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## **On the modulation of kinetic energy transfer by internal gravity waves**

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# On the modulation of kinetic energy transfer by internal gravity waves

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## Key Points:

- Twin submesoscale-permitting simulations of the North Atlantic Ocean were run with and without tidal forcing.
  - The tidally-forced run shows enhanced internal gravity wave (IGW) signal during summer while the non-tidally forced run during winter.
  - Tidally-forced IGWs (i.e. internal tides) enhance forward cascade of kinetic energy particularly during summertime.

21 **Abstract**

22 Understanding how kinetic energy (KE) is exchanged across scales and eventually dissipated  
 23 remains a key question in physical oceanography. Recent theoretical works suggests that  
 24 both quasi-balanced submesoscale motions and internal gravity waves (IGWs) could play  
 25 a role in fluxing KE towards dissipation. How these classes of motions actually provide a  
 26 route to dissipation in the ocean is still debated. This study investigates the impact of IGWs  
 27 generated by tidal motions on cross-scale KE exchanges at mid-latitude. Our analysis is  
 28 based on the output of two realistic submesoscale permitting ocean model simulations of the  
 29 North Atlantic Ocean, run respectively with and without tidal forcing. These twin experi-  
 30 ments permit investigation of how tidally-forced IGWs modify the KE variance, cross-scale  
 31 exchanges, and associated seasonality. Our results show that, in the presence of externally-  
 32 forced IGWs, KE transfer towards dissipative scales is enhanced in summertime both at the  
 33 surface and in the ocean interior.

34 **Plain Language Summary**

35 Energetic oceanic currents on the scales of tens of kilometers that emerge as a re-  
 36 sult of the chaotic nature of the ocean, known as (sub)mesoscale flows, have been of great  
 37 scientific interest to the oceanographic community. These currents are associated with  
 38 the time scales of roughly a day, which overlap with the astronomical tidal frequency of  
 39 the ocean. Due to their similar time scales, it has been argued that the (sub)mesoscale  
 40 flows and tides would interact with each other. Here, by running a twin numerical sim-  
 41 ulation of the North Atlantic Ocean, one with tides and the other without, we examine  
 42 the physical interaction between the two and show that the tides stimulate a loss of en-  
 43 ergy from the (sub)mesoscale flows particularly during the summer.

44 **1 Introduction**

45 The ocean's kinetic energy is mostly concentrated in motions close to geostrophic balance,  
 46 with frequencies lower than the Coriolis frequency ( $f$ ) and spatial scales larger than the first  
 47 Rossby deformation radius ( $R_d$ ; Vallis, 2017). These balanced motions (BMs) are largely  
 48 energized through baroclinic instability which extracts energy from large scale stratification.  
 49 Balanced motions include large-scale motions ( $> 300$  km), mesoscale motions (50–300 km)  
 50 and submesoscale balanced motions ( $< 50$  km) (McWilliams, 2016). Balanced motions are  
 51 characterized by an inverse cascade of energy (Scott & Wang, 2005; Scott & Arbic, 2007;  
 52 Eden, 2007; Aluie et al., 2017), so they do not provide a route to dissipation by themselves.  
 53 Therefore, energy has to be transferred from balanced motions to high-frequency unbalanced

54 motion for dissipation to occur. To equilibrate the well-known inverse cascade of energy, the  
55 ocean requires ageostrophic processes to extract energy from balanced motions. Documented  
56 mechanisms that might trigger a forward transfer of energy from balanced motions down  
57 to dissipate scale include for instance (but are not limited to) (i) bottom boundary-layer  
58 turbulence (Wunsch & Ferrari, 2004; Sen et al., 2008; Arbic et al., 2009), (ii) generation  
59 of lee waves by mesoscale eddies interacting with topography (Nikurashin & Ferrari, 2010;  
60 Nikurashin et al., 2013; Trossman et al., 2013, 2016), (iii) generation of internal waves by  
61 upper-ocean frontal instabilities (Danioux et al., 2012; Shakespeare & Taylor, 2014) and  
62 (iv) direct cascade of energy by energetic submesoscale motions (Capet et al., 2008; Ferrari  
63 & Wunsch, 2009; Molemaker et al., 2010; McWilliams, 2016). Overall, these studies have  
64 shown that, at fine-scale, there are two classes of motions that can provide efficient transfer  
65 of energy to dissipative scales; submesoscale motions and internal gravity waves.

66 Submesoscale motions are flow structures in the form of density fronts, filaments and to-  
67 graphic wakes at the surface and throughout the interior at scale smaller than  $O(100\text{ km})$   
68 (McWilliams, 2016). Internal gravity waves (IGWs) are a particular class of fast propagating  
69 unbalanced motions with frequencies equal to or higher than the Coriolis frequency  $f$  and  
70 spatial scale ranging from  $O(10\text{ m})$  to  $O(100\text{ km})$ . IGWs include wind-induced near-inertial  
71 waves with frequency near the Coriolis frequency and internal tides generated by large scale  
72 barotropic tidal flow over topographic features. Near-inertial waves are usually stronger in  
73 winter than in summer because they are driven by surface winds (D'Asaro, 1985) while the  
74 signature of internal tides (namely IGWs at tidal frequencies) in ocean fields is typically  
75 stronger in summer time (Gerkema et al., 2004). Recent works have highlighted that these  
76 two classes of motions—submesoscale motions and IGWs—are out of phase seasonally (Rocha  
77 et al., 2016). Submesoscale motions tend to be stronger than IGWs in wintertime. The  
78 emergence of submesoscales is due to winter favored mechanisms such as mixed layer in-  
79 stability, wind-induced frontal instability among other processes (Qiu et al., 2014; Sasaki  
80 et al., 2014; Callies, Flierl, et al., 2015; Brannigan et al., 2015; McWilliams, 2016; Uchida  
81 et al., 2019; Khatri et al., 2021). On the other hand, the kinetic energy associated with  
82 IGWs shows stronger amplitude in summertime. This is due to the intensification of vertical  
83 normal modes and shallow mixed layer (Callies, Ferrari, et al., 2015; Rocha et al., 2016).

