

## Schwarzschild Solution - Static, Spherical Symmetric and Vacuum

static: time independent!

$$ds^2 = -U(r) c^2 dt^2 + V(r) dr^2 +$$

$$W(r) r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \Rightarrow -U(r) c^2 dt^2 + V(r) dr^2 + W(r) r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

$$\blacksquare \text{ let } U(r) \rightarrow e^{2\nu(r)}, V(r) \rightarrow e^{2\lambda(r)}$$

Metric Tensor :

$$g_{\mu\nu} = \begin{pmatrix} -e^{2\nu} & 0 & 0 & 0 \\ 0 & e^{2\lambda} & 0 & 0 \\ 0 & 0 & r^2 & 0 \\ 0 & 0 & 0 & r^2 \sin^2 \theta \end{pmatrix}, \quad g^{\mu\nu} = \begin{pmatrix} -e^{-2\nu} & 0 & 0 & 0 \\ 0 & e^{-2\lambda} & 0 & 0 \\ 0 & 0 & 1/r^2 & 0 \\ 0 & 0 & 0 & 1/r^2 \sin^2 \theta \end{pmatrix}$$

```
In[1]:= dim = 4;
coord = {ct, r, θ, ϕ};
metric = {{-e^{2ν[r]}, 0, 0, 0}, {0, e^{2λ[r]}, 0, 0}, {0, 0, r^2, 0}, {0, 0, 0, r^2 Sin[θ]^2}};
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**metric // MatrixForm**

$$\begin{pmatrix} -e^{2\nu[r]} & 0 & 0 & 0 \\ 0 & e^{2\lambda[r]} & 0 & 0 \\ 0 & 0 & r^2 & 0 \\ 0 & 0 & 0 & r^2 \sin^2[\theta] \end{pmatrix}$$

$$\blacksquare R_{\mu\nu} = R^\rho_{\mu\rho\nu} = \Gamma^\rho_{\mu\nu,\rho} - \Gamma^\rho_{\mu\rho,\nu} + \Gamma^k_{\mu\nu} \Gamma^\rho_{k\rho} - \Gamma^k_{\mu\rho} \Gamma^\rho_{k\nu}$$

Let's get the Christoffel symbol first!

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In[4]:= inversemetric = Simplify[Inverse[metric]];
affine := affine = Simplify[Table[(1/2) * Sum[(inversemetric[[i, s]] *
  (D[metric[[s, j]], coord[[k]]] +
    D[metric[[s, k]], coord[[j]]] - D[metric[[j, k]], coord[[s]]]),
  {s, 1, dim}],
  {i, 1, dim}, {j, 1, dim}, {k, 1, dim}]];
listaffine :=
  Table[{ToString[i]_j_k, ToString["="], affine[[i, j, k]]},
    {i, 1, dim}, {j, 1, dim}, {k, 1, dim}];
TableForm[listaffine, TableSpacing -> {3, 3}]

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Out[7]/TableForm=

$\Gamma^1_{1,1} = 0$	$\Gamma^1_{2,1} = v'[r]$	$\Gamma^1_{3,1} = 0$	$\Gamma^1_{4,1} = 0$
$\Gamma^1_{1,2} = v'[r]$	$\Gamma^1_{2,2} = 0$	$\Gamma^1_{3,2} = 0$	$\Gamma^1_{4,2} = 0$
$\Gamma^1_{1,3} = 0$	$\Gamma^1_{2,3} = 0$	$\Gamma^1_{3,3} = 0$	$\Gamma^1_{4,3} = 0$
$\Gamma^1_{1,4} = 0$	$\Gamma^1_{2,4} = 0$	$\Gamma^1_{3,4} = 0$	$\Gamma^1_{4,4} = 0$
$\Gamma^2_{1,1} = e^{2v[r]-2\lambda[r]} v'[r]$	$\Gamma^2_{2,1} = 0$	$\Gamma^2_{3,1} = 0$	$\Gamma^2_{4,1} = 0$
