



NOVAC

NOVAC Instrument v3.5 User Manual

**Optical Remote Sensing Group,
Department of Radio and Space Sciences,
Chalmers University of Technology,
Göteborg, Sweden**

Last updated 2022-12-15

Santiago Arellano, Mattias Johansson
Chalmers University Optical Remote Sensing Group
santiago.arellano@chalmers.se

Christoph Kern, Diana Norgaard, Allan Lerner
USGS Volcano Disaster Assistance Program
ckern@usgs.gov, alerner@usgs.gov

Table of Contents

1	OVERVIEW	3
2	SYSTEM REQUIREMENTS	4
3	SETUP AND CONFIGURATION OF THE INSTRUMENTS IN THE FIELD	5
3.1	TECHNICAL SETUP	5
3.1.1	<i>3rd Gen NOVAC instrument (MOXA and Ocean Optics)</i>	5
3.1.2	<i>4th Gen NOVAC instrument (Axiomtek and Avantes)</i>	10
3.1.3	<i>DOAS Scanner</i>	12
3.1.4	<i>Example setup</i>	15
3.2	CONFIGURING THE INSTRUMENT IN THE FIELD.....	17
3.2.1	<i>Give your computer a static IP address</i>	17
3.2.2	<i>Connecting to the instrument to configure it (with WinSCP)</i>	19
3.2.3	<i>Configuring the network settings (e.g. IP address) of the embedded computer</i>	19
3.2.4	<i>Configuring the instrument and measurement routine (cfg.txt)</i>	22
3.3	TROUBLESHOOTING OR REBOOTING THE INSTRUMENT WITH PUTTY, TELNET, OR HYPERTERMINAL	28
3.3.1	<i>Connecting to the instrument with an SSH client</i>	28
3.3.2	<i>Useful commands for troubleshooting.....</i>	29
3.3.3	<i>Description of important files and their locations.....</i>	30
4	CONFIGURING THE NOVAC SOFTWARE AT THE OBSERVATORY.....	31
4.1	SCANNING INSTRUMENT SETTINGS.....	31
4.1.1	<i>Instrument information 'Info'</i>	31
4.1.2	<i>Evaluation Setting</i>	33
4.1.3	<i>Communication Setting</i>	34
4.1.4	<i>Wind Configuration.....</i>	38
4.1.5	<i>Instrument Calibration.....</i>	39
4.2	GLOBAL SETTINGS	42
4.3	SAVING THE SETTINGS.....	44
5	DATA COLLECTION.....	44
5.1	RUNNING, SLEEPING	44
5.1.1	<i>Main Window</i>	45
5.1.2	<i>Overview Tab</i>	47
5.1.3	<i>Instrument Tab</i>	47
5.1.4	<i>Column History Tab</i>	49
5.1.5	<i>Flux History Tab</i>	49
5.1.6	<i>Wind Measurement Tab</i>	50
5.2	STOPPING THE DATA COLLECTION	51
5.3	EXPLANATION OF MESSAGES IN NOVAC SOFTWARE	51
5.4	OUTPUT FILES	53
5.4.1	<i>Local Folders Structure</i>	53
5.4.2	<i>Evaluation log</i>	54
5.4.3	<i>GeometryLog.....</i>	57
5.4.4	<i>FluxLog</i>	58
5.4.5	<i>StatusLog</i>	59
5.4.6	<i>Wind field file</i>	59
6	POST-PROCESSING SO₂ DATA.....	61
6.1	POST FLUX CALCULATION	61
6.2	RE-EVALUATION OF SPECTRA.....	64
6.2.1	<i>Re-Evaluation – Selecting Scans.....</i>	65
6.2.2	<i>Re-Evaluation – Setting up the fit windows.....</i>	65
6.2.3	<i>Re-Evaluation – Misc. Settings</i>	68

6.2.4	<i>Re-Evaluation – The final evaluation.....</i>	69
6.3	SETUP INSPECTION.....	71
6.4	POST-PROCESSING WIND SPEED AND DIRECTION.....	72
6.5	DATA BROWSER DIALOG.....	74
7	ADVANCED FILE HANDLING OPTIONS	75
7.1	IMPORTING FILES	75
7.2	EXPORTING FILES	76
7.2.1	<i>Exporting spectra.....</i>	76
7.2.2	<i>Exporting Evaluation logs.....</i>	77
7.3	CHECKING PAK-FILES	78
7.4	MERGE EVALUATION LOG FILES	78
7.5	MERGING/SPLITTING PAK-FILES	79
7.5.1	<i>Splitting pak-files into scans.....</i>	79
7.5.2	<i>Merging pak-files</i>	80
8	INSTRUMENT CALIBRATION.....	81
8.1	WAVELENGTH CALIBRATION AND INSTRUMENT LINE SHAPE FROM MEASURED SCAN DATA	81
8.1.1	<i>Required input.....</i>	82
8.1.2	<i>Setting up the instrument calibration</i>	82
8.1.3	<i>Running the instrument calibration procedure</i>	83
8.2	CREATING REFERENCES FROM INSTRUMENT CALIBRATION	85
8.2.1	<i>Creating a single reference</i>	86
8.2.2	<i>Creating standard references</i>	87
9	ADVANCED UPGRADING AND TROUBLESHOOTING	87
9.1	INSTALLING OR RE-INSTALLING KONGO ON A MOXA COMPUTER	88
9.1.1	<i>Formatting the MOXA to the factory settings</i>	88
9.1.2	<i>Killing kongo on a MOXA computer prior to re-installation</i>	89
9.1.3	<i>Installing kongo on a MOXA.....</i>	89
9.1.4	<i>Connecting to the Moxa with a serial cable</i>	91
9.2	INSTALLING OR RE-INSTALLING KONGO ON AXIOMTEK.....	92
9.3	SETTING THE DEFAULT BAUD RATE OF AN SD2000 SPECTROMETER.....	93
9.4	INSTRUMENT CONFIGURATION NOT RECOGNIZED	94
APPENDIX A: SPECIAL MEASUREMENT MODES		96
APPENDIX B: CONFIGURING POINT-TO-MULTIPOINT FREEWAVE RADIO MODEMS.....		97
APPENDIX C: PREPARING AN OCEAN OPTICS MAYAPRO SPECTROMETER FOR USE IN NOVAC INSTRUMENTS		104

1 Overview

This manual describes the NOVAC Scanning DOAS system (Scanning Dual-beam miniature Differential Optical Absorption Spectroscopy system). The system is designed to measure volcanic gas emissions by UV absorption spectroscopy. It is composed of a miniaturized spectrometer (Ocean Optics Inc.), an embedded PC, an optical fibre, a telescope, and a GPS-receiver.

- The optical spectrometer detects ultraviolet light and converts it to digital numbers. These numbers are the spectra for later gas evaluation.
- The telescope collects the ultraviolet light scattered from aerosols and molecules in the atmosphere. Controlled by the motor, it is able to turn in 360 degrees.
- The optical fibre transfers light from the telescope to the spectrometer.
- The motor controls the angle of the telescope. It moves the telescope certain steps according to a command from the embedded PC.
- The GPS receiver records the position of the system and gives out the universal standard time.
- The embedded PC collects the spectra and manages communication.

Two software packages are needed for the instrument to work. The ‘Kongo’ software runs on the embedded PC in the field. It runs the measurements and saves the data to a flash memory card in the instrument. The amount of data that the standard flash memory can save depends on the number of channels in the spectra and the sampling rate, but typically up to 3 months of data can be saved in the embedded PC. When the flash memory is full, the oldest data is replaced by new data.

The NOVAC software runs on the computer back at the observatory. It connects to the instruments in the field at regular intervals and checks if there is new data to download. If so, it downloads the data, deletes it from the instrument’s flash memory, and performs a real-time preliminary evaluation of the data, and creates folders where data and results files are archived.

The most accurate SO₂ emission rate results are obtained by post-processing the data. This is also done using the NOVAC software at the observatory.

This manual describes the technical setup of the instruments, the software running on the embedded PC and the software running at the observatory. It is updated to describe the 3rd Generation (MOXA computer) and 4th Generation (Axiomtek computer) of the NOVAC instrumentation, but the concepts apply to the older instruments as well. For more details on troubleshooting old instrument versions (Beck and AXIS computers), see the Appendix.

2 System Requirements

The NOVAC program does not need any other drivers or programs. The platform requirements are as follows:

- CPU - Pentium III 550 MHz or higher
- Memory - 64 MB or more
- Operating System - Windows
- Screen Resolution - 1024 by 768 pixels or higher.

3 Setup and configuration of the instruments in the field

3.1 Technical setup

The NOVAC scanning DOAS instruments should be installed at a location on the volcano where the plume is expected to be overhead. It is recommended to install at least 2 instruments within a few km of one another (on the same side of the volcano) such that the plume height can be triangulated when the plume is overhead.

The instrument runs off 12V power. The power consumption is approximately 8W while on. If run for 8 hours per day, the power daily average power consumption is approximately 3W. This does not include the power consumption of the telemetry. A FreeWave HT-PE radio is often used for telemetry. This option uses another approx. 3W while on. An appropriate solar power system or line power should be used to power the instrument and telemetry.

The NOVAC scanning DOAS has undergone several updates since the first instruments were assembled in 2005. Currently, most instruments in the network are either 3rd generation (3rd Gen) or 4th Generation (4th Gen) NOVAC scanning DOAS. Though both instrument versions function equally well and are fully compatible with the NOVAC Program software, updates to the system were required as certain spectrometers and embedded computers were discontinued by their manufacturers. The 3rd Gen NOVAC instruments are controlled by MOXA UC-7112 computers and use Ocean Optics spectrometers. The 4th Gen NOVAC instruments are controlled by Axiomtek IFB122 computers and mostly use Avantes spectrometers.

In the following sections, we describe the 3rd Gen and 4th Gen NOVAC scanning DOAS instruments separately. Check your embedded computer to identify your instrument version and follow the appropriate instructions for setup. There are also a few different types of scanners, but these can be used with any version of the NOVAC instrument regardless of the type. The scanners are described separately in section 3.1.3.

3.1.1 3rd Gen NOVAC instrument (MOXA and Ocean Optics)

Most new installations of NOVAC equipment will use the 4th Gen NOVAC instruments. These contain an Axiomtek embedded computer and are described in the next section. Please skip to that section unless you are looking for information on the older, 3rd Gen instruments that use the MOXA computer.

Version 3 of the NOVAC instrument consists of 3 main parts: The Ocean Optics SD2000 spectrometer, the MOXA UC-7112 embedded PC, and a custom electronics box built by Manne Kihlman (sometimes called ‘Octopus’ for obvious reasons). Though the MOXA is the brain of the instrument, the Octopus is the center of most wiring connections so it’s usually easiest to use this element as the starting point for connections. The following wiring diagram and photographs show how to connect the various components to one another. **Remember that the optical fiber is very fragile and may not be bent more than about a 15 cm radius! Be very careful when connecting the fiber.**

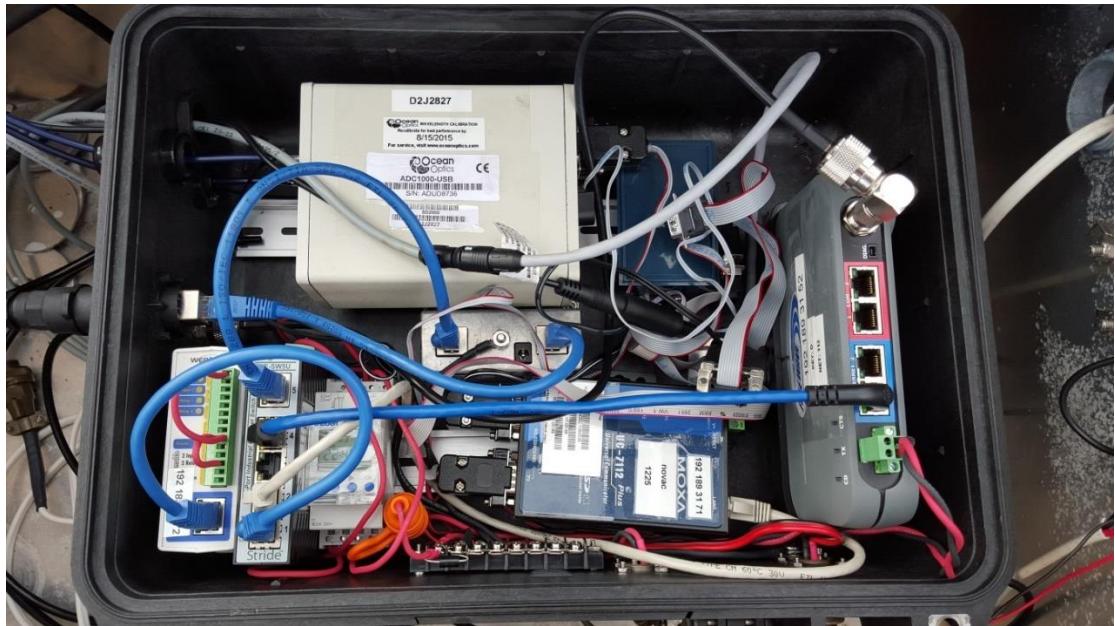


Figure 1- Picture of the NOVAC instrument. The main components are the Ocean Optics SD2000 spectrometer, the MOXA UC-7112 embedded PC, and the custom ‘Octopus’ electronics. Other optional components shown are a Theben TR610top2 timer, a FreeWave HT-PE radio modem, a ControlByWeb Ethernet relay, and a Stride SE-SW5U Ethernet switch.

Wiring diagram NOVAC version 3 (MOXA) instrument

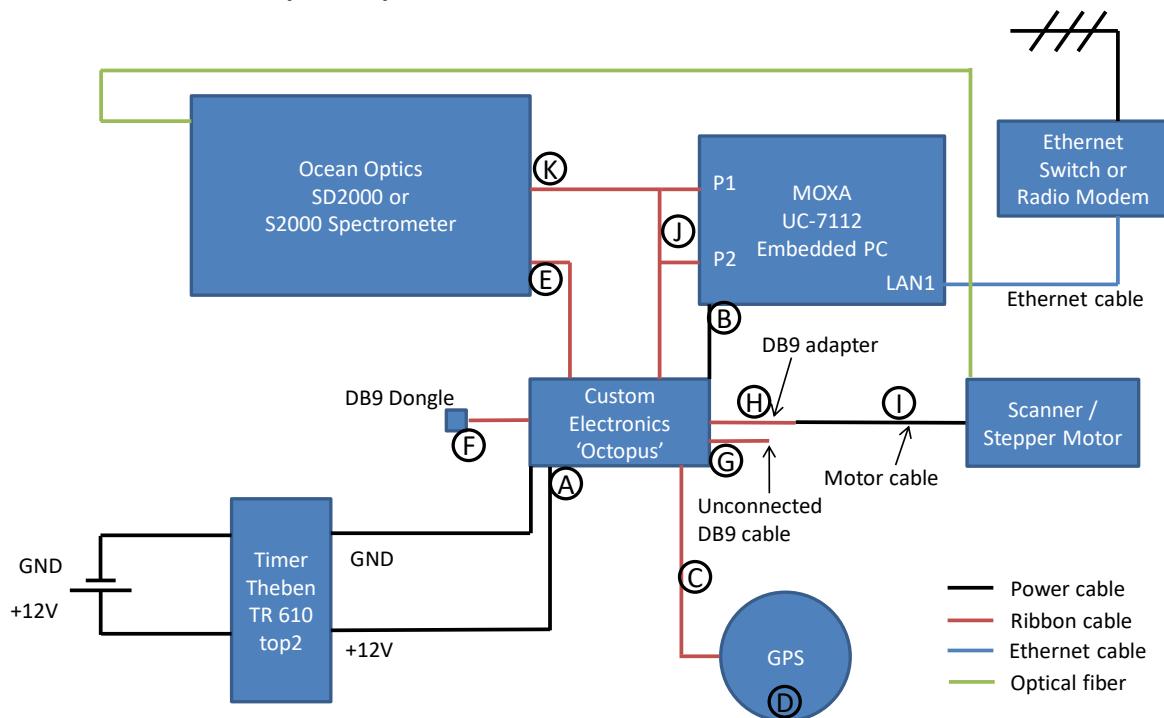


Figure 2 – Wiring diagram for the NOVAC version 3 (MOXA) instrument. The circled letters correspond to the following photographs of respective connectors and components.

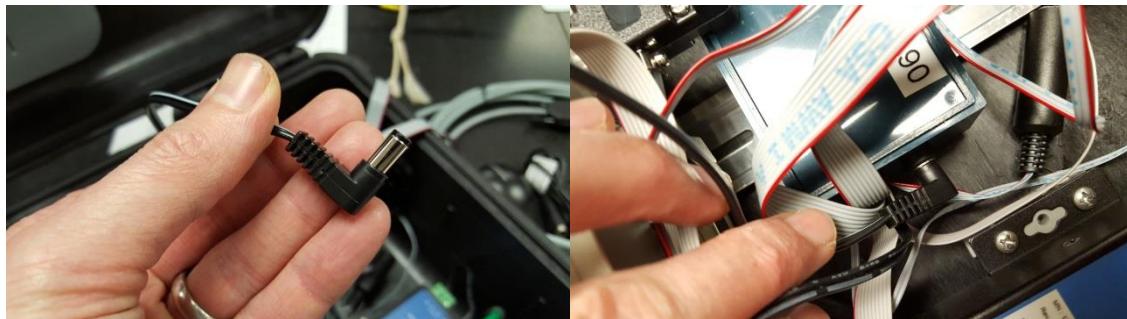


Figure 3 – (A) Power connector supplying 12 V to custom electronics ‘Octopus’



Figure 4 – (B) The Octopus supplies the MOXA embedded PC with power



Figure 5 – (C) PS2 connection from the Octopus to the GPS



Figure 6 – (D) There are 2 types of GPS units currently in use. Both have PS2 connectors.



Figure 7 – (E) The Octopus supplies the Spectrometer with power through a 2 lead ribbon cable.



Figure 8 – (F) One DB9 cable on the Octopus has a dongle on it enclosing a temperature sensor. Don't remove the dongle.

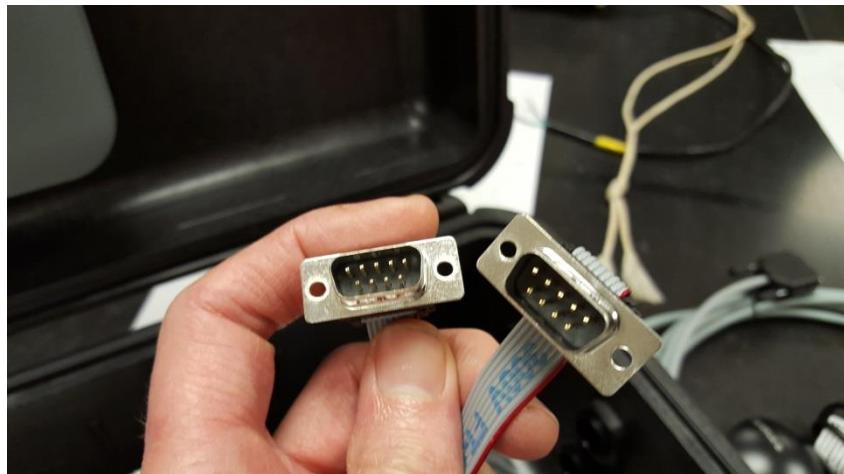


Figure 9 – (G) The Octopus has two identical male DB9 connectors protruding from it. These can be used to power stepper motors. The current NOVAC system (v. 3) only uses 1 motor. The correct DB9 connector must be identified by trial and error. Connect the motor cable to one of them and see if the motor works. If not, try the other connector.



Figure 10 – (H) The motor connector is DB9. This adapter is required to connect the motor cable.



Figure 11 – (I) The motor cable is long enough to reach the scanner (typically on a pole).



Figure 12 – (J) One ribbon cable coming off the Octopus is split into 2 female DB9 connectors, with an additional 3-lead ribbon cable splitting off and going to a male DB9 plug. The two female connectors go into serial ports P1 and P2 on the MOXA. The plug with the extra cable coming off it goes in P1.



Figure 13 – (K) The extra 3-lead cable coming off the MOXA P1 plug goes to the spectrometer. It only fits in the one serial port.

3.1.2 4th Gen NOVAC instrument (Axiomtek and Avantes)

The 4th Gen NOVAC scanning DOAS instrument always uses an Axiomtek IFB122 embedded computer. This computer is more powerful than the MOXA while at the same time also using a bit less power. Unlike the MOXA, it does not have a slot for an SD card but instead has 4GB eMMC internal flash storage. The Axiomtek computer also has a USB interface, which the MOXA did not have.

Since the Ocean Optics SD2000 spectrometers are no longer being manufactured, we tested several other spectrometers to find a suitable replacement. We found that the Avantes AvaSpec-Mini2048CL will provide the best performance per cost in most applications, so this spectrometer is the standard choice for the 4th Gen instruments. The spectrometer makes use of the Axiomtek's USB interface for communication. However, a USB hub is needed as the GPS also now connects to the embedded computer by USB (using a serial to USB converter supplied with the GPS). Tests have shown that the Anker A7516011 is a reliable choice for the USB hub, but other products may also work well.

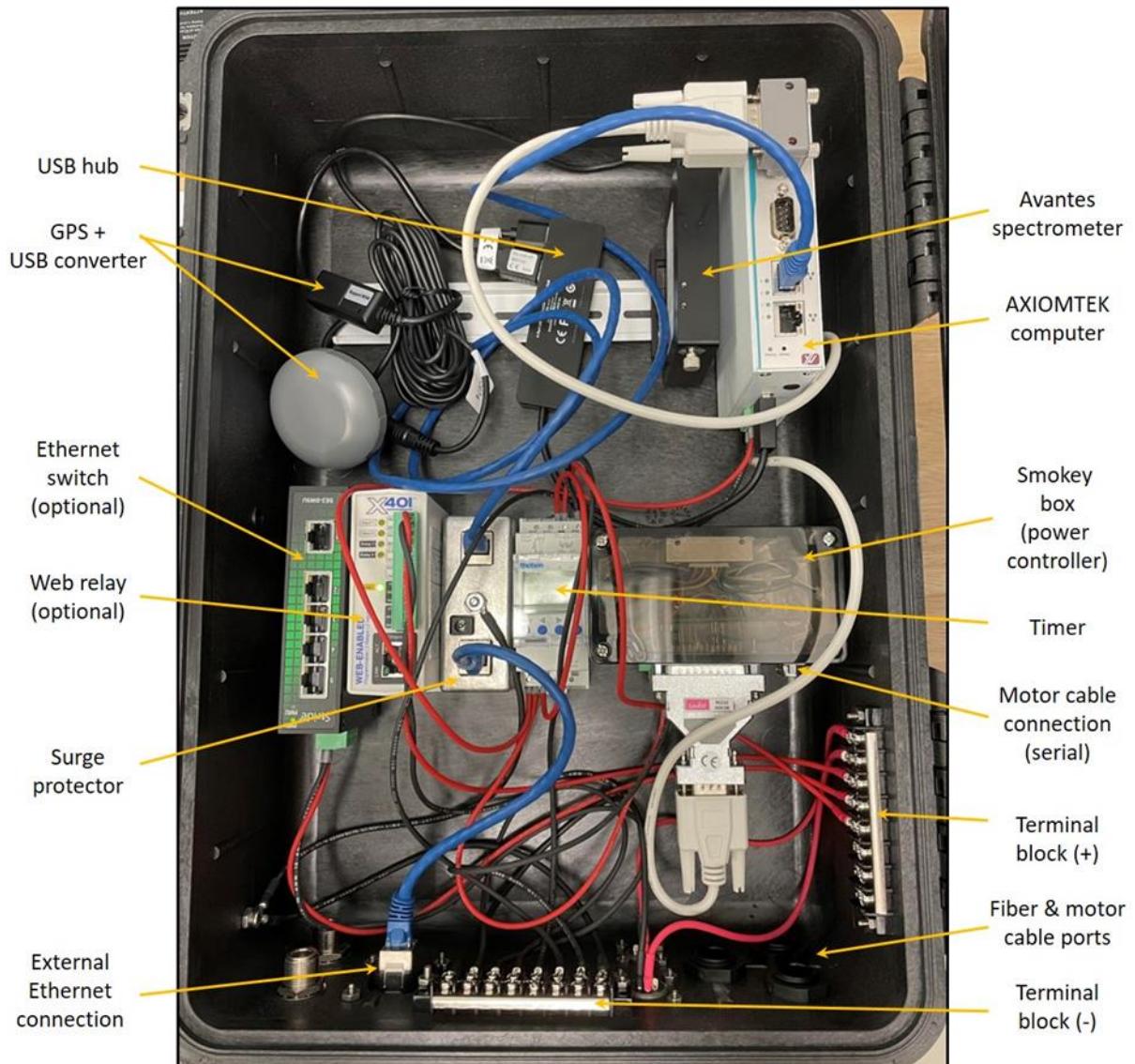


Figure 14 – Inside a 4th Gen NOVAC instrument. Equipment is inside a Pelican case (style 1500 protector case; dimensions: 16.8 × 11.2 × 6.1" interior, 18.5 × 14.1 × 6.9" exterior). Not pictured are input motor cable and 1x-2x fiber optic cables that lead to the scanner. A radio can also be integrated, and the equipment can be used as both a power and Ethernet hub for connecting additional instruments.

The 'octopus' electronics from the 3rd Gen instruments have been replaced by the so-called 'smokey box' electronics in the 4th Gen instruments. In contrast to its multi-tentacled predecessor, the smokey box only has 2 connections: A standard serial cable is used to connect the metallic "taskit" ADC on the smokey box to serial port 2 on the Axiomtek computer. A 90-degree serial connector is used for this purpose. The other connector on the smokey box is connected to the RS232 plug on the scanner motor cable.

The Avantes spectrometer only has a single channel, so wind speed measurements are currently not supported with this spectrometer. If a dual-beam scanner is used, 2 fibers will be connected to it. It is recommended that both fibers be run into the instrument box to protect the fiber ends, but only 1 fiber will be connected to the spectrometer. Either fiber can be used. If a fiber breaks for some reason, simply switching to the other fiber may solve the problem.

As was the case for previous versions of the instrument, a “theben” timer is used to switch the instrument on during the day and off at night. Please check the program of this timer before installing an instrument in the field to assure that the instrument schedule is correct for your time zone and geographic location.

The instrument box may contain some additional components such as a Web Relay and an Ethernet switch. The Ethernet switch can be used to add additional data streams from other, co-located instruments (e.g. seismometers, GPS, cameras) to the same telemetry network. The Web Relay can be configured to allow users to switch the instrument on and off via commands sent through the telemetry network. If not using these components, please unplug them from the power so that they are not unnecessarily using valuable power. Please refer to the manufacturer’s instructions on how to configure these optional components.

In the rare case that an Ocean Optics spectrometer is used with an Axiomtek computer, the Ocean Optics is connected to the Axiomtek via serial connection on port 1, again using a 90-degree serial connector.

3.1.3 DOAS Scanner

The NOVAC DOAS scanner is used to scan the sky from horizon to horizon, scanning a cross-section of the volcanic plume if it is overhead. The scanner is assembled from a number of individual components.

Coming from the spectrometer, one or two optical fibers end in the focal plane of a simple telescope. The telescope contains an optical bandpass filter that only transmits UV wavelengths, filtering out visible and infrared radiation (very little visible light is let through, so it’s not possible to look through the telescope). A quartz lens defines a relatively narrow field of view of about 0.6° (11 mrad), which can be approximated as a single beam.

The telescope is pointed at a scanning element. This is either a mirror or a prism, depending on whether the scanner is a ‘cone’ or ‘flat’ scanner. A flat scanner has a 90-degree scan angle and scans from one horizon to the other passing through the zenith. Flat scanners use a prism to deflect the light into the telescope. Cone scanners have a 60-degree scan angle and therefore surround the volcano a bit during the scan. They never pass through zenith, but rather define a cone around the volcano. The cone scanners use a mirror to deflect light into the telescope. There are also two types of moving scanner hoods, and the correct one needs to be employed for each type of scanner (see images below). The hood rotates together with the mirror or prism; thus light is transmitted through a quartz window in the hood that is always aligned with the internal mirror (prism).

As an alternative to the rotating-hood system, we have also developed a scanner with no external moving parts. Instead of a rotating hood, this scanner employs a piece of quartz tube to allow light to enter the system. Though the quartz tube has a slight widening effect on the instrument’s field of view, the design is particularly useful for applications where fine ash may otherwise cause problems for rotating scanners, or in situations where rime-ice might build up and block a rotating hood. An example of the quartz-tube scanner is also shown in the images below.

If the plume is located overhead, wind measurements can be made by using a ‘dual-beam’ scanner. (**Though this mode of operation is currently not functional in the instrument version 4 (Axiomtek), we hope to restore this functionality in the future.**) Dual-beam scanners use two fibers leading to an SD2000 double spectrometer. During normal scanning operation, only one spectrometer channel (typically the Master channel) is used; if the conditions are right for a wind measurement, the second spectrometer channel (Slave) is also used. The main difference between single and dual beam scanners is that the instrument uses two fiber optics cables, and therefore the telescope has two holes that accept the fiber ends (see pictures below). Dual-beam scanners are by default flat.

When the instrument is powered on, the computer does not know in which direction the scanner is pointing. Therefore, the first thing that the scanner does is find its ‘home’ position. This position is defined by a small reference switch inside the scanner. When the scanner is powered off, it can be rotated by hand. You should be able to hear a little ‘click’ when the scanner passes through the reference switch. Once this switch is reached, the computer knows the scanner is at ‘home’ and the measurements can begin. In the past, the reference switch has sometimes come loose. If this happens, the instrument will no longer be able to find ‘home’, and no measurements will be made. Check to ensure that the reference switch is activated properly. If not, the scanner hood should be removed, the four screws holding the stepper motor plate taken out and the switch adjusted as close as possible to the motor axis to ensure proper contact with the motor switch.

When installing the scanners in the field, the axis of the scanner is typically lined up with the volcanic vent. The scanner hood should be **pointing away** from the volcano. The mirror in the cone scanners will collect light coming from the direction of the volcano only if the scanner is pointing away from the volcano. In the standard configuration of the scanner, the rotation proceeds clockwise as seen from the scanner hood in direction of the volcano.



Figure 15 – NOVAC scanner assembly. This is a ‘cone’ scanner. Note how the cone scanner points back towards the right in this image. This would be the proper mounting direction if the volcano is on the right side of the image. This positioning also holds for flat scanners.

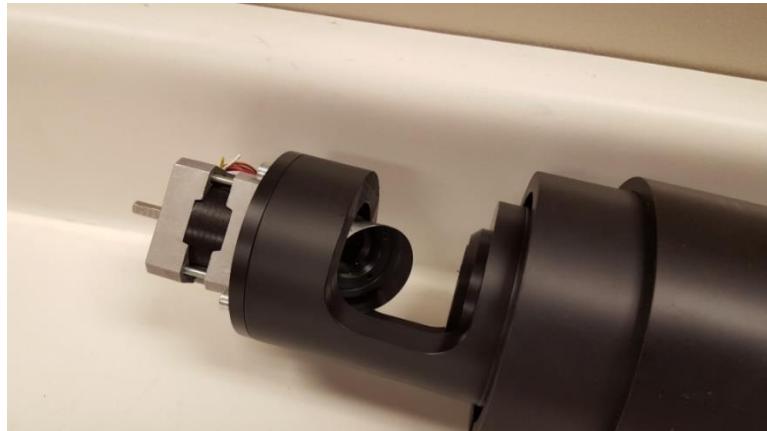


Figure 16 – The optical elements in the NOVAC scanner. A telescope is pointing at a moving mirror. The mirror is mounted on a stepper motor. The motor shaft protrudes out the far side as well – this is where the scanner hood is mounted. On both sides of the motor shaft there are flattened sections in opposite positions. These indicates where the screws of the hood (external shaft) and the mirror or prism (internal shaft) have to be fastened.



Figure 17 – The double scanner (on the left) has two holes for fiber optic cables to be connected and secured with corresponding screws (one of which is visible on the left scanner). The single scanner (on the right) has only one fiber connection and screw. The other holes on the plate are intended for securing it with a special tool.



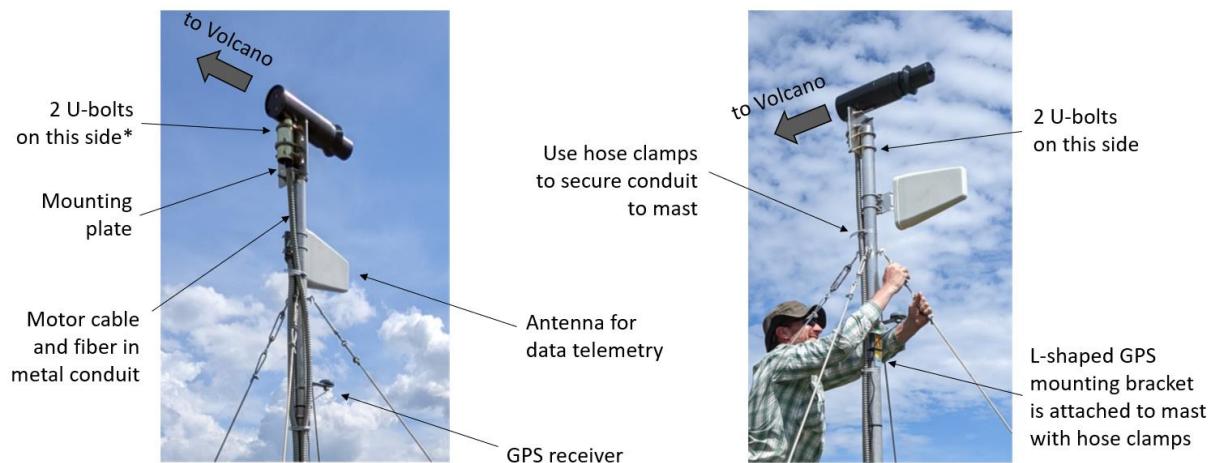
Figure 18 – When switching from a cone to a flat scanner, the mirror must be replaced by a prism and the cone scanner hood must be replaced by a flat hood like the one in this image. Flat scanners are preferable for dual-beam instruments.



Figure 19 – A new scanner designed for locations where rime-ice or blowing ash might cause problems uses a quartz tube in place of the rotating hood. Seen here on Mount St. Helens, USA.

