

General LSI imaging system

$p(x)$ = unknown source image

$$I(x) = \int_{-\infty}^{\infty} p(u) \cdot \frac{1}{r(u, x; z)} p(u) du$$

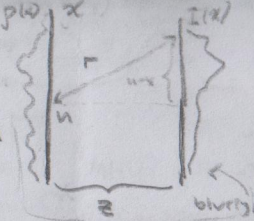
$$w(u, x; z) = ((u-x)^2 + z^2)^{-1/2}$$

$$I(x) = (p * h)(x) \quad w/h(k) = (k^2 + z^2)^{-1/2}$$

$$FWHM: \frac{1}{2}(z^2)^{-1/2} = (\bar{x} + \bar{z})^{-1/2}$$

$$\bar{x} = \pm \sqrt{z^2(2/p-1)}$$

$$\Rightarrow FWHM = 2(\bar{z}) = 2 \times \sqrt{z^2(2/p-1)}$$



Cascade N-LSI sys.

Single PSF $g_n(x) = e^{-\pi(x/w_n)^2}$

overall spatial response is then

$$h(x) = g_1(x) * g_2(x) \dots * g_N(x)$$

$$H(k) = \prod_{i=1}^N |w_i| e^{-\pi(k w_i)^2}$$

$$\dots = \left[\prod |w_i| \right] \times \exp\{-\pi k^2 \sum w_i^2\}$$

$$\therefore FWHM = \sqrt{\sum w_i^2}$$

$$\therefore h(x) = \frac{w_1 \dots w_N}{\sqrt{\sum w_i^2}} N(x/\sqrt{\sum w_i^2})$$

Deconvolution in k-space Filters

Additive noise: $m(x) = (F * h)(x) + n(x)$

$$M(k) = F(k) H(k) + N(k)$$

\Rightarrow Dividing both by $H(k)$ will lead to $N(k)/H(k)$ term which grows up

For big k $\Rightarrow H(k) \rightarrow 0$

\Rightarrow Could crop when noise dominates, but this'll marginally improve SNR ($\approx 50\%$) but not resolution;

- FWHM improved by high-freq, which'll narrow transfer Fctn
- But noise dominates high-freq, and noise SNR limits how wide we can crop

\Rightarrow Fundamentally ill-posed problem; could infer high- k values using constraints, but this means the output is no longer determined by just data measured (possibly introducing artifacts)

WTIS $\delta(x) = |a| \delta(x)$

Inter: scale width effectively, center height

Sampling Fctn with $\delta(x)$:

$$\int_{-\infty}^{\infty} F(x) \delta(x/a) dx = \int_{-\infty}^{\infty} F(ax) \delta(y) \times a dy = a F(ax)|_{y=0} = a F(0)$$

$$= a \int_{-\infty}^{\infty} F(x) \delta(x) dx$$

IF $a < 0$ then the integration limits are reversed, so $\delta(x) = -a \delta(x) \square$

Sampling

$$F_n = \int_{n-\frac{1}{2}\Delta}^{n+\frac{1}{2}\Delta} F(x) dx$$

$$= \int_{-\infty}^{\infty} F(x) \Pi\left(\frac{x-n\Delta}{\Delta}\right) dx$$

sub: $\Pi(u) = \Pi(-u)$

$$= F(n\Delta) * \Pi\left(\frac{x-n\Delta}{\Delta}\right)$$

$$= [F(x) * \Pi\left(\frac{x}{\Delta}\right)]_{x=n\Delta} = (F(x) * \Pi\left(\frac{x}{\Delta}\right)) * \delta(x-n\Delta)$$

$$F_s[x] = \sum_{n=-\infty}^{\infty} F_n$$

$$= [F(x) * \Pi\left(\frac{x}{\Delta}\right)] * \sum_{n=-\infty}^{\infty} \delta(x-n\Delta)$$

$$\approx \int_{\text{Sampling area}} \frac{1}{\Delta} \sum \delta\left(\frac{x-n\Delta}{\Delta}\right) W(x/a) dx$$

$$F_s[k] = [F(k) * \Delta \text{sinc}(k\Delta)] * \frac{1}{\Delta} W(k\Delta) \Delta$$

$$= \Delta \sum_m F(k - \frac{m}{\Delta}) \text{sinc}\left(\frac{k-m}{\Delta} \Delta\right)$$

\Rightarrow Copies every $1/\Delta$

No overlap if $2K_0 \leq 1/\Delta$

$F(k)$ bandlimit is $\pm K_0$

Recovery:

$$F(n) = F_s[k] \cdot \Pi(k\Delta) / \Delta \text{sinc}(k\Delta)$$

Constraints:

Measured: Using phantom, we can approximate $\delta(x)$ and thus $h(k)$.

