

# Quarks, Leptons, and the Big Bang

## 44-1 GENERAL PROPERTIES OF ELEMENTARY PARTICLES

### Learning Objectives

After reading this module, you should be able to . . .

- 44.01** Identify that a great many different elementary particles exist or can be created and that nearly all of them are unstable.
- 44.02** For the decay of an unstable particle, apply the same decay equations as used for the radioactive decay of nuclei.
- 44.03** Identify spin as the intrinsic angular momentum of a particle.
- 44.04** Distinguish fermions from bosons, and identify which are required to obey the Pauli exclusion principle.
- 44.05** Distinguish leptons and hadrons, and then identify the two types of hadrons.
- 44.06** Distinguish particle from antiparticle, and identify that if they meet, they undergo annihilation and are transformed into photons or into other elementary particles.
- 44.07** Distinguish the strong force and the weak force.
- 44.08** To see if a given process for elementary particles is physically possible, apply the conservation laws for charge, linear momentum, spin angular momentum, and energy (including mass energy).

### Key Ideas

- The term fundamental particles refers to the basic building blocks of matter. We can divide the particles into several broad categories.
- The terms particles and antiparticles originally referred to common particles (such as the electrons, protons, and neutrons in your body) and their antiparticle counterparts (the positrons, antiprotons, and antineutrons), but for most of the rarely detected particles, the distinction between particles and antiparticles is made largely to be consistent with experimental results.
- Fermions (such as the particles in your body) obey the Pauli exclusion principle; bosons do not.

## What Is Physics?

Physicists often refer to the theories of relativity and quantum physics as “modern physics,” to distinguish them from the theories of Newtonian mechanics and Maxwellian electromagnetism, which are lumped together as “classical physics.” As the years go by, the word “modern” seems less and less appropriate for theories whose foundations were laid down in the opening years of the 20th century. After all, Einstein published his paper on the photoelectric effect and his first paper on special relativity in 1905, Bohr published his quantum model of the hydrogen atom in 1913, and Schrödinger published his matter wave equation in 1926. Nevertheless, the label of “modern physics” hangs on.

In this closing chapter we consider two lines of investigation that are truly “modern” but at the same time have the most ancient of roots. They center around two deceptively simple questions:

*What is the universe made of?*

*How did the universe come to be the way it is?*

Progress in answering these questions has been rapid in the last few decades.

Many new insights are based on experiments carried out with large particle accelerators. However, as they bang particles together at higher and higher

energies using larger and larger accelerators, physicists come to realize that no conceivable Earth-bound accelerator can generate particles with energies great enough to test the ultimate theories of physics. There has been only one source of particles with these energies, and that was the universe itself within the first millisecond of its existence.

In this chapter you will encounter a host of new terms and a veritable flood of particles with names that you should not try to remember. If you are temporarily bewildered, you are sharing the bewilderment of the physicists who lived through these developments and who at times saw nothing but increasing complexity with little hope of understanding. If you stick with it, however, you will come to share the excitement physicists felt as marvelous new accelerators poured out new results, as the theorists put forth ideas each more daring than the last, and as clarity finally sprang from obscurity. The main message of this book is that, although we know a lot about the physics of the world, grand mysteries remain.

## Particles, Particles, Particles

In the 1930s, there were many scientists who thought that the problem of the ultimate structure of matter was well on the way to being solved. The atom could be understood in terms of only three particles—the electron, the proton, and the neutron. Quantum physics accounted well for the structure of the atom and for radioactive alpha decay. The neutrino had been postulated and, although not yet observed, had been incorporated by Enrico Fermi into a successful theory of beta decay. There was hope that quantum theory applied to protons and neutrons would soon account for the structure of the nucleus. What else was there?

The euphoria did not last. The end of that same decade saw the beginning of a period of discovery of new particles that continues to this day. The new particles have names and symbols such as *muon* ( $\mu$ ), *pion* ( $\pi$ ), *kaon* (K), and *sigma* ( $\Sigma$ ). All the new particles are unstable; that is, they spontaneously transform into other types of particles according to the same functions of time that apply to unstable nuclei. Thus, if  $N_0$  particles of any one type are present in a sample at time  $t = 0$ , the number  $N$  of those particles present at some later time  $t$  is given by Eq. 42-15,

$$N = N_0 e^{-\lambda t}, \quad (44-1)$$

the rate of decay  $R$ , from an initial value of  $R_0$ , is given by Eq. 42-16,

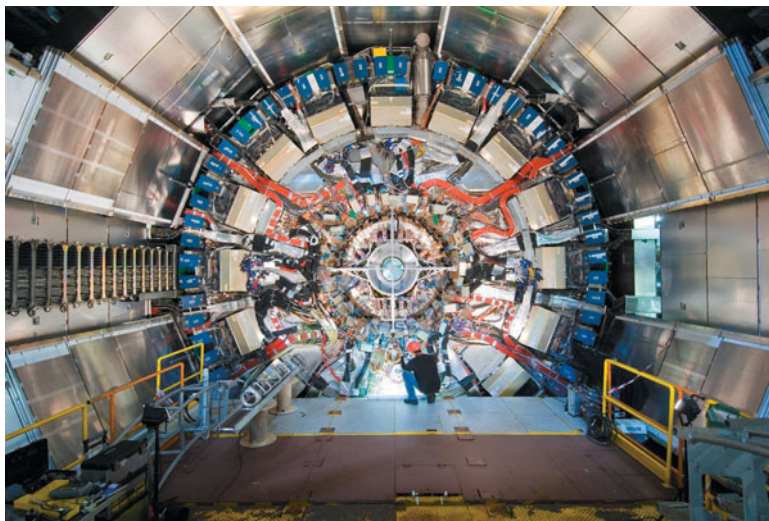
$$R = R_0 e^{-\lambda t}, \quad (44-2)$$

and the half-life  $T_{1/2}$ , decay constant  $\lambda$ , and mean life  $\tau$  are related by Eq. 42-18,

$$T_{1/2} = \frac{\ln 2}{\lambda} = \tau \ln 2. \quad (44-3)$$

The half-lives of the new particles range from about  $10^{-6}$  s to  $10^{-23}$  s. Indeed, some of the particles last so briefly that they cannot be detected directly but can only be inferred from indirect evidence.

These new particles have been commonly produced in head-on collisions between protons or electrons accelerated to high energies in accelerators at places like Brookhaven National Laboratory (on Long Island, New York), Fermilab (near Chicago), CERN (near Geneva, Switzerland), SLAC (at Stanford University in California), and DESY (near Hamburg, Germany). They have been discovered with particle detectors that have grown in sophistication until they rival the size and complexity of entire accelerators of only a few decades ago.



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One of the detectors at the Large Hadron Collider at CERN, where the Standard Model of the elementary particles is being put to the test.

Today there are several hundred known particles. Naming them has strained the resources of the Greek alphabet, and most are known only by an assigned number in a periodically issued compilation. To make sense of this array of particles, we look for simple physical criteria by which we can place the particles in categories. The result is known as the **Standard Model** of particles. Although this model is continuously challenged by theorists, it remains our best scheme of understanding all the particles discovered to date.

To explore the Standard Model, we make the following three rough cuts among the known particles: fermion or boson, hadron or lepton, particle or antiparticle? Let's now look at the categories one by one.

### Fermion or Boson?

All particles have an intrinsic angular momentum called **spin**, as we discussed for electrons, protons, and neutrons in Module 32-5. Generalizing the notation of that section, we can write the component of spin  $\vec{S}$  in any direction (assume the component to be along a  $z$  axis) as

$$S_z = m_s \hbar \quad \text{for } m_s = s, s - 1, \dots, -s, \quad (44-4)$$

in which  $\hbar$  is  $h/2\pi$ ,  $m_s$  is the *spin magnetic quantum number*, and  $s$  is the *spin quantum number*. This last can have either positive half-integer values ( $\frac{1}{2}, \frac{3}{2}, \dots$ ) or nonnegative integer values ( $0, 1, 2, \dots$ ). For example, an electron has the value  $s = \frac{1}{2}$ . Hence the spin of an electron (measured along any direction, such as the  $z$  direction) can have the values

$$S_z = \frac{1}{2}\hbar \quad (\text{spin up})$$

or

$$S_z = -\frac{1}{2}\hbar \quad (\text{spin down}).$$

Confusingly, the term *spin* is used in two ways: It properly means a particle's intrinsic angular momentum  $\vec{S}$ , but it is often used loosely to mean the particle's spin quantum number  $s$ . In the latter case, for example, an electron is said to be a spin- $\frac{1}{2}$  particle.

Particles with half-integer spin quantum numbers (like electrons) are called **fermions**, after Fermi, who (simultaneously with Paul Dirac) discovered the statistical laws that govern their behavior. Like electrons, protons and neutrons also have  $s = \frac{1}{2}$  and are fermions.

Particles with zero or integer spin quantum numbers are called **bosons**, after Indian physicist Satyendra Nath Bose, who (simultaneously with Albert Einstein) discovered the governing statistical laws for *those* particles. Photons, which have  $s = 1$ , are bosons; you will soon meet other particles in this class.

This may seem a trivial way to classify particles, but it is very important for this reason:

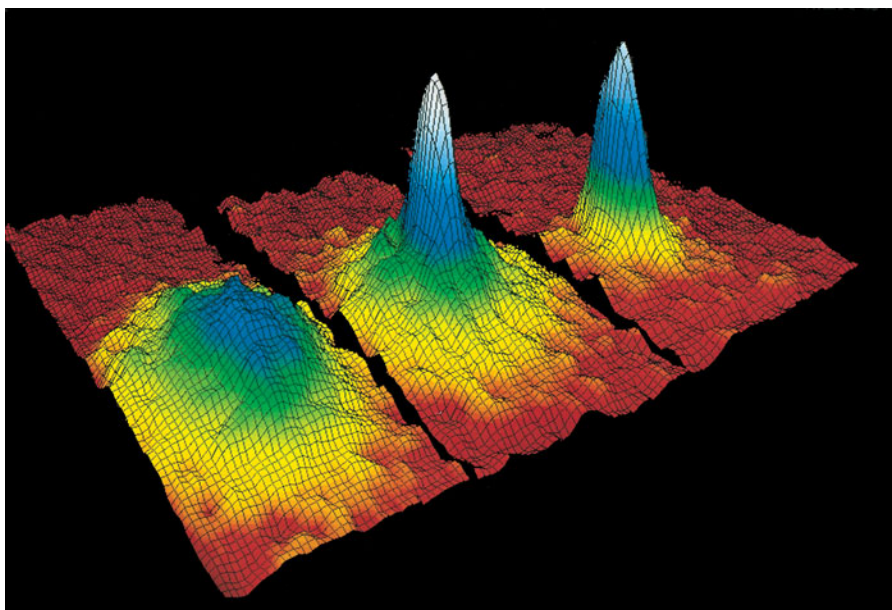


Fermions obey the Pauli exclusion principle, which asserts that only a single particle can be assigned to a given quantum state. Bosons *do not* obey this principle. Any number of bosons can occupy a given quantum state.

We saw how important the Pauli exclusion principle is when we “built up” the atoms by assigning (spin- $\frac{1}{2}$ ) electrons to individual quantum states. Using that principle led to a full accounting of the structure and properties of atoms of different types and of solids such as metals and semiconductors.

Because bosons do *not* obey the Pauli principle, those particles tend to pile up in the quantum state of lowest energy. In 1995 a group in Boulder, Colorado, succeeded in producing a condensate of about 2000 rubidium-87 atoms—they are bosons—in a single quantum state of approximately zero energy.

For this to happen, the rubidium has to be a vapor with a temperature so low and a density so great that the de Broglie wavelengths of the individual atoms are greater than the average separation between the atoms. When this condition is met, the wave functions of the individual atoms overlap and the entire assembly becomes a single quantum system (one big atom) called a *Bose–Einstein condensate*. Figure 44-1 shows that, as the temperature of the rubidium vapor is lowered to about  $1.70 \times 10^{-7}$  K, the atoms do indeed “collapse” into a single sharply defined state corresponding to approximately zero speed.



Courtesy Michael Mathews

(a)

(b)

(c)

**Figure 44-1** Three plots of the particle speed distribution in a vapor of rubidium-87 atoms. The temperature of the vapor is successively reduced from plot (a) to plot (c). Plot (c) shows a sharp peak centered around zero speed; that is, all the atoms are in the same quantum state. The achievement of such a Bose–Einstein condensate, often called the Holy Grail of atomic physics, was finally recorded in 1995.



### Hadron or Lepton?

We can also classify particles in terms of the four fundamental forces that act on them. The *gravitational force* acts on *all* particles, but its effects at the level of subatomic particles are so weak that we need not consider that force (at least not in today's research). The *electromagnetic force* acts on all *electrically charged* particles; its effects are well known, and we can take them into account when we need to; we largely ignore this force in this chapter.

We are left with the *strong force*, which is the force that binds nucleons together, and the *weak force*, which is involved in beta decay and similar processes. The weak force acts on all particles, the strong force only on some.

We can, then, roughly classify particles on the basis of whether the strong force acts on them. Particles on which the strong force acts are called **hadrons**. Particles on which the strong force does *not* act, leaving the weak force and the electromagnetic force as the dominant forces, are called **leptons**. Protons, neutrons, and pions are hadrons; electrons and neutrinos are leptons.

We can make a further distinction among the hadrons because some of them are bosons (we call them **mesons**); the pion is an example. The other hadrons are fermions (we call them **baryons**); the proton is an example.

### Particle or Antiparticle?

In 1928 Dirac predicted that the electron  $e^-$  should have a positively charged counterpart of the same mass and spin. The counterpart, the *positron*  $e^+$ , was discovered in cosmic radiation in 1932 by Carl Anderson. Physicists then gradually realized that *every* particle has a corresponding **antiparticle**. The members of such pairs have the same mass and spin but opposite signs of electric charge (if they are charged) and opposite signs of quantum numbers that we have not yet discussed.

At first, *particle* was used to refer to the common particles such as electrons, protons, and neutrons, and *antiparticle* referred to their rarely detected counterparts. Later, for the less common particles, the assignment of *particle* and *antiparticle* was made so as to be consistent with certain conservation laws that we shall discuss later in this chapter. (Confusingly, both particles and antiparticles are sometimes called particles when no distinction is needed.) We often, but not always, represent an antiparticle by putting a bar over the symbol for the particle. Thus,  $p$  is the symbol for the proton, and  $\bar{p}$  (pronounced “p bar”) is the symbol for the antiproton.

**Annihilation.** When a particle meets its antiparticle, the two can *annihilate* each other. That is, the particle and antiparticle disappear and their combined energies reappear in other forms. For an electron annihilating with a positron, this energy reappears as two gamma-ray photons:

$$e^- + e^+ \rightarrow \gamma + \gamma. \quad (44-5)$$

If the electron and positron are stationary when they annihilate, their total energy is their total mass energy, and that energy is then shared equally by the two photons. To conserve momentum and because photons cannot be stationary, the photons fly off in opposite directions.

Antihydrogen atoms (each with an antiproton and positron instead of a proton and electron in a hydrogen atom) are now being manufactured and studied at CERN. The Standard Model predicts that a transition in an antihydrogen atom (say, between the first excited state and the ground state) is identical to the same transition in a hydrogen atom. Thus, any difference in the transitions would clearly signal that the Standard Model is erroneous; no difference has yet been spotted.