3.1.4 Example setup

To ensure a good view of the volcanic plume and the horizon on either side of it, the NOVAC scanners are typically mounted on a mast of some sort, thus lifting it above any local obstructions such as vegetation or buildings. The scanner must be aimed away from the volcano so that it scans the sky from left to right when facing the volcano (see Figure 20). The telemetry antenna can be mounted on the same mast as the scanner if it does not obstruct the instrument's view of the sky. The rest of the NOVAC instrument will perform best if it's not exposed to the elements (heat, precipitation, volcanic gas) more than necessary. Therefore, the instruments are typically placed in a separate box located at the base of the mast, possibly shaded by the solar panels. It is best to protect the optical fiber(s) and motor cable, e.g. by running them inside a metal conduit or inside the mast itself.



* Alternatively, all 4 U-bolts can be mounted on one side such that the metal part of the scanner is positioned directly over the top of the mast. In this configuration, the conduit is then run down the inside of the mast to the instrument box below.

Figure 20 – Example showing how the NOVAC scanner can be mounted to the top of a mast, with the motor cable and fiber running through metal conduit attached to the mast. The GPS and radio antenna are also mounted to the same mast as the scanner, but not in the scanner’s field of view.



Figure 21 – Test setup of a NOVAC instrument in the parking lot of the Cascades Volcano Observatory. The scanner is mounted on a pole (it would normally be mounted higher to ensure clear view of the sky), and the instrument is in the metal box under the solar panels.



Figure 22 – Recently, the FreeWave HT-PE radio modems have been shown to work very well with the NOVAC instruments. These radios offer two Ethernet and two serial ports, so additional instrumentation such as a seismometer or webcam can be collocated with the NOVAC scanning DOAS and telemetered over the same radio.

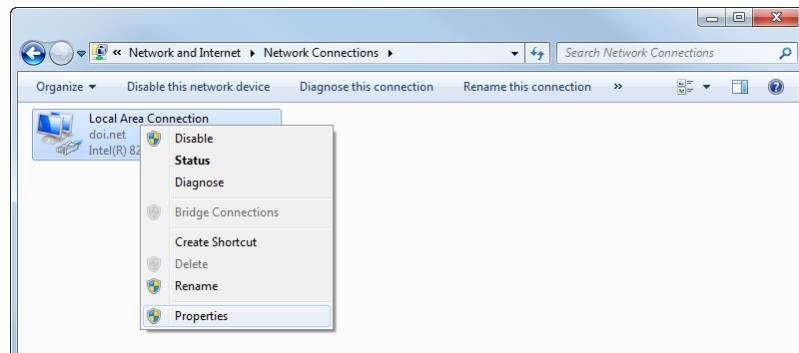
3.2 Configuring the instrument in the field

The ‘cfg.txt’ is located on the embedded PC. This file defines how the measurements by the instrument should be done. The cfg.txt file is read by the data collection software kongo.exe. There is no difference in the formatting of the cfg.txt file between the different instrument versions (Beck, Axis or MOXA, Axiomtek), though the file is located in slightly different locations on the different embedded computers. This is detailed below.

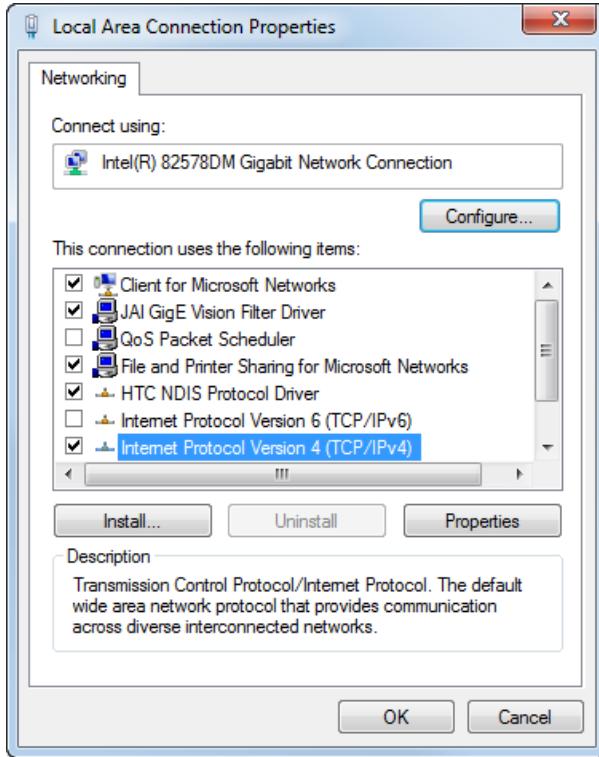
3.2.1 Give your computer a static IP address

To establish a connection to the embedded, the computer you are using needs to be set to a fixed IP address. In Windows 7, this is done in the ‘Network and Sharing Center’ by clicking on ‘Change adapter settings’ in the top left. In Windows 10, this is done by going to “Network status”, then clicking on “Ethernet” in the menu at left. This brings up the menu item “Change adapter options”.

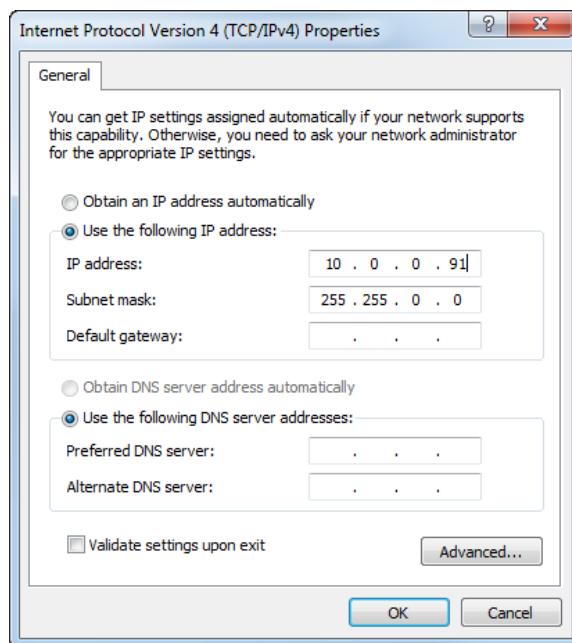
After clicking ‘Change adapter settings’, right-click on the Local Area Connection and select ‘Properties’.



Click on ‘Internet Protocol Version 4 (TCP/IPv4)’ and select ‘Properties’.



Now select ‘Use the following IP address’ and enter an IP address that is on the same Subnet mask as the instrument. **Do not use the IP address of the instrument!** Note that the instrument computer should always be accessible at IP address 10.0.0.90, subnet 255.255.0.0 on LAN2. **These LAN2 settings should never be changed.** If you are connecting to the embedded computer via the LAN2 port, you might set your laptop to the IP address 10.0.0.91 and the Subnet 255.255.0.0



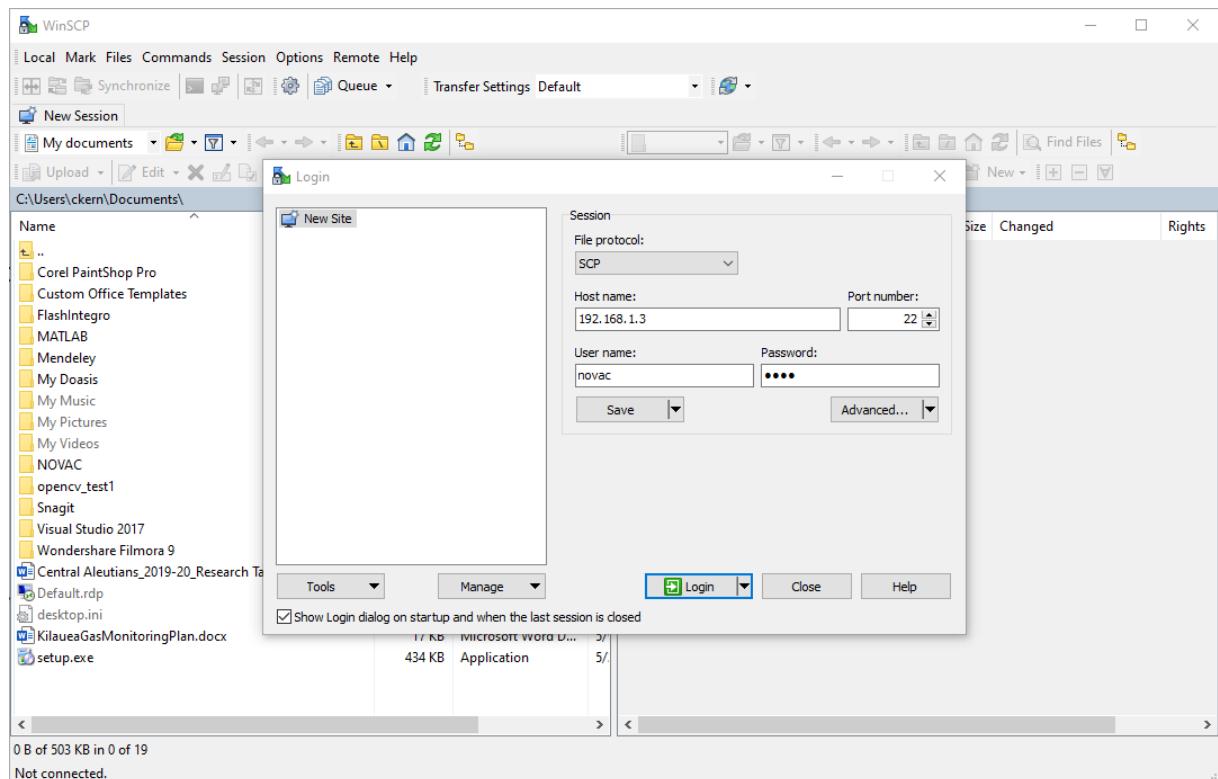
Click OK. You have now finished configuring your computer.

3.2.2 Connecting to the instrument to configure it (with WinSCP)

WinSCP is a free SFTP / SCP client that can be used to access the files on an Axiomtek or MOXA embedded computer. The software is available for download at <https://winscp.net/eng/download.php>.

Plug a network cable into the Axiomtek or MOXA computer and connect your laptop (alternatively, this procedure can be performed over an existing telemetry link).

After starting the WinSCP software, a Login screen will appear. The settings you use for “File protocol” depend on the embedded computer you are connecting to. **For Axiomtek (4th Gen), use SCP. For MOXA (3rd Gen), use FTP.**



In the ‘Host’ field, type the IP address of the embedded computer. If you are connected to the LAN2 port on the computer, the IP address should be 10.0.0.90. The Username is **novac**, and the password is **1225**. Click ‘Login’.

If the connection is successful, a directory listing will appear on the right side of the screen. This is the directory structure of the embedded PC. The left side of the screen shows the directory listing of your laptop computer.

3.2.3 Configuring the network settings (e.g. IP address) of the embedded computer

The network settings for the embedded PC are saved in the file `/etc/network/interfaces`. To configure the network settings of your instrument, first download this file in WinSCP. Next, open the interfaces file with a text editor. Default configurations are shown for Axiomtek and MOXA below:

/etc/network/interfaces example on Axiomtek:

```
# /etc/network/interfaces -- configuration file for ifup(8), ifdown(8)

# The loopback interface
auto lo
iface lo inet loopback

#auto wlan0
#auto ra0
# Wireless interfaces
iface wlan0 inet dhcp
#iface ra0 inet dhcp
    wireless_mode managed
    wireless_essid any
    wpa-driver wext
    wpa-conf /etc/wpa_supplicant.conf

iface atml0 inet dhcp

# Wired or wireless interfaces
auto eth0
iface eth0 inet static
    address 192.168.1.3
    netmask 255.255.255.0
    network 192.168.1.0

auto eth1
iface eth1 inet static
    address 10.0.0.90
    netmask 255.255.255.0
    network 10.0.0.0

# Ethernet/RNDIS gadget (g_ether)
# ... or on host side, usbnnet and random hwaddr
iface usb0 inet static
    address 192.168.7.2
    netmask 255.255.255.0
    network 192.168.7.0
    gateway 192.168.7.1

# Bluetooth networking
iface bnep0 inet dhcp
```

/etc/network/interfaces example on MOXA:

```
#####
## /etc/network/interfaces -- configuration file for ifup(8), ifdown(8)
#
# A "#" character in the very first column makes the rest of the line
# be ignored. Blank lines are ignored. Lines may be indented freely.
# A "W" character at the very end of the line indicates the next line
# should be treated as a continuation of the current one.
#
# The "pre-up", "up", "down" and "post-down" options are valid for all
# interfaces, and may be specified multiple times. All other options
# may only be specified once.
#
# See the interfaces(5) manpage for information on what options are
# available.
#####
#####

# We always want the loopback interface.

auto eth0 eth1 lo
iface lo inet loopback

# embedded ethernet LAN1
iface eth0 inet static
    address 192.168.1.2
    network 192.168.1.0
    netmask 255.255.255.0

# embedded ethernet LAN2
iface eth1 inet static
    address 10.0.0.90
    network 10.0.0.0
    netmask 255.255.255.0
```

The two blocks that begin with *iface eth0 inet static* and *iface eth1 inet static* are where the two LAN ports can be configured. It is our strategy to **never change the LAN2 settings. They should always be kept at IP 10.0.0.90 and Subnet 255.255.255.0.** This way we ensure that whenever we go to a NOVAC instrument in the field, we will always be able to log in on LAN2. The settings for LAN1 can be changed to whatever settings are needed for the telemetry network at a given volcano. **However, always make sure that LAN1 and LAN2 are on different subnets. This means that the third number in the IP address must be different. Setting both Ethernet ports to the same subnet may give unexpected results.**

Once the interfaces file has been modified, upload it to the `\etc\network\` directory on the embedded computer using the same technique as before. Replace the existing interfaces file, but first backup the original file in case you need to revert to it later. Reboot the embedded computer for the changes to take effect.

3.2.4 Configuring the instrument and measurement routine (cfg.txt)

Aside from the instrument network settings (e.g., IP address) which are configured in a separate file (see the previous section), the main configuration of the NOVAC scanning DOAS instruments is set in a file called “**cfg.txt**”. This file contains information on the hardware used in the instrument, defines geometry of the measurement, and sets the measurement schedule, i.e. which directions should be included in each scan. The cfg.txt file is read each time the instrument is powered on.

The cfg.txt file is saved on the embedded computer in each instrument so that the instruments can function independently, even in if data telemetry is lost. The location of the cfg.txt file depends on the version of the instrument. In the 3rd Gen instruments using the MOXA computer, the cfg.txt file is located in the `/mnt/flash/` directory. In the 4th Gen instruments using the Axiomtek computer, the cfg.txt file is located in the `/home/root/` directory.

To make changes to the cfg.txt file, first connect to the instrument using WinSCP. This step is explained in sections 3.2.1 and 3.2.2 of this manual. Then, navigate to the appropriate directory and find the cfg.txt file. Download this file to your local computer and open it with a text editor to view the contents or make any changes. Once you are finished updating the file, upload it back to the embedded computer, overwriting the previous version. It is recommended to make a backup of the file in case you wish to revert to the original version at a later time. Finally, reboot the instrument to apply the changes in the configuration.

Below is a copy of an example cfg.txt file. There are two sections of the cfg.txt file. In the first part of the file, several settings are specified. These will be described in more detail below. The general syntax is SETTING=###, where ### is a numerical value or pre-defined string. The meaning of each individual setting is described below. A line that begins with a ‘%’ is a comment and is not read by the kongo.exe software. The settings indicated below can be used as a default when configuring a new NOVAC scanner.

In the second part of the cfg.txt file, the angles at which the scanner should take measurements are defined. Each entry that begins with MEAS=... represents one position of the scanner at which a measurement will be performed. Details are discussed below.

```
% This is the cfg.txt file
% The settings given below can be used as a default when configuring a new scanner

% The 4th Gen NOVAC instruments mostly use Avantes spectrometers
% If this instrument uses an Avantes, you must set the spectrometer type as indicated below.
% If this instrument uses an OceanOptics SD2000 spectrometer comment out the next line:
SPECTROMETERTYPE=AVANTES

% The following channels defines which channels in the spectra that will be transferred
% This should normally be all pixels on the detector (0 -> 2047)
STARTCHN=0
STOPCHN=2047

% If Realtime=1 then the spectra will be added to work.pak one at a time.
% If RealTime=0 then the spectra will be added to work.pak one scan at a time
% The default is 0.
REALTIME=0

% The Cone-Angle of the system (half the opening angle of the cone)
% Should be 90 if this is a flat instrument and 60 if this is a conical scanner.
CONEANGLE=60

% The geometry: compassDirection tiltX(=roll) tiltY(=pitch) temperature
% Typically, only the first number in this set is changed to reflect the direction to the volcano
COMPASS=123.4 0.0 0.0

% StepsPerRound defines the number of steps the steppermotor divides one round into
STEPSPERROUND=200

% The position of the end switch in the scanner
MOTORSTEPSCOMP=119

% The time delay between motor steps.
% 100 can be used for 4th Gen instruments (with Axiomtek and Smokey Box)
% For 3rd Gen instruments (MOXA and Octopus), use 500.
DELAY=100
POWERSAVE=0

% The minimum solar elevation angle for which to collect data
% The instrument will enter power saving mode and stop scanning below this solar angle.
MINELEVANGLE=20

%If Skipmotor=1 then the scanner will not be used. ONLY FOR TESTING PURPOSES
SKIPMOTOR=0

% The computer will shut down all activity if the voltage of the battery is below this limit
BATTERYLIMIT=10

% Percent defines how much of the spectrometers dynamic range we want to use
% The exposure time will be set accordingly
PERCENT=0.7
```

```

% Channel is the pixel around which the intensity of the spectra will be measured
% a value of -1 means the maximum value in the spectrum
CHANNEL=-1

% The max and min exposure time that we allow the spectrometer to use. In milliseconds
MAXINTTIME=1000
MININTTIME=50

% The debug-level, the higher number the more output will be created
DEBUG=1

% sum1 is inside the spectrometer [1 to 15]
%----pos---time-sum1-sum2--chn--basename---- repetitions--flag
MEAS=0 -1 15 1 0 sky 1 0
MEAS=100 0 15 1 0 dark 1 0
MEAS=-50 0 15 1 0 scan 1 0
MEAS=-48 0 15 1 0 scan 1 0
MEAS=-46 0 15 1 0 scan 1 0
MEAS=-44 0 15 1 0 scan 1 0
MEAS=-42 0 15 1 0 scan 1 0
MEAS=-40 0 15 1 0 scan 1 0
MEAS=-38 0 15 1 0 scan 1 0
MEAS=-36 0 15 1 0 scan 1 0
MEAS=-34 0 15 1 0 scan 1 0
MEAS=-32 0 15 1 0 scan 1 0
MEAS=-30 0 15 1 0 scan 1 0
MEAS=-28 0 15 1 0 scan 1 0
MEAS=-26 0 15 1 0 scan 1 0
MEAS=-24 0 15 1 0 scan 1 0
MEAS=-22 0 15 1 0 scan 1 0
MEAS=-20 0 15 1 0 scan 1 0
MEAS=-18 0 15 1 0 scan 1 0
MEAS=-16 0 15 1 0 scan 1 0
MEAS=-14 0 15 1 0 scan 1 0
MEAS=-12 0 15 1 0 scan 1 0
MEAS=-10 0 15 1 0 scan 1 0
MEAS=-8 0 15 1 0 scan 1 0
MEAS=-6 0 15 1 0 scan 1 0
MEAS=-4 0 15 1 0 scan 1 0
MEAS=-2 0 15 1 0 scan 1 0
MEAS=0 0 15 1 0 scan 1 0
MEAS=2 0 15 1 0 scan 1 0
MEAS=4 0 15 1 0 scan 1 0
MEAS=6 0 15 1 0 scan 1 0
MEAS=8 0 15 1 0 scan 1 0
MEAS=10 0 15 1 0 scan 1 0
MEAS=12 0 15 1 0 scan 1 0
MEAS=14 0 15 1 0 scan 1 0
MEAS=16 0 15 1 0 scan 1 0
MEAS=18 0 15 1 0 scan 1 0
MEAS=20 0 15 1 0 scan 1 0
MEAS=22 0 15 1 0 scan 1 0
MEAS=24 0 15 1 0 scan 1 0

```

```
MEAS=26 0 15 1 0 scan 1 0
MEAS=28 0 15 1 0 scan 1 0
MEAS=30 0 15 1 0 scan 1 0
MEAS=32 0 15 1 0 scan 1 0
MEAS=34 0 15 1 0 scan 1 0
MEAS=36 0 15 1 0 scan 1 0
MEAS=38 0 15 1 0 scan 1 0
MEAS=40 0 15 1 0 scan 1 0
MEAS=42 0 15 1 0 scan 1 0
MEAS=44 0 15 1 0 scan 1 0
MEAS=46 0 15 1 0 scan 1 0
MEAS=48 0 15 1 0 scan 1 0
MEAS=50 0 15 1 0 scan 1 0
```

Here follows a description of all the fields in the cfg.txt file.

STARTCHN= and STOPCHN=

This determines the start channel and stop channel of the spectrum polled from the spectrometer. If both values are zero or if these parameters are not given, then all the channels of the spectrum are used.

Default:

REALTIME=

This setting determines how spectra are saved and transmitted to the observatory. Use REALTIME=0 for version 3 and 4 (MOXA and Axiomtek) instruments.

CONEANGLE=

Defines the cone-angle of the instrument. This can either be 60 (cone scanner) or 90 (flat scanner). This value is written in the header of each spectrum (version 4 and newer). **It is very important to use the correct value here, or the evaluation will give incorrect results!**

COMPASS=

This parameter takes 3 numbers separated by spaces. The first number is the compass direction of the scanner. Normally, this is the direction from the instrument to the volcano. E.g. if the volcano is west of the instrument, the compass direction would be 270.0. An instrument installed on the north side of a volcano would typically have a compass direction of 180.0, since the direction from instrument to volcano is south. Note that when the scanner is mounted on the pole, it protrudes *away from* the volcano, not towards it. However, the compass direction is towards the volcano.

The second and 3rd numbers in this line are usually not used and should be set to 0.0. The correct entry for an instrument installed on the east side of a volcano would be COMPASS=270.0 0.0 0.0

This setting must be correctly configured when installing a new instrument!

STEPSPERROUND=

This gives the number of steps required to make one full turn of the scanning system. For the stepper motors we currently use, the value is 200. Very few NOVAC instruments were built with more than 1 motor. If you are using a ‘Heidelberg’ type instrument with 2 motors, this

string should have two numbers separated by a space. But in most cases, it should be set to
STEPSPERROUND=200

MOTORSTEPSCOMP=

This is the number of steps between the location of the scanner reference switch and the position at which the scanner is looking straight up at the zenith position. Usually it is about 120. This parameter is best set by watching what the scanner does when the instrument power is turned on. When the power comes on, the scanner will first go to the reference switch location. Then, it should go straight up. If it doesn't go straight up, modify the MOTORSTEPSCOMP and try again until it does go straight up. In the usual way that the rotation is implemented, this number should be increased if the motor stops ahead of the zenith position and decreased if it stops behind. If two motors are used (Heidelberg type instrument) then also two values must follow this string. This is rare.

DELAY=

This determines how many milliseconds to pause between each pulse when pulsing the stepper-motor. If this is set to 0 then the fastest possible speed will be used. The DELAY in 4th Gen instruments (with Axiomtek and Smokey Box) can be set to 100. For 3rd Gen instruments (MOXA and Octopus), recommended values are 200 for the flat scanners and 500 for the cone scanners.

POWERSAVE=

In the current version of the instrument, this setting should be set to 0.

SKIPMOTOR=

If this value is not zero, then all motor-activity is skipped. **This should only be used for testing!**

BATTERYLIMIT=

This setting tells the instrument when to shut off if the battery voltage gets too low. Typically, we use 10, which means the instrument shuts off if the battery voltage goes below 10V.

PERCENT=

A decimal number between 0 and 1 should follow this string. This entry determines the desired relative saturation of a spectrum recorded using an automatic exposure time. It should normally be about 0.7, which means the exposure time will be chosen such that the spectral intensity is 70% of the saturation value.

CHANNEL=

This determines the channel where the max-value used for the automatic exposure is determined. The light intensity will be averaged over 10 pixels centered on the channel given by this argument. If CHANNEL=-1 or if it is not defined, then the max-value over all pixels will be used to determine automatic exposure time.

MAXINTTIME=

This is the maximum integration time that the spectrometer is allowed to use, in ms. Usually set to 1000.

MININTTIME=

This is the minimum exposure time that the spectrometer is allowed to use, in ms. Usually set to 10.

MINELEVANGLE=

This is the minimum solar elevation angle for which to collect data. The instrument will enter power saving mode and stop scanning when the sun moves to smaller solar elevation angles than this. Typically set to 20 but may be set lower if measurements in the early morning or late afternoon are desired.

DEBUG=

The higher value given here, the more status messages will Kongo.exe produce for debugging purposes. If DEBUG=0 then no text is written. Usually DEBUG=1 is a good value.

SERVER=

This setting is not usually used. Only used if the instrument is connected directly to the internet. This determines the FTP server that will be contacted when the IP number of the embedded computer changes. For example, if “SERVER= 192.168.1.20 novac 1225” is written in cfg.txt file, the remote PC will contact an FTP server with IP address 192.168.1.20 username novac and password 1225. When IP address is changed, the embedded computer will try to upload a file to this FTP server with the name given by the INSTRUMENTNAME= parameter. If SERVER= is not given at all or if the IP nr=0.0.0.0 then no attempt to connect is made.

FTPTIMEOUT=

This setting is not usually used. This determines the timeout in milliseconds for accessing the server given by the SERVER= parameter.

MEAS=

This determines the different measurement position that should be used.

Typically, many MEAS= lines are following each other with increasing position in the end of the cfg.txt see the example in the beginning of this chapter.

Note: STEPSPERROUND= should always be define before the first MEAS= line.

- pos1, pos 2 – Position in motor steps. The reference position is zenith position.
If only one motor is used, then only one parameter indicating the motor position is expected.

One row of measurement position definition looks like this:

MEAS=	pos1	exptime	sum1	sum2	chn	basename	repetitions	flag
MEAS= 0	-1	1	1	0	up_ang		1	0

If two motors are used, then two parameters indicating the two motor-positions are expected (pos1 and pos2):

MEAS=	pos1	pos2	exptime	sum1	sum2	chn	basename	repetitions	flag
MEAS= 0	0	-1	1	1		0	up_ang	1	0

- exptime – Exposure time in milliseconds. It has three kinds of values.

Positive – Use this value as the exposure time for the spectra collection at this position

0 – Use same exposure time as the first MEAS= line

Negative - an automatic exposure time is calculated for the spectra collection at this position.

The first MEAS= line is used for the zenith measurement and using automatic exposure time.

- sum1 - the number of spectrums added together inside the spectrometer. This number is limited to a value between 1 and 15.
- sum2 - the number of spectrums added together in the embedded computer. This number is limited to a value between 1 and 32767. The total number of added spectrums are sum1*sum2. It is better to use a high sum1 instead of a high sum2 since it reduces the time for transfer spectrum from the spectrometer.
- chn - the channel number for the spectrometer. When a double-spectrometer is used then chn=0 will use the master-channel and chn=1 will use the slave channel. Values above 256 have special meaning, see the documentation for the OceanOptics spectrometers. With a single spectrometer chn=0 should be used.
- basename - gives the name of this direction. This is copied into the header of the spectrum file.
- repetitions - determine how many times this MEAS= line will be repeated. If this value is 0 then it will not be executed at all. This is useful when doing wind-measurements, where the system should stand for a long time taking many spectrums in the same direction. If the repetitions number is missing in the MEAS= line then repetitions will be set to 1. For fixed angle measurements, set repetition between 25 and 49. The NOVAC Program software will interpret fixed angle measurements with repetitions greater than 50 to be wind measurements.
- flag – it is used for new scanner. It determines the state of the solenoids.

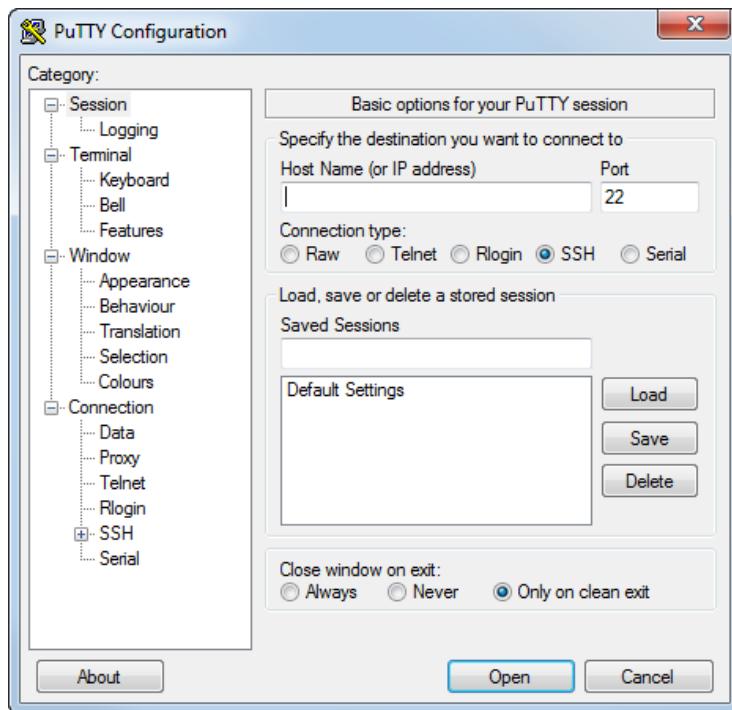
3.3 Troubleshooting or rebooting the instrument with PuTTY, Telnet, or Hyperterminal

3.3.1 Connecting to the instrument with an SSH client

After configuring an instrument (see previous section), it is typically necessary to reboot the instrument. Or perhaps you would like to check what the instrument is doing because it does not appear to be running properly. For this, it is useful to log in to the embedded computer using an SSH client. PuTTY is a freely available software that can be downloaded here:

www.putty.org

To connect to the embedded computer (Axiomtek or MOXA), first give your computer a static IP address on the same subnet as the instrument. This process is described in section 3.2.1 of this manual. Next, connect your computer to the instrument using an Ethernet cable. Alternatively, this process can be performed through an existing telemetry link. When PuTTY is started, you will see the below window:



Type in the IP address of the instrument in the ‘Host Name’ field (default 10.0.0.90 for LAN2 port) on the MOXA or Axiomtek computers. Select ‘SSH’ for the ‘Connection Type’. The Port should be set to 22. Click ‘Open’. You will be prompted for a username and password. All NOVAC instruments have the user **novac** and the password **1225**.

You should now be connected to the embedded computer, where the Kongo.exe program should be running.

3.3.2 Useful commands for troubleshooting

To show the output of the Kongo – program the commands **showlog** and **nolog** are used. Typing them in at the command prompt and pressing Enter should have the following effect:

showlog: giving this command will make the remote PC show the output from the kongo program for some minutes. These messages can be quite helpful for troubleshooting the instrument. For example, if you see the message ‘Spectrometer refuses to scan’, you might check to make sure that the spectrometer is properly connected and has power. It is still possible to type commands while the kongo output is being displayed. If no messages are displayed after giving the command **showlog**, the remote PC might have stopped working. Then one solution is to reboot the remote PC.

nolog: giving this command will stop showing the output from the kongo program.

You can browse the directory structure of the embedded computer and view text files using the following Linux commands:

ls

Shows the files and directories in the current directory.

ls -l

A more extensive list, including file sizes and access rights can be obtained by adding the flag ‘-l’.

more

This built-in program can be used to view the content of a file. E.g. if you want to read the content of the file cfg.txt the command to use is the following (if cfg.txt can be found in the current directory): *more cfg.txt*

rm

This command removes a file found in the current directory. E.g. *rm U001.pak*

reboot

Reboots the remote PC.

cd

Change the current working directory. This can be an absolute path or a relative path. E.g. changing to the data-storage directory: *cd /mnt/flash/novac/*

Changing to the parent directory of the current working directory by the command (notice the space): *cd ..*

3.3.3 Description of important files and their locations

cfg.txt is the configuration file that tells the NOVAC scanner in which directions to measure and for how long. Details of this file are discussed in the previous section ‘Configuring the instrument in the field’. In 4th Gen instruments (Axiomtek), you will find the cfg.txt in the */home/root* directory. For 3rd Gen instruments (MOXA), the ‘cfg.txt’ file is in */mnt/flash/*.

Kongo is the program which controls the whole Scanning DOAS. It downloads data from the spectrometer and creates the file upload.pak upon the completion of one full scan. The kongo software is in */home/*

Upload.pak is a compressed spectra file which is ready to be transferred to our local computer. In 4th Gen instruments (Axiomtek), these data files are stored in */home/root/*. In the previous version (MOXA), they are in */mnt/flash/novac/*

U001.pak, U002.pak, ... these are compressed spectrum files ready to be transferred to our local computer. These are stored in the same directory as upload.pak, or may also be stored in subdirectories labelled *r001, r002, etc.* Every time the instrument is rebooted, all existing pak files are moved to a new rXXX directory.

4 Configuring the NOVAC software at the observatory

The most stable present version of the NovacProgram is v3.4. To know which version of the program you are running, click on ‘Help->About Novac Program...’ in the main window of the software.

You can just copy NovacProgram.exe, the .dll files which are supplied with it (libcurl.dll, libssl-1_1.dll and libcrypto-1_1.dll) and configuration.xml into any folder in the computer (all these files must be copied to the same directory). The program is then ready to run. If you do not have a configuration file, just copy the NOVAC program and create a configuration through the configuration dialog as described below.

Before the NovacProgram can download and evaluate spectra from connected instruments, it is necessary that the configuration is correct. The configuration is stored in the file ‘configuration.xml’ in the same directory as the NovacProgram and is modified through the configuration dialog inside the NovacProgram. This is reached from the main menu in the main-window of the program. Choose ‘Configuration’→‘Configuration’ to open the configuration dialog.

To close the configuration dialog press either of the two buttons ‘Save’ or ‘Cancel’ in the lower right of the dialog. Pressing ‘Save’ will save all changes made, pressing ‘cancel’ will discard all changes made in the dialog. **Note** that the NovacProgram will have to be restarted for changes to come into effect.

4.1 Scanning Instrument Settings

The settings found under the ‘Scanning instrument’ – tab are individual settings for each of the connected scanning instruments in the local network.

In the Scanning Instrument list to the left in the configuration dialog (see Figure 23) choose one instrument. The configurations will be special for that instrument.

If the list of scanning instruments is empty, press the button ‘Add’ below the list (see Figure 23) to insert one more instrument into the list. More instruments can be added to the list by again pressing the button ‘Add’.