\Rightarrow This informs us of the system-level contributions to blur (hardware, geometry, radiation, etc) and thus how could improve our sys \Rightarrow images

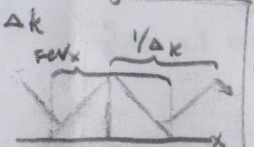
\Rightarrow Can also measure Fov, letting us determine spatial sampling rate

Impaired: Depend on judgement

- k_1/k_2 are a priori math constraints, rather than calibration or physically motivated constraints (e.g. Maxwell's eqn holds, $\mu \geq 0$ for x-ray, $\text{img} \in \mathbb{R}$)
- Inherit real-world effects, i.e. magnetic field imperfections, may make these physical assumptions invalid
- Leads to False credibility: MD may assume img determined 100% by data, when $\exists \infty$ solns and we choose one based on (arbitrary) weights chosen

FWHM Heuristic: $\Delta_x \approx \frac{1}{2} FWHM(h)$

Nyquist Sampling: $\Delta_k = \frac{1}{Fov_x}$



Ex: 1x1mm res over 10x10cm Fov

$$\Delta_k = \frac{1}{10\text{cm}} = \frac{1\text{cm}}{10\text{mm}} = \frac{1}{100}\text{mm}^{-1}$$

$$N_x = \frac{K_0}{\Delta_k} = \frac{1\text{mm}^{-1}}{1/100\text{mm}^{-1}} = 100$$

\therefore Total samples = 10,000 inside $K_0 \times K_0$ box in k-space

Ex: Sampling $m(k) = H * F(k) + N(k)$ in k-space

$F(k)$ bw $L \Rightarrow \Delta_k = 1/L$ (relying $\Delta_k \leq 1/L$)

$F(k)$ bw $K_0 \Rightarrow N = \frac{K_0}{\Delta_k} = L K_0$

Undersample: k-space bw = Fov_x

Sampling in k-space at interval Δ_k

Alias region = $Fov_x = Fov_x - \frac{1}{\Delta_k}$

Non-0 $\Rightarrow Fov_{NA} = \frac{1}{\Delta_k} - Fov_x$

Nyq: $Fov_x = 1/\Delta_k \Rightarrow \frac{2}{\Delta_k} - Fov_x$

LST IF sys LST, they can only blur input image

1] Lin: $\mathcal{L}\{\alpha f_1 + \beta f_2\} = \alpha g_1 + \beta g_2$
 2] SI: $\mathcal{L}\{f(x-\tau)\} = g(\tau) \Leftrightarrow \mathcal{L}\{f(x-\tau)\} = g(\tau-\tau)$

Ex: $\mathcal{L}\{f\} = F(x)$
 L: $\mathcal{L}\{\alpha f_1 + \beta f_2\} = (\alpha f_1 + \beta f_2)' = \alpha f_1' + \beta f_2' \neq \alpha f_1' + \beta f_2'$

SI: $\mathcal{L}\{f(x-\tau)\} = F(2x) - \text{shrink by 2}$
 $\mathcal{L}\{f(x-\tau)\} = F(2(x-\tau)) = F(2x-2\tau)$

Ex: $\mathcal{L}\{f\} = f'(x)$
 LI: $\mathcal{L}\{\alpha f_1 + \beta f_2\} = \frac{d}{dx}[\alpha f_1(x) + \beta f_2(x)] = \alpha f_1'(x) + \beta f_2'(x)$
 SI: $\mathcal{L}\{f(x-\tau)\} = \frac{d}{dx} F(x-\tau) = \frac{d}{dx} F(x) = f'(x)$

Ex: Let input $u(x)$, output $= F(x)$
 IF LSC sys. is invertible, it's inverse must also be LST (swap in & out)

