

# Physics – Vật Lý

Nguyễn Quân Bá Hồng\*

Ngày 16 tháng 11 năm 2024

## Tóm tắt nội dung

This text is a part of the series *Some Topics in Advanced STEM & Beyond*:

URL: [https://nqbh.github.io/advanced\\_STEM/](https://nqbh.github.io/advanced_STEM/).

Latest version:

- *Physics – Vật Lý*.

PDF: URL: [https://github.com/NQBH/advanced\\_STEM\\_beyond/blob/main/physics/NQBH\\_physics.pdf](https://github.com/NQBH/advanced_STEM_beyond/blob/main/physics/NQBH_physics.pdf).

TeX: URL: [https://github.com/NQBH/advanced\\_STEM\\_beyond/blob/main/physics/NQBH\\_physics.tex](https://github.com/NQBH/advanced_STEM_beyond/blob/main/physics/NQBH_physics.tex).

## Mục lục

<b>1</b>	<b>Wikipedia</b>	<b>2</b>
1.1	Wikipedia/advection	2
1.1.1	Distinction between advection & convection	3
1.1.2	Meteorology	3
1.1.3	Other quantities	3
1.1.4	Mathematics of advection	3
1.2	Wikipedia/computational physics	4
1.2.1	Overview	4
1.2.2	Challenges in computational physics	5
1.2.3	Methods & algorithms	5
1.2.4	Divisions	5
1.2.5	Applications	6
1.3	Wikipedia/diffusion	6
1.3.1	Diffusion vs. bulk flow	6
1.3.2	Diffusion in the context of different disciplines	7
1.3.3	History of diffusion in physics	7
1.3.4	Basic models of diffusion	8
1.3.5	Diffusion in physics	8
1.3.6	Random walk (random motion)	8
1.4	Wikipedia/dynamics (mechanics)	8
1.4.1	Principles	8
1.4.2	Linear & rotational dynamics	8
1.4.3	Force	8
1.4.4	Newton's laws	8
1.5	Wikipedia/fluid mechanics	9
1.5.1	Brief history	9
1.5.2	Main branches	9
1.5.3	Relationship to continuum mechanics	10
1.5.4	Assumptions	10
1.5.5	NSEs	10
1.5.6	Inviscid & viscous fluids	11
1.5.7	Newtonian vs. non-Newtonian fluids	11
1.6	Wikipedia/external flow	12
1.7	Wikipedia/Gauss's principle of least constraint	12
1.7.1	Statement	12
1.7.2	Hertz's principle of least curvature	13
1.8	Wikipedia/hydrostatics	13
1.8.1	History	13
1.8.2	Pressure in fluids at rest	14

---

\*A Scientist & Creative Artist Wannabe. E-mail: [nguyenquanbahong@gmail.com](mailto:nguyenquanbahong@gmail.com). Bến Tre City, Việt Nam.

1.8.3	Liquids (fluids with free surfaces)	15
1.9	Wikipedia/mathematical physics	16
1.9.1	Scope	16
1.9.2	Usage	16
1.9.3	Prominent mathematical physicists	17
1.10	Wikipedia/matrix mechanics	17
1.10.1	Development of matrix mechanics	17
1.10.2	Mathematical development	17
1.10.3	Further results	18
1.11	Wikipedia/numerical methods in fluid mechanics	18
1.11.1	Discretization	18
1.11.2	FDM	18
1.11.3	FEM	19
1.11.4	Spectral method	19
1.11.5	FVM	19
1.11.6	Computational cost	19
1.11.7	Forward Euler approximation	19
1.11.8	Backward difference	19
1.12	Wikipedia/physics	19
1.13	Wikipedia/stationary-action principle	20
1.13.1	General statement	20
1.13.2	Origins, statements, & controversy	20
1.13.3	Further development	21
1.13.4	Disputes about possible teleological aspects	21
1.14	Wikipedia/statistical mechanics	22
1.14.1	History	22
1.14.2	Principles: mechanics & ensembles	22
1.14.3	Statistical thermodynamics	23
1.14.4	Non-equilibrium statistical mechanics	23
1.14.5	Applications	23
1.15	Wikipedia/universal (metaphysics)	24
1.15.1	Problem of universals	24
1.15.2	Particular	24
1.15.3	Platonic realism	25
1.15.4	Nominalism	25
1.15.5	Ness-ity-hood principle	25
2	RICHARD P. FEYNMAN, ROBERT B. LEIGHTON, MATTHEW SANDS. <b>6 Easy Pieces: Essentials of Physics Explained By Its Most Brilliant Teacher</b>	25
2.1	FEYNMAN's Preface from The Feynman Lectures on Physics	30
2.2	Atoms in Motion	31
2.2.1	Introduction	31
2.2.2	Matter is made of atoms	32
2.3	Basic Physics	33
2.4	The Relation of Physics to Other Sciences	33
2.5	Conservation of Energy	33
2.6	The Theory of Gravitation	33
2.7	Quantum Behavior	33
3	RICHARD FEYNMAN. <b>The Feynman Lectures on Physics</b>	33
4	<b>Miscellaneous</b>	33
	<b>Tài liệu</b>	33

# 1 Wikipedia

## 1.1 Wikipedia/advection

“In the field of **physics**, **engineering**, & **earth sciences**, *advection* is the **transport** of a substance or quantity by bulk motion. The properties of that substance are carried with it. Generally the majority of the advected substance is a fluid. The properties that are carried with the advected substance are **conserved** properties e.g. **energy**. An example of advection is the transport of **pollutants** or **silt** in a **river** by bulk water flow downstream.

Another commonly advected quantity is energy or **enthalpy**.

Here the fluid may be any material that contains thermal energy, e.g. **water** or **air**.

In general, any substance or conserved, **extensive** quantity can be advected by a fluid that can hold or contain the quantity or substance.

During advection, a fluid transports some conserved quantity or material via bulk motion.

The fluid's motion is described mathematically as a **vector field**, & the transported material is described by a **scalar field** showing its distribution over space.

Advection requires currents in the fluid, & so cannot happen in rigid solids.

It does not include transport of substances by **molecular diffusion**.

Advection is sometimes confused with the more encompassing process of **convection** which is the combination of advective transport & diffusive transport.

In **meteorology** & **physical oceanography**, advection often refers to the transport of some property of the atmosphere or **ocean**, e.g., **heat**, humidity (see **moisture**) or **salinity**.

Advection is important for the formation of **orographic** clouds & the precipitation of water from clouds, as part of the **hydrological cycle**.

### 1.1.1 Distinction between advection & convection

The term *advection* often serves as a synonym for **convection**, & this correspondence of terms is used in the literature.

More technically, convection applies to the movement of a fluid (often due to density gradients created by thermal gradients), whereas advection is the movement of some material by the velocity of the fluid.

Thus, somewhat confusingly, it is technically correct to think of momentum being advected by the velocity field in the Navier-Stokes equations, although the resulting motion would be considered to be convection.

Because of the specific use of the term convection to indicate transport in association with thermal gradients, it is probably safer to use the term advection if one is uncertain about which terminology best describes their particular system.

### 1.1.2 Meteorology

In **meteorology** & **physical oceanography**, advection often refers to the horizontal transport of some property of the atmosphere or **ocean**, e.g., **heat**, humidity or salinity, & convection generally refers to vertical transport (vertical advection).

Advection is important for the formation of **orographic clouds** (terrain-forced convection) & the precipitation of water from clouds, as part of the **hydrological cycle**.

### 1.1.3 Other quantities

The advection equation also applies if the quantity being advected is represented by a **probability density function** at each point, although accounting for diffusion is more difficult.

### 1.1.4 Mathematics of advection

The *advection equation* is the **PDE** that governs the motion of a conserved **scalar field** as it is advected by a known **velocity vector field**.

It is derived using the scalar field's **conservation law**, together with **Gauss's theorem**, & taking the **infinitesimal** limit.

One easily visualized example of advection is the transport of ink dumped into a river.

As the river flows, ink will move downstream in a “pulse” via advection, as the water's movement itself transports the ink.

If added to a lake without significant bulk water flow, the ink would simply disperse outwards from its source in a **diffusive** manner, which is not advection.

Note that as it moves downstream, the “pulse” of ink will also spread via diffusion.

The sum of these processes is called **convection**.

**The advection equation.** In Cartesian coordinates the advection **operator** is

$$\mathbf{u} \cdot \nabla = u_x \partial_x + u_y \partial_y + u_z \partial_z,$$

where  $\mathbf{u} = (u_x, u_y, u_z)$  is the **velocity field**, &  $\nabla$  is the **del** operator (note that **Cartesian coordinates** are used here).

The advection equation for a conserved quantity described by a **scalar field**  $\psi$  is expressed mathematically by a **continuity equation**:

$$\partial_t \psi + \nabla \cdot (\psi \mathbf{u}) = 0,$$

where  $\nabla \cdot$  is the divergence operator & again  $\mathbf{u}$  is the **velocity vector field**.

Frequently, it is assumed that the flow is **incompressible**, i.e., the **velocity field** satisfies

$$\nabla \cdot \mathbf{u} = 0.$$

In this case,  $\mathbf{u}$  is said to be **solenoidal**.

If this is so, the above equation can be rewritten as

$$\partial_t \psi + \mathbf{u} \cdot \nabla \psi = 0.$$

In particular, if the flow is steady, then

$$\mathbf{u} \cdot \nabla \psi = 0,$$

which shows that  $\psi$  is constant along a streamline.

Hence,  $\partial_t \psi = 0$ , so  $\psi$  does not vary in time.

If a vector quantity  $\mathbf{a}$  (e.g., a **magnetic field**) is being advected by the **solenoidal velocity field**  $\mathbf{u}$ , the advection equation above becomes:

$$\partial_t \mathbf{a} + (\mathbf{u} \cdot \nabla) \mathbf{a} = 0.$$

Here,  $\mathbf{a}$  is a vector field instead of the scalar field  $\psi$ .

**Solving the equation.** A simulation of the advection equation where  $\mathbf{u} = (\sin t, \cos t)$  is solenoidal.

The advection equation is not simple to solve **numerically**: the system is a **hyperbolic partial differential equation**, & interest typically centers on **discontinuous** “shock” solutions (which are notoriously difficult for numerical schemes to handle).

Even with 1 space dimension & a constant **velocity field**, the system remains difficult to simulate.

