

Interpretable Gaussian Processes for Stellar Light Curves

RODRIGO LUGER^{1,2,*}

¹*Center for Computational Astrophysics, Flatiron Institute, New York, NY*

²*Virtual Planetary Laboratory, University of Washington, Seattle, WA*

Keywords: methods: analytic

1. THE SPOT EXPANSION

1.1. Preliminaries

We adopt the following expression for the spherical harmonic coefficient of degree l and order m in the expansion of a spot at $\theta = \varphi = 0$:

$$y_{lm}(r, \delta) = \begin{cases} 1 - \frac{\delta cr}{2(1 + cr)} & l = m = 0 \\ -\frac{\delta cr (2 + cr)}{2\sqrt{2l+1}(1 + cr)^{l+1}} & l > 0, m = 0 \\ 0 & m \neq 0 \end{cases} \quad (1)$$

where $\delta \in [-\infty, 1]$ is the spot contrast (the fractional decrease in the brightness at the center of the spot) and $r \in (0, 1]$ is the normalized spot radius. The quantity $c \in (0, \infty)$ is a normalization constant for the radius (more details below). The expression in Equation (1) is convenient because it satisfies four important properties:

- 1 The surface intensity monotonically increases away from the spot center
- 2 The surface intensity at the spot center is $1 - \delta$
- 3 The surface intensity at the antipode of the spot center is unity
- 4 The size of the spot increases monotonically with r

These properties may be demonstrated by considering the expression for the surface intensity at a given point (θ, φ) :

$$\begin{aligned} I(\theta, \varphi) &= \sum_{l=0}^{\infty} \sum_{m=-l}^l y_{lm} Y_{lm}(\theta, \varphi) \\ &= \sum_{l=0}^{\infty} y_{l0} \sqrt{2l+1} P_l(\cos \theta) \end{aligned} \quad (2)$$

where P_l is the Legendre polynomial of degree l and we have implicitly assumed a normalization such that the integral of our expansion over the unit sphere is 4π . Combining this with Equation (1) and rearranging, we may write

$$I(\theta, \varphi) = 1 + \frac{\delta cr}{2} - \frac{\delta cr(2+cr)}{2(1+cr)} \sum_{l=0}^{\infty} \left(\frac{1}{1+cr} \right)^l P_l(\cos \theta). \quad (3)$$

The summation in Equation (3) has a closed-form expression in terms of the generating function of the Legendre polynomials:

$$\sum_{l=0}^{\infty} t^l P_l(\cos \theta) = \frac{1}{\sqrt{1+t^2-2t\cos\theta}}, \quad (4)$$

so we may express the intensity in the fairly simple form

$$I(\theta, \varphi) = A - \frac{B}{\sqrt{C - \cos \theta}}, \quad (5)$$

where

$$\begin{aligned} A &= 1 + \frac{\delta cr}{2} \\ B &= \delta cr(2+cr) \sqrt{\frac{1}{8+8cr}} \\ C &= \frac{1+(1+cr)^2}{2+2cr} \end{aligned} \quad (6)$$

are positive constants.

Differentiating Equation (5) with respect to θ , we have

$$\frac{dI(\theta, \varphi)}{d\theta} = -\frac{B \sin \theta}{2(C - \cos \theta)^{\frac{3}{2}}}, \quad (7)$$

which is zero only for $\theta = 0$ (for which $I(\theta, \varphi)$ is minimized) and $\theta = \pi$ (for which it is maximized). The intensity therefore increases monotonically from the spot center to the antipode, as stated in 1. The value at the minimum is

$$\begin{aligned} I_{\min} &= A - \frac{B}{\sqrt{C-1}} \\ &= 1 - \delta, \end{aligned} \quad (8)$$

as stated in **2**, and the value at the maximum is

$$\begin{aligned} I_{\max} &= A - \frac{B}{\sqrt{C+1}} \\ &= 1, \end{aligned} \quad (9)$$

as stated in **3**. Finally, to show **4**, let us compute the half width at half minimum $\Delta\theta$ of the intensity profile:

$$1 - \left(A - \frac{B}{\sqrt{C - \cos \Delta\theta}} \right) = \frac{1}{2}\delta \quad (10)$$

