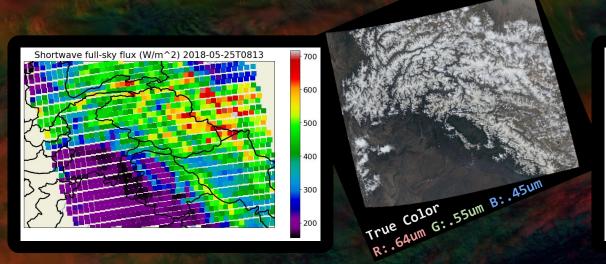
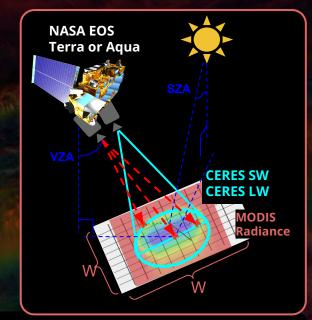
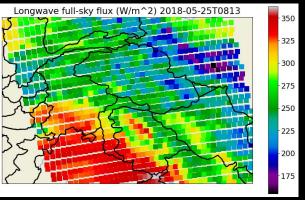
A Deep Learning approach to Estimate CERES Broadband Flux with 1km MODIS Pixels.

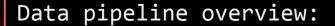
AES 690 - Project Update 2 - Mitchell Dodson

<u>Goal</u>: develop a data-driven deep learning model to characterize the relationship between EOS-MODIS' 36 spectral radiance bands, the sun/pixel/satellite viewing geometry, and the full-sky broadband radiative flux observed by the EOS-CERES instrument over a diverse set of regions.







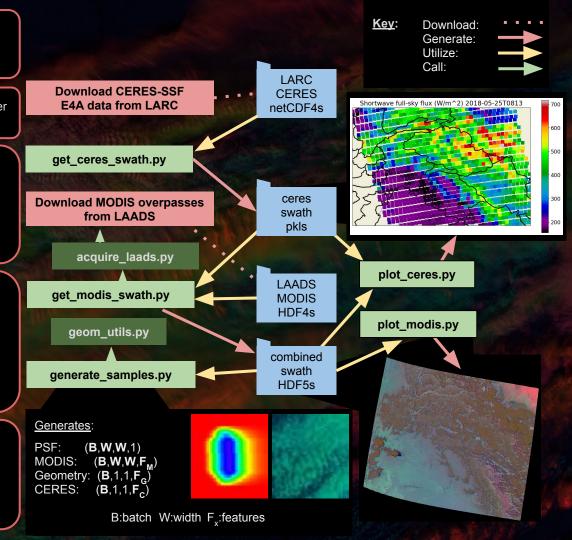


1 Use the LARC downloader to acquire CERES SSF data over a series of desired regions and time periods.

Select fields, restrict valid viewing geometry, and apply other constraints, then call **get_ceres_swath.py**, which extracts, subsets, and preprocesses CERES footprints, then groups them into individual overpasses alongside their meta-info in a new pickle file.

Select a series of MODIS bands to extract, and set file storage parameters in **get_modis_swath.py**, then run the script to download MxD021KM and MxD03 L1b granules from LAADS corresponding to each swath. Extract the desired bands in range of the CERES data, and combine the MODIS and CERES data for each overpass in a custom 'swath' HDF5.

generate_samples.py loads several swath HDF5s concurrently, selects CERES footprints and surrounding MODIS tiles, then calculates the CERES point spread function over each sub-domain. The PSF, MODIS/CERES data, and viewing geometry are interleaved and yielded as tensors



Acquired regional data:

Using the described data pipeline, I acquired every third available EOS Terra and Aqua swath in the 3 years between Jan 1, 2018 and Dec 31, 2020, within the regional boundaries described below, for a total of about 2,600 overpasses, each of which contain about 300 individual samples on average.