The equation becomes

$$\partial_t \psi + u_x \partial_x \psi = 0,$$

where  $\psi = \psi(t, x)$  is the scalar field being advected &  $u_x$  is the  $x$  component of the vector  $\mathbf{u} = (u_x, 0, 0)$ .

**Treatment of the advection operator in the compressible NSEs.** According to Zang, [Zang, Thomas (1991). “On the rotation & skew-symmetric forms for incompressible flow simulations”. Applied Numerical Mathematics. 7: 27–40. Bib-code:1991ApNM....7...27Z. doi:10.1016/0168-9274(91)90102-6.] numerical simulation can be aided by considering the **skew symmetric** form for the advection operator.

$$\frac{1}{2} \mathbf{u} \cdot \nabla \mathbf{u} + \frac{1}{2} \nabla(\mathbf{u}\mathbf{u}) \text{ where } \nabla(\mathbf{u}\mathbf{u}) = [\nabla(\mathbf{u}u_x), \nabla(\mathbf{u}u_y), \nabla(\mathbf{u}u_z)]$$

and  $\mathbf{u}$  is the same as above.

Since skew symmetry implies only imaginary eigenvalues, this form reduces the “blow up” & “spectral blocking” often experienced in numerical solutions with sharp discontinuities (see Boyd [Boyd, John P. (2000). *Chebyshev & Fourier Spectral Methods* 2nd edition. Dover. p. 213.]).

Using **vector calculus identities**, these operators can also be expressed in other ways, available in more software packages for more coordinate systems.

$$\begin{aligned} \mathbf{u} \cdot \nabla \mathbf{u} &= \nabla \left( \frac{\|\mathbf{u}\|^2}{2} \right) + (\nabla \times \mathbf{u}) \times \mathbf{u}, \\ \frac{1}{2} \mathbf{u} \cdot \nabla \mathbf{u} + \frac{1}{2} \nabla(\mathbf{u}\mathbf{u}) &= \nabla \left( \frac{\|\mathbf{u}\|^2}{2} \right) + (\nabla \times \mathbf{u}) \times \mathbf{u} + \frac{1}{2} \mathbf{u}(\nabla \cdot \mathbf{u}). \end{aligned}$$

This form also makes visible that the **skew symmetric** operator introduces error when the velocity field diverges.

Solving the advection equation by numerical methods is very challenging & there is a large scientific literature about this.”  
– **Wikipedia/advection**

## 1.2 Wikipedia/computational physics

“*Computational physics* is the study & implementation of numerical analysis to solve problems in physics. Historically, computational physics was the 1st application of modern computers in science, & is now a subset of **computational science**. It is sometimes regarded as a subdiscipline (or offshoot) of **theoretical physics**, but others consider it an intermediate branch between theoretical & **experimental physics** – an area of study which supplements both theory & experiment.

### 1.2.1 Overview

In physics, different **theories** based on mathematical models provide very precise predictions on how system behave. Unfortunately, it is often the case that solving the mathematical model for a particular system in order to produce a useful prediction is not feasible. This can occur, e.g., when the solution does not have a **closed-form expression**, or is too complicated. In such cases, numerical approximations are required. Computational physics is the subject that deals with these numerical approximations: the approximation of the solution is written as a finite (& typically large) number of simple mathematical operations (**algorithm**), & a computer is used to perform these operations & compute an approximated solution & respective **approximation error**.

**Status in physics.** A representation of the multidisciplinary nature of computational physics both as an overlap of physics, applied mathematics, & computer science & as a bridge among them. There is a debate about the status of computation within the scientific method. Sometimes it is regarded as more akin to theoretical physics; some others regard computer simulation as “**computer experiments**”, yet still others consider it an intermediate or different branch between theoretical & **experimental physics**, a 3rd way that supplements theory & experiment. While computers can be used in experiments for the measurement & recording (& storage) of data, this clearly does not constitute a computational approach.

### 1.2.2 Challenges in computational physics

Computational physics problems are in general very difficult to solve exactly. This is due to several (mathematical) reasons: lack of algebraic &/or analytic solvability, **complexity**, & chaos. E.g., even apparently simple problems, e.g., calculating the **wavefunction** of an electron orbiting an atom in a strong **electric field** (**Stark effect**, may require great effort to formulate a practical algorithm (if one can be found); other cruder or brute-force techniques, e.g. **graphical methods** or **root finding**, may be required. On the more advanced side, mathematical **perturbation theory** is also sometimes used (a working is shown for this particular **example**). In addition, the **computational cost** & **computational complexity** for **many-body problems** (& their **classical counterparts**) tend to grow quickly. A macroscopic system typically has a size of the order of  $10^{23}$  constituent particles, so it is somewhat of a problem. Solving quantum mechanical problems is generally of **exponential order** in the size of the system & for classical  $N$ -body it is of order  $N$ -squared. Finally, many physical systems are inherently nonlinear at best, & at worst **chaotic**: i.e., it can be difficult to ensure any **numerical errors** do not grow to the point of rendering the ‘solution’ useless.

### 1.2.3 Methods & algorithms

Because computational physics uses a broad class of problems, it is generally divided amongst the different mathematical problems it numerically solves, or the methods it applies. Between them, one can consider:

- **root finding** (using e.g. **Newton–Raphson method**)
- **system of linear equations** (using e.g. **LU decomposition**)
- **ODEs** (using e.g. **Runge–Kutta methods**)
- **integration** (using e.g. **Romberg method** & **Monte Carlo integration**)
- **PDEs** (using e.g. **FDM** & **relaxation method**)
- **matrix eigenvalue problem** (using e.g. **Jacobi eigenvalue algorithm** & **power iteration**)

All these methods (& several others) are used to calculate physical properties of the modeled systems.

Computational physics also borrows a number of ideas from **computational chemistry** – e.g., the **density functional theory** used by computational solid state physicists to calculate properties of solids is basically the same as that used by chemists to calculate the properties of molecules.

Furthermore, computational physics encompasses the **tuning** of the **software/hardware structure** to solve the problems (as the problems usually can be very large, in **processing power need** or in **memory requests**).

### 1.2.4 Divisions

It is possible to find a corresponding computational branch for every major field in physics:

1. **Computational mechanics** consists of CFDs, computational **solid mechanics** & computational **contact mechanics**.
2. **Computational electrodynamics** is the process of modeling the interaction of **electromagnetic fields** with physical objects & the environment. 1 subfield at the confluence between CFD & electromagnetic modeling is **computational magnetohydrodynamics**.
3. **Computational chemistry** is a rapidly growing field that was developed due to the **quantum many-body problem**.
4. Computational **solid state physics** is a very important division of computational physics dealing directly with **material science**.
5. Computational **statistical mechanics** is a field related to computational **condensed matter** which deals with the simulation of models & theories (e.g. **percolation** & **spin models**) that are difficult to solve otherwise.
6. Computational **statistical physics** make heavy use of Monte Carlo-like methods. More broadly, (particularly through the use of **agent based modeling** & **cellular automata**) it also concerns itself with (& finds application in, through the use of its techniques) in the **social sciences**, **network theory**, & mathematical models for the propagation of disease (most notably, the **SIR Model**) & the **spread of forest fires**.
7. **Numerical relativity** is a (relatively) new field interested in finding numerical solutions to the field equations of both **special relativity** & **general relativity**.
8. **Computational particle physics** deals with problems motivated by particle physics.
9. **Computational astrophysics** is the application of these techniques & methods to astrophysical problems & phenomena.
10. **Computational biophysics** is a branch of biophysics & **computational biology** itself, applying methods of computer science & physics to large complex biological problems.

### 1.2.5 Applications

Due to the broad class of problems computational physics deals, it is an essential component of modern research in different areas of physics, namely: **accelerator physics**, **astrophysics**, **general theory of relativity** (through **numerical relativity**), **fluid mechanics** (CFDs), **lattice field theory/lattice gauge theory** (especially **lattice quantum chromodynamics**), **plasma physics** (see **plasma modeling**), simulating physical systems (using e.g. **molecular dynamics**), **nuclear engineering computer codes**, **protein structure prediction**, **weather prediction**, **solid state physics**, **soft condensed matter physics**, hypervelocity impact physics etc.

Computational solid state physics, e.g., uses **density functional theory** to calculate properties of solids, a method similar to that used by chemists to study molecules. Other quantities of interest in solid state physics, e.g. electronic band structure, magnetic properties & charge densities can be calculated by this & several methods, including the **Luttinger-Kohn/k.p method** & **ab-initio** methods.

On top of advanced physics software, there are also a myriad of tools of analytics available for beginning students of physics e.g. PASCO Capstone software.” – [Wikipedia/computational physics](#)

## 1.3 Wikipedia/diffusion

Some particles are **dissolved** in a glass of water. At 1st, the particles are all near 1 top corner of the glass. If the particles randomly move around (“diffuse”) in the water, they eventually become distributed randomly & uniformly from an area of high concentration to an area of low, & organized (diffusion continues, but with no net **flux**).

“*Diffusion* is the net movement of anything (e.g., atoms, ions, molecules, energy) generally from a region of higher **concentration** to a region of lower concentration. Diffusion is driven by a gradient in **Gibbs free energy** or **chemical potential**. It is possible to diffuse “uphill” from a region of lower concentration to a region of higher concentration, as in **spinodal decomposition**. Diffusion is a stochastic process due to the inherent randomness of the diffusing entity & can be used to model many real-life stochastic scenarios. Therefore, diffusion & the corresponding mathematical models are used in several fields beyond physics, e.g. **statistics**, **probability theory**, **information theory**, **neural networks**, **finance**, & **marketing**.

The concept of diffusion is widely used in many fields, including physics (**particle diffusion**), chemistry, biology, sociology, economics, statistics, data science, & finance (diffusion of people, ideas, data, & price values). The central idea of diffusion, however, is common to all of these: a substance or collection undergoing diffusion spreads out from a point or location at which there is a higher concentration of that substance or collection.

A **gradient** is the change in the value of a quantity; e.g., concentration, **pressure**, or **temperature** with the change in another variable, usually **distance**. A change in concentration over a distance is called a **concentration gradient**, a change in pressure over a distance is called a **pressure gradient**, & a change in temperature over a distance is called a **temperature gradient**.

The word *diffusion* derives from the Latin word, *diffundere*, which means “to spread out”.

A distinguishing feature of diffusion is that it depends on particle **random walk**, & results in mixing or mass transport without requiring directed bulk motion. Bulk motion, or bulk flow, is the characteristic of **advection**. The term **convection** is used to describe the combination of both **transport phenomena**.

