

1a) $3.3708567 \times 10^{-19} \text{ J}$

b) $5.9332098 \times 10^{20} \text{ photons}$

2a) $7.4245799 \times 10^{-19}$

b) 4.6340583 eV

3) 3.38945138 m/s

4) Consider a blackbody radiator at temperature T. We know the energy density, $\frac{E_{total}(T)}{V}$, is given by

$$\int_0^\infty \rho(\nu, T) d\nu = \int_0^\infty \frac{8\pi h \nu^3}{c^3 (e^{\frac{h\nu}{k_B T}} - 1)} d\nu. \quad \text{Let } x = \frac{h\nu}{k_B T}, \text{ then we have } dx = \frac{h}{k_B T} d\nu \Rightarrow d\nu = \frac{k_B T}{h} dx \text{ and}$$

then $x^3 = \left(\frac{h}{k_B T}\right)^3 \nu^3 \Rightarrow \nu^3 = \left(\frac{h}{k_B T}\right)^3 x^3$. Substituting these formulae into the original equation we arrive at

$$\frac{E_{total}(T)}{V} = \frac{8\pi h}{c^3} \int_0^\infty \frac{k_B^3 T^3 h^{-3} x^3}{e^x - 1} (k_B T h^{-1}) dx = \frac{8\pi k_B^4 T^4}{(hc)^3} \int_0^\infty \frac{x^3}{e^x - 1} dx = \left(\frac{8\pi k_B^4 T^4}{(hc)^3}\right) \left(\frac{\pi^4}{15}\right) = \sigma T^4 \text{ where}$$

$$\sigma = \left(\frac{8}{15}\right) \frac{\pi^5 k_B^4}{(hc)^3} \text{ is a constant. Therefore the energy depends on temperature as } T^4.$$

5a) $5.2728591 \times 10^{-26} \text{ J}$

b) $0.0317538996 \text{ J} \cdot \text{mol}^{-1}$

6a) 6329.9109 \AA

b) $15798.011 \text{ cm}^{-1}$

c) 299709770 m/s

7a) For X-Rays of wavelength $3.00 \times 10^{-10} \text{ m}$ (3.00 \AA) incident at an angle of $90^\circ = \frac{\pi}{2}$ radians the change in wavelength is described by the equation

$$\Delta \lambda = \frac{(1 - \cos \theta) h}{m_e c} = \frac{(1 - \cos(\frac{\pi}{2})) h}{m_e c} = \frac{h}{m_e c} = 2.4263102 \times 10^{-12} \text{ m} \text{ so the final wavelength of the scattered}$$

$$\text{X-Rays is } \lambda_{final} = \lambda_{initial} + \Delta \lambda = 3.0242631 \times 10^{-10} \text{ m}.$$

b) The initial energy of the incoming X-Ray photon is given by $E = h\nu = \frac{hc}{\lambda} = 6.6214862 \times 10^{-16} \text{ J}$, by the same relation we know that the final energy of the scattered photon is $6.5683632 \times 10^{-16} \text{ J}$ and because energy is conserved (even if momentum is not because this is an inelastic collision) we know that the energy lost by the scattered X-Ray photon will be equal to the kinetic energy gained by the scattered electron. Thus we

have $KE_{electron} = \Delta E_{photon} = E_f - E_i = 5.3122957 \times 10^{-18} J = \frac{1}{2} m_{electron} v^2 \Rightarrow v = \sqrt{\frac{2(KE_{electron})}{m_{electron}}}$ which is equal
to $3.41516443 \times 10^6 \frac{m}{s}$.

8a)