

Chapter 1

Introduction to Quantum Optics by Immanuel

1.1 Introduction

- classicaal: classical atom and light
- semiclassical: quantized atom and classical light
- quantum mechanical: quantized atom and light

Light-Atom Interaction Hamiltonian

- classical dipole in electric field: dipole moment $\vec{d} = q\vec{r}$, $U_I = -\vec{d} \cdot \vec{E}$. We have

$$\hat{H}_I = -\hat{d} \cdot \vec{E}(\vec{v}_0, t), \quad (1.1)$$

where $\hat{d} = q\hat{v}$ is the dipole operator.

- induced atomic dipole

1.2 Light Atom Quantum Evolution

Time Evolution We have the Schrodinger equation (both sides) as

$$i\hbar \frac{\partial}{\partial t} |\Psi(t)\rangle = (\hat{H}_0 + \hat{H}_I(t)) |\Psi(t)\rangle, \quad (1.2)$$

where the general ansatz (assumption) is

$$|\Psi(t)\rangle = \sum_n c_n(t) e^{-iE_n t/\hbar} |n\rangle, \quad (1.3)$$

and

$$\hat{H}_0 |n\rangle = E_n |n\rangle \quad (1.4)$$

is the atomic eigenstates. Inserting $|\Psi(t)\rangle$ and $\hat{H}_0|n\rangle$ into Schrodinger equation, we get

$$i\hbar \sum_n \left\{ \dot{c}_n e^{-iE_n t/\hbar} |n\rangle - \frac{iE_n}{\hbar} c_n e^{-iE_n t/\hbar} |n\rangle \right\} = \sum_n \left\{ c_n e^{-iE_n t/\hbar} |n\rangle + c_n e^{-iE_n t/\hbar} \hat{H}_I |n\rangle \right\} \quad (1.5)$$

$$\implies i\hbar \sum_n \dot{c}_n e^{-iE_n t/\hbar} |n\rangle = \sum_n c_n e^{-iE_n t/\hbar} \hat{H}_I |n\rangle \quad (1.6)$$

$$\implies i\hbar \dot{c}_n e^{-iE_n t/\hbar} = \sum_n c_n(t) e^{-iE_n t/\hbar} \langle n | \hat{H}_I(t) | n \rangle \quad (1.7)$$

$$\implies i\hbar \dot{c}_k = \sum_n c_n(t) e^{-iE_{n,k} t/\hbar} \langle k | \hat{H}_I(t) | n \rangle, \quad (1.8)$$

where we use

$$\langle k | n \rangle = \delta_{kn}, \quad (1.9)$$

$$E_{n,k} = E_n - E_k, \quad (1.10)$$

$$\omega_{nk} = (E_n - E_k)/\hbar. \quad (1.11)$$

and $\langle k | \hat{H}_I(t) | n \rangle$ is the matrix element.

1.3 Time Dependent Perturbation Theory

Recall the time evolution:

$$i\hbar \dot{c}_k = \sum_n c_n(t) e^{-i\omega_{nk} t} \langle k | \hat{H}_I(t) | n \rangle, \quad (1.12)$$

and

$$\omega_{nk} = (E_n - E_k)/\hbar. \quad (1.13)$$

Consider the Simplification (Perturbation Theory)

- System only in state $|1\rangle$ at $t = 0 \implies c_1|0\rangle = 1$ (only the ground state $|1\rangle$),
- Perturbative treatment of interaction term: weak perturbation $\forall |c_k(t)|^2 \ll 1$.

We then have

$$i\hbar \dot{c}_k = e^{i\omega_{1k} t} \langle k | \hat{H}_I(t) | 1 \rangle, \quad (1.14)$$

with $c_k(0) = 0$, we obtain:

$$c_k(t) = \frac{1}{i\hbar} \int_0^t e^{-i\omega_{1k} t'} \langle k | \hat{H}_I(t') | 1 \rangle dt'. \quad (1.15)$$

Example 1.3.1 (Sinusoidal perturbation). Define

$$\hat{H}(t) = \hat{H}_I e^{-i\omega t}. \quad (1.16)$$

Given the figure in the video, we have

$$c_k(T) = \frac{1}{i\hbar} \int_0^T e^{i\Delta\omega t} \langle k | \hat{H}_I | 1 \rangle dt \quad (1.17)$$

$$\implies \text{Transition probability } P_{k1}(T) = |c_k(T)|^2 = \frac{1}{\hbar^2} |\langle k | \hat{H}_I | 1 \rangle|^2 Y(\Delta\omega, T), \quad (1.18)$$

with

$$Y(\Delta\omega, T) = \frac{\sin^2(\Delta\omega T/2)}{(\Delta\omega/2)^2} \quad (1.19)$$

$$\sim \text{sinc}^2 x, \quad (1.20)$$

where $\Delta\omega = \omega - \omega_{1k}$ is the detuning.