– 1 đặc điểm nổi bật của sự khuếch tán là nó phụ thuộc vào bước đi ngẫu nhiên của hạt và dẫn đến sự trộn lẫn hoặc vận chuyển khối lượng mà không cần chuyển động khối có hướng. Chuyển động khối, hay dòng chảy khối, là đặc điểm của sự tiến bộ. Thuật ngữ đối lưu được sử dụng để mô tả sự kết hợp của cả hai hiện tượng vận chuyển.

In a diffusion process can be described by **Fick's laws**, it is called a normal diffusion (or Fickian diffusion), otherwise, it is called an **anomalous diffusion** (or non-Fickian diffusion).

When talking about the extent of diffusion, 2 length scales are used in 2 different scenarios:

1. **Brownian motion** of an **impulsive** point source (e.g., 1 single spray of perfume) – the square root of the **mean squared displacement** from this point. In Fickian diffusion, this is  $\sqrt{2nDt}$ , where  $n$  is the dimension of this Brownian motion;
2. **Constant concentration source** in 1D – the diffusion length. In Fickian diffusion, this is  $2\sqrt{Dt}$ .

Diffusion from a microscopic & a macroscopic point of view. Initially, there are **solute** molecules on the left side of a barrier & none on the right. The barrier is removed, & the solute diffuses to fill the whole container. Top: A single molecule moves around randomly. Middle: With more molecules, there is a statistical trend that the solute fills the container more & more uniformly. Bottom: With an enormous number of solute molecules, all randomness is gone: The solute appears to move smoothly & deterministically from high-concentration areas to low-concentration areas. There is no microscopic **force** pushing molecules rightward, but there *appears* to be one in the bottom panel. This apparent force is called an **entropic force**.

### 1.3.1 Diffusion vs. bulk flow

“Bulk flow” is the movement/flow of an entire body due to a pressure gradient (e.g., water coming out of a tap). “Diffusion” is the gradual movement/dispersion of concentration within a body with no net movement of matter. An example of a process where both **bulk motion** & diffusion occur is human breathing.

1st, there is a “bulk flow” process. The **lungs** are located in the **thoracic cavity**, which expands as the 1st step in external respiration. This expansion leads to an increase in volume of the **alveoli** in the lungs, which causes a decrease in pressure in the alveoli. This creates a pressure gradient between the **air** outside the body at relatively high pressure & the alveoli<sup>1</sup> at relatively

---

<sup>1</sup>phế nang.



low pressure. The air moves down the pressure gradient through the airways of the lungs & into the alveoli until the pressure of the air & that in the alveoli are equal, i.e., the movement of air by bulk flow stops once there is no longer a pressure gradient.

2nd, there is a “diffusion” process. The air arriving in the alveoli has a higher concentration of oxygen than the “stale” air in the alveoli. The increase in oxygen concentration creates a concentration gradient for oxygen between the air in the alveoli & the blood in the **capillaries** that surround the alveoli. Oxygen then moves by diffusion, down the concentration gradient, into the blood. The other consequence of the air arriving in alveoli is that the concentration of **carbon dioxide** CO<sub>2</sub> in the alveoli decreases. This creates a concentration gradient for carbon dioxide to diffuse from the blood into the alveoli, as fresh air has a very low concentration of carbon dioxide compared to the **blood** in the body.

3rd, there is another “bulk flow” process. The pumping action of the **heart** then transports the blood around the body. As the left ventricle of the heart contracts, the volume decreases, which increases the pressure in the ventricle. This creates a pressure gradient between the heart & the capillaries, & blood moves through **blood vessels** by bulk flow down the pressure gradient.

### 1.3.2 Diffusion in the context of different disciplines

Diffusion furnaces used for **thermal oxidation**. “There are 2 ways to introduce the notion of *diffusion*: either a **phenomenological approach** starting with **Fick’s laws of diffusion** & their mathematical consequences, or a physical & atomistic one, by considering the **random walk** of the diffusion particles.

In the phenomenological approach, *diffusion is the movement of a substance from a region of high concentration to a region of low concentration without bulk motion*. According to Fick’s laws, the diffusion **flux** is proportional to the negative gradient  $-\nabla u$  of concentrations. It goes from regions of higher concentration to regions of lower concentration. Sometime later, various generalizations of Fick’s laws were developed in the frame of **thermodynamics** & **non-equilibrium thermodynamics**.

From the *atomistic point of view*, diffusion is considered as a result of the random walk of the diffusion particles. In **molecular diffusion**, the moving molecules in a gas, liquid, or solid are self-propelled by kinetic energy. Random walk of small particles in suspension in a fluid was discovered in 1827 by **ROBERT BROWN**, who found that minute particle suspended in a liquid medium & just large enough to be visible under an optical microscope exhibit a rapid & continually irregular motion of particles known as Brownian movement. The theory of the **Brownian motion** & the atomistic backgrounds of diffusion were developed by **ALBERT EINSTEIN**. The concept of diffusion is typically applied to any subject matter involving random walks in **ensembles** of individuals.

In chemistry & **materials science**, diffusion also refers to the movement of fluid molecules in porous solids. Different types of diffusion are distinguished in porous solids. **Molecular diffusion** occurs when the collision with another molecule is more likely than the collision with the pore walls. Under such conditions, the diffusivity is similar to that in a non-confined space & is proportional to the mean free path. **Knudsen diffusion** occurs when the pore diameter is comparable to or smaller than the mean free path of the molecule diffusing through the pore. Under this condition, the collision with the pore walls becomes gradually more likely & the diffusivity is lower. Finally there is configurational diffusion, which happens if the molecules have comparable size to that of the pore. Under this condition, the diffusivity is much lower compared to molecular diffusion & small differences in the kinetic diameter of the molecule cause large differences in .

**Biologists** often use the terms “net movement” or “net diffusion” to describe the movement of ions or molecules by diffusion. E.g., oxygen can diffuse through cell membranes so long as there is a higher concentration of oxygen outside the cell. However, because the movement of molecules is random, occasionally oxygen molecules move out of the cell (against the concentration gradient). Because there are more oxygen molecules outside the cell, the probability that oxygen molecules will enter the cell is higher than the probability that oxygen molecules will leave the cell. Therefore, the “net” movement of oxygen molecules (the difference between the number of molecules either entering or leaving the cell) is into the cell. I.e., there is a *net movement* of oxygen molecules down the concentration gradient.

### 1.3.3 History of diffusion in physics

In the scope of time, diffusion in solids was used long before the theory of diffusion was created. E.g., **PLINY THE ELDER** had previously described the **cementation process**, which produces steel from the element **iron** (Fe) through carbon diffusion. Another example is well known for many centuries, the diffusion of colors of **stained glass** or **earthenware** & **Chinese ceramics**.

In modern science, the 1st systematic experimental study of diffusion was performed by **THOMAS GRAHAM**. He studied diffusion in gases, & the main phenomenon was described by him in 1831–1833:

“... gases of different nature, when brought into contact, do not arrange themselves according to their density, the heaviest undermost, & the lighter uppermost, but they spontaneously diffuse, mutually & equally, through each other, & so remain in the intimate state of mixture for any length of time.”

The measurements of Graham contributed to **JAMES CLERK MAXWELL** deriving, in 1867, the coefficient of diffusion for CO<sub>2</sub> in the air. The error rate is < 5%.

In 1855, **ADOLF FICK**, the 26-year-old anatomy demonstrator from Zürich, proposed **Fick’s law of diffusion**. He used Graham’s research, stating his goal as “the development of a fundamental law, for the operation of diffusion in a single element of space”. He asserted a deep analogy between diffusion & conduction of heat or electricity, creating a formalism similar to **Fourier’s law for heat conduction** (1822) & **Ohm’s law** for electric current (1827).

**ROBERT BOYLE** demonstrated diffusion in solids in the 17th century by penetration of zinc Zn into a copper Cu coin. Nevertheless, diffusion in solids was not systematically studied until the 2nd part of the 19th century. **WILLIAM CHANDLER ROBERTS-AUSTEN**, the well-known British metallurgist & former assistant of THOMAS GRAHAM studied systematically solid state diffusion on the example of gold in lead in 1896:

“... My long connection with Graham’s researches made it almost a duty to attempt to extend his work on liquid diffusion to metals.”

In 1858, **Rudolf Clausius** introduced the concept of the **mean free path**. In the same year, **JAMES CLERK MAXWELL** developed the 1st atomistic theory of transport processes in gases. The modern atomistic theory of diffusion & Brownian motion was developed by **ALBERT EINSTEIN**, **Marian Smoluchowski**, & **Jean-Baptiste Perrin**. **LUDWIG BOLTZMANN**, in the development of the atomistic backgrounds of the macroscopic **transport processes**, introduced the **Boltzmann equation**, which has served mathematics & physics with a source of transport process ideas & concerns for > 140 years.

In 1920–1921, **GEORGE DE HEVESY** measured **self-diffusion** using **radioisotopes**. He studied self-diffusion of radioactive isotopes of lead in the liquid & solid lead.

**YAKOV FRENKEL** (sometimes, Jakov/Jacob Frenkel) proposed, & elaborated in 1926, the idea of diffusion in crystals through local defects (vacancies & **interstitial** atoms). He concluded, the diffusion process in condensed matter is an ensemble of elementary jumps & quasichemical interactions of particles & defects. He introduced several mechanisms of diffusion & found rate constants from experimental data.

Sometime later, **CARL WAGNER** & **WALTER H. SCHOTTKY** developed Frenkel’s ideas about mechanisms of diffusion further. Presently, it is universally recognized that atomic defects are necessary to mediate diffusion in crystals.

**HENRY EYRING**, with co-authors, applied his theory of **absolute reaction rates** to Frenkel’s quasichemical model of diffusion. The analogy between **reaction kinetics** & diffusion leads to various nonlinear versions of Fick’s law.

