## NUR hand-in 2

Meng Yao; s2308266

May 28, 2019

#### Abstract

The source code and the outputs of Numerical Recipes for Astrophysics Hand-in exercise 2 are shown in this report.

# 1 Question 1 : Normally distributed pseudorandom numbers

The shared modules

```
2 q1(a) main to show
5 Question 1: Normally distributed pseudo-random numbers
6 1(a) random number generator
7 Combine at least MWC and 64-bit shift
{\tiny 9~sys.stdout\ =\ open\,(\,\,'outputs1.txt\,\,'\,,\quad 'w')}
10 print ('Question 1 : Normally distributed pseudo-random numbers')
print('\n 1(a) Random number generator')
12 def generator(n, seed):
13
       combined number generator: \rm XOR\text{-}shift\ \hat{\ }MWC,\ also\ called\ Ranq2. In the text book, the
14
       parameters are given as A3(right)^B1.
       period = period of xorshift * period of MWC. Accoring to the
       \texttt{text} \ \texttt{book} = 8.5*10^37
       the parameters will be given later
       n = the amount of numbers
       seed = initial seed
19
20
       #parameters of each generator
       ##XOR-shift 64-bit
22
23
       XOR_a1=17
       XOR_a2=31
24
       XOR_a3=8
25
       bit64 = 2**64 - 1
       bit32 = 2**32-1
27
       ##MWC
28
       MWC_a = 4294957665
       #initial seed
30
31
       x = seed
32
       number = np.zeros(n)
33
       for i in range(n):
           #XORshift
35
```

```
x=x^{(x)} \times XOR_a1) x=x^{(x)} \times XOR_a2) & bit64 \# do a logical 'and' to cut
36
37
          the number to 64 bits.
               x = x
                           (x \gg XOR_a3)
38
               #MWC
39
               m = \left( \text{MWC.a*} \left( \text{m \& bit32} \right) + \left( \text{m} \right. > > 32 \right) \right) \quad \# \ \text{use all 64 bits of}
40
          updated state in bit mix
               #combine them
41
         mumber [i] = (x \hat{m})

#normalise in (0,1) /maxnumber of 64.

#Note that 'period' shows the repeating information (how long
42
43
44
          the sequence is), not the range of radom number
          number=np.array(number)/(2**64-1)
45
          return number
46
```

### 1.1 1.a Pseudo-random numbers

```
, ,,
2 q1(a) to show
3 ,,
4 #set the seed
 5 \text{ seed} = 123456789
_{6} print('seed = ', seed)
7 #first 1000 numbers and plot
 s n1k_uni = generator(1000, seed)
9 fig1 = plt.figure(1)
ax1_1 = fig1.add_subplot(2,1,1)
ax1_{-1} . scatter(n1k_uni[0:-2], n1k_uni[1:-1])
12 ax1_1.set_xlabel("$X_i$")
13 ax1_1.set_ylabel("$X_{i+1}$")
14 ax1_1.set_title('Sequential 1000 numbers')
ax1_{-2} = fig1.add_subplot(2,1,2)
16 ax1_2.plot(n1k_uni,'.')
17 ax1_2.set_title('1000 numbers vs indices')
18 ax1_2.set_xlabel('index')
19 #fig1.savefig("lalk.png")
20 fig1.tight_layout()
21 # 1 million numbers and plot
22 \text{ fig } 4 = \text{plt. figure } (4)
n1m\_uni=generator(10**6, seed)
24 ax4_1=fig4.add_subplot(1,1,1)
25 \text{ hist_n1m_uni} = \text{ax4_1.hist(n1m_uni, bins=np.linspace(0.0, 1.0, 21)},
       edgecolor='black') #plot the histogram and save the elements
26 ax4_1.set_xlabel('bins')
27 ax4_1.set_ylabel('quantity of numbers')
28 ax4_1.set_title('Histogram of 1 million numbers')
29 fig4.tight_layout()
30 #fig4.save('1alm.png')
31 fig4.show()
32 print ('figure of random number generator please see fig.1')
33 print ('roughly test the quality of RNG by show the largest and
       smallest number among bins:
print ('max = ',max(hist_n1m_uni[0]),'; min = ',min(hist_n1m_uni[0])
       )
      seed = 123456789
```

figure of random number generator please see fig.1

roughly test the quality of RNG by show the largest and smallest number among bins:

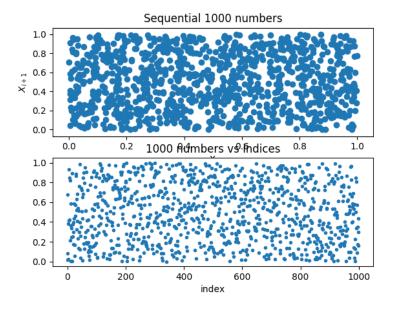


Figure 1: The results of the first 1000 numbers of RNG

 $\max = 50341.0$ ;  $\min = 49714.0$ 

## 1.2 1.b Normally distributed random number

```
, , ,
2 1(b) Normally distributed random number
3 generate normally distributed numbers whose mean=3 sigma=2.4.
  Then, compare them with Gaussian probability density function
  print('\n 1(b) Normally distributed random number')
  def Box_Muller(random_uni):
      in put a uniformly distributed random number sequence
      out put a nomarlly distributed one
10
      note that the sequence is still in [0,1)
11
12
      #split into 2 sequences
13
                                #sequence 1 is even-th elements
      u1 = random_uni[0::2]
      u2 = random_uni [1::2]
                                # odd-th elements
15
      coe = np. sqrt(-2*np. log(u1))
16
      s1 = coe*np.cos(2*np.pi*u2)
                                     # already normal
      s2 = coe*np.sin(2*np.pi*u2)
18
      random\_normal = np.concatenate([s1, s2])
19
      return \ random\_normal
20
21
_{22} def Gaussian_PDF(x,mu,sigma):
23
      define Gaussian fuction
24
25
      return 1/(np.sqrt(2*np.pi*sigma**2)) * np.exp(-0.5*(x-mu)**2/
26
      sigma**2)
```

