OODF: Optimized Opacity Distribution Functions for a New Generation of Solar and Stellar Brightness Variability Models

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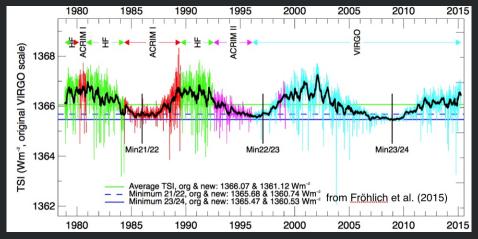






Total Solar Irradiance

• TSI – spectrally integrated solar radiative flux at 1 AU from the sun

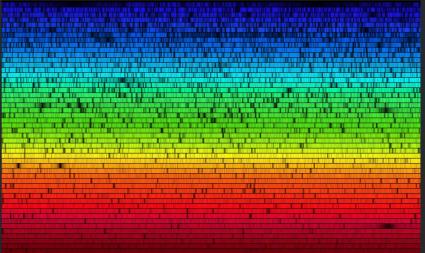






Spectra of the individual components

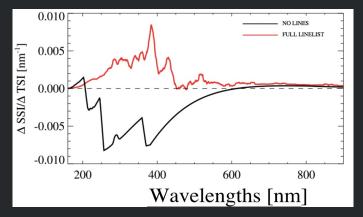
Solar spectra





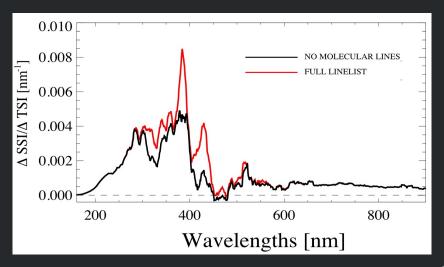


- TSI Total Solar Irradiance, i.e. integrated over wavelengths
- SSI Spectral Solar Irradiance, depends on wavelength
- \triangle difference between the solar minima and maxima



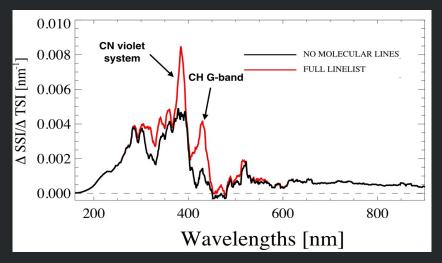








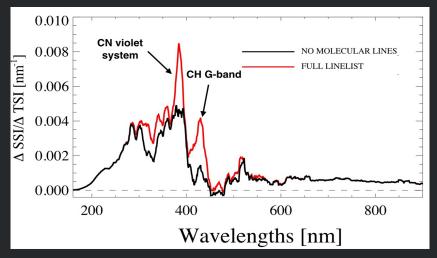








ullet 25% of the variability comes from molecular lines o accurate linelists are required

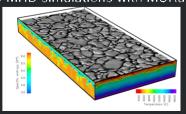






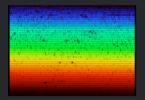
1.5D simulations

3D MHD simulations with MURaM



Structure of the magnetic features

1.5D radiative transfer



Spectra of the magnetic features

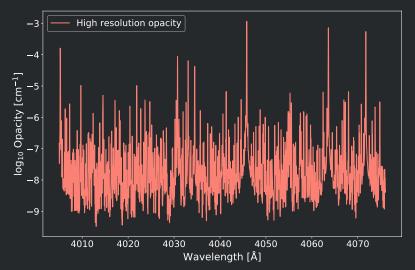




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Generating ODFs

Start with high resolution opacity

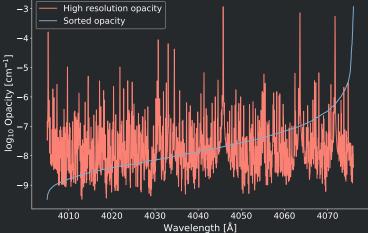






Generating ODFs

- Sort wavelength points by corresponding values of opacity; monotonically increasing opacity
- Integral is preserved by sorting