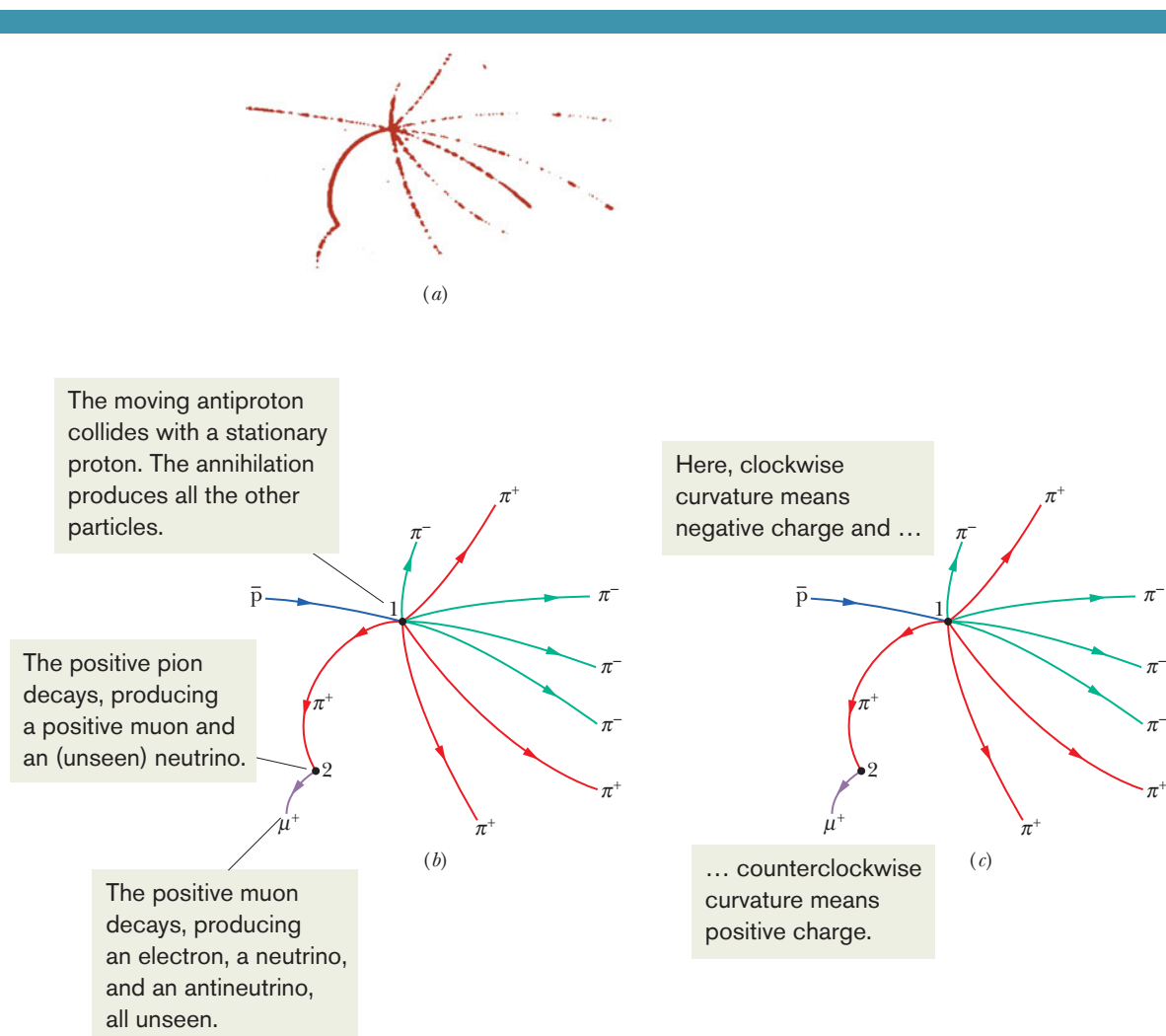
An assembly of antiparticles, such as an antihydrogen atom, is often called *antimatter* to distinguish it from an assembly of common particles (*matter*). (The terms can easily be confusing when the word “matter” is used to describe anything that has mass.) We can speculate that future scientists and engineers may construct objects of antimatter. However, no evidence suggests that nature has

already done this on an astronomical scale because all stars and galaxies appear to consist largely of matter and not antimatter. This is a perplexing observation because it means that when the universe began, some feature biased the conditions toward matter and away from antimatter. (For example, electrons are common but positrons are not.) This bias is still not well understood.

## An Interlude

Before pressing on with the task of classifying the particles, let us step aside for a moment and capture some of the spirit of particle research by analyzing a typical particle event—namely, that shown in the bubble-chamber photograph of Fig. 44-2a.

The tracks in this figure consist of bubbles formed along the paths of electrically charged particles as they move through a chamber filled with liquid hydrogen. We can identify the particle that makes a particular track by—among other means—measuring the relative spacing between the bubbles. The chamber lies in a uniform magnetic field that deflects the tracks of positively



Part (a): Courtesy Lawrence Berkeley Laboratory

**Figure 44-2** (a) A bubble-chamber photograph of a series of events initiated by an antiproton that enters the chamber from the left. (b) The tracks redrawn and labeled for clarity. (c) The tracks are curved because a magnetic field present in the chamber exerts a deflecting force on each moving charged particle.

Table 44-1 The Particles or Antiparticles Involved in the Event of Fig. 44-2

Particle	Symbol	Charge $q$	Mass (MeV/ $c^2$ )	Spin Quantum Number $s$	Identity	Mean Life (s)	Antiparticle
Neutrino	$\nu$	0	$\approx 1 \times 10^{-7}$	$\frac{1}{2}$	Lepton	Stable	$\bar{\nu}$
Electron	$e^-$	-1	0.511	$\frac{1}{2}$	Lepton	Stable	$e^+$
Muon	$\mu^-$	-1	105.7	$\frac{1}{2}$	Lepton	$2.2 \times 10^{-6}$	$\mu^+$
Pion	$\pi^+$	+1	139.6	0	Meson	$2.6 \times 10^{-8}$	$\pi^-$
Proton	p	+1	938.3	$\frac{1}{2}$	Baryon	Stable	$\bar{p}$

charged particles counterclockwise and the tracks of negatively charged particles clockwise. By measuring the radius of curvature of a track, we can calculate the momentum of the particle that made it. Table 44-1 shows some properties of the particles and antiparticles that participated in the event of Fig. 44-2a, including those that did not make tracks. Following common practice, we express the masses of the particles listed in Table 44-1—and in all other tables in this chapter—in the unit MeV/ $c^2$ . The reason for this notation is that the rest energy of a particle is needed more often than its mass. Thus, the mass of a proton is shown in Table 44-1 to be 938.3 MeV/ $c^2$ . To find the proton’s rest energy, multiply this mass by  $c^2$  to obtain 938.3 MeV.

The general tools used for the analysis of photographs like Fig. 44-2a are the laws of conservation of energy, linear momentum, angular momentum, and electric charge, along with other conservation laws that we have not yet discussed. Figure 44-2a is actually one of a stereo pair of photographs so that, in practice, these analyses are carried out in three dimensions.

The event of Fig. 44-2a is triggered by an energetic antiproton ( $\bar{p}$ ) that, generated in an accelerator at the Lawrence Berkeley Laboratory, enters the chamber from the left. There are three separate subevents; one occurs at point 1 in Fig. 44-2b, the second occurs at point 2, and the third occurs out of the frame of the figure. Let’s examine each:

1. *Proton–Antiproton Annihilation.* At point 1 in Fig. 44-2b, the initiating antiproton (blue track) slams into a proton of the liquid hydrogen in the chamber, and the result is mutual annihilation. We can tell that annihilation occurred while the incoming antiproton was in flight because most of the particles generated in the encounter move in the forward direction—that is, toward the right in Fig. 44-2. From the principle of conservation of linear momentum, the incoming antiproton must have had a forward momentum when it underwent annihilation. Further, because the particles are charged and moving through a magnetic field, the curvature of the paths reveal whether the particles are negatively charged (like the incident antiproton) or positively charged (Fig. 44-2c).

The total energy involved in the collision of the antiproton and the proton is the sum of the antiproton’s kinetic energy and the two (identical) rest energies of those two particles ( $2 \times 938.3$  MeV, or 1876.6 MeV). This is enough energy to create a number of lighter particles and give them kinetic energy. In this case, the annihilation produces four positive pions (red tracks in Fig. 44-2b) and four negative pions (green tracks). (For simplicity, we assume that no gamma-ray photons, which would leave no tracks because they lack electric charge, are produced.) Thus we conclude that the annihilation process is

$$p + \bar{p} \rightarrow 4\pi^+ + 4\pi^- \tag{44-6}$$

We see from Table 44-1 that the positive pions ( $\pi^+$ ) are *particles* and the negative pions ( $\pi^-$ ) are *antiparticles*. The reaction of Eq. 44-6 is a *strong interaction*.

tion (it involves the strong force) because all the particles involved are hadrons.

Let us check whether electric charge is conserved in the reaction. To do so, we can write the electric charge of a particle as  $qe$ , in which  $q$  is a **charge quantum number**. Then determining whether electric charge is conserved in a process amounts to determining whether the initial net charge quantum number is equal to the final net charge quantum number. In the process of Eq. 44-6, the initial net charge number is  $1 + (-1)$ , or 0, and the final net charge number is  $4(1) + 4(-1)$ , or 0. Thus, charge is conserved.

For the energy balance, note from above that the energy available from the  $p\bar{p}$  annihilation process is at least the sum of the proton and antiproton rest energies, 1876.6 MeV. The rest energy of a pion is 139.6 MeV, which means the rest energies of the eight pions amount to  $8 \times 139.6$  MeV, or 1116.8 MeV. This leaves at least about 760 MeV to distribute among the eight pions as kinetic energy. Thus, the requirement of energy conservation is easily met.

2. **Pion Decay.** Pions are unstable particles and decay with a mean lifetime of  $2.6 \times 10^{-8}$  s. At point 2 in Fig. 44-2b, one of the positive pions comes to rest in the chamber and decays spontaneously into an antimuon  $\mu^+$  (purple track) and a neutrino  $\nu$ :

$$\pi^+ \rightarrow \mu^+ + \nu. \quad (44-7)$$

The neutrino, being uncharged, leaves no track. Both the antimuon and the neutrino are leptons; that is, they are particles on which the strong force does not act. Thus, the decay process of Eq. 44-7, which is governed by the weak force, is described as a *weak interaction*.

Let's consider the energies in the decay. From Table 44-1, the rest energy of an antimuon is 105.7 MeV and the rest energy of a neutrino is approximately 0. Because the pion is at rest when it decays, its energy is just its rest energy, 139.6 MeV. Thus, an energy of  $139.6 \text{ MeV} - 105.7 \text{ MeV}$ , or 33.9 MeV, is available to share between the antimuon and the neutrino as kinetic energy.

Let us check whether spin angular momentum is conserved in the process of Eq. 44-7. This amounts to determining whether the net component  $S_z$  of spin angular momentum along some arbitrary  $z$  axis can be conserved by the process. The spin quantum numbers  $s$  of the particles in the process are 0 for the pion  $\pi^+$  and  $\frac{1}{2}$  for both the antimuon  $\mu^+$  and the neutrino  $\nu$ . Thus, for  $\pi^+$ , the component  $S_z$  must be  $0\hbar$ , and for  $\mu^+$  and  $\nu$ , it can be either  $+\frac{1}{2}\hbar$  or  $-\frac{1}{2}\hbar$ .

The net component  $S_z$  is conserved by the process of Eq. 44-7 if there is *any* way in which the initial  $S_z$  ( $= 0\hbar$ ) can be equal to the final net  $S_z$ . We see that if one of the products, either  $\mu^+$  or  $\nu$ , has  $S_z = +\frac{1}{2}\hbar$  and the other has  $S_z = -\frac{1}{2}\hbar$ , then their final net value is  $0\hbar$ . Thus, because  $S_z$  can be conserved, the decay process of Eq. 44-7 *can* occur.

From Eq. 44-7, we also see that the net charge is conserved by the process: before the process the net charge quantum number is  $+1$ , and after the process it is  $+1 + 0 = +1$ .

3. **Muon Decay.** Muons (whether  $\mu^-$  or  $\mu^+$ ) are also unstable, decaying with a mean life of  $2.2 \times 10^{-6}$  s. Although the decay products are not shown in Fig. 44-2, the antimuon produced in the reaction of Eq. 44-7 comes to rest and decays spontaneously according to

$$\mu^+ \rightarrow e^+ + \nu + \bar{\nu}. \quad (44-8)$$

The rest energy of the antimuon is 105.7 MeV, and that of the positron is only 0.511 MeV, leaving 105.2 MeV to be shared as kinetic energy among the three particles produced in the decay process of Eq. 44-8.

You may wonder: Why *two* neutrinos in Eq. 44-8? Why not just one, as in the pion decay in Eq. 44-7? One answer is that the spin quantum numbers of the



antimuon, the positron, and the neutrino are each  $\frac{1}{2}$ ; with only one neutrino, the net component  $S_z$  of spin angular momentum could not be conserved in the antimuon decay of Eq. 44-8. In Module 44-2 we shall discuss another reason.



### Sample Problem 44.01 Momentum and kinetic energy in a pion decay

A stationary positive pion can decay according to

$$\pi^+ \rightarrow \mu^+ + \nu.$$

What is the kinetic energy of the antimuon  $\mu^+$ ? What is the kinetic energy of the neutrino?

#### KEY IDEA

The pion decay process must conserve both total energy and total linear momentum.

**Energy conservation:** Let us first write the conservation of total energy (rest energy  $mc^2$  plus kinetic energy  $K$ ) for the decay process as

$$m_\pi c^2 + K_\pi = m_\mu c^2 + K_\mu + m_\nu c^2 + K_\nu.$$

Because the pion was stationary, its kinetic energy  $K_\pi$  is zero. Then, using the masses listed for  $m_\pi$ ,  $m_\mu$ , and  $m_\nu$  in Table 44-1, we find

$$\begin{aligned} K_\mu + K_\nu &= m_\pi c^2 - m_\mu c^2 - m_\nu c^2 \\ &= 139.6 \text{ MeV} - 105.7 \text{ MeV} - 0 \\ &= 33.9 \text{ MeV}, \end{aligned} \quad (44-9)$$

where we have approximated  $m_\nu$  as zero.