Let's take a look at the sinc function $Y(\Delta\omega, T) = \text{sinc}^2 x$. Transition for $\Delta\omega \leq \frac{2\pi}{T}$, we have $\Delta\omega \cdot T \leq 2\pi$, which implies

$$\Delta E \cdot T \leq h, \quad (1.21)$$

which is the time-frequency uncertainty. (The expression in the video seems wrong, so I make corrections above.) We have the following case

$$\frac{1}{2\pi T} Y(\Delta\omega, T) \xrightarrow{T \rightarrow \infty} \delta(\Delta\omega), \quad (1.22)$$

then we have

$$P_{k1}(T \rightarrow \infty) = \frac{2\pi}{\hbar^2} |\langle k | \hat{H}_I | 1 \rangle|^2 \delta(\Delta\omega) T. \quad (1.23)$$

Fermi's Golden Rule $|k\rangle$ Quasi continuum of final states. We have the transition probability

$$P_{k1} = \Gamma_{k1} T, \quad (1.24)$$

where

$$\Gamma_{k1} = \frac{2\pi}{\hbar} |\langle k | \hat{H}_I | 1 \rangle|^2 \rho(E_k = E_1 + \hbar\omega) \quad (1.25)$$

is called the Fermi's Golden Rule,

$$|\langle k | \hat{H}_I | 1 \rangle|^2 \quad (1.26)$$

is the coupling strength $\propto E_0^2$ and $\propto I$,

$$\rho(E_k = E_1 + \hbar\omega) \quad (1.27)$$

is the density states which is number of available final states to the system,

$$\Gamma_{k1} \hat{=} \text{Transition Rate} = \frac{dP_{k1}}{dT}, \quad (1.28)$$

and density states

$$\rho(E) = \frac{dN}{dE}, \quad (1.29)$$

where ΔN is the number of states in an energy interval ΔE around energy E_k and we let ΔE approaches 0.

1.4 Two Level Atom (TLA)

Given by the figure, in state $|1\rangle$, we have $E_1 = \hbar\omega_1$ and in state $|2\rangle$, we have $E_2 = \hbar\omega_2$ and $E_2 - E_1 = \hbar(\omega_2 - \omega_1) = \omega_{21}$. We have the Hamiltonian

$$\hat{H} = \hat{H}_0 - \hat{d} \cdot E(t), \quad (1.30)$$

where

$$E(t) = \varepsilon E_0 \cos(\omega t), \quad (1.31)$$

where ε is the polarization vector, E_0 is the field amplitude, and ω is the frequency of the light field.

Ansatz for Solving TLA We have

$$|\Psi(t)\rangle = c_1(t)e^{-i\omega_1 t}|1\rangle + c_2(t)e^{-i\omega_2 t}|2\rangle. \quad (1.32)$$

Time Evolution Amplitude We have

$$\dot{c}_1(t) = i \frac{d_{12}^\varepsilon E_0}{\hbar} e^{-\omega_{21} t} \cos(\omega t) c_2(t) \quad (1.33)$$

$$\dot{c}_2(t) = i \frac{d_{12}^\varepsilon E_0}{\hbar} e^{+\omega_{21} t} \cos(\omega t) c_1(t), \quad (1.34)$$

where

$$d_{12}^\varepsilon = \langle 1 | \hat{d} \cdot \varepsilon | 2 \rangle \quad (1.35)$$

$$= \langle 1 | \hat{d} | 2 \rangle \cdot \varepsilon \quad (1.36)$$

$$= \langle 1 | \hat{d}_x | 2 \rangle \cdot \varepsilon_x + \langle 1 | \hat{d}_y | 2 \rangle \cdot \varepsilon_y + \langle 1 | \hat{d}_z | 2 \rangle \cdot \varepsilon_z. \quad (1.37)$$

is the Dipole Matrix Element, which is the atomic property and we assume it's real. We also define

$$\Omega_0 = \frac{d_{12}^\varepsilon E_0}{\hbar} \quad (1.38)$$

as the Rabi frequency.