$\Gamma^2_{1,2} = 0$	$\Gamma^2_{2,2} = \lambda'[r]$	$\Gamma^2_{3,2} = 0$	$\Gamma^2_{4,2} = 0$
$\Gamma^2_{1,3} = 0$	$\Gamma^2_{2,3} = 0$	$\Gamma^2_{3,3} = -e^{-2\lambda[r]} r$	$\Gamma^2_{4,3} = 0$
$\Gamma^2_{1,4} = 0$	$\Gamma^2_{2,4} = 0$	$\Gamma^2_{3,4} = 0$	$\Gamma^2_{4,4} = -e^{-2\lambda[r]} r \sin[\theta]$
$\Gamma^3_{1,1} = 0$	$\Gamma^3_{2,1} = 0$	$\Gamma^3_{3,1} = 0$	$\Gamma^3_{4,1} = 0$
$\Gamma^3_{1,2} = 0$	$\Gamma^3_{2,2} = 0$	$\Gamma^3_{3,2} = \frac{1}{r}$	$\Gamma^3_{4,2} = 0$
$\Gamma^3_{1,3} = 0$	$\Gamma^3_{2,3} = \frac{1}{r}$	$\Gamma^3_{3,3} = 0$	$\Gamma^3_{4,3} = 0$
$\Gamma^3_{1,4} = 0$	$\Gamma^3_{2,4} = 0$	$\Gamma^3_{3,4} = 0$	$\Gamma^3_{4,4} = -\cos[\theta] \sin[\theta]$
$\Gamma^4_{1,1} = 0$	$\Gamma^4_{2,1} = 0$	$\Gamma^4_{3,1} = 0$	$\Gamma^4_{4,1} = 0$
$\Gamma^4_{1,2} = 0$	$\Gamma^4_{2,2} = 0$	$\Gamma^4_{3,2} = 0$	$\Gamma^4_{4,2} = \frac{1}{r}$
$\Gamma^4_{1,3} = 0$	$\Gamma^4_{2,3} = 0$	$\Gamma^4_{3,3} = 0$	$\Gamma^4_{4,3} = \cot[\theta]$
$\Gamma^4_{1,4} = 0$	$\Gamma^4_{2,4} = \frac{1}{r}$	$\Gamma^4_{3,4} = \cot[\theta]$	$\Gamma^4_{4,4} = 0$

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In[8]:= Riemann := Riemann =
  Simplify[Table[D[affine[[i, j, l]], coord[[k]]] - D[affine[[i, j, k]], coord[[l]]] +
    Sum[affine[[i, s, k]] affine[[s, j, l]] - affine[[i, s, l]] affine[[s, j, k]],
    {s, 1, dim}],
  {i, 1, dim}, {j, 1, dim}, {k, 1, dim}, {l, 1, dim}]];
Ricci := Table[Sum[Riemann[[u, a, u, b]], {u, 1, dim}], {a, 1, dim}, {b, 1, dim}];
RicciR := Sum[Ricci[[a, b]] inversemetric[[a, b]], {a, 1, dim}, {b, 1, dim}]

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In[18]:= For[i = 1, i < 5, i++, 1,
  For[j = 1, j < 5, j++, 1,
    Print[ToString[R[i, j]], ToString["="], Ricci[[i, j]]]]]

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$$R[1, 1] = \frac{2 e^{2v[r] - 2\lambda[r]} v'[r]}{r} + e^{2v[r] - 2\lambda[r]} (v'[r]^2 - v'[r] \lambda'[r] + v''[r])$$

$$R[1, 2] = 0$$

$$R[1, 3] = 0$$

$$R[1, 4] = 0$$

$$R[2, 1] = 0$$

$$R[2, 2] = -v'[r]^2 + \frac{2\lambda'[r]}{r} + v'[r] \lambda'[r] - v''[r]$$

$$R[2, 3] = 0$$

$$R[2, 4] = 0$$

$$R[3, 1] = 0$$

$$R[3, 2] = 0$$

$$R[3, 3] = 1 - e^{-2\lambda[r]} - e^{-2\lambda[r]} r v'[r] + e^{-2\lambda[r]} r \lambda'[r]$$

$$R[3, 4] = 0$$