**Momentum conservation:** We cannot solve Eq. 44-9 for either  $K_\mu$  or  $K_\nu$  separately, and so let us next apply the principle of conservation of linear momentum to the decay process. Because the pion is stationary when it decays, that principle requires that the muon and neutrino move in opposite directions after the decay. Assume that their motion is along an axis. Then, for components along that axis, we can write the conservation of linear momentum for the decay as

$$p_\pi = p_\mu + p_\nu,$$

which, with  $p_\pi = 0$ , gives us

$$p_\mu = -p_\nu. \quad (44-10)$$

**Relating  $p$  and  $K$ :** We want to relate these momenta  $p_\mu$  and  $-p_\nu$  to the kinetic energies  $K_\mu$  and  $K_\nu$  so that we can solve for the kinetic energies. Because we have no reason to believe that classical physics can be applied, we use Eq. 37-54, the momentum–kinetic energy relation from special relativity:

$$(pc)^2 = K^2 + 2Kmc^2. \quad (44-11)$$

From Eq. 44-10, we know that

$$(p_\mu c)^2 = (p_\nu c)^2. \quad (44-12)$$

Substituting from Eq. 44-11 for each side of Eq. 44-12 yields

$$K_\mu^2 + 2K_\mu m_\mu c^2 = K_\nu^2 + 2K_\nu m_\nu c^2.$$

Approximating the neutrino mass to be  $m_\nu = 0$ , substituting  $K_\nu = 33.9 \text{ MeV} - K_\mu$  from Eq. 44-9, and then solving for  $K_\mu$ , we find

$$\begin{aligned} K_\mu &= \frac{(33.9 \text{ MeV})^2}{(2)(33.9 \text{ MeV} + m_\mu c^2)} \\ &= \frac{(33.9 \text{ MeV})^2}{(2)(33.9 \text{ MeV} + 105.7 \text{ MeV})} \\ &= 4.12 \text{ MeV}. \end{aligned} \quad (\text{Answer})$$

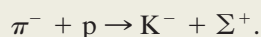
The kinetic energy of the neutrino is then, from Eq. 44-9,

$$\begin{aligned} K_\nu &= 33.9 \text{ MeV} - K_\mu = 33.9 \text{ MeV} - 4.12 \text{ MeV} \\ &= 29.8 \text{ MeV}. \end{aligned} \quad (\text{Answer})$$

We see that, although the magnitudes of the momenta of the two recoiling particles are the same, the neutrino gets the larger share (88%) of the kinetic energy.

### Sample Problem 44.02 $Q$ in a proton-pion reaction

The protons in the material filling a bubble chamber are bombarded with a beam of high-energy antiparticles known as negative pions. At collision points, a proton and a pion transform into a negative kaon and a positive sigma in this reaction:



The rest energies of these particles are

$$\begin{array}{llll} \pi^- & 139.6 \text{ MeV} & K^- & 493.7 \text{ MeV} \\ p & 938.3 \text{ MeV} & \Sigma^+ & 1189.4 \text{ MeV} \end{array}$$

What is the  $Q$  of the reaction?

#### KEY IDEA

The  $Q$  of a reaction is

$$Q = \left( \begin{array}{c} \text{initial total} \\ \text{mass energy} \end{array} \right) - \left( \begin{array}{c} \text{final total} \\ \text{mass energy} \end{array} \right).$$

**Calculation:** For the given reaction, we find

$$\begin{aligned} Q &= (m_\pi c^2 + m_p c^2) - (m_K c^2 + m_\Sigma c^2) \\ &= (139.6 \text{ MeV} + 938.3 \text{ MeV}) \\ &\quad - (493.7 \text{ MeV} + 1189.4 \text{ MeV}) \\ &= -605 \text{ MeV}. \end{aligned} \quad (\text{Answer})$$

The minus sign means that the reaction is *endothermic*; that is, the incoming pion ( $\pi^-$ ) must have a kinetic energy greater than a certain threshold value if the reaction is to occur. The threshold energy is actually greater than 605 MeV because linear momentum must be conserved. (The incoming pion

has momentum.) This means that the kaon ( $K^-$ ) and the sigma ( $\Sigma^+$ ) not only must be created but also must be given some kinetic energy. A relativistic calculation whose details are beyond our scope shows that the threshold energy for the reaction is 907 MeV.



Additional examples, video, and practice available at WileyPLUS



## 44-2 LEPTONS, HADRONS, AND STRANGENESS

### Learning Objectives

After reading this module, you should be able to . . .

- 44.09** Identify that there are six leptons (with an antiparticle each) in three families, with a different type of neutrino in each family.
- 44.10** To see if a given process for elementary particles is physically possible, determine whether it conserves lepton number and whether it conserves the individual family lepton numbers.
- 44.11** Identify that there is a quantum number called baryon number associated with the baryons.
- 44.12** To see if a given process for elementary particles is physically possible, determine whether the process conserves baryon number.
- 44.13** Identify that there is a quantum number called strangeness associated with some of the baryons and mesons.
- 44.14** Identify that strangeness must be conserved in an interaction involving the strong force, but this conservation law can be broken for other interactions.
- 44.15** Describe the eightfold-way patterns.

### Key Ideas

- We can classify particles and their antiparticles into two main types: leptons and hadrons. The latter consists of mesons and baryons.
- Three of the leptons (the electron, muon, and tau) have electric charge equal to  $-1e$ . There are also three uncharged neutrinos (also leptons), one corresponding to each of the charged leptons. The antiparticles for the charged leptons have positive charge.
- To explain the possible and impossible reactions of these particles, each is assigned a lepton quantum number, which must be conserved in a reaction.
- The leptons have half-integer spin quantum numbers and are thus fermions, which obey the Pauli exclusion principle.
- Baryons, including protons and neutrons, are hadrons with half-integer spin quantum numbers and thus are also fermions.
- Mesons are hadrons with integer spin quantum numbers and thus are bosons, which do not obey the Pauli exclusion principle.
- To explain the possible and impossible reactions of these particles, baryons are assigned a baryon quantum number, which must be conserved in a reaction.
- Baryons are also assigned a strangeness quantum number, but it is conserved only in reactions involving the strong force.

### The Leptons

In this module, we discuss some of the particles of one of our classification schemes: lepton or hadron. We begin with the leptons, those particles on which the strong force does *not* act. So far, we have encountered the familiar electron and the neutrino that accompanies it in beta decay. The muon, whose decay is described in Eq. 44-8, is another member of this family. Physicists gradually learned that the neutrino that appears in Eq. 44-7, associated with the production of a muon, is *not the same particle* as the neutrino produced in beta decay, associated with the appearance of an electron. We call the former the **muon neutrino** (symbol  $\nu_\mu$ ) and the latter the **electron neutrino** (symbol  $\nu_e$ ) when it is necessary to distinguish between them.

Table 44-2 The Leptons<sup>a</sup>

Family	Particle	Symbol	Mass (MeV/c <sup>2</sup> )	Charge <i>q</i>	Antiparticle
Electron	Electron	e <sup>−</sup>	0.511	−1	e <sup>+</sup>
	Electron neutrino <sup>b</sup>	ν <sub>e</sub>	≈ 1 × 10 <sup>−7</sup>	0	$\bar{\nu}_e$
Muon	Muon	μ <sup>−</sup>	105.7	−1	μ <sup>+</sup>
	Muon neutrino <sup>b</sup>	ν <sub>μ</sub>	≈ 1 × 10 <sup>−7</sup>	0	$\bar{\nu}_\mu$
Tau	Tau	τ <sup>−</sup>	1777	−1	τ <sup>+</sup>
	Tau neutrino <sup>b</sup>	ν <sub>τ</sub>	≈ 1 × 10 <sup>−7</sup>	0	$\bar{\nu}_\tau$

<sup>a</sup>All leptons have spin quantum numbers of  $\frac{1}{2}$  and are thus fermions.

<sup>b</sup>The neutrino masses have not been well determined. Also, because of neutrino oscillations, we might not be able to associate a particular mass with a particular neutrino.

These two types of neutrino are known to be different particles because, if a beam of muon neutrinos (produced from pion decay as in Eq. 44-7) strikes a solid target, *only muons*—and never electrons—are produced. On the other hand, if electron neutrinos (produced by the beta decay of fission products in a nuclear reactor) strike a solid target, *only electrons*—and never muons—are produced.

Another lepton, the **tau**, was discovered at SLAC in 1975; its discoverer, Martin Perl, shared the 1995 Nobel Prize in physics. The tau has its own associated neutrino, different still from the other two. Table 44-2 lists all the leptons (both particles and antiparticles); all have a spin quantum number *s* of  $\frac{1}{2}$ .

There are reasons for dividing the leptons into three families, each consisting of a particle (electron, muon, or tau), its associated neutrino, and the corresponding antiparticles. Furthermore, there are reasons to believe that there are *only* the three families of leptons shown in Table 44-2. Leptons have no internal structure and no measurable dimensions; they are believed to be truly pointlike fundamental particles when they interact with other particles or with electromagnetic waves.

The Conservation of Lepton Number

According to experiment, particle interactions involving leptons obey a conservation law for a quantum number called the **lepton number** *L*. Each (normal) particle in Table 44-2 is assigned *L* = +1, and each antiparticle is assigned *L* = −1. All other particles, which are not leptons, are assigned *L* = 0. Also according to experiment,



In all particle interactions, the net lepton number is conserved.

This experimental fact is called the law of **conservation of lepton number**. We do not know *why* the law must be obeyed; we only know that this conservation law is part of the way our universe works.

There are actually three types of lepton number, one for each lepton family: the electron lepton number *L<sub>e</sub>*, the muon lepton number *L<sub>μ</sub>*, and the tau lepton number *L<sub>τ</sub>*. In nearly all observed interactions, these three quantum numbers are separately conserved. An important exception involves the neutrinos. For reasons that we cannot explore here, the fact that neutrinos are not massless means that they can “oscillate” between different types as they travel long distances. Such oscillations were proposed to explain why only about a third of the expected number of electron neutrinos arrive at Earth from the proton-proton fusion mechanism in the Sun (Fig. 43-11). The rest change on the way. The oscilla-

tions, then, mean that the individual family lepton numbers are not conserved for neutrinos. In this book we shall not consider such violations and shall always conserve the individual family lepton numbers.

Let's illustrate such conservation by reconsidering the antimuon decay process shown in Eq. 44-8, which we now write more fully as

$$\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu. \quad (44-13)$$

Consider this first in terms of the muon family of leptons. The  $\mu^+$  is an antiparticle (see Table 44-2) and thus has the muon lepton number  $L_\mu = -1$ . The two particles  $e^+$  and  $\nu_e$  do not belong to the muon family and thus have  $L_\mu = 0$ . This leaves  $\bar{\nu}_\mu$  on the right which, being an antiparticle, also has the muon lepton number  $L_\mu = -1$ . Thus, both sides of Eq. 44-13 have the same net muon lepton number—namely,  $L_\mu = -1$ ; if they did not, the  $\mu^+$  would not decay by this process.

No members of the electron family appear on the left in Eq. 44-13; so there the net electron lepton number must be  $L_e = 0$ . On the right side of Eq. 44-13, the positron, being an antiparticle (again see Table 44-2), has the electron lepton number  $L_e = -1$ . The electron neutrino  $\nu_e$ , being a particle, has the electron number  $L_e = +1$ . Thus, the net electron lepton number for these two particles on the right in Eq. 44-13 is also zero; the electron lepton number is also conserved in the process.

Because no members of the tau family appear on either side of Eq. 44-13, we must have  $L_\tau = 0$  on each side. Thus, each of the lepton quantum numbers  $L_\mu$ ,  $L_e$ , and  $L_\tau$  remains unchanged during the decay process of Eq. 44-13, their constant values being  $-1$ ,  $0$ , and  $0$ , respectively.



### Checkpoint 1

(a) The  $\pi^+$  meson decays by the process  $\pi^+ \rightarrow \mu^+ + \nu$ . To what lepton family does the neutrino  $\nu$  belong? (b) Is this neutrino a particle or an antiparticle? (c) What is its lepton number?

## The Hadrons

We are now ready to consider hadrons (baryons and mesons), those particles whose interactions are governed by the strong force. We start by adding another conservation law to our list: conservation of baryon number.

To develop this conservation law, let us consider the proton decay process

$$p \rightarrow e^+ + \nu_e. \quad (44-14)$$

This process *never* happens. We should be glad that it does not because otherwise all protons in the universe would gradually change into positrons, with disastrous consequences for us. Yet this decay process does not violate the conservation laws involving energy, linear momentum, or lepton number.

We account for the apparent stability of the proton—and for the absence of many other processes that might otherwise occur—by introducing a new quantum number, the **baryon number**  $B$ , and a new conservation law, the **conservation of baryon number**:



To every baryon we assign  $B = +1$ . To every antibaryon we assign  $B = -1$ . To all particles of other types we assign  $B = 0$ . A particle process cannot occur if it changes the net baryon number.

In the process of Eq. 44-14, the proton has a baryon number of  $B = +1$  and the positron and neutrino both have a baryon number of  $B = 0$ . Thus, the process does not conserve baryon number and cannot occur.

**Checkpoint 2**

This mode of decay for a neutron is *not* observed:

$$n \rightarrow p + e^-.$$

Which of the following conservation laws does this process violate: (a) energy, (b) angular momentum, (c) linear momentum, (d) charge, (e) lepton number, (f) baryon number?

The masses are  $m_n = 939.6 \text{ MeV}/c^2$ ,  $m_p = 938.3 \text{ MeV}/c^2$ , and  $m_e = 0.511 \text{ MeV}/c^2$ .

**Still Another Conservation Law**

Particles have intrinsic properties in addition to the ones we have listed so far: mass, charge, spin, lepton number, and baryon number. The first of these additional properties was discovered when researchers observed that certain new particles, such as the kaon (K) and the sigma ( $\Sigma$ ), always seemed to be produced in pairs. It seemed impossible to produce only one of them at a time. Thus, if a beam of energetic pions interacts with the protons in a bubble chamber, the reaction

$$\pi^+ + p \rightarrow K^+ + \Sigma^+ \quad (44-15)$$

often occurs. The reaction

$$\pi^+ + p \rightarrow \pi^+ + \Sigma^+, \quad (44-16)$$

which violates no conservation law known in the early days of particle physics, never occurs.

It was eventually proposed (by Murray Gell-Mann in the United States and independently by K. Nishijima in Japan) that certain particles possess a new property, called **strangeness**, with its own quantum number  $S$  and its own conservation law. (Be careful not to confuse the symbol  $S$  here with the symbol for spin.) The name *strangeness* arises from the fact that, before the identities of these particles were pinned down, they were known as “strange particles,” and the label stuck.

The proton, neutron, and pion have  $S = 0$ ; that is, they are not “strange.” It was proposed, however, that the  $K^+$  particle has strangeness  $S = +1$  and that  $\Sigma^+$  has  $S = -1$ . In the reaction of Eq. 44-15, the net strangeness is initially zero and finally zero; thus, the reaction conserves strangeness. However, in the reaction shown in Eq. 44-16, the final net strangeness is  $-1$ ; thus, that reaction does not conserve strangeness and cannot occur. Apparently, then, we must add one more conservation law to our list—the **conservation of strangeness**:



Strangeness is conserved in interactions involving the strong force.

Strange particles are produced only (rapidly) by strong interactions and only in pairs with a net strangeness of zero. They then decay (slowly) through weak interactions without conserving strangeness.

It may seem heavy-handed to invent a new property of particles just to account for a little puzzle like that posed by Eqs. 44-15 and 44-16. However, strangeness soon solved many other puzzles. Still, do not be misled by the whimsical name. Strangeness is no more mysterious a property of particles than is charge. Both are properties that particles may (or may not) have; each is described by an appropriate quantum number. Each obeys a conservation law. Still other properties of particles have been discovered and given even more whimsical names, such as *charm* and *bottomness*, but all are perfectly legitimate properties. Let us see, as an example, how the new property of strangeness “earns its keep” by leading us to uncover important regularities in the properties of the particles.



