

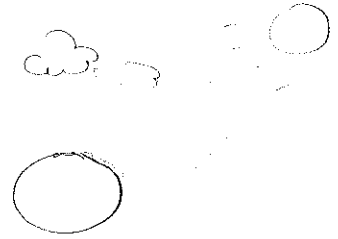
II.4 Convective instability

①

II.4.1 Rayleigh-Bénard instability

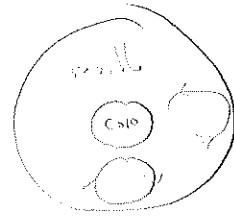
Instability caused by a vertical gradient of T
(cold on top of warm).

= atmosphere heated from below by the ground



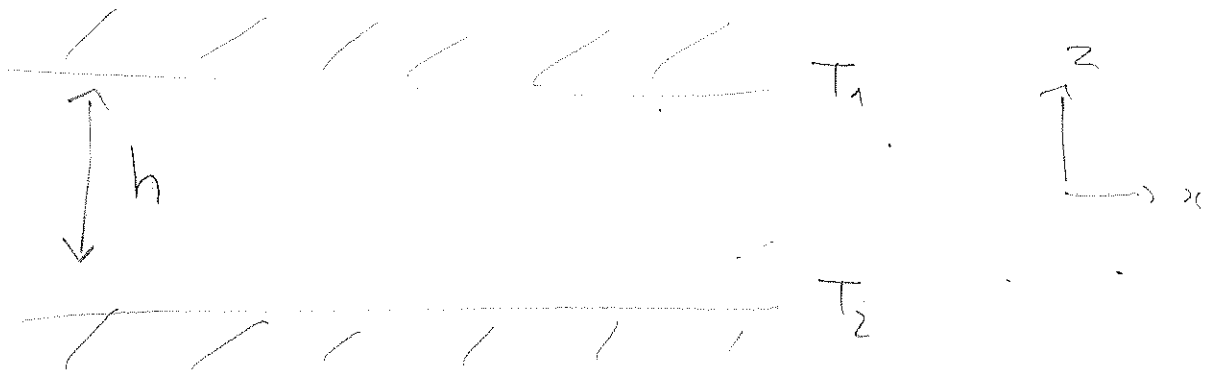
= mantle convection:

= oil in a pan



- the flow is destabilized by buoyancy forces.
- in competition with viscosity and diffusivity.

ex: Rayleigh-Bénard (Detailed analysis by Rayleigh)



② Some dimensionless numbers:

physical parameters are

ν = kinematic viscosity

k = thermal diffusivity

h = height of fluid.

(heat equation $\frac{\partial T}{\partial t} = k \nabla^2 T$)

• acceleration due to buoyancy:?

Linear expansion of state for $e(T)$, with $e(T_0) = e_0$

$\rightarrow \underline{e} = e_0 (1 - \alpha(T - T_0))$ $\alpha = \frac{1}{e} \left(\frac{de}{dT} \right)_p$

the thermal expansion coef.

So the acceleration is $\left| g \frac{\Delta e}{e_0} = g \alpha \Delta T \right|$

With these 4 parameters, we can define 2 numbers:

① The Prandtl number $\left[Pr = \frac{\nu}{k} = \frac{\text{viscosity}}{\text{diffusivity}} \right]$

metal $Pr \ll 1 \rightarrow$ heat diffuses quickly = conductor

oil $Pr \gg 1 \rightarrow$ heat diffuses slowly = insulator

② The Rayleigh number $\left[Ra = \frac{g \alpha \Delta T h^3}{\nu k} \right]$

experimentally: fluid is at rest if $Ra < Ra_c$
a critical value. if $Ra > Ra_c$, the
flow became unstable (bifurcation \rightarrow full turbulence)

- if a fluid parcel is accelerated over the height h during a time τ_0 by buoyancy forces

$$g\tau_0/\Delta T \sim \frac{h}{\tau_0^2}$$

- effect of viscosity (smoothing velocity gradients) corresponds to a time-scale $\tau_v \sim \frac{h^2}{\nu}$
- effect of diffusivity $\rightarrow \tau_T \sim \frac{h^2}{k}$

So the Rayleigh number corresponds to

$$Ra = \frac{\tau_v \tau_T}{\tau_0^2}$$

- if Ra is small, the diffusion of momentum and T is being compared to time the accelerate.
= fluctuations are damped
viscous and diffusive effects dominate the flow.
- if Ra is large, it is the opposite...