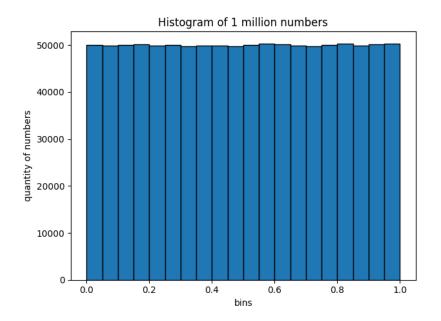


Figure 2: Histogram of 1 million random numbers in 20 bins with a width of 0.05

```
28 n1k_normal = Box_Muller(n1k_uni)
                                                                                                                                                                           normally
                     distributed whose mu = 0; sigma =1
_{29} _{n1k\_normal\_1b} = 2.4 * _{n1k\_normal}
                                                                                                                                                                    # target sigma is
                     2.4
30 \text{ n1k\_normal\_1b} += 3.
                                                                                                                                                                    # target mean is 3
31
32 #plot the normal random number histogram and corresponding Gaussian
                         line
33 fig2 = plt.figure(2)
34 #hist
ax2_1 = fig2.add_subplot(1,1,1)
a_{b} = a_{b} + a_{b
                      (3-2.4*5, 3+2.4*5, 21),
                                                                                                         density='true', label='hist')
37
38 ax2_1.set_xlabel('value of the numbers')
39 ax2_1.set_ylabel('probability')
40 #Gaussian line
_{41} \text{ x\_GPDF} = \text{np.linspace} \left( 3 - 2.4 * 5 \,, \ 3 + 2.4 * 5 \,, 101 \right)
42 y_GPDF = Gaussian_PDF(x_GPDF, 3, 2.4)
43 ax2_1.plot(x_GPDF, y_GPDF, label='Gausian PDF')
44 #indicated lines
45 for i in range (1,6):
                      ax2_1.axvline(x=3+2.4*i, color='k', linestyle='--')
46
                     ax2_1.axvline(x=3-2.4*i, color='k', linestyle='--')
47
48 #show the figure
49 ax2_1.legend(loc='best')
50 fig2.suptitle('normally distributed random number test')
51 #fig2.savefig('1b.png')
52 fig 2. show()
```

## 1.3 1.c KS-test

### normally distributed random number test

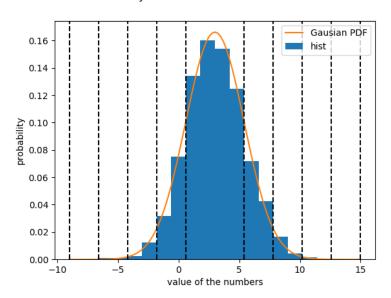


Figure 3: Histogram of 1000 normally distributed random number compared with Guassian distribution

```
, ,,
2 1(c) KS-test
3 KS-test is based on Cumulative distribution function. In our case,
        we use Gaussian CDF.
 4 Gaussian CDF can be given by a error function 'erf(z)' which is a
        special funtion.
_{5}\ m= 0; sigma = 1 gives a standard normal distribution , where
6 Gaussian_CDF(x) = 0.5*(1+erf(x/sqrt(2)), erf(z) = 2/sqrt(pi) *
        integral_0^z (e^-t2^dt)
9 print('\n 1(c) KS-test')
10 # Now do the integral, using trapeziodal rules
_{11} def integrator(function, lower ,upper, n_intervals):
12
13
        trapeziodal rule
        function =
14
15
        lower = lower limit of the integral
        upper = upper limit
16
        interbvals \, = \, n \ of \ n\_intervals
17
        \begin{array}{ll} h = (upper - lower) \ / \ n\_intervals \\ S = 0.5*(function(lower) + function(upper)) \end{array}
19
20
        for i in range(1, n_intervals):
21
        S += function(lower + i*h)
integral = h * S
22
23
        return integral
24
_{25} #define the error function in order to calculate Gaussian CDF
26 def erf(z):
        erf_{integral} = lambda t: np.exp(-t**2)
27
        \mathtt{erf} \, = \, 2. \, / \, \mathtt{np.\,sqrt} \, \big( \mathtt{np.\,pi} \, \big) \, \, * \, \, \mathtt{integrator} \, \big( \, \mathtt{erf\_integral} \, \, , 0 \, , \mathtt{z} \, , \! 10 \! * \! * \! 3 \big)
28
        return erf
```

```
30 # Gaussian cumulative distribution funciton
def Gaussian\_CDF(x):
       CDF = 0.5*(1 + erf(x/np.sqrt(2.)))
32
33
       return CDF
34 # KS-test needs to sort the array
  def quick_sort(array,i,j):
35
       i and j are the two elements we want to start with
37
38
39
           pivot = quick\_sort\_process(array, i, j)
40
41
            quick_sort (array, i, pivot)
42
            quick\_sort(array, pivot+1, j) # do several times
       return array
43
  def quick_sort_process(array,i,j):
       pivot = array[i]
45
       while i < j:
46
            while i < j and array[j] >= pivot:
47
               j -= 1
48
            while i < j and array[j] < pivot:
49
                array\,[\,i\,]\,=\,array\,[\,j\,]
50
51
                i += 1
52
                array[j] = array[i]
           array[i]=pivot
53
54
       return i
55 #KS-test function
56
57
  def KS_test(array,CDF):
58
59
       array is the data that we want to test.
60
       index is the index of this array where we want to get the
61
       percentage.
       WARN: index should already be integers
62
63
       #calculate CDF for data
64
       array = sorted (array)
65
       N = len(array)
66
67
       pECDF = np.arange(0,1,1/N) + 1/N
       pCDF = np.zeros(N)
68
       for i in range (N):
69
           pCDF[i] = CDF(array[i])
70
       D = [abs(x) for x in (pECDF-pCDF)]
71
72
       return D
73
74
75 # generate 100,000 normallu distributed numbers
n100k\_normal = Box\_Muller(n1m\_uni[:10**5])
77 #calculate D for every single point, 10**5 in total
78 #D100k = KS_test (n100k_normal, Gaussian_CDF)
                                                           # this step is a
bit slow and cause my laptop heating......

79 #np.save('D100k', D100k)
so D100k = np.load('D100k.npy')
\texttt{s1 index\_ks} = [\,\texttt{int}(\texttt{x})\ \texttt{for x in np.logspace}\,(\texttt{1,5,num} = \texttt{41,base} = \texttt{10})\,]
        to 10**5
82 Dmax = np.zeros(len(index_ks))
ss for i in range(len(index_ks)):
       Dmax[i] = max(D100k[:index_ks[i]])
85 # calculate P value
86 def KS\_CDF(D,N):
       use D value to calculate P_ks(z)
```

```
89
90
        z = (np. sqrt(N) + 0.12 + 0.11/np. sqrt(N))*D
91
92
        if z < 1.18:
             \exp = np. \exp(-np. pi**2/(8*z**2))
93
             P_ks = np. sqrt(2*np.pi)/z * (exp + exp**9 + exp**25)
94
95
96
             \exp = \operatorname{np.exp}(-2*z**2)
97
             P_ks = 1-2*(exp-exp**4+exp**9)
98
99
100
        return P_ks
P_ks_z = p.zeros(len(Dmax))
102 for i in range (len(Dmax)):
        P_ks_z[i] = KS_CDF(Dmax[i], index_ks[i])
104 D_sci = np.zeros(len(index_ks))
   P_val_sci = np.zeros(len(index_ks))
105
106 for i in range(len(index_ks)):
        D\_sci\left[\,i\,\right]\,,\ P\_val\_sci\left[\,i\,\right]\,=\,stats\,.\,kstest\,(\,n100k\_normal\,\left[\,:\,index\_ks\,\right[\,i]\,\right]
107
        ]], 'norm')
        #plot consistentcy
108
_{109} fig3 = plt.figure(3)
_{\mbox{\scriptsize 110}} #plot Dmax vs the amount of numbers I used to test
111 ax3_1 = fig3.add_subplot(2,2,1)
ax3_1.scatter(index_ks, Dmax)
113 ax3_1.set_xscale('log')
114 ax3_1.set_ylim(-10**-4,0.004)
_{\mbox{\scriptsize 115}} ax3\mbox{\scriptsize \_1}.set\mbox{\scriptsize \_ylabel}('maxium D')
116 ax3_1.set_title('my Dmax')
ax3_2 = fig3.add_subplot(2,2,2)
118 ax3_2.scatter(index_ks, D_sci)
119 ax3_2.set_xscale('log')
120 ax3_2.set_title('Dmax from scipy')
121 # plot P value
122 \text{ ax} 3.3 = \text{fig} 3. \text{add_subplot} (2, 2, 3)
123 ax3_3.scatter(index_ks, P_ks_z)
124 ax3_3.set_ylabel('P value')
ax3_3.set_title('my P of z')
126 ax3_3.set_xscale('log')
127 \text{ ax} 3.4 = \text{fig} 3. \text{ add_subplot} (2, 2, 4)
{\tt 128}~{\tt ax3\_4.scatter(index\_ks} , {\tt P\_val\_sci)}
129 ax3_4.set_title('P value from scipy')
130 ax3_4.set_xscale('log')
131
132 fig3.tight_layout()
fig 3. suptitle ('KS-test from 10 to 10000 numbers', y = 1)
134 #fig3.savefig('1c.png')
135 fig3.show()
136
137
138 I thought D is the maxium value from all the points that I put into
         the test.
                     I calculated 10<sup>5</sup> distances between ECDF and
139 In this case,
        Gaussion CDF.
140 Then I select the maximum from first 10 distance, 10^1.1 distances
        , 10^1.2 distances.....10^5 distances.
141 My maximum D goes up, besides, values of D also cause a unnormal
   result to P value.
```