**Table 44-3** Eight Spin- $\frac{1}{2}$  Baryons

Particle	Symbol	Mass (MeV/c <sup>2</sup> )	Quantum Numbers	
			Charge $q$	Strangeness $S$
Proton	p	938.3	+1	0
Neutron	n	939.6	0	0
Lambda	$\Lambda^0$	1115.6	0	-1
Sigma	$\Sigma^+$	1189.4	+1	-1
Sigma	$\Sigma^0$	1192.5	0	-1
Sigma	$\Sigma^-$	1197.3	-1	-1
Xi	$\Xi^0$	1314.9	0	-2
Xi	$\Xi^-$	1321.3	-1	-2

**Table 44-4** Nine Spin-Zero Mesons<sup>a</sup>

Particle	Symbol	Mass (MeV/c <sup>2</sup> )	Quantum Numbers	
			Charge $q$	Strangeness $S$
Pion	$\pi^0$	135.0	0	0
Pion	$\pi^+$	139.6	+1	0
Pion	$\pi^-$	139.6	-1	0
Kaon	$K^+$	493.7	+1	+1
Kaon	$K^-$	493.7	-1	-1
Kaon	$K^0$	497.7	0	+1
Kaon	$\bar{K}^0$	497.7	0	-1
Eta	$\eta$	547.5	0	0
Eta prime	$\eta'$	957.8	0	0

<sup>a</sup>All mesons are bosons, having spins of 0, 1, 2, . . . . The ones listed here all have a spin of 0.

## The Eightfold Way

There are eight baryons—the neutron and the proton among them—that have a spin quantum number of  $\frac{1}{2}$ . Table 44-3 shows some of their other properties. Figure 44-3a shows the fascinating pattern that emerges if we plot the strangeness of these baryons against their charge quantum number, using a sloping axis for the charge quantum numbers. Six of the eight form a hexagon with the two remaining baryons at its center.

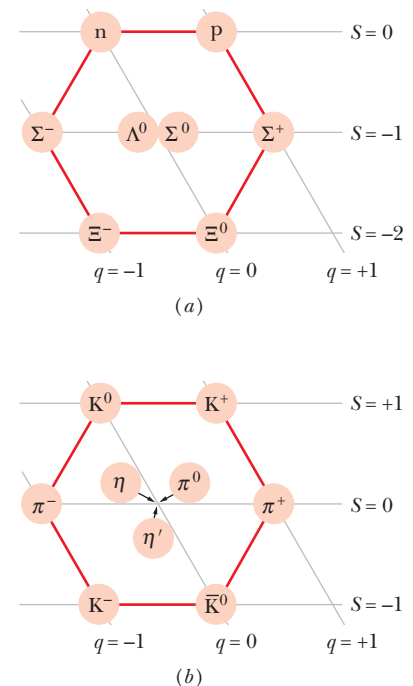
Let us turn now from the hadrons called baryons to the hadrons called mesons. Nine with a spin of zero are listed in Table 44-4. If we plot them on a sloping strangeness–charge diagram, as in Fig. 44-3b, the same fascinating pattern emerges! These and related plots, called the **eightfold way** patterns,\* were proposed independently in 1961 by Murray Gell-Mann at the California Institute of Technology and by Yuval Ne’eman at Imperial College, London. The two patterns of Fig. 44-3 are representative of a larger number of symmetrical patterns in which groups of baryons and mesons can be displayed.

The symmetry of the eightfold way pattern for the spin- $\frac{3}{2}$  baryons (not shown here) calls for ten particles arranged in a pattern like that of the tenpins in a bowling alley. However, when the pattern was first proposed, only nine such particles were known; the “headpin” was missing. In 1962, guided by theory and the symmetry of the pattern, Gell-Mann made a prediction in which he essentially said:

*There exists a spin- $\frac{3}{2}$  baryon with a charge of -1, a strangeness of -3, and a rest energy of about 1680 MeV. If you look for this omega minus particle (as I propose to call it), I think you will find it.*

A team of physicists headed by Nicholas Samios of the Brookhaven National Laboratory took up the challenge and found the “missing” particle, confirming all its predicted properties. Nothing beats prompt experimental confirmation for building confidence in a theory!

The eightfold way patterns bear the same relationship to particle physics that the periodic table does to chemistry. In each case, there is a pattern of organization in which vacancies (missing particles or missing elements) stick out like sore thumbs, guiding experimenters in their searches. In the case of the periodic table, its very existence strongly suggests that the atoms of the elements are not fundamental particles but have an underlying structure.



**Figure 44-3** (a) The eightfold way pattern for the eight spin- $\frac{1}{2}$  baryons listed in Table 44-3. The particles are represented as disks on a strangeness–charge plot, using a sloping axis for the charge quantum number. (b) A similar pattern for the nine spin-zero mesons listed in Table 44-4.

\*The name is a borrowing from Eastern mysticism. The “eight” refers to the eight quantum numbers (only a few of which we have defined here) that are involved in the symmetry-based theory that predicts the existence of the patterns.

Similarly, the eightfold way patterns strongly suggest that the mesons and the baryons must have an underlying structure, in terms of which their properties can be understood. That structure can be explained in terms of the *quark model*, which we now discuss.



### Sample Problem 44.03 Proton decay: conservation of quantum numbers, energy, and momentum

Determine whether a stationary proton can decay according to the scheme

$$p \rightarrow \pi^0 + \pi^+.$$

Properties of the proton and the  $\pi^+$  pion are listed in Table 44-1. The  $\pi^0$  pion has zero charge, zero spin, and a mass energy of 135.0 MeV.

#### KEY IDEA

We need to see whether the proposed decay violates any of the conservation laws we have discussed.

**Electric charge:** We see that the net charge quantum number is initially +1 and finally  $0 + 1$ , or +1. Thus, charge is conserved by the decay. Lepton number is also conserved, because none of the three particles is a lepton and thus each lepton number is zero.

**Linear momentum:** Because the proton is stationary, with zero linear momentum, the two pions must merely move in opposite directions with equal magnitudes of linear momentum (so that their total linear momentum is also zero) to conserve linear momentum. The fact that linear momentum *can* be conserved means that the process does not violate the conservation of linear momentum.

**Energy:** Is there energy for the decay? Because the proton is stationary, that question amounts to asking whether the proton's mass energy is sufficient to produce the mass

energies and kinetic energies of the pions. To answer, we evaluate the  $Q$  of the decay:

$$\begin{aligned} Q &= \left( \begin{array}{c} \text{initial total} \\ \text{mass energy} \end{array} \right) - \left( \begin{array}{c} \text{final total} \\ \text{mass energy} \end{array} \right) \\ &= m_p c^2 - (m_0 c^2 + m_+ c^2) \\ &= 938.3 \text{ MeV} - (135.0 \text{ MeV} + 139.6 \text{ MeV}) \\ &= 663.7 \text{ MeV}. \end{aligned}$$

The fact that  $Q$  is positive indicates that the initial mass energy exceeds the final mass energy. Thus, the proton *does* have enough mass energy to create the pair of pions.

**Spin:** Is spin angular momentum conserved by the decay? This amounts to determining whether the net component  $S_z$  of spin angular momentum along some arbitrary  $z$  axis can be conserved by the decay. The spin quantum numbers  $s$  of the particles in the process are  $\frac{1}{2}$  for the proton and 0 for both pions. Thus, for the proton the component  $S_z$  can be either  $+\frac{1}{2}\hbar$  or  $-\frac{1}{2}\hbar$  and for each pion it is  $0\hbar$ . We see that there is no way that  $S_z$  can be conserved. Hence, spin angular momentum is not conserved, and the proposed decay of the proton cannot occur.

**Baryon number:** The decay also violates the conservation of baryon number: The proton has a baryon number of  $B = +1$ , and both pions have a baryon number of  $B = 0$ . Thus, nonconservation of baryon number is another reason the proposed decay cannot occur.

### Sample Problem 44.04 Xi-minus decay: conservation of quantum numbers

A particle called xi-minus and having the symbol  $\Xi^-$  decays as follows:

$$\Xi^- \rightarrow \Lambda^0 + \pi^-.$$

The  $\Lambda^0$  particle (called lambda-zero) and the  $\pi^-$  particle are both unstable. The following decay processes occur in *cascade* until only relatively stable products remain:

$$\begin{aligned} \Lambda^0 &\rightarrow p + \pi^- & \pi^- &\rightarrow \mu^- + \bar{\nu}_\mu \\ \mu^- &\rightarrow e^- + \nu_\mu + \bar{\nu}_e. \end{aligned}$$

(a) Is the  $\Xi^-$  particle a lepton or a hadron? If the latter, is it a baryon or a meson?

#### KEY IDEAS

(1) Only three families of leptons exist (Table 44-2) and none include the  $\Xi^-$  particle. Thus, the  $\Xi^-$  must be a

hadron. (2) To answer the second question we need to determine the baryon number of the  $\Xi^-$  particle. If it is +1 or -1, then the  $\Xi^-$  is a baryon. If, instead, it is 0, then the  $\Xi^-$  is a meson.

**Baryon number:** To see, let us write the overall decay scheme, from the initial  $\Xi^-$  to the final relatively stable products, as

$$\Xi^- \rightarrow p + 2(e^- + \bar{\nu}_e) + 2(\nu_\mu + \bar{\nu}_\mu). \quad (44-17)$$

On the right side, the proton has a baryon number of +1 and each electron and neutrino has a baryon number of 0. Thus, the net baryon number of the right side is +1. That must then be the baryon number of the lone  $\Xi^-$  particle on the left side. We conclude that the  $\Xi^-$  particle is a baryon.

(b) Does the decay process conserve the three lepton numbers?

**KEY IDEA**

Any process must separately conserve the net lepton number for each lepton family of Table 44-2.

**Lepton number:** Let us first consider the electron lepton number  $L_e$ , which is  $+1$  for the electron  $e^-$ ,  $-1$  for the anti-electron neutrino  $\bar{\nu}_e$ , and  $0$  for the other particles in the overall decay of Eq. 44-17. We see that the net  $L_e$  is  $0$  before the decay and  $2[+1 + (-1)] + 2(0 + 0) = 0$  after the decay. Thus, the net electron lepton number *is* conserved. You can similarly show that the net muon lepton number and the net tau lepton number are also conserved.

(c) What can you say about the spin of the  $\Xi^-$  particle?



Additional examples, video, and practice available at WileyPLUS

**KEY IDEA**

The overall decay scheme of Eq. 44-17 must conserve the net spin component  $S_z$ .

**Spin:** We can determine the spin component  $S_z$  of the  $\Xi^-$  particle on the left side of Eq. 44-17 by considering the  $S_z$  components of the nine particles on the right side. All nine of those particles are spin- $\frac{1}{2}$  particles and thus can have  $S_z$  of either  $+\frac{1}{2}\hbar$  or  $-\frac{1}{2}\hbar$ . No matter how we choose between those two possible values of  $S_z$ , the net  $S_z$  for those nine particles must be a *half-integer* times  $\hbar$ . Thus, the  $\Xi^-$  particle must have  $S_z$  of a *half-integer* times  $\hbar$ , and that means that its spin quantum number  $s$  must be a half-integer. (It is  $\frac{1}{2}$ .)



## 44-3 QUARKS AND MESSENGER PARTICLES

### Learning Objectives

After reading this module, you should be able to . . .

**44.16** Identify that there are six quarks (with an antiparticle for each).

**44.17** Identify that baryons contain three quarks (or antiquarks) and mesons contain a quark and an antiquark, and that many of these hadrons are excited states of the basic quark combinations.

**44.18** For a given hadron, identify the quarks it contains, and vice versa.

**44.19** Identify virtual particles.

**44.20** Apply the relationship between the violation of energy by a virtual particle and the time interval allowed for that violation (an uncertainty principle written in terms of energy).

**44.21** Identify the messenger particles for electromagnetic interactions, weak interactions, and strong interactions.

### Key Ideas

- The six quarks (up, down, strange, charm, bottom, and top, in order of increasing mass) each have baryon number  $+\frac{1}{3}$  and charge equal to either  $+\frac{2}{3}$  or  $-\frac{1}{3}$ . The strange quark has strangeness  $-1$ , whereas the others all have strangeness  $0$ . These four algebraic signs are reversed for the antiquarks.

- Leptons do not contain quarks and have no internal structure. Mesons contain one quark and one antiquark. Baryons contain three quarks or antiquarks. The quantum numbers of the quarks and antiquarks are assigned to be consistent with the quantum numbers of the mesons and baryons.

- Particles with electric charge interact through the electromagnetic force by exchanging virtual photons.

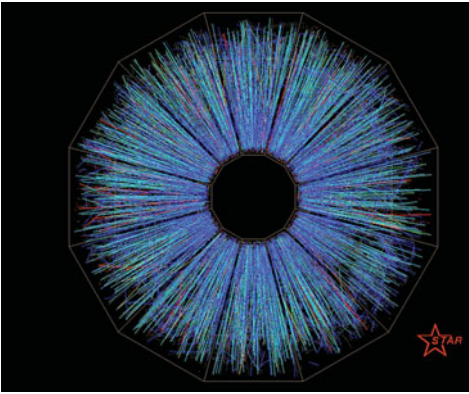
- Leptons can also interact with each other and with quarks through the weak force, via massive W and Z particles as messengers.

- Quarks primarily interact with each other through the color force, via gluons.

- The electromagnetic and weak forces are different manifestations of the same force, called the electroweak force.

### The Quark Model

In 1964 Gell-Mann and George Zweig independently pointed out that the eightfold way patterns can be understood in a simple way if the mesons and the baryons are built up out of subunits that Gell-Mann called **quarks**. We deal first with three of them, called the *up quark* (symbol  $u$ ), the *down quark* (symbol  $d$ ), and the *strange quark* (symbol  $s$ ). The names of the quarks, along with those assigned to three other quarks that we shall meet later, have no meaning other than



Courtesy Brookhaven National Laboratory  
The violent head-on collision of two 30 GeV beams of gold atoms in the RHIC accelerator at the Brookhaven National Laboratory. In the moment of collision, a gas of individual quarks and gluons was created.

Table 44-5 The Quarks<sup>a</sup>

Particle	Symbol	Mass (MeV/c <sup>2</sup> )	Quantum Numbers			Antiparticle
			Charge <i>q</i>	Strangeness <i>S</i>	Baryon Number <i>B</i>	
Up	u	5	$+\frac{2}{3}$	0	$+\frac{1}{3}$	$\bar{u}$
Down	d	10	$-\frac{1}{3}$	0	$+\frac{1}{3}$	$\bar{d}$
Charm	c	1500	$+\frac{2}{3}$	0	$+\frac{1}{3}$	$\bar{c}$
Strange	s	200	$-\frac{1}{3}$	-1	$+\frac{1}{3}$	$\bar{s}$
Top	t	175 000	$+\frac{2}{3}$	0	$+\frac{1}{3}$	$\bar{t}$
Bottom	b	4300	$-\frac{1}{3}$	0	$+\frac{1}{3}$	$\bar{b}$

<sup>a</sup>All quarks (including antiquarks) have spin  $\frac{1}{2}$  and thus are fermions. The quantum numbers *q*, *S*, and *B* for each antiquark are the negatives of those for the corresponding quark.

as convenient labels. Collectively, these names are called the *quark flavors*. We could just as well call them vanilla, chocolate, and strawberry instead of up, down, and strange. Some properties of the quarks are displayed in Table 44-5.

The fractional charge quantum numbers of the quarks may jar you a little. However, withhold judgment until you see how neatly these fractional charges account for the observed integer charges of the mesons and the baryons. In all normal situations, whether here on Earth or in an astronomical process, quarks are always bound up together in twos or threes (and perhaps more) for reasons that are still not well understood. Such requirements are our normal rule for quark combinations.

An exciting exception to the normal rule occurred in experiments at the RHIC particle collider at the Brookhaven National Laboratory. At the spot where two high-energy beams of gold nuclei collided head-on, the kinetic energy of the particles was so large that it matched the kinetic energy of particles that were present soon after the beginning of the universe (as we discuss in Module 44-4). The protons and neutrons of the gold nuclei were ripped apart to form a momentary gas of individual quarks. (The gas also contained gluons, the particles that normally hold quarks together.) These experiments at RHIC may be the first time that quarks have been set free of one another since the universe began.

