

Physics Constrained Neural Networks for Designing Gravity Assist Trajectories

Master Thesis Midterm Review

Mitchell van Doorn

4-11-2024



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1

Recap Research Proposal

Research questions

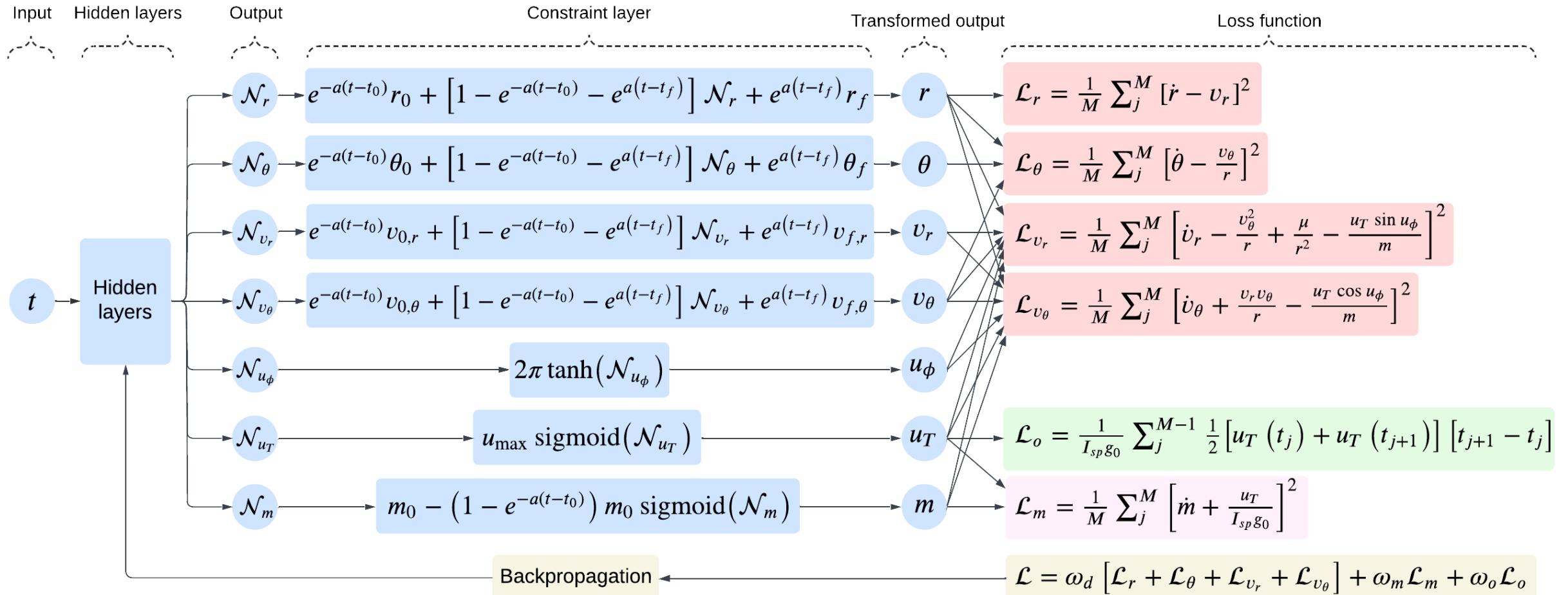
Main Research Question:

“How can Physics-Constrained Neural Networks (PCNNs) be effectively utilized to design optimal trajectories involving gravity assist for spacecraft with low-thrust propulsion systems?”

Sub questions:

- How can the dynamics of gravity assist trajectories be integrated in PCNNs?
- How can PCNNs learn to perform gravity assist maneuvers?
- How can suitable neural network architectures be established?
- How can the PCNN model be trained effectively?
- What are the benefits and limitations of using PCNNs for optimizing low-thrust gravity assists trajectories?

Basic PCNN design



Subproblems to be solved

Subproblem 1: Earth → Jupiter rendezvous

- Fixed x_0, x_f, ToF

Subproblem 2: Earth → Fixed Time Jupiter GA → Pluto

- Fixed $x_0, x_f, \text{ToF}, t_{GA}$
- Variable x_{GA}

Subproblem 3: Earth → Jupiter GA → Pluto

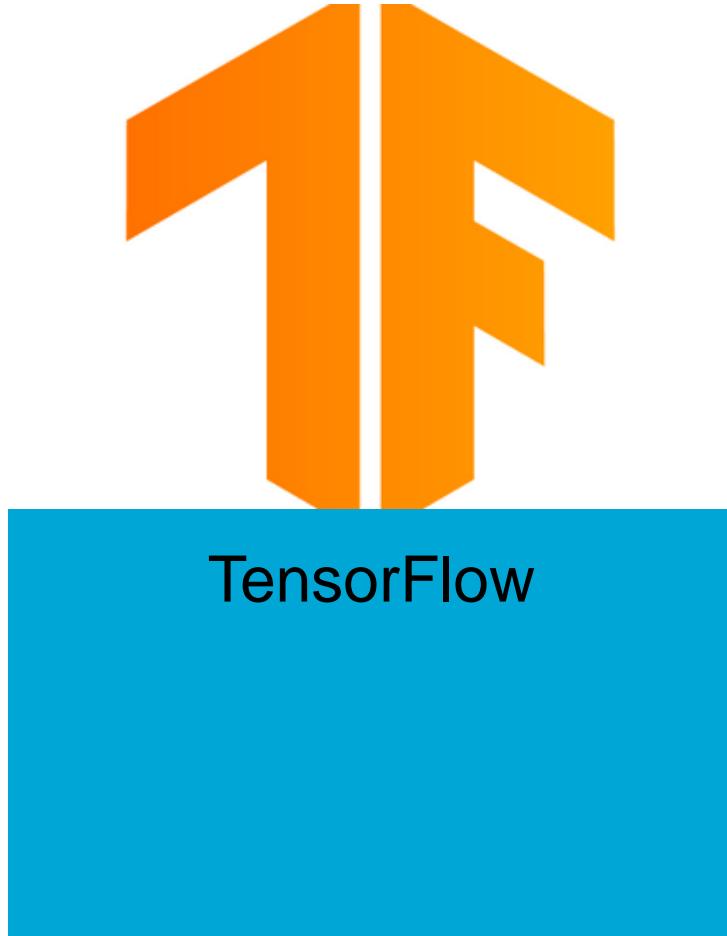
- Fixed x_0, x_f, ToF
- Variable x_{GA}, t_{GA}



2

Results so far

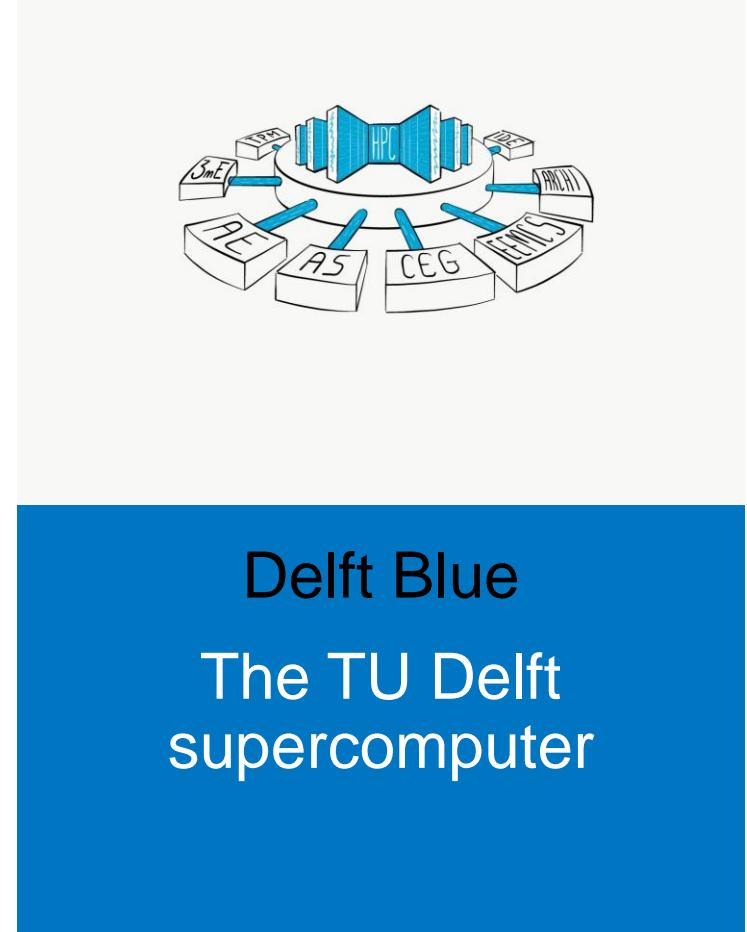
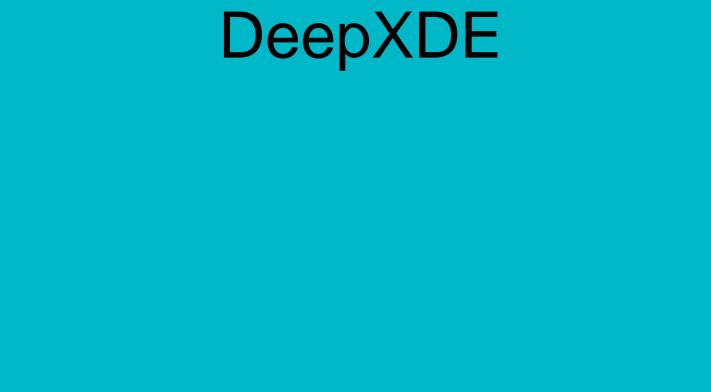
Software/hardware Exploration



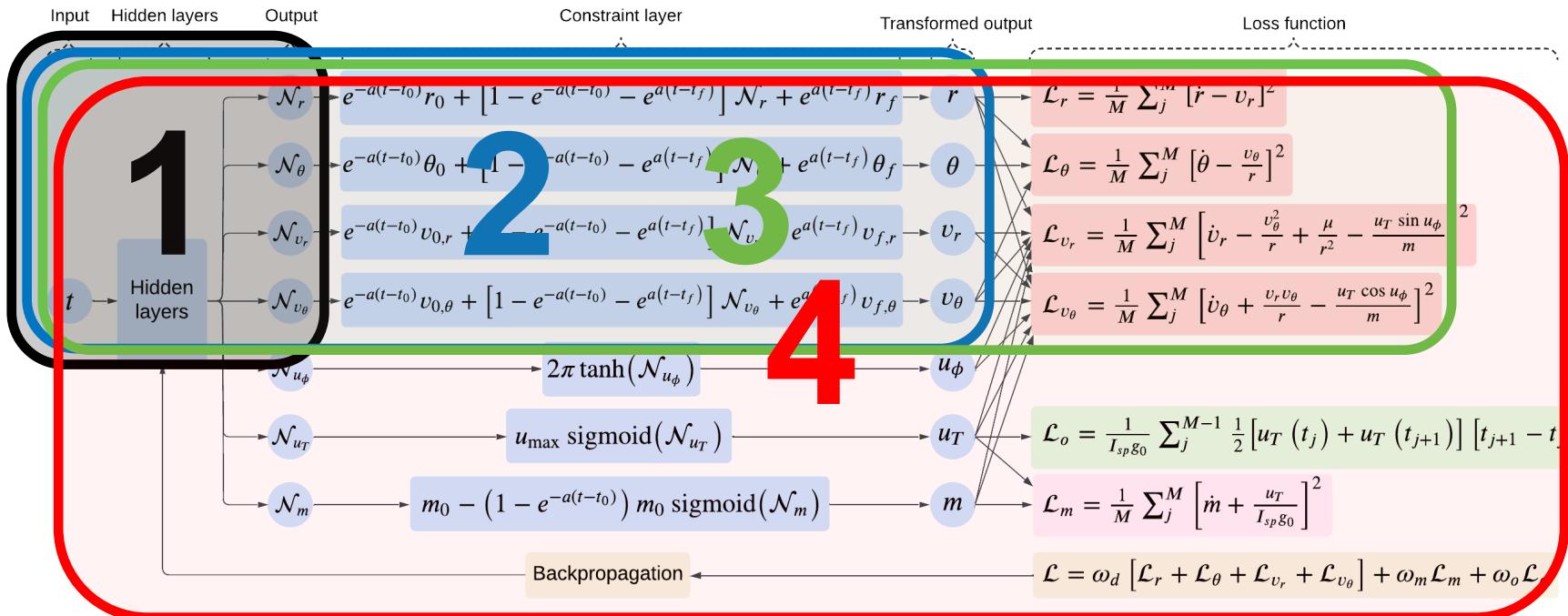
DeepXDE: A Deep Learning Library for Solving Differential Equations*