Time Evolution Using Euler' form, we have

$$\dot{c}_1(t) = i \frac{\Omega_0}{2} e^{-\omega_{21} t} (e^{i\omega t} + e^{-i\omega t}) c_2(t) \quad (1.39)$$

$$\dot{c}_2(t) = i \frac{\Omega_0}{2} e^{+\omega_{21} t} (e^{i\omega t} + e^{-i\omega t}) c_1(t) \quad (1.40)$$

by

$$\cos \alpha = \frac{1}{2}(e^{i\alpha} + e^{-i\alpha}) \quad (1.41)$$

and

$$e^{i\alpha} = \cos \alpha + i \sin \alpha. \quad (1.42)$$

Rotating Wave Approximation We have

$$\dot{c}_1(t) = i\frac{\Omega_0}{2}(e^{+i(\omega-\omega_{21})t} + e^{-i(\omega+\omega_{21})t})c_2(t) \quad (1.43)$$

$$\dot{c}_2(t) = i\frac{\Omega_0}{2}(e^{-i(\omega-\omega_{21})t} + e^{+i(\omega+\omega_{21})t})c_1(t), \quad (1.44)$$

and we ignore the sum frequency term and get

$$\dot{c}_1(t) = i\frac{\Omega_0}{2}e^{+i(\omega-\omega_{21})t}c_2(t) \quad (1.45)$$

$$\dot{c}_2(t) = i\frac{\Omega_0}{2}e^{-i(\omega-\omega_{21})t}c_1(t), \quad (1.46)$$

which is a good approximation for detuning $\delta = \omega - \omega_{21} \approx 0$. We introduce

$$\tilde{c}_1(t) = c_1(t)e^{-i\frac{\delta}{2}t} \quad (1.47)$$

$$\tilde{c}_2(t) = c_2(t)e^{+i\frac{\delta}{2}t}. \quad (1.48)$$

$$(1.49)$$

Ansatz Wavefunctions for TLA Whole time evolution in state amplitudes

$$|\Psi(t)\rangle = c'_1(t)|1\rangle + c'_2(t)|2\rangle. \quad (1.50)$$

Time evolution when field is off

$$|\Psi(t)\rangle = c'_1(0)e^{-i\omega_1 t}|1\rangle + c'_2(0)e^{-i\omega_2 t}|2\rangle. \quad (1.51)$$

However, this is boring. We chose different ansatz as

$$|\Psi(t)\rangle = c_1(t)e^{-i\omega_1 t}|1\rangle + c_2(t)e^{-i\omega_2 t}|2\rangle \quad (1.52)$$

$$\iff |\Psi(t)\rangle = c_1(t)|1\rangle + c_2(t)e^{-i\omega_{21} t}|2\rangle, \quad (1.53)$$

where $c_1(t)$ and $c_2(t)$ capture time evolution on top of eigenstate evolution! We now have

$$|\Psi(t)\rangle = c_1(t)|1\rangle + c_2(t)e^{-i\omega_{21} t}|2\rangle, \quad (1.54)$$

which is called the rotating frame of atom. We also have Rotating frame of light field as

$$|\Psi(t)\rangle = \tilde{c}_1(t)|1\rangle + \tilde{c}_2(t)e^{-i\omega t}|2\rangle, \quad (1.55)$$

where ω is the light frequency, \tilde{c}_1 and \tilde{c}_2 describe time evolution on top of fast light field oscillation.

Solving the TLA Dynamics We have the following equations:

$$\frac{d}{dt} \begin{pmatrix} \tilde{c}_1(t) \\ \tilde{c}_2(t) \end{pmatrix} = \frac{i}{2} \begin{pmatrix} -\delta & \Omega_0 \\ \Omega_0 & +\delta \end{pmatrix} \begin{pmatrix} \tilde{c}_1(t) \\ \tilde{c}_2(t) \end{pmatrix}. \quad (1.56)$$

Considering the simplest case $\delta = 0$

$$\frac{d}{dt}\tilde{c}_1(t) = \frac{i}{2}\Omega_0\tilde{c}_2(t) \quad (1.57)$$

$$\frac{d}{dt}\tilde{c}_2(t) = \frac{i}{2}\Omega_0\tilde{c}_1(t). \quad (1.58)$$

Take time derivative of the first equation, then we have

$$\ddot{\tilde{c}}_1(t) = -\frac{\Omega_0^2}{4}\tilde{c}_1(t), \quad (1.59)$$

the solutions of which are

$$\tilde{c}_1(t) = \cos(\Omega_0 t/2) \quad (1.60)$$

$$\tilde{c}_2(t) = i \sin(\Omega_0 t/2) \quad (1.61)$$

for $\tilde{c}_1(0) = 1$ and $\tilde{c}_2(0) = 0$. Also we can obtain the excited state probability as

$$P_2(t) = |c_2(t)|^2 \quad (1.62)$$

$$= |\tilde{c}_2(t)|^2. \quad (1.63)$$

Rabi Oscillations (Resonant Case) Nonlinear Response can be seen from the figure.

General Rabi Oscillations (with detuning) Given the figure.

$$|\tilde{c}_2(t)|^2 = \frac{\Omega_0^2}{\Omega} \sin^2\left(\frac{1}{2}\Omega t\right) \quad (1.64)$$

$$= \frac{\Omega_0^2}{2\Omega^2} \{1 - \cos(\Omega t)\}, \quad (1.65)$$

where $\Omega = \sqrt{\Omega_0^2 + \delta^2}$ is the effective Rabi frequency.

