

## The Sommerfeld Theory of Metals

In the Drude model it is implicitly assumed that the electron velocity distribution is given by the Maxwell-Boltzmann's:

$$f(\vec{v}) = n \left( \frac{m}{2\pi k_B T} \right)^{3/2} e^{-\frac{\frac{1}{2}mv^2}{k_B T}}$$

### 1. The Maxwell-Boltzmann velocity distribution:

Consider a classical gas with total energy  $K(\vec{p}_1, \vec{p}_2, \dots, \vec{p}_N) + U(\vec{r}_1, \vec{r}_2, \dots, \vec{r}_N)$

The kinetic part is a quadratic function of the momenta, while the potential part depends only on the particle positions. The probability of occurrence of a microstate at equilibrium at  $T$  is given by the canonical distribution:

$$P(\vec{r}_1, \dots, \vec{r}_N, \vec{p}_1, \dots, \vec{p}_N) = A e^{-[K+U]/kT} = A e^{-\frac{K}{kT}} e^{-\frac{U}{kT}}$$

This factorization implies that the probabilities of position and momenta are independent. The probability for the momenta are:

$$f(\vec{p}_1, \dots, \vec{p}_N) d\vec{p} = C e^{-\frac{K}{kT}} d\vec{p}$$

Because  $K$  is the sum of the energy of each particle, this probability is still factorized: the momenta distribution of the various particles are independent (and do not depend on the interactions)

For a single particle:

$$f(p_x, p_y, p_z) dp_x dp_y dp_z = C e^{-\frac{(p_x^2 + p_y^2 + p_z^2)/2m}{kT}} dp_x dp_y dp_z$$

$$C = \frac{1}{\int_{-\infty}^{+\infty} e^{-\frac{(p_x^2 + p_y^2 + p_z^2)}{2m kT}} dp_x dp_y dp_z} = \frac{1}{\left[ \int_{-\infty}^{+\infty} e^{-\frac{p^2}{2m kT}} dp \right]^3} =$$

$$= (2\pi m kT)^{-3/2}$$

Hence:

$$d^3p \quad f(p_x, p_y, p_z) = \left( \frac{1}{2\pi m kT} \right)^{3/2} e^{-\frac{(p_x^2 + p_y^2 + p_z^2)}{2m kT}} dp_x dp_y dp_z$$

$$f(v_x, v_y, v_z) dv_x dv_y dv_z = \left( \frac{m}{2\pi kT} \right)^{3/2} e^{-\frac{v_x^2 + v_y^2 + v_z^2}{2kT}} dv_x dv_y dv_z$$

□

## 2. The Fermi-Dirac distribution

Occupation numbers: we divide the system into sub-systems, each of which is the set of all particles that are in a given single particle microstate. Then, the number of the particles in a sub-system can vary  $\Rightarrow$  the use of the grand-canonical ensemble: (more to the next page)

$$Z_G = \sum_S e^{-\beta(E_S - \mu N_S)} \quad ; \quad \Omega = -kT \ln Z_G$$

$S$  runs over the microstates; The mean number of particles is

then

$$\bar{N} = - \left( \frac{\partial \Omega}{\partial \mu} \right)_{T, V}$$

The probability of occupation of a particular microstate is:

$$p_S = \frac{1}{Z_G} e^{-\beta(E_S - \mu N_S)}$$

For a single particle microstate  $\epsilon_k$  (a sub-system), the grand-partition function is

$$Z_{G,k} = \sum_{n_k} e^{-\beta n_k (\epsilon_k - \mu)} = \sum_{n_k} e^{-\beta (\epsilon_k - \mu) n_k}$$

For fermions,  $n_k = 0, \text{ or } 1$ :  $\rightarrow Z_{G,k} = 1 + e^{-\beta (\epsilon_k - \mu)}$

Then, the Landau potential becomes:

$$\Omega_k = -KT \ln Z_{G,k} = -KT \ln [1 + e^{-\beta (\epsilon_k - \mu)}]$$

Then,

$$\bar{n}_k = - \frac{\partial \Omega_k}{\partial \mu} = \frac{e^{-\beta (\mu - \epsilon_k)}}{1 + e^{-\beta (\mu - \epsilon_k)}} = \frac{1}{e^{\beta (\epsilon_k - \mu)} + 1}$$