### 1.3.4 Basic models of diffusion

### 1.3.5 Diffusion in physics

### 1.3.6 Random walk (random motion)

” – [Wikipedia/diffusion](#)

## 1.4 Wikipedia/dynamics (mechanics)

“*Dynamics* is the **branch** of **classical mechanics** that is concerned with the study of **force** & their effects on **motion**. **Isaac Newton** was the 1st to formulate the fundamental **physical laws** that govern dynamics in classical non-relativistic physics, especially his **2nd law of motion**.” – [Wikipedia/dynamics \(mechanics\)](#)

### 1.4.1 Principles

“Generally speaking, researchers involved in dynamics study how a physical system might develop or alter over time & study the causes of those changes. In addition, Newton established the fundamental physical laws which govern dynamics in physics. By studying his system of mechanics, dynamics can be understood. In particular, dynamics is mostly related to Newton’s 2nd law of motion. However, all 3 laws of motion are taken into account because these are interrelated in any given observation or experiment.” – [Wikipedia/dynamics \(mechanics\)/principles](#)

### 1.4.2 Linear & rotational dynamics

“The study of dynamics falls under 2 categories: linear & rotational. Linear dynamics pertains to objects moving in a line & involves such quantities as **force**, **mass/inertia**, **displacement** (in units of distance), **velocity** (distance per unit time), **acceleration** (distance per unit of time squared) & **momentum** (mass times unit of velocity). Rotational dynamics pertains to objects that are rotating or moving in a curved path & involves such quantities as **torque**, **moment of inertia/rotational inertia**, **angular displacement** (in radians or less often, degrees), **angular velocity** (radians per unit time), **angular acceleration** (radians per unit of time squared) & **angular momentum** (moment of inertia times unit of angular velocity). Very often, objects exhibit linear & rotational motion.

For classical **electromagnetism**, **Maxwell’s equations** describe the kinematics. The dynamics of classical systems involving both mechanics & electromagnetism are described by the combination of Newton’s laws, Maxwell’s equations, & the **Lorentz force**.” – [Wikipedia/dynamics \(mechanics\)/linear & rotational dynamics](#)

### 1.4.3 Force

“Main article: [Wikipedia/force](#). From Newton, force can be defined as an exertion or **pressure** which can cause an object to **accelerate**. The concept of force is used to describe an influence which causes a **free body** (object) to accelerate. It can be a push or a pull, which causes an object to change direction, have new **velocity**, or to **deform** temporarily or permanently. Generally speaking, force causes an object’s **state of motion** to change.” – [Wikipedia/dynamics \(mechanics\)/force](#)

### 1.4.4 Newton’s laws

“Main article: [Wikipedia/Newton’s laws of motion](#). Newton described force as the ability to cause a mass to accelerate. His 3 laws can be summarized as follows:



1. *1st law*: if there is no net force on an object, then its velocity is constant: either the object is at rest (if its velocity is equal to zero), or it moves with constant speed in a single direction.
2. *2nd law*: The rate of change of linear momentum  $\mathbf{P}$  of an object is equal to the net force  $\mathbf{F}_{\text{net}}$ , i.e.,  $\frac{d\mathbf{P}}{dt} = \mathbf{F}_{\text{net}}$ .
3. *3rd law*: When a 1st body exerts a force  $\mathbf{F}_1$  on a 2nd body, the 2nd body simultaneously exerts a force  $\mathbf{F}_2 = -\mathbf{F}_1$  on the 1st body. I.e.,  $\mathbf{F}_1$  &  $\mathbf{F}_2$  are equal in magnitude & opposite in direction.

Newton's laws of motion are valid only in an **inertial frame of reference**." – Wikipedia/dynamics (mechanics)

## 1.5 Wikipedia/fluid mechanics

*Fluid mechanics* is the branch of **physics** concerned with the mechanics of **fluids** (**liquids**, **gases**, & **plasmas**) & the **forces** on them.[White, Frank M. (2011). Fluid Mechanics (7th ed.). McGraw-Hill. ISBN 978-0-07-352934-9.]:3

It has applications in a wide range of disciplines, including **mechanical**, **civil**, **chemical** & **biomedical engineering**, **geophysics**, **oceanography**, **meteorology**, **astrophysics**, & **biology**.

It can be divided into **fluid statics**, the study of fluids at rest; & **fluid dynamics**, the study of the effect of forces on fluid motion.[White, Frank M. (2011). Fluid Mechanics (7th ed.). McGraw-Hill. ISBN 978-0-07-352934-9.]:3

It is a branch of **continuum mechanics**, a subject which models matter without using the information that it is made out of atoms; i.e., it models matter from a *macroscopic* viewpoint rather than from *microscopic*.

Fluid mechanics, especially fluid dynamics, is an active field of research, typically mathematically complex.

Many problems are partly or wholly unsolved & are best addressed by **numerical methods**, typically using computers.

A modern discipline, called **computational fluid dynamics** (CFD), is devoted to this approach.[2]

**Particle image velocimetry**, an experimental method for visualizing & analyzing fluid flow, also takes advantage of the highly visual nature of fluid flow.

### 1.5.1 Brief history

Main article: **History of fluid mechanics**.

The study of fluid mechanics goes back at least to the days of **ancient Greece**, when **Archimedes** investigated fluid statics & **buoyancy** & formulated his famous law known now as the **Archimedes' principle**, which was published in his work **On Floating Bodies** - generally considered to be the 1st major work on fluid mechanics.

Rapid advancement in fluid mechanics began with **Leonardo da Vinci** (observations & experiments), **Evangelista Torricelli** (invented the **barometer**), **Isaac Newton** (investigated **viscosity**) & **Blaise Pascal** (researched **hydrostatics**, formulated **Pascal's law**), & was continued by **Daniel Bernoulli** with the introduction of mathematical fluid dynamics in *Hydrodynamica* (1739).

Inviscid flow was further analyzed by various mathematicians (**Jean le Rond d'Alembert**, **Joseph Louis Lagrange**, **Pierre-Simon Laplace**, **Siméon Denis Poisson**) & viscous flow was explored by a multitude of **engineers** including **Jean Léonard Marie Poiseuille** & **Gotthilf Hagen**.

Further mathematical justification was provided by **Claude-Louis Navier** & **George Gabriel Stokes** in the **Navier-Stokes equations**, & **boundary layers** were investigated (**Ludwig Prandtl**, **Theodore von Kármán**), while various scientists such as **Osborne Reynolds**, **Andrey Kolmogorov**, & **Geoffrey Ingram Taylor** advanced the understanding of fluid viscosity & **turbulence**.

### 1.5.2 Main branches

1. **Fluid statics**. Main article: **Fluid statics**.

**Fluid statics** or *hydrostatics* is the branch of fluid mechanics that studies **fluids** at rest.

It embraces the study of the conditions under which fluids are at rest in stable equilibrium; & is contrasted with fluid dynamics, the study of fluids in motion.

Hydrostatics offers physical explanations for many phenomena of everyday life, such as why atmospheric pressure changes with altitude, why wood & oil float on water, & why the surface of water is always level whatever the shape of its container.

Hydrostatics is fundamental to hydraulics, the engineering of equipment for storing, transporting & using fluids.

It is also relevant to some aspects of **geophysics** & **astrophysics** (e.g., in understanding **plate tectonics** & anomalies in the **Earth's gravitational field**), to **meteorology**, to **medicine** (in the context of **blood pressure**), & many other fields.

2. **Fluid dynamics**. Main article: **Fluid dynamics**.

**Fluid dynamics** is a subdiscipline of fluid mechanics that deals with *fluid flow* - the science of liquids & gases in motion.[3]

Fluid dynamics offers a systematic structure - which underlies these **practical disciplines** - that embraces empirical & semi-empirical laws derived from **flow measurement** & used to solve practical problems.

The solution to a fluid dynamics problem typically involves calculating various properties of the fluid, such as **velocity**, **pressure**, **density**, & **temperature**, as functions of space & time.

It has several subdisciplines itself, including **aerodynamics**[4][5][6][7] (the study of air & other gases in motion) & *hydrodynamics*[8][9] (the study of liquids in motion).

Fluid dynamics has a wide range of applications, including calculating **forces** & **movements** on **aircraft**, determining the **mass flow rate** of **petroleum** through pipelines, predicting evolving **weather** patterns, understanding **nebulae** in **interstellar space** & modeling **explosions**.

Some fluid-dynamical principles are used in **traffic engineering** & crowd dynamics.

### 1.5.3 Relationship to continuum mechanics

Fluid mechanics is a subdiscipline of **continuum mechanics**, as illustrated in the following table.

- **Continuum mechanics**. The study of the physics of continuous materials
  - **Solid mechanics**. The study of the physics of continuous materials with a defined rest shape.
    - \* **Elasticity**. Describes materials that return to their rest shape after applied **stresses** are removed.
    - \* **Plasticity**. Describes materials that permanently deform after a sufficient applied stress.
      - **Viscoplasticity**. The study of materials with both solid & fluid characteristics.
  - **Fluid mechanics**. The study of the physics of continuous materials which deform when subjected to a force.
    - \* **Non-Newtonian fluids** do not undergo strain rates proportional to the applied shear stress.
      - **Viscoelasticity**. The study of materials with both solid & fluid characteristics.
    - \* **Newtonian fluids** undergo strain rates proportional to the applied shear stress.

In a mechanical view, a fluid is a substance that does not support **shear stress**; that is why a fluid at rest has the shape of its containing vessel.

A fluid at rest has no shear stress.

### 1.5.4 Assumptions

Balance for some integrated fluid quantity in a **control volume** enclosed by a **control surface**.

The assumptions inherent to a fluid mechanical treatment of a physical system can be expressed in terms of mathematical equations.

Fundamentally, every fluid mechanical system is assumed to obey:

- **Conservation of mass**
- **Conservation of energy**
- **Conservation of momentum**
- The continuum assumption.

E.g., the assumption that mass is conserved means that for any fixed **control volume** (e.g., a spherical volume) - enclosed by a **control surface** - the **rate of change** of the mass contained in that volume is equal to the rate at which mass is passing through the surface from *outside* to *inside*, minus the rate at which mass is passing from *inside* to *outside*.

This can be expressed as an **equation in integral form** over the control volume.[Batchelor, George K. (1967). *An Introduction to Fluid Dynamics*. Cambridge University Press. p. 74. ISBN 0-521-66396-2.]:74

The *continuum assumption* is an idealization of **continuum mechanics** under which fluids can be treated as **continuous**, even though, on a microscopic scale, they are composed of **molecules**.

Under the continuum assumption, macroscopic (observed/measurable) properties such as density, pressure, temperature, & bulk velocity are taken to be well-defined at “infinitesimal” volume elements - small in comparison to the characteristic length scale of the system, but large in comparison to molecular length scale.

Fluid properties can vary continuously from 1 volume element to another & are average values of the molecular properties.

