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# **Tutorial 3**

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TIME SERIES & DATA ANALYSIS

NASSP – UCT

DUE 11AM ON 31 MARCH 2020

You may use any programming language.

Submit both the code (can be as Jupyter notebook) and a pdf by email.

## Question 1: Phase Dispersion Minimization

- 1.1 Write your own implementation of the PDM algorithm
- 1.2 Run your implementation on the sunspot data (`zuerich-monthly-sunspot-numbers-.csv`) and plot the PDM statistic vs period.
- 1.3 Run the PyAstronomy implementation of the PDM on the sunspot data and compare results of the two implementations in terms of any periodicity in the dataset.

## Question 2: Event data

- 2.1 Read in the event data of the magnetar Swift J1834.9–0846 and plot a light curve.  
(see Kargaltsev et al. 2012 – `acisf14329N002_evt2_reg_filtered.txt`)
- 2.2 Calculate the  $Z_1^2$  statistic for this time series, and plot the statistic vs the input period.
- 2.3 Can you identify the 2.48 s spin period of the neutron star?

## Question 3: Non-Fourier methods

Explain the differences and similarities between the epoch-folding, phase dispersion minimisation and the minimum string length methods for period finding.

## Question 4: Dynamic periodogram

Use the data on the unusual X-ray binary Her X-1 (`herx1_asm.txt`) for this question. The columns are Modified Julian Date, Count rate, and an error on the count rate.

- 4.1 Create a series of Lomb-Scargle periodograms for overlapping stretches of the Her X-1 lightcurve, using the frequency grid from 0.01 to 0.1 with a frequency resolution of 0.001. Use a data window of 200 d with an overlap of 50d.
- 4.2 Visualize the dynamic periodogram by constructing a three-dimensional map in (time, frequency, power) space with spectral power being represented by the colour scale.
- 4.3 What can you say about the evolution of the 35 d period?