Solving for $\Delta\theta$ yields

$$\Delta\theta = \cos^{-1} \left[\frac{2 + 3cr(2 + cr)}{2(1 + cr)^3} \right]. \quad (11)$$

Differentiation with respect to r yields

$$\frac{d\Delta\theta}{dr} = \frac{3c^2r(2 + cr)}{(1 + cr)^4 \sqrt{\frac{c^2r^2(2+cr)^2(1+2cr)(3+2cr)}{(1+cr)^6}}}, \quad (12)$$

which is positive definite for all $c, r > 0$.

1.2. Distributions over r and δ

We model the normalized spot radius r as a random variable drawn from a Beta distribution, whose PDF is given by

$$p(r | \alpha, \beta) = \frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} r^{\alpha-1} (1 - r)^{\beta-1}, \quad (13)$$

where Γ is the Gamma function and α and β are hyperparameters characterizing the shape of the distribution. Since we require the brightness $b = 1 - \delta$ at the center of the spot to be non-negative, it is convenient to parametrize the distribution of b instead of δ . We model b as a random variable drawn from a log-normal distribution:

$$p(b | \mu, \nu) = \frac{1}{b\sqrt{2\pi\nu}} \exp \left[-\frac{(\ln b - \mu)^2}{2\nu} \right], \quad (14)$$

where μ and ν are the mean and variance of the log-normal, respectively. The joint distribution over r and b is therefore separable, by construction:

$$p(r, b | \alpha, \beta, \mu, \nu) = p(r | \alpha, \beta) p(b | \mu, \nu). \quad (15)$$

1.3. First moment

The first moment of the distribution of y_{lm} over r and b is

$$\begin{aligned} \mathbb{E}[y_{lm}] &= \int_0^\infty \int_0^1 y_{lm}(r, 1-b) p(r, 1-b | \alpha, \beta, \mu, \nu) dr db \\ &= \int_0^\infty \mathbb{E}_r[y_{lm}] p(b | \mu, \nu) db, \end{aligned} \quad (16)$$

where

$$\mathbb{E}_r[y_{lm}] = \int_0^1 y_{lm}(r, 1-b) p(r | \alpha, \beta) dr. \quad (17)$$

Expanding Equation (17), we obtain

$$\mathbb{E}_r[y_{lm}] = \begin{cases} \frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \left(1 - \frac{(1-b)cr}{2(1+cr)}\right) r^{\alpha-1} (1-r)^{\beta-1} dr & l = m = 0 \\ -\frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \frac{(1-b)cr(2+cr)r^{\alpha-1}(1-r)^{\beta-1}}{2\sqrt{2l+1}(1+cr)^{l+1}} dr & l > 0, m = 0 \\ 0 & m \neq 0. \end{cases} \quad (18)$$

Thanks to the dependence of the spherical harmonic coefficients on only powers of r and $(1+cr)$, the integrals above may be expressed in closed form in terms of the hypergeometric function ${}_2F_1$:

$$\mathbb{E}_r[y_{lm}] = \begin{cases} 1 - \frac{(1-b)\alpha c}{2(\alpha + \beta)} F_l & l = m = 0 \\ -\frac{(1-b)\alpha c}{\sqrt{2l+1}(\alpha + \beta)} \left[F_l + \frac{(\alpha + 1)c}{2(\alpha + \beta + 1)} G_l \right] & l > 0, m = 0 \\ 0 & m \neq 0 \end{cases} \quad (19)$$

where we define

$$\begin{aligned} F_l &\equiv {}_2F_1(l+1, \alpha+1; \alpha+\beta+1; -c) \\ G_l &\equiv {}_2F_1(l+1, \alpha+2; \alpha+\beta+2; -c). \end{aligned} \quad (20)$$

Note that for a given value of α , β , and c , we need only compute F_0 , F_1 , G_0 , and G_1 directly, since the remaining terms may be obtained recursively:

$$\begin{aligned} F_l &= \frac{\alpha + \beta + 1 - l}{2l} F_{l-2} - \frac{2\alpha + \beta + 2 - 3l}{2l} F_{l-1} \\ G_l &= \frac{\alpha + \beta + 2 - l}{2l} G_{l-2} - \frac{2\alpha + \beta + 4 - 3l}{2l} G_{l-1}. \end{aligned} \quad (21)$$