Label		Max	Avg	Std
	37.378	46.063	41.750	2.078
swflux	90.989	611.449		120.906
lwflux		285.003		
oct_clr		93.280		28.979
oct_l1		99.300		
oct_l2	0.000	93.300		23.803
l1_cod			12.805	13.646
l2_cod	0.000	72.005		8.280
aer_land_pct	0.071			
aod_land	0.000	0.380	0.037	0.063
aer_ocean_pct	0.006			
aod_ocean	0.000	0.304	0.024	0.045
aod ocean small	0.000	0.180	0.015	0.028

Northeast United States lat=[34,43], lon=[-69,-79] Swath count: 532

(idn) Label			Avg	Std
		4.570	-0.170	2.646
swflux				
lwflux		291.400		
pct_clr				
	0.000			
l1_cod				
	0.000			
aer_land_pct	0.000			
aod_land	0.000			
ser_ocean_pct	0.000			
aod_ocean	0.000			0.056
aod ocean small	0.000	0.199	0.022	0.034

Indonesia lat=[-5,5], lon=[120,130] Swath count: 493

(alk)				
			Avg	
	58.017			0.834
	50.844			
swflux				
		252.280		15.006
pct_clr		76.016		
pct_l1				
pct_l2	0.000			20.830
	0.000			
aer_land_pct	0.008			
aod_land	0.000		0.043	
aer_ocean_pct	0.000			
aod_ocean	0.000		0.008	
aod_ocean_small	0.000		0.006	

Alaska lat=[58,61], lon=[-160,-154] Swath count: 383

Label			Avg	
lat	34.040			
lon				
swflux				79.409
lwflux				20.083
	0.003			
l1_cod				10.614
l2_cod	0.000			
aer_land_pct				
aod_land	0.002		0.082	0.078
aer_ocean_pct	0.000	0.000	0.000	0.000
aod_ocean	0.000	0.000	0.000	0.000
aod_ocean_small	0.000	0.000	0.000	0.000

Southeast United States lat=[34,38], lon=[-90,-85] Swath count: 356

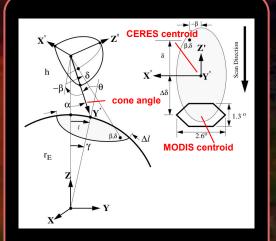
			Avg	
			38.904	
				113.011
	168.895			
pct_clr				
pct_l1				
pct_l2	0.000			
l2_cod	0.000			
aer_land_pct	0.000			
aod_land	0.000			
aer_ocean_pct	0.000	0.000		0.000
aod_ocean	0.000	0.000	0.000	0.000
and ocean small	0.000	0.000	0.000	0.000

Hindu Kush Himalayas lat=[32,38], lon=[69,79] Swath count: 520

			Avg	
				6.608
swflux		693.881	342.504	110.928
lwflux		277.341		
pct_clr		84.800	30.260	
pct_l1	1.094			
pct_l2	0.000	91.002		
l1_cod	0.966	70.073		
12_cod	0.000			
aer_land_pct			16.845	
aod_land	0.002			
aer_ocean_pct	0.000	0.000	0.000	0.000
aod_ocean	0.000	0.000	0.000	0.000
aod_ocean_small	0.000	0.000	0.000	0.000

Amazon Rainforest lat=[-3,-8], lon=[-65,-70] Swath count: 305

Challenge: Point Spread Function



Equatorial geometry of the observation, and MODIS/CERES centroid offset schematic

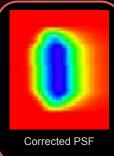
Implementing the CERES point spread function (PSF) over the co-located MODIS domain has presented a major development challenge. The algorithm to do so is specified in CERES ATBD subsystem 4.4, which I vectorized and implemented in generate samples.py and geom utils.py.

One issue I encountered is that the analytic equation approximating the PSF (specified to the right) was producing a result with very negative values near the aft border of the oblong CERES centroid.

I found that the cause of the error was that the 'xi' angle in the formula is assumed to be in radians rather than degrees. despite having been bounded and scaled previously by values in degrees.