84 A forward cascade is seen in spectral kinetic energy fluxes computed from gridded  
85 satellite altimeter products (Scott & Wang, 2005) at scales smaller than  $O(100\text{ km})$ . Later  
86 work (Arbic et al., 2013) argued that this forward cascade could be physically based, and  
87 related to energy transfer between mesoscale eddies and IGWs, but could also be due to  
88 the spatial and temporal smoothing inherent in the construction of present-day gridded  
89 altimeter products. In this study, we partly focus on the energy transfers involving IGWs.

Indeed, due to high-frequency winds and tidal motions, energy extraction via IGWs seems to be a highly probable mechanism of kinetic energy sinks for balanced motions (J. Thomas & Daniel, 2021). From a theoretical standpoint, there are so far essentially two well known and documented energy transfer mechanisms from balanced motions to IGWs away from topography, namely the *stimulated generation* of a forward cascade of kinetic energy by near-inertial waves from balanced flows (Gertz & Straub, 2009; Rocha et al., 2018; Barkan et al., 2021) and the *spontaneous generation* of near-inertial waves from balanced flows (Nagai et al., 2015; Shakespeare & Hogg, 2017). In stimulated generation, near-inertial waves are first introduced by external forcing (e.g. wind) at the inertial frequency and then grow by extracting energy from the balanced flow (Gertz & Straub, 2009; Barkan et al., 2017; L. N. Thomas, 2017) while spontaneous generation is the emission of waves by unbalanced, large Rossby number flow at density fronts without external forcing. These waves then radiate vertically downwards into the interior and amplify by extracting energy from deep balanced flow (Shakespeare & Hogg, 2017). Spontaneous generation is localized at sharp submesoscale fronts and is not very efficient at small Rossby numbers (Danioux et al., 2012; Nagai et al., 2015; Shakespeare & Hogg, 2017). On the other hand, stimulated generation is efficient at small Rossby numbers provided that the waves are forced externally.

As listed above, several studies have focused on the impact of wind-generated near-inertial waves on energy dissipation (e.g. Barkan et al., 2017; J. Thomas & Arun, 2020). In contrast, little is known as to the role of internal tides on kinetic energy exchanges (cf. Barkan et al., 2021). We know that internal tides contribute to the building up of internal gravity waves continuum (Garret-Munk spectra) and that their energy eventually contributes to diapycnal mixing in the ocean interior (Arbic et al., 2018). Whether internal tides could play a significant role in the down-scale transfer of mesoscale eddy kinetic energy is yet to be fully explored. In this context, we here focus on investigating the role of IGWs (and particularly internal tides) on cross-scale kinetic energy exchanges in a regime with active submesoscale motions. Our results show that externally forced IGWs enhance KE dissipation in summertime by catalyzing energy transfer from balanced motion to dissipative scale. We do this by using a twin submesoscale resolving numerical simulations (with/without tides) of the North Atlantic Ocean with a horizontal resolution of  $1/60^\circ$ . The only difference between the two simulations is the inclusion of tidal forcing. This permits us to investigate how IGWs affect kinetic energy exchanges in the presence of active submesoscales motions.

This paper is organized as follows: in the next section, we describe the numerical simulations. Section 3 presents the seasonality of balanced motions and IGWs characterized by the kinetic energy frequency-wavenumber spectral density. The contribution of balanced

126 motions and IGWs to the kinetic energy transfer is presented in section 4. We discuss the  
 127 impacts of this observed seasonality on the kinetic energy spectral flux in section 5, and we  
 128 identify two different mechanisms that support a direct cascade of energy in a dynamical  
 129 regime with/without tidal motions.