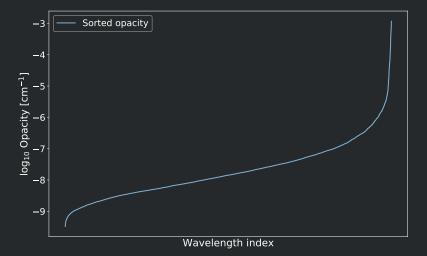




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Generating ODFs

• All wavelength information within the bin is lost







Generating ODFs - Example with 10 uniform sub bins

Approximate the sorted opacity with a step-wise function







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ODF generation process

Mean is skewed by extreme values

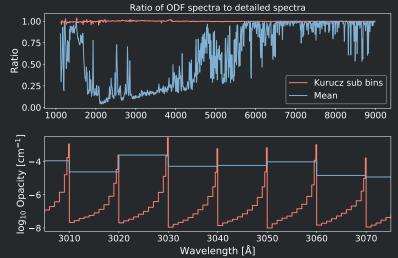






ODF performance analysis

- Synthesize spectrum using ODFs from 1000-9000Å with 10Å bins
- Compare the fluxes from the ODF spectrum with the high resolution spectrum in the bins







Possible solutions

Number of calculations goes as: $n_{bins} \times n_{subbins}$



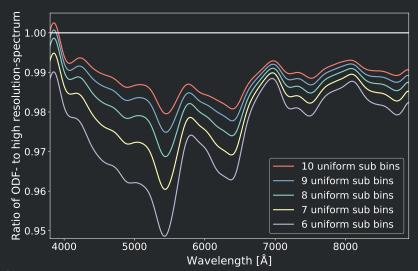




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Analysis of different ODFs

Uniform ODFs



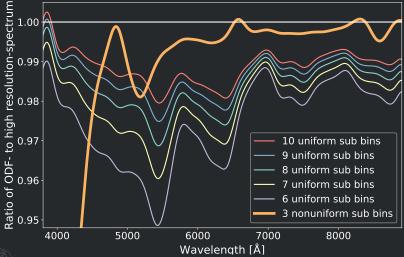




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Analysis of different ODFs

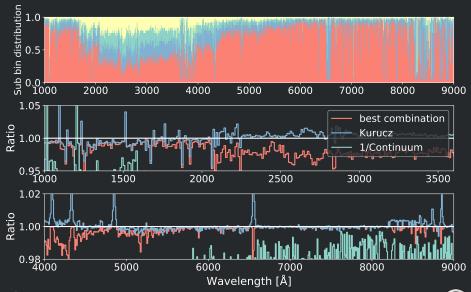
- Nonuniform ODFs
- The last sub bin is crucial after 5000Å





16/23 **WIPS**

Best sub bin combinations using 4 sub bins

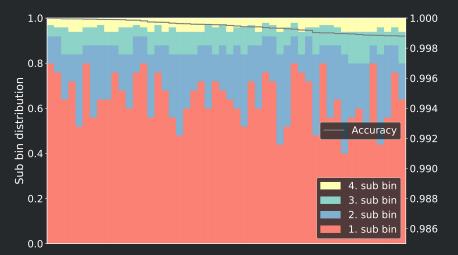






Best combination of 4 sub bins for Strömgren b

ullet Total line contribution \sim 15%







Speedups in the case of Strömgren b

Interval length: ~400Å

High resolution: 80 points per Å \sim 32 000 points

ODF: 12 points per 10Å \sim 480 points speedup 67 times

OODF: 3 points for the whole binÅ \sim 3 points speedup \sim 11 000 times





Conclusions

- We developed a novel method for fast spectral synthesis.
- Found optimal sub bins for different wavelength regimes.
- Can be tailored for different filters: Strömgren b + y, Kepler, PLATO and others.
- Significant speed up relative to standard methods by a factor of at least two orders of magnitude.





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Thank you for your attention!





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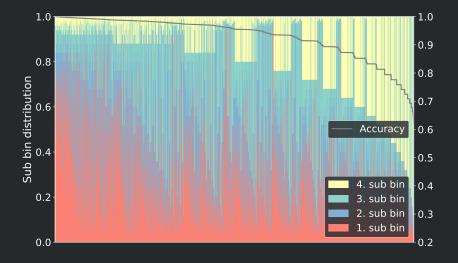






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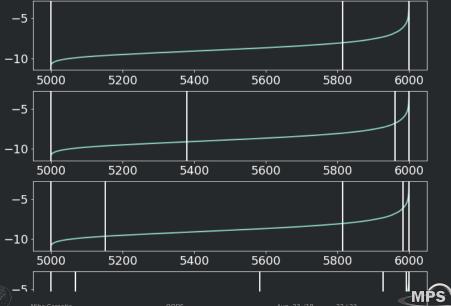
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Formula: value weighted by the derivative



Ascending vs descending sort