□

For an ideal Fermi gas:  $\epsilon_k = \frac{1}{2} m v^2$

$$V = \frac{h^3}{m^3} \quad \frac{d^3 k}{(2\pi)^3} = \left(\frac{m}{h}\right)^3 \frac{1}{(2\pi)^3} d^3 v$$

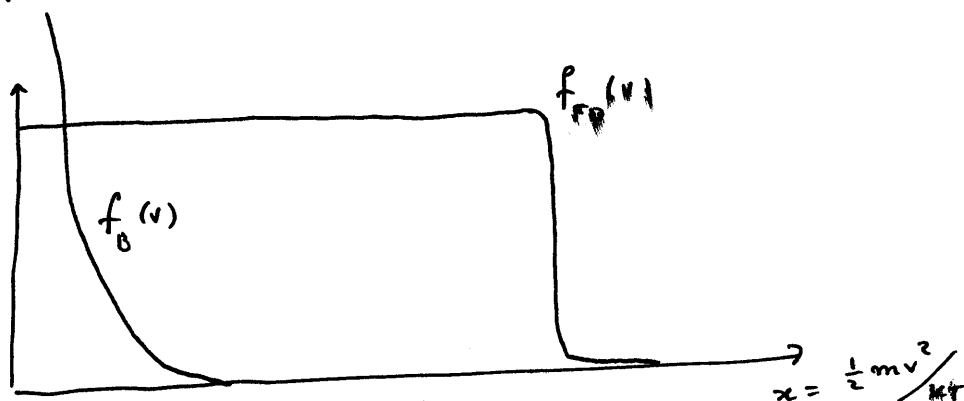
or:

$$d^3 v f(\vec{v}) = \left(\frac{m}{h}\right)^3 \frac{1}{4\pi^3} \frac{1}{e^{\beta (\frac{1}{2} m v^2 - \mu)} + 1} d^3 v$$

(including spin)

□

What are the consequences of this <sup>drastic</sup> change in terms of the electron gas?



3. The ground-state of  $N$  electrons confined to a volume  $V$ :  
(ideal gas: electrons do not interact with each other)

$$-\frac{\hbar^2}{2m} \nabla^2 \psi = E \psi \quad (1)$$

Boundary conditions (Periodic)  $\psi(x, y, z+L) = \psi(x, y, z) \quad \text{etc}$   
 $\psi(x, y+L, z) = \psi(x, y, z) \quad \text{etc}$

The general solution of (1) is  $\psi_{\vec{k}} = \frac{1}{\sqrt{V}} e^{i\vec{k} \cdot \vec{r}}$ , with  $E(\vec{k}) = \frac{\hbar^2 k^2}{2m}$

$\hat{p} = -i\hbar \nabla_{\vec{r}} \Rightarrow \hat{p} \psi = \hbar \vec{k} \psi$  (each eigenstate has a well defined momentum  $\vec{p} = \hbar \vec{k}$ , and hence  $\vec{v} = \frac{\vec{p}}{m} = \frac{\hbar \vec{k}}{m}$  ( $E(\vec{k}) = \frac{p^2}{2m}$ ))

We now invoke the boundary conditions:

$$e^{i k_x L} = e^{i k_y L} = e^{i k_z L} = 1 \Rightarrow k_x = \frac{2\pi m_x}{L}$$

Each allowed  $k$ -point occupies a volume  $\left(\frac{2\pi}{L}\right)^3$  in  $k$ -space

Therefore, a volume  $\Omega$  of the  $k$ -space will contain  $\frac{\Omega}{(2\pi/L)^3}$

points, or the number of allowed  $k$ -points per unit volume is:

$$\frac{V}{(2\pi)^3}$$

The ground state, will then correspond to occupy all the  $k$ -states within a sphere of radius  $k_F$ : Hence:

$$\frac{4}{3}\pi k_F^3 \frac{V}{(2\pi)^3} = \frac{N}{2} \text{ electrons}$$

$\Rightarrow \text{spin}$

$$\frac{N}{V} = \frac{k_F^3}{3\pi^2} = n \quad ; \quad E_F = \frac{\hbar^2 k_F^2}{2m}$$

that means that at  $T=0K$ , all states up to  $E_F$  are occupied and all states above  $E_F$  are empty.