$$R[4, 1] = 0$$

$$R[4, 2] = 0$$

$$R[4, 3] = 0$$

$$R[4, 4] = (1 - e^{-2\lambda[r]}) \sin[\theta]^2 - e^{-2\lambda[r]} r \sin[\theta]^2 v'[r] + e^{-2\lambda[r]} r \sin[\theta]^2 \lambda'[r]$$

Vacuum Einstein Field Equation:

$$R_{\mu\nu} = 0$$

From  $R_{00} = 0$ ,  $R_{11} = 0$ ,

$$\lambda'(r) + v'(r) = 0$$

From  $R_{22} = 0$ ,

$$e^{2\lambda(r)} = 1 + 2r v'(r) = e^{-2v(r)}$$

$$1 = (e^{2v(r)} + 2r v'(r) e^{2v(r)}) = (r e^{2v(r)})'$$

$$r + C = r e^{2v(r)}, \text{ let } C = -2m$$

$$e^{2v(r)} = 1 - \frac{2m}{r} \text{ Hence. } e^{2\lambda(r)} = \left(1 - \frac{2m}{r}\right)^{-1}$$

$$\text{In}[24]:= \text{metric} = \text{metric} /. \{e^{2v[r]} \rightarrow \left(1 - \frac{2m}{r}\right), e^{2\lambda[r]} \rightarrow \left(1 - \frac{2m}{r}\right)^{-1}\};$$

**metric // MatrixForm**

Out[25]//MatrixForm=

$$\begin{pmatrix} -1 + \frac{2m}{r} & 0 & 0 & 0 \\ 0 & \frac{1}{1 - \frac{2m}{r}} & 0 & 0 \\ 0 & 0 & r^2 & 0 \\ 0 & 0 & 0 & r^2 \sin[\theta]^2 \end{pmatrix}$$

Above metric is called "Schwarzschild metric". ( $m = GM/c^2$ )

Note that  $g_{11}$  becomes singular at  $r \rightarrow 2GM/c^2 = r_s$ : the emergence of Black Hole

**Birkhoff's theorem-** For vacuum equation, any spherical symmetric solution of the field equation is static

Now suppose that metric evolves as a function of time. Then, unlike static case, here comes new

friend, cross-term,  $c dt dr$

$$ds^2 = -P(r, t) c^2 dt^2 + Q(r, t) dr^2 + 2R(r, t) c dt dr + S(r, t) r^2 d\Omega^2$$

let  $f(r, t) (P(r, t) c dt - R(r, t) dr) = c dF(r, t) = \partial_t F dt + \partial_r F dr \rightarrow$  total derivative

Since  $\partial_r \partial_t F = \partial_t \partial_r F$ ,  $\partial_r (f(r, t) P(r, t)) = \partial_t (-f(r, t) R(r, t))$ , namely,

$$-\frac{\partial f}{\partial t} = \frac{1}{R} \left( \frac{\partial f}{\partial r} P + f \frac{\partial P}{\partial r} + f \frac{\partial R}{\partial t} \right)$$

From  $c^2 dF^2 = f^2 (P c dt - R dr)^2 = f^2 (P^2 c^2 dt^2 - 2PR c dt dr + R^2 dr^2)$ ,

$$ds^2 = -\frac{c^2}{f^2 P} dF^2 + \left( \frac{R^2}{f^2 P} + Q \right) dr^2 + r^2 d\Omega^2 = -U(r, t') c^2 dt'^2 + V(r, t) dr^2 + r^2 d\Omega^2, (\text{substitute } t' \text{ for } F)$$

Above form is pretty analogous to that of Schwarzschild metric! - the only difference is time dependency.