4.1.1 Instrument information ‘Info’

Under the ‘Info’ tab are the general settings about the scanning instrument, which are necessary to handle the data coming from the instrument (see Figure 23).

In the Site group, you can set the name of the volcano (choose one from the list), the name of the site where the instrument is located, and the name of the observatory performing the measurements. All fields must be filled in for the program to function properly. If your volcano is not listed, select ‘Other’ and you will be prompted for the source name and coordinates.

In the Spectrometer group, there are settings that give general information about the spectrometer used. Select the model of the spectrometer used. If the spectrometer is not listed, select the ‘Custom...’ option and manually add the spectrometer name and maximum intensity. The serial number is the serial number of the spectrometer, as given by the supplier of spectrometers. This identifies the spectrometer and is necessary for performing evaluations of the spectra. If the spectrometer contains several channels, it is necessary to select the number of channels from the drop-down list for the program to be able to handle all spectra. **Note** it is important that the configuration contain the serial number of all spectrometers used in the local network. Any data coming in from a non-configured spectrometer will not be handled correctly and the data will not be shown in the interface.

In the Instrument group, there is an option to modify the instrument type. Please select Axiomtek (4th Gen), Moxa (3rd Gen), or Axis (2nd Gen), depending on the version of the instrument you are using (where these identifiers refer to the embedded computer inside the instrument).

In the Plot Options group, you can choose to display columns instead of fluxes on the main window (under ‘Last 24 hours’) by checking ‘Plot columns instead of flux’. User may want to select this option for scanners which are configured for fixed view. Users can also choose to display a column history tab by checking ‘Plot Column History’. The Min Column and Max Column are the Y-axis range for the plots on column history tab. Similarly, users can display a flux history tab by checking ‘Plot Flux History’ and adjusting the corresponding min/max flux values.

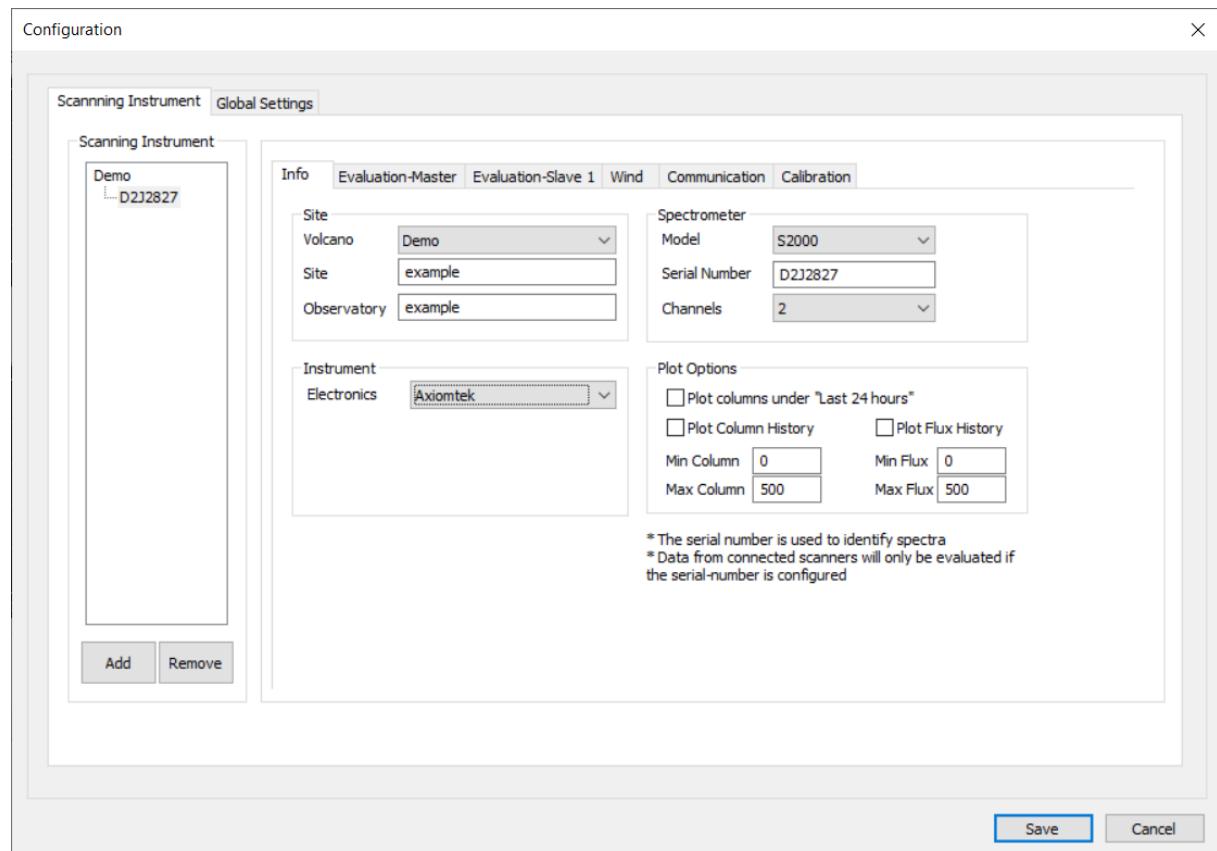


Figure 23 The instrument information page in the configuration dialog.

4.1.2 Evaluation Setting

In the Evaluation-page, you can set the settings for how to perform the real-time evaluation of the spectra as they come in to the observatory.

Note: if these settings are not filled in, no spectra will be evaluated, and no result obtained.

Note: these settings will relate to the instrument which is currently marked in the list of scanning instruments to the left.

For spectrometers with several channels, there is one evaluation-page for every channel in the spectrometer (these will be labelled ‘Evaluation-Master’, ‘Evaluation-Slave1’, ‘Evaluation-Slave2’, etc). Be sure to specify at least one reference-file for each channel.

The fit low and fit high values determine the pixels in the spectrum that will be used to evaluate the spectrums column of gas. These are given in pixels on the detector and are not affected if the spectra are readout as interlaced or only as partial spectra.

To evaluate the spectra there need to be at least one reference-file, typically one for SO₂.

To add a reference-file press the ‘Add reference’ button in the lower part of the dialog.

To remove one reference, select the reference you want to remove in the table and press the ‘Remove Reference’ button.

To inspect or change the properties of an already included reference file press the ‘Properties’ button, this will bring up the ‘Reference Properties’-dialog which lets you change the settings for the references.

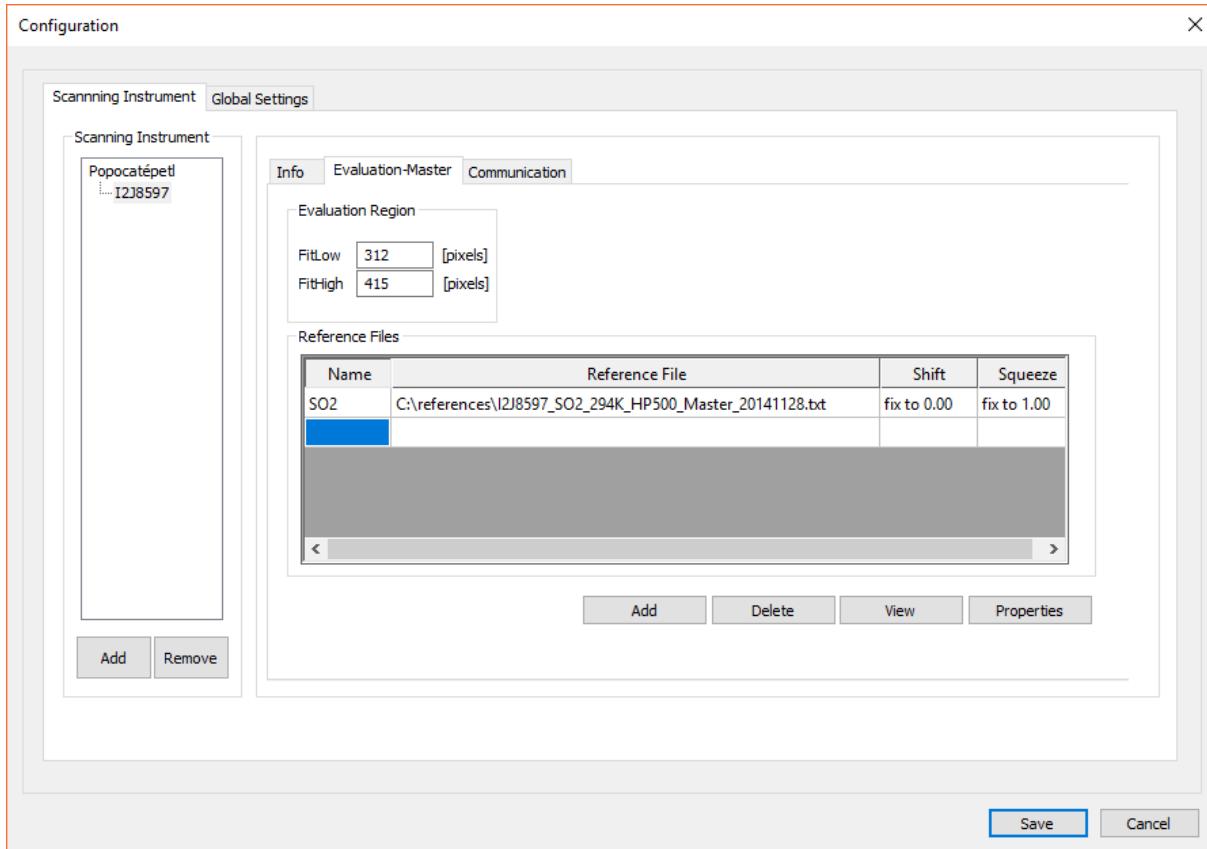


Figure 24. The evaluation settings page in the configuration dialog.

4.1.3 Communication Setting

On the left tree list, select the scanning instrument that you are going to configure. Then you can set the communication parameters for this instrument. There are three communication types to choose from – “Serial Point-to-Point”, “FreeWave – Serial Point-to-Multipoint”, and “FTP”.

In these three communication settings, “Sleep from” and “To” are common parameters, which are used to set the sleeping time for the scanner so that the scanner stops working at night and restart in the next morning. They are common for all the settings. “Sleep From” is the time to quit scanning. “Sleep To” is the time to restart the scanning.

(1) Serial Point-to-Point

The first setting is “Serial Point-to-Point”. This setting is used when there is only one scanner working and the communication is via serial cable or a pair of radio modems.

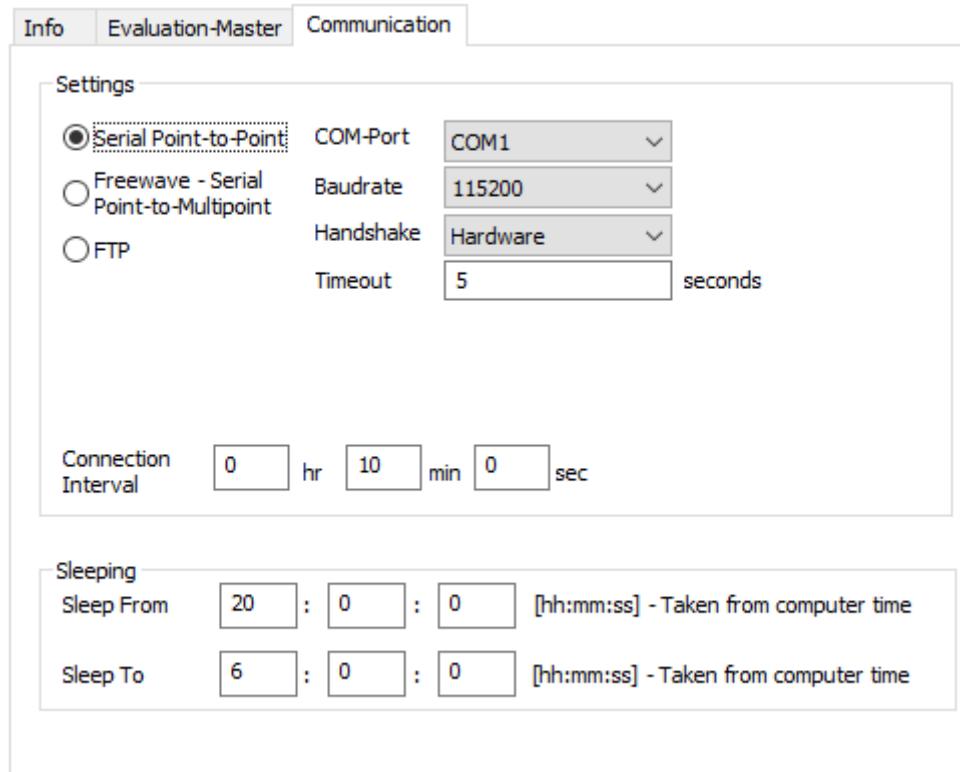


Figure 25. Communication Setting - Serial Point-to-Point

COM-Port – the COM port on the computer that connects the scanner by serial RS232 cable or by radio modem.

Baudrate – the baud rate for the communication between the computer and the scanner or two radio modems which connects the computer and the scanner, in the default setup this is 115200 bps.

Handshake – must be “Hardware” according to the present system.

Timeout – timeout is to set the timeout for the connection. If you find out that connection is not good, you can try to increase timeout value.

Connection interval – the time between two attempts to download files from the instrument. Since one scan takes 5 minutes or more (depending on the conditions), it is good to have a connection interval somewhere from 2 to 10 minutes.

(2) FreeWave – Serial Point-to-Multipoint

The second setting is “FreeWave - Serial Point-to-Multipoint” (it is only applied for FreeWave radio modems). This setting is used when there is more than one scanner working. The computer at observatory communicates with several FreeWave radio modems. In this setting panel, COM-PORT, baud rate, timeout and connection interval have same meanings as the “Serial Point-to-Point” settings.

RadioID / Callbook number – this parameter can be either the ID of the radio modem which connects with that scanner, or the number in the master radio modem’s callbook. It must refer to the configuration of the radio modem settings (more details in Appendix III).

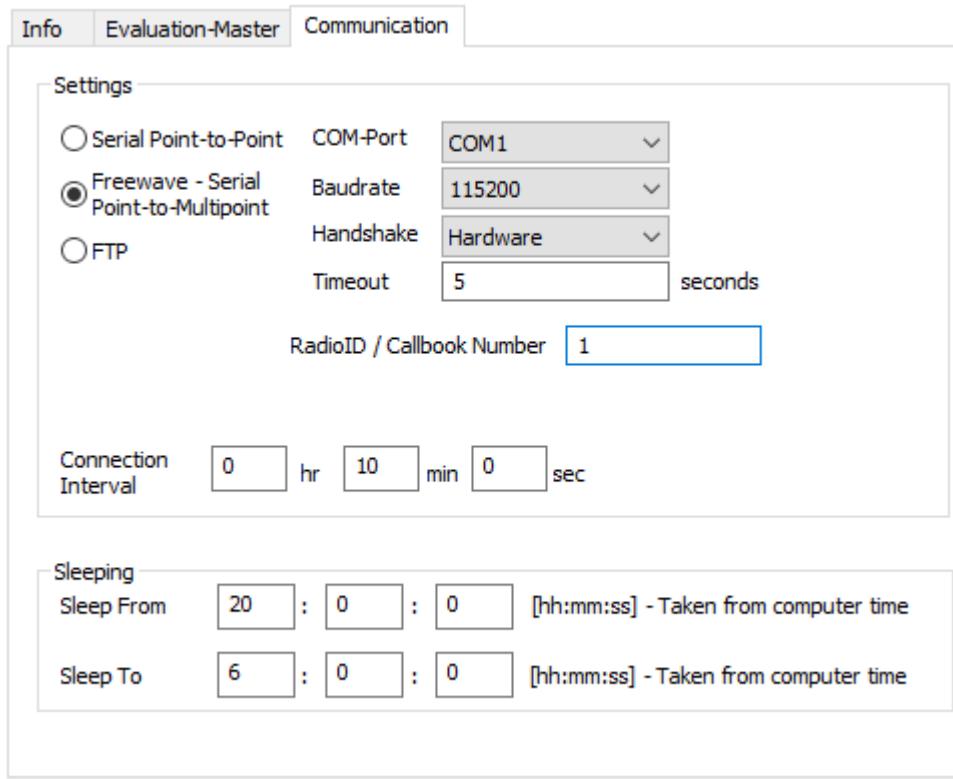


Figure 26. Communication Configuration – Serial Point-to-Multipoint with FreeWave Radio Modem

(3) FTP

This setting is used when the scanner works as an FTP server and when the communication is done through the Ethernet port of the instrument. **This is the default operating mode used with the current version of the NOVAC instruments.**

IP-Address – the IP address of the scanner.

Username – the user name required to login to the FTP server inside the instrument. The default is “novac”

Password – the password required to login to the FTP server inside the instrument. The default is “1225”

Timeout – timeout is to set the timeout for the connection. If you find out that connection is not good, you can try to increase timeout value.

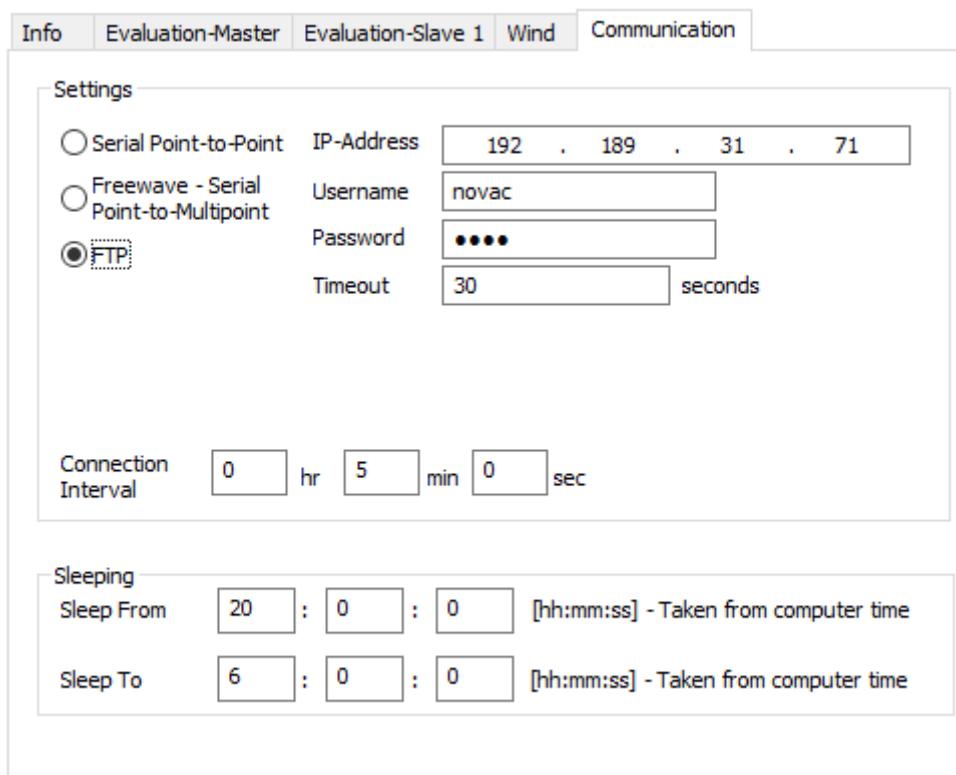
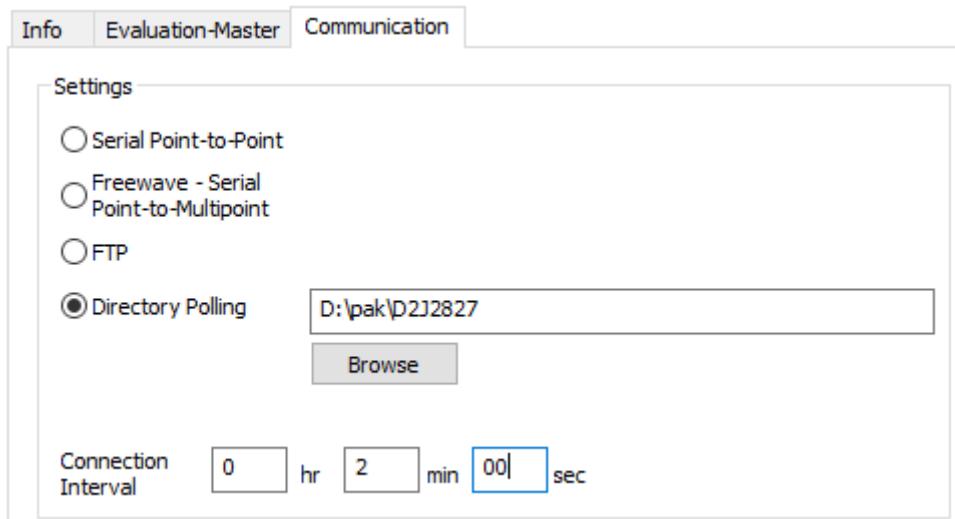


Figure 27. Communication Setting – FTP

In the current version of the NOVAC instrument, the Ethernet configuration of the embedded PC is performed by editing the ‘interfaces’ file in the network directory. See section 3.2.3

(4) Directory Polling



This is a new feature introduced in version 3.3. It allows users to poll pak files from a directory in cases where it is desirable to perform the acquisition of the pak files from the instrument computer separately by an external program. To use this feature, select the

“Directory Polling” option. Then browse for the directory where the pak files for the scanner will be sent so that it is populated in the text field next to the ‘Directory Polling’ radio button. It is recommended that each scanner has a separate directory that is checked to prevent potential conflicts of file names from different scanners having the same file name.

Pak files in the directory are expected to have the “.pak” extension with up to 9 characters for the file name excluding the extension portion (e.g. u001.pak, 123456789.pak). The program will also look for files in subdirectories that begin with “r” up to 10 chars long (e.g. r001, r123456789). The program is not case sensitive.

While pak files are removed when processed (they go to the Output folder just like it does in other modes), the subdirectories are not deleted even when emptied. This is because the program cannot know when the external program is done writing to the subdirectory. It will be up to the external program to clean up subdirectories no longer in use.

The external program performing the acquisition and writing to the polling directory should also be aware of the possibility of a duplicate file and subdirectory names. On the instrument computer, new subdirectories are created each day and upon restart. The pak file names in the subdirectories may be the same as pak file names in other subdirectories. Similarly, once a subdirectory is deleted, it is possible for the instrument pc to reuse that subdirectory name later. For this reason, it is recommended that the external program not overwrite any subdirectories or files already in the polling directories. If necessary, rename the file or subdirectories when copying (e.g. if r001 already exists, rename to r001_1 and copy). The extra characters allowed in file and directory names are to support this.

The Connection Interval in this option will determine the sleep time between directory checks for new pak files. Consider how often data is available for each scanner when configuring this setting.

4.1.4 Wind Configuration

Unfortunately, wind measurements are currently not available in the current instrument version. However, we hope to re-establish this functionality in the near future. Once this is resolved, the below instructions will apply.

In systems where the spectrometer has two or more channels configured, there will also be one ‘Wind’ page in the configuration dialog. If the system is such that the channels are configured to be able to make wind speed measurements, the program is able to automatically tell the NOVAC Scanning instrument to start these wind-measurements and to evaluate the result from them.

The check-box ‘Do wind speed measurements automatically’ enables or disables the use of automatic wind-speed measurements. This is by default turned off.

In this page it is possible to set how often the automatic wind-speed measurements should be performed (by default every 60 minutes) and how long each measurement should be (by default 20 minutes).

There are some conditions for when to perform the wind-speed measurement. No wind measurement can be performed by an instrument which does not see the plume from the volcano.

In addition, the plume must be close to the zenith-position of the instrument for the measurement to give any result. It is in this page possible to specify how many degrees from zenith the centre of the gas-plume can be for the system to perform automatic measurements of the wind-speed.

It is also required that the position of the plume is relatively stable, i.e. there cannot be large fluctuations in the wind-direction when performing wind speed measurements. An automatic wind-measurement is therefore only started if the position of the centre of the plume has not changed considerably during some period. By default, this period is set to 3 scans.

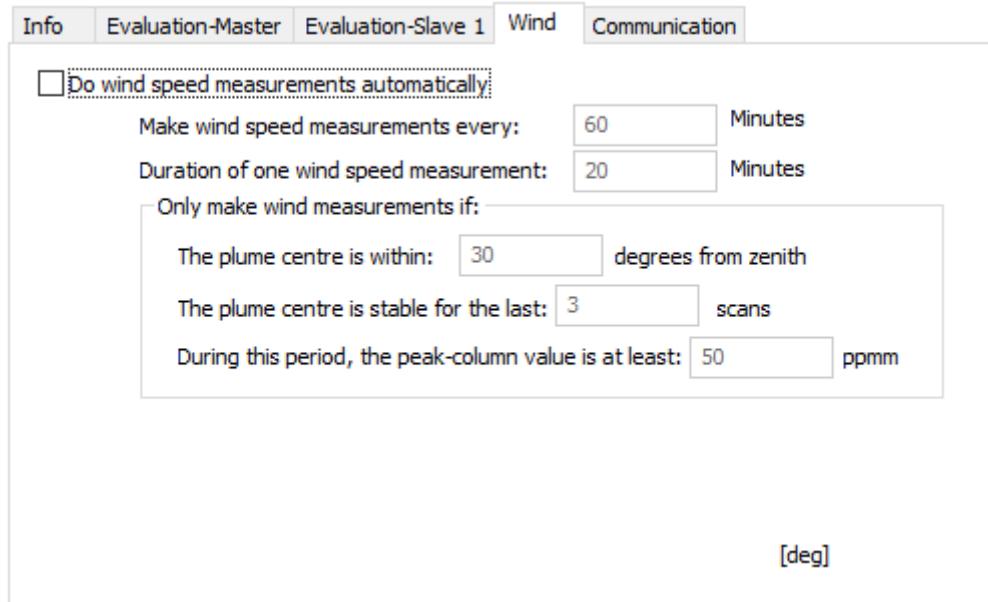


Figure 28. The Configuration page for automatic wind-measurements in the configuration dialog. This is only available for systems with one spectrometer with two or more channels.

4.1.5 Instrument Calibration

From version 3.4 of the NovacProgram, there is also a tab named ‘Calibration’ in the configuration. This allows for performing automatic calibrations of the instrument at regular intervals. The instrument calibrations can optionally be used to create new references for the instrument, replacing the settings in Figure 24 above.

To enable instrument calibrations to be performed, check the check box titled *Enable automatic instrument calibration from measured spectra* on of the dialog page. This will enable the controls below. Notice that calibrations are only performed on the master channel of a spectrometer containing multiple channels.

Initial Calibration

Click on the button *Browse* next to this text box to select an initial calibration which can act as a starting point for the calibration procedure. Clicking on this button will bring up the dialog seen in Figure 63. There are two options for input here

1. *Extended Std File*: This is an STD file which must at least contain a wavelength calibration for the device. This can be saved from an earlier instrument calibration using NovacProgram or MobileDOAS. Only the first field named *Calibration* can be filled in here, the other field will be ignored if it is filled in.
2. *Wavelength Calibration & Measured Instrument Line Shape*: The file selected as the *Calibration* should here point to a file containing a wavelength calibration for the device. This must be a text file containing one or two columns of data where the first (or only) column is the wavelength calibration of the device. A reference file used for evaluation will do here if the file contains two columns of data. The file selected as *Instrument Line Shape* must point to a file containing the instrument line shape in the SLF used by QDOAS, preferably saved from the *Instrument Line Shape* tab in the Instrument calibration dialog. Only the field named *Calibration* is required.

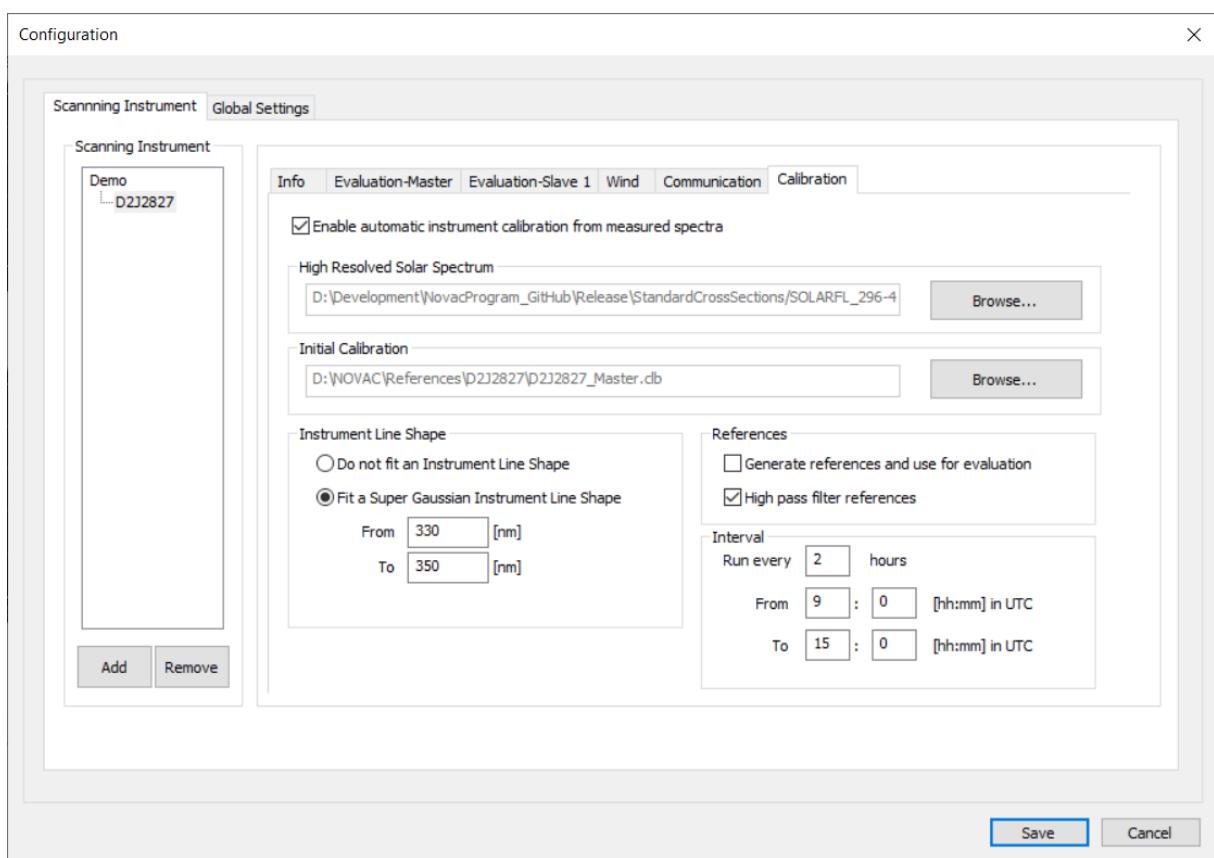


Figure 29. Setting up real time instrument calibration

Instrument Line Shape

This part of the Calibration settings page defines what type of instrument line shape should be fitted to the measured spectrum. It is recommended to fit a super-gaussian instrument line shape unless you are certain that you have provided an accurate instrument line shape when setting up the initial calibration.

- *Do not fit an Instrument Line Shape* select this option if you have supplied an instrument line shape to the initial calibration setup above, either through an extended STD file or through an SLF file. If no initial instrument line shape is supplied and this option is selected, then a gaussian instrument line shape will be approximated which

will produce a lower quality result than fitting a super gaussian instrument line shape (the second option).

- *Fit a Super Gaussian Instrument Line Shape* select this option to fit an instrument line shape approximated by a Super Gaussian profile. This is the recommended option.

References

Generate references and use for evaluation: select this option to replace the references used for evaluation (i.e. the setup in Figure 24 above) with the references which are created by the instrument calibration routine. If this option is not selected then reference files will be generated and saved to disk but will not be automatically used for evaluation.

High pass filter references: Check this box if the created references should be high-pass filtered before being used in the evaluation. This is standard in NovacProgram, *do not uncheck this box unless you really know what you are doing*.

Interval

This part of the dialog defines when and how often the automatic instrument calibrations should be performed. It is recommended to only perform the instrument calibrations during times of the day when the sun is relatively high on the horizon and to avoid these for at least an hour after sunrise or before sunset for the best result.

Run every ... hours makes it possible to define how often the instrument calibrations should be performed at most.

From and To: defines the time when to start performing the instrument calibrations in the morning and when to stop in the afternoon. NOTICE these times are not local time but instead UTC since this timestamp is compared to the timestamp from the GPS which is written into the measured spectrum files. It is therefore possible for these two timestamps to wrap midnight, make sure that that *From* corresponds to morning in your local timezone and *To* corresponds to afternoon in your local timezone.

Conditions for when the calibrations are done

When the above dialog has been setup then automatic instrument calibrations will be performed on the sky spectrum in downloaded pak files which fulfill the following conditions:

- The downloaded spectrum file is measured today.
- The timestamp of the sky spectrum in the file lies between *From* and *To* as setup in the Interval section.
- The timestamp of the sky spectrum in the downloaded file differs with more than the configured number of hours from the timestamp of the last calibrated sky spectrum.
- The measured sky spectrum has an intensity which falls in the range 20% to 85% of the full intensity range of the spectrometer (i.e. the spectrum isn't very dark or saturated).

When the calibrations are enabled, downloaded .pak files will be calibrated at the configured intervals. The instrument calibration procedure will output an instrument calibration file (in .std file format) and references for SO₂, O₃ and Ring as well as a Fraunhofer reference spectrum into the output directory with the current date. If the option *Generate references and*

use for evaluation is checked then the SO₂, O₃ and Ring references will be used in the future evaluations.

If the instrument calibration has been setup but you do not see any calibrations being performed, there is a file named *Debug_InstrumentCalibration.txt* saved in todays output directory. This contains helpful information into understanding why calibrations are not performed.

4.2 Global Settings

The ‘Global Settings’ tab contains settings for the properties of the program which are not connected to any individual instrument.