## 130 2 North Atlantic Ocean Simulation

131 In this study, we use numerical outputs from a NEMO-based submesoscale eddy-permitting  
 132 simulations of the North Atlantic with a horizontal resolution of  $1/60^\circ$  (eNATL60; Brodeau  
 133 et al., 2020). eNATL60 is a spatially-extended version of NATL60 (Ajayi et al., 2020, 2021).  
 134 The simulation spans the North Atlantic ocean from about  $6^\circ\text{N}$  up to the polar circle. In  
 135 particular, this simulation has a horizontal grid spacing ranging from 1.6 km at  $6^\circ\text{N}$  to 0.9 km  
 136 at  $65^\circ\text{N}$ . The model has 300 vertical levels with a resolution of 1 m at the top-most layers to  
 137 better resolve a realistic surface boundary layer. In practice, the model effective resolution is  
 138 about 10–15 km in wavelength, the same as the resolution of the anticipated Surface Water  
 139 and Ocean Topography (SWOT) altimetry mission (Fu & Ubelmann, 2014). The initial  
 140 and open boundary conditions are based on GLORYS2v3 ocean reanalysis with a relaxation  
 141 zone at the northern boundary for sea-ice concentration and thickness. The atmospheric  
 142 forcing is based on ERA-Interim (ECMWF; Dee et al., 2011), the grid and bathymetry follow  
 143 (Ducousoo et al., 2017). A third-order upwind advection scheme is used for both momentum  
 144 and tracers in the model simulation to implicitly adapt lateral viscosity and diffusivity to  
 145 flow properties. The model is spin-up for a period of 18 months, and a one-year simulation  
 146 output from July 2009 to June 2010 is used in this study. eNATL60 has two identical runs (i)  
 147 eNATL60 with tidal forcing herein referred to as eN60-WT and (ii) eNATL60 with no tidal  
 148 forcing eN60-NT. The two simulations have perfectly the same configuration except for the  
 149 inclusion of tidal motions in eN60-WT. In the rest of this article, we use eNATL60 to refer to  
 150 the two simulations while individual runs are addressed as eN60-WT (with tides) or eN60-  
 151 NT (no tides). The inclusion of tidal forcing in eN60-WT run provides conversion of tidal  
 152 energy into the internal wave field through, both, flow-topography interactions and wave-  
 153 balanced motions interactions (Arbic et al., 2008, 2018). Consequently, the comparison of  
 154 resolved dynamics between the two simulations allows diagnosing the impacts of the internal  
 155 tides on cross-scale energy transfer in the North Atlantic.

156 To investigate cross-scale energy exchanges between different scales of motions, we es-  
 157 timate kinetic energy spectral density in frequency-wavenumber space as a proxy to under-  
 158 stand the energetic nature of balanced/unbalanced motions in regimes with/without tidal  
 159 motions. Also, we estimate the rate at which nonlinear mechanisms exchange energy across  
 160 temporal and spatial scales in the two scenarios. In what follows, our analysis of kinetic

161 energy density and transfer is based on the hourly output of horizontal total velocity field  
 162 and are computed using the following equation;

$$\frac{\partial \widehat{KE}}{\partial t} = T_{KE} + \widehat{\mathbf{u}}^* \cdot OT \quad (1)$$

$$\widehat{KE} = \frac{1}{2} \widehat{\mathbf{u}}^* \cdot \widehat{\mathbf{u}} \quad (2)$$

$$T_{KE} = -\widehat{\mathbf{u}}^* \cdot [\mathbf{u} \cdot \widehat{\nabla} \mathbf{u}] \quad (3)$$

163 Equations (1)–(3) are derived from the Fourier transform of momentum equation multi-  
 164 plied by horizontal velocity field (Scott & Wang, 2005; Capet et al., 2008; Müller et al.,  
 165 2015). In the momentum equation (Equation 1),  $KE$  and  $T_{KE}$  represents the kinetic energy  
 166 density and kinetic energy transfer respectively while  $OT$  stands “Other Terms”.  $\widehat{[ ]}$  refers  
 167 to Fourier transform and  $*$  represents the complex conjugate. Before performing spectral  
 168 analysis the 2D time series were de-trended and windowed in space and time. The procedure  
 169 performed in this study is consistent with standard procedures previously used in Rocha et  
 170 al. (2016), Müller et al. (2015) and Torres et al. (2018).

171 The eNATL60 simulation resolves well to a reasonable extent, mesoscale motions, sub-  
 172 mesoscale motions, and IGWs (see Figure 1 in SI). The comparison of eNATL60 sea surface  
 173 height (SSH) spectral density with SARAL AltiKa (Figure 1a) shows that the predicted  
 174 SSH variance by the model compares well with the satellite observation for scales  $> 100$  km.  
 175 There are differences at scales  $< 100$  km that are due to the satellite instrument noise level.  
 176 There seems to be quite a robust agreement between the two runs of eNATL60 simulations in  
 177 wintertime. However, of particular interest is the difference between the runs in summertime  
 178 where variance at fine-scales is of higher magnitude in eN60-WT than to eN60-NT.

179 A similar analysis of the kinetic energy spectral density in the same region (Figure 1b),  
 180 shows that the variance associated with fine-scale motions smaller than 100 km is higher in  
 181 the eN60-WT compared to eN60-NT. So what are the mechanism/dynamics at fine-scales  
 182 in eN60-WT that could be responsible for this higher variance? A possible answer to this is  
 183 that the inclusion of tidal motion in eN60-WT simulation is responsible for enhanced wave  
 184 activity, and this is why we see higher variance at fine scales in the SSH and KE spectra  
 185 density plot. To qualitatively investigate this, we separate the flow into its rotational and  
 186 divergent part, which represents the balanced and the unbalanced wave motions, respec-  
 187 tively. Figure 1c presents the spectral density for these two components. The spectra of

the rotational part for the two runs are almost indistinguishable, indicating that both simulations have nearly equal energy levels in geostrophically balanced motions. However, the divergence spectra of the kinetic energy is very different between the two simulations. This difference is obvious at scales less than 500 km, and indeed, the divergent motions are more energetic in eN60-WT by a factor of 2 with two interesting peaks. We can conclude that the higher variance in eN60-WT at fine-scales compared to eN60-NT is primarily due to stronger divergent motions in eN60-WT, which is caused by the inclusion of tidal forcing in this simulation.