$$ds^2 = -U(r, t') c^2 dt'^2 + V(r, t) dr^2 + r^2 d\Omega^2 = -e^{2\nu(r,t)} c^2 dt^2 + e^{2\lambda(r,t)} dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

Metric Tensor

$$g_{\mu\nu} = \begin{pmatrix} -e^{2\nu(r,t)} & 0 & 0 & 0 \\ 0 & e^{2\lambda(r,t)} & 0 & 0 \\ 0 & 0 & r^2 & 0 \\ 0 & 0 & 0 & r^2 \sin^2 \theta \end{pmatrix}, \quad g^{\mu\nu} = \begin{pmatrix} -e^{-2\nu(r,t)} & 0 & 0 & 0 \\ 0 & e^{-2\lambda(r,t)} & 0 & 0 \\ 0 & 0 & 1/r^2 & 0 \\ 0 & 0 & 0 & 1/r^2 \sin^2 \theta \end{pmatrix}$$

In[86]:= **dim = 4;**

**coord = {t, r,  $\theta$ ,  $\phi$ };**

**metric = {{-e^{2\nu[r,t]}, 0, 0, 0}, {0, e^{2\lambda[r,t]}, 0, 0}, {0, 0, r^2, 0}, {0, 0, 0, r^2 Sin[\theta]^2}};**

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In[89]:= inversemetric = Simplify[Inverse[metric]];
affine := affine = Simplify[Table[(1/2) * Sum[(inversemetric[[i, s]] *
  (D[metric[[s, j]], coord[[k]]] +
    D[metric[[s, k]], coord[[j]]] - D[metric[[j, k]], coord[[s]])],
  {s, 1, dim}],
  {i, 1, dim}, {j, 1, dim}, {k, 1, dim}]];
listaffine :=
  Table[{ToString[i-1]_j-1,k-1, ToString["="], affine[[i, j, k]]},
    {i, 1, dim}, {j, 1, dim}, {k, 1, dim}];
TableForm[listaffine, TableSpacing -> {3, 3}]

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Out[92]//TableForm=

$\Gamma_{0,0}^0 = v^{(0,1)}[r, t]$	$\Gamma_{1,0}^0 = v^{(1,0)}[r, t]$	$\Gamma_{2,0}^0 = 0$
$\Gamma_{0,1}^0 = v^{(1,0)}[r, t]$	$\Gamma_{1,1}^0 = e^{-2v[r,t]+2\lambda[r,t]} \lambda^{(0,1)}[r, t]$	$\Gamma_{2,1}^0 = 0$
$\Gamma_{0,2}^0 = 0$	$\Gamma_{1,2}^0 = 0$	$\Gamma_{2,2}^0 = 0$
$\Gamma_{0,3}^0 = 0$	$\Gamma_{1,3}^0 = 0$	$\Gamma_{2,3}^0 = 0$
$\Gamma_{0,0}^1 = e^{2v[r,t]-2\lambda[r,t]} v^{(1,0)}[r, t]$	$\Gamma_{1,0}^1 = \lambda^{(0,1)}[r, t]$	$\Gamma_{2,0}^1 = 0$
$\Gamma_{0,1}^1 = \lambda^{(0,1)}[r, t]$	$\Gamma_{1,1}^1 = \lambda^{(1,0)}[r, t]$	$\Gamma_{2,1}^1 = 0$
$\Gamma_{0,2}^1 = 0$	$\Gamma_{1,2}^1 = 0$	$\Gamma_{2,2}^1 = -e^{-2\lambda[r,t]}$
$\Gamma_{0,3}^1 = 0$	$\Gamma_{1,3}^1 = 0$	$\Gamma_{2,3}^1 = 0$
$\Gamma_{0,0}^2 = 0$	$\Gamma_{1,0}^2 = 0$	$\Gamma_{2,0}^2 = 0$
$\Gamma_{0,1}^2 = 0$	$\Gamma_{1,1}^2 = 0$	$\Gamma_{2,1}^2 = \frac{1}{r}$
$\Gamma_{0,2}^2 = 0$	$\Gamma_{1,2}^2 = \frac{1}{r}$	$\Gamma_{2,2}^2 = 0$
$\Gamma_{0,3}^2 = 0$	$\Gamma_{1,3}^2 = 0$	$\Gamma_{2,3}^2 = 0$
$\Gamma_{0,0}^3 = 0$	$\Gamma_{1,0}^3 = 0$	$\Gamma_{2,0}^3 = 0$
$\Gamma_{0,1}^3 = 0$	$\Gamma_{1,1}^3 = 0$	$\Gamma_{2,1}^3 = 0$
$\Gamma_{0,2}^3 = 0$	$\Gamma_{1,2}^3 = 0$	$\Gamma_{2,2}^3 = 0$