### 1.4 1.d Kuiper's test

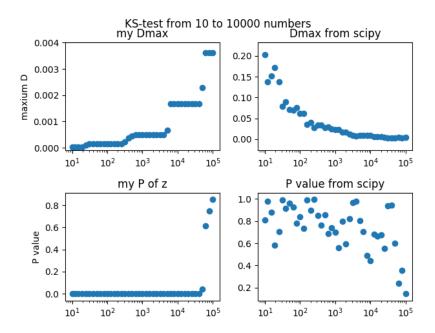


Figure 4: The upper panels show the maxium distance of KS-test for my test and scipy test; the lower panels show the P value

```
1 (d) Kuiper test:
  I wll present the statistic of Kuiper test
  print('\n 1(d) Kuiper test')
  def Kuiper_test(array, CDF):
7
       array : array we want to test
      CDF: targert distribution
10
11
12
13
14
      N = len(array)
      array = sorted(array)
15
      # empirical cdf
16
17
      pECDF = np.arange(0,1,1/N) + 1/N
      pCDF = np.zeros(N)
18
       for i in range(N):
19
           pCDF[i] = CDF(array[i])
20
      # Maximum distance when p_data > p CDF
21
      D_plus = max(pECDF-pCDF)
      # p_data < p CDF
23
      D_minus = max(CDF-ECDF)
24
      V = D_plus + D_minus
                               # Kuiper statistic
25
26
      return V
```

My algorithm takes too long during my testing, I will just show it without resluts.

## 1.5 1.e Compare random numbers with examples

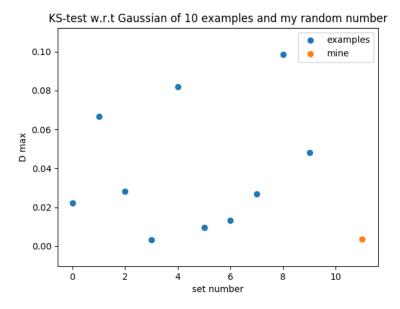


Figure 5: Comparision of my random numbers and examples based on D values of KS-test

According to the figure, my set of random numbers has a lower D value of KS-test compared to the examples. This means it shows more consistency with corresponding Gaussian distribution.

# 2 Question 2: Gaussian random field

I am sorry I don't understand the symmetric thing very well. I asked TA about this but still not very clear. I know that if I take a 2-D matrix of real numbers and do Fourier transform, I will find conjugate symmetry in Fourier space. However, when I did this, I generated  $k^{**}2 = kx^{**}2 + ky^{**}2$  for right half of Fourier space and take  $P(k) = k^{**}$  power as an amplitude. I did this separetly in two (512,512) array. Then I have four (512,512) Gaussian distributed random numbers arrays which were generated and saved in question. Those Gaussian

arrays represent real part and imaginary part for two amplitude arrays. I took conjugate symmetric matrix, with respect to the center, of region 1 and region 2. Finally I combined those 4 arrays to get a (1024,1024) matrix which represented the field in Fourier space.

```
1 #!/usr/bin/env python
2 # coding: utf-8
4 # In[185]:
7 import numpy as np
s import matplotlib.pyplot as plt
9 from scipy import fftpack
10 plt.ioff()
11
12 print ('\n Question 2 : Gaussian random field')
_{\rm 13} def gaussian_random_field(power, size):
14
       power : power of the spectrum
15
       size : image size in per axis
16
17
      N=int(size/2)
18
      #because of conjuate symmetry, I only generate N=\,\mathrm{size}\,/2
19
       Grn = np.load('normal_rn_q2.npy').reshape((4,N,N))
20
21
       Fourier_p1_real = Grn[0]
       Fourier_pl_imag = Grn[1]
22
       Fourier_p2_real = Grn[2]
23
       Fourier_p2_imag = Grn[3]
       Fourier_p1 = Fourier_p1_real + 1j*Fourier_p1_imag
25
       Fourier_p2 = Fourier_p2_real + 1j*Fourier_p2_imag
26
       #because of conjugate symmetry, I only generate N = size/2
27
       matrix for sub-plane 1 and sub-plane 2
28
      #and sub-plane 3 is the conjugate symmetric matrix of 1. 4 is
       of 2.
       kx = np.linspace(np.pi/N, np.pi, N)
29
       ky = kx
30
       k = np.zeros((N,N))
31
32
       for i in range (N):
33
           for j in range(N):
               k[i,j] = (kx[i]**2+ky[j]**2)**0.5
34
35
      Pk = k ** power
                           # I get Pk now.
      # Gaussian random number will be used to scale the Pk
36
       Fourier_p1 *= Pk
37
       Fourier_p2 *= Pk
      #Fourier_p2 = Fourier_p1.conjugate()
39
       Fourier\_p2 = np.\,flip\,(\,Fourier\_p2\,\,,\ axis\,=\,0)
40
       Fourier_p12 = np.concatenate((Fourier_p2, Fourier_p1))
41
      # take conjuagate
42
       Fourier_p3 = Fourier_p1.conjugate()
43
       Fourier_p3 = np.flip(Fourier_p3, axis = 0)
44
       Fourier_p3 = np.flip(Fourier_p3, axis = 1)
45
       Fourier_p4 = Fourier_p2.conjugate()
46
       Fourier_p4 = np.flip(Fourier_p4, axis = 0)
47
48
       Fourier_p4 = np. flip (Fourier_p4, axis = 1)
       Fourier_p34 = np.concatenate((Fourier_p4, Fourier_p3))
49
       Fourier_p = np.concatenate((Fourier_p34, Fourier_p12), axis = 1)
50
      GRF = fftpack.ifft2(Fourier_p)
51
       return GRF, Fourier_p
52
53
54 for n in [-1.0, -2.0, -3.0]:
```