Lu Lu[†]
Xuhui Meng[‡]
Zhiping Mao[§]
George Em Karniadakis[¶]

Abstract. Deep learning has achieved remarkable success in diverse applications; however, its use in solving partial differential equations (PDEs) has emerged only recently. Here, we present an overview of physics-informed neural networks (PINNs), which embed a PDE into the loss of the neural network using automatic differentiation. The PINN algorithm is simple, and it can be applied to different types of PDEs, including integro-differential equations, fractional PDEs, and stochastic PDEs. Moreover, from an implementation point of view, PINNs solve inverse problems as easily as forward problems. We propose a new residual-based adaptive refinement (RAR) method to improve the training efficiency of PINNs. For pedagogical reasons, we compare the PINN algorithm to a standard finite element method. We also present a Python library for PINNs, DeepXDE, which is designed to serve both as an educational tool to be used in the classroom as well as a research tool for solving problems in computational science and engineering. Specifically, DeepXDE can



Basic PCNN implementation



Toy problem:
Circular Earth orbit (500km altitude)

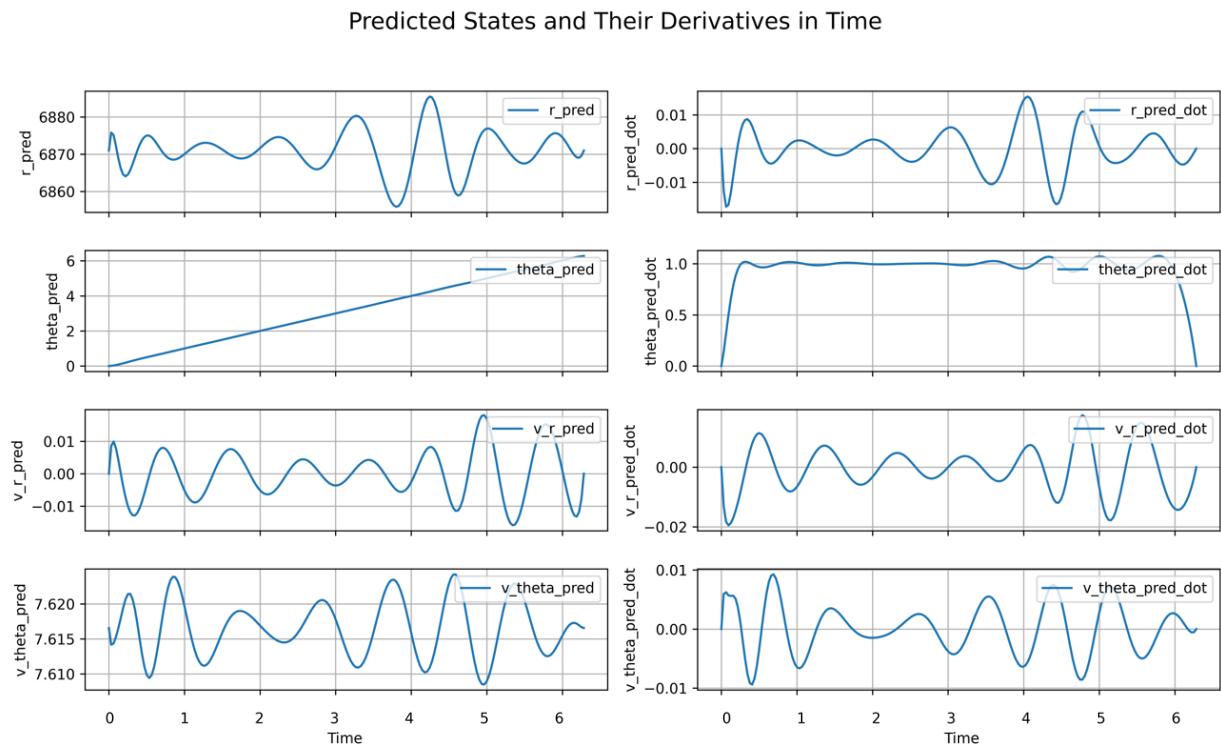
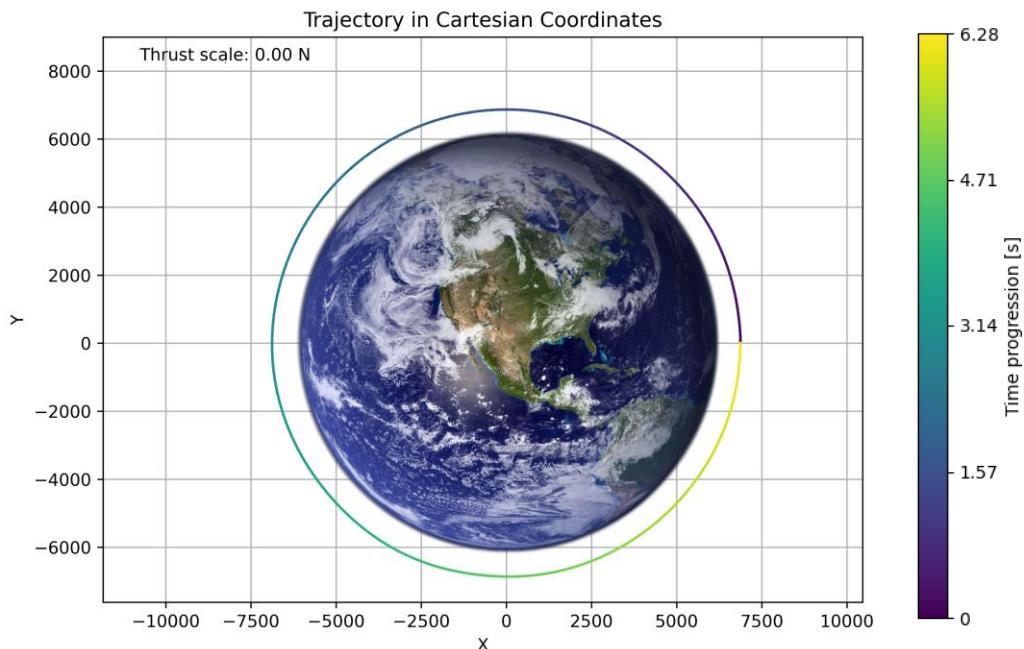


1. NN supervised
2. NN supervised + constraint layer
3. NN unsupervised + constraint layer + custom loss function (dynamics only)
4. NN unsupervised + constraint layer + custom loss function (dynamics + objective + mass)

Setting	Baseline Configuration
Neurons	20
Hidden Layers	5
Activation	sin
Training points M	200
Learning rate Schedule	1
$[\omega_d \quad \omega_m \quad \omega_o]$	$[1 \quad 10^{-5} \quad 10^{-7}]$
a	10

Results

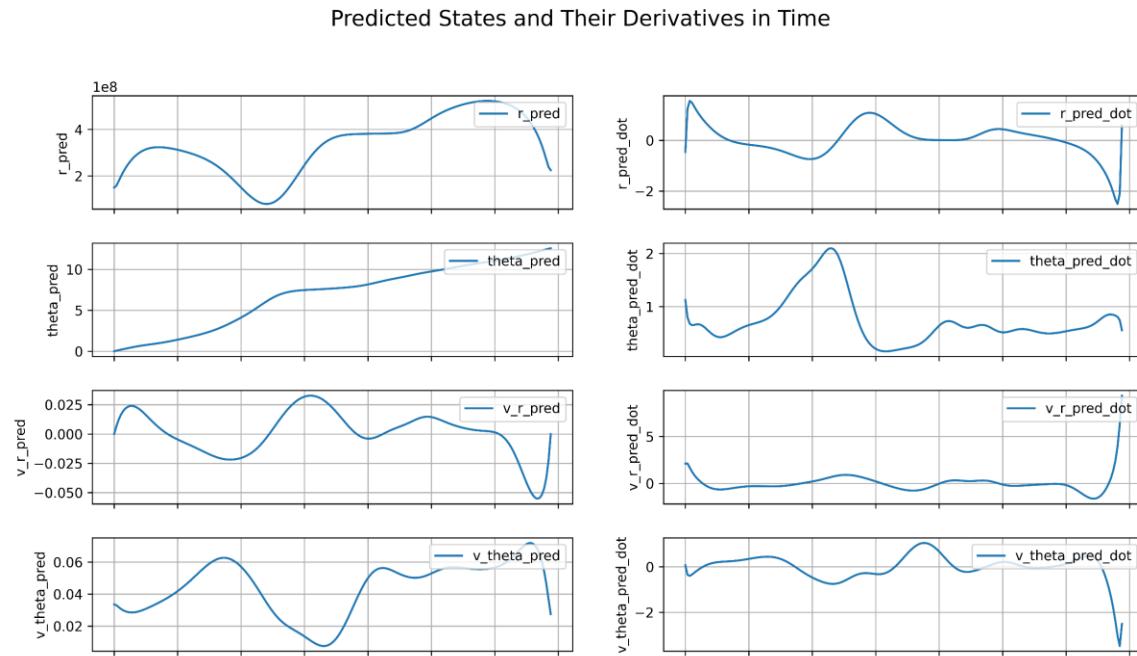
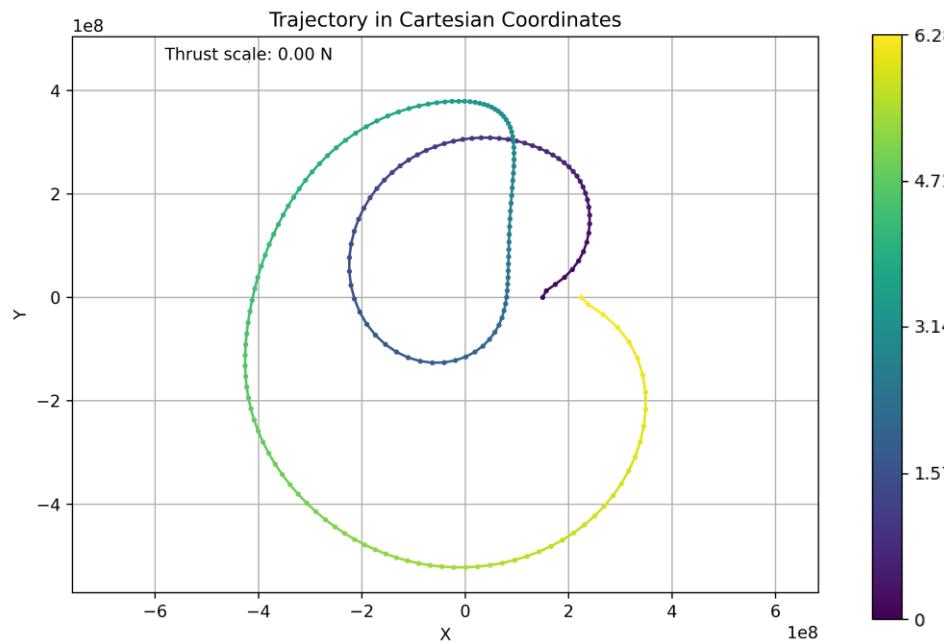
1. NN supervised
2. NN supervised + constraint layer



