

# Assignment 1

AST4320: Cosmology and Extragalactic Astronomy

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## Exercise 1

The continuity equation for an unperturbed universe driven by hubble expansion is given as

$$\frac{d\bar{\rho}}{dt} + \bar{\rho} \nabla \cdot \mathbf{v}_0 = 0, \quad (1)$$

where  $\bar{\rho}$  is the average density of the universe, and  $\mathbf{v}_0$  comes from the Hubble law today which states that

$$\mathbf{v}_0 = H_0 \mathbf{r}. \quad (2)$$

Here  $H_0$  is the Hubble parameter today, given as  $H_0 = \dot{a}(t = t_0)/a(t = t_0)$ , where  $a$  is the scale factor, and is defined so that  $a(t = t_0) = 1$ .

Inserting for (2) into the continuity equation we find

$$\frac{d\bar{\rho}}{dt} + \bar{\rho} \nabla \cdot \mathbf{r} \frac{da}{dt} \frac{1}{a} = 0,$$

where the del operator only works on  $\mathbf{r}$  as its the scale factor is only a function of time. The operation results in  $\nabla \cdot \mathbf{r} = 3$ . We can then swap the right term over to the right hand side of the equation which leaves us with

$$\frac{d\bar{\rho}}{dt} = -3\bar{\rho} \frac{da}{dt} \frac{1}{a}.$$

Further, we can separate the equation

$$\frac{d\bar{\rho}}{\bar{\rho}} = -3 \frac{da}{a}.$$

This is then solved in the following way

$$\begin{aligned} \int_{\bar{\rho}(t=t_0)}^{\bar{\rho}(t)} \frac{d\bar{\rho}}{\bar{\rho}} &= \int_{a(t=t_0)}^{a(t)} -3 \frac{da}{a}, \\ \Rightarrow \ln \left( \frac{\bar{\rho}(t)}{\bar{\rho}(t=t_0)} \right) &= \ln \left( \frac{a(t)}{a(t=t_0)} \right)^{-3}. \end{aligned}$$

This can then be reduced to

$$\bar{\rho}(t) = \bar{\rho}(t=t_0) a(t)^{-3}, \quad (3)$$

where we have used that  $a(t=t_0) = 1$ .

## Exercise 2

The Poisson equation is given as

$$\nabla^2 \quad (4)$$