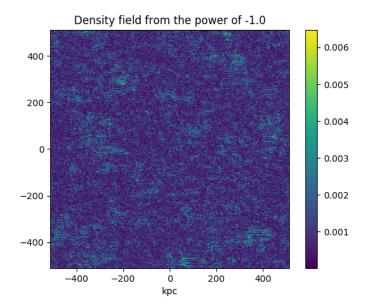


Figure 6: Density field 1

```
field = gaussian_random_field(n, size=1024)[0]
55
56
        plt.imshow(np.absolute(field), extent = [-512,512,-512,512]) plt.title('Density field from the power of \{0\}'.format(n))
57
        plt.xlabel('kpc')
59
        plt.xlabel('kpc')
60
        plt.colorbar()
        plt.savefig(,q_2_{-}\{0\}.png'.format(int(-n)))
62
63
        #plt.show()
65 \#print('\n show a row of Fourier plane = ', gaussian_random_field(n,
         size=1024)[1][0])
```

The mistakes are caused by my misunderstanding on symmetry.

# 3 Question 3: Linear structure growth

Basically, this question is to solve a ODE, we will use 4th order Runge-Kutta method. Since  $\Omega_m = 1$ , the differential equation becomes : D"+4/(3t)\*D' =  $2/(3t^{**}2)$ \*D.

In terms of Runge-Kutta, for every iteration, I use a pair of D and D' as an input to calculate D". Then I will have all information of this present state in order to calculate next state. The pair of D and D' will be updated by R-K method and used in next iteration.

```
1 #!/usr/bin/env python
2 # coding: utf-8
3
4 # In[1]:
5
6
```

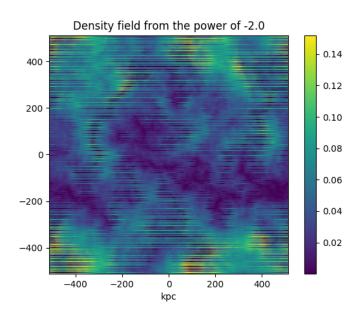


Figure 7: Density field 2

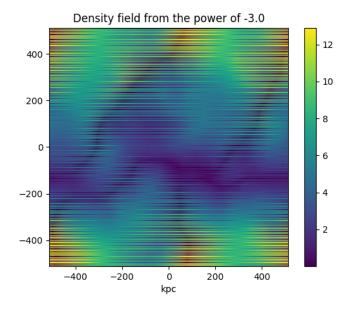


Figure 8: Density field 3

```
7 import numpy as np
8 import matplotlib.pyplot as plt
9 import sys
10 from scipy import stats
11 plt.ioff()
12
13
14 # In[]:
15
17 ,,,
18 Question 3: Linear structure growth
19
20 print ('Question 3: Linear structure growth')
22
23 def R_K_4 (ODE, stepsize, t, D_duo):
       Runge_Kutta methor (4th order)
25
       ODE: the ODE we want to solve
26
       initial : initial state. In our case D(1) \& D'(1)
27
28
29
       h= stepsize
       \# calculate yn+1 - yn = ?
30
       # This will use the old state D_duo which is so-called yn
31
32
       # D_duo cantains D0 and D1, which can be used to calculate D2
       \# D1 and D2 determine the k and growth of D0 and D1 \,
33
34
       k1 = h * ODE(t, D_duo)
                                                       \# k1 = h* f(x,y,y1,
       y2 . . . . . . )
       k2 = h * ODE(t + h*0.5, D_duo + k1*0.5)
35
36
       k3 = h * ODE(t + h*0.5, D_duo + k2*0.5)
       k4 = h * ODE(t + h, D_duo + k3)
37
       growth = k1/6. + k2/3. + k3/3. + k4/6.
38
       \#update the state to yn+1 = D_duo
39
       D_{-}duo += growth
40
       return D_duo
41
42 def ODE(t, D_duo):
43
       as same as the f(x,y) in slides
44
       note that we have second order derivative here.
45
       D\_duo\ will\ include\ both\ D\ and\ D'\,,\ will\ call\ them\ D0\ and\ D1
46
47
       and be propagated togather.
48
       #initial state
49
50
       D0 = D_duo[0]
       D1 = D_duo[1]
51
52
       D2 = 2./(3.*t**2)*D0 - 4./(3.*t)*D1
53
       return np.array((D1,D2))
54
55 def D_analytical(D0,D1,t):
       term1 = 3/5 * (D0+D1)

term2 = D0 - term1
56
57
       return term1*t**(2/3) + term<math>2*t**(-1)
58
59
60
61 # In [28]:
62
63
64 # solve case 1
65 D_duo_case1 = np.array((3.,2.))
66 \text{ stepsize} = 0.1
t = np.arange(1,1000,stepsize)
```

```
0.68 D_{case1_rk} = np.zeros(len(t))
 D_{case1_a} = np.zeros(len(t))
 70 for i in range(len(t)):
           \texttt{D\_case1\_rk}\left[\,i\,\right] \;=\; \texttt{R\_K\_4}\left(\texttt{ODE}, \;\; 0.1\,, t\left[\,i\,\right]\,, \texttt{D\_duo\_case1}\right)\left[\,0\,\right]
71
           D_{case1_a[i]} = D_{analytical(3,2,t[i])}
73 \text{ fig1} = \text{plt.figure}(1)
 74 \text{ ax1}_{-1} = \text{fig1.add}_{-\text{subplot}}(111)
75 ax1.1.loglog(t, D_case1_rk, label = 'Runge-Kutta')
76 ax1.1.loglog(t, D_case1_a, label = 'Analytical')
 ax1_1.legend(loc='best')
78 ax1_1.set_xlabel('t')
 79 ax1_1.set_ylabel(',D(t)')
 so fig1.suptitle('case1')
 si fig1.savefig('q3_case1'.png')
 82 #fig1.show()
83
 84
 85 # In [32]:
 86
 87
88 \# case 2
 89 D_duo_case2 = np.array((10., -10.))
 90 \text{ stepsize} = 0.1
91 t = np.arange(1,1000,stepsize)
 p_2 D_case2_rk = np.zeros(len(t))
 D_{case2_a} = np.zeros(len(t))
94 for i in range(len(t)):
           \texttt{D\_case2\_rk}\left[\,i\,\right] \;=\; \texttt{R\_K\_4}\left(\texttt{ODE}, \;\; 0.1\,, t\left[\,i\,\right]\,, \texttt{D\_duo\_case2}\,\right)\left[\,0\,\right]
95
           D_{case2_a[i]} = D_{analytical}(10., -10., t[i])
97 \operatorname{fig} 2 = \operatorname{plt} . \operatorname{figure} (2)
 98 ax2_1 = fig2.add_subplot(111)
99 ax2_1.loglog(t,D_case2_rk, label = 'Runge-Kutta')
100 ax2_1.loglog(t,D_case2_a, label = 'Analytical')
101 ax2_1.legend(loc='best')
102 ax2-1.set_xlabel('t')
103 ax2-1.set_ylabel('D(t)')
104 fig2.suptitle('case2')
105 \operatorname{fig} 2. \operatorname{savefig} ('q3 \operatorname{case} 2.\operatorname{png}')
106 #fig2.show()
107
108
109 # In [34]:
110
111
_{112} #case _{3}
113
D_{duo\_case3} = np.array((5.,0.))
115 \text{ stepsize} = 0.1
_{116}\ t = np.arange(1,1000,stepsize)
D_{case3_rk} = np.zeros(len(t))
D_{case3_a} = np.zeros(len(t))
119 for i in range(len(t)):
           D_{case3_rk[i]} = R_K_4(ODE, 0.1, t[i], D_duo_case3)[0]
120
           D_{case3_a[i]} = D_{analytical(5.,0.,t[i])}
121
122 \text{ fig3} = \text{plt.figure}(3)
123 \text{ ax} 3.1 = \text{fig} 3. \text{add\_subplot} (111)
\begin{array}{lll} _{124} & ax3\_1.\log\log\left(t\,,D\_case3\_rk\,,\ label = \ 'Runge-Kutta\,'\right) \\ _{125} & ax3\_1.\log\log\left(t\,,D\_case3\_a\,,\ label = \ 'Analytical\,'\,) \end{array}
126 ax3_1.legend(loc='best')
127 ax3_1.set_xlabel('t')
128 ax3_1.set_ylabel('D(t)')
129 fig3.suptitle('case3')
```

