OODF: Optimized Opacity Distribution Functions for a New Generation of Solar and Stellar Brightness Variability Models

Miha Cernetic, Alexander I. Shapiro, Veronika Witzke, Natalia A. Krivova, Sami K. Solanki, Rinat V. Tagoirov

Aug. 23, '18











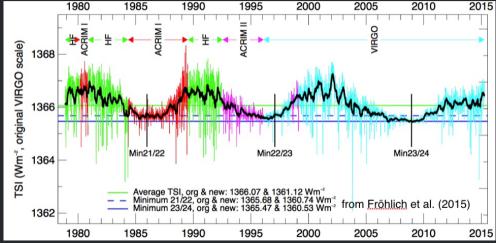




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Total Solar Irradiance

• TSI – spectrally integrated solar radiative flux at 1 AU from the sun

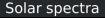


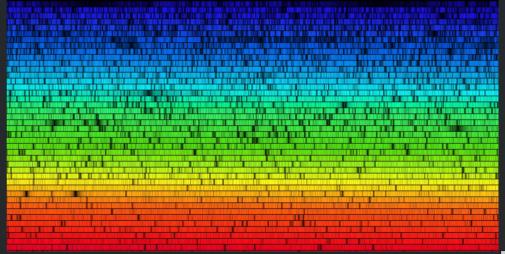


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Spectra of the individual components

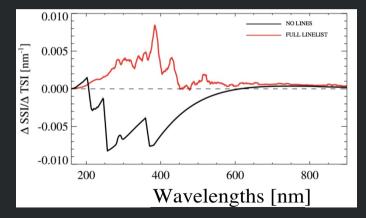






Importance of lines for variability

- TSI Total Solar Irradiance, i.e. integrated over wavelengths
- SSI Spectral Solar Irradiance, depends on wavelength
- \triangle difference between the solar minima and maxima

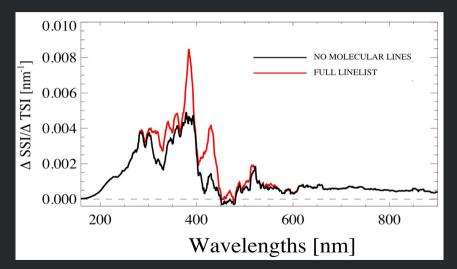






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Importance of lines for variability



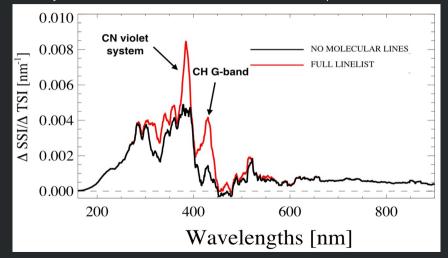




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Importance of lines for variability

25% of the variability comes from molecular lines → accurate linelists are required



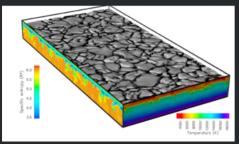




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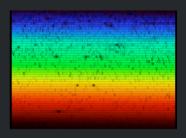
1.5D simulations

3D MHD simulations with MURaM



Structure of the magnetic features

1.5D radiative transfer



Spectra of the magnetic features

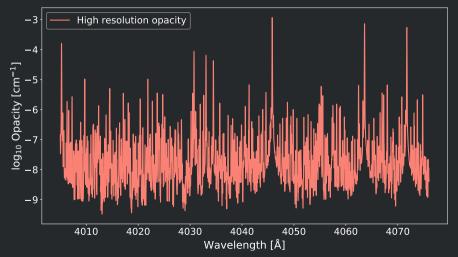




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Generating ODFs

Start with high resolution opacity

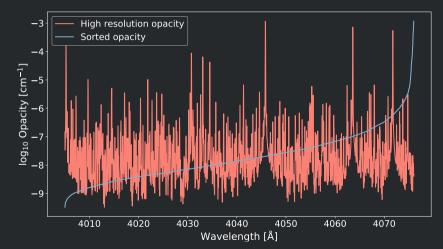






Generating ODFs

- Sort wavelength points by corresponding values of opacity; monotonically increasing opacity
- Integral is preserved by sorting



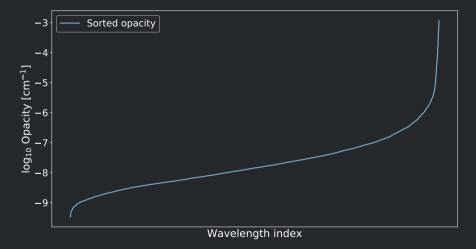




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Generating ODFs

• All wavelength information within the bin is lost



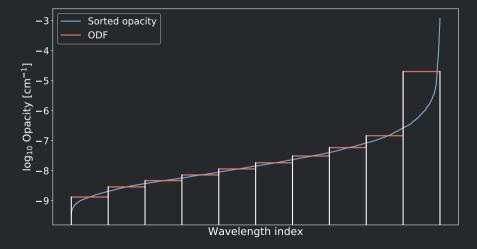




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Generating ODFs - Example with 10 uniform sub bins