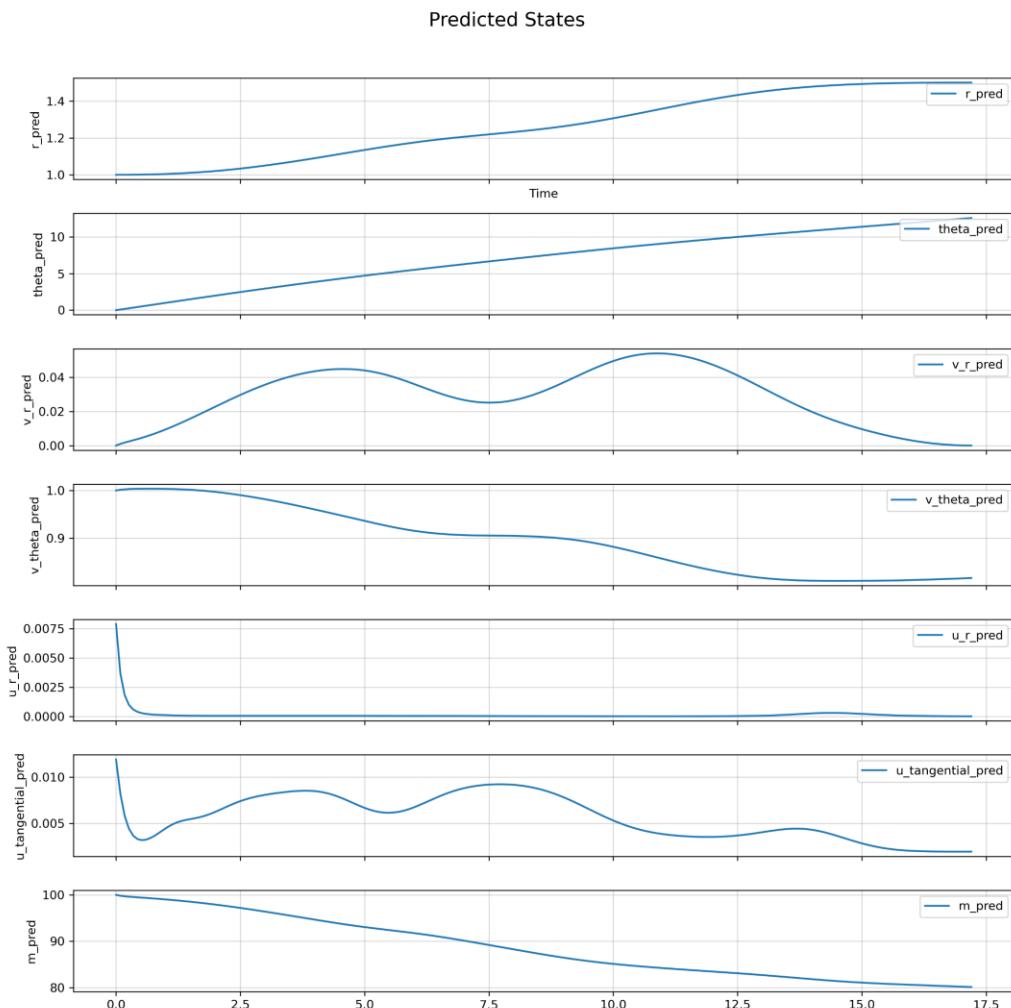
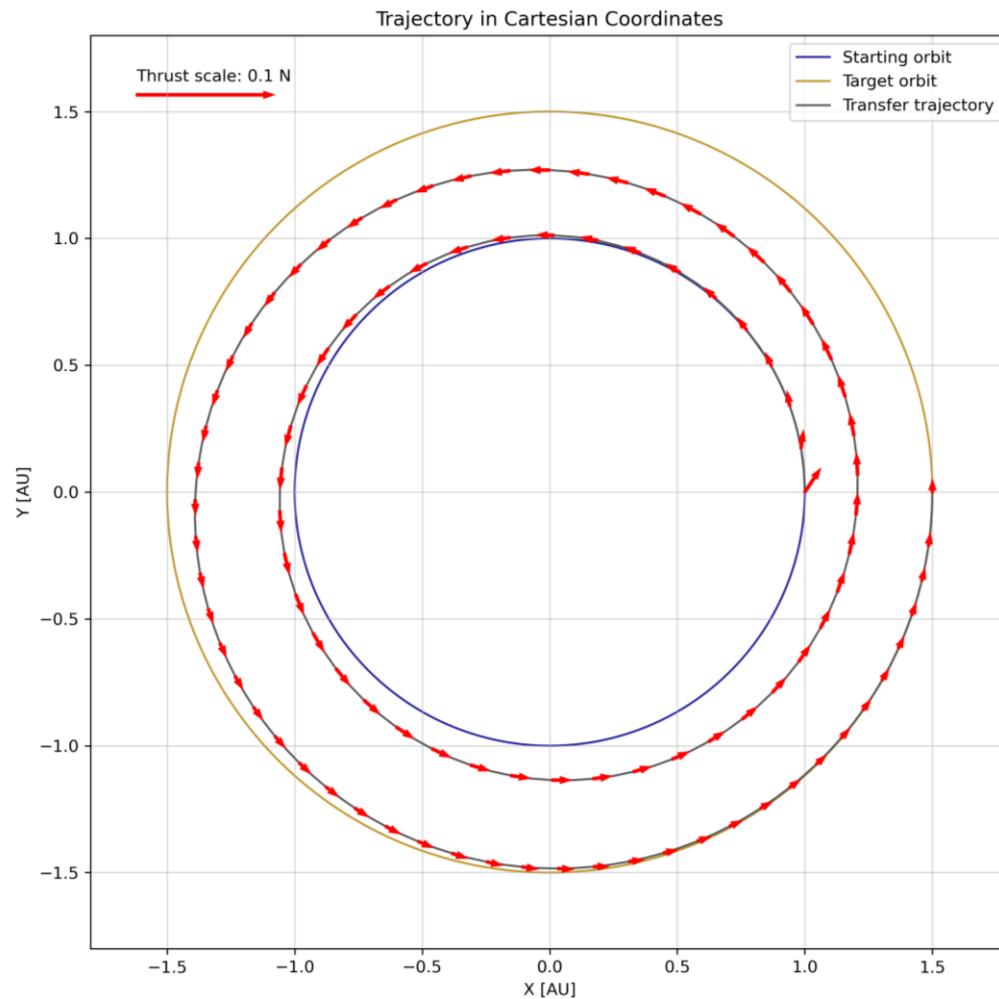
$$v = \sqrt{\frac{GM}{r}} = \sqrt{\frac{\mu}{r}}$$

Results

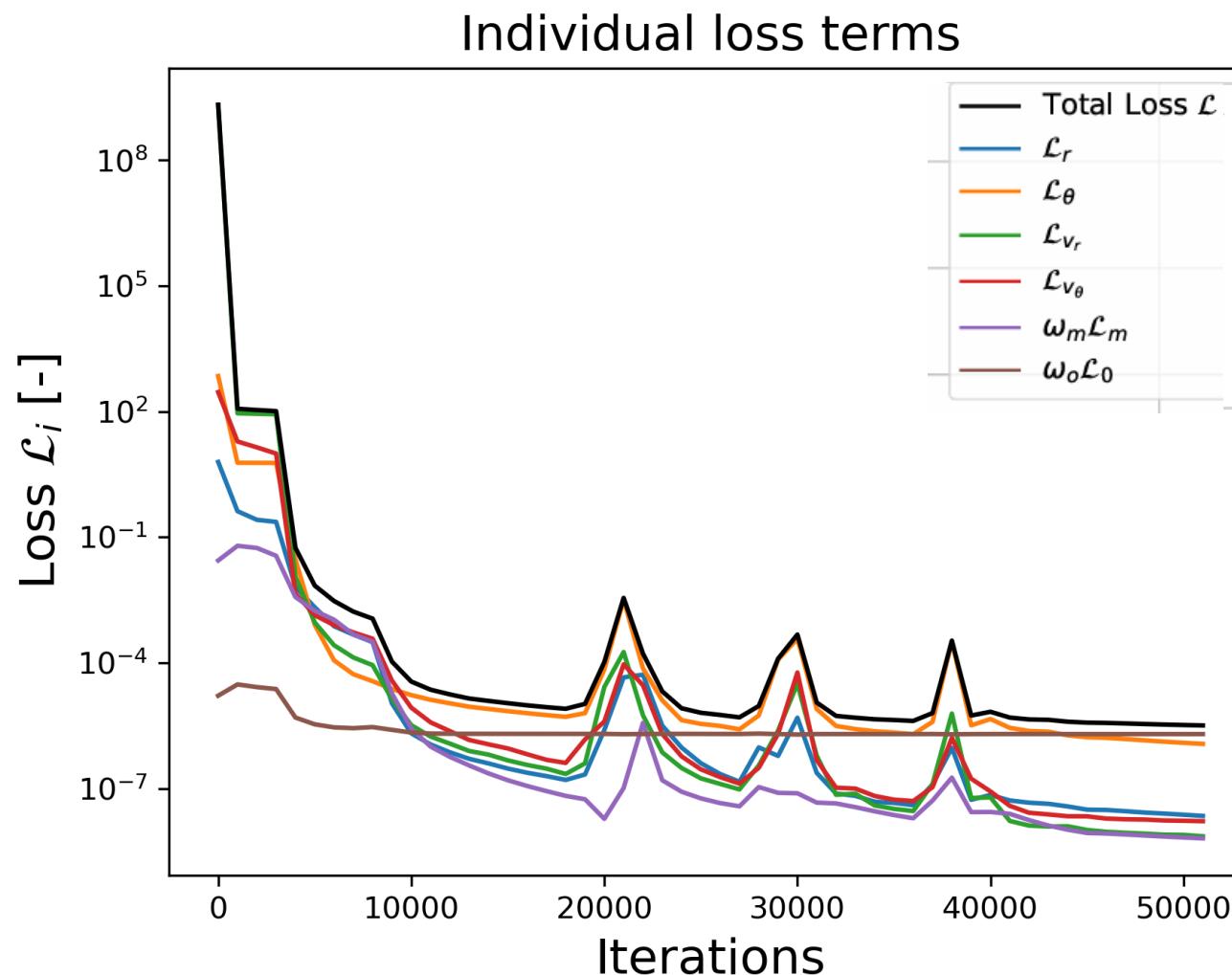
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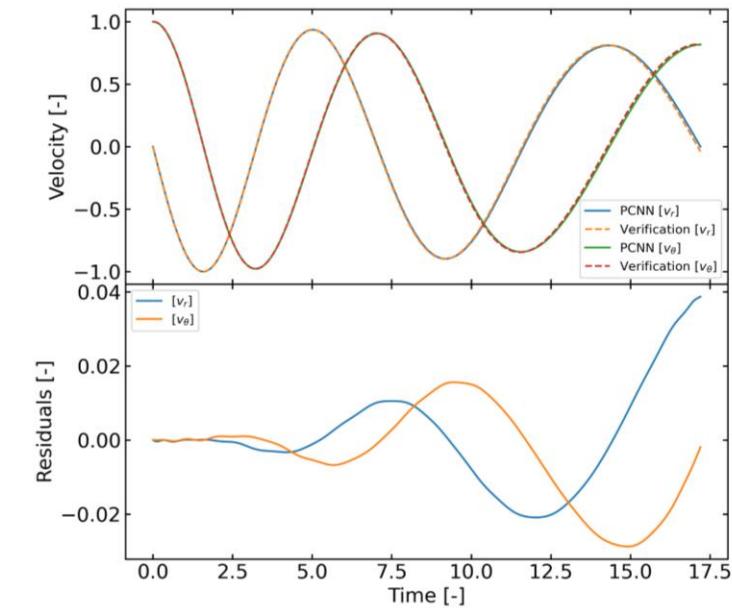
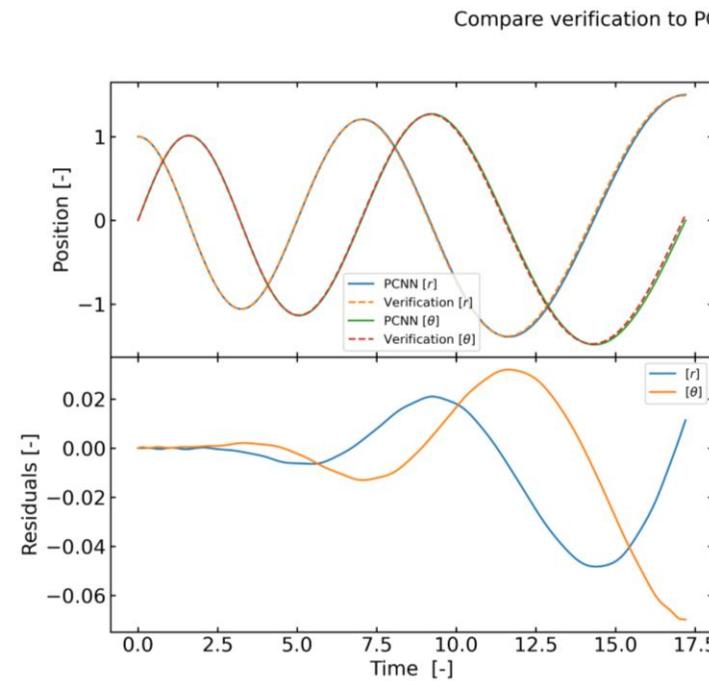
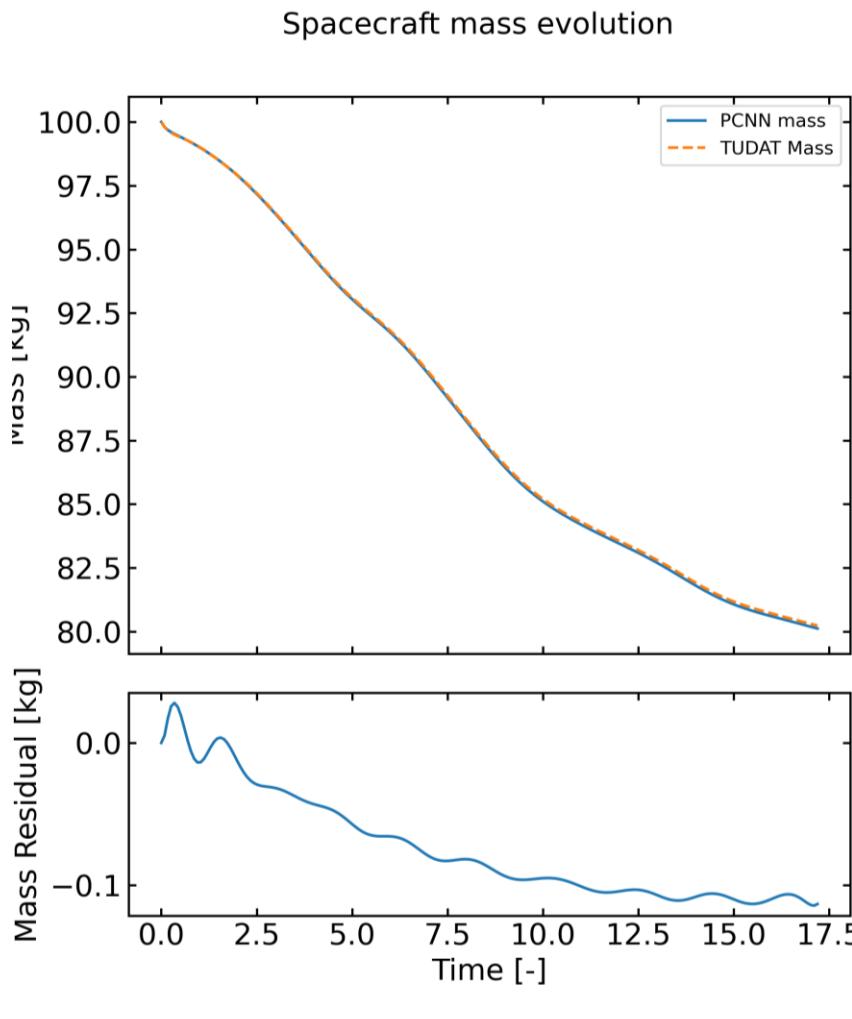
Results with DeepXDE - Trajectory