### case1

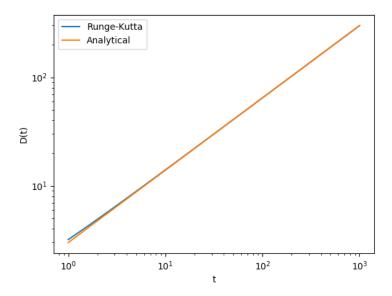


Figure 9: Result provied by Runge-Kutta method and analytical solution of case 1.

```
130 fig3.savefig('q3_case3.png')
131 #fig3.show()
132
133
134 # In[]:
```

The fugures show that my results fit with analytical solution very well.

# 4 Question 4 : Zeldovich approximation

```
_{1} \#!/usr/bin/env python _{2} \# coding: utf-8
4 # In[15]:
7 import numpy as np
s import matplotlib.pyplot as plt
9 import sys
10
11
_{12} # In [39]:
13
14
15
16 Question 4 : Zeldovich approximation
17 4(a) : calculte the growth factor
18
19 print ('\n Question 4 : Zeldovich approximation')
20 print ('\n 4(a) calculte the growth factor')
```

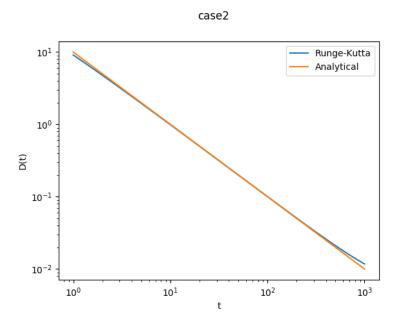


Figure 10: Result provied by Runge-Kutta method and analytical solution of case  $\!2$ .

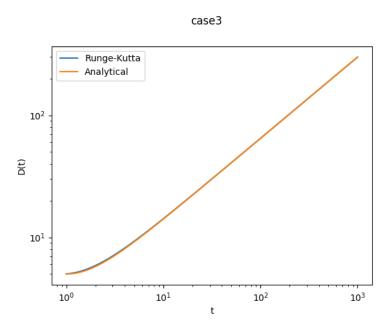


Figure 11: Result provied by Runge-Kutta method and analytical solution of case 3.

```
21 def integrator (function , lower , upper , n_i intervals):
22
      trapeziodal rule
23
24
      function =
      lower = lower limit of the integral
25
      upper = upper limit
26
27
      interbvals = n of n_intervals
28
      h = (upper - lower) / n_intervals
29
      S = 0.5*(function(lower) + function(upper))
      for i in range(1, n_intervals):
31
32
          S += function(lower + i*h)
33
      integral = h * S
      return integral
34
35 def D_of_a_int(a_prime):
36
      use a, because 1/1+z will become a, which is more simpler.
37
      hereby I define the integral part of D(a)
39
40
      Omega\_m\!=\!0.3
41
      Omega\_lambda = 0.7
42
43
      return 1/(Omega\_m/a\_prime + Omega\_lambda*a\_prime**2)**(3/2)
44 z = 50
a = 1/(1+z)
_{46} Omega_m=0.3
_{47} Omega_lambda = 0.7
48 \#calculte the coefficient part
49 coeffi = 5*Omega_m/2. *np.sqrt(Omega_m*(1+z)**3 + Omega_lambda) #
      for accuracy, use z to calculte.
50 integral = integrator(D_of_a_int, 10**-8, a, 10**6)
51 growth = coeffi * integral
52 print ('growth factor =
54
55 # In[44]:
57
58 ,,,
59 4 (b)
60 dD/dt = 5*Omega_m*H0**2/(2*a**3*H(a)) * (-3*Omega_m*H(a)*integral
      /2*H0 + 1)
62 print('\n 4(b)')
63 z = 50
a = 1/(1+z)
65 \text{ H0} = 70
_{66} Ha = 70*(Omega_m*(1+z)**3+Omega_lambda)**0.5
integral/(2*H0) + 1
68 print('analytical derivative at (z=50) = ', analytical_d)
69
71 # In[ ]:
        Result
```

```
(a) calculte the growth factor growth factor = 0.01960778042827206 . (b) analytical derivative at (z=50) = 34499.31702139651.
```

