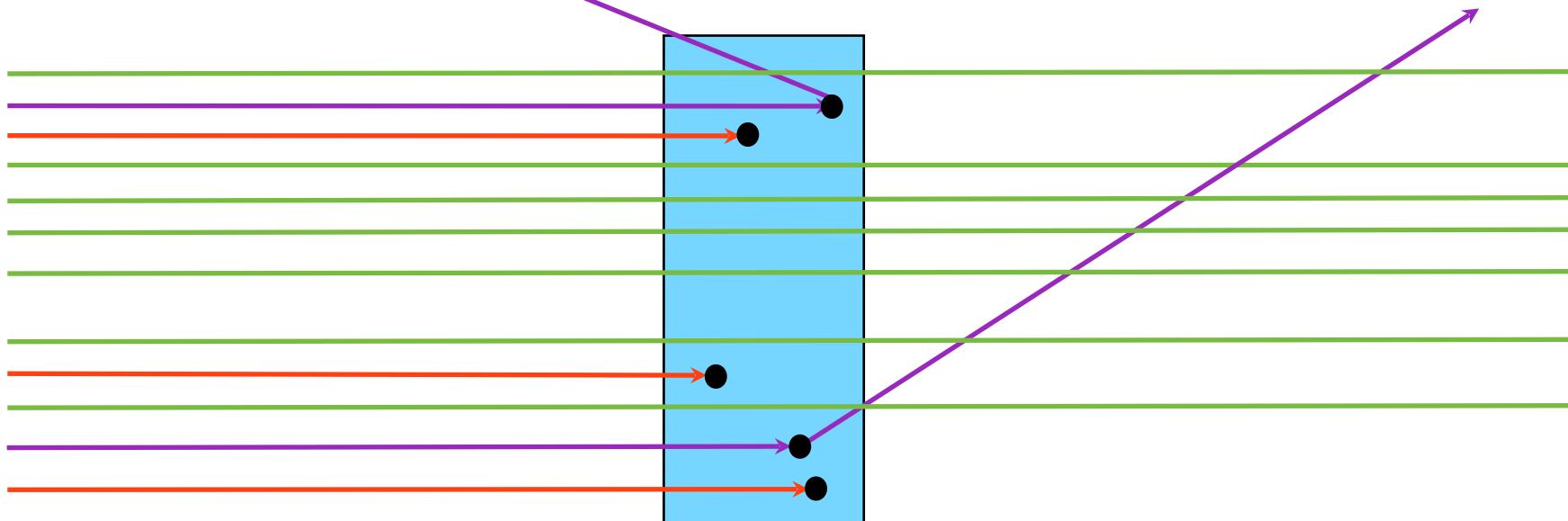
$$R_0 \cdot \left(1 - \frac{\tanh(Q - mQ_c)}{W}\right) \cdot (1 - \alpha(Q - Q_c))$$



(i.e.  $Q$  is calculated for given neutron, reflectivity encoded in **changed  $p$  value**)