(b) Use the equations to find critical Ra

We use the NS equations

$$\frac{\partial \vec{u}}{\partial t} = - \frac{\vec{\nabla} P}{\rho} + \vec{g} + \nu \nabla^2 \vec{u}$$

$$\frac{\partial T}{\partial t} = k \nabla^2 T$$

$$\vec{\nabla} \cdot \vec{u} = 0 \quad (\text{we neglect compressibility})$$

with $\rho = \rho_0 (1 - \alpha (T - T_0))$ linear expansion

The balanced state is

$$\left\{ \begin{array}{l} \vec{U}_0 = \vec{0} \\ - \frac{1}{\rho_0} \frac{\partial P_0}{\partial z} - g = 0 \end{array} \right.$$

$$\nabla^2 T_0 = 0$$

$$\Rightarrow T_0 = \Lambda_1 z + \Lambda_2$$

$$\text{with } \left\{ \begin{array}{l} T_0(0) = T_2 \\ T_0(h) = T_1 \end{array} \right.$$

$$\Rightarrow T_0 = \frac{T_1 - T_2}{h} z + T_2$$

$$\left\{ T_0(z) = T_2 - \frac{\Delta T z}{h} \right.$$

We look at small perturbations around this state:

$$\rho = \rho_0 + \rho'$$

$$P = P_0 + P'$$

Using the Boussinesq approximation $\left(\begin{array}{l} e' \ll e_0 \\ \alpha \Delta T \ll 1 \end{array} \right)$ (3)

we get
$$\frac{D\vec{u}'}{Dt} = - \frac{\vec{\nabla} p'}{e_0} + \vec{g} \frac{e'}{e_0} + \nu \nabla^2 \vec{u}'$$

$$\Rightarrow \frac{D\vec{u}'}{Dt} = - \frac{\vec{\nabla} p'}{e_0} + \vec{g} \frac{e'}{e_0} + \nu \nabla^2 \vec{u}' - \underbrace{\frac{\vec{\nabla} p_0}{e_0} + g \frac{e_0}{e_0}}_{=0}$$

we linearize:

$$\frac{\partial \vec{u}'}{\partial t} = - \frac{\vec{\nabla} p'}{e_0} + \vec{g} \frac{e'}{e_0} + \nu \nabla^2 \vec{u}'$$

$$\textcircled{1} \quad \frac{\partial \vec{u}'}{\partial t} = - \frac{\vec{\nabla} p'}{e_0} + g \alpha T \vec{k} + \nu \nabla^2 \vec{u}'$$

$$\textcircled{2} \quad T_h' + \omega \frac{\partial T_h'}{\partial z} = T_h' - \omega \frac{\Delta T}{h} = k \nabla^2 T'$$

$$\textcircled{3} \quad \vec{\nabla} \cdot \vec{u}' = 0$$

we nondimensionalize:

$$x \rightarrow \frac{x}{h}$$

$$t \rightarrow \frac{t}{\tau} = \frac{t k}{h^2}$$

$$u \rightarrow u \frac{h}{\nu}$$

$$T \rightarrow \frac{T'}{\Delta T}$$

$$p \rightarrow \frac{\rho \alpha^2 \Delta T^2}{e_0 \nu^2}$$

The equations become:

$$\left\{ \begin{array}{l} \frac{\partial \bar{\psi}'}{\partial t} = -Pr \bar{\nabla} \bar{p}' + Ra T \bar{k}' + Pr \nabla'^2 \bar{\psi}' \\ \frac{\partial T'}{\partial t} - Pr w = \nabla'^2 T' \\ \bar{\nabla}' \cdot \bar{\psi}' = 0 \end{array} \right. \quad \text{with} \quad Pr = \frac{\nu \rho c_p}{k}$$

$$Ra = \frac{g \alpha \Delta T h^3}{\nu k}$$

we also define boundary conditions:

$$w = 0 \quad \text{at} \quad z = 0, 1$$

no normal velocity

$$T' = 0 \quad \text{at} \quad z = 0, 1$$

no T fluctuations
at the boundary

we can now reduce the system to one variable (4)

momentum eqns:

$$\textcircled{1} \quad w_t = -P_r \rho_z + R_0 T + P_r \nabla^2 w$$

$$\textcircled{2} \quad u_t = -P_r \rho_x + P_r \nabla^2 u$$

$$\textcircled{3} \quad v_t = -P_r \rho_y + P_r \nabla^2 v$$

$$\nabla_x \textcircled{2} + \nabla_y \textcircled{3} + \nabla_z \textcircled{1}$$

$$\Rightarrow \underbrace{(u_{xt} + v_{yt} + w_{zt})}_{=0} = -P_r \nabla^2 \rho + P_r \underbrace{\nabla^2 (u_x + v_y + w_z)}_{=0} + R_0 T_z$$

$$\Rightarrow R_0 T_z = P_r \nabla^2 \rho$$

$$\text{So } \nabla^2 \textcircled{1} \Rightarrow \nabla^2 w_t = -P_r \nabla^2 \rho_z + R_0 \nabla^2 T + P_r \nabla^4 w \\ = -R_0 T_{zz} + R_0 \nabla^2 T + P_r \nabla^4 w$$

$$\nabla^2 w_t = R_0 \nabla_h^2 T + P_r \nabla^4 w$$

$$\Rightarrow (\nabla_t - P_r \nabla^2) \nabla^2 w = R_0 \nabla_h^2 T \quad \textcircled{5}$$

$$\text{using } T_t - P_r w = \nabla^2 T$$

$$\Rightarrow (\nabla_t - \nabla^2) T = P_r w$$

to eliminate T in $\textcircled{5}$, we get.

$$\begin{aligned}
 (\Delta_r - \nabla^2)(\Delta_r - P_r \nabla^2) \nabla^2 \omega &= R_0 \nabla_h^2 ((\Delta_r - \nabla^2) T) \\
 &= R_0 P_r \nabla_h^2 \omega \quad (6)
 \end{aligned}$$

with boundary conditions $T=0 \Rightarrow \nabla_h^2 T=0 \mid_{z=0,1}$

so $\nabla^2 \omega_r = P_r \nabla_h^2 \omega \mid_{z=0,1}$

we look for solutions in the form

$$\omega = \hat{\omega}(z) e^{i k x - s t}$$

with $|k| = k$

\Rightarrow using (6) :

$$\begin{aligned}
 \left(i s + k^2 - \frac{\nabla^2}{S_2^2} \right) \left(s + P_r k^2 - P_r \frac{\nabla^2}{S_2^2} \right) \left(-k^2 + \frac{\nabla^2}{S_2^2} \right) \omega \\
 = -R_0 P_r k^2 \omega \quad (7)
 \end{aligned}$$

we can show by multiplying (7) by ω^* and integrating between $z=0,1$ that

s is a real if $R_0 > 0$.

This is called the exchange of stabilities (3)

The marginal stability corresponds to the case $s=0$
this will give us the minimum R_0 such that
the flow is stable.

$$\Rightarrow \left(k^2 - \frac{\gamma^2}{\gamma_2^2}\right) \left(k^2 - \frac{\gamma^2}{\gamma_1^2}\right) \left(\frac{\gamma^2}{\gamma_1^2} - k^2\right) \omega = -R_0 k^2 \omega$$

$$\Rightarrow \left(k^2 - \frac{\gamma^2}{\gamma_1^2}\right)^3 \omega = R_0 k^2 \omega$$

(independent of P_0)

• if we choose free slip conditions: (not realistic but easy to solve)