$\Gamma_{0,3}^3 = 0$	$\Gamma_{1,3}^3 = \frac{1}{r}$	$\Gamma_{2,3}^3 = \text{Cot}[\theta]$

```

In[93]:= Riemann := Riemann =
  Simplify[Table[D[affine[[i, j, l]], coord[[k]]] - D[affine[[i, j, k]], coord[[l]]] +
    Sum[affine[[i, s, k]] affine[[s, j, l]] - affine[[i, s, l]] affine[[s, j, k]],
  {s, 1, dim}],
  {i, 1, dim}, {j, 1, dim}, {k, 1, dim}, {l, 1, dim}]];
Ricci := Table[Sum[Riemann[[u, a, u, b]], {u, 1, dim}], {a, 1, dim}, {b, 1, dim}];
RicciR := Sum[Ricci[[a, b]] inversemetric[[a, b]], {a, 1, dim}, {b, 1, dim}]

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In[96]:= For[i = 1, i < 5, i++ 1,
  For[j = 1, j < 5, j++ 1,
    Print[ToString[R[i-1, j-1]], ToString["="], Ricci[[i, j]]]]]

```

$$R[0, 0] = v^{(0,1)}[r, t] \lambda^{(0,1)}[r, t] - \lambda^{(0,1)}[r, t]^2 - \lambda^{(0,2)}[r, t] + \frac{2 e^{2v[r, t] - 2\lambda[r, t]} v^{(1,0)}[r, t]}{r} + e^{2v[r, t] - 2\lambda[r, t]} (v^{(1,0)}[r, t]^2 - v^{(1,0)}[r, t] \lambda^{(1,0)}[r, t] + v^{(2,0)}[r, t])$$

$$R[0, 1] = \frac{2 \lambda^{(0,1)}[r, t]}{r}$$

$$R[0, 2] = 0$$

$$R[0, 3] = 0$$

$$R[1, 0] = \frac{2 \lambda^{(0,1)}[r, t]}{r}$$

$$R[1, 1] = -e^{-2v[r, t] + 2\lambda[r, t]} v^{(0,1)}[r, t] \lambda^{(0,1)}[r, t] + e^{-2v[r, t] + 2\lambda[r, t]} \lambda^{(0,1)}[r, t]^2 + e^{-2v[r, t] + 2\lambda[r, t]} \lambda^{(0,2)}[r, t] - v^{(1,0)}[r, t]^2 + \frac{2 \lambda^{(1,0)}[r, t]}{r} + v^{(1,0)}[r, t] \lambda^{(1,0)}[r, t] - v^{(2,0)}[r, t]$$

$$R[1, 2] = 0$$

$$R[1, 3] = 0$$

$$R[2, 0] = 0$$

$$R[2, 1] = 0$$

$$R[2, 2] = 1 - e^{-2\lambda[r, t]} - e^{-2\lambda[r, t]} r v^{(1,0)}[r, t] + e^{-2\lambda[r, t]} r \lambda^{(1,0)}[r, t]$$

$$R[2, 3] = 0$$

$$R[3, 0] = 0$$

$$R[3, 1] = 0$$

$$R[3, 2] = 0$$

$$R[3, 3] = (1 - e^{-2\lambda[r, t]}) \sin[\theta]^2 - e^{-2\lambda[r, t]} r \sin[\theta]^2 v^{(1,0)}[r, t] + e^{-2\lambda[r, t]} r \sin[\theta]^2 \lambda^{(1,0)}[r, t]$$

Note that  $\lambda^{(0,1)}[r, t]$  means time derivative of  $\lambda[r, t]$

Einstein field equation of vacuum condition :  $R_{\mu\nu} = 0$

$$v^{(0,1)}[r, t] \lambda^{(0,1)}[r, t] - \lambda^{(0,1)}[r, t]^2 - \lambda^{(0,2)}[r, t] + \frac{2 e^{2v[r, t] - 2\lambda[r, t]} v^{(1,0)}[r, t]}{r} + \text{-----(A)}$$

$$e^{2v[r, t] - 2\lambda[r, t]} (v^{(1,0)}[r, t]^2 - v^{(1,0)}[r, t] \lambda^{(1,0)}[r, t] + v^{(2,0)}[r, t]) = 0$$

$$\frac{2 \lambda^{(0,1)}[r, t]}{r} = 0 \text{ -----(B) tells us that } \lambda \text{ is independent of time. } \partial_t \lambda = 0$$