Let us express these quantities in terms of the electron density

$$k_F = (3\pi^2 n)^{1/3} \quad E_F = \frac{\hbar^2 (3\pi^2 n)^{2/3}}{2m}$$

or in terms of a dimensionless parameter  $\frac{r_s}{a_0}$

$$r_s = \left( \frac{3}{4\pi n} \right)^{1/3} \equiv \text{Radius of the "free space" sphere per electron}$$

$$\frac{3}{4\pi r_s^3} = \frac{3}{4\pi n} \rightarrow n = \frac{3}{4\pi r_s^3} \rightarrow k_F = \left( \frac{3\pi^2 \cdot 3}{4\pi r_s^3} \right)^{1/3} = \left( \frac{9\pi}{4} \right)^{1/3} \frac{1}{r_s}$$

$$k_F \sim \frac{3,63}{r_s/a_0} \text{ \AA}^{-1} \Rightarrow \text{de Broglie wavelength} \sim \text{\AA}$$

As a result, the velocity  $v_F = \frac{\hbar k_F}{m} = 10^8 \text{ cm.s}^{-1}$  (1% c)

(Note that this is at  $T=0K$ ).

$$E_F = \frac{\hbar^2 k_F^2}{2m} = \left( \frac{e^2}{2a_0} \right) (k_F a_0)^2$$

↓

$R_y$  (Rydberg)  $\equiv$  ground state binding energy of the hydrogen atom

$\sim 13,6 \text{ eV}$

$$E_F = (13,6 \text{ eV}) (k_F a_0)^2 =$$

$$= \frac{50,1 \text{ eV}}{(\lambda_s/a_0)^2}$$

(Something in the range of  
1,5 — 15 eV)

Because the  $k$ -volume per point is  $\frac{8\pi^3}{V}$

$$\sum_{\vec{k}} F(\vec{k}) = \frac{V}{8\pi^3} \sum_{\vec{k}} F(\vec{k}) \Delta \vec{k}$$

In the limit  $V \rightarrow \infty$  ( $\Delta \vec{k} \rightarrow 0$ ) this goes to:

$$\int \frac{d\vec{k}}{8\pi^3} F(\vec{k}) = \lim_{V \rightarrow \infty} \frac{1}{V} \sum_{\vec{k}} F(\vec{k})$$

$$\bullet \quad \frac{E}{V} = \frac{1}{4\pi^3} \int_0^{k_F} d\vec{k} \frac{\hbar^2 k^2}{2m} = \frac{4\pi}{4\pi^3} \frac{\hbar^2}{2m} \int_0^{k_F} k^4 dk = \frac{1}{\pi^2} \frac{\hbar^2}{10m} k_F^5$$

$$\bullet \quad \frac{N}{V} = \frac{4\pi}{4\pi^3} \int_0^{k_F} k^2 dk = \frac{1}{\pi^2} \frac{1}{3} k_F^3$$

Hence  $\frac{E}{N} = \frac{3}{5} E_F = \frac{3}{5} k_B T_F \rightarrow T_F \sim 10^4 K!$

$$P = - \left( \frac{\partial E}{\partial V} \right)_N = \left[ \frac{\partial}{\partial V} \left( \frac{3}{5} N E_F \right) \right]_N = \frac{\partial}{\partial V} \left( \frac{3}{5} N \cdot \frac{\hbar^2 (3\pi^2 n)^{2/3}}{2m} \right)$$

$$n^{2/3} = \left( \frac{N}{V} \right)^{2/3} \rightarrow P = \frac{2}{3} \frac{E}{V}$$

$$B = -V \frac{\partial P}{\partial V} \text{ (bulk modulus)} = \frac{2}{3} n E_F$$

□

4. Thermal properties of the free electron gas:

$$\frac{E}{V} = u = \int \frac{d\vec{k}}{4\pi^3} \epsilon(\vec{k}) f_{FD}(\epsilon(\vec{k}))$$

$$\frac{N}{V} = n = \int \frac{d\vec{k}}{4\pi^3} f(\epsilon)$$

$$d\vec{k} = k^2 4\pi dk \quad \frac{\hbar^2 k^2}{2m} = E \quad dE = \frac{2\hbar^2 k}{2m} dk$$

$$k^2 dk = \frac{2mE}{\hbar^2} \cdot \frac{2m}{2\hbar^2 \left(\frac{2mE}{\hbar^2}\right)^{1/2}} dE =$$