Interesting Special Cases a) Pi-Puls $\Omega_0\tau = \pi$: swap population

$$|1\rangle \rightarrow i|2\rangle \quad (1.66)$$

$$|2\rangle \rightarrow i|1\rangle. \quad (1.67)$$

b) 2Pi-Puls $\Omega_0\tau = 2\pi$: flip the sign

c) Pi/2-Puls $\Omega_0\tau = \pi/2$: superposition state

1.5 Oscillating Dipoles

Atomic Eigenstates

$$|\Psi_{nlm}(t)\rangle = e^{-iE_{nlm}t/\hbar}|\Psi_{nlm}(0)\rangle, \quad (1.68)$$

$$\hat{H}_0|\Psi_{nlm}(0)\rangle = E_{nlm}|\Psi_{nlm}\rangle, \quad (1.69)$$

and the electron density is

$$\rho(r, \theta, \phi) = |\Psi(r, \theta, \phi, t = 0)|^2. \quad (1.70)$$

Atomic Dipole Calculate (Oscillating) Dipole Moment for Atomic Eigenstate. We denote $|1\rangle = |\Psi_{nlm}\rangle$. We have

$$d(t) = \langle 1(t) | \hat{d} | 1(t) \rangle \quad (1.71)$$

$$= \langle \hat{d} | 1 \rangle \quad (1.72)$$

$$= -e \langle 1 | \hat{r} | 1 \rangle. \quad (1.73)$$

Then,

$$-e \langle 1 | \hat{r} | 1 \rangle = -e \langle 1 | \hat{P} \hat{P}^{-1} \hat{r} \hat{P} \hat{P}^{-1} | 1 \rangle \quad (1.74)$$

$$= +e \langle 1 | \hat{r} | 1 \rangle, \quad (1.75)$$

which implies

$$\langle 1 | \hat{r} | 1 \rangle = 0. \quad (1.76)$$

Atomic Dipole - Superposition States Calculate (Oscillating) Dipole Moment for Atomic Superposition State

$$|\Psi(0)\rangle = \frac{1}{\sqrt{2}}(|1\rangle + i|2\rangle). \quad (1.77)$$

Evolution

$$|\Psi(t)\rangle = \frac{1}{\sqrt{2}}(|1\rangle + ie^{-i\omega_{21}t}|2\rangle). \quad (1.78)$$

We have

$$d(t) = \langle \Psi(t) | \hat{d} | \Psi(t) \rangle \quad (1.79)$$

$$= \frac{1}{2} \left\{ \langle 1 | \hat{d} | 1 \rangle + \langle 2 | \hat{d} | 2 \rangle + ie^{-i\omega_{21}t} \langle 1 | \hat{d} | 2 \rangle - ie^{-i\omega_{21}t} \langle 2 | \hat{d} | 1 \rangle \right\} \quad (1.80)$$

$$= d_{12} i \frac{1}{2} \{ e^{-i\omega_{21}t} - e^{i\omega_{21}t} \} \quad (1.81)$$

$$= d_{12} \sin(\omega_{21}t), \quad (1.82)$$

where d_{12} is the dipole moment amplitude, ω_{21} is the natural resonance frequency.

Electron Density - Superposition States Calculate Electron Probability Density for Superposition State. The superposition state is

$$\Psi(r, t) = \frac{1}{\sqrt{2}} (\Psi_1(r) + ie^{-i\omega_{21}t} \Psi_2(r)). \quad (1.83)$$

The Electron Probability Density is

$$\rho(r, t) = |\Psi(r, t)|^2 \quad (1.84)$$

$$= \Psi^* \Psi \quad (1.85)$$

$$= \frac{1}{2} \{ |\Psi_1(r)|^2 + |\Psi_2(r)|^2 + 2\text{Re} (ie^{-i\omega_{21}t} \Psi_1^*(r) \Psi_2(r)) \}, \quad (1.86)$$

where $2\text{Re} (ie^{-i\omega_{21}t} \Psi_1^*(r) \Psi_2(r))$ is the interference term.

Examples This is shown by animation and figure in the video.

1.6 The Bloch Sphere

General Two-Level State

- General State Description

$$|\Psi\rangle = c'_1|1\rangle + c'_2|2\rangle \quad (1.87)$$

$$\text{Up to a global phase} \quad (1.88)$$

$$= |c'_1||1\rangle + e^{i\phi}|c'_2|2\rangle \quad (1.89)$$

satisfying $|c'_1|^2 + |c'_2|^2 = 1$.

- Alternative way

$$|\Psi\rangle = \cos(\theta/2)|1\rangle + e^{i\phi}\sin(\theta/2)|2\rangle, \quad (1.90)$$

since $\cos^2(\theta/2) + \sin^2(\theta/2) = 1$.