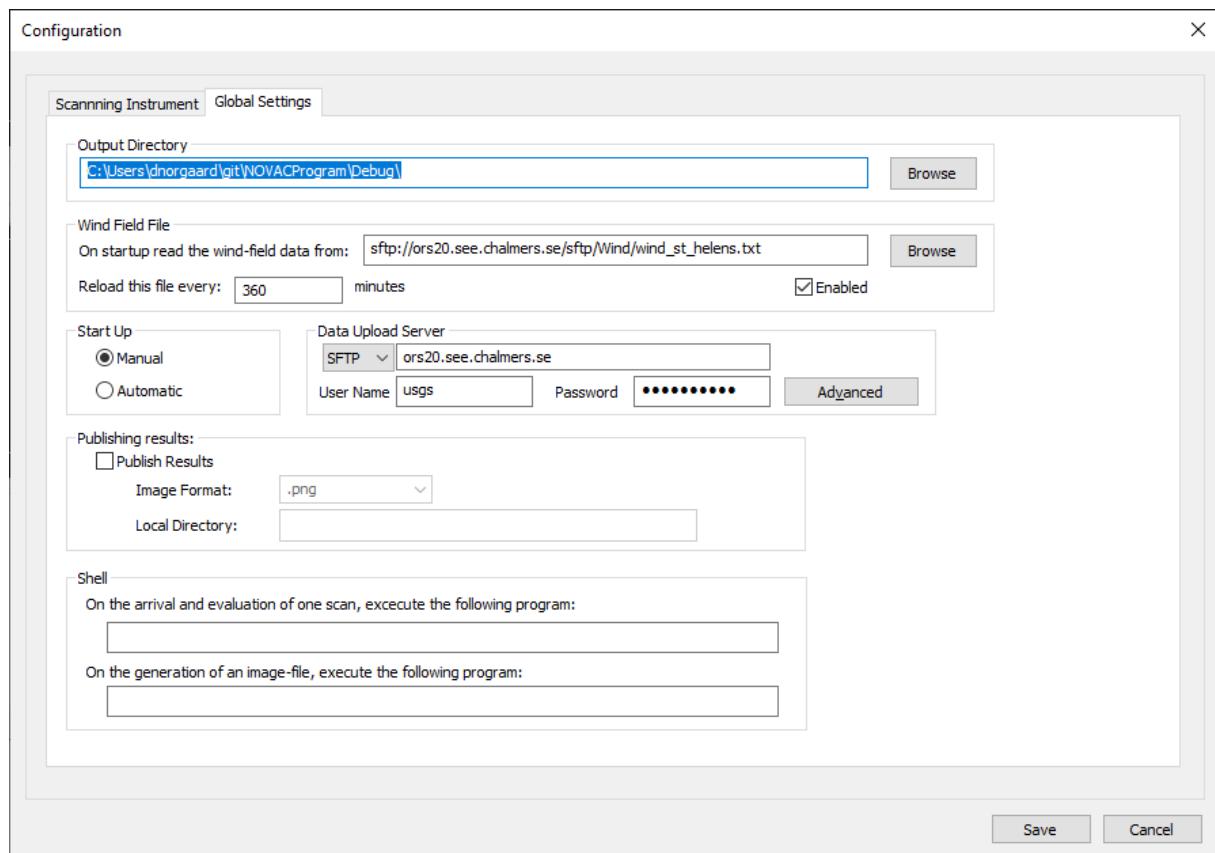


Figure 30. The Global Settings page in the configuration dialog.

Output Directory:

In configuration dialog, the “Global Setting” tab is to set the output directory for all the results that the program gets. All data that the system collects will be saved in this directory, and all temporary files generated will be put here. It is necessary that the user running the program has write-access to the specified directory.

Wind Field File

The NovacProgram can read in a wind-field from a file in a specific format. The wind-field file contains the wind field at specific times of day and for scans collected during the day(s) specified in the file, the flux will be calculated using a wind-field interpolated between the specified times. This file can be located either on the local computer or on the NOVAC SFTP-

server and can be reloaded at any given interval. The location of the file and the number of minutes between each re- reading of this file should be specified in the ‘Wind Field File’ group. For the format of these wind-field files, see the section ‘Wind field file’ below.

Start Up

The program’s start-up-mode can also be specified here in the ‘Start up’ group. If ‘Manual’ is selected the user must press the start button in the toolbar on the main-page to start the measurements. If ‘Automatic’ is selected, the program will automatically start the measurements when started.

Data Upload Server

In the ‘Data Upload Server’ group the user can specify a server to which to upload the spectral and evaluated data in real-time using FTP or SFTP. The username and password to log in to this server must also be specified. This is typically used to upload the data to the NOVAC database in Gothenburg.

Starting in version 3.1, the FTP upload server username and password are no longer stored in configuration.xml. It is stored in ftplogin.xml.

Starting in version 3.2 SFTP protocol is supported for uploads but requires each observatory to acquire an account (login and password) which supports this. FTP uploads will eventually be disabled. If you have not already been contacted with your observatory’s account info, request one by e-mailing dnorgaard@usgs.gov.

To support SFTP, libcurl.dll, libcrypto-1_1.dll, and libssl-1_1.dll are now shipped with the software. This DLL is required to be in the same directory as the NOVAC Program executable for the software to work.

Data Upload Server - Advanced

If the local internet connection is slow or shared with some other instruments, it is possible to restrict the time when the program can upload data. By pressing the ‘Advanced’ button it is possible to define a time interval during which the NovacProgram will try to upload data. The times are in local time of the computer that the NovacProgram is running on.

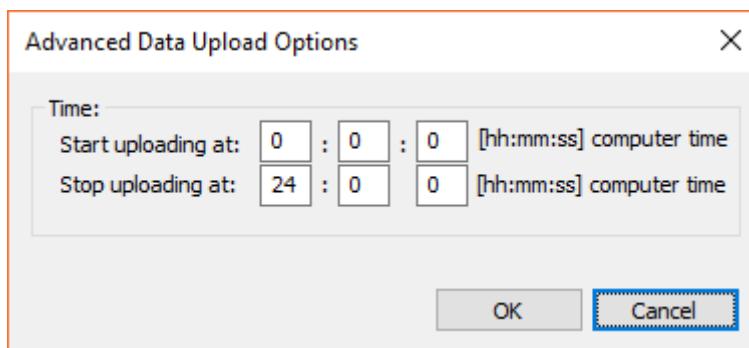


Figure 31. The Advanced Data Upload settings make it possible to restrict the time of day during which the NovacProgram will use the internet connection.

Publishing Results:

The NovacProgram can be made to save an image with the contents of the column plot and the flux plot in the main window each time a scan is received from one connected instrument.

In the ‘Publishing Results’ group the settings for these savings can be specified. The available file-formats for the saved images is Portable Network Graphics (.png), Windows Bitmap (.bmp), JPEG (.jpg) or GIF (.gif). The images will be saved in the directory specified under ‘Local Directory’ which has to be on the local computer. Each image-file will have a file-name which corresponds to the file-name of the saved .pak-files in the output-directory and they will be stored under a sub-folder containing the current date.

Shell:

For reasons of automatization, the NovacProgram can be made to call an external software or script for each time a scan has been downloaded from one connected instrument.

It is possible to specify two scripts/external software:

- One script which will be called when one scan has been downloaded and evaluated. The script will be called with two parameters, the first is the full filename and path of the .pak-file which has been stored in the output-directory, the second is the full filename and path of the .txt-file containing the evaluation results of the scan.
- One script which will be called when one image-file has been generated. The script will be called with one parameter; the file name and full path of the image file. If the check-box ‘Publish Result’ is not checked or an illegal path is supplied under ‘Local Directory’ then this script will never be called.

4.3 Saving the Settings

When pressing the ‘Save’ button in the bottom part of the configuration dialog, the dialog is closed, and the settings are saved to the file ‘configuration.xml’ in the same directory as the NovacProgram. The login (username and password) for the FTP Upload Server is stored in a separate configuration file ‘ftplogin.xml’.

Note: the changed settings will not take effect until the program is restarted!

5 Data collection

There are two ways to start data collection.

- Automatic start. The data collection starts when the program starts. This can be set at configuration dialog. Click menu “Configuration → Configuration”. The configuration dialog will show up. Change the “Start up” under “Global Settings” to ‘automatic’. Restarting the program will automatically start the data collection.
- Manual start. The procedure is as following:

Manually start: Click the menu “Control → start” or click  on the toolbar.

5.1 Running, sleeping

When starting the program, the main window is shown. For each spectrometer that is configured in the configuration file ‘configuration.xml’ there is one page in the main window, where the name of the tag is the name of the site of the instrument.

5.1.1 Main Window

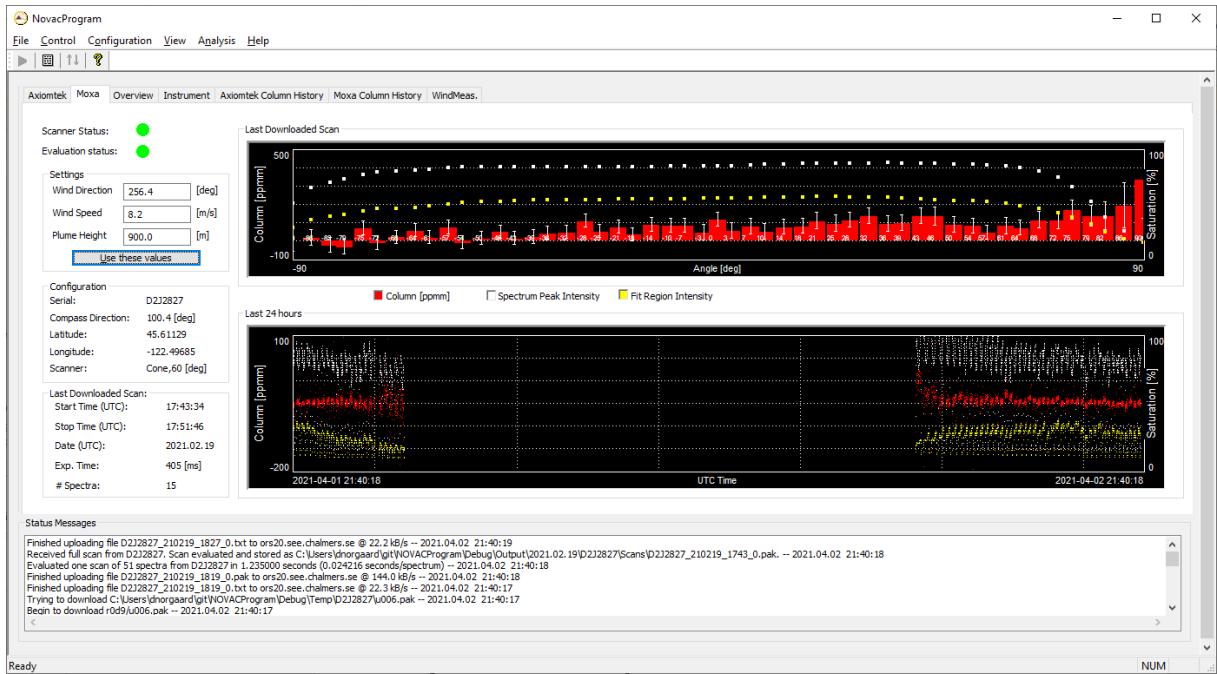


Figure 32 Main window on start up

When the program is running and collecting data, the evaluated results will be shown in the page with the corresponding serial-number and in the ‘Flux Overview’-page.

Note: results from scanning instruments which are connected to the network but not configured through the configuration dialog will not be shown at all.

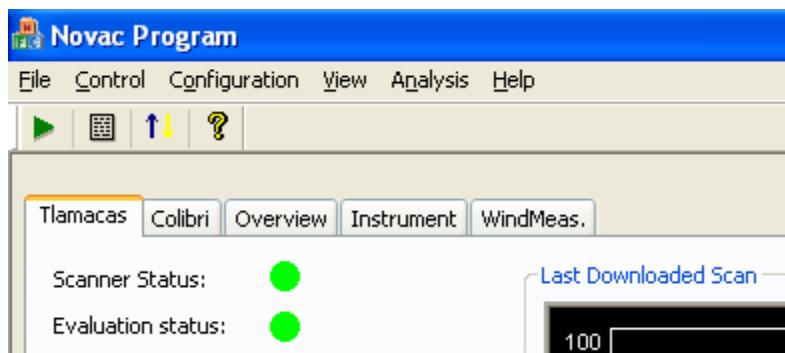


Figure 33. The different pages in the main window. There is one page for each spectrometer configured in the configuration dialog and one ‘Overview’ page which shows a summary of all configured spectrometers. The name of each page is the same as the serial-number of the spectrometer.

The colour of the “Scanner Status” shows the current status of the scanner

- Green light: the instrument is ok
- Yellow light: the instrument is sleeping
- Red light: the program cannot contact the scanner

The colour of the “Evaluation Status” shows the current status of the spectral evaluation

- Green light: the data are ok

- Yellow light: either no spectral data received yet **or** there were some corrupted spectra in the last scan downloaded
- Red light: either the spectral evaluation failed **or** there were several corrupted spectra in the last scan downloaded

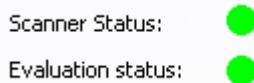


Figure 34. The status indicators in the main window shows the status of the scanner itself and the status of the spectral evaluation.

The program's routine work is to begin to work in the morning and will continuously check the connected scanners for spectrum files. When there is a spectra file named *.pak in one scanning instrument, it will be downloaded.

At night, the NovacProgram stops running according to the sleep time which is set in the communication configuration. It sends message to the scanner so that the scanner will exit at night. When the scanner is not running, the “Scanner status” turns to yellow.

As soon as the spectrum file is downloaded, it will be processed. If the spectra file contains a full scan, it will be evaluated. The evaluation result will be shown on the screen. If the file does not contain a full scan, it will not be evaluated. The incomplete scan will be moved to “Temp” folder under the output directory.

The last spectra file downloaded will be displayed in the ‘Last Scan’ plot area. The red bars are the column values in ppmm of the last measured scan. The white and yellow squares are the maximum intensities of the whole spectrum and the maximum intensity in the fit-region (between fit low and fit high) respectively. The vertical yellow line indicates the software’s guess for the position of the centre of mass of the plume.

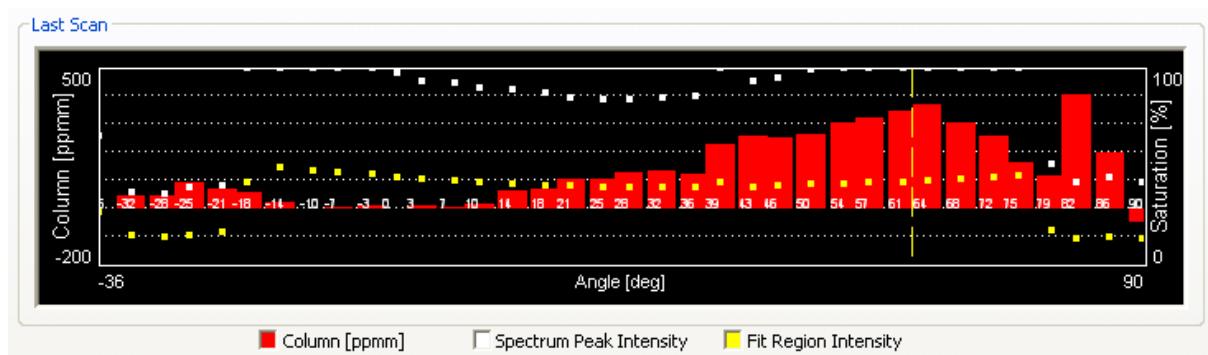


Figure 35. The column plot in the main window.

The ‘Last 24 hours’ plot area contains the flux values from the last 24 hours in UTC time. If the ‘Plot columns instead of fluxes’ option is selected in the Scanning Instrument -> Plot Options configuration, then columns from the last 24 hours will be shown.

The status message box on the window will show the updating messages. The top message is the latest message. All these messages are saved as “Output\DATE>StatusLog.txt” (where

DATE is today's date) under the output directory. The status message box is shown under all tabs.



Figure 36 Status Message Box

5.1.2 Overview Tab

The ‘Overview’ tab shows the flux-results received in the last 24 hours from the different scanning instruments.

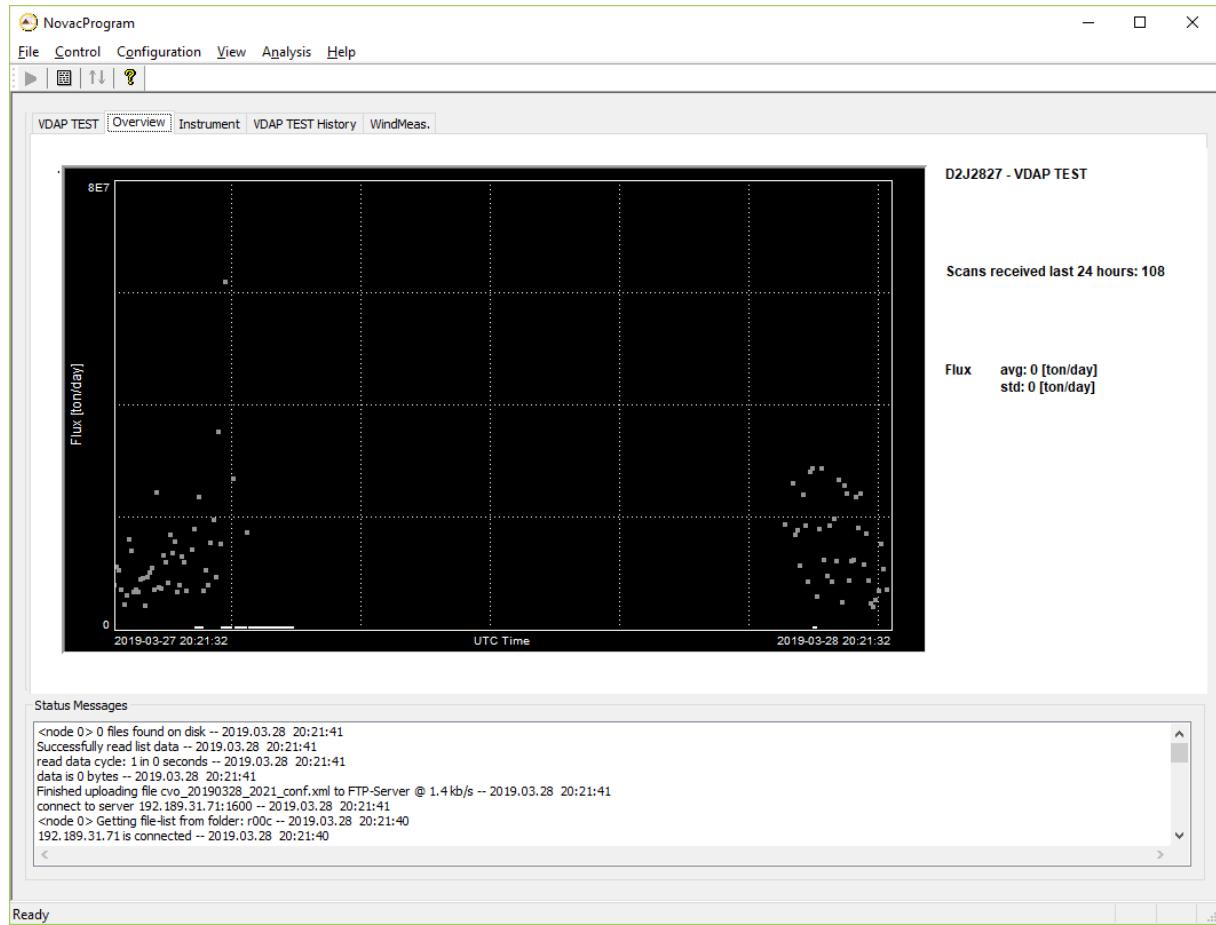


Figure 37 Overview tab

5.1.3 Instrument Tab

There is one page named ‘Instrument’ which shows part of the available instrumental data. Currently the graphs show the temperature, battery voltage and exposure-times today for each of the instruments plus the link speed for each of the spectral file downloads done in last 24 hours.

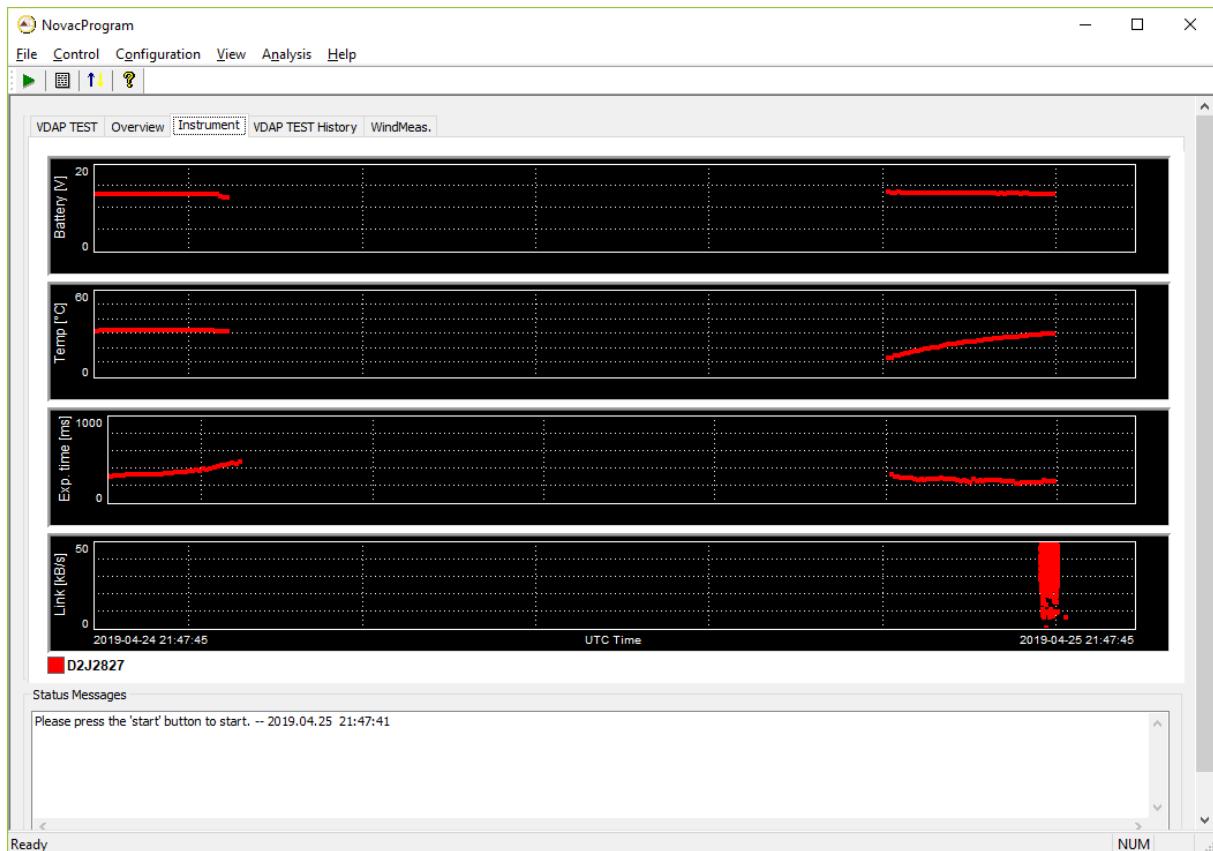


Figure 38 Instrument tab

5.1.4 Column History Tab

For each spectrometer configured, if the ‘Plot Column History’ option is selected in the Scanning Instrument->Info->Plot Options configuration, then a column history tab is shown. It shows column plot for last 24 hours, 10 days, and 30 days. The history data is read from the evaluation logs found in the Output Directory specified under Global Settings configuration. It may therefore take a few seconds to load this tab initially. The ‘Last 24 hours’ plot will update as new data is received. The 10-day and 30-day plots will update once a day when new data is first received on the new UTC day.

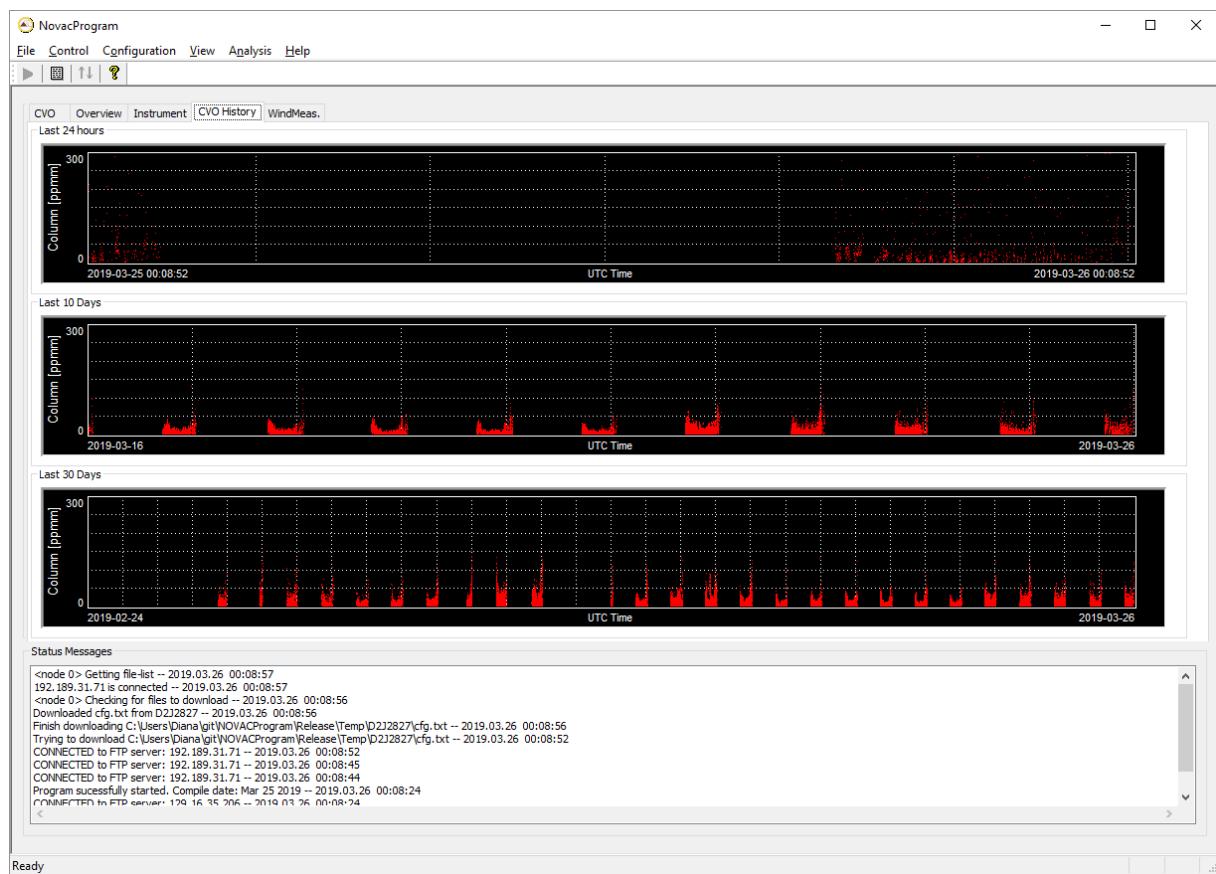


Figure 39 - Column History tab

5.1.5 Flux History Tab

For each spectrometer configured, if the ‘Plot Flux History’ option is selected in the Scanning Instrument->Info->Plot Options configuration, then a flux history tab is shown. It shows flux plot for last 24 hours, 10 days, and 30 days. The history data is read from the flux logs found in the Output Directory specified under Global Settings configuration. The ‘Last 24 hours’ plot will update as new data is received. The 10-day and 30-day plots will update once a day when new data is first received on the new UTC day.



Figure 40 Flux History tab

5.1.6 Wind Measurement Tab

If any connected instrument contains a double-spectrometer, then there is also one page named ‘WindMeas’. The main graph in this page shows the result of the wind-measurements performed today by each of the instruments that can perform wind speed measurements.

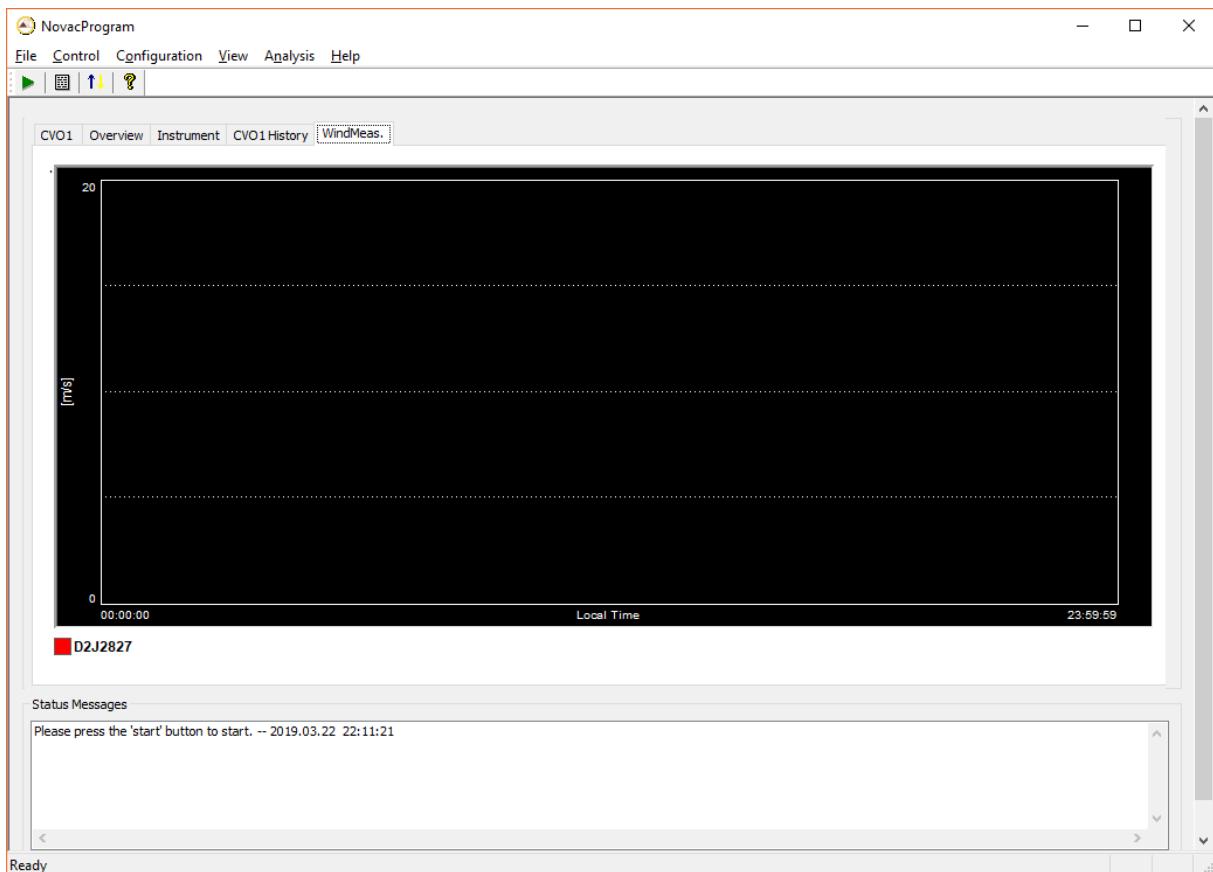


Figure 41 Wind Measurement tab

5.2 Stopping the data collection

Click the menu “Control->Exit” or the red button with a white cross in the upper right part of the window.

5.3 Explanation of messages in NOVAC software

The meaning of the messages shown in the status message box of the NOVAC Software can be found here. They are listed in alphabetic order.

Message	Description	Solution
All instruments sleep for 10 hours. The scanner will start working at 12:00:00	The scanners will stop working for 10 hours. At 12:00:00 they will wake up.	—
Begin to download STATUS.DAT	Begin to download STATUS.DAT file. This file is to check the working status of the scanner.	—
Begin Evaluation of Spectrum File	One downloaded spectrum-file will be looked at. If it contains complete scans, they will be evaluated	—
Cannot switch to shell	Cannot switch to command mode. The communication might be broken. Or the	Restart the program. If it cannot recover, use

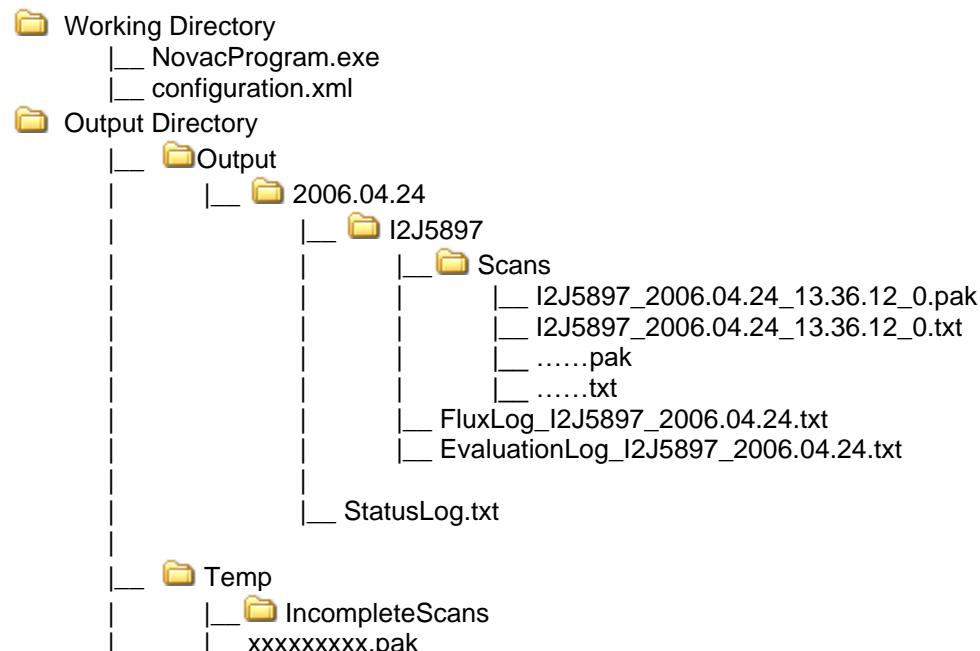
	scanner has no power.	HyperTerminal to see whether the scanner is running.
Check file size	Check the size of the file to be downloaded	—
Could not calculate flux, no good data points in measurement	The spectra in one scan have too low intensity. The flux in this scan is not calculated and is set to 0.	—
Could not read sky-spectrum in file...	There was no spectrum in the scan-file which is a sky-spectrum. The scan-file is corrupt	If problem persist, consider re-formatting the Compact Flash card in the embedded computer
Could not read dark-spectrum in file...	There was no spectrum in the scan-file which is a sky-spectrum. The scan-file is corrupt	If problem persist, consider re-formatting the Compact Flash card in the embedded computer
Downloaded file does not contain a complete scan	There was one incomplete scan in the downloaded file. The incomplete scan will be moved to the 'Incomplete' folder in 'Temp' and later to 'Lost' if no matching end of the scan can be found	—
Download finished	File is downloaded.	—
File size of upload.pak OK	The downloaded file has correct size, which means that it is a complete file.	—
Found spectra which are not evaluated. Will evaluate it now.	The program has found downloaded spectra files at start up. The spectra will be evaluated	—
Have exited Kongo. Will wake up at 12:00:00	The scanner has stopped working. It will be restarted at 12:00:00.	—
Initiated serial port	The serial port is initiated.	—
Kongo is running	The scanner is working.	—
Kongo has exited	The scanner has stopped.	—
Moved incomplete scan file to folder for lost scans	The program has given up on finding the continuation of one downloaded incomplete scan. The file will be moved to the 'Lost' folder	—
Program successfully started	Program is started. But it doesn't mean communication is initiated now.	—
Received new scan from: I2J5897	One scan from I2J5897 spectrometer is evaluated.	—
Received full scan from %s. Scan evaluated and stored as %s	One full scan has been received. The evaluation was successful.	—
Received incomplete scan from %s. Scan evaluated	One incomplete scan has been received. The scan was evaluated anyway and the	—

and stored	evaluation was successful.	
Reboot remote PC	Reboot remote PC so that the scanner starts to work	—
Scan evaluated, the scan file was moved to the storage folder	The downloaded spectra is evaluated. Scan files will be moved to folder "Scan"	—
Set remote PC time	The program will set remote PC time.	—
set remote PC to local PC time 2006-4-25 09:46:10	The program has set remote PC time same as local computer's time.	—
Spectrum XX in pak file is corrupt, checksum mismatch	There was an error in one of the spectra in one downloaded spectrum file. This spectrum cannot be used.	If problem persist, consider re-formatting the Compact Flash card in the embedded computer
Switch remote PC to command mode, No. 0	Switch remote PC's mode to command mode so that command can be received.	—
UPLOAD.PAK is not found	UPLOAD.PAK in the scanner is not ready to download yet.	—
Wake up remote PC	Restart the scanner to begin to work.	—
Will sleep 4.4 seconds	The program will query the scanner in 4.4 seconds.	—

5.4 Output Files

5.4.1 Local Folders Structure

The program works under the output path that you have set in “Global Setting” configuration dialog (see Figure 30). Under this folder the program generates three folders – “Output”, “Lost”, and “Temp”. For example, the serial number of the spectrometer inside the Scanning DOAS is I2J5897. The working directory structure is as follows:



```

|   __ STATUS.DAT
|   __ cfg.txt
|   __ fileList.txt
|   Lost
|       __ xxxxxxx_0_0.pak
|       ...