The continuum hypothesis can lead to inaccurate results in applications like supersonic speed flows, or molecular flows on nano scale.[11]

Those problems for which the continuum hypothesis fails can be solved using **statistical mechanics**.

To determine whether or not the continuum hypothesis applies, the **Knudsen number**, defined as the ratio of the molecular **mean free path** to the characteristic length **scale**, is evaluated.

Problems with Knudsen numbers below 0.1 can be evaluated using the continuum hypothesis, but molecular approach (statistical mechanics) can be applied to find the fluid motion for larger Knudsen numbers.

### 1.5.5 NSEs

Main article: **Navier-Stokes equations**.

The *Navier-Stokes equations* (named after **Claude-Louis Navier** & **George Gabriel Stokes**) are **differential equations** that describe the force balance at a given point within a fluid.

For an **incompressible fluid** with vector velocity field **u**, the Navier-Stokes equations are[12][13][14][15]

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\frac{1}{\rho} \nabla P + \nu \Delta \mathbf{u}.$$

These differential equations are the analogues for deformable materials to Newton's equations of motion for particles - the Navier-Stokes equations describe changes in **momentum** (**force**) in response to **pressure**  $P$  & viscosity, parameterized by the **kinematic viscosity**  $\nu$  here.

Occasionally, **body forces**, such as the gravitational force or Lorentz force are added to the equations.

Solutions of the Navier-Stokes equations for a given physical problem must be sought with the help of **calculus**.

In practical terms, only the simplest cases can be solved exactly in this way.

These cases generally involve non-turbulent, steady flow in which the **Reynolds number** is small.

For more complex cases, especially those involving **turbulence**, such as global weather systems, aerodynamics, hydrodynamics & many more, solutions of the Navier-Stokes equations can currently only be found with the help of computers.

This branch of science is called **computational fluid dynamics**. [16][17][18][19][20]

### 1.5.6 Inviscid & viscous fluids

An *inviscid fluid* has no **viscosity**,  $\nu = 0$ .

In practice, an inviscid flow is an **idealization**, one that facilitates mathematical treatment.

In fact, purely inviscid flows are only known to be realized in the case of **superfluidity**.

Otherwise, fluids are generally *viscous*, a property that is often most important within a **boundary layer** near a solid surface, [21] where the flow must match onto the **no-slip condition** at the solid.

In some cases, the mathematics of a fluid mechanical system can be treated by assuming that the fluid outside of boundary layers is inviscid, & then **matching** its solution onto that for a thin **laminar** boundary layer.

For fluid flow over a porous boundary, the fluid velocity can be discontinuous between the free fluid & the fluid in the porous media (this is related to the Beavers & Joseph condition).

Further, it is useful at low **subsonic** speeds to assume that gas is **incompressible** - i.e., the density of the gas does not change even though the speed & **static pressure** change.

### 1.5.7 Newtonian vs. non-Newtonian fluids

A *Newtonian fluid* (named after **Isaac Newton**) is defined to be a **fluid** whose **shear stress** is linearly proportional to the velocity gradient in the direction **perpendicular** to the plane of shear.

This definition means regardless of the forces acting on a fluid, it *continues to flow*.

E.g., water is a Newtonian fluid, because it continues to display fluid properties no matter how much it is stirred or mixed.

A slightly less rigorous definition is that the **drag** of a small object being moved slowly through the fluid is proportional to the force applied to the object.

(Compare **friction**).

Important fluids, like water as well as most gases, behave - to good approximation - as a Newtonian fluid under normal conditions on Earth. [Batchelor, George K. (1967). *An Introduction to Fluid Dynamics*. Cambridge University Press. p. 74. ISBN 0-521-66396-2.]:145

By contrast, stirring a **non-Newtonian fluid** can leave a "hole" behind.

This will gradually fill up over time - this behavior is seen in materials such as pudding, **oobleck**, or **sand** (although sand isn't strictly a fluid).

Alternatively, stirring a non-Newtonian fluid can cause the viscosity to decrease, so the fluid appears "thinner" (this is seen in non-drip **paints**).

There are many types of non-Newtonian fluids, as they are defined to be something that fails to obey a particular property - e.g., most fluids with long molecular chains can react in a non-Newtonian manner. [Batchelor, George K. (1967). *An Introduction to Fluid Dynamics*. Cambridge University Press. p. 74. ISBN 0-521-66396-2.]:145

**Equations for a Newtonian fluid.** Main article: **Newtonian fluid**.

The constant of proportionality between the viscous stress tensor & the velocity gradient is known as the **viscosity**.

A simple equation to describe incompressible Newtonian fluid behavior is

$$\tau = -\mu \frac{dv}{dy},$$

where

- $\tau$  is the shear stress exerted by the fluid ("**drag**")
- $\mu$  is the fluid viscosity - a constant of proportionality
- $\frac{dv}{dy}$  is the velocity gradient perpendicular to the direction of shear.

For a Newtonian fluid, the viscosity, by definition, depends only on **temperature** & **pressure**, not on the forces acting upon it.

If the fluid is **incompressible** the equation governing the viscous stress (in **Cartesian coordinates**) is

$$\tau_{ij} = \mu (\partial_{x_j} v_i + \partial_{x_i} v_j)$$

where

- $\tau_{ij}$  is the shear stress on the  $i$ th face of a fluid element in the  $j$ th direction
- $v_i$  is the velocity in the  $i$ th direction
- $x_j$  is the  $j$ th direction coordinate.

If the fluid is not incompressible the general form for the viscous stress in a Newtonian fluid is

$$\tau_{ij} = \mu \left( \partial_{x_j} v_i + \partial_{x_i} v_j - \frac{2}{3} \delta_{ij} \nabla \cdot \mathbf{v} \right) + \kappa \delta_{ij} \nabla \cdot \mathbf{v}$$

where  $\kappa$  is the 2nd viscosity coefficient (or bulk viscosity).

If a fluid does not obey this relation, it is termed a **non-Newtonian fluid**, of which there are several types.

Non-Newtonian fluids can be either plastic, Bingham plastic, pseudoplastic, dilatant, thixotropic, rheopectic, viscoelastic.

In some applications, another rough broad division among fluids is made: ideal & non-ideal fluids.

An *ideal fluid* is non-viscous & offers no resistance whatsoever to a shearing force.

An ideal fluid really does not exist, but in some calculations, the assumption is justifiable.

1 example of this is the flow far from solid surfaces.

In many cases, the viscous effects are concentrated near the solid boundaries (such as in boundary layers) while in regions of the flow field far away from the boundaries the viscous effects can be neglected & the fluid there is treated as it were inviscid (ideal flow).

When the viscosity is neglected, the term containing the viscous stress tensor  $\tau$  in the Navier-Stokes equation vanishes.

The equation reduced in this form is called the **Euler equation**.

## 1.6 Wikipedia/external flow

In **fluid mechanics**, *external flow* is such a flow that **boundary layers** develop freely, without constraints imposed by adjacent surfaces.

Accordingly, there will always exist a region of the flow outside the boundary layer in which velocity, temperature, and/or **concentration gradients** are negligible.

It can be defined as the flow of a fluid around a body that is completely submerged in it.

An example includes fluid motion over a flat plate (inclined or parallel to the free stream velocity) & flow over curved surfaces such as a sphere, cylinder, **airfoil**, or **turbine blade**, air flowing around an airplane & water flowing around the submarines.” – **Wikipedia/external flow**

## 1.7 Wikipedia/Gauss’s principle of least constraint

“The *principle of least constraint* is 1 **variational formulation** of **classical mechanics** enunciated by **Carl Friedrich Gauss** in 1829, equivalent to all other formulations of analytical mechanics. Intuitively, it says that the acceleration of a **constrained physical system** will be as similar as possible to that of the corresponding unconstrained system.” – **Wikipedia/Gauss’s principle of least constraint**

### 1.7.1 Statement

“The principle of least constraint is a **least squares** principle stating that the true accelerations of a mechanical system of  $n$  masses is the minimum of the quantity

$$Z := \sum_{j=1}^n m_j \left| \ddot{\mathbf{r}}_j - \frac{\mathbf{F}_j}{m_j} \right|^2,$$

where the  $j$ th particle has **mass**  $m_j$ , **position vector**  $\mathbf{r}_j$ , & applied non-constraint force  $\mathbf{F}_j$  acting on the mass.

The notation  $\dot{\mathbf{r}}$  indicates **time derivative** of a vector function  $\mathbf{r}(t)$ , i.e., position. The corresponding **accelerations**  $\ddot{\mathbf{r}}_j$  satisfy the imposed constraints, which in general depends on the current state of the system,  $\{\mathbf{r}_j(t), \dot{\mathbf{r}}_j(t)\}$ .

It is recalled the fact that due to active  $\mathbf{F}_j$  & reactive (constraint)  $\mathbf{F}_{c_j}$  forces being applied, with resultant  $\mathbf{R} = \sum_{j=1}^n \mathbf{F}_j + \mathbf{F}_{c_j}$ , a system will experience an acceleration  $\ddot{\mathbf{r}}_j = \sum_{j=1}^n \frac{\mathbf{F}_j}{m_j} + \frac{\mathbf{F}_{c_j}}{m_j} = \sum_{j=1}^n \mathbf{a}_j + \mathbf{a}_{c_j}$ .

**Connections to other formulations.** Gauss’s principle is equivalent to **D’Alembert’s principle**. The principle of least constraint is qualitatively similar to **Hamilton’s principle**, which states that the true path taken by a mechanical system is an extremum of the **action**. However, Gauss’s principle is a true (local) *minimal* principle, whereas the other is an *extremal* principle.” – **Wikipedia/Gauss’s principle of least constraint/statement**

### 1.7.2 Hertz's principle of least curvature

“Hertz's principle of least curvature is a special case of Gauss's principle, restricted by the 2 conditions that there are no externally applied forces, no interactions (which can usually be expressed as a **potential energy**), & all masses are equal. W.l.o.g., the masses may be set equal to 1. Under these conditions, Gauss's minimized quantity can be written  $Z = \sum_{j=1}^n |\ddot{\mathbf{r}}_j|^2$ . The **kinetic energy**  $T$  is also conserved under these conditions  $T := \frac{1}{2} \sum_{j=1}^n |\dot{\mathbf{r}}_j|^2$ . Since the **line element**  $ds^2$  in the  $3n$ -dimensional space of the coordinates is defined  $ds^2 := \sum_{j=1}^n |d\mathbf{r}_j|^2$ , the **conservation of energy** may also be written  $\left(\frac{ds}{dt}\right)^2 = 2T$ . Dividing  $Z$  by  $2T$  yields another minimal quantity  $K := \sum_{j=1}^n \left|\frac{d^2\mathbf{r}_j}{ds^2}\right|^2$ . Since  $\sqrt{K}$  is the local **curvature** of the trajectory in the  $3n$ -dimensional space of the coordinates, minimization of  $K$  is equivalent to finding the trajectory of least curvature (a **geodesic**) that is consistent with the constraints. Hertz's principle is also a special case of **Jacobi's** formulation of **the least-action principle**.” – [Wikipedia/Gauss's principle of least constraint](#)

## 1.8 Wikipedia/hydrostatics

Table of Hydraulics & Hydrostatics, from the 1728 Cyclopædia.