Geometric Description - Bloch Sphere We then have

$$|\Psi\rangle = \cos(\theta/2)|1\rangle + e^{i\phi}\sin(\theta/2)|2\rangle \quad (1.91)$$

with $0 \leq \theta \leq \pi$ as the latitude and $0 \leq \phi \leq 2\pi$ as the longitude. This is the Bloch Sphere representation. The definition of θ and ϕ and their ranges are different from my familiar coordinate system.

Special States on Bloch Sphere

Analogy to Spin -1/2 States Is shown in the figure.

1.7 Density Operator and Density Matrix

The Problem How do we describe "imperfect state preparation" in an experiment? For example, 50% $|1\rangle$ and 50% $|2\rangle$. We may think of

$$|\Psi\rangle = \frac{1}{\sqrt{2}}(|1\rangle + |2\rangle) \text{ .???} \quad (1.92)$$

This is 100% $|\Psi\rangle$ pure state. We need stable relative phase between the two states!

Optical Analogy - Controlled Phase The double slit problem is shown in the video.

Intensity on Detection Screen:

$$I \propto |E|^2 = |E_1 + e^{i\phi} E_2|^2 \quad (1.93)$$

$$= |E_1|^2 + |E_2|^2 + 2\text{Re} \left(E_1 E_2 e^{i\phi} \right). \quad (1.94)$$

As ϕ varies, Interference pattern "washed out"!

We need new formalism to describe mixed states!(imperfect state preparation, spontaneous emission,...)

Density Operator and Matrix The description of mixed states can be handled by the density operator (matrix) formalism!

- Density operator (hermitian)

$$\hat{\rho} = \sum p_i |\Psi_i\rangle \langle \Psi_i| \quad (1.95)$$

$$\hat{\rho} = I \hat{\rho} I \quad (1.96)$$

$$= \sum_{i,j} |i\rangle \langle i| \hat{\rho} |j\rangle \langle j| \quad (1.97)$$

$$= \rho_{11}|1\rangle\langle 1| + \rho_{12}|1\rangle\langle 2| + \rho_{21}|2\rangle\langle 1| + \rho_{22}|2\rangle\langle 2|, \quad (1.98)$$

where $I = \sum_i |i\rangle \langle i|$.

- Density matrix

$$\rho = \begin{bmatrix} \rho_{11} & \rho_{12} \\ \rho_{21} & \rho_{22} \end{bmatrix}, \quad (1.99)$$

where ρ_{11} and ρ_{22} are the populations, ρ_{12} and ρ_{21} are the coherence. Since ρ is hermitian, we have

$$\rho_{12} = \rho_{21}^*. \quad (1.100)$$

Example 1.7.1 (Example: Density Matrix of Pure State). *We have*

$$|\Psi\rangle = |c_1||1\rangle + e^{i\phi}|c_2||2\rangle. \quad (1.101)$$

*The corresponding density operator of the **pure state** is $\hat{\rho} = |\Psi\rangle \langle \Psi|$. Then the corresponding density matrix is*

$$\rho = \begin{bmatrix} |c_1|^2 & |c_1||c_2|e^{-i\phi} \\ |c_1||c_2|e^{i\phi} & |c_2|^2 \end{bmatrix}, \quad (1.102)$$

where $|c_1||c_2|e^{-i\phi}$ and $|c_1||c_2|e^{i\phi}$ are relative phase between states $|1\rangle$ and $|2\rangle$.

specific example:

$$|\Psi\rangle = \frac{1}{\sqrt{2}} (|1\rangle + |2\rangle), \quad (1.103)$$

so

$$\rho = \begin{bmatrix} 1/2 & 1/2 \\ 1/2 & 1/2 \end{bmatrix}. \quad (1.104)$$

Example 1.7.2 (Example: Fully Incoherent Mixture).

$$\hat{\rho} = \frac{1}{2}|1\rangle\langle 1| + \frac{1}{2}|2\rangle\langle 2| \quad (1.105)$$

with

$$\rho = \begin{bmatrix} 1/2 & 0 \\ 0 & 1/2 \end{bmatrix}, \quad (1.106)$$

where vanishingly coherence and the phase varies from 0 to 2π . It means that we did not control phase.