Results with DeepXDE – Loss evolution



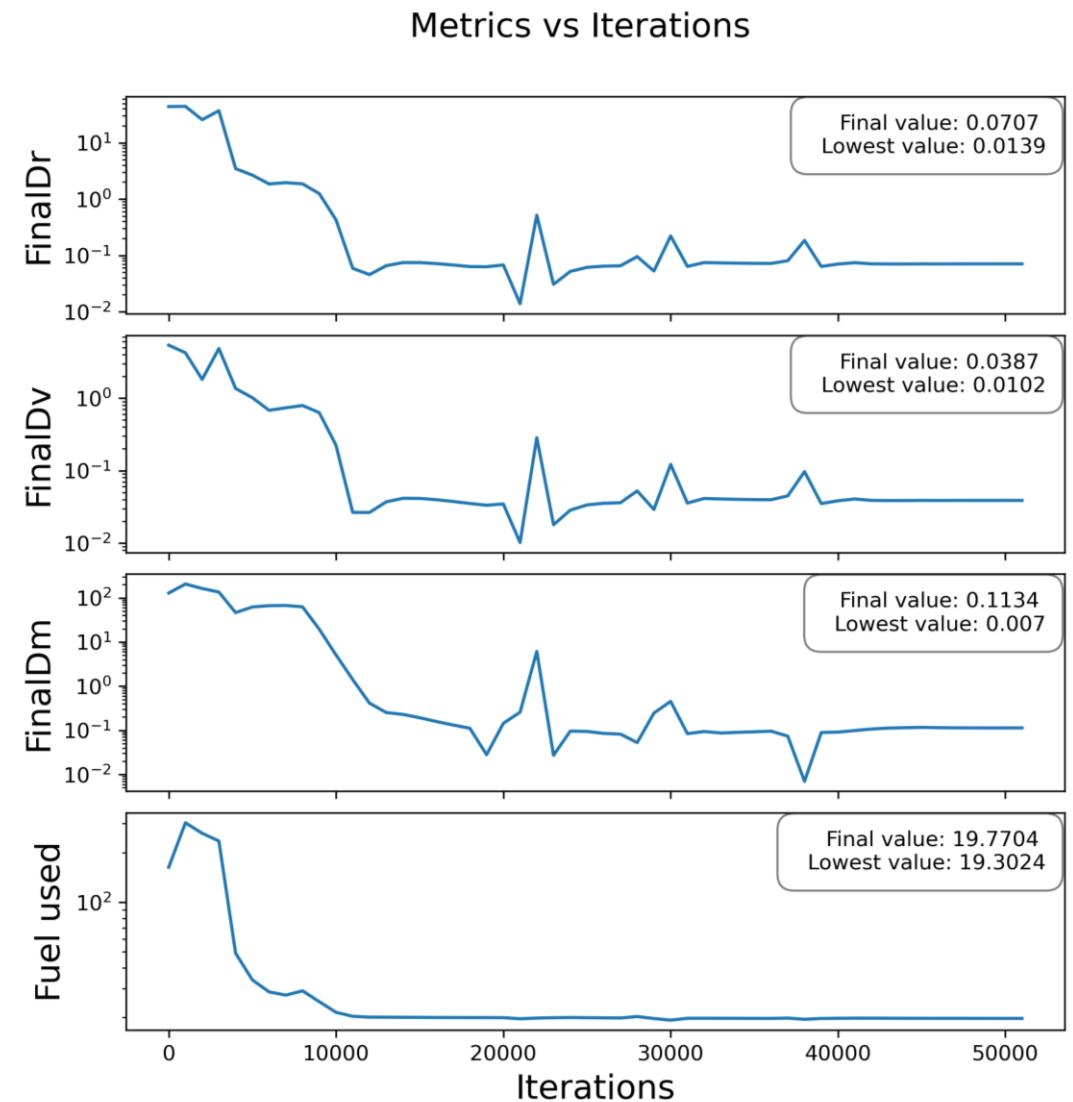
Results with DeepXDE – Comparison with TUDat



Numerical integrator used:
Runge-Kutta 7(8) variable step-size integrator
Rel. & abs. tolerances: 10e-10

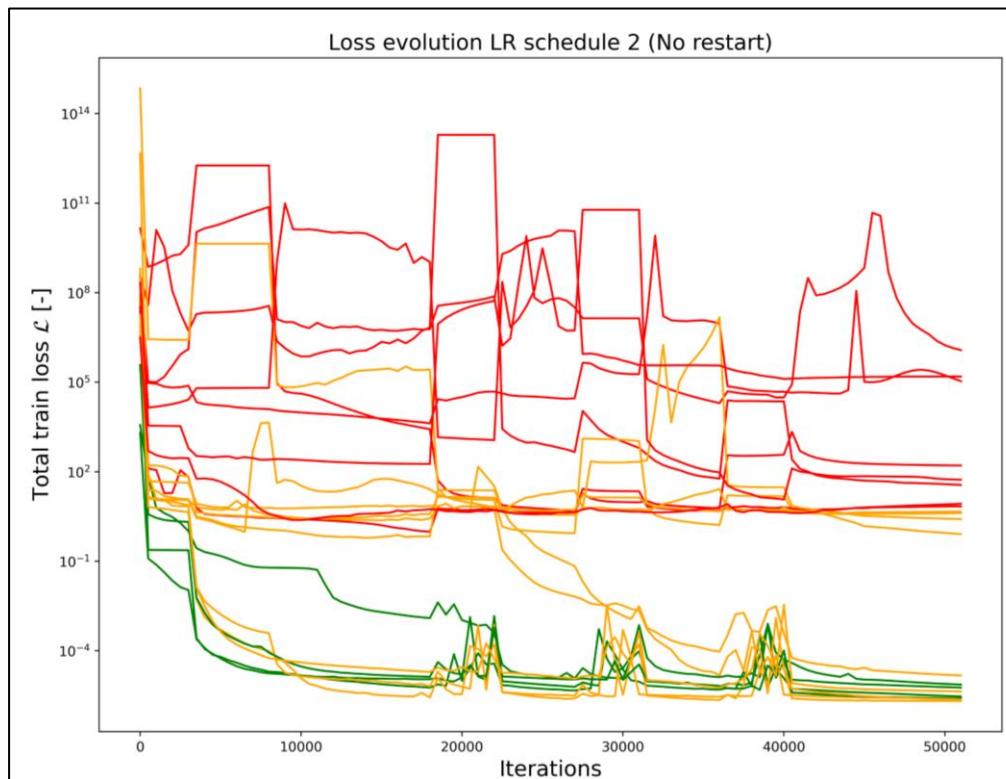
Results with DeepXDE – Metrics

$$dx = \sqrt{(x(t_f) - x_v(t_f))^2 + (y(t_f) - y_v(t_f))^2}$$
$$dv = \sqrt{(v_x(t_f) - v_{x,v}(t_f))^2 + (v_y(t_f) - v_{y,v}(t_f))^2}$$
$$dm = m(t_f) - m_v(t_f)$$



Verification & Validation

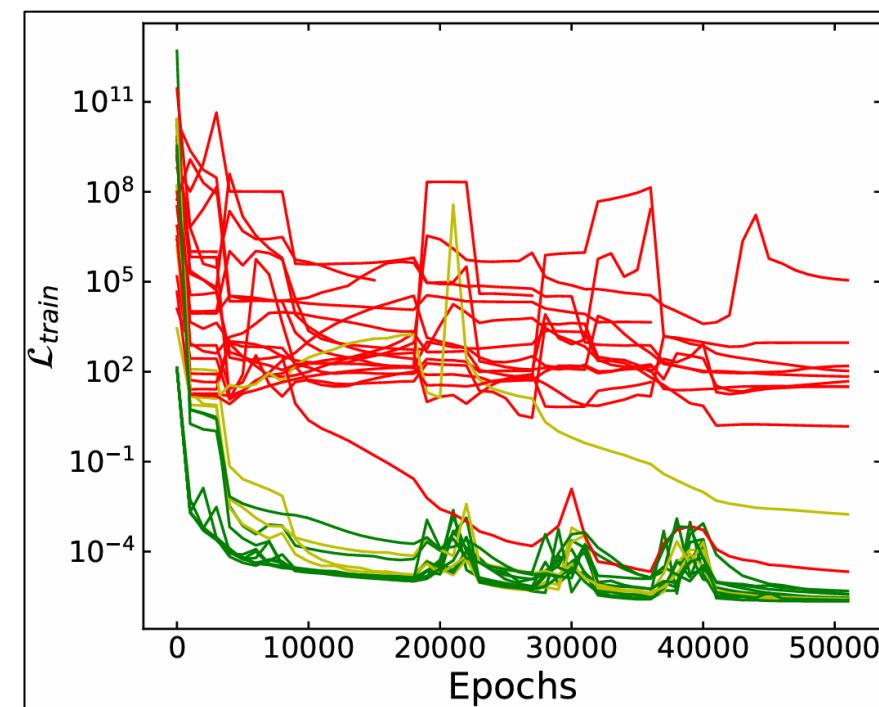
Loss evolutions



This work

Schedule 2

$[10^{-2}, 3k] \rightarrow [10^{-3}, 5k] \rightarrow [10^{-4}, 10k]$
$\rightarrow [5 \times 10^{-3}, 4k] \rightarrow [10^{-4}, 5k]$
$\rightarrow [5 \times 10^{-3}, 4k] \rightarrow [10^{-4}, 5k]$
$\rightarrow [5 \times 10^{-3}, 4k] \rightarrow [10^{-4}, 5k] \rightarrow [10^{-5}, 6k]$