### 3 Seasonality of BMs and IGWs

This section presents the different classes of motions and their seasonality based on frequency-wavenumber spectral density. This diagnostic will help us better understand how the difference in wave activity between the two simulations affects oceanic motions' spectral signature across different temporal and spatial scales. For simplicity, we refer to frequency-horizontal wavenumber spectra as  $\omega$ - $K$  spectra. Following Torres et al. (2018), we begin by presenting a schematic (Figure 2a) showing the different observable dynamical regimes in the ocean as a function of their temporal and spatial scale. These classes of motions starting with low-frequency, low-wavenumber motions to high-frequency, high-wavenumber motions are Rossby waves (RW), mesoscale balanced motions (MBM), submesoscale balanced motions (SBM), unbalanced submesoscale motions (USM) and internal gravity waves (IGW). In this study, we focus on understanding how IGWs and BMs (in the presence of tidal motions) affect cross-scale energy exchanges. Due to the computational cost of this diagnostic tool, we perform the  $\omega$ - $K$  spectra analysis in a  $5^\circ \times 5^\circ$  ( $-40^\circ$  to  $-35^\circ$ ,  $40^\circ$  to  $45^\circ$ ) box located inside the previous large box (see Figure 1 of SI).

We also show, in Figure 2, the winter and summer averages of surface KE  $\omega$ - $K$  spectra for the two runs. In the figure, the upper panel corresponds to eN60-NT and the lower panel to eN60-WT. The classes of motions described previously in Figure 2a are identifiable in the figure. The winter-summer contrast shows a strong seasonality in the spectral density of submesoscale BMs and IGWs. In wintertime, for eN60-WT (simulation with tidal forcing), energy is mostly concentrated in BMs, near-inertial waves, and the dispersion curve of IGWs (Figure 2b), while in summertime, energy is mostly concentrated in the mesoscale BMs, near-inertial waves, and internal tides (Figure 2d). In particular, the variance associated with submesoscale BMs is stronger in winter, while that of IGWs is stronger in summer. Our understanding is that IGWs are stronger in summer due to shallow mixed layer depth and vertical normal modes' intensification. At the same time, submesoscale BMs are stronger in winter because they are driven by winter-favored mechanisms such as frontogenesis, wind-

induced frontal instabilities, and mixed layer instability, among other processes (Qiu et al., 2014; Sasaki et al., 2014; Callies, Flierl, et al., 2015; Brannigan et al., 2015; McWilliams, 2016; Uchida et al., 2019; Khatri et al., 2021). This out of phase seasonality of submesoscale BMs and IGWs is consistent with the findings of Rocha et al. (2016) and Torres et al. (2018).

Similarly, eN60-NT (simulation without tidal forcing) resolves the same classes of motion as eN60-WT except that internal tides are absent and supertidal IGWs are less energetic in this simulation. In wintertime, energy is mostly concentrated in BMs, near-inertial waves, and along the dispersion curve of IGWs (Figure 2a). This is consistent with the winter dynamics in eN60-WT. In summertime, energy is concentrated in mesoscale balanced motions and near-inertial waves (Figure 2c). Unlike eN60-WT, the seasonality observed in eN60-NT is associated with stronger submesoscale BMs and IGWs in winter. The seasonality of IGWs is reversed in eN60-NT when compared to eN60-WT (simulation with tidal forcing). How can this be? We know that the classical paradigm for the generation of the supertidal IGW continuum is that winds produce near-inertial waves, barotropic tidal flow over topographic features creates internal tides, and the energy along the IGW dispersion curve is due to nonlinear interactions. Both simulations are forced with realistic high-frequency winds with 3-hourly outputs (Brodeau et al., 2020). These winds are stronger in winter, hence there is a well-resolved near-inertial wave and IGWs dispersion curve in winter. The dynamics in summertime are different between the two runs. For the eN60-WT simulation, internal tides generated by tidal motions are amplified by a shallow mixed layer in summertime, and nonlinearity produces energy in the IGW dispersion curve. Thus for eN60-NT, the mechanism for generating waves in summertime is relatively weak; no tidal forcing and weaker winds, hence relatively weak wave motions in summer. The difference in the spectral energy density of different dynamical regimes between the two simulations extends below the surface (see Figure 3 of SI).

To obtain frequency spectra, we integrate the  $\omega$ - $K$  spectra over all wavenumbers for the two runs (Figure 3). In summertime, the variance at high frequencies ( $M_2$  and supertidal frequencies) is higher in eN60-WT compared to eN60-NT. We believe that this is likely due to the amplification of internal gravity waves by tidal motions. eN60-WT spectra approximately follow the estimated Garrett-Munk spectra (Garrett & Munk, 1975; Cairns & Williams, 1976; Müller et al., 2015) in summertime. Visible in eN60-WT spectra are the peaks at the inertia frequency  $f$  and the  $M_2$  tidal frequency. In contrast, only the near-inertial peak is visible in eN60-NT. To a large extent, we now understand the dynamics responsible for the differences in kinetic energy density that we see in the two simulations. In the following sections, we shall discuss how the enhanced IGWs arising from tidal forcing affect the redistribution of energy in different dynamical regimes.