### Quarks and Baryons

Each baryon is a combination of three quarks; some of the combinations are given in Fig. 44-4a. With regard to baryon number, we see that any three quarks (each with  $B = +\frac{1}{3}$ ) yield a proper baryon (with  $B = +1$ ).

Charges also work out, as we can see from three examples. The proton has a quark composition of uud, and so its charge quantum number is

$$q(uud) = \frac{2}{3} + \frac{2}{3} + \left(-\frac{1}{3}\right) = +1.$$

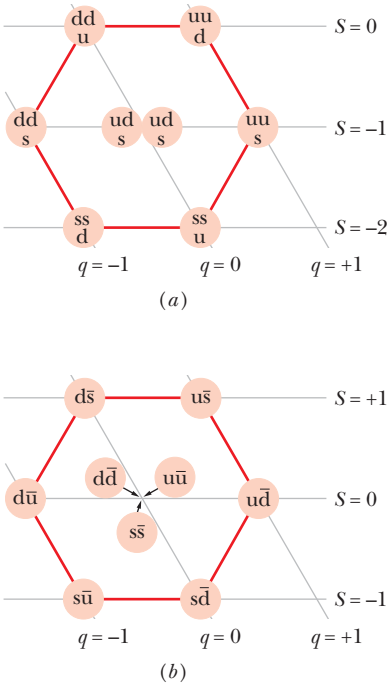
The neutron has a quark composition of udd, and its charge quantum number is therefore

$$q(udd) = \frac{2}{3} + \left(-\frac{1}{3}\right) + \left(-\frac{1}{3}\right) = 0.$$

The  $\Sigma^-$  (sigma-minus) particle has a quark composition of dds, and its charge quantum number is therefore

$$q(dds) = -\frac{1}{3} + \left(-\frac{1}{3}\right) + \left(-\frac{1}{3}\right) = -1.$$

The strangeness quantum numbers work out as well. You can check this by using Table 44-3 for the  $\Sigma^-$  strangeness number and Table 44-5 for the strangeness numbers of the dds quarks.



**Figure 44-4** (a) The quark compositions of the eight spin- $\frac{1}{2}$  baryons plotted in Fig. 44-3a. (Although the two central baryons share the same quark structure, they are different particles. The sigma is an excited state of the lambda, decaying into the lambda by emission of a gamma-ray photon.) (b) The quark compositions of the nine spin-zero mesons plotted in Fig. 44-3b.

Note, however, that the mass of a proton, neutron,  $\Sigma^-$ , or any other baryon is *not* the sum of the masses of the constituent quarks. For example, the total mass of the three quarks in a proton is only 20 MeV/c<sup>2</sup>, woefully less than the proton's mass of 938.3 MeV/c<sup>2</sup>. Nearly all of the proton's mass is due to the internal energies of (1) the quark motion and (2) the fields that bind the quarks together. (Recall that mass is related to energy via Einstein's equation, which we can write as  $m = E/c^2$ .) Thus, because most of your mass is due to the protons and neutrons in your body, your mass (and therefore your weight on a bathroom scale) is primarily a measure of the energies of the quark motion and the quark-binding fields within you.

### Quarks and Mesons

Mesons are quark–antiquark pairs; some of their compositions are given in Fig. 44-4b. The quark–antiquark model is consistent with the fact that mesons are not baryons; that is, mesons have a baryon number  $B = 0$ . The baryon number for a quark is  $+\frac{1}{3}$  and for an antiquark is  $-\frac{1}{3}$ ; thus, the combination of baryon numbers in a meson is zero.

Consider the meson  $\pi^+$ , which consists of an up quark  $u$  and an antidown quark  $\bar{d}$ . We see from Table 44-5 that the charge quantum number of the up quark is  $+\frac{2}{3}$  and that of the antidown quark is  $+\frac{1}{3}$  (the sign is opposite that of the down quark). This adds nicely to a charge quantum number of +1 for the  $\pi^+$  meson; that is,

$$q(u\bar{d}) = \frac{2}{3} + \frac{1}{3} = +1.$$

All the charge and strangeness quantum numbers of Fig. 44-4b agree with those of Table 44-4 and Fig. 44-3b. Convince yourself that all possible up, down, and strange quark–antiquark combinations are used. Everything fits.



### Checkpoint 3

Is a combination of a down quark ( $d$ ) and an antiup quark ( $\bar{u}$ ) called (a) a  $\pi^0$  meson, (b) a proton, (c) a  $\pi^-$  meson, (d) a  $\pi^+$  meson, or (e) a neutron?

### A New Look at Beta Decay

Let us see how beta decay appears from the quark point of view. In Eq. 42-24, we presented a typical example of this process:



After the neutron was discovered and Fermi had worked out his theory of beta decay, physicists came to view the fundamental beta-decay process as the changing of a neutron into a proton inside the nucleus, according to the scheme

$$n \rightarrow p + e^- + \bar{\nu}_e,$$

in which the neutrino is identified more completely. Today we look deeper and see that a neutron ( $udd$ ) can change into a proton ( $uud$ ) by changing a down quark into an up quark. We now view the fundamental beta-decay process as

$$d \rightarrow u + e^- + \bar{\nu}_e.$$

Thus, as we come to know more and more about the fundamental nature of matter, we can examine familiar processes at deeper and deeper levels. We see too that the quark model not only helps us to understand the structure of particles but also clarifies their interactions.

### Still More Quarks

There are other particles and other eightfold way patterns that we have not discussed. To account for them, it turns out that we need to postulate three more quarks, the *charm quark*  $c$ , the *top quark*  $t$ , and the *bottom quark*  $b$ . Thus, a total of six quarks exist, as listed in Table 44-5.



Note that three quarks are exceptionally massive, the most massive of them (top) being almost 190 times more massive than a proton. To generate particles that contain such quarks, with such large mass energies, we must go to higher and higher energies, which is the reason that these three quarks were not discovered earlier.

The first particle containing a charm quark to be observed was the  $J/\psi$  meson, whose quark structure is  $c\bar{c}$ . It was discovered simultaneously and independently in 1974 by groups headed by Samuel Ting at the Brookhaven National Laboratory and Burton Richter at Stanford University.

The top quark defied all efforts to generate it in the laboratory until 1995, when its existence was finally demonstrated in the Tevatron, a large particle accelerator at Fermilab. In this accelerator, protons and antiprotons, each with an energy of 0.9 TeV ( $= 9 \times 10^{11}$  eV), were made to collide at the centers of two large particle detectors. In a very few cases, the colliding particles generated a top–antitop ( $t\bar{t}$ ) quark pair, which *very* quickly decays into particles that can be detected and thus can be used to infer the existence of the top–antitop pair.

Look back for a moment at Table 44-5 (the quark family) and Table 44-2 (the lepton family) and notice the neat symmetry of these two “six-packs” of particles, each dividing naturally into three corresponding two-particle families. In terms of what we know today, the quarks and the leptons seem to be truly fundamental particles having no internal structure.



### Sample Problem 44.05 Quark composition of a xi-minus particle

The  $\Xi^-$  (xi-minus) particle is a baryon with a spin quantum number  $s$  of  $\frac{1}{2}$ , a charge quantum number  $q$  of  $-1$ , and a strangeness quantum number  $S$  of  $-2$ . Also, it does not contain a bottom quark. What combination of quarks makes up  $\Xi^-$ ?

**Reasoning:** Because the  $\Xi^-$  is a baryon, it must consist of three quarks (not two as for a meson).

Let us next consider the strangeness  $S = -2$  of the  $\Xi^-$ . Only the strange quark  $s$  and the antistrange quark  $\bar{s}$  have nonzero values of strangeness (see Table 44-5). Further, because only the strange quark  $s$  has a *negative* value of strangeness,  $\Xi^-$  must contain that quark. In fact, for  $\Xi^-$  to have a strangeness of  $-2$ , it must contain two strange quarks.

To determine the third quark, call it  $x$ , we can consider the other known properties of  $\Xi^-$ . Its charge quantum number  $q$  is  $-1$ , and the charge quantum number  $q$  of each

strange quark is  $-\frac{1}{3}$ . Thus, the third quark  $x$  must have a charge quantum number of  $-\frac{1}{3}$ , so that we can have

$$\begin{aligned} q(\Xi^-) &= q(ssx) \\ &= -\frac{1}{3} + (-\frac{1}{3}) + (-\frac{1}{3}) = -1. \end{aligned}$$

Besides the strange quark, the only quarks with  $q = -\frac{1}{3}$  are the down quark  $d$  and bottom quark  $b$ . Because the problem statement ruled out a bottom quark, the third quark must be a down quark. This conclusion is also consistent with the baryon quantum numbers:

$$\begin{aligned} B(\Xi^-) &= B(ssd) \\ &= \frac{1}{3} + \frac{1}{3} + \frac{1}{3} = +1. \end{aligned}$$

Thus, the quark composition of the  $\Xi^-$  particle is  $ssd$ .



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## The Basic Forces and Messenger Particles

We turn now from cataloging the particles to considering the forces between them.

### The Electromagnetic Force

At the atomic level, we say that two electrons exert electromagnetic forces on each other according to Coulomb's law. At a deeper level, this interaction is described by a highly successful theory called **quantum electrodynamics** (QED). From this point of view, we say that each electron senses the presence of the other by exchanging photons with it.

We cannot detect these photons because they are emitted by one electron and absorbed by the other a very short time later. Because of their undetectable existence, we call them **virtual photons**. Because they communicate between the two interacting charged particles, we sometimes call these photons *messenger particles*.

If a stationary electron emits a photon and remains itself unchanged, energy is not conserved. The principle of conservation of energy is saved, however, by an uncertainty principle written in the form

$$\Delta E \cdot \Delta t \approx \hbar. \quad (44-18)$$

Here we interpret this relation to mean that you can “overdraw” an amount of energy  $\Delta E$ , violating conservation of energy, *provided* you “return” it within an interval  $\Delta t$  given by  $\hbar/\Delta E$  so that the violation cannot be detected. The virtual photons do just that. When, say, electron *A* emits a virtual photon, the overdraft in energy is quickly set right when that electron receives a virtual photon from electron *B*, and the violation is hidden by the inherent uncertainty.

### The Weak Force

A theory of the weak force, which acts on all particles, was developed by analogy with the theory of the electromagnetic force. The messenger particles that transmit the weak force between particles, however, are not (massless) photons but massive particles, identified by the symbols *W* and *Z*. The theory was so successful that it revealed the electromagnetic force and the weak force as being different aspects of a single **electroweak force**. This accomplishment is a logical extension of the work of Maxwell, who revealed the electric and magnetic forces as being different aspects of a single *electromagnetic* force.

The electroweak theory was specific in predicting the properties of the messenger particles. In addition to the massless photon, the messenger of the electromagnetic interactions, the theory gives us three messengers for the weak interactions:

Particle	Charge	Mass
W	$\pm e$	80.4 GeV/ $c^2$
Z	0	91.2 GeV/ $c^2$

Recall that the proton mass is only 0.938 GeV/ $c^2$ ; these are massive particles! The 1979 Nobel Prize in physics was awarded to Sheldon Glashow, Steven Weinberg, and Abdus Salam for their electroweak theory. The theory was confirmed in 1983 by Carlo Rubbia and his group at CERN, and the 1984 Nobel Prize in physics went to Rubbia and Simon van der Meer for this brilliant experimental work.

Some notion of the complexity of particle physics in this day and age can be found by looking at an earlier particle physics experiment that led to the Nobel Prize in physics—the discovery of the neutron. This vitally important discovery was a “tabletop” experiment, employing particles emitted by naturally occurring radioactive materials as projectiles; it was reported in 1932 under the title “Possible Existence of a Neutron,” the single author being James Chadwick.

The discovery of the *W* and *Z* messenger particles in 1983, by contrast, was carried out at a large particle accelerator, about 7 km in circumference and operating in the range of several hundred billion electron-volts. The principal particle detector alone weighed 20 MN. The experiment employed more than 130 physicists from 12 institutions in 8 countries, along with a large support staff.

### The Strong Force

A theory of the strong force—that is, the force that acts between quarks to bind hadrons together—has also been developed. The messenger particles in this case

are called **gluons** and, like the photon, they are predicted to be massless. The theory assumes that each “flavor” of quark comes in three varieties that, for convenience, have been labeled *red*, *yellow*, and *blue*. Thus, there are three up quarks, one of each color, and so on. The antiquarks also come in three colors, which we call *antired*, *antiyellow*, and *antibblue*. You must not think that quarks are actually colored, like tiny jelly beans. The names are labels of convenience, but (for once) they do have a certain formal justification, as you will see.

The force acting between quarks is called a **color force** and the underlying theory, by analogy with quantum electrodynamics (QED), is called **quantum chromodynamics** (QCD). Apparently, quarks can be assembled only in combinations that are *color-neutral*.

There are two ways to bring about color neutrality. In the theory of actual colors, red + yellow + blue yields white, which is color-neutral, and we use the same scheme in dealing with quarks. Thus we can assemble three quarks to form a baryon, provided one is a yellow quark, one is a red quark, and one is a blue quark. Antired + antiyellow + antibblue is also white, so that we can assemble three antiquarks (of the proper anticolors) to form an antibaryon. Finally, red + antired, or yellow + antiyellow, or blue + antibblue also yields white. Thus, we can assemble a quark–antiquark combination to form a meson. The color-neutral rule does not permit any other combination of quarks, and none are observed.

The color force not only acts to bind together quarks as baryons and mesons, but it also acts between such particles, in which case it has traditionally been called the strong force. Hence, not only does the color force bind together quarks to form protons and neutrons, but it also binds together the protons and neutrons to form nuclei.

### The Higgs Field and Particle

The Standard Model of the fundamental particles consists of the theory for the electroweak interactions and the theory for the strong interactions. A key success in the model has been to demonstrate the existence of the four messenger particles in the electroweak interactions: the photon, and the Z and W particles. However, a key puzzle has involved the masses of those particles. Why is the photon massless while the Z and W particles are extremely massive?

In the 1960s, Peter Higgs and, independently, Robert Brout and François Englert suggested that the mass discrepancy is due to a field (now called the *Higgs field*) that permeates all of space and thus is a property of the vacuum. Without this field, the four messenger particles would be massless and indistinguishable—they would be *symmetric*. The Brout–Englert–Higgs theory demonstrates how the field breaks that symmetry, producing the electroweak messengers with one being massless. It also explains why all other particles, except for the gluon, have mass. The quantum of that field is the **Higgs boson**. Because of its pivotal role for all particles and because the theory behind its existence is compelling (even beautiful), intense searches for the Higgs boson were conducted on the Tevatron at Brookhaven and the Large Hadron Collider at CERN. In 2012, tantalizing experimental evidence was announced for the Higgs boson, at a mass of  $125 \text{ GeV}/c^2$ .

### Einstein’s Dream

The unification of the fundamental forces of nature into a single force—which occupied Einstein’s attention for much of his later life—is very much a current focus of research. We have seen that the weak force has been successfully combined with electromagnetism so that they may be jointly viewed as aspects of a single *electroweak force*. Theories that attempt to add the strong force to this combination—called *grand unification theories* (GUTs)—are being pursued actively. Theories that seek to complete the job by adding gravity—sometimes called *theories of everything* (TOE)—are at a speculative stage at this time. *String theory* (in which particles are tiny oscillating loops) is one approach.