# Neutron (ray)-matter interaction in General



absorbed, transmitted, or scattered



absorption transmission

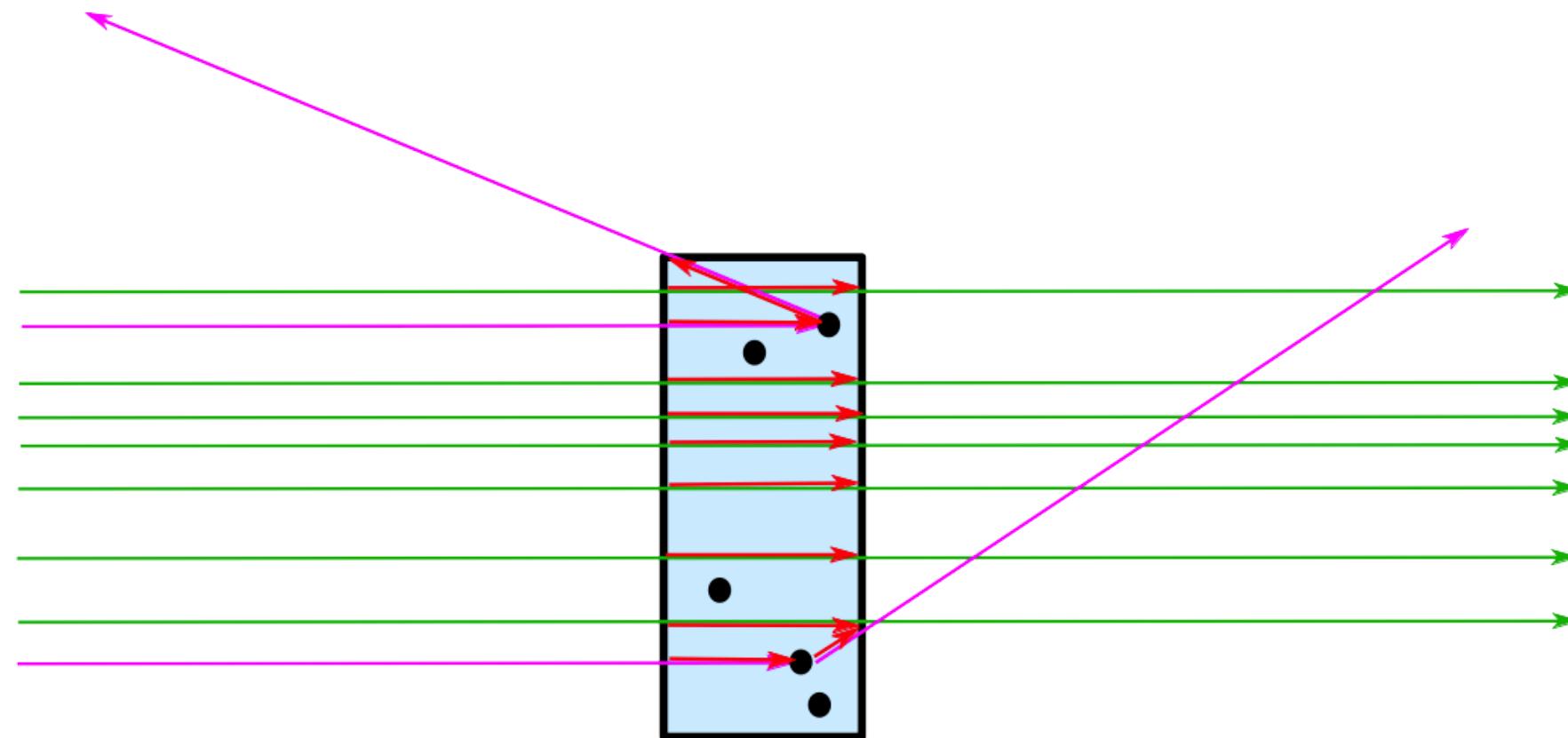
$$p_A = (1 - e^{-\Sigma_T t})(\Sigma_A / \Sigma_T)$$