## 5 Question 5 : Mass assignment schemes

### 5.1 a. Nearest Grid Point

I loop over the particles and assign their densties to the cells and the fraction of particle's mass assigned to a cell 'ijk' is the S(x) averaged over this cell. For 3 dimensions, W(ijk) = W(X)\*W(Y)\*W(Z). NGP is the simplest PM algorithm that assume particles are point-like and all of particles's mass is assigned to the single grid cell that contains it.

```
1 print ('Question 5 : Mass assignment schemes')
2 print ('\n 5(a)')
3 def NGP(cell_size, positions):
       cell\_size = the size of the cells ; (n,n,n)
       position: positions of the particles
6
       grid = np.zeros(cell_size)
       #assign the particles to the cell, indices is the cell that the
9
        particle belongs to, is int()
       indices = positions.astype(np.int)
                                                    #for every particles
       for i in range (indices.shape [1]):
11
            grid[indices[:,i][0],indices[:,i][1],indices[:,i][2]] += 1
12
          # located in a grid and count it
       return grid
13
14
15 # Particles ' positions
16 np.random.seed (121)
positions = np.random.uniform(low=0, high=16, size=(3,1024))
18 grid = NGP((16, 16, 16), positions)
20 # Plot x-y slices of the grid
z = [4,9,11,14]
22 for i in range (4):
       plt.title('z={0} layer'.format(z[i]))
23
       {\tt plt.imshow} \, (\, {\tt ngp} \, [\, : \, , : \, , z \, [\, i \, ]\, ] \,\, , \, {\tt extent} = [\, 0 \, , 16 \, , 16 \, , 0\, ]\, )
24
       plt.colorbar()
       #plt.savefig('q5_a{}.png'.format(i))
26
       plt.show()
```

### 5.2 b. Test NGP

I moved one particle along x axis and test the number in cell0 and cell4. When x is from 0 to 1, particle number in cell0 should be equal to 1. When x is from 4 to 5, particle number in cell4 should be equal to 1.

```
1 '''
2 (b) test the robustness
3 '''
4 print('\n 5(b) test the robustness')
5 x_test = np.linspace(0.1,16,30)  # move the particle along x-axis
6 cell4 = np.zeros(30)
7 cell0 = np.zeros(30)
8 for i in range (16):
9     position_test = np.array(([x_test[i]],[0],[0]))
10     grid_test = NGP((16,16,16), position_test)
11     cell4[i] = grid_test[4,0,0]
12     cell0[i] = grid_test[0,0,0]
13
14 #plot numbers in cell 4 first, when x = 4 to 5 , cell 4 should be 1
```

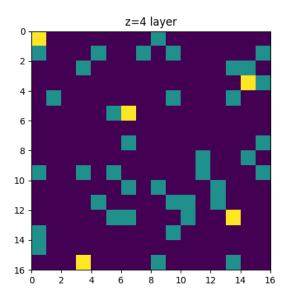


Figure 12: Slice of NGP at z=4.

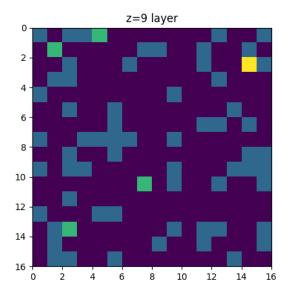


Figure 13: Slice of NGP at z = 9.

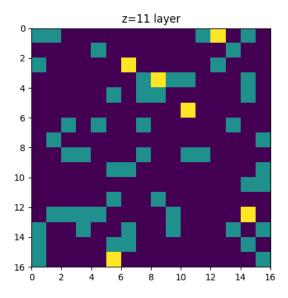


Figure 14: Slice of NGP at z=11.

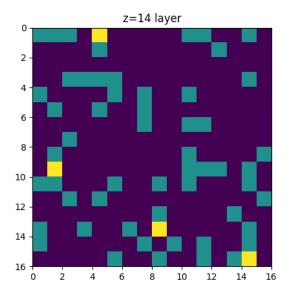


Figure 15: Slice of NGP at z = 14.

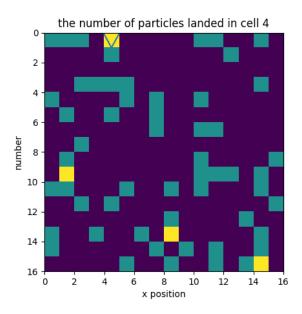


Figure 16: Robustness test on cell0

```
15 plt.plot(x_test,cell4)
16 plt.title('the number of particles landed in cell 4')
17 plt.ylabel('number')
18 plt.xlabel('x position')
19 #plt.savefig('q5_b1.png')
20 plt.show()
21
22
23 # repeat for cell 0, when x = 0 to 1, cell 0 should be equal to 1
24 plt.plot(x_test,cell0)
25 plt.title('the number of particles landed in cell 0')
26 plt.ylabel('number')
27 plt.xlabel('x position')
28 #plt.savefig('q5_b2.png')
29 plt.show()
```

### 5.3 d. Fast Fourier transform

```
1 , , ,
2 5(d) FFT
3 ,,,
4 def DFT_slow(x):
      #1-D discrete Fourier Transform
      x = np.array(x, dtype=float)
      N = len(x)
      n = np.arange(N)
      k = n.reshape((N, 1))
9
      M = np.\exp(-2\,j \ * np.pi \ * k \ * n \ / \ N)
                                # use vector multiplication
      return np.dot(M, x)
11
_{12} def FFT_1D(x):
      \# 1–D Fast FT
13
      x = np.array(x, dtype=float)
14
```

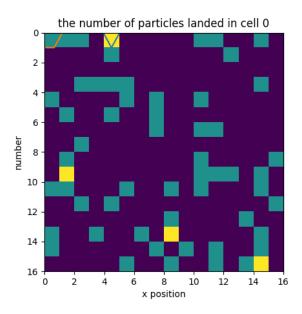


Figure 17: Robustness test on cell4

```
N = len(x)
15
16
       if N \% 2 != 0:
17
           raise ValueError("size of x must be a power of 2")
18
                       # end of the recurse
       elif N \le 4:
19
           return DFT_slow(x)
20
21
           X_{even} = FFT_1D(x[::2])
22
           X_{\text{odd}} = FFT_{\text{1D}}(x[1::2])
           24
25
                                    X_{even} + factor[int(N*0.5):] * X_{odd}
26
      ])
27 #test
28 #use sin(t)
29 \operatorname{fun1} = \operatorname{lambda} t : \operatorname{np.sin}(t)
30 N = 64
t_test = np.linspace(0,4*np.pi,N)
32 \text{ fun1t} = \text{fun1}(t_{\text{-}}\text{test})
33 Xk = FFT_1D(fun1t)
_{34} fft_np = np.fft.fft(fun1t)
35
з6 #plot
37 fs = 64/(4*np.pi)
                         # sampling frequency
                                     # fs/N = interval
38 fk = fs/N*np.arange(0,N*0.5,1)
39 plt.figure()
40 plt.plot(fk,np.abs(Xk)[:int(N*0.5)],label='My FFT')
plt.axvline(x=1/(2*np.pi), color='k', linestyle='--', label = '
      analytical')
42 plt.plot(fk,np.abs(fft_np[:int(N*0.5)]), linestyle='--', label = '
      FFT by numpy')
43 plt.xlim(0,1)
```