Useful Facts

- Expectation values: $\langle \hat{A} \rangle = \text{Tr}(\hat{\rho}\hat{A}) = \text{Tr}(\rho A)$
- Time evolution (von Neumann equation)

$$i\hbar \frac{\partial \hat{\rho}}{\partial t} = [\hat{H}, \hat{\rho}] \quad (1.107)$$

- Pure state: $\text{Tr}(\rho^2) = 1$
- Mixed states: $\text{Tr}(\rho^2) < 1$

1.8 Optical Bloch Equations

Time Evolution of Density Matrix How to calculate time evolution of density matrix?

$$i\hbar \frac{\partial \hat{\rho}}{\partial t} = [\hat{H}, \hat{\rho}]. \quad (1.108)$$

Assume pure state

$$\frac{d}{dt}\rho_{11} = \frac{d}{dt}(c_1 c_1^*) \quad (1.109)$$

$$= \dot{c}_1 c_1^* + c_1 \dot{c}_1^* \quad (1.110)$$

$$= i\frac{\Omega_0}{2} \left(e^{i\delta t} \rho_{21} - e^{-i\delta t} \rho_{12} \right) \quad (1.111)$$

Transformation to rotality frame of light

$$= i\frac{\Omega_0}{2} (\tilde{\rho}_{21} - \tilde{\rho}_{12}), \quad (1.112)$$

where

$$\dot{c}_1(t) = i\frac{\Omega_0}{2} e^{+i\delta t} c_2(t) \quad (1.113)$$

$$\dot{c}_2(t) = i\frac{\Omega_0}{2} e^{-i\delta t} c_1(t) \quad (1.114)$$

$$\tilde{\rho}_{12} = e^{-i\delta t} \rho_{12} \quad (1.115)$$

$$\tilde{\rho}_{21} = e^{+i\delta t} \rho_{21}. \quad (1.116)$$

Other elements obtained in analogy!

$$\frac{d}{dt}\rho_{11} = i\frac{\Omega_0}{2}(\tilde{\rho}_{21} - \tilde{\rho}_{12}) \quad (1.117)$$

$$\frac{d}{dt}\rho_{22} = i\frac{\Omega_0}{2}(\tilde{\rho}_{12} - \tilde{\rho}_{21}) \quad (1.118)$$

$$\frac{d}{dt}\tilde{\rho}_{12} = -i\delta\tilde{\rho}_{12} + i\frac{\Omega_0}{2}(\rho_{22} - \rho_{11}) \quad (1.119)$$

$$\frac{d}{dt}\tilde{\rho}_{21} = +i\delta\tilde{\rho}_{21} + i\frac{\Omega_0}{2}(\rho_{11} - \rho_{22}). \quad (1.120)$$

Noting that $\tilde{\rho}_{12} = \tilde{\rho}_{21}$ due to hermitian matrix, the third and the forth equations are the same. So we have

$$\frac{d}{dt}\rho_{11} = i\frac{\Omega_0}{2}(\tilde{\rho}_{21} - \tilde{\rho}_{12}) \quad (1.121)$$

$$\frac{d}{dt}\rho_{22} = i\frac{\Omega_0}{2}(\tilde{\rho}_{12} - \tilde{\rho}_{21}) \quad (1.122)$$

$$\frac{d}{dt}\tilde{\rho}_{12} = -i\delta\tilde{\rho}_{12} + i\frac{\Omega_0}{2}(\rho_{22} - \rho_{11}). \quad (1.123)$$

Optical Bloch Equations with Damping Phenomenological damping and spontaneous emission in the figure. Combine the decay, we have

$$\frac{d}{dt}\rho_{11} = i\frac{\Omega_0}{2}(\tilde{\rho}_{21} - \tilde{\rho}_{12}) + \gamma\rho_{22} \quad (1.124)$$

$$\frac{d}{dt}\rho_{22} = i\frac{\Omega_0}{2}(\tilde{\rho}_{12} - \tilde{\rho}_{21}) - \gamma\rho_{22} \quad (1.125)$$

$$\frac{d}{dt}\tilde{\rho}_{12} = -i\delta\tilde{\rho}_{12} + i\frac{\Omega_0}{2}(\rho_{22} - \rho_{11}) - (\gamma/2)\tilde{\rho}_{12}. \quad (1.126)$$

We now define the inversion $w = \rho_{22} - \rho_{11}$. We have Optical Bloch Equations with Damping

$$\frac{d}{dt}\tilde{\rho}_{21} = -(\gamma/2 - i\delta)\tilde{\rho}_{21} - \frac{i\omega\Omega_0}{2} \quad (1.127)$$

$$\frac{d}{dt}\omega = -\gamma(\omega + 1) - i\Omega_0(\tilde{\rho}_{21} - \tilde{\rho}_{12}) \quad (1.128)$$

in the Density Matrix Form.

1.9 Optical Bloch Equations - Dynamics and Steady State

Dynamical Evolution of System Shown in the figure in the picture.