$$p_S = (1 - e^{-\Sigma_T t})(\Sigma_S / \Sigma_T)$$

$$p_T = 1 - p_S - p_A = e^{-\Sigma_T t}$$

$$\Sigma_* = \rho \sigma_*$$

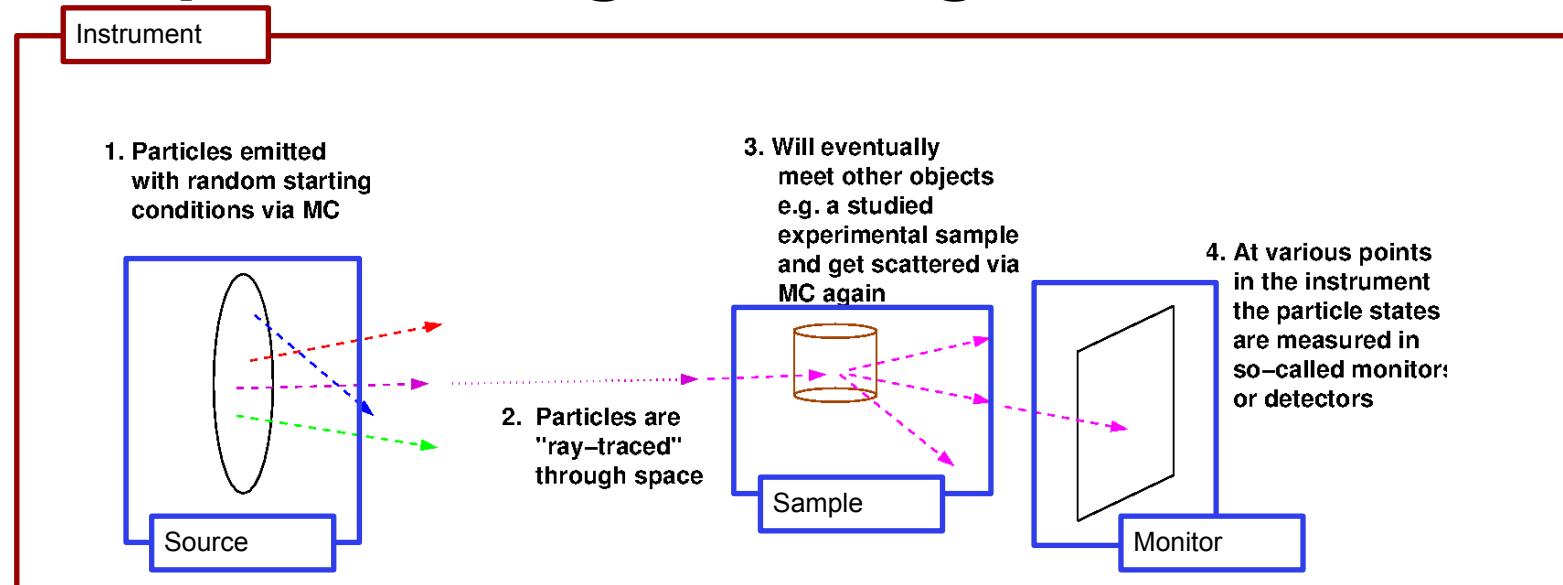
1 barn =  $10^{-24} \text{cm}^2$



transmitted+absorption, or scattered+absorption



# Transport of weight through the instrument...



$p_0$

$p_j$

$p_n$

$$p_j = w_j p_{j-1}$$

$$p_j = p_0 \prod_{k=1}^j w_k$$

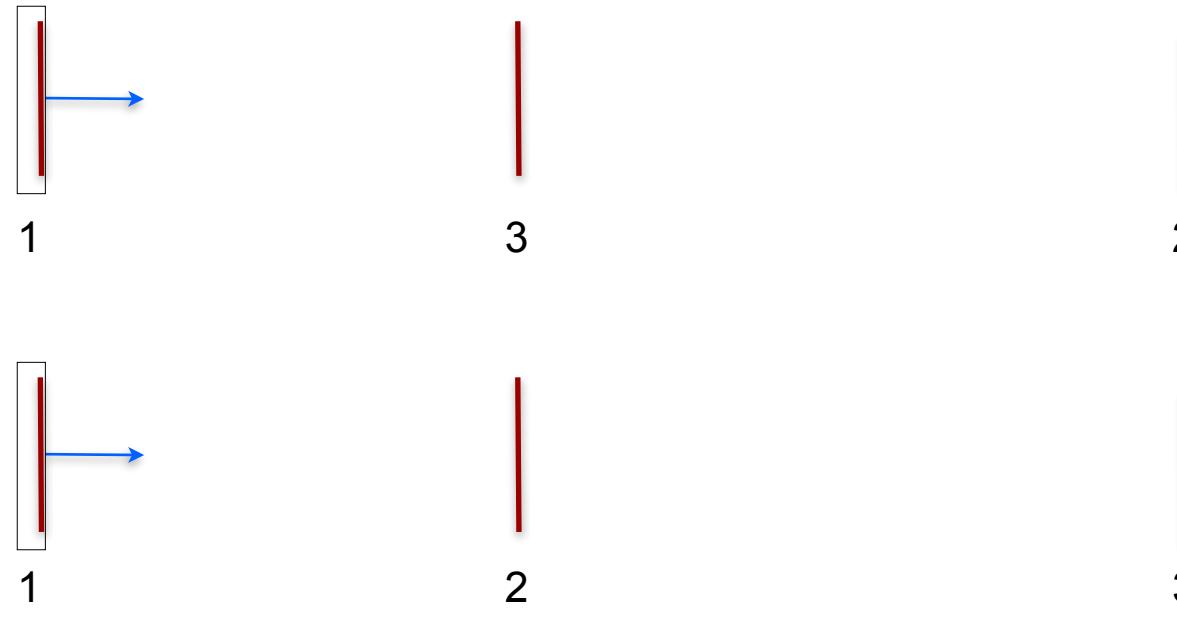
The weight multiplier of the  $j$ 'th component,  $w_j$ , is calculated by the probability rule  $f_{MC,b}w_j = P_b$  where  $P_b$  is the physical probability for the event "b", and  $f_{MC,b}$  is the probability that the Monte Carlo simulation selects this event.