## 44-4 COSMOLOGY

### Learning Objectives

After reading this module, you should be able to . . .

- 44.22** Identify that the universe (all of spacetime) began with the big bang and has been expanding ever since.
- 44.23** Identify that all distant galaxies (and thus their stars, black holes, etc.), in all directions, are receding from us because of the expansion.
- 44.24** Apply Hubble's law to relate the recession speed  $v$  of a distant galaxy, its distance  $r$  from us, and the Hubble constant  $H$ .
- 44.25** Apply the Doppler equation for the red shift of light to relate the wavelength shift  $\Delta\lambda$ , the recession speed  $v$ , and the proper wavelength  $\lambda_0$  of the emission.
- 44.26** Approximate the age of the universe using the Hubble constant.
- 44.27** Identify the cosmic background radiation and explain the importance of its detection.
- 44.28** Explain the evidence for the dark matter that apparently surrounds every galaxy.
- 44.29** Discuss the various stages of the universe from very soon after the big bang until atoms began to form.
- 44.30** Identify that the expansion of the universe is being accelerated by some unknown property dubbed dark energy.
- 44.31** Identify that the total energy of baryonic matter (protons and neutrons) is only a small part of the total energy of the universe.

### Key Ideas

- The universe is expanding, which means that empty space is continuously appearing between us and any distant galaxy.
- The rate  $v$  at which a distance to a distant galaxy is increasing (the galaxy appears to be moving at speed  $v$ ) is given by the Hubble law:

$$v = Hr,$$

where  $r$  is the current distance to the galaxy and  $H$  is the Hubble constant, which we take to be

$$H = 71.0 \text{ km/s} \cdot \text{Mpc} = 21.8 \text{ mm/s} \cdot \text{ly}.$$

- The expansion causes a red shift in the light we receive from distant galaxies. We can assume that the wavelength shift  $\Delta\lambda$  is given (approximately) by the Doppler shift equation for light discussed in Module 37-5:

$$v = \frac{|\Delta\lambda|}{\lambda_0} c,$$

where  $\lambda_0$  is the proper wavelength as measured in the frame of the light source (the galaxy).

- The expansion described by Hubble's law and the presence of ubiquitous background microwave radiation reveal that the universe began in a “big bang” 13.7 billion years ago.
- The rate of expansion is increasing due to a mysterious property of the vacuum called dark energy.
- Much of the energy of the universe is hidden in dark matter that apparently interacts with normal (baryonic) matter through the gravitational force.

### A Pause for Reflection

Let us put what you have just learned in perspective. If all we are interested in is the structure of the world around us, we can get along nicely with the electron, the neutrino, the neutron, and the proton. As someone has said, we can operate “Spaceship Earth” quite well with just these particles. We can see a few of the more exotic particles by looking for them in the cosmic rays; however, to see most of them, we must build massive accelerators and look for them at great effort and expense.

The reason we must go to such effort is that—measured in energy terms—we live in a world of very low temperatures. Even at the center of the Sun, the value of  $kT$  is only about 1 keV. To produce the exotic particles, we must be able to accelerate protons or electrons to energies in the GeV and TeV range and higher.

Once upon a time the temperature everywhere *was* high enough to provide such energies. That time of extremely high temperatures occurred in the **big bang** beginning of the universe, when the universe (and both space and time) came

into existence. Thus, one reason scientists study particles at high energies is to understand what the universe was like just after it began.

As we shall discuss shortly, *all* of space within the universe was initially tiny in extent, and the temperature of the particles within that space was incredibly high. With time, however, the universe expanded and cooled to lower temperatures, eventually to the size and temperature we see today.

Actually, the phrase “we see today” is complicated: When we look out into space, we are actually looking back in time because the light from the stars and galaxies has taken a long time to reach us. The most distant objects that we can detect are **quasars** (*quasistellar* objects), which are the extremely bright cores of galaxies that are as much as  $13 \times 10^9$  ly from us. Each such core contains a gigantic black hole; as material (gas and even stars) is pulled into one of those black holes, the material heats up and radiates a tremendous amount of light, enough for us to detect in spite of the huge distance. We therefore “see” a quasar not as it looks today but rather as it once was, when that light began its journey to us billions of years ago.

## The Universe Is Expanding

As we saw in Module 37-5, it is possible to measure the relative speeds at which galaxies are approaching us or receding from us by measuring the shifts in the wavelength of the light they emit. If we look only at distant galaxies, beyond our immediate galactic neighbors, we find an astonishing fact: They are *all* moving away (receding) from us! In 1929 Edwin P. Hubble connected the recession speed  $v$  of a galaxy and its distance  $r$  from us—they are directly proportional:

$$v = Hr \quad (\text{Hubble's law}), \quad (44-19)$$

in which  $H$  is called the **Hubble constant**. The value of  $H$  is usually measured in the unit kilometers per second-megaparsec ( $\text{km/s} \cdot \text{Mpc}$ ), where the megaparsec is a length unit commonly used in astrophysics and astronomy:

$$1 \text{ Mpc} = 3.084 \times 10^{19} \text{ km} = 3.260 \times 10^6 \text{ ly}. \quad (44-20)$$

The Hubble constant  $H$  has not had the same value since the universe began. Determining its current value is extremely difficult because doing so involves measurements of very distant galaxies. However, the Hubble constant is now known to be

$$H = 71.0 \text{ km/s} \cdot \text{Mpc} = 21.8 \text{ mm/s} \cdot \text{ly}. \quad (44-21)$$

We interpret the recession of the galaxies to mean that the universe is expanding, much as the raisins in what is to be a loaf of raisin bread grow farther apart as the dough expands. Observers on all other galaxies would find that distant galaxies were rushing away from them also, in accordance with Hubble's law. In keeping with our analogy, we can say that no raisin (galaxy) has a unique or preferred view.

Hubble's law is consistent with the hypothesis that the universe began with the big bang and has been expanding ever since. If we assume that the rate of expansion has been constant (that is, the value of  $H$  has been constant), then we can estimate the age  $T$  of the universe by using Eq. 44-19. Let us also assume that since the big bang, any given part of the universe (say, a galaxy) has been receding from our location at a speed  $v$  given by Eq. 44-19. Then the time required for the given part to recede a distance  $r$  is

$$T = \frac{r}{v} = \frac{r}{Hr} = \frac{1}{H} \quad (\text{estimated age of universe}). \quad (44-22)$$

For the value of  $H$  in Eq. 44-21,  $T$  works out to be  $13.8 \times 10^9$  y. Much more sophisticated studies of the expansion of the universe put  $T$  at  $13.7 \times 10^9$  y.





### Sample Problem 44.06 Using Hubble's law to relate distance and recessional speed

The wavelength shift in the light from a particular quasar indicates that the quasar has a recessional speed of  $2.8 \times 10^8$  m/s (which is 93% of the speed of light). Approximately how far from us is the quasar?

#### KEY IDEA

We assume that the distance and speed are related by Hubble's law.

**Calculation:** From Eqs. 44-19 and 44-21, we find

$$r = \frac{v}{H} = \frac{2.8 \times 10^8 \text{ m/s}}{21.8 \text{ mm/s} \cdot \text{ly}} (1000 \text{ mm/m}) = 12.8 \times 10^9 \text{ ly.} \quad (\text{Answer})$$

This is only an approximation because the quasar has not always been receding from our location at the same speed  $v$ ; that is,  $H$  has not had its current value throughout the time during which the universe has been expanding.

### Sample Problem 44.07 Using Hubble's law to relate distance and Doppler shift

A particular emission line detected in the light from a galaxy has a detected wavelength  $\lambda_{\text{det}} = 1.1\lambda$ , where  $\lambda$  is the proper wavelength of the line. What is the galaxy's distance from us?

#### KEY IDEAS

(1) We assume that Hubble's law ( $v = Hr$ ) applies to the recession of the galaxy. (2) We also assume that the astronomical Doppler shift of Eq. 37-36 ( $v = c|\Delta\lambda|/\lambda$ , for  $v \ll c$ ) applies to the shift in wavelength due to the recession.

**Calculations:** We can then set the right side of these two equations equal to each other to write

$$Hr = \frac{c|\Delta\lambda|}{\lambda}, \quad (44-23)$$

which leads us to

$$r = \frac{c|\Delta\lambda|}{H\lambda}. \quad (44-24)$$

In this equation,

$$\Delta\lambda = \lambda_{\text{det}} - \lambda = 1.1\lambda - \lambda = 0.1\lambda.$$

Substituting this into Eq. 44-24 then gives us

$$\begin{aligned} r &= \frac{c(0.1\lambda)}{H\lambda} = \frac{0.1c}{H} \\ &= \frac{(0.1)(3.0 \times 10^8 \text{ m/s})}{21.8 \text{ mm/s} \cdot \text{ly}} (1000 \text{ mm/m}) \\ &= 1.4 \times 10^9 \text{ ly.} \quad (\text{Answer}) \end{aligned}$$



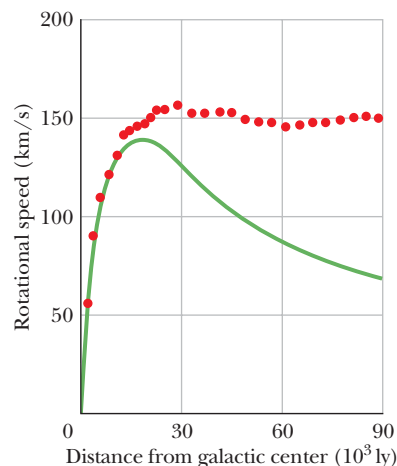
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## The Cosmic Background Radiation

In 1965 Arno Penzias and Robert Wilson, of what was then the Bell Telephone Laboratories, were testing a sensitive microwave receiver used for communications research. They discovered a faint background “hiss” that remained unchanged in intensity no matter where their antenna was pointed. It soon became clear that Penzias and Wilson were observing a **cosmic background radiation**, generated in the early universe and filling all space almost uniformly. Currently this radiation has a maximum intensity at a wavelength of 1.1 mm, which lies in the microwave region of electromagnetic radiation (or light, for short). The wavelength distribution of this radiation matches the wavelength distribution of light that would be emitted by a laboratory enclosure with walls at a temperature of 2.7 K. Thus, for the cosmic background radiation, we say that the enclosure is the entire universe and that the universe is at an (average) temperature of 2.7 K. For their discovery of the cosmic background radiation, Penzias and Wilson were awarded the 1978 Nobel Prize in physics.

The cosmic background radiation is now known to be light that has been in flight across the universe since shortly after the universe began billions of years ago. When the universe was even younger, light could scarcely go any significant distance without being scattered by all the individual, high-speed particles along its path. If a light ray started from, say, point  $A$ , it would be scattered in so many directions that if you could have intercepted part of it, you would have not been



**Figure 44-5** The rotational speed of stars in a typical galaxy as a function of their distance from the galactic center. The theoretical solid curve shows that if a galaxy contained only the mass that is visible, the observed rotational speed would drop off with distance at large distances. The dots are the experimental data, which show that the rotational speed is approximately constant at large distances.

able to tell that it originated at point *A*. However, after the particles began to form atoms, the scattering of light greatly decreased. A light ray from point *A* might then be able to travel for billions of years without being scattered. This light is the cosmic background radiation.

As soon as the nature of the radiation was recognized, researchers wondered, “Can we use this incoming radiation to distinguish the points at which it originated, so that we then can produce an image of the early universe, back when atoms first formed and light scattering largely ceased?” The answer is yes, and that image is coming up in a moment.

## Dark Matter

At the Kitt Peak National Observatory in Arizona, Vera Rubin and her co-worker Kent Ford measured the rotational rates of a number of distant galaxies. They did so by measuring the Doppler shifts of bright clusters of stars located within each galaxy at various distances from the galactic center. As Fig. 44-5 shows, their results were surprising: The orbital speed of stars at the outer visible edge of the galaxy is about the same as that of stars close to the galactic center.

As the solid curve in Fig. 44-5 attests, that is not what we would expect to find if all the mass of the galaxy were represented by visible light. Nor is the pattern found by Rubin and Ford what we find in the solar system. For example, the orbital speed of Pluto (the “planet” most distant from the Sun) is only about one-tenth that of Mercury (the planet closest to the Sun).

The only explanation for the findings of Rubin and Ford that is consistent with Newtonian mechanics is that a typical galaxy contains much more matter than what we can actually see. In fact, the visible portion of a galaxy represents only about 5 to 10% of the total mass of the galaxy. In addition to these studies of galactic rotation, many other observations lead to the conclusion that the universe abounds in matter that we cannot see. This unseen matter is called **dark matter** because either it does not emit light or its light emission is too dim for us to detect.

Normal matter (such as stars, planets, dust, and molecules) is often called **baryonic matter** because its mass is primarily due to the combined mass of the protons and neutrons (baryons) it contains. (The much smaller mass of the electrons is neglected.) Some of the normal matter, such as burned-out stars and dim interstellar gas, is part of the dark matter in a galaxy.

However, according to various calculations, this dark normal matter is only a small part of the total dark matter. The rest is called **nonbaryonic dark matter** because it does not contain protons and neutrons. We know of only one member of this type of dark matter—the neutrinos. Although the mass of a neutrino is very small relative to the mass of a proton or neutron, the number of neutrinos in a galaxy is huge and thus the total mass of the neutrinos is large. Nevertheless, calculations indicate that not even the total mass of the neutrinos is enough to account for the total mass of the nonbaryonic dark matter. In spite of over a hundred years in which elementary particles have been detected and studied, the particles that make up the rest of this type of dark matter are undetected and their nature is unknown. Because we have no experience with them, they must interact only gravitationally with the common particles.

## The Big Bang

In 1985, a physicist remarked at a scientific meeting:

*It is as certain that the universe started with a big bang about 15 billion years ago as it is that the Earth goes around the Sun.*

This strong statement suggests the level of confidence in which the big bang theory, first advanced by Belgian physicist Georges Lemaître, is held by those

who study these matters. However, you must not imagine that the big bang was like the explosion of some gigantic firecracker and that, in principle at least, you could have stood to one side and watched. There was no “one side” because the big bang represents the beginning of spacetime itself. From the point of view of our present universe, there is no position in space to which you can point and say, “The big bang happened there.” It happened everywhere.

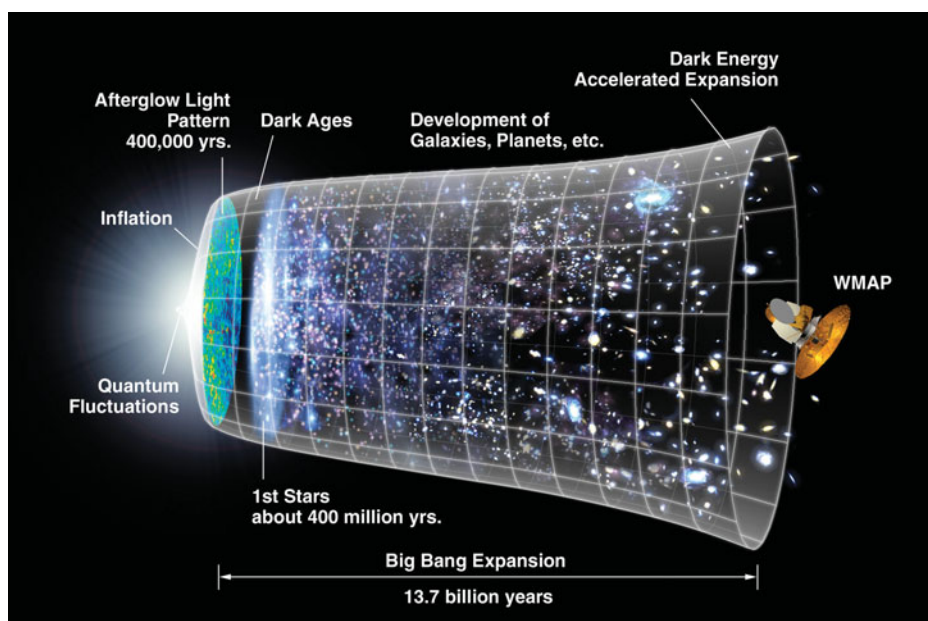
Moreover, there was no “before the big bang,” because time *began* with that creation event. In this context, the word “before” loses its meaning. We can, however, conjecture about what went on during succeeding intervals of time after the big bang (Fig. 44-6).