Ex: $\frac{dF}{dx} + bF(x) = u(x)$
 L: $\mathcal{L}\{x f_1 + x f_2\} = \frac{d}{dx} [x f_1 + x f_2] = x f_1' + x f_2' + f_1 + f_2$
 $\mathcal{L}\{f_1\} = u_1, \mathcal{L}\{f_2\} = u_2$

SI: $\mathcal{L}\{h(x)\} = \frac{d}{dx} (h(x) + b h(x))$ for $h(x) = f(x-\tau)$
 $= \frac{d}{dx} f(x-\tau) + b f(x-\tau) = u(x-\tau) = \mathcal{L}\{f(x)\}$

Ex: $\frac{dF}{dx} = -b u(x) F(x) \equiv u(x) = \frac{dF}{dx} / -b F(x)$
 L: $\mathcal{L}\{x f_1 + x f_2\} = \frac{d}{dx} [x f_1 + x f_2] = x f_1' + x f_2' + f_1 + f_2$
 SI: $\mathcal{L}\{f(x-\tau)\} = \frac{d}{dx} [h(x)] / -b h(x) = \frac{d}{dx} f(x-\tau) / -b f(x-\tau) = u(x-\tau) = \mathcal{L}\{f(x)\}$

$\mathcal{L}\{F(x)\} = \int F(x) h(x-u) du = (F * h)(x)$
 $= \lim_{\Delta \rightarrow 0} \sum_n F(n\Delta) \mathcal{L}\left\{\frac{1}{\Delta} \Pi\left(\frac{x-n\Delta}{\Delta}\right)\right\} \Delta$

Shift: IF $g(x) = (F * h)(x)$
 then $g(x-\tau) = F(x-\tau) * h(x-\tau)$
 $\Rightarrow \int F(u-\tau) h(x-\tau-u) du$
 Let $x \Rightarrow x-\tau-(x-\tau)$
 $\Rightarrow \int F(u) h((x-\tau)-u) du$

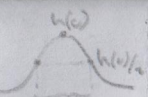
Scale: $|b| K(\frac{x}{b}) = (F * g)(\frac{x}{b})$
 as $u = \int F(u/b) g(\frac{x}{b} - u/b) du$
 Let $u \Rightarrow \frac{x}{b} - l; dl = \frac{1}{b} du$
 $\Rightarrow \int F(l/b) g(\frac{x}{b} - l/b) |b| dl$

Area: $\int (F * g)(x) dx = \int [\int F(u) g(x-u) du] dx$
 swap $\Rightarrow \int [F(u) \int g(x-u) dx] du$
 $\dots = \int F(u) \int g(l) dl du = \int F(u) du \int g(l) dl$

Fourier Properties:
 Defn: $F(x) = \int_{-\infty}^{\infty} F(u) e^{i2\pi kx} dx$
 $F(k) = \int_{-\infty}^{\infty} F(x) e^{-i2\pi kx} dx$
 Shift: $F(x-\tau) \Leftrightarrow e^{-i2\pi k\tau} F(k)$
 Scale: $f(x/a) \Leftrightarrow |a| F(k)$

$\frac{1}{\omega} \Pi(\frac{x}{\omega}) \Leftrightarrow \text{sinc}(\omega k) \mid W(x) \Leftrightarrow W(k)$
 $\exp(-\pi(\frac{x}{\omega})^2) \Leftrightarrow |w| \exp(-\pi(kw)^2) \equiv \frac{1}{W} N(\frac{k}{W})$
 $\delta(x-\tau) \Leftrightarrow 1 \cdot \exp(-j2\pi\tau k)$
 $\cos 2\pi\tau x \Leftrightarrow \frac{1}{2} [\delta(k-\tau) + \delta(k+\tau)]$
 $\sin 2\pi\tau x \Leftrightarrow \frac{1}{2j} [\delta(k-\tau) - \delta(k+\tau)]$

Alt: Let $y = f(x)$, for $\frac{dy}{dx} = -b u(x) y$
 LI: $\mathcal{L}\{h(x)\} = \int y' dy = \int -b u(x) y dx$
 $= \exp(-b \int u(x) dx) = \exp(-b \int u(x) dx)$
 $= \exp(k) \mathcal{L}\{u_1\} = \exp(k) \mathcal{L}\{u_2\}$
 $\neq \alpha_1 F_1(x) + \alpha_2 F_2(x)$