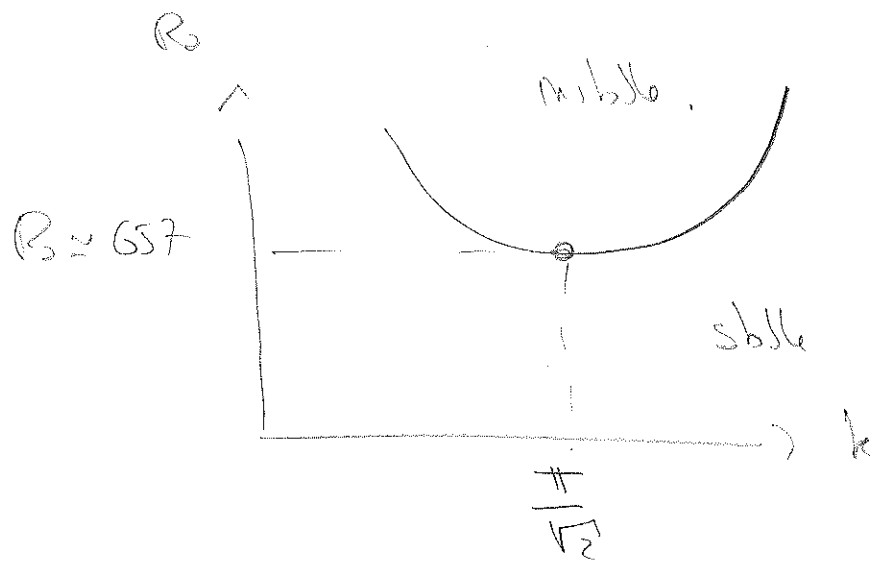
$$u_z = v_z = 0 \quad \text{at } z=0,1$$

we get $w_{zz} = 0$ at $z=0,1$

$$\Rightarrow \text{solution is } w(z) = \sin(n\pi z)$$

which gives marginal stability curve:

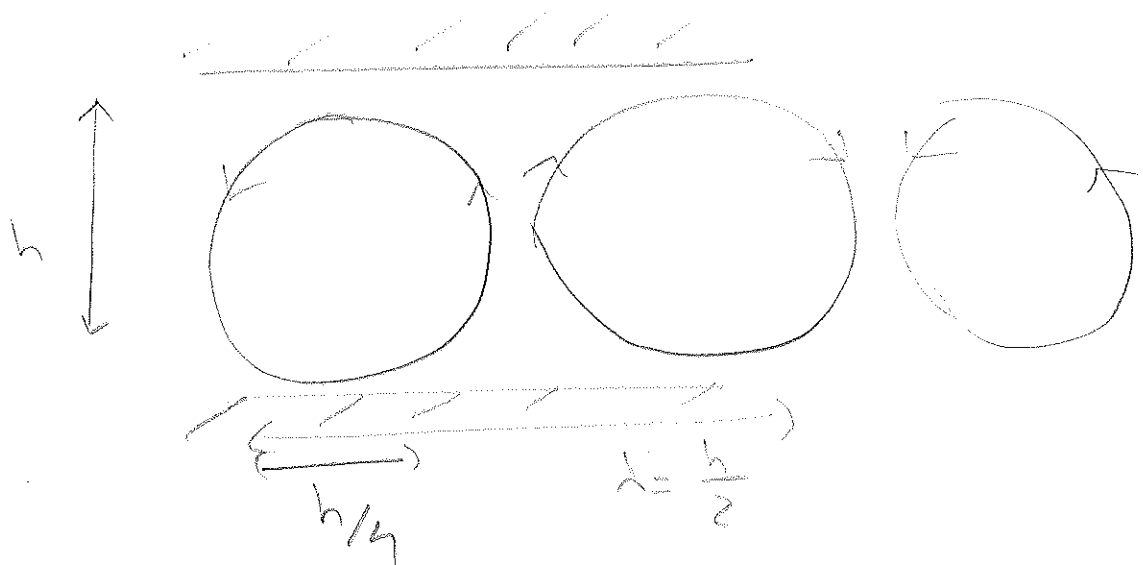
$$R_0 = \frac{(n^2 \pi^2 + k^2)^3}{k^2}$$



For a given k , smaller values of Ra_c corresponds to $n=1$. and the minimum of the curve is for $k = \frac{\pi}{\sqrt{2}} \approx 2$

The aspect ratio is close to 1.

(For larger k , the flow will be stabilized by viscous effects)



• For larger Rayleigh it will become unstable.