```

“Output” – stores all the output data files and log files. In this folder, there are folders named by the dates, which stores log files – Evaluation_ErrorLog.txt, and StatusLog.txt. There is also a folder named by serial number, for example I2J5897. That folder stores all uncompressed spectra.

“Temp” – stores temporary incomplete scans to be completed, STATUS.DAT and cfg.txt.

“Lost” – stores incomplete scans which are not possible to be recovered.

5.4.2 Evaluation log

The evaluation result of all spectra collected for one day are stored in the Evaluation log file. The name of the evaluation-log contains the date when the spectra were collected and the name of the spectrometer that collected them. The evaluation result of one scan can be found in two locations, in the same directory as the .pak-files are small .txt files with the same name as one .pak-file, except for the file ending. These small text files contains the evaluation result of the .pak-file with the same name. In the parent directory of the ‘Scans’ directory one large Evaluation log can be found, this contains the evaluation results of all scans collected during the day and is merely a combination of all the small text files.

The evaluation log file consists of three segments, the *scaninformation*, *fluxinfo* and *spectraldata*.

5.4.2.1 Scan Information

The scan-information segment contains general information about the scan. This segment is found in the top part of the evaluation log and starts with the line ‘<scaninformation>’ and ends with the line ‘</scaninformation>’. As of version 2.1 of the evaluation log files, the following items can be found in this section;

<i>date</i>	The date when this scan was collected. In UTC.
<i>starttime</i>	The time of day when the collection of this scan began. In UTC
<i>compass</i>	The compass direction of the instrument that was used to collect the scan.
<i>tilt</i>	The tilt of the instrument at the time of the collection of this scan.
<i>lat</i>	The latitude of the instrument. Information taken from the GPS of the instrument. Unit is degrees and decimal degrees.
<i>long</i>	The longitude of the instrument. Information taken from the GPS of the instrument. Unit is degrees and decimal degrees.
<i>alt</i>	The altitude of the instrument. Information taken from the GPS of the instrument. Unit is meters above sea level.
<i>volcano</i>	The name of the volcano that this scan monitored.
<i>site</i>	The name of the site where the instrument collecting this scan was located at the time the scan was downloaded.
<i>observatory</i>	The name of the observatory responsible for the instrument

<i>serial</i>	collecting this scan.
<i>spectrometer</i>	The serial number of the spectrometer that collected this scan.
<i>channel</i>	The model of the spectrometer that collected this scan.
	The channel of the spectrometer used. This is normally 0, however for SD2000 spectrometers, two spectrometers are present in one box and then 0 corresponds to the ‘Master’ spectrometer and 1 corresponds to the ‘Slave’ spectrometer.
<i>coneangle</i>	The cone angle of the instrument that collected this scan. 90 means a flat scanner and 60 is a cone scanner.
<i>interlacesteps</i>	Which pixels of the detector of the spectrometer that was used to collect this scan. A value of 1 corresponds to every pixel of the detector. A value of 2 corresponds to every other pixel of the detector, etc.
<i>startchannel</i>	The first pixel of the detector of the spectrometer that was used to collect the spectra in this scan. A value of 0 corresponds to starting at the first pixel on the detector.
<i>spectrumlength</i>	The length of the spectra collected in this scan.
<i>flux</i>	The calculated flux of this scan, in kg/s. See the section ‘fluxinfo’ below for more information.
<i>battery</i>	The input voltage of the ManneBox at the start of this scan.
<i>temperature</i>	The temperature of the instrument at the start of this scan.
<i>mode</i>	The measurement mode of this scan. Can be either; ‘plume’ – normal flux measurement ‘wind’ – dual-beam wind speed measurement ‘composition’ – composition measurement ‘direct_sun’ – direct sun measurement ‘stratosphere’ – measurement in stratospheric mode.
<i>instrumenttype</i>	The type of this instrument, can be either ‘gothenburg’ or ‘heidelberg’
<i>version</i>	The version of this evaluation log file. Currently equal to 2.1.
<i>softwareversion</i>	The version of the NovacProgram that was used to create this evaluation log file.
<i>compiledate</i>	The date the NovacProgram creating this evaluation log file was compiled.

5.4.2.2 Fluxinfo

This section contains information on the calculated flux from this scan and some of the information that was used to calculate this scan. As of version 2.1 of the evaluation log files, the following items can be found in this section;

<i>flux</i>	The calculated flux from this scan. In kg/s.
<i>windspeed</i>	The wind speed that was used to calculate the flux from this scan. In m/s.
<i>winddirection</i>	The wind direction that was used to calculate the flux from this scan. In degrees from north with increasing numbers clockwise.
<i>plumeheight</i>	The altitude of the plume above the instrument (in meters) that was used to calculate this flux.
<i>windspeedsource</i>	The source of the wind speed used to calculate this flux.
<i>winddirectionsource</i>	The source of the wind direction used to calculate this

<i>plumeheightsource</i>	flux. The source of the plume height used to calculate this flux.
<i>compasssource</i>	The source of the compass direction of the instrument. Can be either ‘user’ or ‘compassreading’

5.4.2.3 Spectraldata

This section contains information of each of the spectra collected in the scan, including the sky and dark spectra. This section contains one line for each spectrum collected in the scan. The header of the table found in this section can be found before the line ‘<spectraldata>’. As of version 2.1 of the evaluation log files, the following columns can be found in this section;

<i>scanangle</i>	Only for instruments of version 1: The scan angle for the current spectrum. Unit is degrees.
<i>observationangle</i>	Only for instruments of version 2: The observation angle for the current spectrum. Unit is degrees.
<i>azimuth</i>	Only for instruments of version 2: The azimuth angle for the current spectrum. Unit is degrees.
<i>starttime</i>	The UTC time when the spectrum collection began.
<i>stoptime</i>	The UTC time when the spectrum collection was finished.
<i>name</i>	The name of the spectrum, from ‘cfg.txt’ in the instrument.
<i>specsaturation</i>	The saturation of the pixel with maximum intensity in the current spectrum. This is a value between 0.0 and 1.0 saying how large portion of the available dynamic range of the spectrometer is used, 1.0 meaning the spectrum is saturated and 0.0 meaning that the spectrum is completely dark.
<i>fitsaturation</i>	The saturation of the pixel with maximum intensity within the range of pixels used in the spectral fit.
<i>counts_ms</i>	The number of counts for the pixel with highest intensity divided by the number of milliseconds of exposure-time. Gives information on the sensitivity of the spectrometer.
<i>delta</i>	The delta (peak-to-peak difference) of the residual of the fit.
<i>chisquare</i>	The chi ² of the fit.
<i>exposuretime</i>	The exposure-time of the current spectrum, in milliseconds.
<i>numspec</i>	The number of detector read-outs that were co-added to generate this spectrum.
<i>column (...)</i>	The evaluated column of the species/cross section specified in the parenthesis. There is one column like this for each trace gas evaluated for in the fit.
<i>columnerror (...)</i>	The column error of the species/cross section specified in the parenthesis. There is one column like this for each trace gas evaluated for in the fit.
<i>shift (...)</i>	The shift of the species/cross section specified in the parenthesis. There is one column like this for each trace gas evaluated for in the fit.
<i>shiftError (...)</i>	The error in the shift of the species/cross section specified in the parenthesis. There is one column like this for each trace gas evaluated for in the fit.
<i>squeeze (...)</i>	The squeeze of the species/cross section specified in the parenthesis. There is one column like this for each trace gas

<i>squeezeError (...)</i>	evaluated for in the fit.
<i>isgoodpoint</i>	The error of the squeeze of the specie/cross section specified in the parenthesis. There is one column like this for each trace gas evaluated for in the fit.
<i>offset</i>	Flag determining of this spectrum has been judged as a good spectrum by the real-time program. 0 corresponds to bad spectrum, 1 corresponds to good spectrum.
<i>flag</i>	The electronic offset of the current spectrum. In counts. Note that this is a sum, i.e. if the spectrum is a sum of 15 read outs then this is 15x the offset of a single spectrum.
	The ‘flag’ that was specified in the cfg.txt file in the instrument for the collection of this spectrum.

5.4.3 GeometryLog

Whenever two complete scans, collected closely in time, see the plume the information in the two scans can be combined to calculate the direction of the plume and the altitude of the plume. The calculated plume heights and wind directions can be found in the GeometryLog file. The file contains the following columns;

<i>Volcano</i>	The name of the volcano that is monitored with the instruments used
<i>EvaluationLog1</i>	The file name of the evaluation log of one of the scans that was used to make this calculation
<i>EvaluationLog2</i>	The file name of the evaluation log of one of the scans that was used to make this calculation
<i>AverageStartTime</i>	The average of the start times of the two scans that were combined.
<i>PlumeCentre1</i>	The scan angle in which the centre of mass of the plume is found in scan 1
<i>PlumeCentre2</i>	The scan angle in which the centre of mass of the plume is found in scan 2
<i>ScannerDistance</i>	The distance between the two scanners that collected the used scans. In meters.
<i>CalculatedPlumeHeight</i>	The calculated altitude of the plume, in meters above the lowest of the two instruments used.
<i>TotalPlumeHeight</i>	The calculated altitude of the plume, in meters above sea level.
<i>PlumeHeightError</i>	A value describing the uncertainty in the calculated plume height. This is a minimum value only!!
<i>WindDirection</i>	The calculated direction of the plume. In degrees from north, increasing clockwise.
<i>WindDirectionError</i>	A value describing the uncertainty in the calculated plume direction. This is a minimum value only!!

5.4.4 FluxLog

The fluxes calculated in real-time are stored in the FluxLogs. The flux-log contains a small header describing the instrument;

<i>serial</i>	The serial number of the spectrometer used to collect the data found in the current flux-log file
<i>volcano</i>	The name of the volcano that these fluxes are for.
<i>site</i>	The name of the site where the instrument is located.

Following the header is a table of calculated fluxes. Each scan generates the one line in the flux log file. The table contains the following columns;

<i>scandate</i>	The date when the scan was collected. In UTC.
<i>scanstarttime</i>	The time when the scan started. In UTC.
<i>scansoptime</i>	The time when the scan was completed. In UTC.
<i>flux_ [kg/s]</i>	The calculated flux of the scan. Unit is kilograms per second
<i>windspeed_ [m/s]</i>	The wind speed that was used to calculate the flux. Unit is meters per second.
<i>winddirection_ [deg]</i>	The wind direction that was used to calculate the flux. Unit is degrees from north, increasing clockwise
<i>windspeedsource</i>	A string describing the source of the wind speed. This can be ‘user’ if given by the user through the interface.
<i>winddirectionsource</i>	A string describing the source of the wind direction. This is typically ‘user’ – when given by the user through the interface.
<i>plumeheight_ [m]</i>	The assumed altitude of the plume (in meters above the instrument collecting the scan) that was used to calculate the flux.
<i>plumeheightsource</i>	A string describing the source of the plume height used. This is typically ‘user’ – i.e. given by the user through the interface.
<i>compassdirection_ [deg]</i>	The compass-direction of the instrument. This is the direction in which the instrument points, in degrees from north, increasing clockwise.
<i>compassource</i>	The source of the compass-direction. Can either be ‘user’ when specified in the cfg.txt file in the instrument or ‘reading’ when read out from a compass attached to the instrument.
<i>plumecentre_ [deg]</i>	The scan angle where the centre of the plume was calculated to be.
<i>plumecompleteness_ [%]</i>	The calculated completeness of the scan. This is a number between 0.5 and 1.0 describing how large portion of the plume is seen. 1.0 means that the entire plume is captured by the scan, 0.5 means that the plume is on the edge of the scan and that the calculated flux is thus underestimating the true flux. This value is however not a percentage of the observed plume, only a parameter that tells if the scan see clean horizons on the

<i>coneangle</i>	sides of the plume or not.
<i>tilt</i>	The cone angle of the instrument. This is either 90, i.e. a flat scanner, or 60 which is a cone scanner.
<i>okflux</i>	The tilt of the instrument that was assumed when calculating the flux.
<i>temperature</i>	Flag describing whether the software judged the scan to be a good measurement or not. 0 corresponds to a bad measurement, 1 corresponds to a good flux measurement.
<i>batteryvoltage</i>	The temperature of the instrument when collecting this scan. This is read out from the temperature sensor connected to the instrument. Value is 0.0 if no sensor is available. Information is supplied for reasons of instrument monitoring.
<i>exposuretime</i>	The input voltage to the ManneBox at the start of the scan. Information is supplied for the reason of instrument monitoring.
	The exposure time used for the sky measurement in this scan. Can be used to monitor the status of the instrument.

5.4.5 StatusLog

A file containing the messages shown in the ‘Status Messages’ list box in the bottom of the main window in the NovacProgram. This file contains useful debugging information and should always be supplied when reporting bugs/strange behaviour of the program.

5.4.6 Wind field file

When re-evaluating scans with the NovacProgram it is convenient to read out the plume parameters (speed, direction and height of the plume) from a single file. This can be called when re-evaluating fluxes for example, by clicking on ‘Analysis->Flux->File->Import Wind Data->Import WindField from file’ and selecting on the browser the respective file.

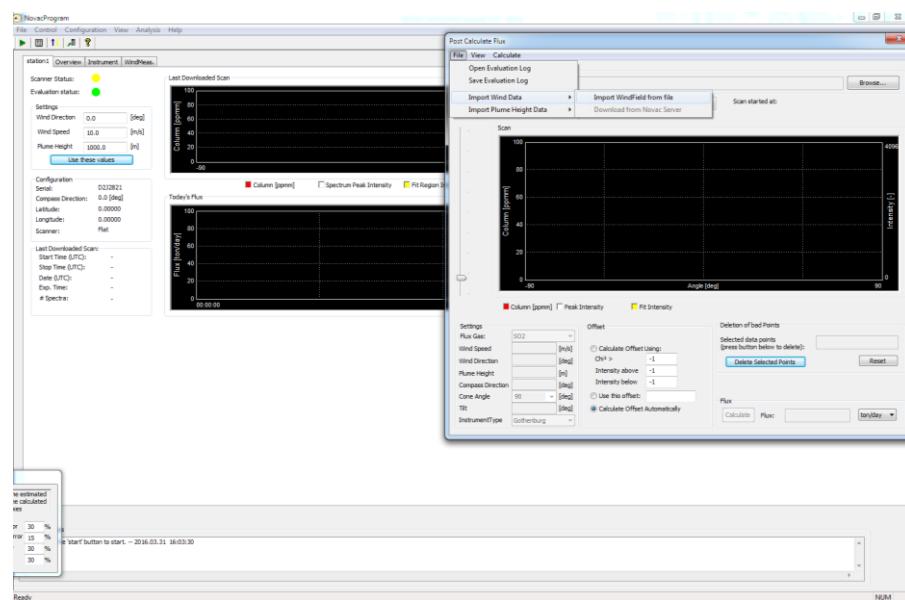


Figure 42. Selecting a wind file for flux-re-evaluation with the NovacProgram.

The name assigned to the file is not important, but it should be a *.txt file containing the following fields in this format:

date	time	ws	wd	ph
2015.10.20	10:00	3	360	1500
2015.10.20	11:00	2.3	350	1200
2015.10.20	12:00	2.8	360	1350
2015.10.20	13:00	3	360	1750
2015.10.20	14:00	3.3	10	2100
2015.10.20	15:00	2.7	350	1900
2015.10.20	16:00	2.8	310	1800

The first column is the date in yyyy.mm.dd format, the second is time in hh:mm format, the third is wind speed in n.n format, the fourth is wind direction in n.n format and the last is plume height in n.n format. As always, date and time should be UTC (to correspond with the measurement time), wind speed is positive, wind direction is measured clockwise from north and plume height is the altitude of plume above the station (which means that for each scanner a new plume height column should be defined). The last column is optional, if it is not defined the NovacProgram will adopt a default plume height value (typically the difference in altitude between the scanner and the volcano's summit).

The information contained in the wind file can be extracted from the GeometryLog (except wind speed), from a meteorological station, a model (ex., NOAA-Ready model: <https://ready.arl.noaa.gov/READYamet.php>), or elsewhere. When read out, the NovacProgram will perform an interpolation to select the closest values in time to each measurement.

6 Post-processing SO₂ data

The NOVAC-Program can also be used to make a second analysis of the generated results, for error checking or for generation of better data if more knowledge about e.g. wind speed has been retrieved.

6.1 Post flux calculation

The post flux calculation dialog can be reached from the menu in the main window by pressing ‘Analysis→Flux’. Through the post-flux calculation dialog, the user can calculate the flux from a generated evaluation (or re-evaluation) – log file.

Open the (re)evaluation –log file for which to calculate the flux by pressing the ‘Browse’-button in the upper part of the dialog or by the menu option ‘File->Open Evaluation Log’ (**not** by typing in a file name in the empty space next to the ‘browse’ button).

When the evaluation-log file has been opened, the column plot in the middle of the dialog will show the variation of gas column with scan angle (angle from zenith) as red bars. The 1-sigma errors of the spectral fit are shown as white error-bars. The maximum intensities of the spectra are shown as white squares. Data points which are regarded as good measurements are shown in a clear red colour. Data points which are regarded as not good spectra, e.g. if the spectrum is too dark or saturated, are shown in a dark red colour and are not used in the flux-calculations.

From the ‘View’ menu it is possible to show more information about the scans, such as maximum intensity in the fit region (‘Fit Intensity’), the column error (which is shown on the same scale as the column-bars), or the delta (peak-to-peak value of the residual) or the chi-square of the fit. However, note that the delta- and chi-square values are shown as normalized values, their absolute numbers are not shown.

There is some information about the currently shown scan shown above the column graph. In the upper left corner is shown the number of scans in the evaluation log and the index of the currently shown scan. In the upper right corner is shown the date when the scan started and the time (in UTC) when the scan started.

In the middle, just above the column graph, is a small spin-control with two arrow buttons. Pressing the right button will show the next scan in the evaluation log and pressing the left will show the previous scan.

It is possible to use a wind-field defined in a wind-field file (for the format see the section ‘Wind field file’). In this case the wind-field used to calculate the fluxes will be interpolated between the data points specified in the wind-field file. To use a wind-field file in the calculations, go to the menu option ‘File->Import WindField’ and select the file to use. On successful parsing of the wind-field field, a dialog will show the number of data points that has been read in from the file.

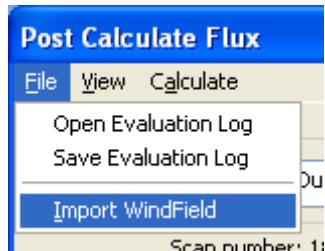


Figure 43. To use a wind-field file to interpolate the wind-field for each scan to calculate, select a wind-field file to use by pressing the menu-option ‘Import WindField’ and select a file to use.

In the ‘Settings’ group in the lower left part of the dialog is numbers which are necessary for calculating the flux for the scans.

- The flux-gas specifies which gas the flux should be calculated for (if several gases are evaluated for in the evaluation log) this is typically SO₂.
- The wind speed and wind direction are necessary parameters and are specified at plume height. Note that there can be a very large difference between the wind speed at ground level and the wind speed at plume height. These fields will be disabled if the wind speed and/or wind direction is taken from a wind-field file.
- The plume height specifies the height of the plume above the instrument. Note that this is not the same as the plume’s altitude above the sea level. This field will be disabled if the plume height is taken from a wind-field file.
- The compass direction is the direction in which the scanning instrument points, in degrees from north counted as positive in the clock-wise direction. If this is specified in the evaluation-log then this field is filled in automatically.

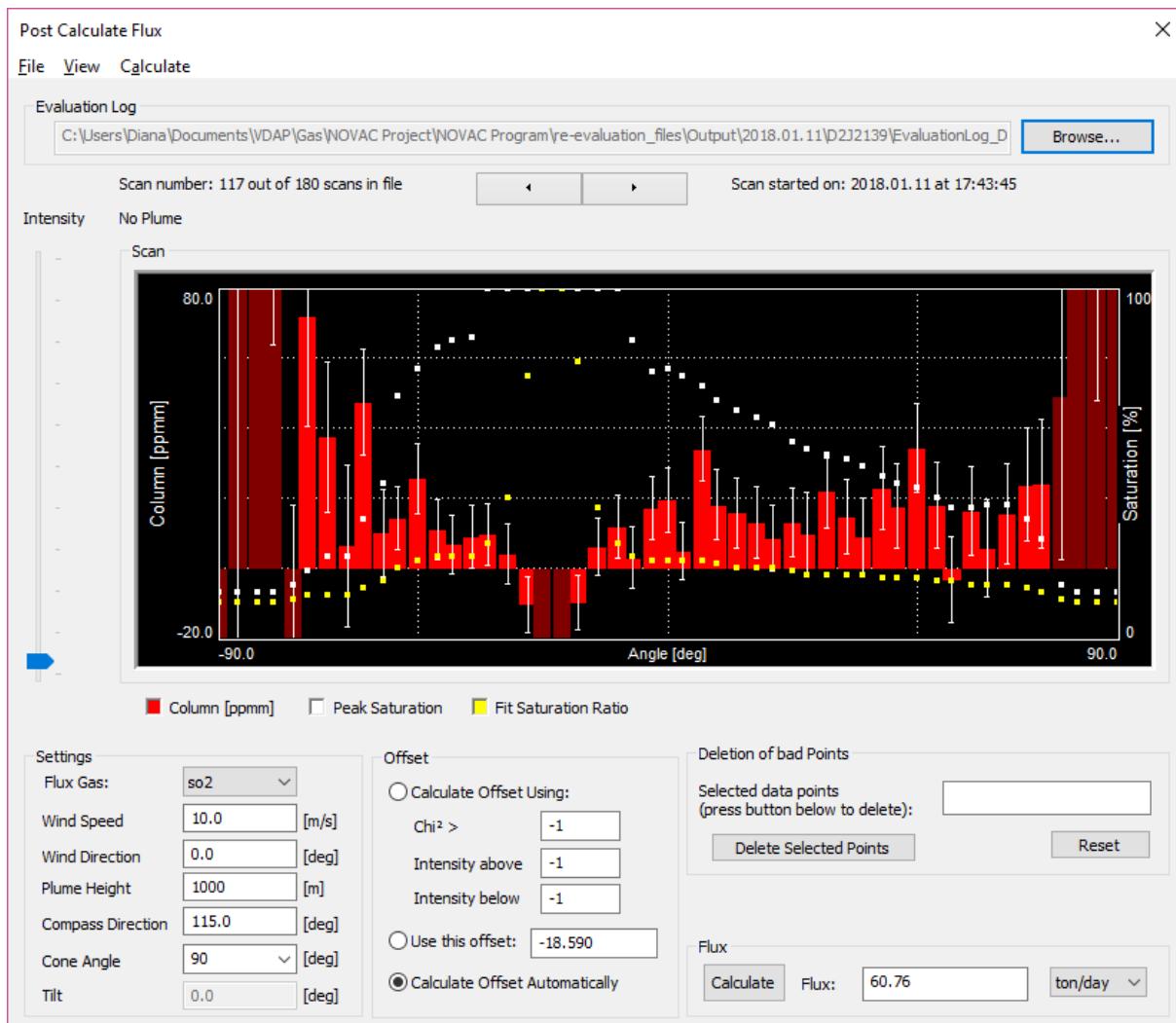


Figure 44. The post-flux calculation dialog. In this dialog, it is possible to perform off-line calculations of the flux.

In the ‘offset’-group are parameters used for deciding the offset-level of the scan. When the gas columns in the spectra are evaluated, they are measured relative to the gas column in the sky spectrum, which is a zenith spectrum collected before the real measurement. If this spectrum has a high gas column, then some of the evaluation will show negative numbers. This is not wrong but simply a result of making a relative measurement. These negative numbers can be compensated for by including an offset-level into the flux-calculation. This offset level tells the program where the ‘real’ zero-level is. The calculated offset-level is shown in the column graph as a dashed red line below the scan. The offset can be calculated automatically (default) or the user can specify an offset-level, or the user can specify the parameters for the calculation of the offset-level.

In the ‘Deletion of bad points’-group, the user has the possibility to remove single measurement points from the flux-calculations. Spectra which are too dark or saturated will have a very high noise-level and could deteriorate the flux-calculations. Therefore, bad measurement points should be removed from the data set before calculating the flux. The program automatically detects bad measurements and marks them in a dark red colour. If the user finds more data points which should be marked as bad, the user can specify them in the

empty space in the ‘Deletion of bad points’ - group. These points will then be selected and appear as blue columns in the column graph. When pressing the ‘Delete Selected Points’- button these selected points will be removed from the data set, and they will appear as dark red in the column graph. Data points can also be selected by intensity by moving the intensity slider on the left side of the column plot up or down. Data points with an intensity below the selected intensity of the column slider will be selected and appear in the list in the ‘Deletion of bad points’ – group and appear as blue columns in the column plot. Pressing the ‘Delete Selected Points’ – button will remove these points from the dataset.

Pressing the button ‘Calculate’ in the ‘Flux’-group in the lower right corner of the dialog will make the program calculate the flux for the currently shown scan using the data points shown in clear red or in blue in the column graph. The result is shown in the empty box to the right of the button. The unit of the flux is selected **before** calculating the flux by selecting the desired unit in the combo-box in the right-most corner.

When the program calculates the flux, it will also generate a post-flux log which is put in the same directory as the opened evaluation log. This log-file contains the resulting flux in kg/s, the wind speed, wind direction and plume height used for calculating the flux. It will also save the offset used and the time (typically GMT) when the spectrum was started.

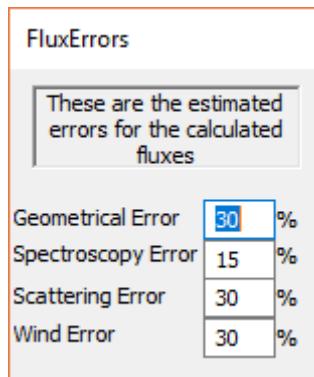


Figure 45.The estimation of errors for post flux calculation.

6.2 Re-Evaluation of spectra

If the user wants to make a second evaluation of spectra collected with the NOVAC-program, this can be done through the ‘Re-Evaluation’-dialog which can be found under the ‘Analysis’ menu in the main window. The re-evaluation dialog lets the user choose the evaluation parameters and can evaluate the spectra with different reference-files, with different settings and make advanced experiments with the spectra.

The pak-files which are used in the re-evaluation dialog **must** be complete, single scans or they will not be evaluated. If the pak-file contains several scans they can be split up using the ‘Split/Merge’ – dialog found under the ‘File’ menu in the main window. Pak-files which are generated by a running NOVAC-program will always be single scans, however they are not always complete.

The re-evaluation dialog contains four pages; in each page one aspect of the evaluation is treated.

Remember that if you want to close the dialog press the ‘Close’ button in the lower right part of the dialog.

NB – Pressing the ‘Close’ – button will close the dialog promptly without doing anything more, no settings will be saved.

6.2.1 Re-Evaluation – Selecting Scans

The first page in the Re-Evaluation dialog is the ‘Scans’ – page. Here you select the pak files you want to evaluate again by pressing the ‘Browse’ – button in the lower part of the dialog. The selected pak-files will be shown in the list in the left part of the window.

In the right part of the window, you can see inspect the spectra in the pak-file and browse through them by pressing the left and right arrows below the spectrum-graph. The lower right part of the dialog shows general information about the currently selected spectrum, such as collection date, exposure time, number of averaged spectra or scan angle.

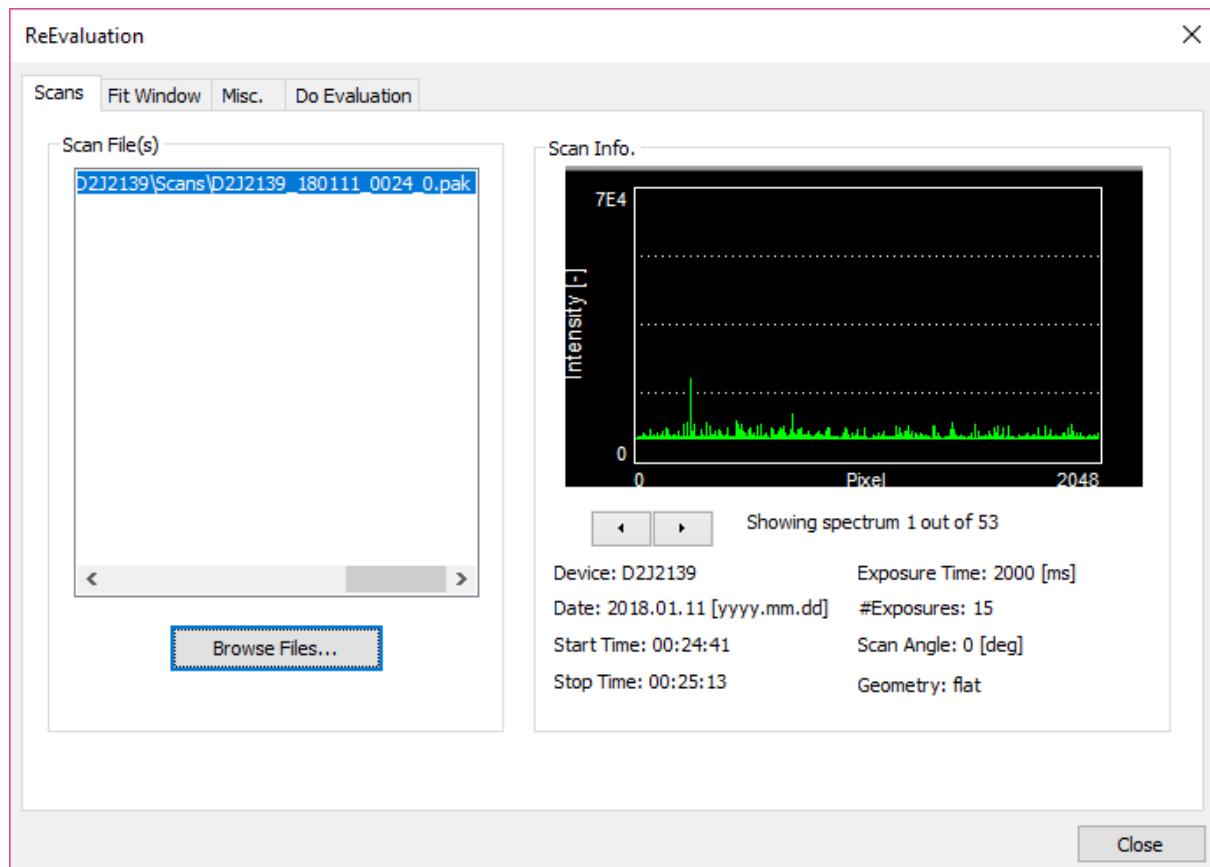


Figure 46. The Scans-page in the Re-Evaluation dialog. Here you can select and inspect each of the the pak-files to evaluate.

6.2.2 Re-Evaluation – Setting up the fit windows

The second page in the Re-Evaluation dialog is the ‘Fit Windows’ – page. The Re-evaluation algorithm is based on the concept of ‘fit windows’. A fit window is a spectral region, where the spectrum is evaluated for one or more species. You can specify more than one fit region in this window, each spectrum will then be evaluated two times, one in each region. If you want, these regions can be overlapping or even be the same. Each fit-window has a name, and all

fit-windows are shown in the fit-window list in the left side of the dialog. The fit-window you are currently working on is the one which is marked by a blue colour in the list.

For each fit window you specify the region of the spectrum that the fit-window should cover. The region is specified by the fit from and fit to parameters, which are same as in the Configuration dialog. **NB** In the same way as in the configuration dialog; the fit from and fit to are given as pixels in the spectrometer and are not affected if the spectra are read out in an interlaced way or as partial spectra.

In the ‘Fit Type’ – group you can specify advanced options on how the evaluation should be performed. The default option is ‘HP, divide by sky’ (which is used in the real-time evaluation). In the first option each measured spectrum will be divided by the sky spectrum before being logarithmated. In the second option, a high-pass filtered and logarithmated spectrum will be included into the fitting procedure, this can be useful if there are large shifts between the sky spectrum and the measured spectra.