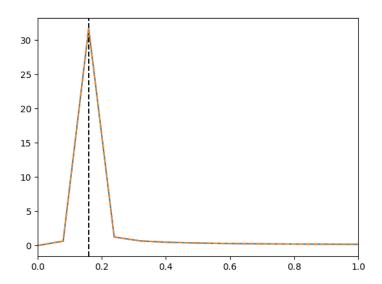


Figure 18: Robustness test on cell0

```
#plt.savefig('q5_d.png')
plt.legend(loc='best')
```

I used  $y = \sin(x)$  to do the test, whose analytical frequency should be 1/2pi. My result's peak shows great consistency with both np.fft and analytical Fourier transform.

## 5.4 e. FTT in 2-D and 3-D

```
<sup>2</sup> 5(e): 2&3-D FFT
  def FFT_2D(x):
5
       2-D FFT. Becuase FFT_2D = FFT(FFT(x), y)
       F\_xy \; = \; np.\,array\,(\,np.\,zeros\,(\,x\,.\,shape\,)\,)
       # 1-D Fourier transform through the rows
       for i in range(len(x)):
10
           F_xy[i,:] = FFT_1D(x[i,:])
11
       # 1-D Fourier transform through the columns
12
       for j in range(len(x[0])):
13
           F_xy[:,j] = FFT_1D(F_xy[:,j])
       return F_xy
15
16
  def FFT_3D(x):
18
       3-D FFT. Becuase FFT_3D = FFT_2D(FFT_2D(FFT_2D(x,y)),(y,z)),(x,
19
       I thought I am wrong here, sorry.
20
21
       F_xyz = np.array(np.zeros(x.shape))
22
       \# 2-D Fourier transform through the x
23
```

```
24
       for i in range (len(x)):
            F_xyz[i,:,:] = FFT_2D(x[i,:,:])
       # 2-D Fourier transform through the y
26
27
       for j in range (len(x[0])):
            F_{xyz}[:, j, :] = FFT_2D(F_{xyz}[:, j, :])
28
       #3-D Fourier transform through the z
29
       for k in range (len(x[1]))
            F_{xyz}[:,:,k] = FFT_{2D}(F_{xyz}[:,:,k])
31
       return F_xyz
32
34
35 # chose function f(x,y) = \sin(x+y)
36 \text{ fun } 2 = \text{lambda } x, y : \text{np.} \sin(x+y)
37 \text{ fun } 2 \text{-xy} = \text{np.zeros} ((64,64))
  x_{test}, y_{test} = t_{test}, t_{test}
39 for i in range (64):
       for j in range (64):
40
            fun2\_xy[i,j] = fun2(x\_test[i],y\_test[j])
42
  F_2D_result = FFT_2D(fun2_xy)
43
44
45 #plot function
46 plt.figure()
47 plt.imshow(fun2_xy)
48 plt.colorbar()
49 plt. title ('Function')
50 #plt.savefig('q5_e1.png')
51 plt.show()
52 #plot fourier space
53 plt.figure()
54 plt.title('Fourier sapce')
55 plt.imshow(F_2D_result)
56 plt.colorbar()
57 #plt.savefig('q5_e2.png')
58 plt.show()
```

I used sin(x+y) as a testing funtion.

# 6 Question 6: Classifying gamma-ray bursts

The work has three part as follow: First, I pre-processing the data including label the data and processing the missing data. Second, I conducting the gradiant ascend algorithm to train the classifer. Conclusion and discussion are followed in the last part.

### 6.1 Data Processing

Data are labelled first according to the parameter 'T90' showing the amount of long(label = 1) object is much more than otherwise. I revise the other features and find that there are many missing values, therefore, handling the missing data is of great importance.

I first plot the histogram to check the feature values.

It's seems that the mass M and metality Z conform to the gauss distribution. Meanwhile, SFR fits the exoponential distribution. Therefore, I choose to fill the non-determinded variables with random number with specific distribution. The data with SSFR and AV are few, so I decided to give up these two features.

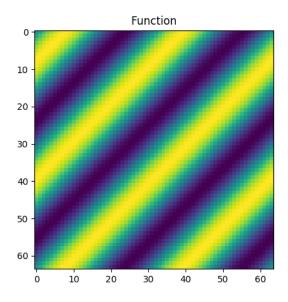


Figure 19: Testing function

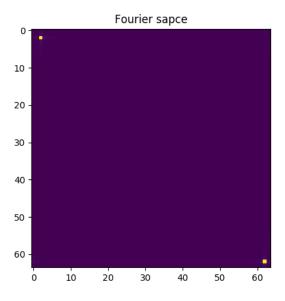


Figure 20: My FFT on testing function

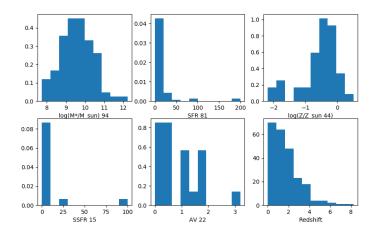


Figure 21: Histograms of the features

### 6.2 Train the classification

Applying the gradiant ascend algorithm, I plot two figures to show my results.

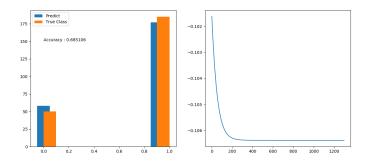


Figure 22: Accuracy and loss

The figures show that the classifer is well trained even if the result is not relative good enough. The histogram displays the similarity between true label and predicted result, however, the accuracy is merely 70 percent, which could be better.

### 6.3 Discussion

I think this lower accuracy is due to my handling of missing values. Because of wrong filled value, the intrinsic feature of some objects changed which lead to a bad result. I can't totally blame it on the bad algorithm in which play a role. Based on this, I can not discuss which features play the key role in classifying the class.