We may now evaluate the integral in Equation (16):

$$E[y_{lm}] = \begin{cases} \frac{1}{b\sqrt{2\pi\nu}} \int_0^\infty \left(1 - \frac{(1-b)\alpha c}{2(\alpha+\beta)} F_l \right) \\ \quad \times \exp \left[-\frac{(\ln b - \mu)^2}{2\nu} \right] db & l = m = 0 \\ -\frac{1}{b\sqrt{2\pi\nu}} \int_0^\infty \frac{(1-b)\alpha c}{\sqrt{2l+1}(\alpha+\beta)} \\ \quad \times \left[F_l + \frac{(\alpha+1)c}{2(\alpha+\beta+1)} G_l \right] \\ \quad \times \exp \left[-\frac{(\ln b - \mu)^2}{2\nu} \right] db & l > 0, m = 0 \\ 0 & m \neq 0 \end{cases} \quad (22)$$

Most of the terms in the integrands above are constants, so these integrals may also be expressed in closed form:

$$E[y_{lm}] = \begin{cases} 1 - \frac{\gamma\alpha c}{2(\alpha+\beta)} F_l & l = m = 0 \\ -\frac{\gamma\alpha c}{\sqrt{2l+1}(\alpha+\beta)} \left[F_l + \frac{(\alpha+1)c}{2(\alpha+\beta+1)} G_l \right] & l > 0, m = 0 \\ 0 & m \neq 0 \end{cases} \quad (23)$$

where we define

$$\gamma \equiv 1 - \exp \left[\mu + \frac{1}{2}\nu \right]. \quad (24)$$

1.4. Second moment

The second moment of the distribution of y_{lm} over r and b is

$$\begin{aligned} \mathbb{E}[y_{lm}y_{l'm'}] &= \int_0^\infty \int_0^1 y_{lm}(r, 1-b)y_{l'm'}(r, 1-b)p(r, 1-b \mid \alpha, \beta, \mu, \nu)dr db \\ &= \int_0^\infty \mathbb{E}_r[y_{lm}y_{l'm'}]p(b \mid \mu, \nu)db, \end{aligned} \quad (25)$$

where

$$\mathbb{E}_r[y_{lm}y_{l'm'}] = \int_0^1 y_{lm}(r, 1-b)y_{l'm'}(r, 1-b)p(r \mid \alpha, \beta)dr. \quad (26)$$

Expanding Equation (26), we obtain

$$\mathbb{E}_r[y_{lm}y_{l'm'}] = \begin{cases} \frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \left(1 - \frac{(1-b)cr}{2(1+cr)}\right)^2 r^{\alpha-1}(1-r)^{\beta-1}dr & l = l', \\ & m = m' = 0 \\ \\ -\frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \left(1 - \frac{(1-b)cr}{2(1+cr)}\right) \\ \times \frac{(1-b)cr(2+cr)r^{\alpha-1}(1-r)^{\beta-1}}{2\sqrt{2l+1}(1+cr)^{l+1}}dr & l > 0, \\ & l' = 0, \\ & m = m' = 0 \\ \\ -\frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \left(1 - \frac{(1-b)cr}{2(1+cr)}\right) \\ \times \frac{(1-b)cr(2+cr)r^{\alpha-1}(1-r)^{\beta-1}}{2\sqrt{2l'+1}(1+cr)^{l'+1}}dr & l = 0, \\ & l' > 0, \\ & m = m' = 0 \\ \\ -\frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 \frac{(1-b)^2c^2r^2(2+cr)^2r^{\alpha-1}(1-r)^{\beta-1}}{4\sqrt{(2l+1)(2l'+1)}(1+cr)^{l+l'+2}}dr & l > 0 \\ & l' > 0 \\ & m = m' = 0 \\ \\ 0 & m, m' \neq 0. \end{cases} \quad (27)$$