In the ‘Misc.’ – group there are two options. The check-box ‘UV’ is checked by default and should be checked for spectra where the starting wavelength is shorter than 300 nm. If the check-box ‘Find Optimal Shift’ is checked, each scan will be evaluated two times. The first time the shift and squeeze of each spectrum will be set to 0 and 1 respectively (no matter what you specify for the references). The spectrum with the highest column value will then be evaluated one time more with the shift set to free, the program can in this way find an ‘optimum’ value for the shift. The whole scan will then be evaluated once more, now with the shift set to the found ‘optimal’ shift. This only works if the scan contains a clear plume with high column values.

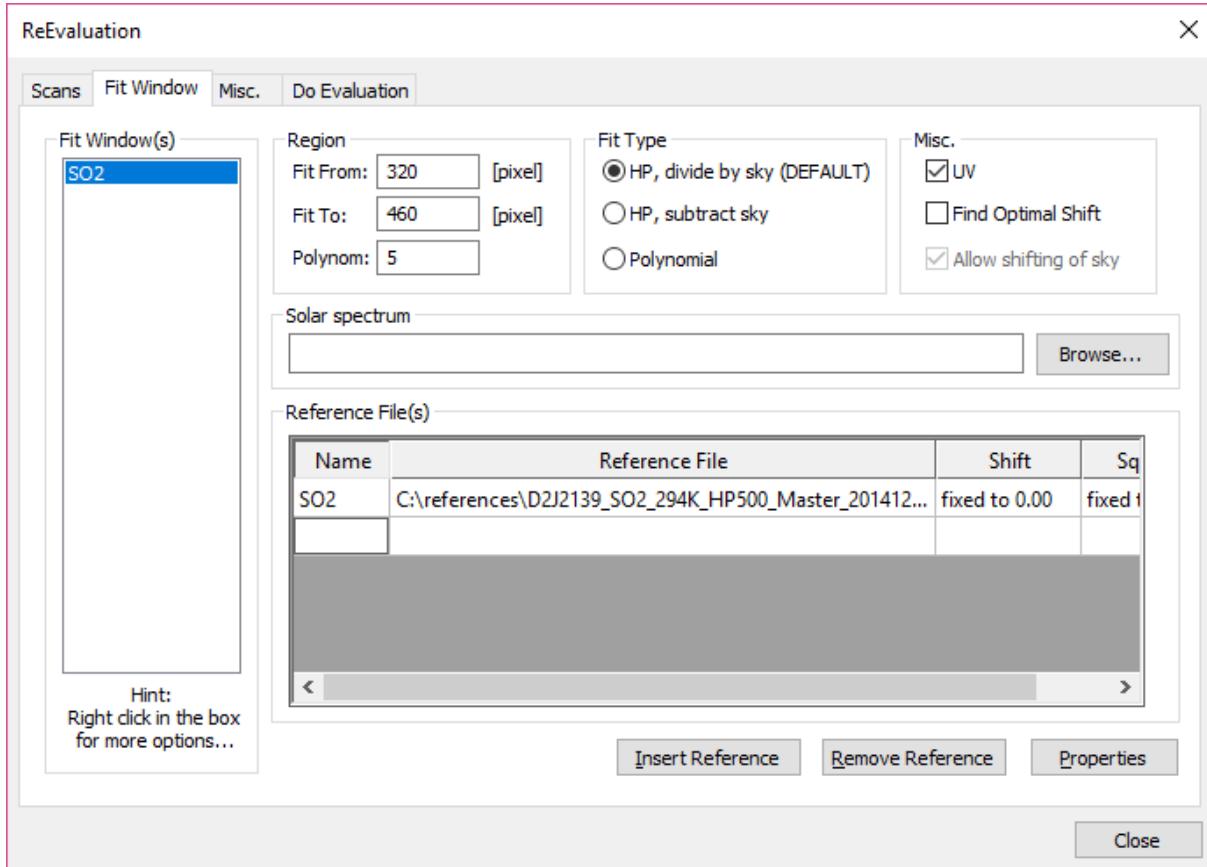


Figure 47. The fit windows page in the Re-Evaluation dialog. This page allows you to setup one or more regions in which the spectra should be evaluated.

In each fit-window you can specify up to 5 references that will be fitted. To include a reference-file press the ‘Insert Reference’ – button, to remove a reference-file press the ‘Remove Reference’ – button. To inspect and/or change the properties and options for a reference-file press the ‘Properties’ – button, this will bring up the ‘Reference Properties’ – dialog in which you can modify the reference-file.

All the fit windows that are currently configured are shown in the list of fit-windows to the left. Right clicking in this list will bring up a context-menu from which you can choose to insert one more window.

By clicking ‘Insert window’ in the context menu a new fit-window will appear in the list and you will be asked for a name for this new fit-window.

By clicking ‘Load Window(s) from File’ you can include fit-windows which have already been configured and saved to a file. **NB** by loading a fit window from a file, the list of fit-windows will be replaced by the windows in the file and the currently existing fit windows in the list will be removed.

To save the current list of fit windows select ‘Save Window(s) to File’ in the context menu. You will be asked for a file name, and the current fit-window setup will be saved in a ‘.nfw’ file. This is useful if you want to make the same evaluation later, using other spectra. **NB** it is

not possible to transfer .nfw-files between computers, since they only store the path to the reference-files, not the files themselves.

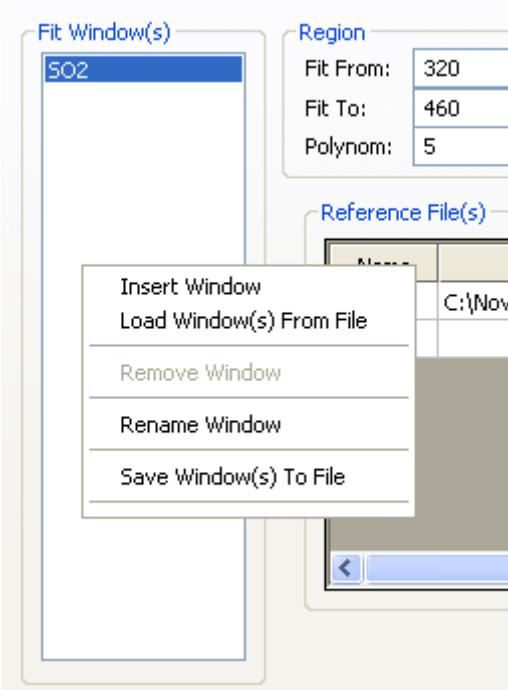


Figure 48. The context-menu in the list of fit windows. The menu is accessed by right-clicking in the list to the left.

6.2.3 Re-Evaluation – Misc. Settings

The third page in the Re-evaluation dialog is the misc. – settings page, which contains miscellaneous settings.

In the ‘Dark Spectra’ – group you can specify what to do with spectra which are dark or could be considered as dark.

1. If the option ‘Ignore Completely Dark spectra’ is chosen, then completely dark spectra will be skipped in the evaluation (they will not be evaluated at all). This is the default option, since they will not give any useful results anyway.
2. If the option ‘Ignore spectra with intensity below:’ is chosen, then you can specify a limit intensity and a pixel number around which to calculate the intensity. Spectra which have a lower intensity at this pixel-number will be ignored in the evaluation.
3. If the option ‘Don’t ignore anything’ is selected then all spectra will be evaluated, no matter how little light they have.

In the ‘Saturated Spectra’ – group you can specify what to do with spectra which have too high intensity.

1. If the option ‘Ignore Saturated spectra’ is chosen, then spectra which have a maximum saturation ratio of 100% will be ignored.
2. If the option ‘Ignore spectra with intensity above’ is chosen, then you can specify a limit intensity and a pixel number around which to calculate the intensity. Spectra which have an intensity higher than the limit intensity around the specified pixel-number will be ignored.
3. If the option ‘Don’t ignore anything’ is selected then all spectra will be evaluated, no matter how saturated they are.

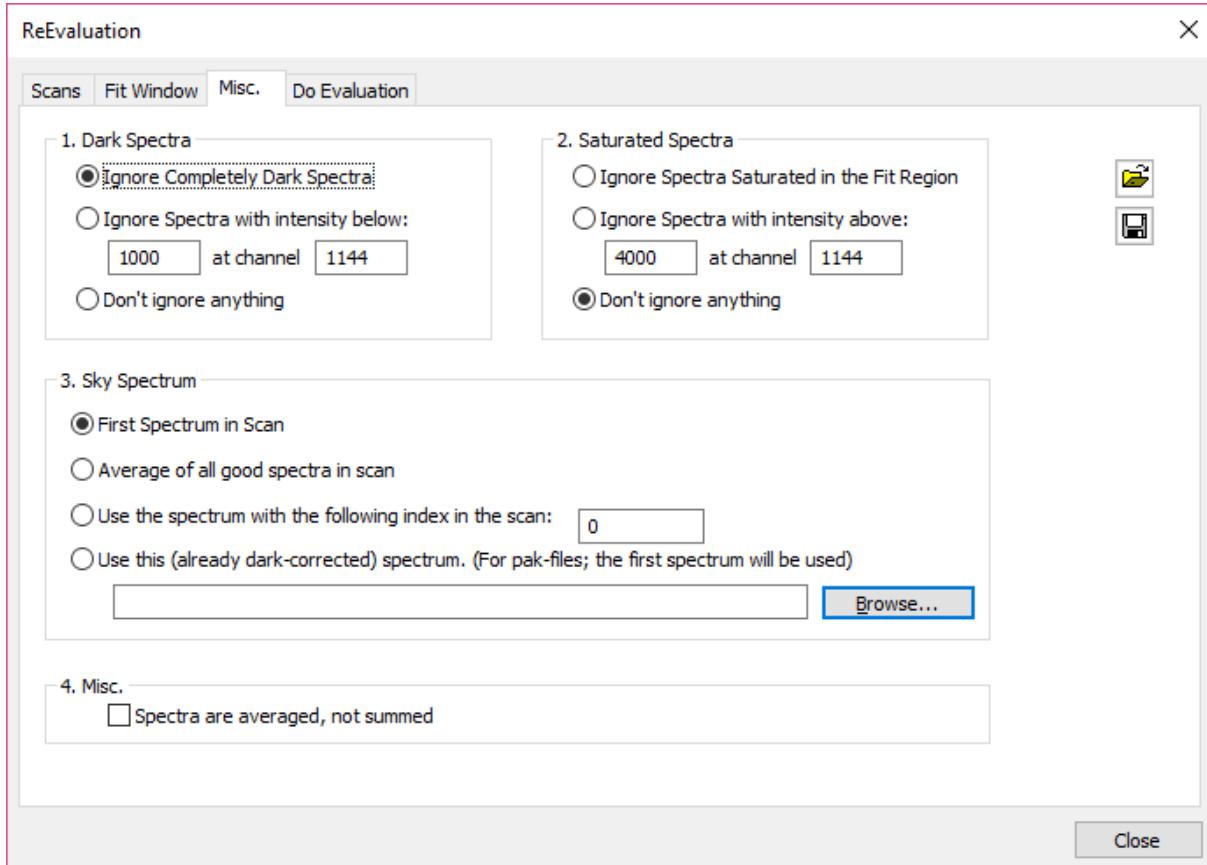


Figure 49. The Misc. page in the Re-Evaluation dialog contains misc. options for the evaluation.

In the ‘Sky Spectrum’ – group you can specify the options for how to select the sky-spectrum. The default option is to use the first spectrum in the scan, but it is also possible to use an average of all good spectra in the scan, which can possibly increase the signal to noise ratio in the evaluation. Another option is to use a specific spectrum number in the scan, this can be useful if e.g. the zenith spectrum is saturated. The fourth option is to specify a spectrum which can be used as sky spectrum. If a pak-file is specified, then only the first spectrum in the pak-file will be used. A user specified sky-spectrum **must** be corrected for dark-current and offset.

In the ‘Misc.’ – group are even more obscure options. If some old ScanDOAS spectra have been imported to a pak-file they may be summed or averaged depending on the version of ScanDOAS used. This button enables the use of spectra which are averaged.

To simplify the use of the re-evaluation dialog, these settings can be stored in a settings file for more easy use on the next occasion. To store the settings, press the button with a picture of a small floppy disk in the upper right corner of the page. To read in these settings from previously stored file, press the ‘open’ button in the upper right corner of the page.

6.2.4 Re-Evaluation – The final evaluation

The final page in the re-evaluation dialog is the ‘Do Evaluation’ – page.

Pressing the ‘Do Evaluation.’ button will start the evaluation of the selected spectra using the options setup in the earlier pages. The result of the evaluation will be shown in the two graphs on this page.

The left graph shows the measured spectrum in the fit-region or the residual of the fit depending on the settings of the radio button below the graph.

The right graph shows the result of the evaluation. It shows, in blue, the reference-file which is selected in the list below the graph (in Figure 50 this is SO₂) scaled to the evaluated column value. Below this is, in red, the same scaled reference-file plus the residual of the fit. This enables inspection of the quality of the fit. If the residual is small and lacks structures, then the evaluation is ok.

Below the right graph is the evaluated column value of the currently selected reference-file shown for the last evaluated spectrum. The shift and squeeze used for the fit are also shown together with the delta and chi-square values which are used as a quality measure of the fitting, smaller values are better.



Figure 50. The final page in the re-evaluation dialog. Here you can perform the evaluation of the selected pak-files using the specified options and inspect the result.

By default, the evaluation runs without making any pause between the spectra. However, if checking the ‘pause’ check button in the left half of the screen, the evaluation will make a pause after evaluating each spectrum. The program will evaluate next spectrum when the user presses the ‘Next’ – button, which is situated where the ‘Do Evaluation’ – button was before.

The evaluation can always be interrupted by pressing the ‘Cancel’ button to the right of the ‘Do Evaluation’/‘Next’ – button. **NB** DO NOT PRESS THE ‘CLOSE’ BUTTON ON THE BOTTOM ROW WHEN THE EVALUATION IS RUNNING.

6.3 Setup inspection

The setup-inspection dialog can be opened by selecting ‘Analysis→Setup’ in the menu in the main window. The setup inspection dialog can handle up to four evaluation logs and show the result of the measurements on a map.

To open an evaluation log, press one of the ‘Browse’ – button. The evaluation log file must contain GPS information, which some older evaluation-log files does not do.

When the evaluation log has been opened, the program will calculate which NOVAC volcano is closest, and calculate the distance from the instrument to the summit of the volcano. If the compass-direction is specified in the evaluation log, the program can make a small map of the scans in the evaluation log file. **NB** the scale of the scan is entirely depending on the plume height which must be supplied by the user.

The single red square on the map is the volcano, the black square is the scanning instrument, and the line of squares of varying colour is the result of the scan. The scan is represented on a relative scale, with blue being the lowest column values and red showing the highest.

For a single scan, the program can calculate the wind direction by knowing the position of the volcano and the position of the highest column values. **NB** This calculation requires knowledge of the plume height, the accuracy of the wind direction is entirely depending on the estimation of the plume height.

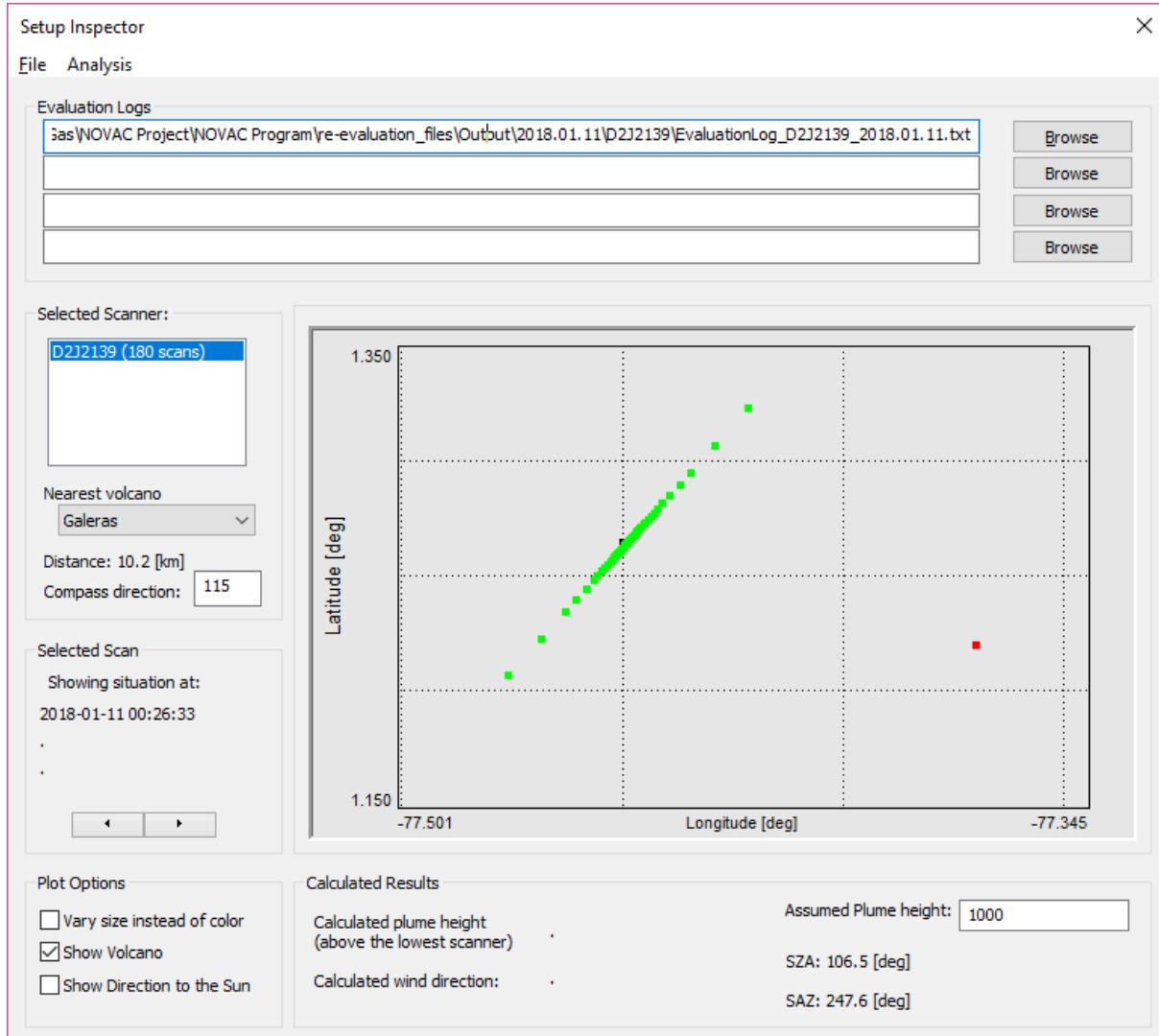


Figure 51. The setup-inspection dialog. Here it is possible to see the position and direction of the scanning instrument on a map and look at the results from a scan.

6.4 Post-processing wind speed and direction

If a connected instrument has performed a measurement of the wind-speed using the dual-beam technique, the resulting wind-speed will be calculated in real-time when the measurement is downloaded and evaluated. This can also be done in off-line mode from an already downloaded scan through the ‘Post Wind speed analysis’ – dialog.

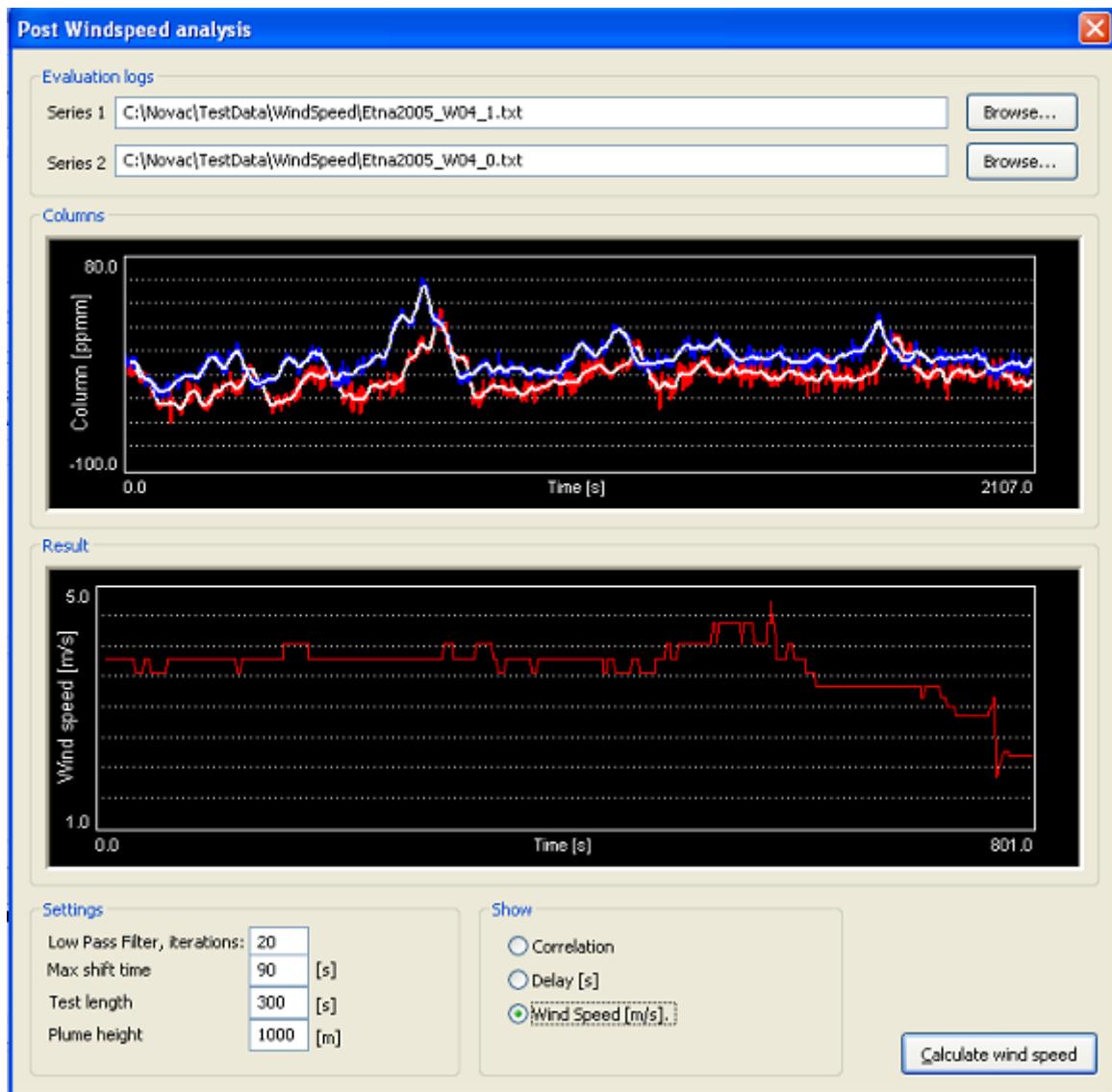


Figure 52. The Offline wind speed analysis dialog. In this dialog it is possible to perform calculations of wind speed from measured dual-beam time series.

6.5 Data Browser Dialog

The Data Browser dialog, accessible from the menu in the main window through; “Analysis” → “View Measured Data” and lets the user browse through the measured data of one day while inspecting the properties of each scan.

To view evaluated data, open an Evaluation log-file by choosing “Data Source” → “Evaluation Log” from the menu in the Data Browser dialog. The evaluation log is then opened, and the first scan is shown in the top graph. The left-hand side of the screen contains several of the parameters of the measurement which allows for easy inspection of the changes of these parameters during the day. The lower graph shows the evolution of the calculated flux of the measurements found in the opened evaluation-log.

By pressing the left and right arrow buttons in the top part of the dialog, the dialog moves to previous or next scan in the evaluation log file.

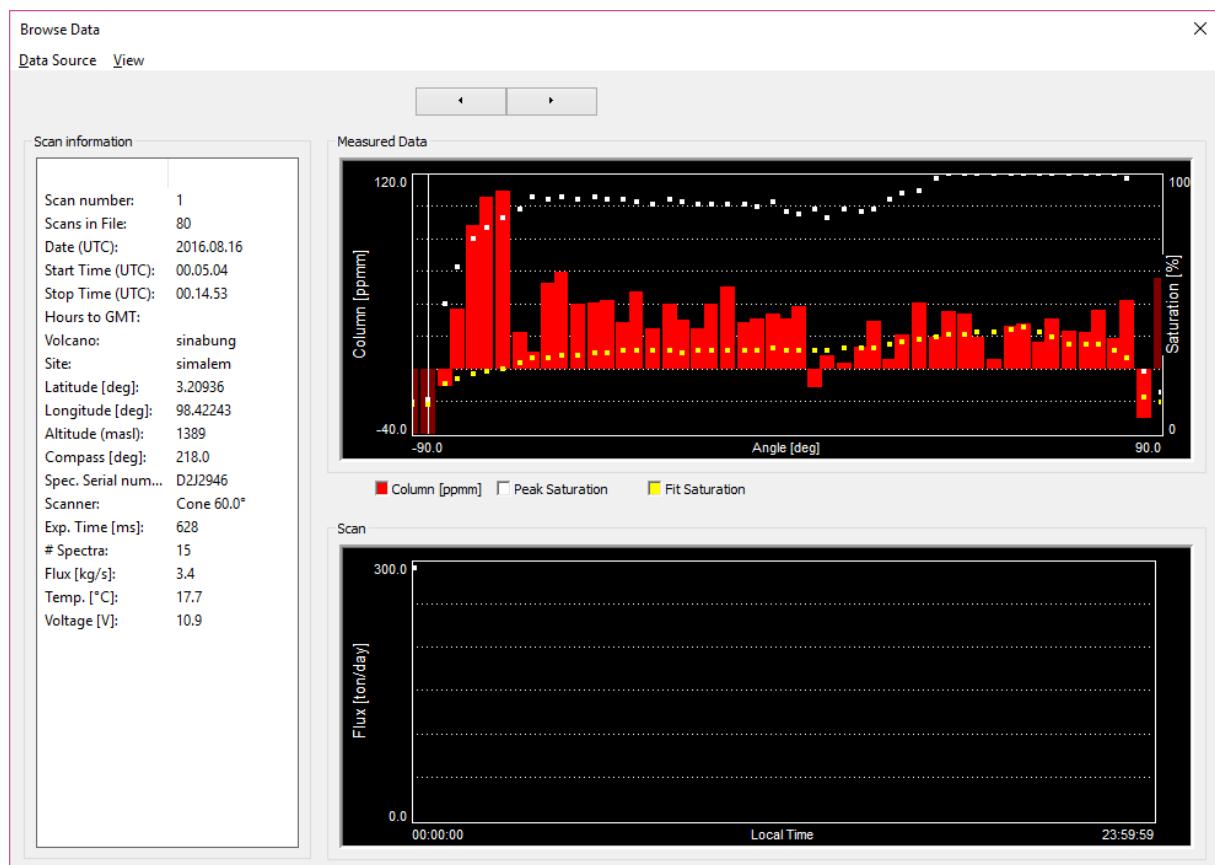


Figure 53. In the data browser dialog it is possible to step through the scans found in one EvaluationLog file and to see the information about the scans and a graphic representation of the calculated fluxes.

7 Advanced file handling options

The program contains a small set of utilities designed to make it easier to analyse the results of the programs real-time analysis. Among these is the possibility to import spectra from other file-formats and to export spectra to other file formats.

Since the spectrum file format that the NOVAC-project use (pak-file format) is capable to contain several spectra in one file, there is also a small function to merge .pak-files together or to split the files in one pak-file into several.

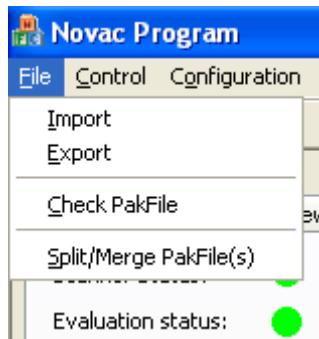


Figure 54. The file-handling functions are accessed from the File menu in the main window.

7.1 Importing files

From the import-dialog, it is possible to convert spectral files from older file-formats to the NOVAC-projects spectral file format, pak. Currently only STD-files generated by the program ScanDOAS made by Chalmers University of Technology can be read-in.

By pressing the ‘Browse’ button, the program brings up a dialog which lets the user choose a number of spectral files to import. In the case for files generated by ScanDOAS, it is important that only full scans are selected.

Under the ‘Generation’ group the user can specify, for spectral-files generated by ScanDOAS how many spectra make up one scan. This is necessary since this information is not specified in the spectrum files.

Under the ‘Additional information’ group, the user can supply further information which is not given in the spectrum file but should be included in the pak-file. Most important is the scan angles for the different spectra in the scan, since this is necessary if one wishes to use the scan for flux-calculations.

Under the ‘Output Directory’ group, the user can type in one directory where the resulting pak-files will be put. The program will generate one pak-file for each scan. It is important that the user has write access to the supplied output directory.

The importing is performed when the user presses the ‘Import’ button in the lower part of the screen. If the ‘Ok’ or ‘Cancel’ buttons are pressed, the dialog will be closed, and nothing will be done.

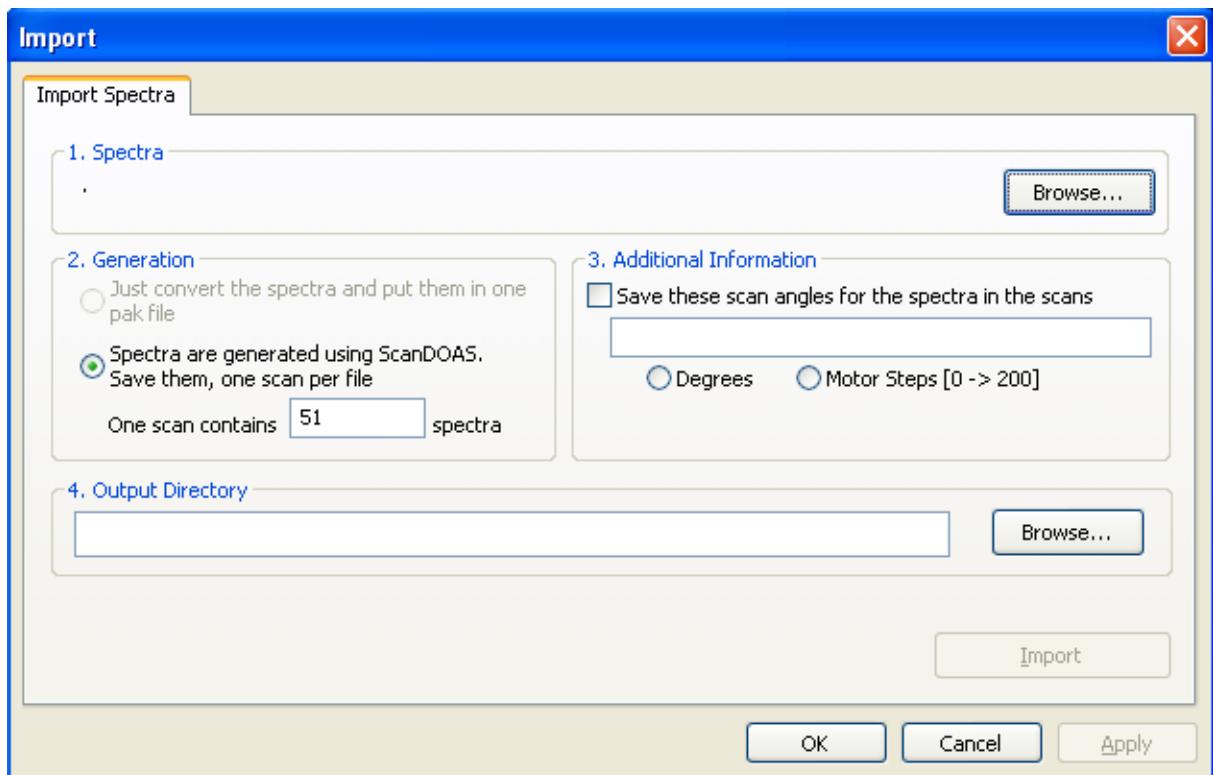


Figure 55. The import dialog which can be found under the File menu in the main window.

7.2 Exporting files

7.2.1 Exporting spectra

From the Export-spectra dialog it is possible to select one or more pak-files and to export them to another spectral file format, which can be read by some other program. Currently only STD-files can be written.

The pak-files to be exported are selected by pressing the ‘Browse’ – button in the ‘.pak – file(s)’ group.

The directory where to put the exported spectral files can be given by pressing the ‘Browse’ button in the ‘Export to:’ group or by directly typing in a directory name in the empty field in the ‘Export to’ group. In the supplied directory, one sub-directory will be created for each pak-file that is selected.

The exporting is done when the user presses the ‘Export’ button. To close without performing any export press the ‘Ok’ or ‘Cancel’ buttons.

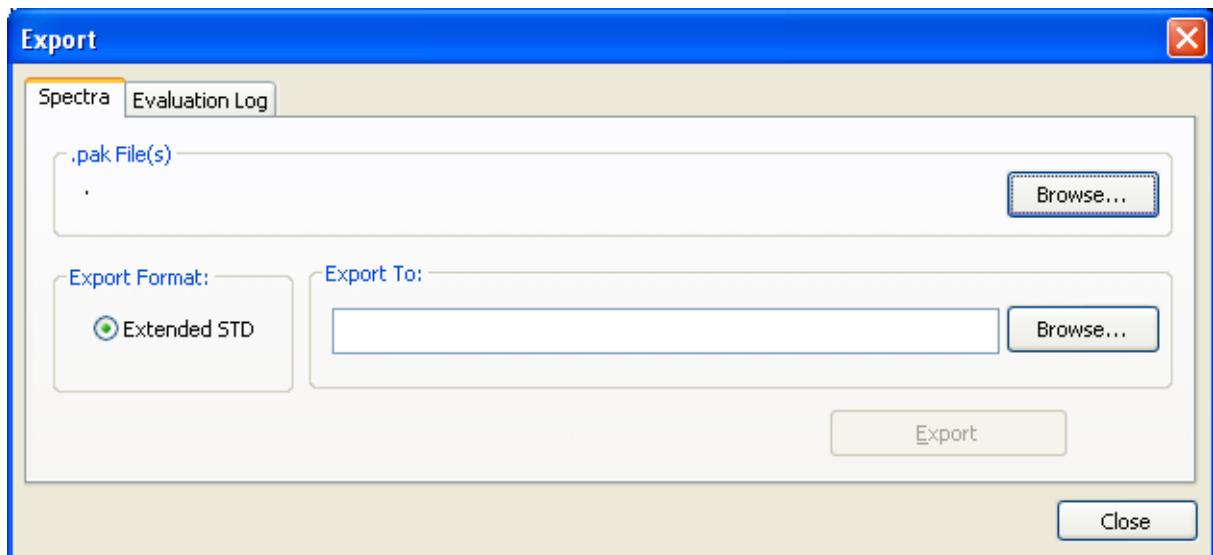


Figure 56. The export spectra dialog, which can be found under the File menu in the main window.