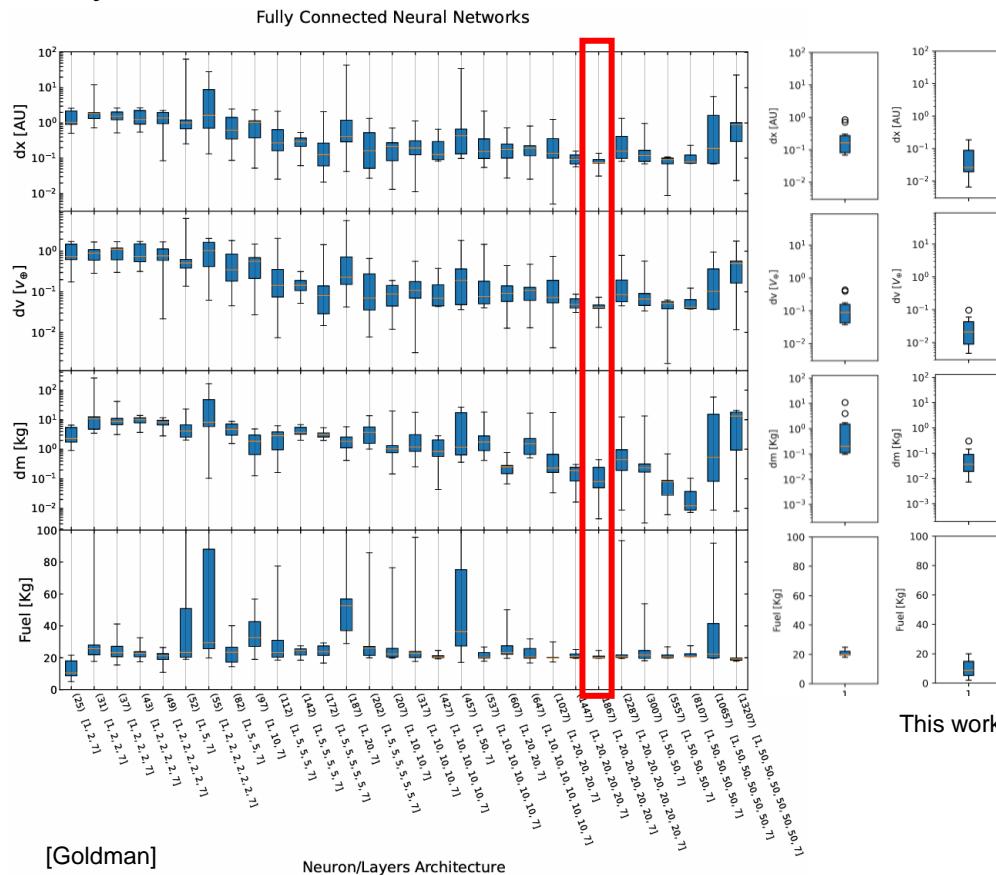


Goldman's work

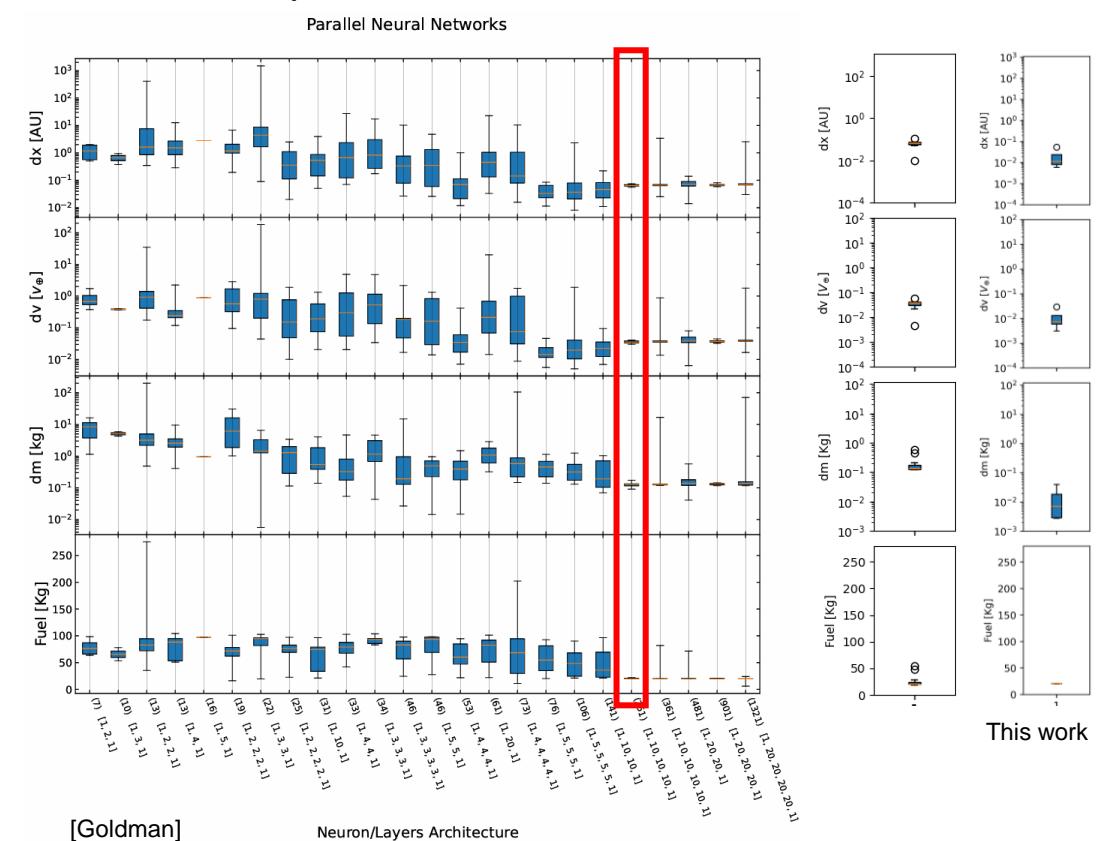
Verification & Validation

Metrics

Fully connected NNs



Bundle of 7 parallel NNs



Verification & Validation

Metrics

Median values comparison

Latest iteration

	FNN Gold-man [1, 20, 20, 20, 20, 20, 7]	FNN this work [1, 20, 20, 20, 20, 20, 7]	Bundle Goldman [1, 10, 10, 10, 10, 1]	Bundle this work [1, 10, 10, 10, 10, 1]
Median dx [AU]	0.076	0.085	0.064	0.071
Median dv [v_{\oplus}]	0.041	0.045	0.035	0.0389
Median dm [kg]	0.081	0.123	0.129	0.1375
Median propellant mass m_p [kg]	19.93	20.04	19.96	20.53

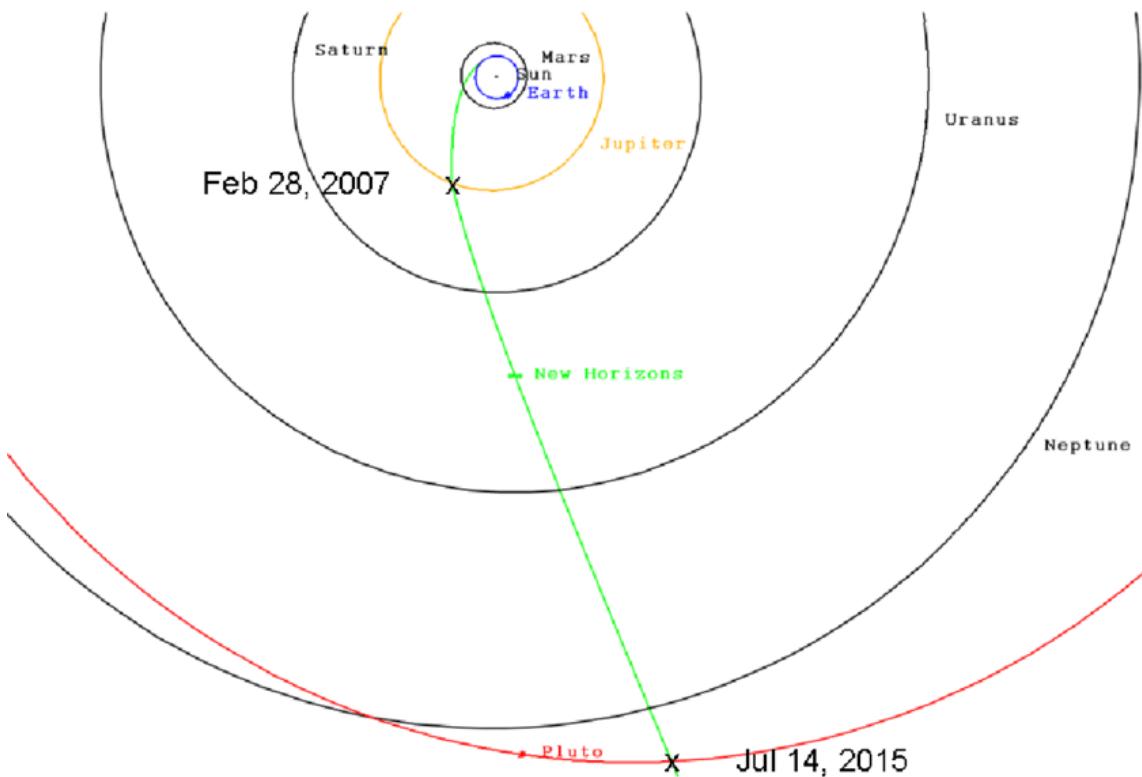
Lowest metrics:

	FNN Gold-man [1, 20, 20, 20, 20, 20, 7]	FNN this work [1, 20, 20, 20, 20, 20, 7]	Bundle Goldman [1, 10, 10, 10, 10, 1]	Bundle this work [1, 10, 10, 10, 10, 1]
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Median dv [v_{\oplus}]	0.041	0.019	0.035	0.0064
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Median propellant mass m_p [kg]	19.93	13.77	19.96	20.27