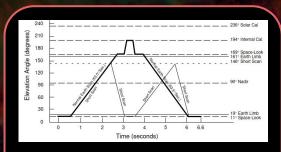






$$\begin{split} F(\xi) &= 1 - (1 + a_1 + a_2) e^{-c_1 \xi} \\ &+ e^{-6.35465 \xi} [a_1 \cos{(1.90282 \xi)} + b_1 \sin{(1.90282 \xi)}] \\ &+ e^{-4.61598 \xi} [a_2 \cos{(5.83072 \xi)} + b_2 \sin{(5.83072 \xi)}] \\ \\ a_1 &= 1.84205 \qquad a_2 = -0.22502 \\ b_1 &= 1.47034 \qquad b_2 = 0.45904 \\ c_1 &= 1.98412 \\ \\ \end{split}$$
 Empirical convolution equation (silently

given in radians)



CERES' alternating scan direction

One related challenge that is still yet to be resolved is that the orientation of the MODIS grid centroid relative to the CERES footprint centroid is determined by the time derivative of CERES' scan elevation (Earth cone angle) α, which alternates between cross-track sweeps, as diagrammed by the figure to the right.

Despite being specified in the ATBD, the derivative quantity isn't included in the CERES SSF data product. Since the above geometry is entirely calculated from relative geodetic values, it also cannot be directly inferred from the data.

The next steps:

The data tensors yielded by **generate_samples.gen_swaths_samples** are structured as tuples: **((M, G, P), C)**, each containing **B** batch samples:

 $\mathbf{M} := (\mathbf{B}, \mathbf{W}, \mathbf{W}, \mathbf{F}_{\mathbf{M}})$ MODIS reflectance or brightness temperature in $\mathbf{F}_{\mathbf{M}}$ bands

 $G := (B,1,1,F_G)^m$ Full-footprint viewing geometry $F_G := (sol_zen, sat_zen, rel_azi)$

P := (B,W,W,1) 2D PSF probability distribution over the MODIS domain

 $C := (B,1,1,F_c)$ CERES footprint over the MODIS domain $F_c := (lw_flux, sw_flux)$

The generator also provides options for linear normalization, and for constraining the parameters and sources of the data it returns.

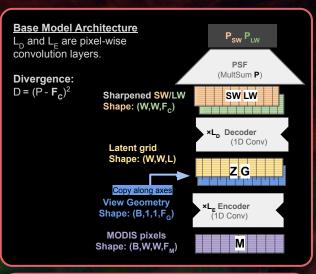
The full dataset includes observations from 2 different sensors, which were collected over 6 distinct regions having widely varying characteristics.

These statistical sub-populations may make the flux prediction significantly more complex, so I will train separate models using stratified sampling techniques on subsets of the data.

The basic model architecture consists of a sequence of 1D convolutions over the MODIS pixels to produce an independent latent representation of the radiance. The latent domain is concatenated with the viewing geometry (which is the same for all the pixels). Next, a second sequence of pixel-wise convolutional layers decodes the intermediate data to pixel-wise SW/LW predictions. Finally, the predictions are aggregated by a weighted sum over the PSF, and the loss is calculated with the result relative to the CERES footprints over the domain.

If time allows, I will experiment with applying the PSF aggregation to the latent grid before concatenating the resulting vector with the viewing geometry, and decoding with a simple multi-layer perceptron.

Since the entire model is pixel-wise, the full range of its predictions can be evaluated by modulating a single vector of input features. As such, it will also be interesting to see how predictions change with respect to the viewing geometry and MODIS spectral response.



Model training plan:

First, do a broad random search of model hyperparameters on the simplest possible task (ie, a single region/season/satellite combo)

Then, using the most promising learning rate schedule and regularization method:

- Train a large and small model on each individual region, using only one satellite.
- Train a large and small model combining all swaths, seasons, and satellites.

Finally, evaluate each of the models on unseen data from all of the input sub-populations.