7.2.2 Exporting Evaluation logs

For back-wards compatibility, it is also possible to convert the evaluation-logs that are generated when the program is running, or the Re-Evaluation logs generated by the spectral re-evaluation routine, to a file format which is readable by ScanDOAS.

Supply an evaluation log to export by pressing the ‘Browse’ button in the ‘Evaluation Log’ group.

Supply a filename for the generated ScanDOAS-log by pressing the ‘Browse’ button in the ‘Save As’ group.

The export will be performed when the user presses the ‘Export’ button. To close without exporting anything, press the ‘Ok’ or ‘Cancel’- buttons.

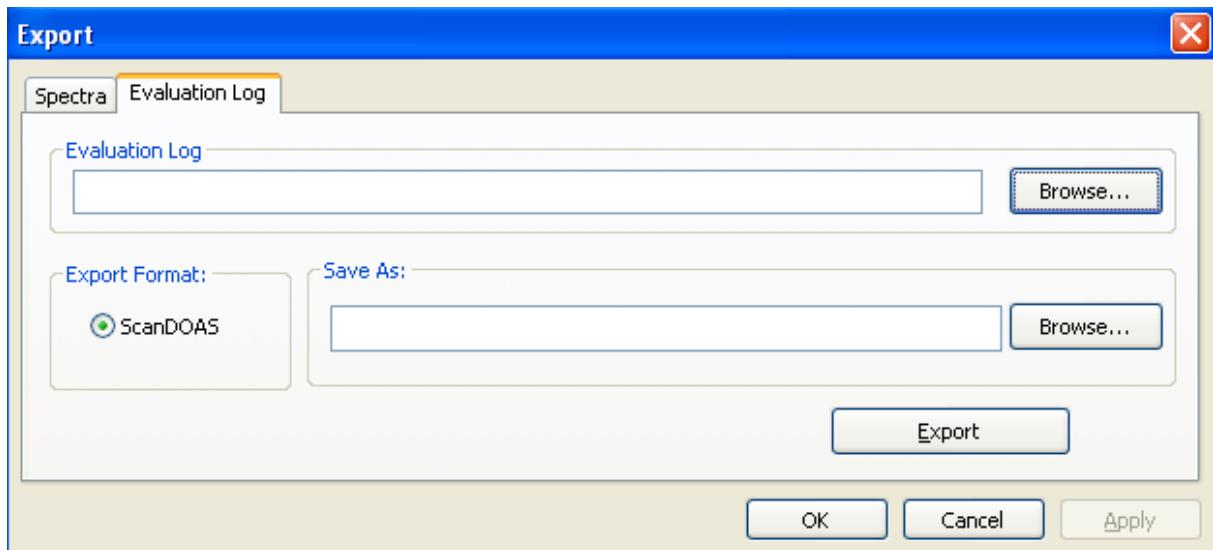


Figure 57. The export Evaluation-log dialog, which can be found under the File menu in the main window.

7.3 Checking pak-files

In the Pak-file inspector dialog (found under ‘File’→‘Check pak file’ in the main menu) the contents of one compressed spectrum file can be inspected. No modifications to the spectrum file can be made.

To get a closer look at the spectrum, the spectrum view can be zoomed by moving the mouse cursor with the left mouse button pressed. To restore the view to the original scale, click once with the right mouse button.

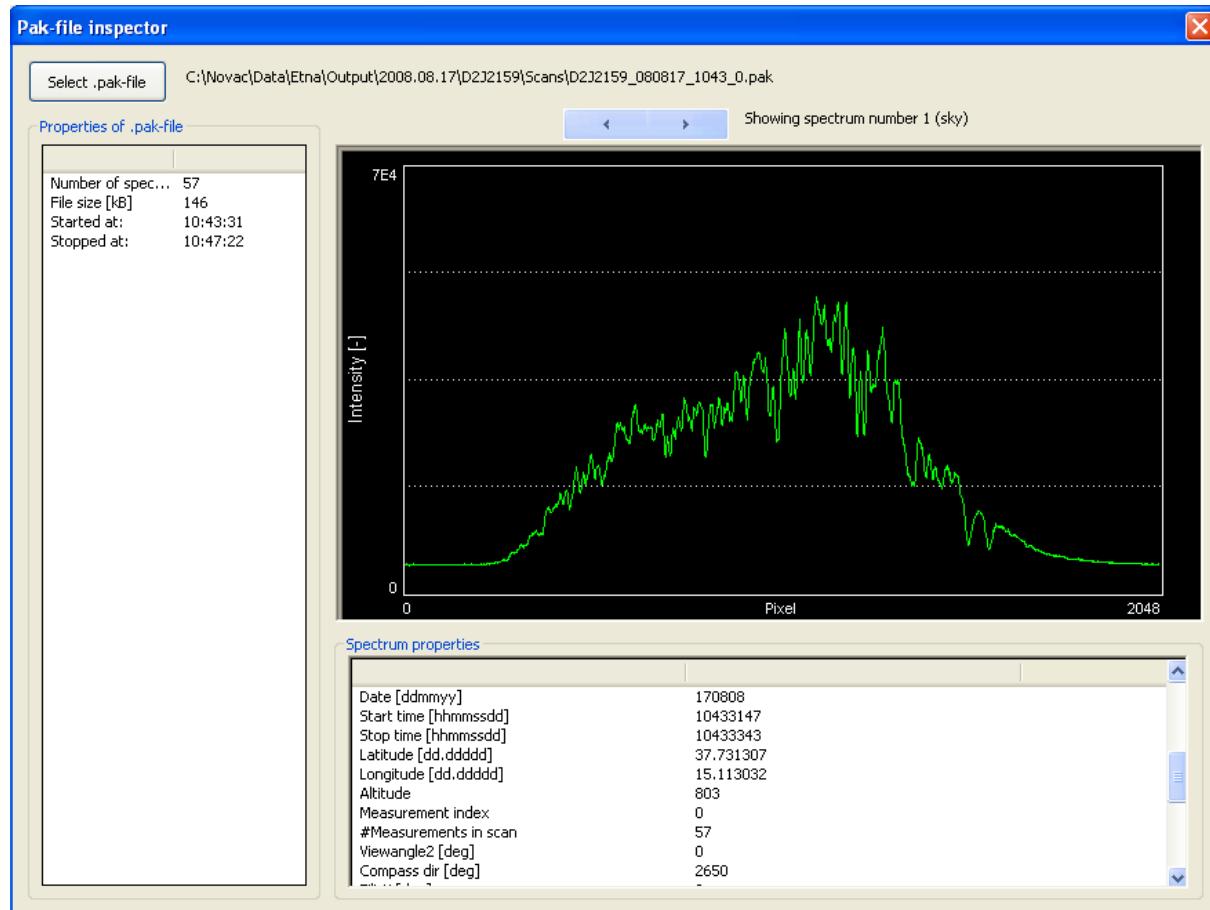


Figure 58. The pak-file inspector dialog shows the contents of one compressed spectrum file (.pak)

7.4 Merge Evaluation Log Files

You can merge evaluation log files from the ‘File’->‘Merge Evaluation Log Files’ menu. Browse for the evaluation log files to merge. The evaluation log files should be in the same directory and you must select multiple evaluation logs by holding down the Ctrl or Shift keys when selecting files. Select the output file name. Then click on merge.

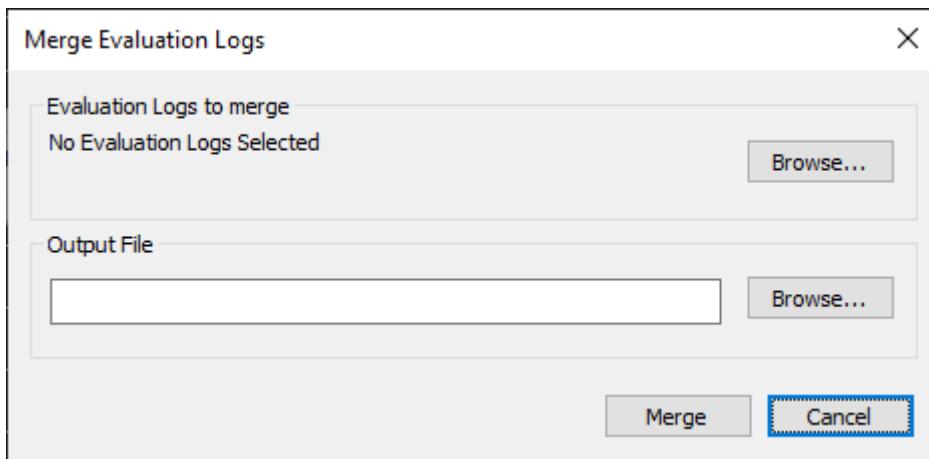


Figure 59 Merge Evaluation Log Files dialog

7.5 Merging/Splitting pak-files

7.5.1 Splitting pak-files into scans

If a pak-file has been downloaded from the instrument manually (not by the running program) or if a pak-file has been uploaded to an FTP-server by the instrument (not by the running program) then the pak-file might not have been divided into separate scans. It is then necessary to separate the pak-file before it can be used in the Re-Evaluation routine.

To do this open the Split/Merge dialog which can be found under ‘File->Split/Merge pak-file(s)’ in the main window and choose the ‘Split pakfile into scans’-page.

The pak-files to be split can be selected by pressing the ‘Browse’ button in the ‘Pak File(s) to split into scans’ group.

The directory, where to put the generated scan-files can be supplied by pressing the ‘Browse’ button in the ‘Output Directory’ – group or by typing in a directory-name in the empty space in the ‘Output Directory’ – group. A subdirectory will be generated in the output directory named ‘Scans’ which will contain the generated pak-files.

The Splitting is started by pressing the ‘Split’ button. To close the dialog **without** splitting any files, press the ‘Ok’ or ‘Cancel’ buttons in the bottom part of the dialog.

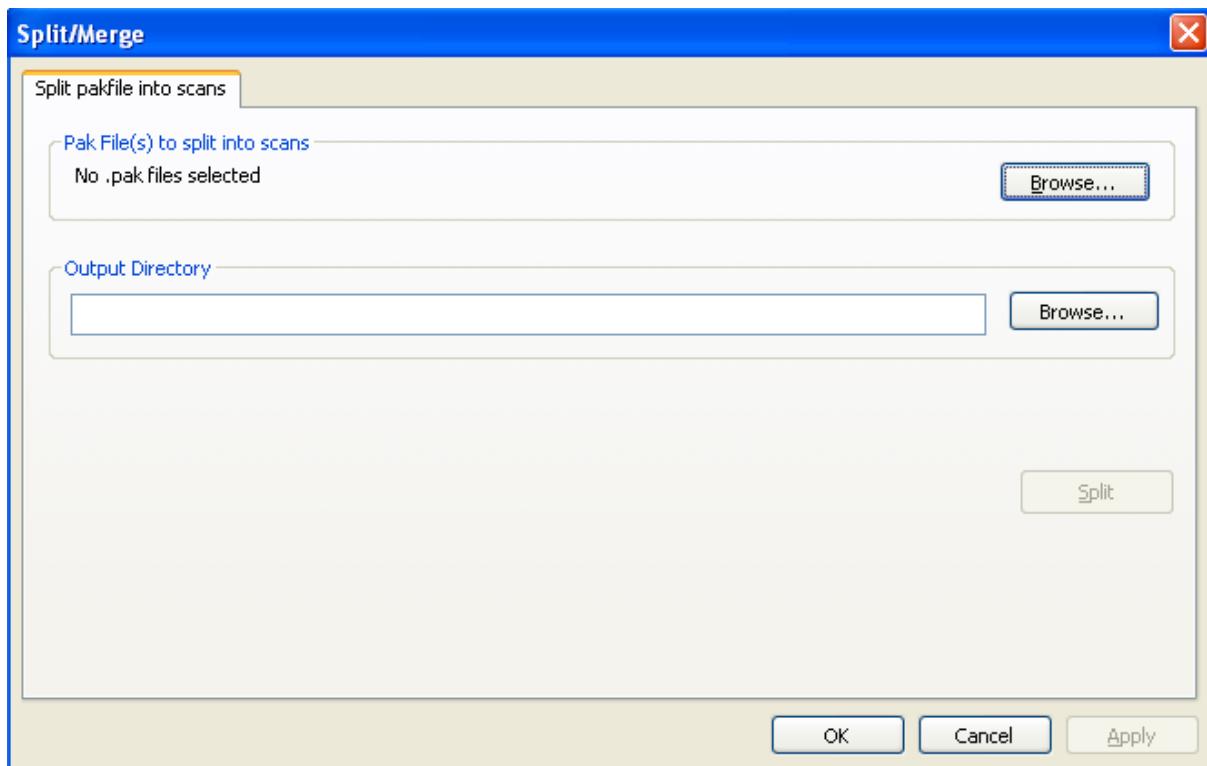


Figure 60. The 'split pak-files into scans' dialog, which is access from the File-menu in the main window. Through this dialog the user can select one or more pak-files and split them into separate scans.

7.5.2 Merging pak-files

Several .pak-files can be combined into one larger pak-file by using the Merge Pak Files page in the Split/Merge dialog. First select the pak-files that should be merged by pressing the upper most 'Browse' button, then select the directory where the merged file by either typing in the directory in the empty box or by pressing the second 'Browse' button. When the files have been selected and an appropriate output directory has been selected the button 'Merge' will appear and pressing it will start the merging process.

There are some options for how to do the merging. The options found in the group 'Sorting' determine how the pak-files should be sorted. The current options are to sort the files in ascending or descending order with respect to the original file name. The option found under 'Options' specified how the pak-files should be merged. The first option (which is the default) is to simply put the spectra one after the other in the same file. The second option is to spectrally merge the files, this only works if pak-files containing scans with the same number of spectra are selected. In this case the spectrum number 1 in the first pak-file will be spectrally added with spectrum number 1 in the second pak-file etc. The number of spectra that will be added together into one is typed into the empty white box just after the second radio-button.

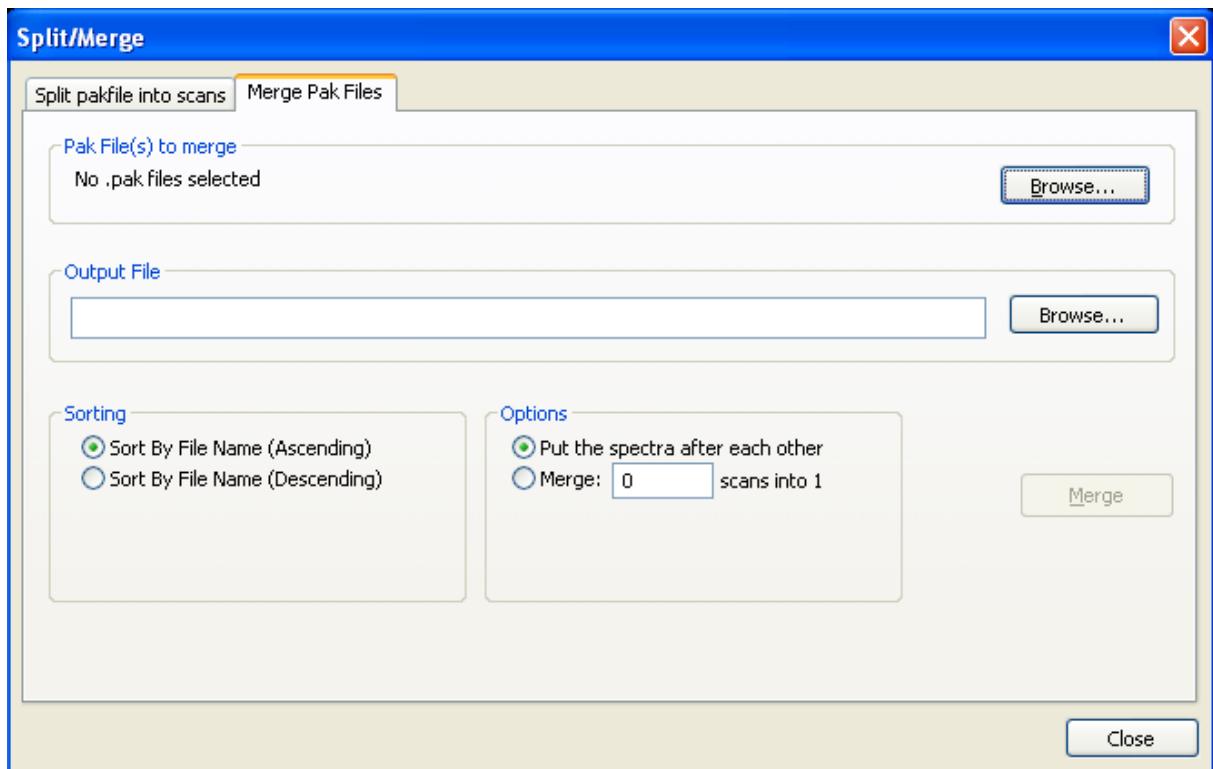


Figure 61. The ‘merge pak-files’ dialog, which is access from the File-menu in the main window. Through this dialog the user can select several pak-files to merge together into one pak-file.

8 Instrument Calibration

As of version 3.4 of the NovacProgram there is an instrument calibration dialog, found by selecting *Analysis*→*Calibrate Spectrometer* in the menu of the main window. This dialog makes it possible to create new references from measured scans in an easy way. This dialog is easy to navigate through but does still require some background insight into the underlying spectroscopy for understanding the final result.

There are two things which needs to be determined in order to create a reference for a particular spectrometer; the wavelength calibration of the device (aka the pixel-to-wavelength mapping) and the instrument line shape. The wavelength calibration says which wavelength (measured in nanometers, nm) each pixel on the spectrometer’s detector corresponds to and makes it possible to determine where in the spectrum the absorption features of e.g., SO₂ should be located. The instrument line shape says how much the spectrometer smoothens the measured spectra and makes it possible to determine the expected shape of the absorption features of e.g., SO₂ which we are measuring.

8.1 Wavelength calibration and instrument line shape from measured scan data

In the *Wavelength Calibration* tab, it is possible to create an instrument calibration from a measured spectrum. The input is a measured scan produced by NovacProgram and collected during good measurement conditions. The calibration procedure will use the measured sky spectrum of the selected scan and for the calibration to not be disturbed by plume gases, the scan should be selected such there should not be a plume directly overhead of the instrument.

8.1.1 Required input

The *Wavelength Calibration* tab makes use of the *StandardCrossSections* setup which should be located in a sub folder named *StandardCrossSections* in the same folder as the NovacProgram executable file. If these files are missing, please contact the supplier of your scanning DOAS system.

The algorithm behind the instrument calibration requires a high-resolution Fraunhofer spectrum as input. This spectrum is already part of the *StandardCrossSections* setup but the user is free to select another high-resolution Fraunhofer spectrum as well. If another spectrum is selected, make sure that the wavelength calibration of this spectrum is in nanometers air.

The algorithm also requires an initial wavelength calibration of the device. This can either be provided by the manufacturer of the device or extracted from an already existing reference.

The algorithm can use an initial instrument line shape as input to the instrument calibration. If no instrument line shape is provided then the instrument calibration routine will create an approximate Gaussian line shape from the structures present in the measured spectrum and use this as an initial instrument line shape. Furthermore, if no instrument line shape is fitted then this Gaussian line shape will be the resulting instrument line shape of the routine.

8.1.2 Setting up the instrument calibration

Start by selecting a measured .pak file by clicking on the button *Browse* next to the Measured Spectrum field in the upper left corner of the dialog. Notice that the plot here will not update until the calibration has been performed.

Once a spectrum has been selected, then setup the required parameters for the calibration routine by clicking on *Setup* in the upper right corner of the dialog. This will bring up the dialog shown in Figure 62

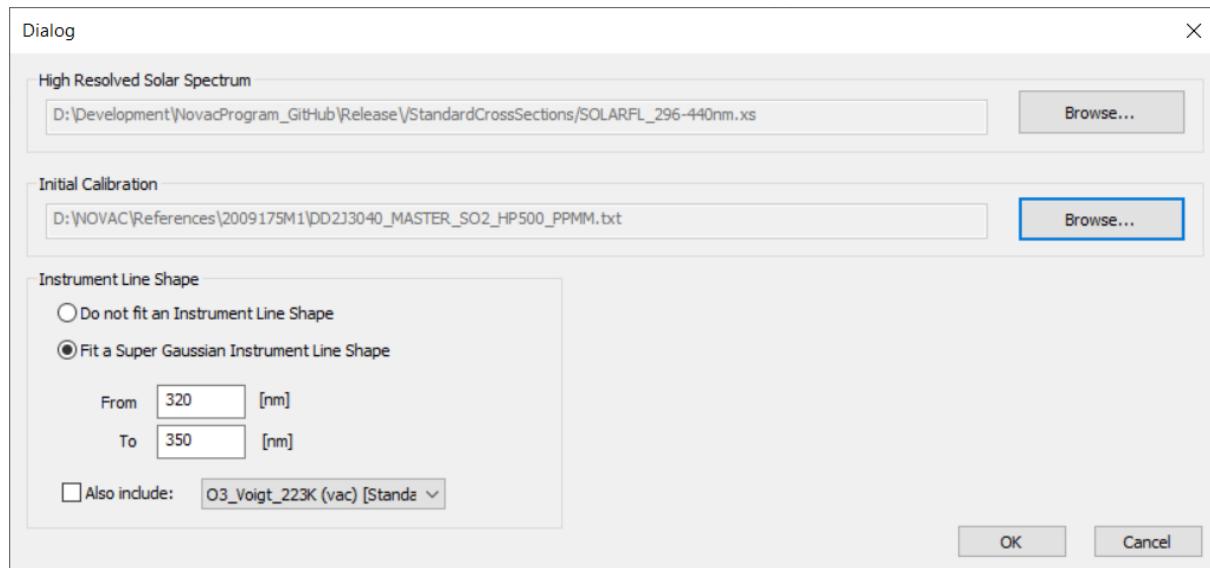


Figure 62. Setup dialog for the instrument calibration

The high resolved solar spectrum should here be filled in automatically by the solar spectrum reference in the supplied *StandardCrossSections*.

The NovacProgram requires an initial wavelength calibration of the device in order to perform the instrument calibration. This is selected by clicking on the button *Browse* next to the Initial Calibration. This will bring up a dialog as shown in Figure 63. There are two options for input here

3. Extended Std File: This is an STD file which must at least contain a wavelength calibration for the device. This can be saved from an earlier instrument calibration using NovacProgram or MobileDOAS. Only the first field named *Calibration* can be filled in here, the other field will be ignored if it is filled in.
4. Wavelength Calibration & Measured Instrument Line Shape: The file selected as the *Calibration* should here point to a file containing a wavelength calibration for the device. This must be a text file containing one or two columns of data where the first (or only) column is the wavelength calibration of the device. A reference file used for evaluation will do here if the file contains two columns of data. The file selected as *Instrument Line Shape* must point to a file containing the instrument line shape in the SLF used by QDOAS, preferably saved from the *Instrument Line Shape* tab in the Instrument calibration dialog. Only the field named *Calibration* is required.

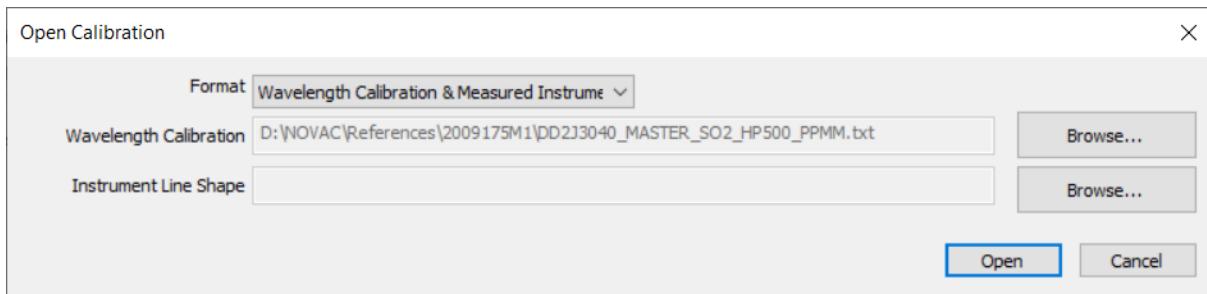


Figure 63. Opening an instrument calibration

Lastly, select if an instrument line shape should be fitted to the measured spectrum in the instrument calibration routine, it is here recommended to select *Fit a Super Gaussian Line Shape* as this gives the best calibration result. If you have provided a measured instrument line shape and are confident that this instrument line shape is correct then the option *Do not fit an Instrument Line Shape* can also be selected.

Finally, there is also an option to include the cross section of O₃ (or other absorbing gases) into the procedure where the instrument line shape is fitted to the measured spectrum. This does not need to be selected in the ordinary case, but is an option to which makes it possible to correct for the effects of ozone absorption in the region where the instrument line shape is fitted. This is expected to only make a difference when the measured spectrum is collected when the sun is close to the horizon and the absorption of O₃ in the measured spectrum is extremely high.

8.1.3 Running the instrument calibration procedure

When all the setup has been done, the instrument calibration procedure can be run by clicking on the button ‘Run’ in the lower right corner. If something is incorrectly setup then a message box will appear with information on what is not correctly setup such that this can be corrected. If not, then the calibration procedure will start. Notice that the instrument calibration can take up to a minute to run and while the procedure is running are the button *Run* disabled.

When the instrument calibration is complete, the buttons in the dialog will be enabled again and a graph will appear in the plot area of the dialog. The graph shows different aspects of the result of the instrument calibration and different types of information can be selected by clicking on the different plot type names in the *Result* selection box to the left of the graph. Select the graph type *Spectra & Polynomial* in the selector in order to view the measured spectrum and the created Fraunhofer reference spectrum in a combined plot. This plot shows several pieces of information which is useful to assess whether the instrument calibration was successful or not:

1. The spectrum drawn in red is the measured, dark-corrected, spectrum.
2. The spectrum drawn in green is the synthetic Fraunhofer spectrum. This has been created using the calculated wavelength calibration and instrument line shape and should match the measured spectrum well over the entire wavelength range.
3. The single solid line drawn across the plot represents the calculated wavelength calibration mapping and the white squares on top of these are the selected points in the spectrum which were used to create the wavelength calibration.



Figure 64. Resulting instrument calibration

Once the calibration routine is complete, the resulting instrument calibration can be saved to file by clicking on the button *Save Calibration* in the lower right corner of the dialog. There are two options for file format here; *Extended STD file* where the calibration is saved into a single STD file, which can be read into DOASIS for reference creation, or *QDOAS Calibrations* where the calibration is saved into a pair of files, a CLB file storing the wavelength calibration and a SLF file storing the instrument line shape. Both of these types of files can be used to create references for the instrument, as described in section 8.2 below.

It is also possible to create a standard set of references for the spectrometer directly by clicking on the button *Save References*. This will bring up the dialog shown in figure Figure 66 where the processing option for creating the references is locked to *high-pass-filter* and the unit is set to *ppmm* as this is the standard for references in NovacProgram. See section 8.2.2 for more information.

8.2 Creating references from instrument calibration

Once an instrument calibration, with both a wavelength calibration and an instrument line shape, have been determined this can be used to create new references for use in evaluation. Open the last tab in the instrument calibration dialog, with the title *Create References*.

This dialog makes use of the *StandardCrossSections* setup which should be located in a sub folder named *StandardCrossSections* in the same folder as the NovacProgram executable file. If these files are missing, please contact the supplier of your DOAS system.

Start by opening an instrument calibration by clicking on the button *Select*. This will bring up the dialog shown in Figure 63 above where an instrument calibration can be opened. Unlike the previous tab, an instrument line shape is required here and when selecting an *Extended STD File*, the file must be produced by the instrument calibration routine in either NovacProgram or MobileDOAS.

Once a valid instrument calibration has been selected the main plot of the dialog will look similar to what is shown in Figure 65 below. The selection box in the upper right corner of the dialog (with the title *High Resolution Cross Section*) controls which absorber the reference is to be created for. This selector will by default list all the settings found in the *StandardCrossSections* setup of the program. If this is empty, please contact the supplier of your mobile DOAS system. If you wish to create a reference for an absorber which is not listed in the standard setup, click on the button *Browse* below the selector to select a high-resolution cross section to be used.

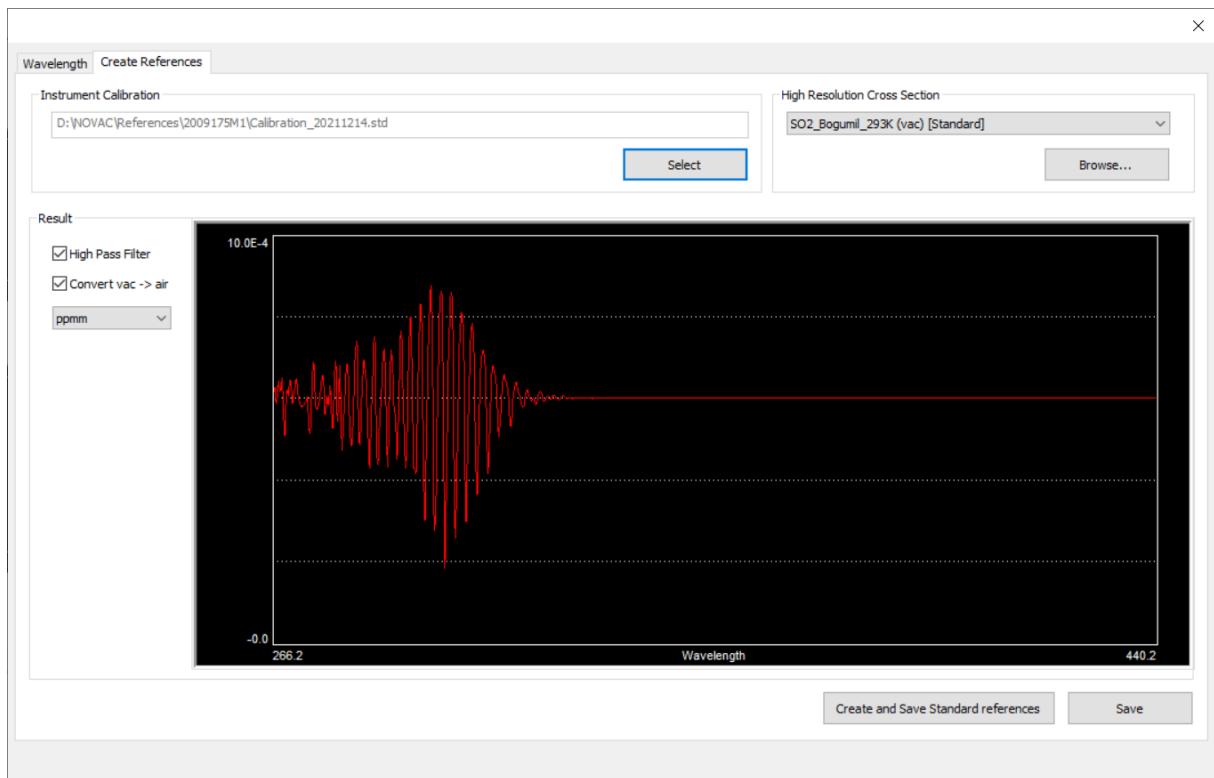


Figure 65. Creating a reference for SO₂

To the left of the graph are three controls which controls how the references are created:

1. *High Pass Filter*: if checked then the created reference will be high-pass filtered before being saved. This is the default setup for references created for the NovacProgram.
2. *Convert vac→air*: references to be used in NovacProgram needs to be in nanometer air, if the high-resolution cross section is in vacuum then the wavelength of the reference needs to be converted to air before it can be used. NOTICE: for the standard cross sections included with NovacProgram the program knows the wavelength unit of the cross section, hence *do not change the state of this check box unless you really know what you are doing*.
3. *Unit* here the selection can be done to create references in ppmm (which is the standard in NovacProgram) or to use molecules/cm². This selector cannot be changed for the included Ring spectrum.

8.2.1 Creating a single reference

The default setup in the *Create References* tab allows for creating and inspecting one reference at a time. Select the high-resolution cross section in the selector in the upper right corner, make sure that the reference looks reasonable and the click on the button *Save* in the lower right corner of the dialog. This saves the reference in a .txt file format which makes it possible to use directly in the evaluation setup (see Figure 24).

Notice that references created for use in real time evaluation in the NovacProgram (i.e., to be inserted into the dialog shown in Figure 24**Fel! Hittar inte referenskälla.**) must be set to use *High Pass Filter* and have the unit *ppmm* selected (with the exception of the Ring spectrum where no unit can be selected).

8.2.2 Creating standard references

To more quickly create a set of the most commonly used references for evaluation, make sure to select *High Pass Filter* and the unit *ppmm* (these are the defaults) and then click the button *Create and Save Standard references* in the bottom right corner of the dialog. This will open a new dialog as seen in Figure 66 below.

Start by selecting the directory where you want to save the new references, enter the name of the instrument in the input field below and optionally a suffix of the file names (e.g., the date the calibration was made, or other information such that the references can be identified later on). Once this is done, verify that the file names listed in the dialog are reasonable and make sure that the names for the SO₂ and O₃ references do include *HP500_PPMM* (if they do not then close the dialog to go back to the *Create References* dialog, select *High Pass filter* and *ppmm* and click on *Create and Save Standard references* again). Clicking on the button *Create and Save* will create the listed reference files and close the dialog.