Pre-liminary results subproblem 1

~~Subproblem 1: Earth → Jupiter rendezvous (New Horizons missions)~~

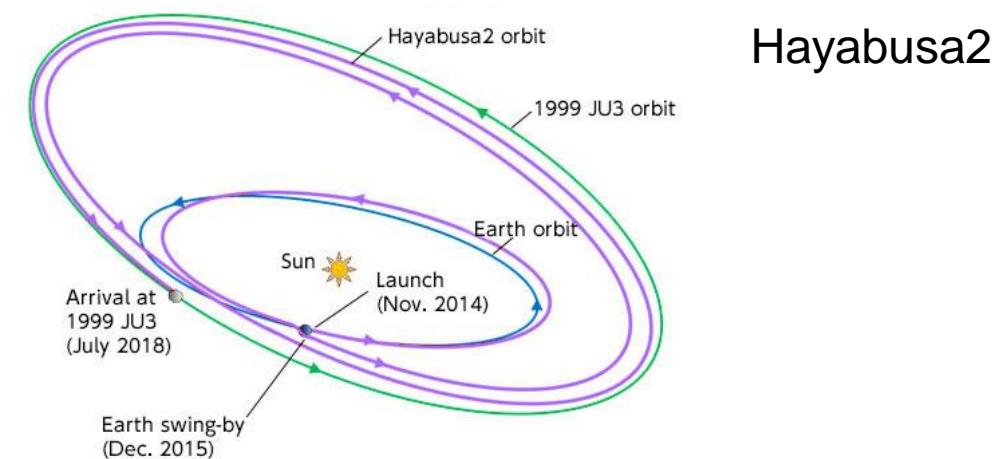
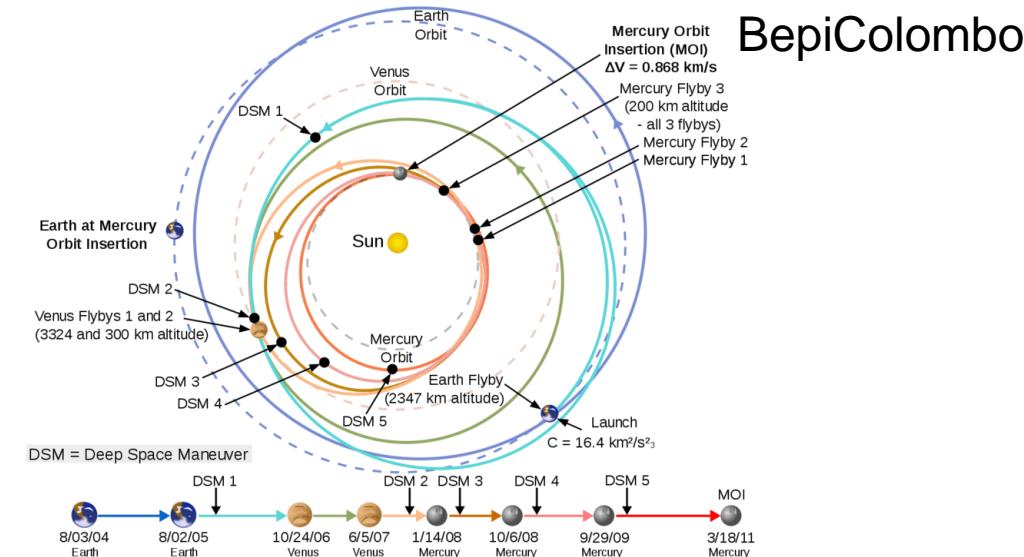
- Fixed x_0 , x_f , ToF



Pre-liminary results subproblem 1

New Subproblem 1:

- Fixed x_0 , x_f , ToF
- Requirements:
 - Involves gravity assists essential to reach the target destination.
 - Utilizes low-thrust propulsion (e.g., ion engines or solar electric propulsion)
 - Has publicly available trajectory data, including the spacecraft's state history
- Choice: BepiColombo



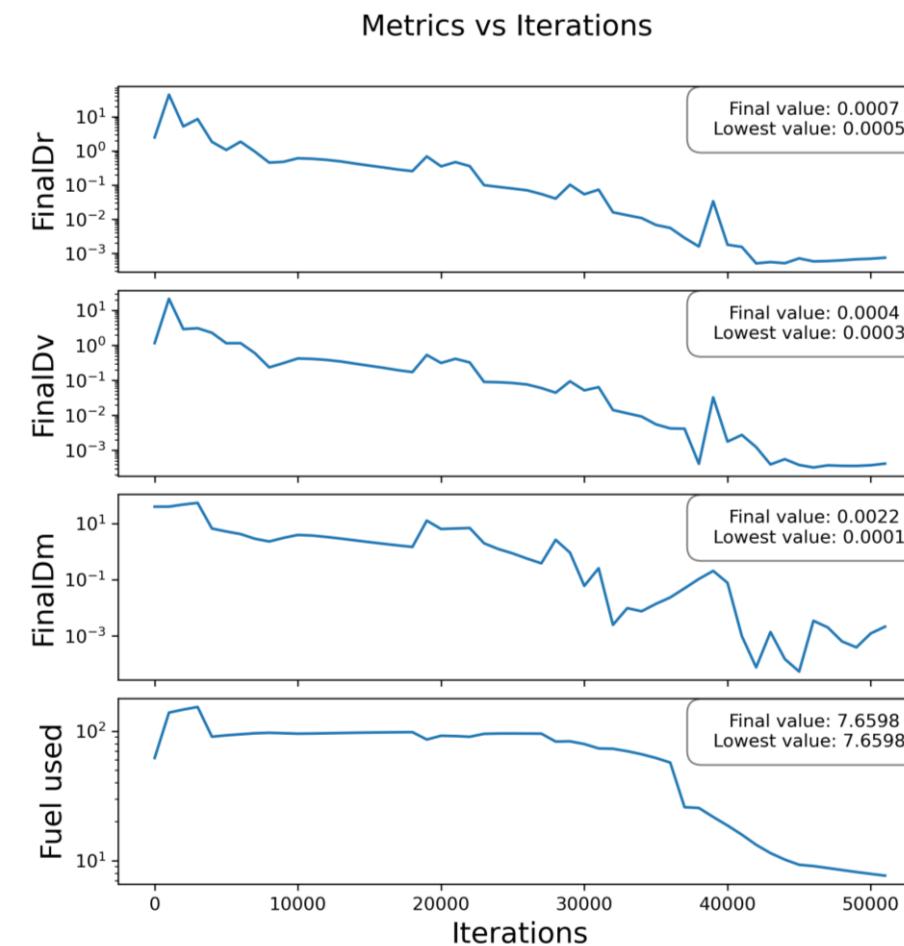
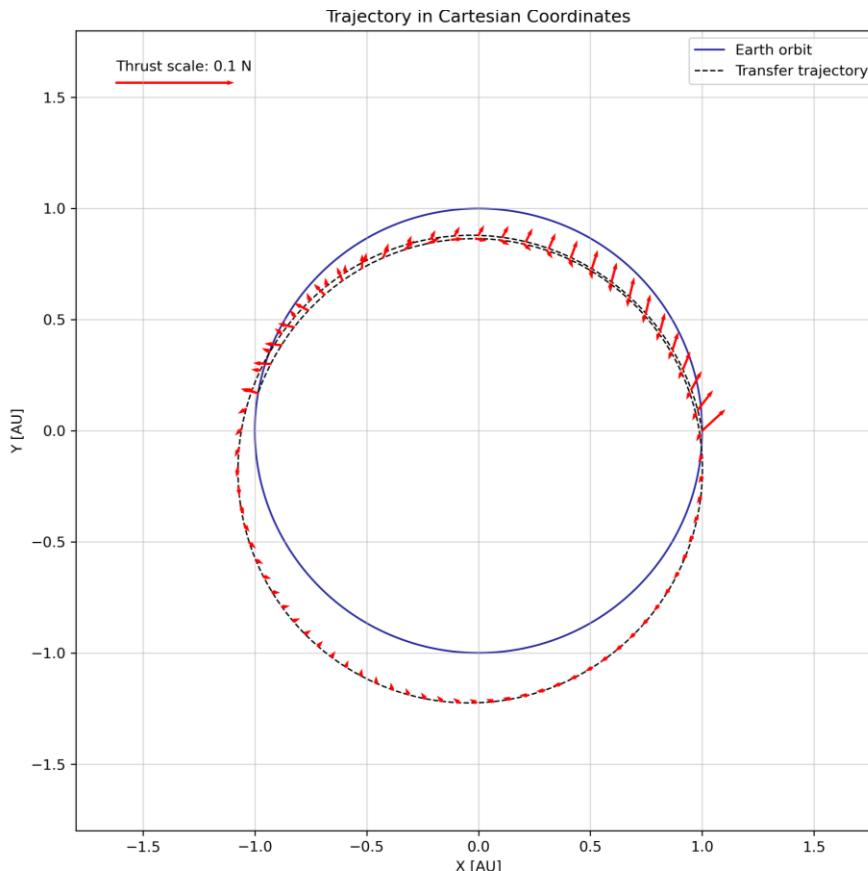
BepiColombo

Hayabusa2

Pre-liminary results subproblem 1

Subproblem 1: Earth → Earth rendezvous

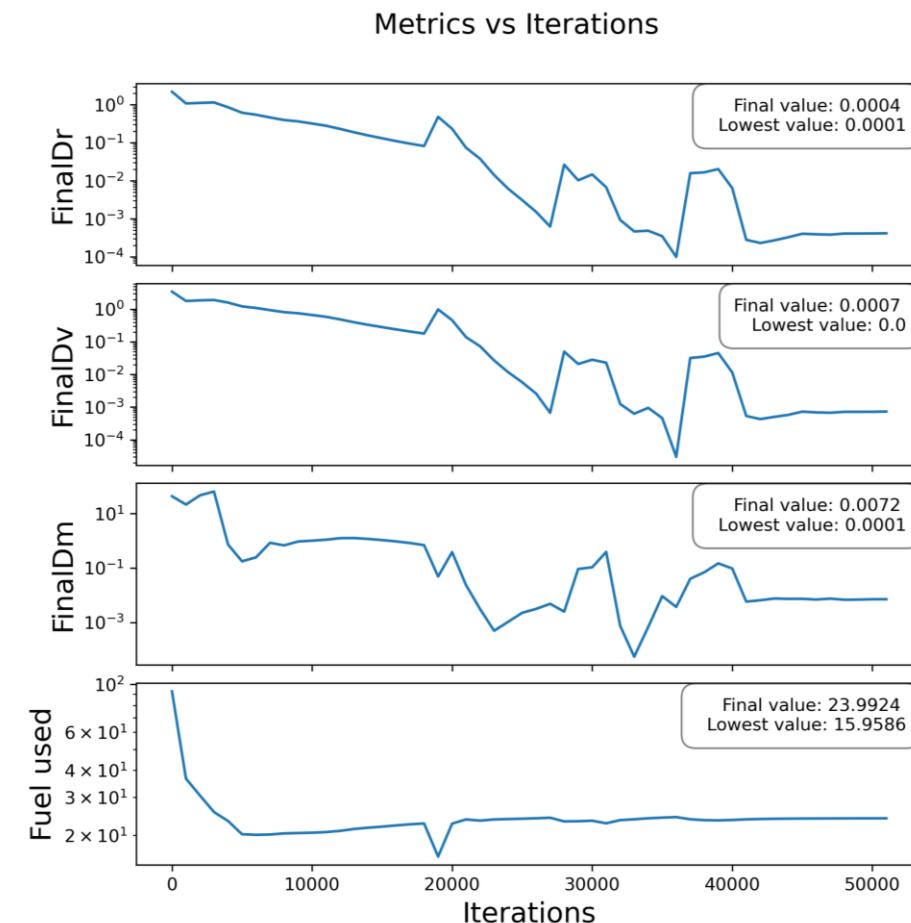
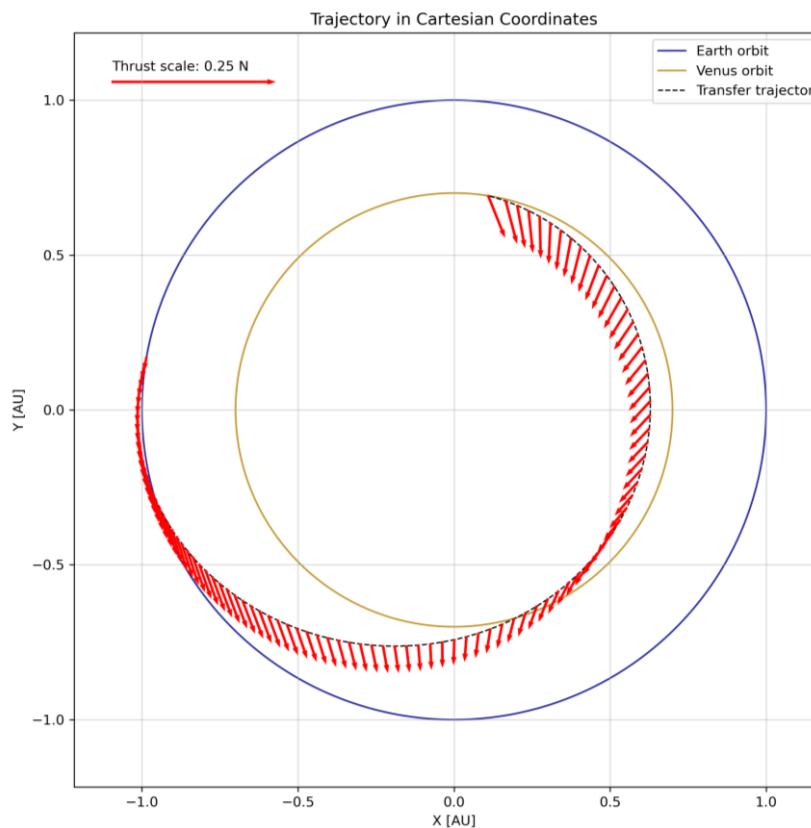
Trajectory leg 1