$$= \sqrt{\frac{2mE}{\hbar^2}} \frac{m}{\hbar^2} dE$$

$$u = \int \frac{d\vec{k}}{4\pi^3} \epsilon(\vec{k}) f_{FD}(\epsilon) =$$

$$= \int \underbrace{\frac{1}{\pi^2} \frac{m}{\hbar^2} \sqrt{\frac{2mE}{\hbar^2}}}_{g(\epsilon) \equiv \text{density of states}} \epsilon f(\epsilon) d\epsilon = \int g(\epsilon) f(\epsilon) \epsilon d\epsilon$$

$$n = \int g(\epsilon) f(\epsilon) d\epsilon \quad \text{etc.}$$

The Sommerfeld expansion:

The above integrals have a general form of

$$\int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon, \quad \text{where } f(\epsilon) = \frac{1}{e^{\beta(\epsilon-\mu)} + 1}$$

$H(\epsilon) \rightarrow 0$  as  $\epsilon \rightarrow -\infty$ , and diverges at most as a power law as  $\epsilon \rightarrow \infty$ .

Define:  $K(\epsilon) = \int_{-\infty}^{\epsilon} H(\epsilon') d\epsilon' ; \quad \frac{dK}{d\epsilon} = H(\epsilon)$

Then,  $\int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon = K(\epsilon) f(\epsilon) \Big|_{-\infty}^{+\infty} - \int_{-\infty}^{+\infty} K(\epsilon) \left( + \frac{\partial f}{\partial \epsilon} \right) d\epsilon$

Consider  $K(\epsilon) = K(\mu) + \sum_{n=1}^{\infty} \frac{(\epsilon-\mu)^n}{n!} \left( \frac{d^n K}{d\epsilon^n} \right)_{\epsilon=\mu}$  (Taylor's series)

So:  $\int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon = \int_{-\infty}^{+\infty} \left[ K(\mu) + \sum_{n=1}^{\infty} \frac{(\epsilon-\mu)^n}{n!} \frac{d^n K}{d\epsilon^n} \right] \left( - \frac{\partial f}{\partial \epsilon} \right) d\epsilon$

$\frac{\partial f}{\partial \epsilon}$  is even  $\Rightarrow$  only  $m = \text{even}$  contribute to the above series  $\uparrow$ .

$\int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon = \int_{-\infty}^{\mu} H(\epsilon) d\epsilon + \sum_{n=1}^{\infty} \left( \frac{(\epsilon-\mu)^{2n}}{2n!} \right)_{\epsilon=\mu} \left( \frac{d^{2n} H}{d\epsilon^{2n-1}} \right) \left( - \frac{\partial f}{\partial \epsilon} \right) d\epsilon$

$\sum_{n=1}^{\infty} \int_{-\infty}^{\mu} \frac{(\epsilon-\mu)^{2n}}{2n!} \left( \frac{d^{2n} H}{d\epsilon^{2n-1}} \right)_{\epsilon=\mu} \left( - \frac{\partial f}{\partial \epsilon} \right) d\epsilon = \sum_n \left( \frac{d^{2n} H}{d\epsilon^{2n-1}} \right)_{\epsilon=\mu} \int_{-\infty}^{\mu} \left( \frac{\epsilon-\mu}{KT} \right)^{2n} \left( - \frac{\partial f}{\partial \epsilon} \right) \frac{d\epsilon}{2n!}$

$= \sum_{n=1}^{\infty} \left( \frac{d^{2n} H}{d\epsilon^{2n-1}} \right)_{\epsilon=\mu} a_n (KT)^{2n}$

with

$a_n = \int \frac{x^{2n}}{2n!} \left( - \frac{d}{dx} \frac{1}{e^{x^2+1}} \right) dx$



Hence: 
$$\int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon = \sum_{n=1}^{\infty} a_n \left( \frac{d^n H}{d\epsilon^{2n-1}} \right)_{\epsilon=\mu} (kT)^{2n} + \int_{-\infty}^{\mu} H(\epsilon) d\epsilon$$