*Fluid statics* or *hydrostatics* is the branch of **fluid mechanics** that studies “**fluids** at rest & the pressure in a fluid or exerted by a fluid on an immersed body”.[“Hydrostatics”. Merriam-Webster. Retrieved Sep 11, 2018.]

It encompasses the study of the conditions under which fluids are at rest in **stable equilibrium** as opposed to **fluid dynamics**, the study of fluids in motion.

Hydrostatics are categorized as a part of the fluid statics, which is the study of all fluids, incompressible or not, at rest.

Hydrostatics is fundamental to **hydraulics**, the engineering of equipment for storing, transporting & using fluids.

It is also relevant to **geophysics** & **astrophysics** (e.g., in understanding **plate tectonics** & the anomalies of the **Earth's gravitational field**), to **meteorology**, to **medicine** (in the context of **blood pressure**), & many other fields.

Hydrostatics offers physical explanations for many phenomena of everyday life, such as why **atmospheric pressure** changes with **altitude**, why wood & oil float on water, & why the surface of still water is always level.

### 1.8.1 History

Some principles of hydrostatics have been known in an empirical & intuitive sense since antiquity, by the builders of boats, **cisterns**, **aqueducts** & **fountains**.

**Archimedes** is credited with the discovery of **Archimedes' Principle**, which relates the **buoyancy** force on an object that is submerged in a fluid to the weight of fluid displaced by the object.

The **Roman** engineer **Vitruvius** warned readers about **lead** pipes bursting under hydrostatic pressure.[Marcus Vitruvius Pollio (ca. 15 BCE), “The Ten Books of Architecture”, Book VIII, Chapter 6. At the University of Chicago's Penelope site. Accessed on 2013-02-25.]

The concept of pressure & the way it is transmitted by fluids was formulated by the **French mathematician** & **philosopher** **Blaise Pascal** in 1647.

### Hydrostatics in ancient Greece & Rome.

**Pythagorean Cup.** Main article: **Pythagorean cup**.

The “fair cup” or **Pythagorean cup**, which dates from about the 6th century BC, is a hydraulic technology whose invention is credited to the Greek mathematician & geometer Pythagoras.

It was used as a learning tool.

The cup consists of a line carved into the interior of the cup, & a small vertical pipe in the center of the cup that leads to the bottom.

The height of this pipe is the same as the line carved into the interior of the cup.

The cup may be filled to the line without any fluid passing into the pipe in the center of the cup.

However, when the amount of fluid exceeds this fill line, fluid will overflow into the pipe in the center of the cup.

Due to the drag that molecules exert on one another, the cup will be emptied.

**Heron's fountain.** Main article: **Heron's fountain**.

**Heron's fountain** is a device invented by **Heron of Alexandria** that consists of a jet of fluid being fed by a reservoir of fluid.

The fountain is constructed in such a way that the height of the jet exceeds the height of the fluid in the reservoir, apparently in violation of principles of hydrostatic pressure.

The device consisted of an opening & 2 containers arranged one above the other.

The intermediate pot, which was sealed, was filled with fluid, & several **cannula** (a small tube for transferring fluid between vessels) connecting the various vessels.

Trapped air inside the vessels induces a jet of water out of a nozzle, emptying all water from the intermediate reservoir.



**Pascal's contribution in hydrostatics.** Main article: **Pascal's Law**.

Pascal made contributions to developments in both hydrostatics & hydrodynamics.

**Pascal's Law** is a fundamental principle of fluid mechanics that states that any pressure applied to the surface of a fluid is transmitted uniformly throughout the fluid in all directions, in such a way that initial variations in pressure are not changed.

### 1.8.2 Pressure in fluids at rest

Due to the fundamental nature of fluids, a fluid cannot remain at rest under the presence of a **shear stress**.

However, fluids can exert **pressure normal** to any contacting surface.

If a point in the fluid is thought of as an infinitesimally small cube, then it follows from the principles of equilibrium that the pressure on every side of this unit of fluid must be equal.

If this were not the case, the fluid would move in the direction of the resulting force.

Thus, the pressure on a fluid at rest is **isotropic**; i.e., it acts with equal magnitude in all directions.

This characteristic allows fluids to transmit force through the length of pipes or tubes; i.e., a force applied to a fluid in a pipe is transmitted, via the fluid, to the other end of the pipe.

This principle was 1st formulated, in a slightly extended form, by Blaise Pascal, & is now called **Pascal's law**.

**Hydro pressure.** See also: **Vertical pressure variation**.

In a fluid at rest, all frictional & inertial stresses vanish & the state of stress of the system is called *hydrostatic*.

When this condition of  $V = 0$  is applied to the **Navier-Stokes equations**, the gradient of pressure becomes a function of body forces only.

For a **barotropic fluid** in a conservative force field like a gravitational force field, the pressure exerted by a fluid at equilibrium becomes a function of force exerted by gravity.

The hydrostatic pressure can be determined from a control volume analysis of an infinitesimally small cube of fluid.

Since pressure is defined as the force exerted on a test area ( $p = \frac{F}{A}$ , with  $p$ : pressure,  $F$ : force normal to area  $A$ ,  $A$ : area), & the only force acting on any such small cube of fluid is the weight of the fluid column above it, hydrostatic pressure can be calculated according to the following formula:

$$p(z) - p(z_0) = \frac{1}{A} \int_{z_0}^z dz' \iint_A dx' dy' \rho(z') g(z') = \int_{z_0}^z dz' \rho(z') g(z'),$$

where:

- $p$  is the hydrostatic pressure (Pa),
- $\rho$  is the fluid **density** (kg/m<sup>3</sup>),
- $g$  is **gravitational** acceleration (m/s<sup>2</sup>),
- $A$  is the test area (m<sup>2</sup>),
- $z$  is the height (parallel to the direction of gravity) of the test area (m),
- $z_0$  is the height of the **zero reference point of the pressure** (m).

For water & other liquids, this integral can be simplified significantly for many practical applications, based on the following 2 assumptions: Since many liquids can be considered **incompressible**, a reasonable good estimation can be made from assuming a constant density throughout the liquid.

(The same assumption cannot be made within a gaseous environment.)

Also, since the height  $h$  of the fluid column between  $z$  &  $z_0$  is often reasonably small compared to the radius of the Earth, one can neglect the variation of  $g$ .

Under these circumstances, the integral is simplified into the formula:

$$p - p_0 = \rho g h,$$

where  $h$  is the height  $z - z_0$  of the liquid column between the test volume & the zero reference point of the pressure.

This formula is often called **Stevin's law**.<sup>[3][4]</sup>

Note that this reference point should lie at or below the surface of the liquid.

Otherwise, one has to split the integral into 2 (or more) terms with the constant  $\rho_{\text{liquid}}$  &  $\rho(z')$  above.

E.g., the **absolute pressure** compared to vacuum is:

$$p = \rho g H + p_{\text{atm}},$$

where  $H$  is the total height of the liquid column above the test area to the surface, &  $p_{\text{atm}}$  is the **atmospheric pressure**, i.e., the pressure calculated from the remaining integral over the air column from the liquid surface to infinity.

This can easily be visualized using a **pressure prism**.

Hydrostatic pressure has been used in the preservation of foods in a process called **pascalization**.<sup>[5]</sup>



**Medicine.** In medicine, hydrostatic pressure in **blood vessels** is the pressure of the blood against the wall.

It is the opposing force to **oncotic pressure**.

**Atmospheric pressure.** **Statistical mechanics** shows that, for a pure **ideal gas** of constant temperature,  $T$ , its pressure,  $p$  will vary with height,  $h$ , as:

$$p(h) = p(0)e^{-\frac{Mgh}{kT}},$$

where:

- $g$  is the **acceleration due to gravity**
- $T$  is the **absolute temperature**
- $k$  is **Boltzmann constant**
- $M$  is the mass of a single **molecule** of gas
- $p$  is the pressure
- $h$  is the height

This is known as the **barometric formula**, & maybe derived from assuming the pressure is **hydrostatic**.

If there are multiple types of molecules in the gas, the **partial pressure** of each type will be given by this equation.

Under most conditions, the distribution of each species of gas is independent of the other species.

**Buoyancy.** Main article: **Buoyancy**.

Any body of arbitrary shape which is immersed, partly or fully, in a fluid will experience the action of a net force in the opposite direction of the local pressure gradient.

If this pressure gradient arises from gravity, the net force is in the vertical direction opposite that of the gravitational force.

This vertical force is termed buoyancy or buoyant force & is equal in magnitude, but opposite in direction, to the weight of the displaced fluid.

Mathematically,

$$F = \rho g V,$$

where  $\rho$  is the density of the fluid,  $g$  is the acceleration due to gravity, &  $V$  is the volume of fluid directly above the curved surface.[Fox, Robert; McDonald, Alan; Pritchard, Philip (2012). *Fluid Mechanics* (8 ed.). John Wiley & Sons. pp. 76–83. ISBN 978-1-118-02641-0.]

In the case of a **ship**, e.g., its weight is balanced by pressure forces from the surrounding water, allowing it to float.

If more cargo is loaded onto the ship, it would sink more into the water - displacing more water & thus receive a higher buoyant force to balance the increased weight.

Discovery of the principle of buoyancy is attributed to **Archimedes**.

**Hydrostatic force on submerged surfaces.** The horizontal & vertical components of the hydrostatic force acting on a submerged surface are given by the following:[6]

$$F_h = p_c A,$$

$$F_v = \rho g V,$$

where:

- $p_c$  is the pressure at the centroid of the vertical projection of the submerged surface
- $A$  is the area of the same vertical projection of the surface
- $\rho$  is the density of the fluid
- $g$  is the acceleration due to gravity
- $V$  is the volume of fluid directly above the curved surface

### 1.8.3 Liquids (fluids with free surfaces)

Liquids can have free surfaces at which they interface with gases, or with a **vacuum**.