In case of "branching", i.e. multiple outcomes, it is clear that

$$\sum_b f_{MC,b} = 1$$

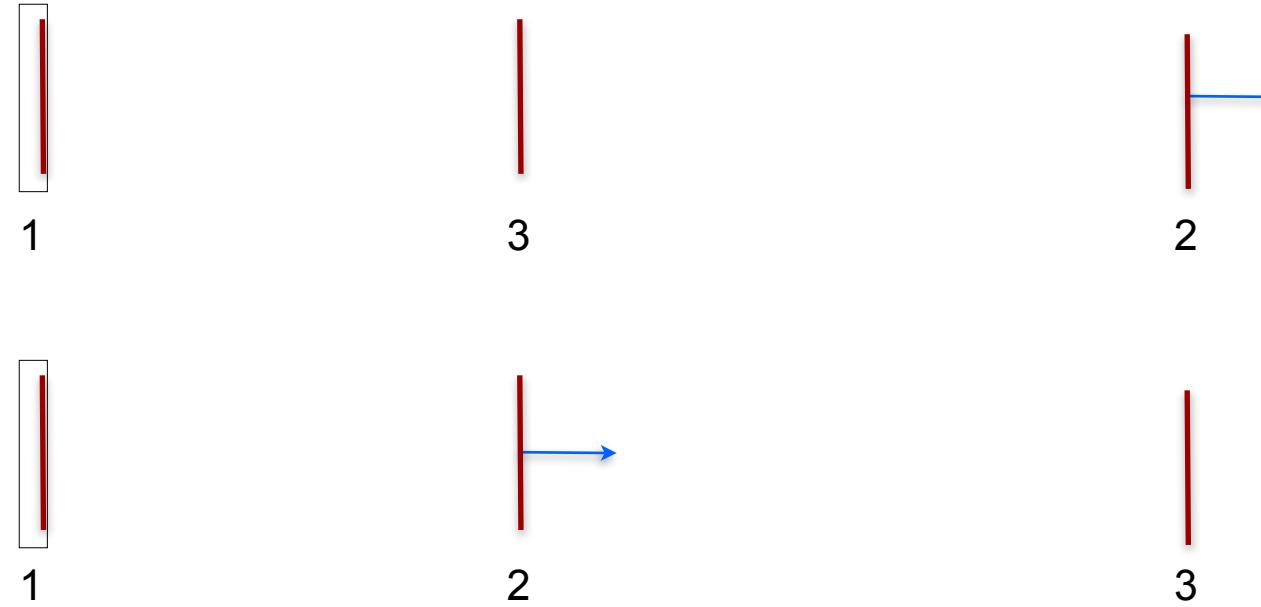


# To first order, McStas is linear and follows sequence of components in your file...





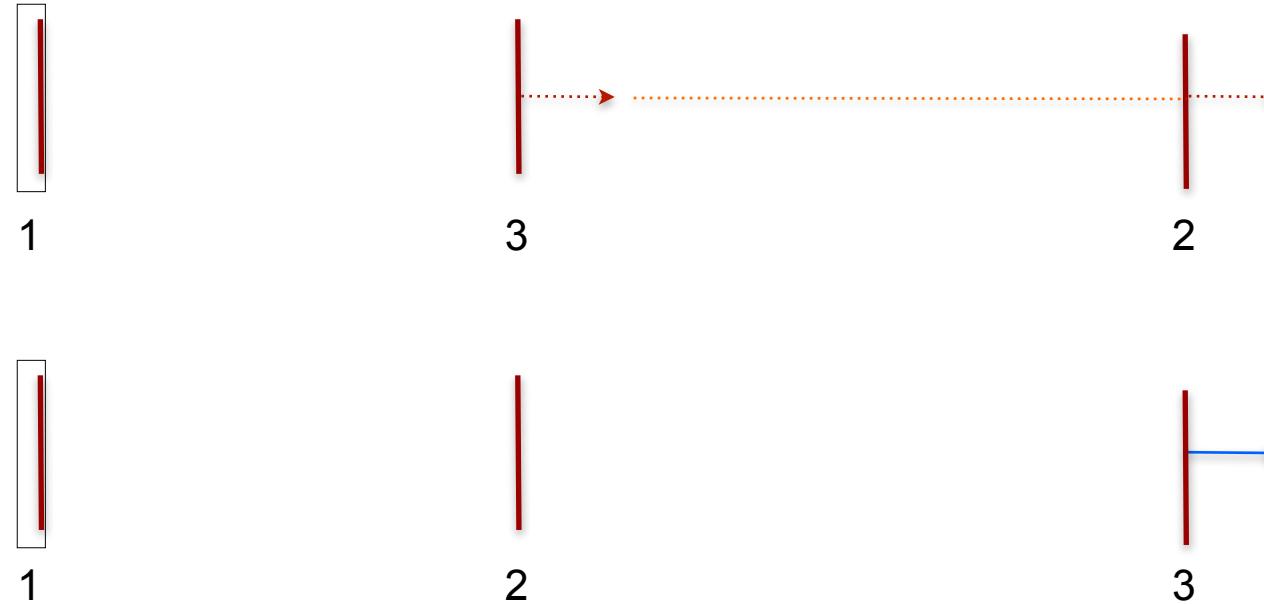
To first order, McStas is linear and follows sequence of components in your file...