259 **4 Modulation of KE forward flux by IGWs**

260 In this section, we will discuss the impact of resolving internal tides on the magnitude  
 261 and direction of KE cascade at fine-scales in the wavenumber and frequency-wavenumber  
 262 domain by comparing the twin eNATL60 runs.

263 **4.1 Wavenumber KE flux**

264 We start by examining the KE flux in the wavenumber domain. We do this by esti-  
 265 mating the net energy (spectral flux) passing through individual wavenumbers in spectral  
 266 space. The spectral flux is obtained by integrating the energy transfer (Equation 3) from a  
 267 particular wavenumber  $K$  to  $K_0$  (the wavenumber corresponding to the box size).

268 We present in Figure 4a,d, the winter and summer averages of kinetic energy spectral  
 269 flux for both simulations. In wintertime and in the two simulations (Figure 4a), the flux  
 270 is nearly identical across all wavenumbers. The forward cascade starts at around 25 km  
 271 and extends down to a kilometric scale. In summertime (Figure 4d), the magnitude of  
 272 the forward cascade at high wavenumbers differs significantly between the two runs. The  
 273 inclusion of tidal forcing in eN60-WT yields a forward flux at high wavenumbers that is a  
 274 factor of 4 higher than the forward high-wavenumber flux in eN60-NT. This difference in  
 275 cascade highlights how internal tides enhance the forward cascade of kinetic energy at high  
 276 wavenumbers.

277 Thus far, the kinetic energy cascade has been estimated only at the surface. Considering  
 278 we have three-dimensional information of the ocean, as opposed to satellite observations, it is  
 279 of great interest to understand the nature of the kinetic energy cascade in the ocean's interior.  
 280 In Figure 4b,c,e,f, we present the spectral flux computed at 32 different vertical levels in  
 281 the upper 1000 m of the water column. In winter and summer, at lower wavenumbers,  
 282 the average KE flux in the two simulations is characterized by a net inverse cascade that  
 283 extends down to around 700 m in the interior. In wintertime, the forward cascade at higher  
 284 wavenumbers in eN60-WT (Figure 4c) is strong both at the surface and in the interior. In  
 285 contrast, in eN60-NT (Figure 4b), the forward cascade is confined mostly to the surface. In  
 286 summertime, the forward cascade in eN60-WT (Figure 4f) span the upper ocean but with  
 287 a gradual decrease in magnitude farther down the water column. In contrast, in eN60-NT  
 288 (Figure 4e), the forward cascade is nearly absent throughout the upper ocean. A stronger  
 289 forward cascade (in summertime for eN60-WT) in the interior is an indication that internal  
 290 tides transfer substantial amounts of kinetic energy to dissipative scales throughout the  
 291 upper ocean water column.

292        **4.2 Frequency-wavenumber KE transfer**

293        We have demonstrated that internal tides enhance the supertidal IGW continuum and,  
 294        in particular, enhance the forward cascade of energy in summertime. To better explain how  
 295        internal tides modify cross-scale energy exchanges among the different classes of motions,  
 296        we present in Figure 5 the winter and summer averages of kinetic energy spectral transfer in  
 297        frequency-wavenumber space. In the  $\omega$ - $K$  spectra, negative values of spectra transfer imply  
 298        that non-linearity extracts energy from these regions to feed other regions with positive  
 299        values. In other words, sinks of energy are characterized by positive values, while sources  
 300        of energy have negative values.

301        We start by discussing the spectral transfer in wintertime (cf. Equation 3; Figures 5a,c).  
 302        In eN60-NT, the spectral transfer's positive values (blue) show that balanced motions serve  
 303        as a source of energy for other motions, namely energy is being extracted from the balanced  
 304        motions. In contrast, near-inertial motions and motions with scales less than 10 km are the  
 305        major sinks of kinetic energy (red). The rate of non-linear exchanges in eN60-WT is similar  
 306        to eN60-NT except for the mild intensification of energy gained by IGWs in eN60-WT. To  
 307        summarize, submesoscale motions and internal gravity waves are sinks of kinetic energy in  
 308        wintertime, with the former playing the major role.

309        The summer spectra (Figure 5b,d) differ significantly between the two runs as expected.  
 310        In eN60-NT, balanced motions represent the major source of energy for other motions,  
 311        while energy is gained mostly by near-inertial motions. The transfer at high frequencies and  
 312        wavenumbers is very small. We can interpret this to signify that high-frequency motions  
 313        and submesoscale motions are less energetic in eN60-NT in summertime. This result is  
 314        consistent with results in the KE spectra density. But in eN60-WT, the situation is different.  
 315        The major energy source for other motions is the mesoscale balanced motions and the  
 316        semi-diurnal tides (blue). Near-inertial waves, submesoscale BMs and the supertidal IGW  
 317        continuum spectrum are seen in Figure 5 to be gaining energy. The extraction of semi-  
 318        diurnal tidal energy, and gain of energy in the IGW continuum, due to nonlinear interactions  
 319        was also seen in (Müller et al., 2015) but is more clear here, perhaps because of the finer  
 320        vertical and horizontal grid spacing in eN60-WT.