In general, the lack of the ability to sustain a **shear stress** entails that free surfaces rapidly adjust towards an equilibrium.

However, on small length scales, there is an important balancing force from **surface tension**.

**Capillary action.** When liquids are constrained in vessels whose dimensions are small, compared to the relevant length scales, **surface tension** effects become important leading to the formation of a **meniscus** through **capillary action**.

This capillary action has profound consequences for biological systems as it is part of 1 of the 2 driving mechanisms of the flow of water in **plant xylem**, the **transpirational pull**.

**Hanging drops.** Without surface tension, **drops** would not be able to form.

The dimensions & stability of drops are determined by surface tension.

The drop's surface tension is directly proportional to the cohesion property of the fluid.” – [Wikipedia/Hydrostatics](#)

## 1.9 Wikipedia/mathematical physics

“*Mathematical physics* refers to the development of mathematical methods for application to problems in physics. The *Journal of Mathematical Physics* defines the field as “the application of mathematics to problems in physics & the development of mathematical methods suitable for such applications & for the formulation of physical theories”. An alternative definition would also include those mathematics that are inspired by physics, known as **physical mathematics**.

### 1.9.1 Scope

There are several distinct branches of mathematical physics, & these roughly correspond to particular historical parts of our world.

1. **Classical mechanics.** Main articles: [Wikipedia/Lagrangian mechanics](#), [Wikipedia/Hamiltonian mechanics](#). Applying the techniques of mathematical physics to classical mechanics typically involves the rigorous, abstract, & advanced reformulation of Newtonian mechanics in terms of Lagrangian mechanics & Hamiltonian mechanics (including both approaches in the presence of constraints). Both formulations are embodied in **analytical mechanics** & lead to an understanding of the deep interplay between the notions of **symmetry** & **conserved quantities** during the dynamical evolution of mechanical system, as embodied within the most elementary formulation of **Noether's theorem**. These approaches & ideas have been extended to other areas of physics, e.g. **statistical mechanics**, **continuum mechanics**, **classical field theory**, & **quantum field theory**. Moreover, they have provided multiple examples & ideas in **differential geometry** (e.g., several notions in **symplectic geometry** & **vector bundles**).
2. **PDEs.** Within mathematics proper, the theory of PDE, **variational calculus**, **Fourier analysis**, **potential theory**, & **vector analysis** are perhaps most closely associated with mathematical physics. These fields were developed intensively from the 2nd half of the 18th century (by, e.g., **D'ALEMBERT**, **EULER**, & **LAGRANGE**) until the 1930s. Physical applications of these developments include **hydrodynamics**, **celestial mechanics**, **continuum mechanics**, **elasticity theory**, **acoustics**, **thermodynamics**, **electricity**, **magnetism**, & **aerodynamics**. An example of mathematical physics: solutions of **Schrödinger's equation** for **quantum harmonic oscillators** with their **amplitude**.
3. **Quantum theory.** Main article: [Wikipedia/quantum mechanics](#). The theory of **atomic spectra** (& later, **quantum mechanics**) developed almost concurrently with some parts of the mathematical fields of linear algebra, the **spectral theory of operators**, **operator algebras** & more broadly, **functional analysis**. Nonrelativistic quantum mechanics includes **Schrödinger operators**, & it has connections to **atomic & molecular physics**. **Quantum information theory** is another subspecialty.
4. **Relativity & quantum relativistic theories.** [Wikipedia/theory of relativity](#), [Wikipedia/quantum field theory](#). The **special** & **general** theories of relativity require a rather different type of mathematics. This was **group theory**, which played an important role in both **quantum field theory** & differential geometry. This was, however, gradually supplemented by **topology** & functional analysis in the mathematical description of **cosmological** as well as **quantum field theory** phenomena. In the mathematical description of these physical areas, some concepts in **homological algebra** & **category theory** are also important.
5. **Statistical mechanics.** Main article: [Wikipedia/statistical mechanics](#). Statistical mechanics forms a separate field, which includes the theory of **phase transitions**. It relies upon the Hamiltonian mechanics (for its quantum versions) & it is closely related with the more mathematical **ergodic theory** & some parts of **probability theory**. There are increasing interactions between **combinatorics** & **physics**, in particular statistical physics.

### 1.9.2 Usage

Relationship between mathematics & physics: Math  $\rightarrow$  [Mathematical Physics, Ontology]  $\rightarrow$  Physics. The usage of the term “mathematical physics” is sometimes **idiosyncratic** (có phong cách riêng). Certain parts of mathematics that initially arose from the development of physics are not, in fact, considered parts of mathematical physics, while other closely related fields are. E.g., **ODEs** & **symplectic geometry** are generally viewed as purely mathematical disciplines, whereas **dynamical systems** & **Hamiltonian mechanics** belong to mathematical physics. **JOHN HERAPATH** used the term for the title of his 1847 text on “mathematical principles of natural philosophy”, the scope at that time being “the causes of heat, gaseous elasticity, gravitation, & other great phenomena of nature”.

1. **Mathematical vs. theoretical physics.** The term “mathematical physics” is sometimes used to denote research aimed at studying & solving problems in physics or **thought experiments** within a mathematically **rigorous** framework. In this sense,

mathematical physics covers a very broad academic realm distinguished only by the blending of some mathematical aspect & theoretical physics aspect. Although related to [theoretical physics](#), mathematical physics in this sense emphasizes the mathematical rigor of the similar type as found in mathematics.

On the other hand, theoretical physics emphasizes the links to observations & [experimental physics](#), which often requires theoretical physicists (& mathematical physicists in the more general sense) to use [heuristic](#), [intuitive](#), or approximate arguments. Such arguments are not considered rigorous by mathematicians.

Such mathematical physicists primarily expand & elucidate (làm sáng tỏ) physics [theories](#). Because of the required level of mathematical rigor, these researches often deal with questions that theoretical physicists have considered to be already solved. However, they can sometimes show that the previous solution was incomplete, incorrect, or simply too naive. E.g., issues about attempts to infer the 2nd law of [thermodynamics](#) from [statistical mechanics](#). Other examples concern the subtleties involved with synchronization procedures in special & general relativity ([Sagnac effect](#) & [Einstein synchronization](#)).

The effort to put physical theories on a mathematically rigorous footing not only developed physics but also has influenced developments of some mathematical areas. E.g., the development of quantum mechanics & some aspects of [functional analysis](#) parallel each other in many ways. The mathematical study of [quantum mechanics](#), [quantum field theory](#), & [quantum statistical mechanics](#) has motivated results in [operator algebras](#). The attempt to construct a rigorous mathematical formulation of quantum field theory has also brought about some progress in fields such as [representation theory](#).

### 1.9.3 Prominent mathematical physicists

1. Before Newton.
2. Newtonian physics & post Newtonian.
3. Relativistic.
4. Quantum.
5. List of prominent contributors to mathematical physics in the 20th century.

” – [Wikipedia/mathematical physics](#)

## 1.10 Wikipedia/matrix mechanics

“*Matrix mechanics* is a formulation of [quantum mechanics](#) created by [WERNER HEISENBERG](#), [MAX BORN](#), & [PASCUAL JORDAN](#) in 1925. It was the 1st conceptually autonomous & logically consistent formulation of quantum mechanics. Its account of [quantum jumps](#) supplanted the [Bohr model](#)’s [electron orbits](#). It did so by interpreting the physical properties of particles as [matrices](#) that evolve in time. It is equivalent to the [Schrödinger wave formulation](#) of quantum mechanics, as manifest in [DIRAC](#)’s [bra-ket notation](#).

In some contrast to the wave formulation, it produces spectra of (mostly energy) operators by purely algebraic, [ladder operator](#) methods. Relying on these methods, [WOLFGANG PAULI](#) derived the hydrogen atom spectrum in 1926, before the development of wave mechanics.

### 1.10.1 Development of matrix mechanics

In 1925, [WERNER HEISENBERG](#), [MAX BORN](#), & [PASCUAL JORDAN](#) formulated the matrix mechanics representation of quantum mechanics.

1. Epiphany at Helgoland. Main article: [Wikipedia/umdeutung paper](#).
2. 3 fundamental papers.
3. Heisenberg’s reasoning.
4. Matrix basics.
5. Nobel Prize.

### 1.10.2 Mathematical development

Once HEISENBERG introduced the matrices for  $X, P$ , he could find their matrix elements in special cases by guesswork, guided by the correspondence principle. Since the matrix elements are the quantum mechanical analogs of Fourier coefficients of the classical orbits, the simplest case is the [harmonic oscillator](#), where the classical position & momentum,  $X(t)$  &  $P(t)$ , are sinusoidal.

1. Harmonic oscillator.
2. Conservation of energy. Main article: [Wikipedia/conservation of energy](#).
3. Differentiation trick – canonical commutation relations.
4. State vectors & Heisenberg equation.

### 1.10.3 Further results

Matrix mechanics rapidly developed into modern quantum mechanics, & gave interesting physical results on the spectra of atoms.

1. Wave mechanics.
2. Ehrenfest theorem.
3. Transformation theory.
4. Selection rules.
5. Sum rules.

” – [Wikipedia/matrix mechanics](#)

## 1.11 Wikipedia/numerical methods in fluid mechanics

“Fluid motion is governed by the [Navier-Stokes equations](#), a set of coupled & nonlinear partial differential equations derived from the basic laws of conservation of [mass](#), [momentum](#) & [energy](#).

The unknowns are usually the [flow velocity](#), the [pressure](#) & [density](#) & [temperature](#).

The [analytical solution](#) of this equation is impossible hence scientists resort to laboratory experiments in such situations. The answers delivered are, however, usually qualitatively different since dynamical & geometric similitude are difficult to enforce simultaneously between the lab experiment & the [prototype](#).

Furthermore, the design & construction of these experiments can be difficult (and costly), particularly for stratified rotating flows.

[Computational fluid dynamics](#) (CFD) is an additional tool in the arsenal of scientists.

In its early days CFD was often controversial, as it involved additional approximation to the governing equations & raised additional (legitimate) issues.

Nowadays CFD is an established discipline alongside theoretical & experimental methods.

This position is in large part due to the exponential growth of computer power which has allowed us to tackle ever larger & more complex problems.