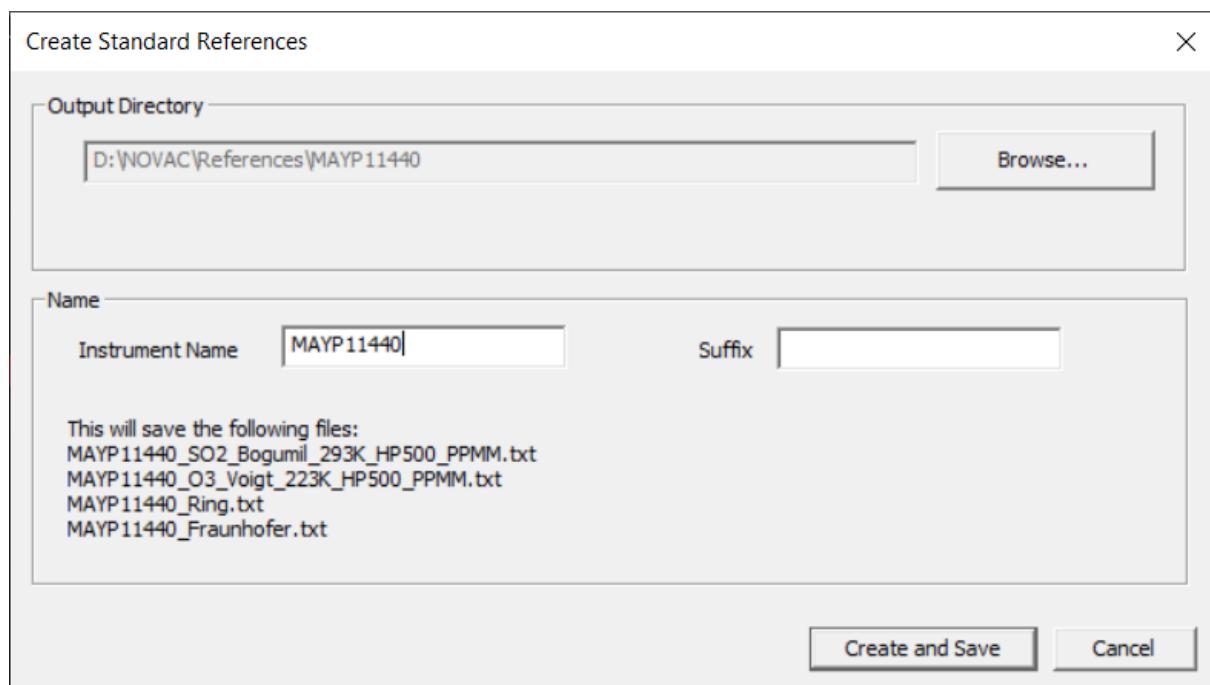


Figure 66. Creating a standard set of references for evaluation

9 Advanced upgrading and troubleshooting

In this section, several advanced techniques for upgrading the software on the embedded PC and troubleshooting the instrument are given. **These techniques are given for reference, but they should only be attempted unless necessary or as a last resort. Using these techniques can render the instrument completely useless if something goes wrong.**

9.1 Installing or re-installing kongo on a MOXA computer

Below are step-by-step instructions on how to install the NOVAC acquisition program (kongo) on a MOXA computer. This procedure is usually performed on new embedded computers. If you are installing kongo on a used MOXA, please be aware that you must first reset the MOXA to factory settings or at least kill the existing kongo software if it is running, otherwise the installation will fail. **Note that formatting / reinstalling should only be done as a last resort if the MOXA is not functioning properly, as this will result in the loss of all data on the embedded computer.**

9.1.1 Formatting the MOXA to the factory settings

If for some reason the system is corrupted may be necessary to reboot the MOXA to the default factory settings. **Notice that you need to backup or have available a copy of the installation file install130614.tar. After the system is formatted, all the configuration files will be deleted including the network settings, and all data will be lost.** The whole procedure requires that the MOXA is connected by a HyperTerminal and an FTP client.

- 1- Reboot the system simply by using the command reboot.
- 2- Immediately, using a clip or a solid wire press the formatting button as shown in the figure below and keep it pressed until the whole process is completed.



The HyperTerminal will show you the following messages:

```
INIT: Switching to runlevel: 6
root@Moxa:/# Stopping portmap daemon:
Stopping internet superserver: inetd.
Sending all processes the TERM signal... done.
Sending all processes the KILL signal... root@Moxa:/# done.
Stopping web server: apache.
Stopping snmpd
Rebooting... Restarting system.
Starting to reboot ...
```

- 3- Still keeping pressed the formatting button, MOXA will beep for 10 seconds before the formatting is completed. In addition to the sound, you will be able to see in HyperTerminal the counting of seconds:

*Press FOR SYSTEM RECOVERY
0,1,2,310*

Before the 10 seconds counting is completed this is your last chance to suspend the procedure

- 4- The following messages will indicate that the procedure is complete:

*Format User Disk...
Erase Flash 0x80a00000-0x81000000 and Length = 0x00600000 bytes
Erase Flash 0x80a00000-0x80a20000 ..OK
Erase Flash 0x80a20000-0x80a40000 ..OK
Erase Flash 0x80a40000-0x80a60000 ..OK
Erase Flash 0x80a60000-0x80a80000 ..OK*

- 5- Then you can release the button and wait for a few seconds until the whole memory is erased. The formatting is completed when you finally get the system's default prompt:

root@Moxa:/#

9.1.2 Killing kongo on a MOXA computer prior to re-installation

If you need to upgrade or re-install kongo on a used MOXA and do not wish to reset it to factory settings (see previous chapter), you must kill the kongo program before proceeding with the installation. Otherwise, the existing software cannot be overwritten.

1. first log in to the MOXA using HyperTerminal, Telnet or PuTTY.
2. Change the working directory to the location /home. This can be done simply by using the command cd /home/
3. Stop the programs that are running in the background - kongo and runsh - by the following commands:

```
killall kongo
killall runsh
```
4. Now that these programs are stopped, follow the steps outlined in the next section to install the new kongo program.

9.1.3 Installing kongo on a MOXA

If the MOXA has been reset to factory settings, note that

- The IP numbers are the factory default:
 Eth1 192.168.3.127 Mask:255.255.255.0
 Eth2: 192.168.3.127 Mask:255.255.255.0

- The ‘novac’ user has been removed and you will need to log in using the following settings:
User: root
Password: root

By FTP, using any of the default Ethernet ports and network/password settings, transfer the compressed installation file as follows: /home/installYYMMDD.tar and continue with the steps below.

For additional questions see:

http://www.moxa.com/doc/man/UC-7112-LX_Plus_7112-LX_7110-LX_HW_Users_Manual_v5.pdf

1. If you didn’t previously install an SD Card, open the MOXA computer and insert a 2GB SecureDigital card. 2 screws need to be removed. Put on the cover again and put back the two screws.
2. Connect all the cables from the Manne-box to the MOXA-computer.
3. Give power to the system.
4. Take the newest install YYMMDD.tar file and transfer that to directory /home in the MOXA box. This can best be done with ftp.
5. Connect to the MOXA using PuTTY or similar program (see section 3).
6. Enter into the /home directory with “cd /home”.
7. Unpack the tar file with “tar –xvf installYYMMDD.tar”.
8. A directory called install will be created. “Enter into that directory with “cd install” .
9. Start the installation by entering “./runme.bat”. This will start the script that’s makes all the installation. After it has finished enter “reboot” to reboot the MOXA computer.
10. The IP address of the MOXA will now be different. It will be set to the NOVAC default. This means that the MOXA will now be accessible at 10.0.0.90 on LAN2 (see section 3.2 for details).
11. When starting up the kongo program will start executing with a test cfg.txt file with a default configuration. Turn on log-printing by entering “showlog”.
12. Make sure that the motor is homing and stops homing when reaching the reference switch.
13. The GPS should be placed where reception is possible. When you see in the log-printing that the system has achieved GPS-time and date then enter “hwclock –s” This will set the internal battery-powered real-time clock in the MOXA box to the received GPS-time. This step is also performed automatically by the kongo program each time the instrument is powered on.
14. Finished

9.1.4 Connecting to the Moxa with a serial cable

The Moxa computer can also be accessed with a serial connection (RS-232). This is an option when the Ethernet connection fails, if serial radio modems are used for telemetry or if the LAN settings need to be modified or accessed. If a serial connection is attempted, no other cable should be simultaneously connected to the Ethernet ports.

For this type of connections, it is necessary to connect with the black serial console cable provided with the instrument, shown below:



Figure 67. Serial console cable of the Moxa computer. The left figure shows the pin distribution of the cable.

An easy way to communicate is via the HyperTerminal program. First the computer used to communicate with the Moxa must have an available COM port (typically provided with older computers, newer computers can use a COM port emulator instead, for example a USB-to-serial adaptor that creates a virtual COM port). Then open HyperTerminal and configure the available port in your computer according to these settings:



Figure 68. Settings for HyperTerminal communication

Then a session opens up and you could type the commands listed in section ‘Useful commands for troubleshooting’. There is no need to provide credentials for login in this case.

For previous versions of the computer check *Appendix A: Previous versions of the NOVAC instrument* on page **Fel! Bokmärket är inte definierat..**

9.2 Installing or re-installing kongo on Axiomtek

Note that formatting / reinstalling should only be done as a last resort if the Axiomtek computer is not functioning properly.

If kongo was previously installed on the Axiomtek computer, you may want to reset it to factory settings before proceeding with the installation. Consult the IFB122 Linux User's Manual for instructions on how to do this. Alternatively, you may kill the kongo and runsh processes before proceeding (killall kongo, killall runsh). If the Axiomtek is new, these steps can be skipped.

https://www.axiomtek.com/Download/download/IFB122/IFB122%20Linux%20User's%20Manual%20VA4_07-18-2018.pdf

- The factory default network settings for the Axiomtek **LAN2** port are:
IP: 192.168.0.254 Mask: 255.255.255.0

- SSH, SCP and SFTP have the following settings:

User: root
no password

Follow the steps below:

1. Connect power cables to the computer and attach Ethernet cable to LAN2 port.
2. Give power to the system (+ override timer, if necessary).
3. Connect via WinSCP using the SCP connection option.
 - a. Eth2: 192.168.0.254 ; Mask: 255.255.255.0;
 - b. Username: root ; no password
4. Take the newest install_YYMMDD_tar file (currently install211105_tar) and transfer that to /home/root directory in the Axiomtek computer.
5. In /home/root directory, also copy the most recent “cfg.txt” file (the configuration file)
6. In /etc/network directory, replace “interfaces” file with appropriate file for the instrument and network (this file specifies the computer’s IP addresses)
7. Connect to the Axiomtek using PuTTY or similar program (see section 3.2).
8. Enter into the /home/root directory with “cd /home/root”.
9. Unpack the tar file with “tar –xvf installYYMMDD_tar”.
10. A directory called installYYMMDD will be created. “Enter into that directory with “cd installYYMMDD” (according to date of tar file)
11. Start the installation by entering “./runme.sh”. This will start the script that does the kongo installation.

- a. Note: don't worry if the following error message appears: "*cp: can't stat 'cfg.txt': no such file or directory*". This just indicates that there is no embedded cfg default file, and one needs to provide a copy separately (which was done in previous steps)
12. After it has finished enter "reboot" to reboot the Axiomtek computer.
13. The IP address of the Axiomtek will now be different. It will be set to the information in the interfaces file you uploaded. As per the NOVAC default, the Axiomtek will now be accessible at the IP address specified in the interfaces file on LAN1 and at the IP address 10.0.0.90 on LAN2 (see section 3.2 for details). You will need to change the IP address of the computer contacting the Axiomtek to be on the new IP network.
14. Log back into the Axiomtel in PuTTY using the standard novac login:
 - a. User: novac
 - b. Passwd: 1225
15. For security, change password of "root" user to be the standard novac password:
 - a. use the command: *passwd root*
 - b. set the new password to be the default novac password: 1225
16. When starting up the kongo program will start executing with the cfg.txt file that you uploaded. Turn on log-printing by entering "showlog" and check to see it's working.
17. The GPS should be placed where reception is possible. When you see in the log-printing that the system has achieved GPS-time and date then enter "hwclock -s" This will set the internal battery-powered real-time clock in the AXIOMTEK box to the received GPS-time. This step is optional, as this process is also automatically run every time the instrument is powered on and obtains GPS lock.
18. Finished

9.3 Setting the default baud rate of an SD2000 spectrometer

In order for the embedded computer to be able to communicate with the Ocean Optics spectrometer (either SD2000 or S2000), the default baud rate of the spectrometer must be set to 115,200. The default baud rate is the baud rate that the spectrometer uses when the power is first turned on. When spectrometers are received from Ocean Optics, their default baud rate is often set to 9600. In this condition, the embedded computer will not be able to communicate with the spectrometer.

To change the default baud rate, connect the spectrometer to a laptop computer with a serial cable (not a null modem cable). Connect power to the spectrometer via the separate power port. Once the green light has turned on on the spectrometer, open PUTTY or a similar program and connect to the spectrometer using the settings

Baud rate: 9600 (unless it has already been changed)

Data bits: 8
Stop bits: 1
Parity: None
Flow control: None

Type **aA** to enter ASCII mode

Check that you are in ASCII mode by performing a scan. Enter **S<Enter>**. This should return 2048 numbers that are clearly legible if you are in ASCII mode.

To change the power-up baud rate:

Type **x43<Enter>**

Type **Baud 6<Enter>**

To check if the parameter change was successful, you can query the parameter by entering **?x43<Enter>**. This should return ‘Baud 6’.

Cycle the power on the device. It should now communicate at 115,200 baud.

Note that if you want to use a different baud rate for any reason, the following table lists the possible options. But the embedded computer only communicates at 115,200 baud.

Baud 0 = 1200
Baud 1 = 2400
Baud 2 = 9600
Baud 3 = 19,200
Baud 4 = 38,400
Baud 5 = 57,600
Baud 6 = 115,200

9.4 Instrument configuration not recognized

The measurement configuration for each instrument is stored in the file cfg.txt on the embedded computer. This file has a very specific format which is described in section 3.2.4 of this manual. If the file format is compromised in some way, the instrument may not be able to recognize the cfg.txt file as a configuration file, and a default configuration will be assumed.

Unfortunately, this default configuration will generally not be useful. Upon reboot, the scanner will move to the end position, then rotate to a fixed direction and measure a single spectrum called 'scan'. This single acquisition will be repeated, each time starting a new *.pak file on the SD card. In other words, each saved *.pak file will only contain a single 'scan' spectrum.

Also, the DELAY parameter will default to 0, which may make the scanner spin at a very high speed, or since the delay is meant to ensure proper stepping of the motor, it may mean that the scanner does not move at all. **If the scanner is spinning very quickly despite the fact that the user has set the DELAY parameter to a reasonable value (e.g. 500), then there is likely a problem with the formatting of the cfg.txt file.**

One common yet very subtle problem with the cfg.txt file formatting is improper use of 'end of line' characters. If the wrong end of line characters are used, the text file may seem completely fine when opened in a text editor on a computer, but may still not be recognized by the computer embedded in the NOVAC instrument.

The embedded computer will recognize the cfg.txt file if each line of the text file ends with both a 'carriage return' (CR) and 'line feed' (LF) character. The text file may not work if only one of these characters is present. The way to check this is to open the cfg.txt file in a text editor and make the special characters visible. One editor capable of this is the freely available Notepad++. In this software, clicking on the 'Show All Characters' button in the quick access bar at the top will make special characters visible. The end of line convention can then be set using the menu. Selecting Edit -> EOL Conversion -> Windows will ensure that each line ends in CR LF. After reformatting the text file, overwrite the existing cfg.txt on the embedded computer. Hopefully, the instrument will now recognize the configuration.

Appendix A: Special Measurement Modes

The performed measurements will, by the NovacProgram, be categorized in any of the following measurement modes. Many of these modes will not be used in a typical measurement setup but are rather for special research purposes only.

Mode	Description
Flux	Normal measurement mode. Full scan from horizon to horizon with variable or fixed exposure-time.
Wind speed	Measurement of plume-speed using the dual-beam technique. Contains at least 50 measurements in the same direction and performed at a time when the Solar Zenith Angle is less than 75°.
Composition	Measurement with a few spectra collected inside the plume (with a high number of co-adds) and variable exposure-time throughout the scan. Has no more than 15 spectra and at least one of them has the name ‘comp’
Stratosphere	Measurement of stratospheric compounds. Contains at least 3 measurements in the zenith and performed at a time when the Solar Zenith Angle is more than or equal to 75°.
Troposphere	--
Max-DOAS	--
Direct-sun	Measurement aimed directly at the sun. Contains at least 5 spectra with the exact name ‘direct_sun’
Lunar	Measurement aimed directly at the moon. Contains at least 5 spectra with the exact name ‘lunar’

In a flux-measurement, all spectra will be used to calculate the flux except for spectra which have names starting with any of the following (case insensitive);

- Sky
- Dark
- Offset
- Sun_
- Moon_
- Strat
- Trop
- Maxdoas
- Wind
- Special

Appendix B: Configuring Point-to-Multipoint FreeWave Radio modems

This part is about how to configure radio modems when multiple remote scanners are applied via a FreeWave Radio Network.

Basically the radio modem at observatory works as a master. The radio modems connecting to the remote scanners work as slaves. According to the geographical environment, choose whether and where repeaters are used. Because in this network structure one radio modem controls all other radio modems, this network is call point-to-multipoint network.

Steps to configure the FreeWave radio modem

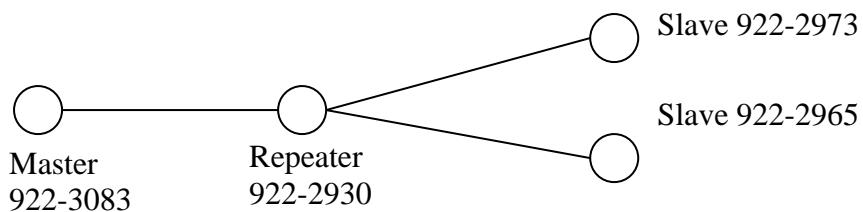
- Connect **6V~ 30V** power to the radio modem.
- Connect a RS232 serial cable to the COM port on your PC and the radio modem.
- Start HyperTerminal. Set data rate to **19,200**, data bits - **8**, Parity- **none**, Stop bits – **1**, Flow control – **none**.
- On the back panel of the FreeWave press the Setup Button for 1 second. The three lamps will become solid green.

Configuration of Master to Multiple Slaves through Repeater

In this network configuration FreeWave transceivers can utilize up to 4 Repeaters. However, there is no limitation on baud rate setting when repeater is used. When a Repeater is used, the RF throughput is cut in half. When there is no repeater used, the sustained data throughput is 115.2 kbps. When repeaters are used, the sustained data throughput is 57.6 kbps.

Master to Multiple Slaves through One Repeater

For example, we have four radio modems with radio ID 922-3083, 922-2973, 922-2965 and 922-2930. We will have 2 slaves, one repeater and one master in this network.



The configurations are as this table:

Radio ID	Modem Mode	Call book		
		Entry Number	Repeater1	Repeater2
922-3083	Point-to-Point Master/Slave Switchable	922-2930 922-2973 922-2965	922-2930 922-2973 922-2965	

922-2930	Point to Point Slave/Repeater	922-3083		
922-2973	Point to Point Slave	922-3083	922-2930	
922-2965	Point to Point Slave	922-3083	922-2930	

I. Configure FreeWave MASTER

The master is working on at observatory site where NOVAC software resides. In the configuration interface in HyperTerminal, configure as follows:

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Master/Slave Switchable”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input serial numbers of radio modems that this radio modem will connect to. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 922-2930		
(1) 922-2973	922-2930	
(2) 922-2965	922-2930	
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

II. Configure Repeater

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Slave/Repeater”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input *only* the serial number of the master radio modem. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 922-3083		
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

III. Configure Remote Stations

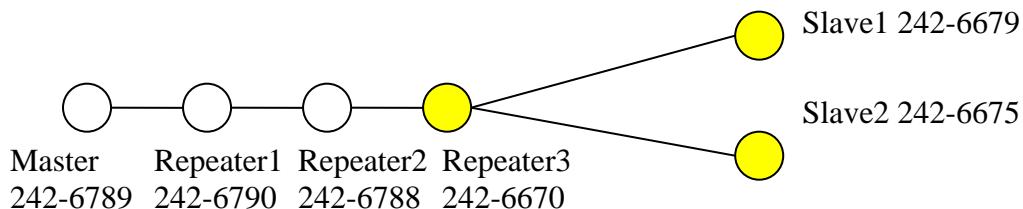
The remote stations are configured as the following:

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Slave”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input the serial number of the master radio modem as “Entry number”. Input the serial number of the repeater respectively. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 922-3083	922-2930	
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

Master to Multiple Slaves through Several Repeaters

For example, we have 6 radio modems with radio ID 242-6789, 242-6790, 242-6788, 242-6670, 242-6679 and 242-6675. We will have 2 slaves, 3 repeaters and one master in this network. Each of Repeater3, Slave1 and Slave2 is connected to one Scanning DOAS.



The configurations of the modems are as this table:

Radio ID	Modem Mode
242-6789	Point-to-Point Master/Slave Switchable
242-6790	Point to Point Repeater
242-6788	Point to Point Repeater
242-6770	Point to Point Slave/Repeater
242-6679	Point to Point Slave
242-6675	Point to Point Slave

I. Configure FreeWave MASTER

The master is working on at observatory site where NOVAC software resides. In the configuration interface in HyperTerminal, configure as follows:

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Master/Slave Switchable”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input serial numbers of radio modems that this radio modem will connect to. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 242-6790		
(1) 242-6788	242-6790	
(2) 242-6670	242-6790	242-6788
(3) 242-6679	242-6790	242-6788
(4) 999-9999	242-6670	
(5) 242-6675	242-6790	242-6788
(6) 999-9999	242-6670	
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

Entry 0 is the repeater1 (242-6790), which is directly linked with the master.

Entry 1 is repeater2 (242-6788), which is linked with the master through repeater1 (242-6790).

Entry 2 is repeater3 (242-6670), which is linked with the master through repeater1(242-6790) and then repeater2 (242-6788).

At Entry 3, slave1 (242-6679) uses 3 repeaters. The master calls slave1 through repeater1 (242-6790), repeater2 (242-6788), and repeater3 (242-6670). There are only 2 positions for repeaters, so the row of entry 4 begins with 999-9999 to continue “Entry 3”. It is same for Entry 5.

II. Configure Repeater1, Repeater2

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Repeater”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”.

Repeater1 is directly connected to the master. So input only the serial number of the master radio modem. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 242-6789		
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

Repeater2's configuration is same in first two steps. Comparing with repeater1 it just has a different call book. It is connected to the master through repeater1. So for repeater2 the call book will look like:

Entry Number	Repeater1	Repeater2
(0) 242-6789		242-6790
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

III. Configure Repeater3

Repeater3 is connected with one Scanning DOAS instrument, so it is not a simple repeater. It uses a different mode than other repeaters.

- Choose “(0) Set Operation Mode”. Set it to “Point-to-Point Slave/Repeater”.
- Choose “(1) Set Baud Rate” Set correct baud rate that you will use in the network, for example 115200.
- Choose “(2) Edit Call Book”. Input the serial numbers of the master and repeaters. It connects to the master (242-6789) through 242-6788 first, and then 242-6790. So the call book will look like:

Entry Number	Repeater1	Repeater2
(0) 242-6789	242-6788	242-6790
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

IV. Configure Remote Stations

A remote station is configured as the following steps using slave1 as example.

- Choose “(0) Set Operation Mode”. Set it to “Point-to-Point Slave”.

- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input the serial number of the master radio modem as “Entry number”. Input the serial number of the repeater respectively. The slave1 calls the master (242-6789) through 242-6670, 242-6788, and then 242-6790. Because there are more than 2 repeaters, the second line continues the first line by put 999-9999 at the first position then the 3rd repeater. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 242-6789	242-6670	242-6788
(1) 999-9999		242-6790
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

Configuration of Master to Multiple Slaves without Repeater

I. Configure FreeWave MASTER

The master is working on at observatory site where NOVAC software resides. Since there is no repeater used, leave the repeater 1 and repeater 2 columns blank. In the configuration interface in HyperTerminal, configure as follows:

- Choose “**(0) Set Operation Mode**”. Set it to “Point-to-Point Master/Slave Switchable”.
- Choose “**(1) Set Baud Rate**” Set correct baud rate that you will use in the network, for example 115200.
- Choose “**(2) Edit Call Book**”. Input serial numbers of radio modems that this radio modem will connect to. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 922-2973		
(1) 922-2965		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

II. Configure Remote Stations

The remote stations are configured as the following:

- Choose “(0) Set Operation Mode”. Set it to “Point-to-Point Slave”.
- Choose “(1) Set Baud Rate” Set correct baud rate that you will use in the network, for example 115200.
- Choose “(2) Edit Call Book”. Input the serial number of the master radio modem as “Entry number”. The call book will look like:

Entry Number	Repeater1	Repeater2
(0) 922-3083		
(1) 000-0000		
(2) 000-0000		
(3) 000-0000		
(4) 000-0000		
(5) 000-0000		
(6) 000-0000		
(7) 000-0000		
(8) 000-0000		
(9) 000-0000		

Test Connection

When all FreeWaves are in the above configuration, you will see only solid Red light on the left (CD) and Blinking Red on the right (CTS) on all radios, which means that the radio modems are not connected.

To test whether your radios are configured correctly, you can use either the HyperTerminal to type the string ATD, at which time you should get an 'OK' from the Master Radio. Then carefully, for you will not see an 'echo' type ATDXXXXXXX where X is the 7 digit ID of the Repeater or a Remote Slave. Immediately, without hitting a C/R, the Master radio should display Green CD, as will the Repeater, and the Slave whose number you typed in. At the

same time you will see, on your screen CONNECTED - if you are in fact connected to the Slave FreeWave. To test another link, disconnect this connection by click  in the HyperTerminal and then connect to another radio modem.

The NOVAC program works same no matter how the point-to-multipoint network is configured, as long as the “radioID/callbook number” is filled in correctly.

Appendix C: Preparing an Ocean Optics MayaPro spectrometer for use in NOVAC instruments

The Ocean Optics MayaPro spectrometer features an uncooled back-thinned CCD detector. These are more sensitive to ultraviolet light than the standard CCD sensors installed in most other Ocean Optics spectrometers. This makes the MayaPro an attractive choice particularly for high-latitude environments where ultraviolet light is limited.

Using the MayaPro in a NOVAC instrument requires specialized firmware to be installed on the spectrometer. This firmware version enables the user to set the power-up baud rate of the spectrometer to 115,200, the speed at which the embedded computer handles serial communication. Below is a description of how to configure a MayaPro for use in a NOVAC instrument.

What you will need:

- Ocean Optics MayaPro Spectrometer
- Ocean Optics HR4-CBL-DB15 adapter cable
- DC power female barrel connector, 1.3mm inside contact, 3.4mm outside contact (e.g. Kobiconn P/N 163-0309-EX available from Mouser Electronics)
- Female DB9 (serial port) connector
- USB cable
- Serial cable
- PuTTY or other serial console software (available for free at www.putty.org)
- USBProgrammer software (available from the Software Downloads section of the Ocean Optics website)
- HRFPGA.HEX (NOVAC configuration file for USBProgrammer)
- MayaPro2000 v1031.iic (spectrometer FX/2 firmware)
- MAYAPROv3000.xsvf (spectrometer FPGA firmware)
- install180617.tar (installer for configuring the embedded PC in the NOVAC instrument)

Step by step instructions:

Install and configure USB Programmer

- Install USBProgrammer on a computer. If using a desktop, beware that a power outage during firmware update could corrupt the spectrometer and render it unusable. Consider using a laptop that is either plugged in or has a charged, reliable battery to avoid this problem. IF USING A LAPTOP, DO NOT LET IT GO TO SLEEP DURING THE UPDATE PROCESS!
- Rename the file in C:\Program Files\Ocean Optics\USBProgrammer\HRFPGA.hex to HRFPGA_OLD.hex
- Copy the NOVAC hex file HRFPGA.HEX into that same directory.

Update the FX/2 firmware

- Connect the MayaPro to the USB port.
- Run the USB Programmer by selecting Start | Programs | Ocean Optics | USB Programmer |USB programmer. This utility will display your Maya2000 Pro in the left hand pane.
- Go to File | Unlock Features. In the new little window type UNLOCKALL | click Ok
- Go to File | Unlock Features. In the new little window type UNLOCKFPGA | click Ok
- Go to File | Unlock Features. In the new little window type UNLOCKMFG | click Ok
- Click on the Maya2000 Pro serial number to select it. Once you select the device, detailed information for that device is displayed in the right pane of the interface.
- Press the 'Program FX/2' button (Not the Program FPGA button!) and select the MayaPro2000 v1031.iic file.
- Programming begins immediately, and the estimated time for completion of the reprogramming process displays in the bottom status bar. DO NOT INTERRUPT the programming process for any reason! If the programming process is interrupted, permanent corruption can occur in the device. Programming the FX/2 will take about 2 minutes.
- Once the programming is complete, a pop up window will appear telling you to power down the spectrometer. Unplug the spectrometer from the computer.
- Shut down the USB Programmer software.
- Wait 10 seconds

Update the FPGA firmware

- Plug the spectrometer back into the USB port.
- Restart USB Programmer
- Go to File | Unlock Features. In the new little window type UNLOCKALL | click Ok
- Go to File | Unlock Features. In the new little window type UNLOCKFPGA | click Ok
- Go to File | Unlock Features. In the new little window type UNLOCKMFG | click Ok
- Click on the Maya2000 Pro serial number to select it. Once you select the device, detailed information for that device is displayed in the right pane of the interface.
- Press the 'Program FPGA' button (Not the Program FX/2 button!) and select the MAYAPROv3000.xsvf file.
- Programming begins immediately, and the estimated time for completion of the reprogramming process displays in the bottom status bar. DO NOT INTERRUPT the programming process for any reason! If the programming process is interrupted, permanent corruption can occur in the device. Programming the FX/2 will take about 15 minutes.

- Once the programming is complete, a pop up window will appear telling you to power down the spectrometer. Unplug the spectrometer from the computer.

Configure the spectrometer's startup baud rate

- Connect the spectrometer to the computer using a serial cable. You might use the HR4000 Break-Out Box to make this connection.
- Power the spectrometer, e.g. by also connecting it to the computer via USB cable.
- Using PuTTY or other serial console, connect to the spectrometer using these settings
 - Baud rate: 9600; Data bits: 8; Stop bits: 1; Flow control: None
- Type 'aA' to set the spectrometer to ASCII mode
- Type 'S' to confirm that the spectrometer is acquiring spectra. If everything is going right, you should receive 2048 spectral intensities, one from each pixel.
- Type '?x18<ENTER>' to query the start up baud rate. The spectrometer should answer '2' indicating a baud rate of 9,600.
- Type 'x18<ENTER>'. The spectrometer will wait for further input. Type '6<ENTER>'. This should change the start up baud rate to 115,200.
- Check to make sure it worked by typing '?x18<ENTER>'. The spectrometer should now answer '6'.
- Unplug the USB cable and close PuTTY.
- Wait 10 seconds
- Plug the USB cable back in.
- Open PuTTY, but this time connect with 115,200 baud.
- Again, type 'aA' followed by 'S'. You should now receive the spectrum much quicker than before.

Connecting the Maya Pro to the NOVAC system

- The Maya Pro needs to be connected to the serial port of the embedded PC. It also needs to be supplied with power from the Octopus electronics. Currently, the best way to accomplish this is by building a custom adapter that plugs into the 30-pin connector next to the Maya's USB port.
- One way to build this custom connector is by starting off with an Ocean Optics HR4-CBL-DB15 adapter. This adapter can be ordered with the Maya Spectrometer.
- Open the housing of the DB15 connector and remove the black heat shrink and any tape on this end of the adapter.
- Identify the 5 pins required for the NOVAC adapter (see pin orientation diagram and connection table below) and solder these to the appropriate pins on the serial DB9 and DC power barrel connectors. Use heat shrink to isolate the conductors.
- Once the adapter has been built, plug the cable coming from serial port 1 of the embedded PC to the female DB9 connector. Plug the power cable coming from the Octopus into the female barrel connector. The Maya Pro should now receive power from the Octopus and be ready to communicate with the embedded PC.



Figure 69 – The original Ocean Optics HR4-CBL-DB15 adapter cable can be ordered from Ocean Optics

Pin orientation	
USB port	2 4 6 8 10 12 14 16 18 20 22 24 26 28 30 1 3 5 7 9 11 13 15 17 19 21 23 25 27 29
Looking at Front of Maya	

Maya 30-pin connector	Function	NOVAC connectors
Pin 1	RS232 Rx	Female DB9 Pin 3
Pin 2	RS232 Tx	Female DB9 Pin 2
Pin 5	Gnd	Female DB9 Pin 5
Pin 12	5V in	Barrel connector inside contact
Pin 29	Gnd	Barrel connector outside contact

Table 1 -Pin orientation and required connections for the custom NOVAC adapter cable



Figure 70 – Adapter cable plugged into Maya Pro spectrometer after modification for use in a NOVAC system. The adapter now has a DB9 serial connector and a DC power socket.

Update the embedded PC

- At the time of this writing (March 2020), the Maya spectrometer has been successfully tested on the MOXA embedded PC with kongo software version 180617. If using a MOXA, it is recommended that a fresh install of this software be made. If using an Axiomtek embedded PC, the most current software version should be used, and the embedded PC may not have to be updated as per the instructions listed in the next few steps.
- If needed, reset the embedded PC to factory settings (see instructions in this manual).
- Using the file install180617.tar on MOXA, configure the embedded PC using the appropriate software version. See section 9.1
- Remember to (re-)configure the CFG.txt and /etc/network/interfaces files after performing this update.

Modifying the CFG.TXT file

- The CFG.TXT is located on the embedded computer at /mnt/flash. Download the file to your computer. See section 3.2 "Configuring the instrument in the field" for detailed instructions.
- Ensure the following parameters are set in the configuration:
 - SPECTROMETERTYPE=MAYAPRO
 - INSTRUMENTNAME=[SERIAL NUMBER] (use your spectrometer's serial number, e.g. MAYP111125)
- Once the CFG.TXT has been modified, upload it back to the embedded PC, overwriting the original file. Restart the PC.

Configuring the NOVAC Program

- The proper configuration of the NOVAC instrument using the MayaPro spectrometer depends on the type of embedded PC (and version of kongo.exe) you are using.
- If using a MOXA embedded PC and kongo.exe version 180617, then the NOVAC instrument will convert each acquired spectrum to a format compatible to an S2000 spectrometer. To do this, the “Number of Spectra” parameter will be multiplied by an additional factor of 16. Thus, when the NOVAC Program receives a spectrum it will divide each spectrum by 16, lowering the maximum pixel intensity from 65,535 to 4095. The spectra can then be handled as if they had been acquired by an S2000 or similar spectrometer. When using this setup, the NOVAC instrument therefore needs to be configured in the NOVAC Program as using an S2000 spectrometer, even though a Maya Pro is actually being used.
- If, on the other hand, a newer Axiomtek embedded PC is being used along with a more current version of kongo.exe, then the NOVAC instrument should be configured as a Maya Pro spectrometer in the NOVAC Program. This newer version of the data acquisition software does not divide each spectrum by 16. Hence, the NOVAC Program can be configured to expect 65,535 counts as a maximum pixel intensity.