321        In summary, there are two mechanisms of energy extraction in summertime. (i) In a  
 322        flow without tidal forcing, near-inertial waves gain energy from balanced motions, and (ii) in  
 323        a flow with tidal forcing, near-inertial waves and the supertidal IGW continuum gain energy  
 324        from internal tides and mesoscale balanced motions. In summary, the forward cascade in  
 325        eN60-WT, is associated with the transfer of energy by nonlinearity from balanced motions  
 326        and internal tides to near-inertial waves and the supertidal IGW continuum. This transfer

327 is possible due to the intensification of IGWs in the presence of tidal motions. The lack of  
 328 tidal motions in eN60-NT (compared to eN60-WT) shows how effective internal tides are  
 329 in providing a route to energy dissipation in summertime, both at the surface and in the  
 330 interior. This result strongly emphasizes the need to include tidal forcing in ocean model  
 331 simulations to accurately predict cross-scale energy exchanges at fine-scales.

## 332 5 Summary and conclusion

333 The role of internal tides in the seasonality of kinetic energy density and spectral kinetic  
 334 energy transfer was investigated in this study. Our analysis was based on the output of a  
 335 realistic NEMO simulation of the North Atlantic Ocean with a horizontal resolution of  
 336  $1/60^\circ$ . We used two outputs of this numerical experiment; one with, and one without,  
 337 tidal forcing. These twin experiments permit investigation of how internal gravity waves (in  
 338 particular internal tides) generated by tidal forcing modify kinetic energy variance, cross-  
 339 scale exchanges, and associated seasonality. We found that the seasonality of IGWs is  
 340 sensitive to tidal forcing. In the simulation without tides, IGWs are stronger in winter,  
 341 whereas, in simulation with tides, they are stronger in summer. The latter condition is  
 342 consistent with the findings of Rocha et al. (2016) and Torres et al. (2018).

343 Our results also show that resolving internal tides in the presence of energetic subme-  
 344 soscale motions has a strong impact on kinetic energy transfer in summertime (cf. Barkan  
 345 et al., 2021, their Figure 3d,f). The magnitude of the estimated forward cascade at high  
 346 wavenumbers (both at the surface and at depth) in the simulation with tidal forcing is a fac-  
 347 tor 4 higher than in the simulation without tidal forcing (Figure 4). Overall, we identified  
 348 that two mechanisms supporting the kinetic energy forward cascade; (i) forward cascade  
 349 due to energetic submesoscale motions in wintertime and (ii) forward cascade due to IGWs  
 350 (enhanced by tidal forcing) in summertime (Figure 5). Our results underscore that at fine-  
 351 scales, internal tides can provide an effective route to kinetic energy dissipation.

352 In light of the SWOT altimeter mission (Morrow et al., 2019; Torres et al., 2019), the  
 353 difference between the runs with tidal forcing and without (e.g. Figures 1 and 3) highlight  
 354 the importance of tidal forcing in emulating the upcoming altimeter observations (Arbic et  
 355 al., 2018; Barkan et al., 2021; Yu et al., 2021).

356

**357 Data Availability Statement**

358 The twin eNATL60 data are permanently stored at the CINES supercomputing center  
359 in Montpellier, France. The regional dataset used for this study is available upon request  
360 to AAI ([aurelie.albert@univ-grenoble-alpes.fr](mailto:aurelie.albert@univ-grenoble-alpes.fr)).