(Sommerfeld expansion)

$$\rightarrow \int_{-\infty}^{+\infty} H(\epsilon) f(\epsilon) d\epsilon = \int_{-\infty}^{\mu} H(\epsilon) d\epsilon + \frac{\pi^2}{6} (k_B T)^2 H'(\mu) + \frac{7\pi^4}{360} (k_B T)^4 H'''(\mu) + \mathcal{O}\left(\frac{k_B T}{\mu}\right)^6$$

$$\bullet u = \int_{-\infty}^{+\infty} d\epsilon \underbrace{g(\epsilon) \epsilon}_{H(\epsilon)} f(\epsilon) = \int_{-\infty}^{\mu} g(\epsilon) \cdot \epsilon d\epsilon + \frac{\pi^2}{6} (kT)^2 [g(\mu) + \mu g'(\mu)] + \mathcal{O}(T^4)$$

$$\bullet n = \int_{-\infty}^{+\infty} d\epsilon g(\epsilon) f(\epsilon) = \int_{-\infty}^{\mu} g(\epsilon) d\epsilon + \frac{\pi^2}{6} (kT)^2 g'(\mu) + \mathcal{O}(T^4)$$

let us consider first this second term; Take  $g(\epsilon) = 0$  if  $\epsilon < 0$ ;

then

$$\begin{aligned} n &= \int_0^{\epsilon_F} d\epsilon g(\epsilon) = \int_0^{\mu(T)} g(\epsilon) d\epsilon + \frac{\pi^2}{6} (kT)^2 g'(\mu) + \dots \\ &= \int_0^{\epsilon_F} g(\epsilon) d\epsilon + g(\epsilon_F) (\mu - \epsilon_F) + \frac{\pi^2}{6} (kT)^2 g'(\epsilon_F) + \dots \end{aligned}$$

Hence

$$g(\epsilon_F) (\mu - \epsilon_F) = - \frac{\pi^2}{6} (kT)^2 g'(\epsilon_F)$$

$$\mu(T) \sim \epsilon_F - \frac{\pi^2}{6} (kT)^2 \left( \frac{g'}{g} \right)_{\epsilon_F}$$

$$\left. \begin{aligned} g &\propto \epsilon^{-1/2} \\ g' &\propto \frac{1}{2} \epsilon^{-3/2} \\ g &\propto \epsilon^{-1/2} \end{aligned} \right)$$

$$\begin{aligned} \mu(T) &\sim \epsilon_F - \frac{\pi^2}{6} (kT)^2 \frac{1}{2\epsilon_F} \\ &\sim \epsilon_F \left[ 1 - \frac{\pi^2}{12} \left( \frac{kT}{\epsilon_F} \right)^2 + \dots \right] \end{aligned}$$

Now,  $u$ :

$$\begin{aligned}
 u &= \int_{-\infty}^{\mu} g(\epsilon) \epsilon d\epsilon + \frac{\pi^2}{6} (kT)^2 [g(\mu) + \mu g'(\mu)] + \dots \\
 &= \int_0^{\epsilon_F} \epsilon g(\epsilon) d\epsilon + \overbrace{(\mu - \epsilon_F) \epsilon_F g(\epsilon_F) + \frac{\pi^2}{6} (kT)^2 \epsilon_F g'(\epsilon_F)}^0 + \\
 &\quad + \frac{\pi^2}{6} (kT)^2 g(\epsilon_F) + \dots
 \end{aligned}$$

$$= \int_0^{\epsilon_F} \epsilon g(\epsilon) d\epsilon + \frac{\pi^2}{6} (kT)^2 g(\epsilon_F) + \dots \propto (T^4)$$

$$u(T) = u_0 + \frac{\pi^2}{6} (kT)^2 g(\epsilon_F) + \dots$$

$$c_v = \frac{du}{dT} = \underbrace{\frac{\pi^2}{3} k_B^2 g(\epsilon_F)}_{\gamma} T$$

For free electrons:  $g(\epsilon_F) = \frac{3}{2} \frac{n}{\epsilon_F}$

$$c_v = \frac{\pi^2}{2} k_B \frac{n}{\epsilon_F} \cdot \frac{k_B T}{\epsilon_F} \rightarrow \text{becomes a function of } T$$

$$= \frac{3}{2} n k_B \cdot \frac{k_B T}{\epsilon_F} \pi^2$$

$$\frac{3}{2} n k_B \cdot \frac{\pi^2}{2} \frac{k_B T}{\epsilon_F}$$