Steady State Solution Conditions: $\frac{d}{dt}\tilde{\rho}_{21} = 0$ and $\frac{d}{dt}\omega = 0$. Then we have the solutions

$$\omega = -\frac{1}{1+S} \quad (1.129)$$

$$\tilde{\rho}_{21} = \frac{2\Omega_0}{2(\gamma/2 - \delta)(1+S)} \quad (1.130)$$

$$S = \frac{\Omega_0^2/2}{\delta^2 + \gamma^2/4} = \frac{S_0}{1 + 4\delta^2/\gamma^2} \quad (1.131)$$

$$S_0 = \frac{2\Omega_0^2}{\gamma^2} = \frac{I}{O_{sat}}, \quad (1.132)$$

where S is called the saturation parameter, S_0 is called resonant saturation parameter.

Limiting Cases:

- $S \leq 1$: $w \rightarrow -1$ where $w = \rho_{22} - \rho_{11}$. Atom is mainly in ground state.
- $S \geq 1$: $S \rightarrow \infty$, $w \rightarrow 0$.
- Excited State Population:

$$\rho_{22} \quad (1.133)$$

Combine with $\rho_{22} + \rho_{11} = 1$

$$= \frac{1}{2}(1+w) \quad (1.134)$$

$$= \frac{S}{2(1+S)} \quad (1.135)$$

$$= \frac{S_0/2}{1 + S_0 + 4\delta^2/\gamma^2} \quad (1.136)$$

$$\xrightarrow{S_0 \rightarrow \infty, \delta=0} \frac{1}{2}. \quad (1.137)$$

- Photon Scattering Rate: $\Gamma_{ph} = \gamma\rho_{22} = \frac{\gamma}{2} \frac{S_0}{1+S_0+4\delta^2/\gamma^2}$. $\Gamma_{ph} \rightarrow \gamma/2$ for $S_0 \rightarrow \infty$ and $\delta = 0$. We rewrite it as

$$\Gamma_{ph} = \left(\frac{S_0}{1+S_0} \right) \left(\frac{\gamma/2}{1+4\delta^2/\gamma'^2} \right) \quad (1.138)$$

$$\gamma' = \gamma\sqrt{1+S_0}. \quad (1.139)$$

It has a figure in the video. The saturation broadening is shown in the figure.

1.10 Lambert-Beer Law

Attenuation of Light It is shown in the figure.

Scattered Light from Slab of Atoms scattered light power by slab of length dz

$$dP_{sc} = \Gamma_{ph} \times nAdz \times \hbar\omega, \quad (1.140)$$

where Γ_{ph} is the single atom photon scattering rate, $\hbar\omega$ is the energy of single atom, $nAdz$ is the number of atoms. Then we have

$$\frac{dP_{sc}}{dz} = \Gamma_{ph} \times nA \times \hbar\omega. \quad (1.141)$$

Scattered Light from Slab of Atoms Energy conservation requires

$$\frac{dP}{dz} = -\frac{dP_{sc}}{dz} \quad (1.142)$$

$$\frac{dP}{dz} = \frac{dI}{dI} A. \quad (1.143)$$

Put every thing together:

$$\frac{dI}{dz} = -\Gamma n \hbar\omega. \quad (1.144)$$

We have

$$\frac{dI(z)}{dz} = -n\sigma I(z), \quad (1.145)$$

where σ is the atomic scattering cross section.

Lambert-Beer Law (no saturation) We compute the solutions

$$I(z) = I(0)e^{-n\sigma z}, \quad (1.146)$$

which is the Lambert-Beer Law of Absorption.

Laser induced Fluorescence Shown in a video.

1.11 Bloch Vector

Density Matrix Revisited Density Matrix of TLA

$$\rho = \begin{pmatrix} \rho_{11} & \rho_{12} \\ \rho_{21} & \rho_{22} \end{pmatrix} \quad (1.147)$$

Density Matrix hermitian

$$\rho = \rho^\dagger = (\rho^T)^*, \quad (1.148)$$

so we have

$$\rho = \begin{pmatrix} \rho_{11} & \text{Re}\rho_{12} + i\text{Im}\rho_{12} \\ \text{Re}\rho_{12} - i\text{Im}\rho_{12} & \rho_{22} \end{pmatrix}. \quad (1.149)$$

Pauli matrices are

$$\sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \quad (1.150)$$

The decomposition of Density matrix into Pauli matrices

$$\rho = \frac{1}{2} (I + b_x \sigma_x + b_y \sigma_y + b_z \sigma_z), \quad (1.151)$$

where $b_x, b_y, b_z \in \mathbb{R}$.