$t \approx 10^{-43}$  s. This is the earliest time at which we can say anything meaningful about the development of the universe. It is at this moment that the concepts of space and time come to have their present meanings and the laws of physics as we know them become applicable. At this instant, the entire universe (that is, the *entire* spatial extent of the universe) is much smaller than a proton and its temperature is about  $10^{32}$  K. Quantum fluctuations in the fabric of spacetime are the seeds that will eventually lead to the formation of galaxies, clusters of galaxies, and superclusters of galaxies.

$t \approx 10^{-34}$  s. By this moment the universe has undergone a tremendously rapid inflation, increasing in size by a factor of about  $10^{30}$ , causing the formation of matter in a distribution set by the initial quantum fluctuations. The universe has become a hot soup of photons, quarks, and leptons at a temperature of about  $10^{27}$  K, which is too hot for protons and neutrons to form.

$t \approx 10^{-4}$  s. Quarks can now combine to form protons and neutrons and their antiparticles. The universe has now cooled to such an extent by continued (but much slower) expansion that photons lack the energy needed to break up these new particles. Particles of matter and antimatter collide and annihilate each other. There is a slight excess of matter, which, failing to find annihilation partners, survives to form the world of matter that we know today.

$t \approx 1$  min. The universe has now cooled enough so that protons and neutrons, in colliding, can stick together to form the low-mass nuclei  $^2\text{H}$ ,  $^3\text{He}$ ,  $^4\text{He}$ , and  $^7\text{Li}$ . The predicted relative abundances of these nuclides are just what we observe in the universe today. Also, there is plenty of radiation present at  $t \approx 1$  min,



**Figure 44-6** An illustration of the universe from the initial quantum fluctuations just after  $t = 0$  (at the left) to the current accelerated expansion,  $13.7 \times 10^9$  y later (at the right). Don't take the illustration literally—there is *no* such “external view” of the universe because there is *no* exterior to the universe.

but this light cannot travel far before it interacts with a nucleus. Thus the universe is opaque.

$t \approx 379\,000\text{ y}$ . The temperature has now fallen to 2970 K, and electrons can stick to bare nuclei when the two collide, forming atoms. Because light does not interact appreciably with (uncharged) particles, such as neutral atoms, the light is now free to travel great distances. This radiation forms the cosmic background radiation that we discussed earlier. Atoms of hydrogen and helium, under the influence of gravity, begin to clump together, eventually starting the formation of galaxies and stars, but until then, the universe is relatively dark (Fig. 44-6).

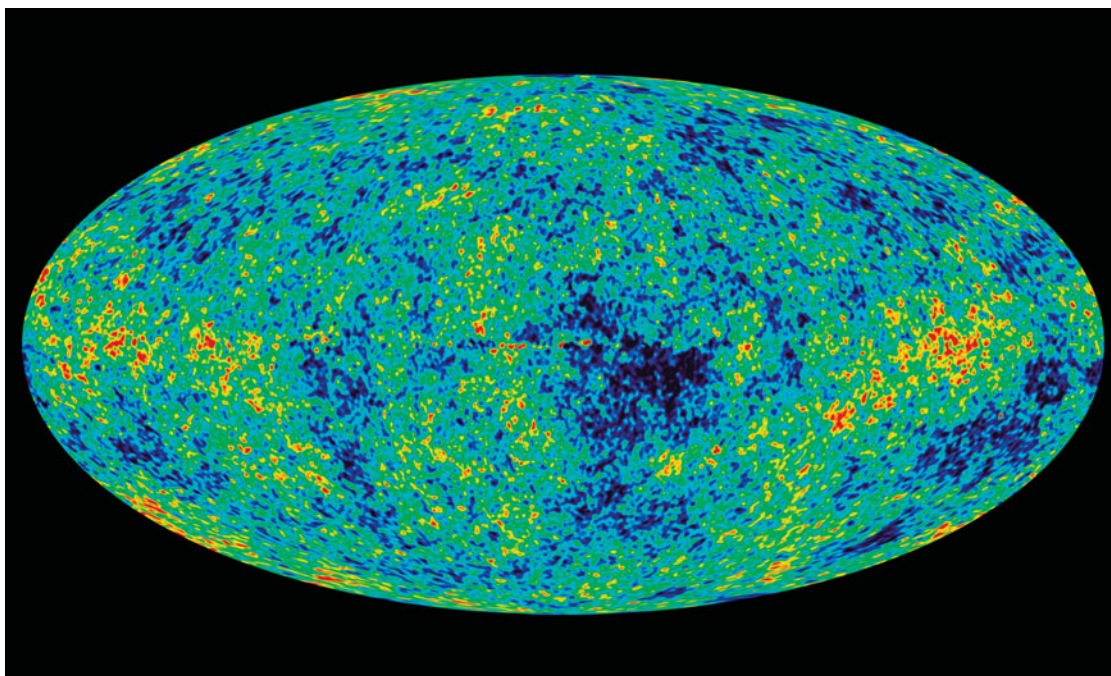
Early measurements suggested that the cosmic background radiation is uniform in all directions, implying that 379 000 y after the big bang all matter in the universe was uniformly distributed. This finding was most puzzling because matter in the present universe is not uniformly distributed, but instead is collected in galaxies, clusters of galaxies, and superclusters of galactic clusters. There are also vast *voids* in which there is relatively little matter, and there are regions so crowded with matter that they are called *walls*. If the big bang theory of the beginning of the universe is even approximately correct, the seeds for this nonuniform distribution of matter must have been in place before the universe was 379 000 y old and now should show up as a nonuniform distribution of the microwave background radiation.

In 1992, measurements made by NASA's Cosmic Background Explorer (COBE) satellite revealed that the background radiation is, in fact, not perfectly uniform. In 2003, measurements by NASA's Wilkinson Microwave Anisotropy Probe (WMAP) greatly increased our resolution of this nonuniformity. The resulting image (Fig. 44-7) is effectively a color-coded photograph of the universe when it was only 379 000 y old. As you can see from the variations in the colors, large-scale collecting of matter had already begun. Thus, the big bang theory and the theory of inflation at  $t \approx 10^{-34}\text{ s}$  are on the right track.

### The Accelerated Expansion of the Universe

Recall from Module 13-8 the statement that mass causes curvature of space. Now that we have seen that mass is a form of energy, as given by Einstein's equa-

**Figure 44-7** This color-coded image is effectively a photograph of the universe when it was only 379 000 y old, which was about  $13.7 \times 10^9\text{ y}$  ago. This is what you would have seen then as you looked away in all directions (the view has been condensed to this oval). Patches of light from collections of atoms stretch across the “sky,” but galaxies, stars, and planets have not yet formed.



Courtesy WMAP Science Team/NASA



tion  $E = mc^2$ , we can generalize the statement: energy can cause curvature of space. This certainly happens to the space around the energy packed into a black hole and, more weakly, to the space around any other astronomical body, but is the space of the universe as a whole curved by the energy the universe contains?

The question was answered first by the 1992 COBE measurements of the cosmic background radiation. It was then answered more definitively by the 2003 WMAP measurements that produced the image in Fig. 44-7. The spots we see in that image are the original sources of the cosmic background radiation, and the angular distribution of the spots reveals the curvature of the universe through which the light has to travel to reach us. If adjacent spots subtend either more than  $1^\circ$  (Fig. 44-8a) or less than  $1^\circ$  (Fig. 44-8b) in the detector's view (or our view) into the universe, then the universe is curved. Analysis of the spot distribution in the WMAP image shows that the spots subtend about  $1^\circ$  (Fig. 44-8c), which means that the universe is *flat* (having no curvature). Thus, the initial curvature the universe presumably had when it began must have been flattened out by the rapid inflation the universe underwent at  $t \approx 10^{-34}$  s.

This flatness poses a very difficult problem for physicists because it requires that the universe contain a certain amount of energy (as mass or otherwise). The trouble is that all estimations of the amount of energy in the universe (both in known forms and in the form of the unknown type of dark matter) fall dramatically short of the required amount.

One theory proposed about this missing energy gave it the gothic name of *dark energy* and predicted that it has the strange property of causing the expansion of the universe to accelerate. Until 1998, determining whether the expansion is, in fact, accelerating was very difficult because it requires measuring distances to very distant astronomical bodies where the acceleration might show up.

In 1998, however, advances in astronomical technology allowed astronomers to detect a certain type of supernovae at very great distances. More important, the astronomers could measure the duration of the burst of light from such a supernova. The duration reveals the brightness of the supernova that would be seen by an observer near the supernova. By measuring the brightness of the supernova as seen from Earth, astronomers could then determine the distance to the supernova. From the redshift of the light from the galaxy containing the supernova, astronomers could also determine how fast the galaxy is receding from us. Combining all this information, they could then calculate the expansion rate of the universe. The conclusion is that the expansion is indeed accelerating as predicted by the theory of dark energy (Fig. 44-6). However, we have no clue as to what this dark energy is.

Figure 44-9 gives our current state of knowledge about the energy in the universe. About 4% is associated with baryonic matter, which we understand fairly well. About 23% is associated with nonbaryonic dark matter, about which we have a few clues that might be fruitful. The rest, a whopping 73%, is associated with dark energy, about which we are clueless. There have been times in the history of physics, even in the 1990s, when pontiffs proclaimed that physics was nearly complete, that only details were left. In fact, we are nowhere near the end.

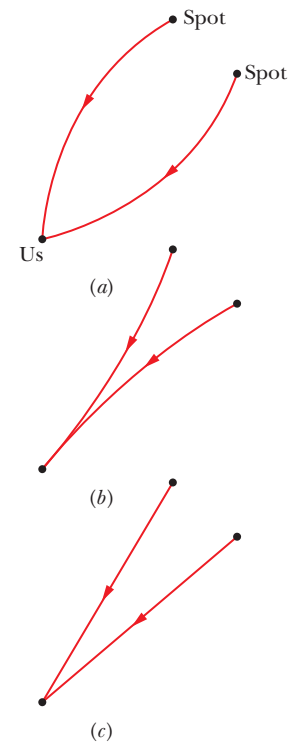
## A Summing Up

In this closing paragraph, let's consider where we are headed as we accumulate knowledge about the universe more and more rapidly. What we have found is marvelous and profound, but it is also humbling in that each new step seems to reveal more clearly our own relative insignificance in the grand scheme of things. Thus, in roughly chronological order, we humans have come to realize that

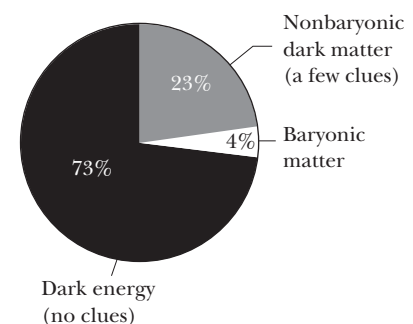
Our Earth is not the center of the solar system.

Our Sun is but one star among many in our galaxy.

Our galaxy is but one of many, and our Sun is an insignificant star in it.



**Figure 44-8** Light rays from two adjacent spots in our view of the cosmic background radiation would reach us at an angle (a) greater than  $1^\circ$  or (b) less than  $1^\circ$  if the space along the light-ray paths through the universe were curved. (c) An angle of  $1^\circ$  means that the space is not curved.



**Figure 44-9** The distribution of energy (including mass) in the universe.



Our Earth has existed for perhaps only a third of the age of the universe and will surely disappear when our Sun burns up its fuel and becomes a red giant.

Our species has inhabited Earth for less than a million years—a blink in cosmological time.

Although our position in the universe may be insignificant, the laws of physics that we have discovered (uncovered?) seem to hold throughout the universe and—as far as we know—have held since the universe began and will continue to hold for all future time. At least, there is no evidence that other laws hold in other parts of the universe. Thus, until someone complains, we are entitled to stamp the laws of physics “Discovered on Earth.” Much remains to be discovered. In the words of writer Eden Phillpotts, “*The universe is full of magical things, patiently waiting for our wits to grow sharper.*” That declaration allows us to answer one last time the question “What is physics?” that we have explored repeatedly in this book. Physics is the gateway to those magical things.

## Review & Summary

**Leptons and Quarks** Current research supports the view that all matter is made of six kinds of **leptons** (Table 44-2), six kinds of **quarks** (Table 44-5), and 12 **antiparticles**, one corresponding to each lepton and each quark. All these particles have spin quantum numbers equal to  $\frac{1}{2}$  and are thus **fermions** (particles with half-integer spin quantum numbers).

**The Interactions** Particles with electric charge interact through the electromagnetic force by exchanging **virtual photons**. Leptons can also interact with each other and with quarks through the **weak force**, via massive W and Z particles as messengers. In addition, quarks interact with each other through the **color force**. The electromagnetic and weak forces are different manifestations of the same force, called the **electroweak force**.

**Leptons** Three of the leptons (the **electron**, **muon**, and **tau**) have electric charge equal to  $-1e$ . There are also three uncharged **neutrinos** (also leptons), one corresponding to each of the charged leptons. The antiparticles for the charged leptons have positive charge.

**Quarks** The six quarks (up, down, strange, charm, bottom, and top, in order of increasing mass) each have baryon number  $+\frac{1}{3}$  and charge equal to either  $+\frac{2}{3}e$  or  $-\frac{1}{3}e$ . The strange quark has strange-

ness  $-1$ , whereas the others all have strangeness 0. These four algebraic signs are reversed for the antiquarks.

**Hadrons: Baryons and Mesons** Quarks combine into strongly interacting particles called **hadrons**. **Baryons** are hadrons with half-integer spin quantum numbers ( $\frac{1}{2}$  or  $\frac{3}{2}$ ). **Mesons** are hadrons with integer spin quantum numbers (0 or 1) and thus are **bosons**. Baryons are fermions. Mesons have baryon number equal to zero; baryons have baryon number equal to  $+1$  or  $-1$ . **Quantum chromodynamics** predicts that the possible combinations of quarks are either a quark with an antiquark, three quarks, or three antiquarks (this prediction is consistent with experiment).

**Expansion of the Universe** Current evidence strongly suggests that the universe is expanding, with the distant galaxies moving away from us at a rate  $v$  given by **Hubble’s law**:

$$v = Hr \quad (\text{Hubble’s law}). \quad (44-19)$$

Here we take  $H$ , the **Hubble constant**, to have the value

$$H = 71.0 \text{ km/s} \cdot \text{Mpc} = 21.8 \text{ mm/s} \cdot \text{ly}. \quad (44-21)$$

The expansion described by Hubble’s law and the presence of ubiquitous background microwave radiation reveal that the universe began in a “big bang” 13.7 billion years ago.