Pre-liminary results subproblem 1

Subproblem 1: Earth → Earth rendezvous

Trajectory leg 2

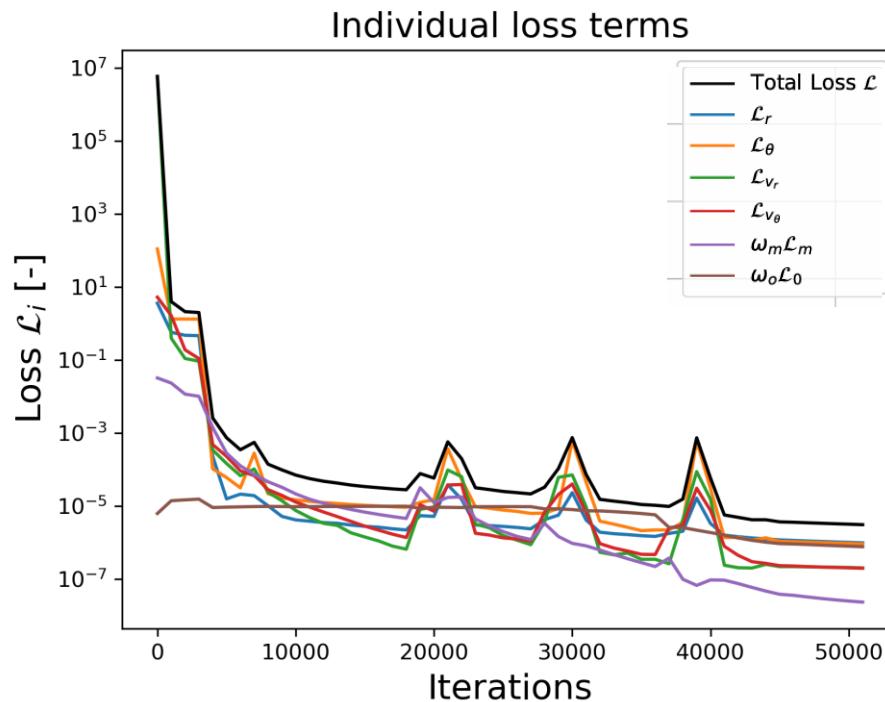


Pre-liminary results subproblem 1

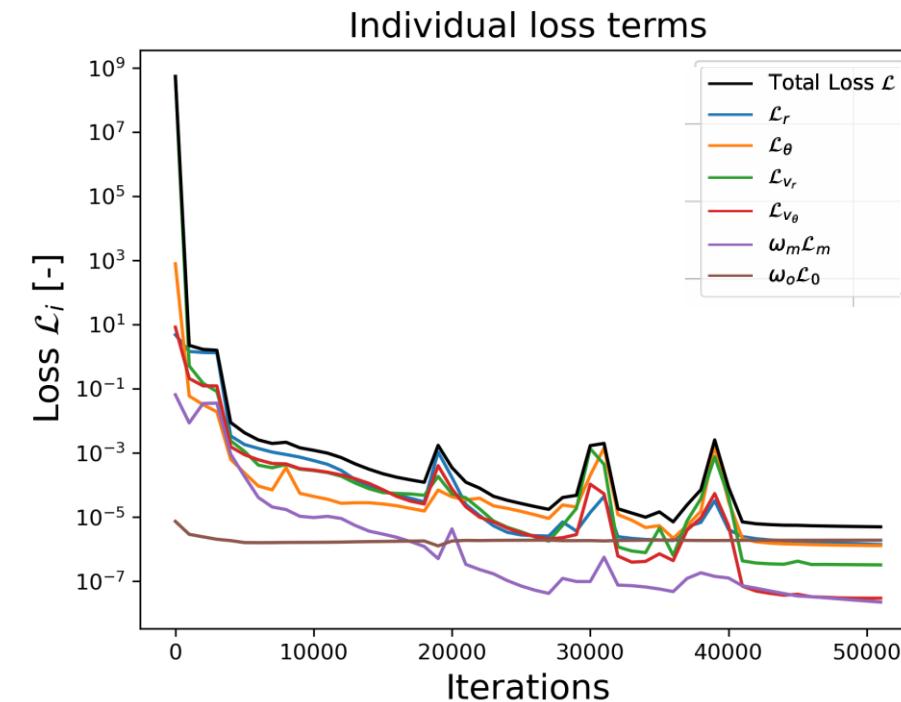
Subproblem 1: Earth → Earth rendezvous

Loss evolutions

Leg 1



Leg 2



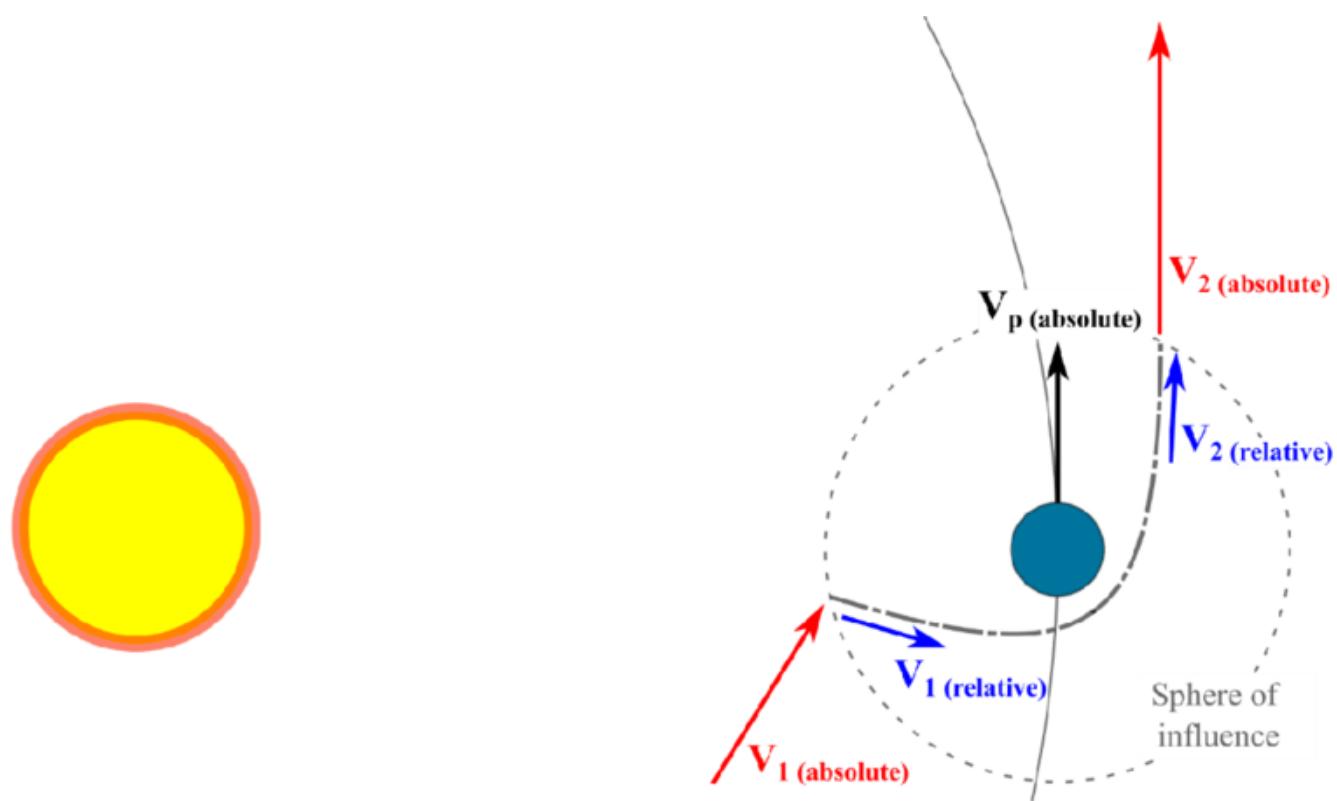


3

Analysis and reflection of results

Analysis and reflection of results

- Accuracy of the method



Analysis and reflection of results

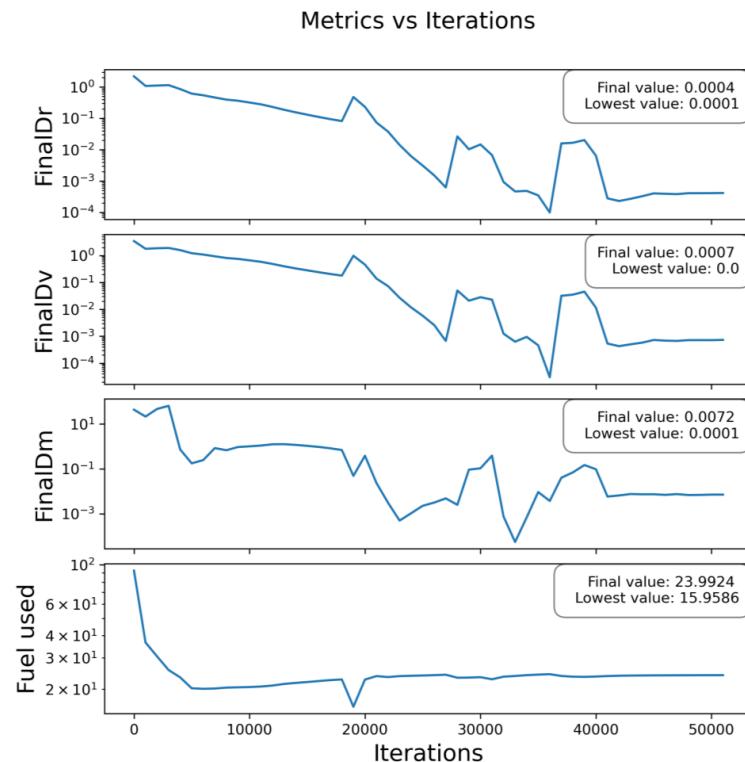
- Accuracy of the method

0.0001 AU = 15 000 km

0.0005 AU = 75 000 km

0.001 AU = 150 000 km

0.01 AU = 1 500 000 km



0.0640 AU = 9 600 000 km

SOI Earth = 924 000 km

SOI Jupiter = 48 200 000

Lowest metrics:

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Latest iteration

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Analysis and reflection of results