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375 **References**

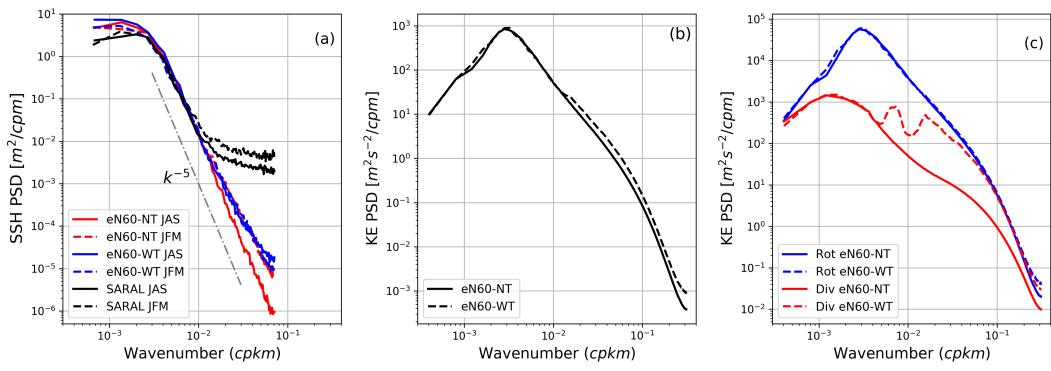
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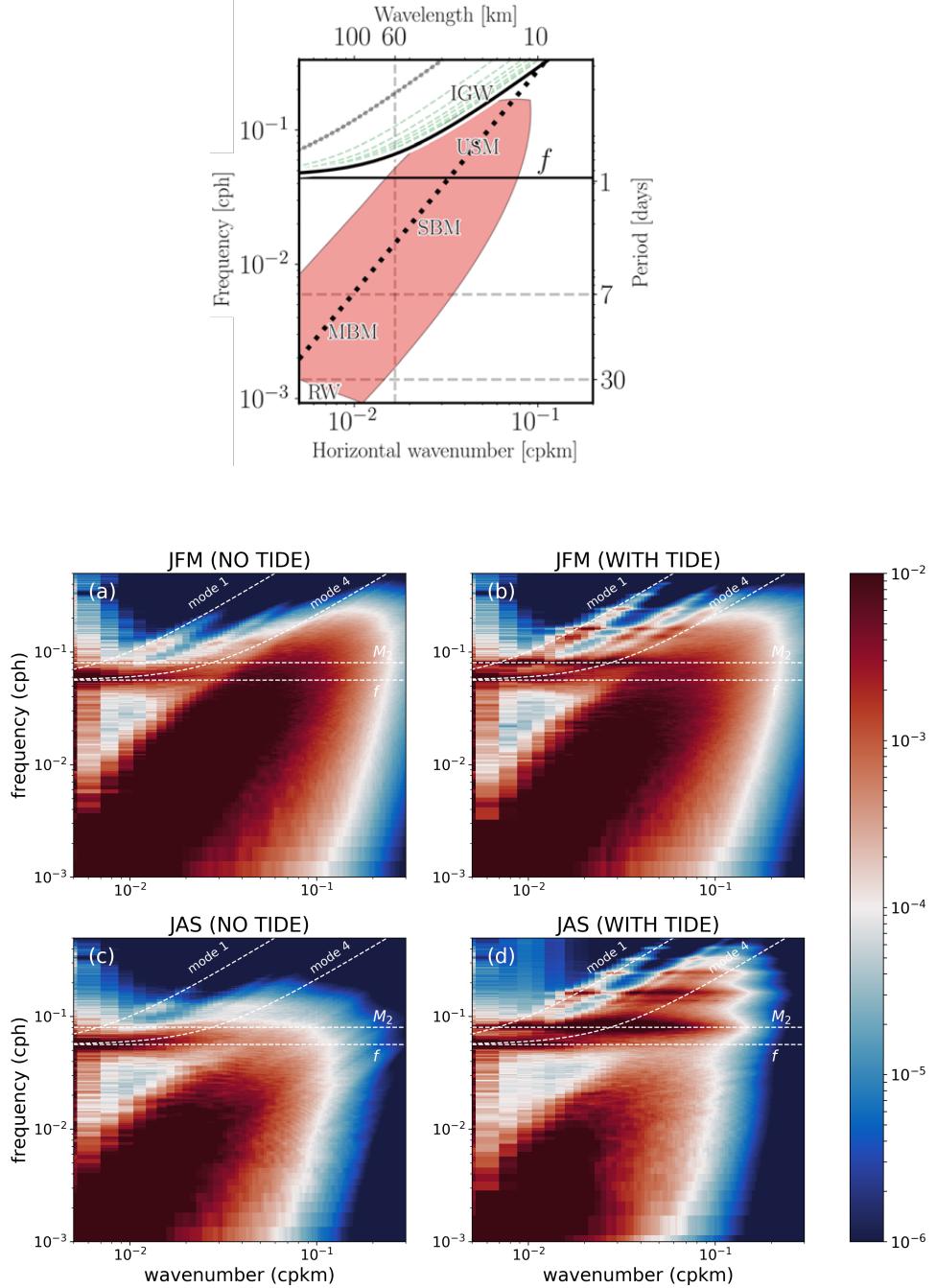
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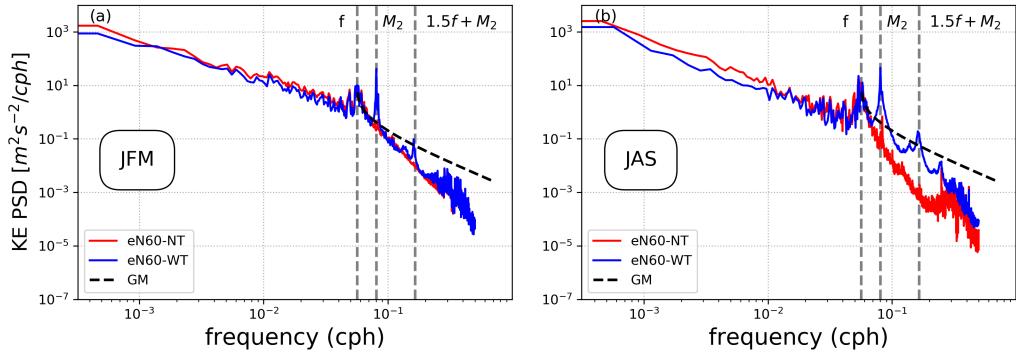
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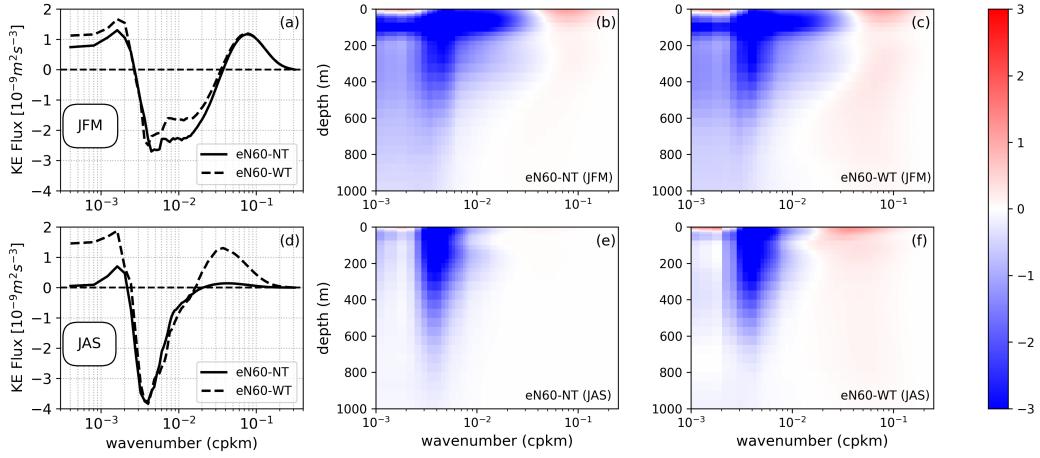
**Figure 1.** (a) Comparison of SSH wavenumber spectra between eNATL60 and SARAL AltiKa satellite. (b) One year average of surface kinetic energy wavenumber spectral density computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides). (c) Helmholtz decomposition of kinetic energy into rotational ( $\zeta$ ) and divergent ( $\delta$ ) spectral components. Thick curves represent the simulation without tides (eN60-NT) and dashed curves represent the simulation with tides (eN60-WT).