Moving to first comp in the list



# To first order, McStas is linear and follows sequence of components in your file...



Moving to 3rd comp in list requires “moving back in time”.

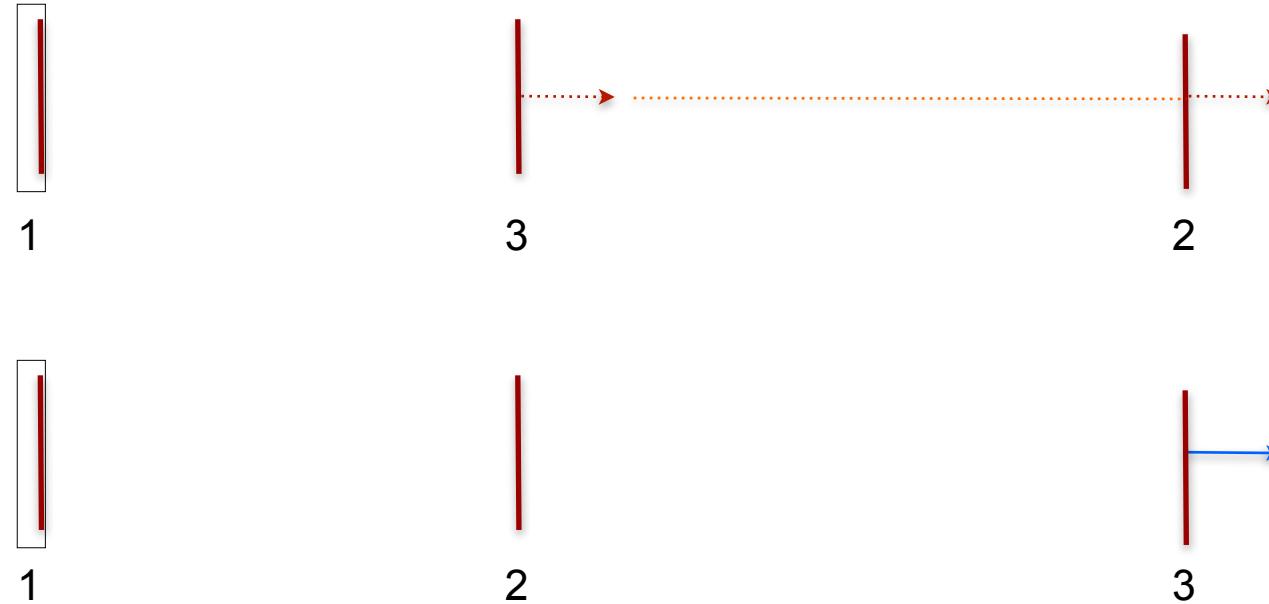
Default behavior is to ABSORB this type of neutron.

For monitors use `restore_neutron=1` in this case.

For homegrown comps use `ALLOW_BACKPROP` macro.



# To first order, McStas is linear and follows sequence of components in your file...



Moving to 3rd comp in list requires “moving back in time”.

Default behavior is to ABSORB this type of neutron.

For monitors use `restore_neutron=1` in this case.

For homegrown comps use `ALLOW_BACKPROP` macro.



**The order of components is important, and in general overlaps should be avoided!**



# Units used - and differences to Vitess

- Generally SI-Units, e.g. **meters** and **seconds** etc.
- Added neutron-scattering meaningful quantities of **E[meV]**,  **$\lambda[\text{\AA}]$**  and cross sections in  **$\sigma$  [barns]**

Main differences / difficulties translating between

- Different length-units for placement / sizes

McStas



[m]

Vitess

[cm]

- Different propagation-coordinate system

z || beam  
y vertical  
x  $\perp$  beam

x || beam  
z vertical  
y  $\perp$  beam

- McStas explicitly and automatically contains “free space”, Vitess has this “inside” the modules or by adding “spacewindow”
- Sources in McStas always propagates “by virtual window”, i.e. does not itself propagate



# McStas $\leftrightarrow$ Vitess compatibility features

- MCPL particle list is supported from McStas 2.3- (2016) and Vitess 3.4 (2018)
- McStas includes Vitess\_output and Vitess\_input, but better use MCPL
- mcstas2vitess can be used to port McStas components to Vitess
- Vitess\_ChopperFermi is the Vitess Fermi chopper ported to McStas

Friendly competition and collaboration for 1.5 decades! - e.g. Klaus Lieutenant was part of McStas team for 2 years