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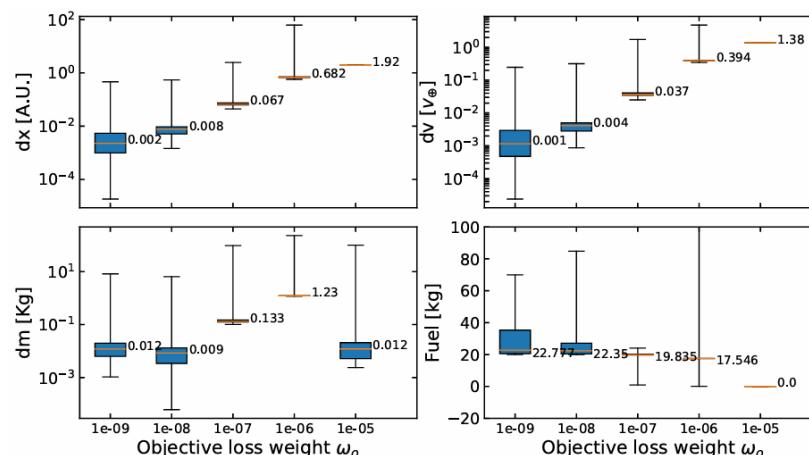
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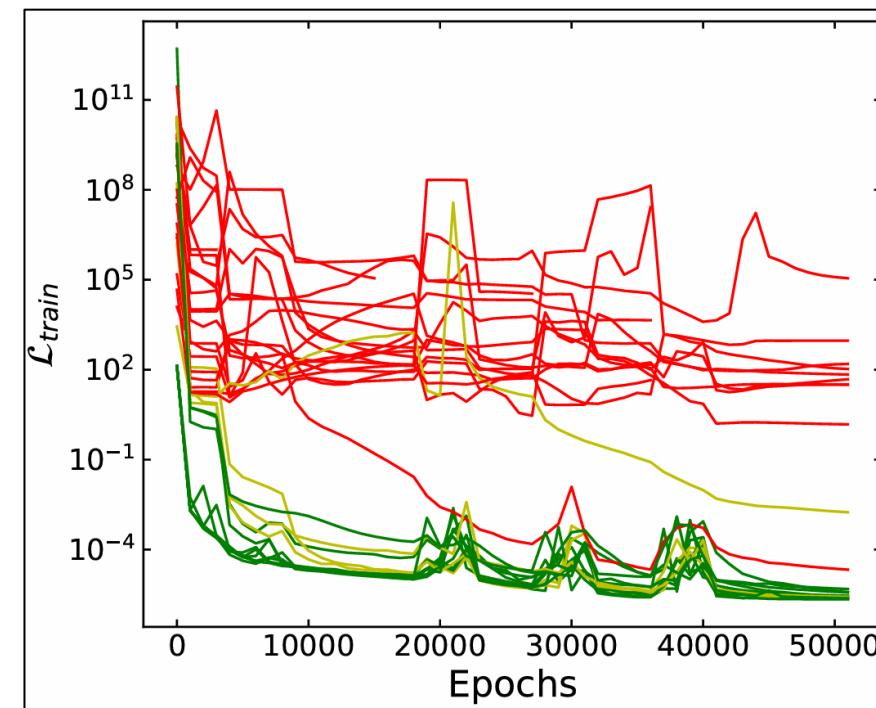
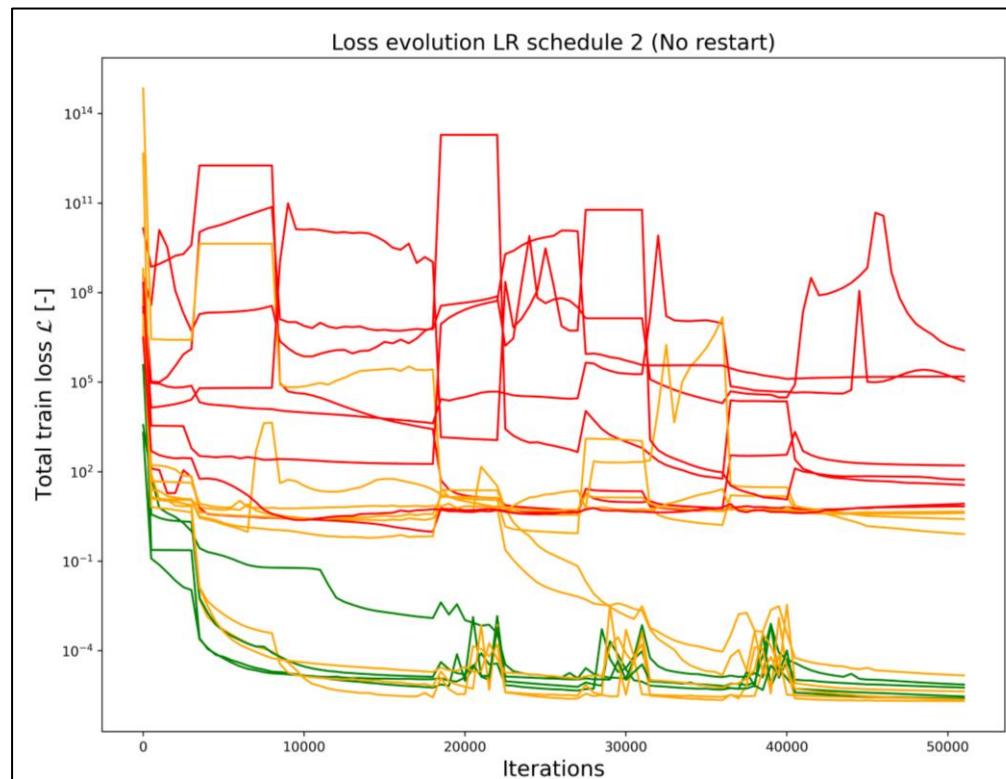
Fixes:

- Lower objective weight
- Try different activation function



Analysis and reflection of results

- Dependence on initialization of NN



Analysis and reflection of results

V&V subproblem 1 + 2

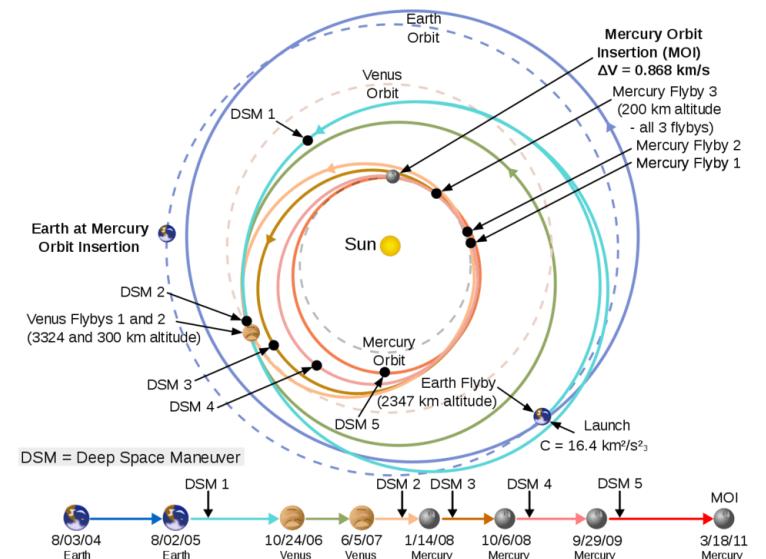
Verification

- Numerical integrations Tudat
- Data BepiColombo

Validation

- Hodographic shaping
- Data BepiColombo?

The screenshot shows the NAIF (Navigation and Ancillary Information Facility) website. At the top, there's a NASA logo and a link to the NASA Portal. Below that is a large banner with the text "NAIF" and "The Navigation and Ancillary Information Facility". On the left, there's a sidebar with links for Home, Announcements, About SPICE, and About NAIF. The main content area has a section titled "SPICE Data (SPICE Kernels)" with links to PDS Archived SPICE Data Sets, Operational Flight Project Kernels and Other Non-archived Project Kernels, and Generic Kernels.



Note

Generated by nbsphinx from a Jupyter notebook. All the examples as Jupyter notebooks are available in the tudatpy-examples repo.

Hodographic-shaping MGA transfer optimization with PyGMO

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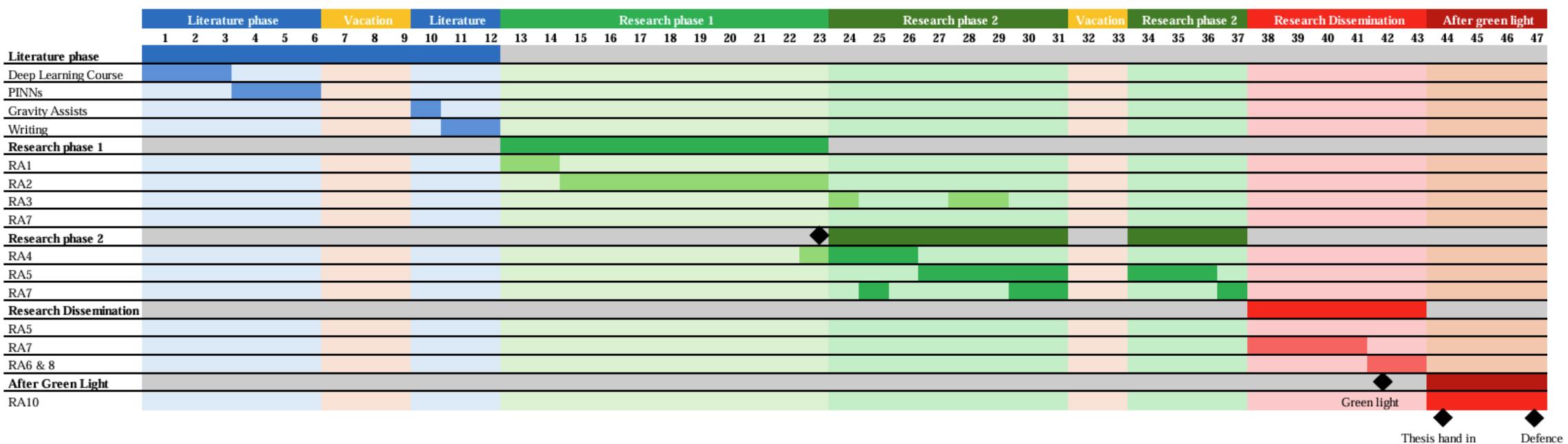
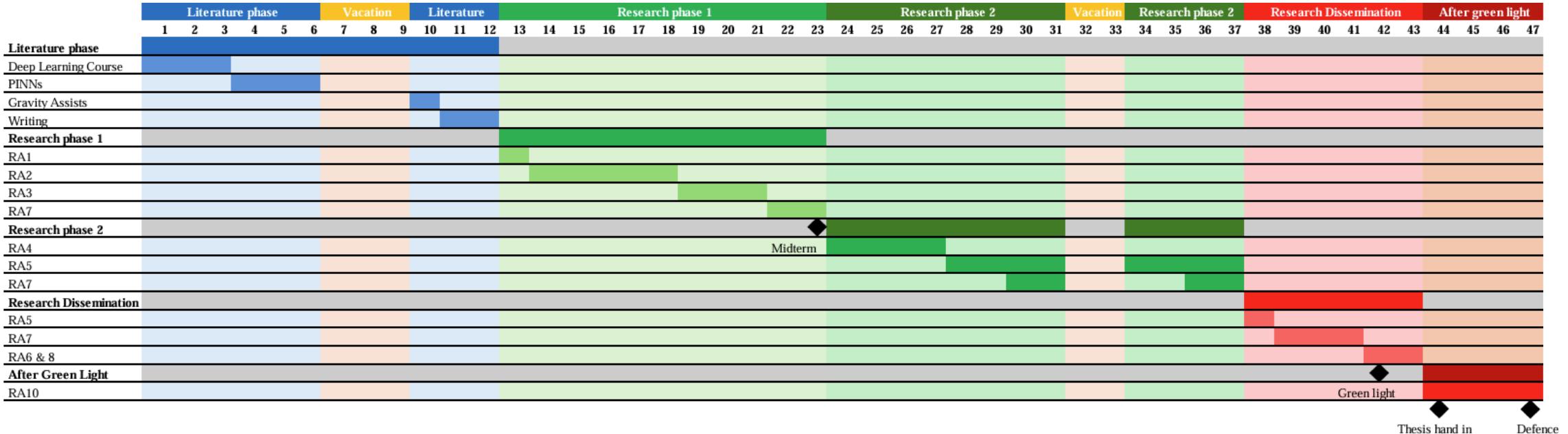
Context

This example illustrates the usage of PyGMO to optimize a low-thrust interplanetary transfer trajectory simulated using the multiple gravity assist (MGA) module of Tudat. The low-thrust legs are simulated using hodographic shaping. The spacecraft is considered to depart from the edge of the sphere of influence (SOI) of the Earth, execute swingbys at the Mars and Earth, and arrive at the edge of Jupiter's SOI. Thus, the transfer includes 3 legs and 2 swingbys.

4

Updated plan for remainder of thesis

Updated plan for remainder of thesis



5

Q&A and Feedback

Q&A and feedback

Personal doubts/ points were I need feedback on:

1. Accuracy of solutions
2. Case scenario
3. V&V
4. Omit mass from output NN

Thank you!