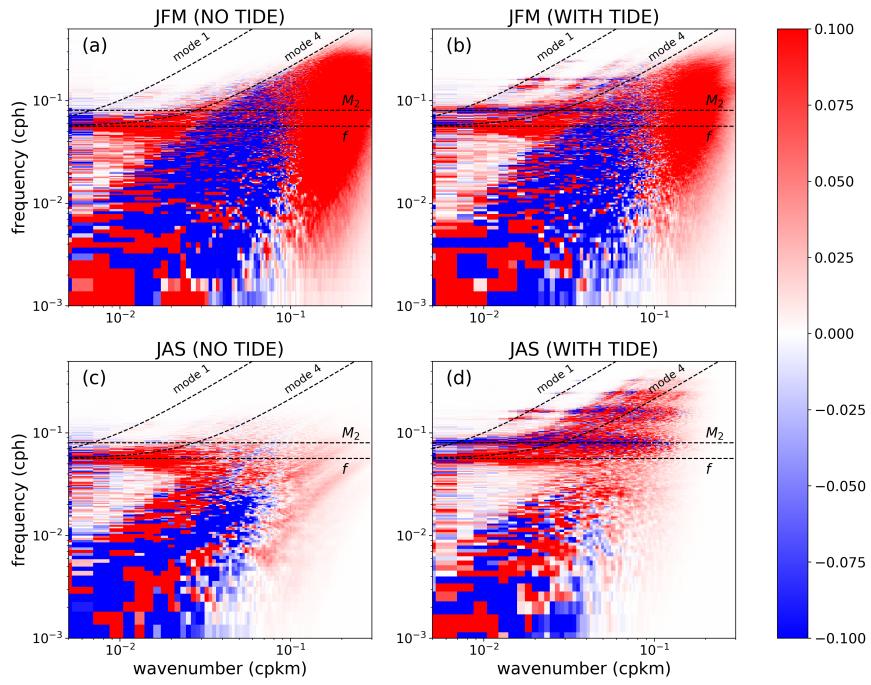
**Figure 2.** Top panel: A schematic of the observable dynamical regimes with different classes of motions in the ocean. These classes of motions starting with low frequency, small wavenumber motions to high frequency frequency, high wavenumber motions are Rossby waves(RW), mesoscale balanced motions (MBM), submesoscale balanced motions (SBM), unbalanced submesoscale motions (USM) and internal gravity waves (IGW). Bottom four plots: Surface kinetic energy frequency-wavenumber spectra computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides) for winter (JFM) and summer (JAS) time.



**Figure 3.** Surface Kinetic energy frequency spectral density, obtained by integrating the  $\omega\text{-}K$  spectra over all wavenumbers for eN60-NT (no tides) and eN60-WT (with tides). The three grey dash lines represent the inertia frequency  $f$ , the  $M_2$  tidal frequency, and  $1.5f+M_2$ . The dashed black line represents the estimate of the Garrett-Munk spectra computed with reference values of total energy of the internal wavefield and stratification set to  $E_0 = 6.3e^{-5} m^2 s^{-2}$  and  $N_0 = 5.2e^{-3} s^{-1}$ , respectively (Garrett & Munk, 1975; Cairns & Williams, 1976; Müller et al., 2015). (a) Winter (JFM) and (b) Summer (JAS).



**Figure 4.** Surface kinetic energy spectral flux computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides). Summer : July, August and September. Winter : January, February and March. (b,c,e,f) Winter and summer averages of kinetic energy spectral flux as a function of depth for eN60-NT and eN60-WT.



**Figure 5.** Surface kinetic energy transfer in frequency-wavenumber space computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides) for winter (JFM) and summer (JAS) time.