$$\frac{2 \lambda^{(0,1)}[r, t]}{r} = 0 \text{ -----(C), (B)=(C)}$$

$$-e^{-2v[r, t] + 2\lambda[r, t]} v^{(0,1)}[r, t] \lambda^{(0,1)}[r, t] + e^{-2v[r, t] + 2\lambda[r, t]} \lambda^{(0,1)}[r, t]^2 + \text{-----(D)}$$

$$e^{-2v[r, t] + 2\lambda[r, t]} \lambda^{(0,2)}[r, t] - v^{(1,0)}[r, t]^2 + \frac{2 \lambda^{(1,0)}[r, t]}{r} + v^{(1,0)}[r, t] \lambda^{(1,0)}[r, t] - v^{(2,0)}[r, t] = 0$$

$$1 - e^{-2\lambda[r, t]} - e^{-2\lambda[r, t]} r v^{(1,0)}[r, t] + e^{-2\lambda[r, t]} r \lambda^{(1,0)}[r, t] = 0 \text{ -----(E)}$$

$$(1 - e^{-2\lambda[r, t]}) \sin[\theta]^2 - e^{-2\lambda[r, t]} r \sin[\theta]^2 v^{(1,0)}[r, t] + e^{-2\lambda[r, t]} r \sin[\theta]^2 \lambda^{(1,0)}[r, t] = 0 \text{ -----(F) } \sim (E)$$

In other words,

$$v^{(1,0)}[r, t]^2 + 2 v^{(1,0)}[r, t] / r - v^{(1,0)}[r, t] \lambda^{(1,0)}[r] + v^{(2,0)}[r, t] = 0 \text{ -----(A)}$$

$$-v^{(1,0)}[r, t]^2 + 2 \lambda^{(1,0)}[r] / r + v^{(1,0)}[r, t] \lambda^{(1,0)}[r] - v^{(2,0)}[r, t] = 0 \text{ -----(D)}$$

$$1 - e^{-2\lambda[r]} (1 + r v^{(1,0)}[r, t] - r \lambda^{(1,0)}[r]) = 0 \text{ -----(E)}$$

$$(A) + (D) \Rightarrow \partial_r v(r, t) + \partial_r \lambda(r) = 0, \quad v(r, t) + \lambda(r) = \kappa(t)$$

$$(E) \Rightarrow 1 - e^{-2\lambda(r)} (1 - 2r \partial_r \lambda(r)) = 0, \quad e^{-2\lambda(r)} (1 - 2r \partial_r \lambda(r)) = 1$$

$$(r e^{-2\lambda(r)})' = 1, \quad (r e^{-2\lambda(r)}) = r + C \rightarrow e^{-2\lambda(r)} = 1 + C/r \quad (\text{let } C = -2m) = 1 - 2m/r$$

$$2\nu(r, t) = 2\kappa(t) - 2\lambda(r), \quad e^{2\nu} = e^{2\kappa} e^{-2\lambda} = e^{2\kappa} \left(1 - \frac{2m}{r}\right)$$

Then,

$$ds^2 = -e^{2\kappa} \left(1 - \frac{2m}{r}\right) c^2 dt^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2) \text{ with } m = \frac{MG}{c^2}$$

Finally, redefining the time coordinate as  $e^{-\kappa(t)} dt \rightarrow dt'$ ,

$$ds^2 = -\left(1 - \frac{2m}{r}\right) c^2 dt'^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2) \text{ just as 'static' Schwarzschild solution!}$$

Birkhoff's theorem tells us that any spherically symmetric solution of the field equations is necessarily static.

e.g. A star pulsating radially has the same external field as a star at rest (that is to say, a radially pulsating star emits no gravitational radiation.)

## Reference

- Lewis Ryder (2009), *Introduction to General Relativity*, New York: Cambridge University Press.