• Approximate the sorted opacity with a step-wise function



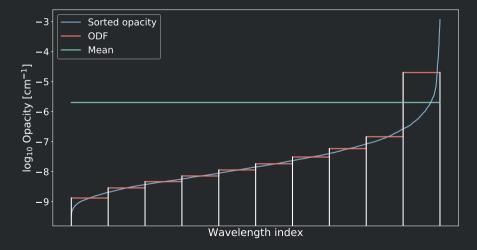




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ODF generation process

Mean is skewed by extreme values



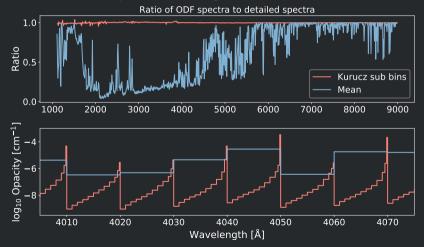




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ODF performance analysis

- Synthesize spectra using ODFs from 1000-9000Å in 10Å bins
- Compare the fluxes from the ODF spectrum with the high resolution spectrum in the bins



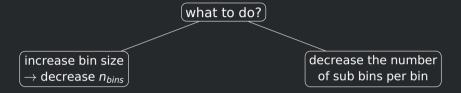




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Possible solutions

Number of calculations goes as: $n_{bins} \times n_{subbins}$



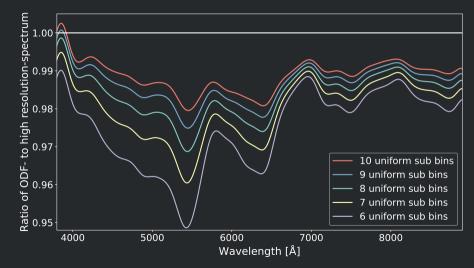




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Analysis of different ODFs

Uniform ODFs



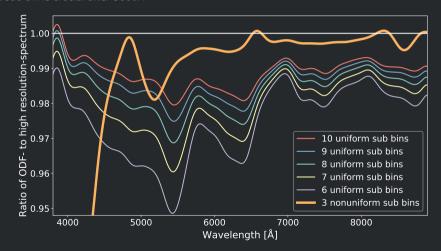




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Analysis of different ODFs

- Nonuniform ODFs
- The last sub bin is crucial after 5000Å

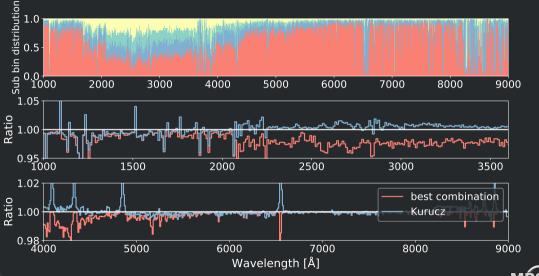






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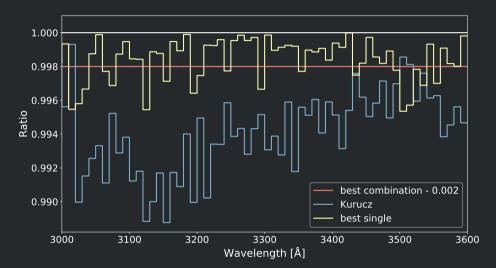
Best sub bin combinations using 4 sub bins





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Best sub bin combinations using 10 sub bins



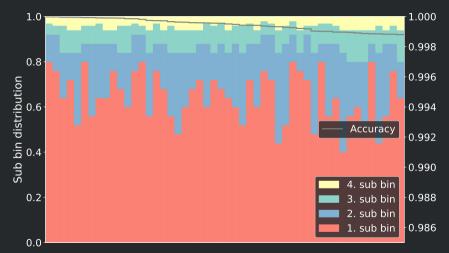




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Best combination of 4 sub bins for Strömgren b

• Total line contribution \sim 15%







Speedups in the case of Strömgren b

ullet Interval length: \sim 400Å

High resolution: 80 points per Å \sim 32 000 points

ODF: 12 points per 10Å \sim 480 points speedup 67 times

OODF: 3 points for the whole binÅ \sim 3 points speedup \sim 11 000 times





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Conclusions

- We developed a novel method for fast spectral synthesis.
- Found optimal sub bins for different wavelength regimes.
- Can be tailored for different filters: Strömgren b + y, Kepler, PLATO and others.
- Significant speed up relative to standard methods by a factor of at least two orders of magnitude.





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Thank you for your attention!

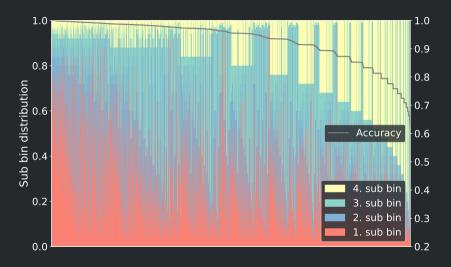






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Best combinations of 4 sub bins for Strömgren b

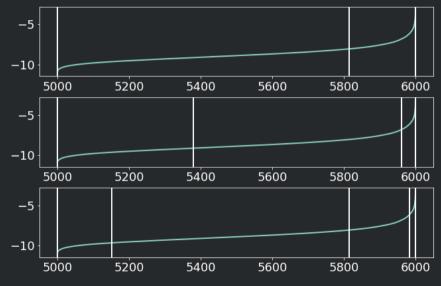






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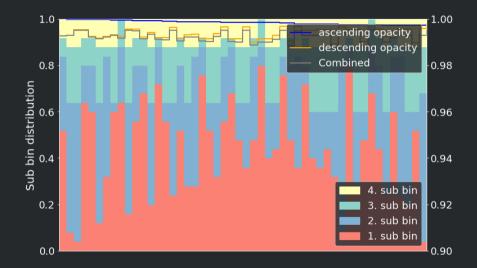
Formula: value weighted by the derivative







Ascending vs descending sort







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