Rayleigh/Heurston Criteria: 
 Two sources are resolvable iff they're separated by more than FWHM, else they blur together, are indistinguishable

\hookrightarrow Heuristic of "real-world" delta FWHM should be at-most half width of FWHM
 Ex: Expected resolution 200nm \Rightarrow 210nm-wide photon
 IF we detect $I(x) = \exp[-\mu x]$ then share $1/\mu \approx 200$ nm

Over-sampling (beyond Nyquist) gains no novel information; blurring occurs before sensors - cannot recover what's already lost, due to falloff

Gauss FWHM: $\frac{1}{2} = \exp(-\pi(\frac{x}{W})^2) \Rightarrow N(\frac{x}{W})$
 $\Rightarrow -\ln 2 = -\pi(\frac{x}{W})^2 \Rightarrow x = W \sqrt{\ln 2 / \pi}$
 $\therefore \text{FWHM} = 2|x| \approx 0.94 W \approx W$

Delta Properties: $\delta(x) * h(x) = h(x)$
 $\delta(x-\tau) * h(x) = h(x-\tau)$
 $\int \delta(x-\tau) f(x) dx = f(\tau)$
 $\delta(x) = \lim_{\omega \rightarrow \infty} \frac{1}{\omega} \Pi(\frac{x}{\omega})$
 $\delta(x) = \lim_{\omega \rightarrow \infty} \frac{1}{\omega} N(\frac{x}{\omega})$
 $\mathcal{L}\{\delta(x)\} = \int \delta(x) e^{-i2\pi kx} dx = e^{-i2\pi k \cdot 0} = 1$

Let $S(x) = W(\frac{x}{\Delta}) \Leftrightarrow S(k) = \Delta W(k\Delta)$
 $S(k) = \sum_n \delta(\frac{x}{\Delta} - n) = \sum_n \delta(x - n\Delta)$
 $= \sum_n \delta(x - n\Delta)$
 $S(x) = \Delta \sum_n \delta(x - n\Delta) e^{i2\pi kx}$
 $\hookrightarrow S(x)$ periodic w/ $1/\Delta$
 $\hookrightarrow S(0) \rightarrow \infty, S(k) \rightarrow 0$ for $k \neq \frac{p}{\Delta}$

Argument to FWHM must be dimensionless; ex rect(x) is actually $\Pi(x/1\text{mm}) = \begin{cases} 1; & |x/1\text{mm}| < 1/2 \\ 0, & \text{o/w} \end{cases}$

Discretization:

ctns int	disc. sum
$\int_a^b u du$	$\sum_{n=a}^b u \Delta x$

 $(f * h)(x) = \sum_n F(n\Delta x) \Delta x \underbrace{h(x - n\Delta x)}_{\text{mod}}$

Units: $H(k)$ is ratio, unitless (V_m/V_{int})
 $h(x) = 1/\text{mm} = \delta(x)$; $V(k) = V \cdot \text{mm}$

Solving FWHM: $\frac{1}{2} h(x_{\text{FWHM}}) = h(x_{\text{FWHM}}/2)$

Sum of Gauss: Let $w_i = \frac{1}{w_i} N(\frac{x}{w_i})$
 $H(k) = H_1 * H_2 = \exp(-\pi k^2 (w_1^2 + w_2^2))$
 $h(x) = \frac{1}{\sqrt{w_1^2 + w_2^2}} N(\frac{x}{\sqrt{w_1^2 + w_2^2}})$
 $\therefore \text{FWHM} = \sqrt{w_1^2 + w_2^2}$
 \hookrightarrow Dominated by blurrer sys. (narrowest in k-space)

Averaging improves SNR $\propto \sqrt{N}$
 Ex: Want to downscale 1mm $\rightarrow \frac{1}{4} \times \frac{1}{4}$ mm while maintaining SNR
 SNR \propto Area, so decreases by 16
 Best SNR by avg'g $16^2 = 256$ samples

IF we can reduce # samples read, then we could increase the sampling rate; this would tackle aliasing in frequency
 \hookrightarrow This is done in X-ray CT, ignoring higher freqs and reducing radiation exposure
 \hookrightarrow Broadest structures in low freqs; aliasing of noise in k-space will be mostly high freqs