Bloch Vector We have the density matrix in rotating frame of light

$$\tilde{\rho} = \begin{pmatrix} \rho_{11} & \tilde{\rho}_{12} \\ \tilde{\rho}_{21} & \rho_{22} \end{pmatrix}, \quad (1.152)$$

where $\tilde{\rho}_{12} = \rho_{12} e^{-i\omega t}$. We use following sign convention and have

$$\tilde{\rho} = \frac{1}{2} (I + u \sigma_x - v \sigma_y - w \sigma_z), \quad (1.153)$$

and the bloch vector is defined as

$$\begin{pmatrix} u \\ v \\ w \end{pmatrix}. \quad (1.154)$$

It can be easily shown that

$$u = 2\text{Re}(\tilde{\rho}_{12}) = \tilde{\rho}_{12} + \tilde{\rho}_{12}^* \quad (1.155)$$

$$v = 2\text{Im}(\tilde{\rho}_{12}) = i(\tilde{\rho}_{12}^* - \tilde{\rho}_{12}) \quad (1.156)$$

$$w = \rho_{22} - \rho_{11}, \quad (1.157)$$

$$(1.158)$$

where u is the dispersive component, v is the absorption component and w is the inversion.

Bloch vector can be used to describe any state of TLA density matrix!

Properties of Bloch Vector

- Mixed State: $u^2 + v^2 + w^2 < 1$
- Pure State: $u^2 + v^2 + w^2 = 1$

1.12 Understanding Bloch Vector

What physical behaviour do the components stand for?

- $w = -1$ atom in ground state. $w = +1$ atom in excited state.

- What about u, v ?

$$\langle \hat{d}_i(t) \rangle = \text{Tr}(\hat{\rho} \hat{d}) \quad (1.159)$$

$$= \text{Tr} \left[\begin{pmatrix} \rho_{11} & \rho_{12} \\ \rho_{12}^* & \rho_{22} \end{pmatrix} \begin{pmatrix} 0 & d_{12}^i \\ d_{12}^i & 0 \end{pmatrix} \right], \quad (1.160)$$

where $d_{12}^x = \langle 1 | -q\hat{x} | 2 \rangle$.

Written in the vector form, we have

$$\langle \hat{d} \rangle(t) = d_{12} (\rho_{12} + \rho_{12}^*) \quad (1.161)$$

$$= d_{12} (\tilde{\rho}_{12} e^{i\omega t} + \tilde{\rho}_{12}^* e^{-i\omega t}) \quad (1.162)$$

$$= d_{12} [u \cos(\omega t) - v \sin(\omega t)], \quad (1.163)$$

where we use $\rho_{12} = \tilde{\rho}_{12} e^{i\omega t}$, u denotes in phase and v denotes 90° out of phase component.

Reminder: $E(t) = \epsilon E_0 \cos(\omega t)$.

- Which component responsible for absorption/emission? We have a figure in the video to show the classical picture.

Average absorbed power per atom (classical ensemble average)

$$\left\langle \frac{dW}{dt} \right\rangle = \epsilon E_0 \cos(\omega t) \left\langle -q \frac{dr}{dt} \right\rangle \quad (1.164)$$

$$= \epsilon E_0 \cos(\omega t) \langle \dot{d} \rangle. \quad (1.165)$$

Quantum mechanical analogue (Ehrenfest)

$$\left\langle \frac{dW}{dt} \right\rangle = \epsilon E_0 \cos(\omega t) \langle \dot{d} \rangle \quad (1.166)$$

$$\langle \dot{d} \rangle(t) = d_{12} [u \cos(\omega t) - v \sin(\omega t)]. \quad (1.167)$$

$$\left\langle \frac{dW}{dt} \right\rangle = -d_{12} \cdot \epsilon E_0 \omega (u \cos(\omega t) \sin(\omega t) + v \sin(\omega t)^2) \quad (1.168)$$

$$\overline{\left\langle \frac{dW}{dt} \right\rangle} = \frac{1}{T} \int dt \left\langle \frac{dW}{dt} \right\rangle \quad (1.169)$$

$$= -\frac{d_{12} \cdot \epsilon E_0 \omega v}{2} \quad (1.170)$$

$$= -\hbar \frac{d_{12} \epsilon E_0}{\hbar} \omega \frac{v}{2} \quad (1.171)$$

$$= -\hbar \Omega_0 \omega \frac{v}{2}, \quad (1.172)$$

which is the absorption.

1.13 Optical Bloch Equations using Bloch Vector

1.14 Interlude: The Mach-Zehnder Interferometer

1.15 Ramsey Interferometer

1.16 Quantization of the e.m. field

Fundamental Idea RadiationMode (k, α)

- **To every radiation mode, we associate a harmonic oscillator!** Creation and annihilation operators can change the degree of excitation of mode (occupation with photons)
- **A photon is an excitation quantum of the harmonic oscillator associated with a mode!**

Creation and Annihilation Operators $\hat{a}_k |n_k\rangle = \sqrt{n_k} |n_k - 1\rangle$: decrease photon number by one photon.

$\hat{a}_k^\dagger |n_k\rangle = \sqrt{n_k + 1} |n_k + 1\rangle$: increase photon number by one photon.

Number operator: $\hat{n}_k |n_k\rangle = n_k |n_k\rangle$.

Fock state: $|n_k\rangle$.