## Questions

- 1 An electron cannot decay into two neutrinos. Which of the following conservation laws would be violated if it did: (a) energy, (b) angular momentum, (c) charge, (d) lepton number, (e) linear momentum, (f) baryon number?
- 2 Which of the eight pions in Fig. 44-2b has the least kinetic energy?
- 3 Figure 44-10 shows the paths of two particles circling in a uniform magnetic field. The particles have the same magnitude of charge but opposite signs. (a) Which path corresponds to the more massive particle? (b) If the magnetic field is directed into the plane of the page, is the more massive particle positively or negatively charged?

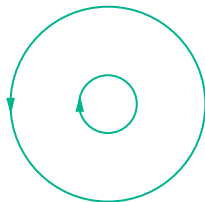


Figure 44-10  
Question 3.

- 4 A proton has enough mass energy to decay into a shower made up of electrons, neutrinos, and their antiparticles. Which of the following conservation laws would necessarily be violated if it did: electron lepton number or baryon number?
- 5 A proton cannot decay into a neutron and a neutrino. Which of the following conservation laws would be violated if it did: (a) energy (assume the proton is stationary), (b) angular momentum, (c) charge, (d) lepton number, (e) linear momentum, (f) baryon number?
- 6 Does the proposed decay  $\Lambda^0 \rightarrow p + K^-$  conserve (a) electric charge, (b) spin angular momentum, and (c) strangeness? (d) If the original particle is stationary, is there enough energy to create the decay products?
- 7 Not only particles such as electrons and protons but also entire

atoms can be classified as fermions or bosons, depending on whether their overall spin quantum numbers are, respectively, half-integral or integral. Consider the helium isotopes  ${}^3\text{He}$  and  ${}^4\text{He}$ . Which of the following statements is correct? (a) Both are fermions. (b) Both are bosons. (c)  ${}^4\text{He}$  is a fermion, and  ${}^3\text{He}$  is a boson. (d)  ${}^3\text{He}$  is a fermion, and  ${}^4\text{He}$  is a boson. (The two helium electrons form a closed shell and play no role in this determination.)

**8** Three cosmologists have each plotted a line on the Hubble-like graph of Fig. 44-11. If we calculate the corresponding age of the universe from the three plots, rank the plots according to that age, greatest first.

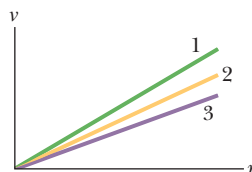


Figure 44-11  
Question 8.

**9** A  $\Sigma^+$  particle has these quantum numbers: strangeness  $S = -1$ , charge  $q = +1$ , and spin  $s = \frac{1}{2}$ . Which of the following quark combinations produces it: (a) dds, (b)  $s\bar{s}$ , (c) uus, (d) ssu, or (e) uu $\bar{s}$ ?

**10** As we have seen, the  $\pi^-$  meson has the quark structure  $d\bar{u}$ . Which of the following conservation laws would be violated if a  $\pi^-$  were formed, instead, from a d quark and a u quark: (a) energy, (b) angular momentum, (c) charge, (d) lepton number, (e) linear momentum, (f) baryon number?

**11** Consider the neutrino whose symbol is  $\bar{\nu}_\tau$ . (a) Is it a quark, a lepton, a meson, or a baryon? (b) Is it a particle or an antiparticle? (c) Is it a boson or a fermion? (d) Is it stable against spontaneous decay?

## Problems



Tutoring problem available (at instructor's discretion) in WileyPLUS and WebAssign



Worked-out solution available in Student Solutions Manual

WWW Worked-out solution is at



Number of dots indicates level of problem difficulty

ILW Interactive solution is at



Additional information available in *The Flying Circus of Physics* and at [flyingcircusofphysics.com](http://flyingcircusofphysics.com)

<http://www.wiley.com/college/halliday>

### Module 44-1 General Properties of Elementary Particles

**•1** A positively charged pion decays by Eq. 44-7:  $\pi^+ \rightarrow \mu^+ + \nu$ . What must be the decay scheme of the negatively charged pion? (Hint: The  $\pi^-$  is the antiparticle of the  $\pi^+$ .)

**•2** Certain theories predict that the proton is unstable, with a half-life of about  $10^{32}$  years. Assuming that this is true, calculate the number of proton decays you would expect to occur in one year in the water of an Olympic-sized swimming pool holding  $4.32 \times 10^5$  L of water.

**•3** An electron and a positron undergo pair annihilation (Eq. 44-5). If they had approximately zero kinetic energy before the annihilation, what is the wavelength of each  $\gamma$  produced by the annihilation?

**•4** A neutral pion initially at rest decays into two gamma rays:  $\pi^0 \rightarrow \gamma + \gamma$ . Calculate the wavelength of the gamma rays. Why must they have the same wavelength?

**•5** An electron and a positron are separated by distance  $r$ . Find the ratio of the gravitational force to the electric force between them. From the result, what can you conclude concerning the forces acting between particles detected in a bubble chamber? (Should gravitational interactions be considered?)

**••6** (a) A stationary particle 1 decays into particles 2 and 3, which move off with equal but oppositely directed momenta. Show that the kinetic energy  $K_2$  of particle 2 is given by

$$K_2 = \frac{1}{2E_1} [(E_1 - E_2)^2 - E_3^2],$$

where  $E_1$ ,  $E_2$ , and  $E_3$  are the rest energies of the particles. (b) A stationary positive pion  $\pi^+$  (rest energy 139.6 MeV) can decay to an antimuon  $\mu^+$  (rest energy 105.7 MeV) and a neutrino  $\nu$  (rest energy approximately 0). What is the resulting kinetic energy of the antimuon?

**••7** The rest energy of many short-lived particles cannot be measured directly but must be inferred from the measured momenta and known rest energies of the decay products. Consider the  $\rho^0$

meson, which decays by the reaction  $\rho^0 \rightarrow \pi^+ + \pi^-$ . Calculate the rest energy of the  $\rho^0$  meson given that the oppositely directed momenta of the created pions each have magnitude 358.3 MeV/c. See Table 44-4 for the rest energies of the pions.

**••8** A positive tau ( $\tau^+$ , rest energy = 1777 MeV) is moving with 2200 MeV of kinetic energy in a circular path perpendicular to a uniform 1.20 T magnetic field. (a) Calculate the momentum of the tau in kilogram-meters per second. Relativistic effects must be considered. (b) Find the radius of the circular path.

**••9** Observations of neutrinos emitted by the supernova SN1987a (Fig. 43-12b) place an upper limit of 20 eV on the rest energy of the electron neutrino. If the rest energy of the electron neutrino were, in fact, 20 eV, what would be the speed difference between light and a 1.5 MeV electron neutrino?

**••10** A neutral pion has a rest energy of 135 MeV and a mean life of  $8.3 \times 10^{-17}$  s. If it is produced with an initial kinetic energy of 80 MeV and decays after one mean lifetime, what is the longest possible track this particle could leave in a bubble chamber? Use relativistic time dilation.

### Module 44-2 Leptons, Hadrons, and Strangeness

**•11** Which conservation law is violated in each of these proposed decays? Assume that the initial particle is stationary and the decay products have zero orbital angular momentum. (a)  $\mu^- \rightarrow e^- + \nu_\mu$ ; (b)  $\mu^- \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$ ; (c)  $\mu^+ \rightarrow \pi^+ + \nu_\mu$ .

**•12** The  $A_2^+$  particle and its products decay according to the scheme

$$\begin{aligned} A_2^+ &\rightarrow \rho^0 + \pi^+, & \mu^+ &\rightarrow e^+ + \nu + \bar{\nu}, \\ \rho^0 &\rightarrow \pi^+ + \pi^-, & \pi^- &\rightarrow \mu^- + \bar{\nu}, \\ \pi^+ &\rightarrow \mu^+ + \nu, & \mu^- &\rightarrow e^- + \nu + \bar{\nu}. \end{aligned}$$

(a) What are the final stable decay products? From the evidence, (b) is the  $A_2^+$  particle a fermion or a boson and (c) is it a meson or a baryon? (d) What is its baryon number?

**•13** Show that if, instead of plotting strangeness  $S$  versus charge  $q$

for the spin- $\frac{1}{2}$  baryons in Fig. 44-3a and for the spin-zero mesons in Fig. 44-3b, we plot the quantity  $Y = B + S$  versus the quantity  $T_z = q - \frac{1}{2}(B + S)$ , we get the hexagonal patterns without using sloping axes. (The quantity  $Y$  is called *hypercharge*, and  $T_z$  is related to a quantity called *isospin*.)

•14 Calculate the disintegration energy of the reactions (a)  $\pi^+ + p \rightarrow \Sigma^+ + K^+$  and (b)  $K^- + p \rightarrow \Lambda^0 + \pi^0$ .

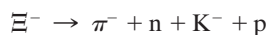
•15 Which conservation law is violated in each of these proposed reactions and decays? (Assume that the products have zero orbital angular momentum.) (a)  $\Lambda^0 \rightarrow p + K^-$ ; (b)  $\Omega^- \rightarrow \Sigma^- + \pi^0$  ( $S = -3, q = -1, m = 1672 \text{ MeV}/c^2$ , and  $m_s = \frac{3}{2}$  for  $\Omega^-$ ); (c)  $K^- + p \rightarrow \Lambda^0 + \pi^+$ .

•16 Does the proposed reaction



conserve (a) charge, (b) baryon number, (c) electron lepton number, (d) spin angular momentum, (e) strangeness, and (f) muon lepton number?

•17 Does the proposed decay process



conserve (a) charge, (b) baryon number, (c) spin angular momentum, and (d) strangeness?

•18 By examining strangeness, determine which of the following decays or reactions proceed via the strong interaction: (a)  $K^0 \rightarrow \pi^+ + \pi^-$ ; (b)  $\Lambda^0 + p \rightarrow \Sigma^+ + n$ ; (c)  $\Lambda^0 \rightarrow p + \pi^-$ ; (d)  $K^- + p \rightarrow \Lambda^0 + \pi^0$ .

•19 The reaction  $\pi^+ + p \rightarrow p + p + \bar{n}$  proceeds via the strong interaction. By applying the conservation laws, deduce the (a) charge quantum number, (b) baryon number, and (c) strangeness of the antineutron.

•20 There are 10 baryons with spin  $\frac{3}{2}$ . Their symbols and quantum numbers for charge  $q$  and strangeness  $S$  are as follows:

	$q$	$S$		$q$	$S$
$\Delta^-$	-1	0	$\Sigma^{*0}$	0	-1
$\Delta^0$	0	0	$\Sigma^{*+}$	+1	-1
$\Delta^+$	+1	0	$\Xi^{*-}$	-1	-2
$\Delta^{++}$	+2	0	$\Xi^{*0}$	0	-2
$\Sigma^{*-}$	-1	-1	$\Omega^-$	-1	-3

Make a charge-strangeness plot for these baryons, using the sloping coordinate system of Fig. 44-3. Compare your plot with this figure.

•21 Use the conservation laws and Tables 44-3 and 44-4 to identify particle  $x$  in each of the following reactions, which proceed by means of the strong interaction: (a)  $p + p \rightarrow p + \Lambda^0 + x$ ; (b)  $p + \bar{p} \rightarrow n + x$ ; (c)  $\pi^- + p \rightarrow \Xi^0 + K^0 + x$ .

•22 GO A 220 MeV  $\Sigma^-$  particle decays:  $\Sigma^- \rightarrow \pi^- + n$ . Calculate the total kinetic energy of the decay products.

•23 GO Consider the decay  $\Lambda^0 \rightarrow p + \pi^-$  with the  $\Lambda^0$  at rest. (a) Calculate the disintegration energy. What is the kinetic energy of (b) the proton and (c) the pion? (Hint: See Problem 6.)

•24 The spin- $\frac{3}{2}$   $\Sigma^{*0}$  baryon (see table in Problem 24) has a rest energy of 1385 MeV (with an intrinsic uncertainty ignored here); the spin- $\frac{1}{2}$   $\Sigma^0$  baryon has a rest energy of 1192.5 MeV. If each of these particles has a kinetic energy of 1000 MeV, (a) which is moving faster and (b) by how much?

### Module 44-3 Quarks and Messenger Particles

•25 The quark makeups of the proton and neutron are uud and udd, respectively. What are the quark makeups of (a) the antiproton and (b) the antineutron?

•26 From Tables 44-3 and 44-5, determine the identity of the baryon formed from quarks (a) ddu, (b) uus, and (c) ssd. Check your answers against the baryon octet shown in Fig. 44-3a.

•27 What is the quark makeup of  $\bar{K}^0$ ?

•28 What quark combination is needed to form (a)  $\Lambda^0$  and (b)  $\Xi^0$ ?

•29 Which hadron in Tables 44-3 and 44-4 corresponds to the quark bundles (a) ssu and (b) dds?

•30 SSM WWW Using the up, down, and strange quarks only, construct, if possible, a baryon (a) with  $q = +1$  and strangeness  $S = -2$  and (b) with  $q = +2$  and strangeness  $S = 0$ .

### Module 44-4 Cosmology

•31 In the laboratory, one of the lines of sodium is emitted at a wavelength of 590.0 nm. In the light from a particular galaxy, however, this line is seen at a wavelength of 602.0 nm. Calculate the distance to the galaxy, assuming that Hubble's law holds and that the Doppler shift of Eq. 37-36 applies.

•32 Because of the cosmological expansion, a particular emission from a distant galaxy has a wavelength that is 2.00 times the wavelength that emission would have in a laboratory. Assuming that Hubble's law holds and that we can apply Doppler-shift calculations, what was the distance (ly) to that galaxy when the light was emitted?

•33 What is the observed wavelength of the 656.3 nm (first Balmer) line of hydrogen emitted by a galaxy at a distance of  $2.40 \times 10^8$  ly? Assume that the Doppler shift of Eq. 37-36 and Hubble's law apply.

•34 An object is  $1.5 \times 10^4$  ly from us and does not have any motion relative to us except for the motion due to the expansion of the universe. If the space between us and it expands according to Hubble's law, with  $H = 21.8 \text{ mm/s} \cdot \text{ly}$ , (a) how much extra distance (meters) will be between us and the object by this time next year and (b) what is the speed of the object away from us?

•35 If Hubble's law can be extrapolated to very large distances, at what distance would the apparent recessional speed become equal to the speed of light?

•36 What would the mass of the Sun have to be if Pluto (the outermost "planet" most of the time) were to have the same orbital speed that Mercury (the innermost planet) has now? Use data from Appendix C, express your answer in terms of the Sun's current mass  $M_S$ , and assume circular orbits.

•37 The wavelength at which a thermal radiator at temperature  $T$  radiates electromagnetic waves most intensely is given by Wien's law:  $\lambda_{\text{max}} = (2898 \mu\text{m} \cdot \text{K})/T$ . (a) Show that the energy  $E$  of a photon corresponding to that wavelength can be computed from

$$E = (4.28 \times 10^{-10} \text{ MeV/K})T.$$

(b) At what minimum temperature can this photon create an electron-positron pair (as discussed in Module 21-3)?

•38 Use Wien's law (see Problem 37) to answer the following questions: (a) The cosmic background radiation peaks in intensity at a wavelength of 1.1 mm. To what temperature does this correspond? (b) About 379 000 y after the big bang, the universe became transparent to electromagnetic radiation. Its temperature then was