#!/usr/bin/env python

```
2 # coding: utf-8
 6 # In[1]:
9 import numpy as np
10 import pandas as pd
11 import matplotlib.pyplot as plt
12 from scipy.optimize import curve_fit
13 plt.ioff()
14
15
16 # ### Read the data file and label the data
17
18 # In [2]:
20
<sub>21</sub> ,,,
22 Question 6 : Classifying gamma-ray bursts
23
24 print ('Question 6 : Classifying gamma-ray bursts')
25 #data = pd.read_csv('GRBs.txt', sep='\s+')
26 data = np.genfromtxt('GRBs.txt', usecols=(2,3,4,5,6,7,8))
27 data = pd.DataFrame(data, columns=['Redshift', 'T90', 'log(M*/M_sun
        )', 'SFR',
                                                 '\log(Z/Z_sun)', 'SSFR', 'AV'])
28
_{29} data ['label'] = None
30 #Assign the label based on T90 parameter
data['label'][data['T90'] < 10] = 0
data['label'][data['T90'] >= 10] = 1
data['label'] = data['label']. convert_objects(convert_numeric=True)
34 print ('Data Shape:', data.shape)
35
_{37} # ### Histogram the features.
_{38}\ \#\ I first plot the histogram to check the feature values.
39
40 # In [3]:
41
42
_{43} # Check the missing data
44 index_M = data['log(M*/M_sun)'] != -1
45 index_SFR = data['SFR'] != -1
46 index_Z = data['log(Z/Z_sun)'] != -1
47 index_SSFR = data['SSFR'] != -1
48 index_AV = data['AV'] != -1
49
fig = plt.figure(figsize = (10,6))
51
ax1 = fig.add_subplot(2,3,1)
53 M = ax1.hist(data['log(M*/M_sun)'][index_M], density=True)
54 ax1.set_xlabel('log(M*/M_sun) %s '%len(data['log(M*/M_sun)'][
        index_M]))
_{56}~ax2=fig.add\_subplot(2,3,2) _{57}~SFR=ax2.hist(data['SFR'][index\_SFR],~density=True)
58 ax2.set_xlabel('SFR %s' %len(data['SFR'][index_SFR]))
ax3 = fig.add\_subplot(2,3,3)
61 Z = ax3. hist (data ['log(Z/Z_sun)'][index_Z], density=True)
```

```
_62 ax3.set_xlabel('log(Z/Z_sun %s)' %len(data['log(Z/Z_sun)'][index_Z
        ]))
63
ax4 = fig.add_subplot(2,3,4)
65 SSFR = ax4.hist(data['SSFR'][index_SSFR], density=True)
66 ax4.set_xlabel('SSFR'%s' %len(data['SSFR'][index_SSFR]))
67
ax5 = fig.add_subplot(2,3,5)
69 AV = ax5. hist (data['AV'][index_AV], density=True)
70 ax5. set_xlabel('AV %s' %len(data['AV'][index_AV]))
71
72 ax6 = fig.add\_subplot(2,3,6)
73 ax6.hist(data['Redshift'])
74 ax6. set_xlabel('Redshift')
75 fig.savefig('q6_1.png')
76
77
78 # It's seems that the mass M and metality Z conform to the gauss
        distribution. Meanwhile, SFR fits the exoponential distribution
        . Therefore, I choose to fill the non-determinded variables
        with random number with specific distribution.
_{79}\ \# The data with SSFR and AV are few, so I decided to give up these
        two features.
80
81 # In [4]:
83
84 #Processing the missing data
\begin{array}{lll} ss & miu\_M = np.mean(data['log(M*/M\_sun)'][index\_M]) \\ sc & sigma\_M = np.std(data['log(M*/M\_sun)'][index\_M]) \end{array}
87 lambda_SFR = np.mean(data['SFR'][index_SFR])
88 miu_Z = np.mean(data['log(Z/Z_sun)'][index_Z])
89 sigma_Z = np.std(data['log(Z/Z_sun)'][index_Z])
91 index = data ['\log (M*/M_sun)'] == -1
92 index_len = len(index)
93 data['log(M*/M_sun)'][index] = np.random.normal(miu_M, sigma_M,
        index_len)
95 index = data['\log(Z/Z_sun)'] == -1
96 \text{ index\_len} = \text{len(index)}
   data['log(Z/Z_sun)'][index] = np.random.normal(miu_Z, sigma_Z, size
        =index_len)
98
99 index = data['SFR'] == -1
index_len = len(index)
101 data['SFR'][index] = np.random.exponential(lambda_SFR, index_len)
102
103
_{104} # ## Part 2, Train the classification applying the gradiant ascend
        algorithm
   \# I plot two figures to show my results.
105
106
107 # In [5]:
108
109
110 def sigmoid(z):
        return 1 / (1 + np.exp(-z))
111
112
113 # def binary_crossentropy(y_true, y_predict):
          m = y_true.shape[0]
114 #
115 #
          return -1/m * ()
```

```
116
117 def load_data(data):
       cols = ['Redshift', 'log(M*/M_sun)', 'SFR', 'log(Z/Z_sun)']
118
       data_Input = pd.DataFrame(data, columns=cols)
119
       data_Input = np.array(data_Input)
120
       data_Label = data['label']
121
122
       data_Label = np.array(data_Label)
       data_Input = np.insert(data_Input, 0, 1, axis=1)
123
       return data_Input, data_Label
124
125
126 def grad_ascent(data_Input, data_Label, alpha, epochs, loss =False)
       data_Mat = np.mat(data_Input)
127
       label_Mat = np.mat(data_Label).transpose()
128
129
       m, n = np.shape(data\_Mat)
       weights = np.random.normal(0.5,0.2,(n,1))
130
       Loss = []
131
       for i in range (epochs):
132
           h = sigmoid(data_Mat * weights)
133
            weights = weights + alpha * data_Mat.transpose() * (h -
134
       label_Mat) / m
            if loss == True:
135
                if i \% 2 == 0:
136
                    Loss.append(-np.sum((np.array(label_Mat) - np.array
137
       (h))**2) / (2*m))
138
       return weights, Loss
139
140
141
142 # In [6]:
144
_{145} data_in , data_lab = load_data(data)
_{146} \text{ epoch} = 1300
147 W, loss = grad_ascent(data_in, data_lab, alpha=0.001, epochs=epoch,
        loss=True)
148
_{149}\;W=\;np\,.\,array\,(\!W\!)
z = np.dot(data_in, W)
z = z.astype(float)
_{152}\ Prediction\ =\ sigmoid\,(\,z\,)
153 Prediction = np.array(Prediction, dtype=int)
154
155 fig = plt.figure(figsize = (14,6))
ax = fig.add\_subplot(121)
ax. hist (Prediction, label='Predict', align='left')
158 ax.hist(data['label'], label='True Class')#, align='right')
ax.legend(loc = 2)
160 \ acc = np.sum(data['label'] = Prediction.flatten()) \ / \ data.shape[0]
161 ax.text(0, 150, 'Accuracy': %f' %acc)
162
ax2 = fig.add\_subplot(122)
ax2.plot(np.arange(0,epoch,2),loss)
_{165} fig.savefig('q6_2.png')
```