### 1.11.1 Discretization

The central process in CFD is the process of [discretization](#), i.e. the process of taking differential equations with an infinite number of [degrees of freedom](#), & reducing it to a system of finite degrees of freedom.

Hence, instead of determining the solution everywhere & for all times, we will be satisfied with its calculation at a finite number of locations & at specified time intervals.

The [PDEs](#) are then reduced to a system of algebraic equations that can be solved on a computer.

Errors creep in during the discretization process.

The nature & characteristics of the errors must be controlled in order to ensure that:

- we are solving the correct equations (consistency property)
- that the error can be decreased as we increase the number of degrees of freedom (stability & convergence).

Once these 2 criteria are established, the power of computing machines can be leveraged to solve the problem in a numerically reliable fashion.

Various discretization schemes have been developed to cope with a variety of issues.

The most notable for our purposes are: [finite difference methods](#), finite volume methods, [finite element methods](#), & [spectral methods](#).

### 1.11.2 FDM

Finite difference replace the infinitesimal limiting process of derivative calculation:

$$\lim_{\Delta x \rightarrow 0} f'(x) = \frac{f(x + \Delta x) - f(x)}{\Delta x},$$

with a finite limiting process, i.e.,

$$f'(x) = \frac{f(x + \Delta x) - f(x)}{\Delta x} + \mathcal{O}(\Delta x).$$

The term  $\mathcal{O}(\Delta x)$  gives an indication of the magnitude of the error as a function of the mesh spacing.

In this instance, the error is halved if the grid spacing,  $\Delta x$  is halved, & we say that this is a 1st order method.

Most FDM used in practice are at least 2nd order accurate except in very special circumstances.

Finite Difference method is still the most popular numerical method for solution of PDEs because of their simplicity, efficiency & low computational cost.

Their major drawback is in their geometric inflexibility which complicates their applications to general complex domains.

These can be alleviated by the use of either mapping techniques and/or masking to fit the computational mesh to the computational domain.

### 1.11.3 FEM

The finite element method was designed to deal with problem with complicated computational regions.

The PDE is 1st recast into a variational form which essentially forces the mean error to be small everywhere.

The discretization step proceeds by dividing the computational domain into elements of triangular or rectangular shape.

The solution within each element is interpolated with a polynomial of usually low order.

Again, the unknowns are the solution at the collocation points.

The CFD community adopted the FEM in the 1980s when reliable methods for dealing with advection dominated problems were devised.

### 1.11.4 Spectral method

Both finite element & finite difference methods are low order methods, usually of 2nd - 4th order, & have local approximation property.

By local we mean that a particular collocation point is affected by a limited number of points around it.

In contrast, spectral method have global approximation property.

The interpolation functions, either polynomials or trigonometric functions are global in nature.

Their main benefits is in the rate of convergence which depends on the smoothness of the solution (i.e. how many continuous derivatives does it admit).

For infinitely smooth solution, the error decreases exponentially, i.e. faster than algebraic.

Spectral methods are mostly used in the computations of homogeneous turbulence, & require relatively simple geometries.

Atmospheric model have also adopted spectral methods because of their convergence properties & the regular spherical shape of their computational domain.

### 1.11.5 FVM

Finite volume methods are primarily used in [aerodynamics](#) applications where strong shocks & discontinuities in the solution occur.

Finite volume method solves an integral form of the governing equations so that local continuity property do not have to hold.

### 1.11.6 Computational cost

The [CPU](#) time to solve the system of equations differs substantially from method to method.

Finite differences are usually the cheapest on a per grid point basis followed by the finite element method & spectral method.

However, a per grid point basis comparison is a little like comparing apple & oranges.

Spectral methods deliver more accuracy on a per grid point basis than either FEM or FDM.

The comparison is more meaningful if the question is recast as “*what is the computational cost to achieve a given error tolerance?*”.

The problem becomes 1 of defining the error measure which is a complicated task in general situations.

### 1.11.7 Forward Euler approximation

$$\frac{u^{n+1} - u^n}{\Delta t} \approx \kappa u^n.$$

Equation is an explicit approximation to the original differential equation since no information about the unknown function at the future time  $(n + 1)t$  has been used on the RHS of the equation.

In order to derive the error committed in the approximation we rely again on Taylor series.

### 1.11.8 Backward difference

This is an example of an implicit method since the unknown  $u(n + 1)$  has been used in evaluating the slope of the solution on the RHS; this is not a problem to solve for  $u(n + 1)$  in this scalar & linear case.

For more complicated situations like a nonlinear RHS or a system of equations, a nonlinear system of equations may have to be inverted.” – [Wikipedia/numerical methods in fluid mechanics](#)

## 1.12 Wikipedia/physics

“

” – [Wikipedia/physics](#)

## 1.13 Wikipedia/stationary-action principle

“The *stationary-action principle* – also known as the *principle of least action* – is a **variational principle** that, when applied to the **action** of a **mechanical** system, yields the **equations of motion** for that system. The principle states that the trajectories (i.e., the solutions of the equations of motions) are **stationary points** of the system’s **action functional**. The term “least action” is a historical misnomer since the principle has no minimality requirement: the value of the action functional need not be minimal (even locally) on the trajectories. Least action refers to the absolute value of the action functional being minimized.

The principle can be used to derive **Newtonian**, **Lagrangian**, & **Hamiltonian equations of motion**, & even **general relativity** (see **Einstein–Hilbert action**). In relativity, a different action must be minimized or maximized.

The classical mechanics & electromagnetic expressions are a consequence of quantum mechanics. The stationary action method helped in the development of quantum mechanics. In 1933, the physicist **Paul Dirac** demonstrated how this principle can be used in quantum calculations by discerning the **quantum mechanical underpinning** of the principle in the **quantum interference** of amplitudes. Subsequently **Julian Schwinger** & **Richard Feynman** independently applied this principle in quantum electrodynamics.

The principle remains central in **modern physics** & mathematics, being applied in **thermodynamics**, **fluid mechanics**, the **theory of relativity**, **quantum mechanics**, **particle physics**, & **string theory** & is a focus of modern mathematical investigation in **Morse theory**. **Maupertuis’ principle** & **Hamilton’s principle** exemplify the principle of stationary action.

The action principle is preceded by earlier ideas in **optics**. In **ancient Greece**, **Euclid** wrote in his *Catoptrica* that, for the path of light reflecting from a mirror, the **angle of incidence** equals the **angle of reflection**. **Hero of Alexandria** later showed that this path was the shortest length & least time.

Scholars often credit **for** formulating the principle of least action because he wrote about it in 1744 & 1746. However, **Leonhard Euler** discussed the principle in 1744, & evidence shows that **Gottfried Leibniz** preceded both by 39 years.” – **Wikipedia/stationary-action principle**

### 1.13.1 General statement

Fig. As the system evolves,  $\mathbf{q}$  traces a path through **configuration space** (only some are shown). The path taken by the system (red) has a stationary action ( $\delta S = 0$ ) under small changes in the configuration of the system ( $\delta \mathbf{q}$ ).

“The **action**, denoted  $S$ , of a physical system is defined as the **integral** of the **Lagrangian**  $L$  between 2 instants of **time**  $t_1$  &  $t_2$  – technically a **functional** of the  $N$  **generalized coordinates**  $\mathbf{q} = (q_1, \dots, q_n)$  which are functions of time & define the **configuration** of the system:

$$\mathbf{q} : \mathbb{R} \rightarrow \mathbb{R}^N,$$
$$S[\mathbf{q}, t_1, t_2] = \int_{t_1}^{t_2} L(\mathbf{q}(t), \dot{\mathbf{q}}(t), t) dt,$$

where the dot denotes the **time derivative**, &  $t$  is time. Mathematically the principle is  $\delta S = 0$ , where  $\delta$  (lowercase Greek **delta**) means a *small* change. In words this reads:

*The path taken by the system between times  $t_1$  &  $t_2$  & configurations  $q_1$  &  $q_2$  is the one for which the **action** is **stationary (no change)** to 1st order.*

Stationary action is not always a minimum, despite the historical name of least action. It is a minimum principle for sufficiently short, finite segments in the path.

In applications the statement & definition of action are taken together:  $\delta \int_{t_1}^{t_2} L(\mathbf{q}, \dot{\mathbf{q}}, t) dt = 0$ . The action & Lagrangian both contain the dynamics of the system for all times. The term “path” is simply refers to a curve traced out by the system in terms of the coordinates in the **configuration space**, i.e., the curve  $\mathbf{q}(t)$ , parametrized by time (see also **parametric equation** for this concept).” – **Wikipedia/stationary-action principle/general statement**

### 1.13.2 Origins, statements, & controversy

1. FERMAT. “Main article: **Wikipedia/Fermat’s principle**. In the 1600s, **Pierre de Fermat** postulated that “*light travels between 2 given points along the path of shortest time*,” which is known as the *principle of least time* or **Fermat’s principle**.
2. MAUPERTUIS. Main article: **Maupertuis principle**. Credit for the formulation of the *principle of least action* is commonly given to **Pierre Louis Maupertuis**, who felt that “Nature is thrifty in all its actions”, & applied the principle broadly:

“The laws of movement & of rest deduced from this principle being precisely the same as those observed in nature, we can admire the application of it to all phenomena. The movement of animals, the vegetative growth of plants ... are only its consequences; & the spectacle of the universe becomes so much the grander, so much more beautiful, the worthier of its Author, when one knows that a small number of laws, most wisely established, suffice for all movements.” – Pierre Louis Maupertuis

This notion of Maupertuis, although somewhat deterministic today, does capture much of the essence of mechanics.

In application to physics, Maupertuis suggested that the quantity to be minimized was the product of the duration (time) of movement within a system by the “**vis viva**”, **Maupertuis’ principle**:  $\delta \int 2T(t) dt = 0$ , which is the integral of twice what we now call the **kinetic energy**  $T$  of the system.



3. EULER. **Leonhard Euler** gave a formulation of the action principle in 1744, in very recognizable terms, in the *Additamentum 2* to his *Methodus Inveniendi Lineas Curvas Maximi Minive Proprietate Gaudentes*. Beginning with the 2nd paragraph:

Let the mass of the projectile be  $M$ , & let its speed be  $v$  while being moved over an infinitesimal distance  $ds$ . The body will have a momentum  $Mv$  that, when multiplied by the distance  $ds$ , will give  $Mvds$ , the momentum of the body integrated over the distance  $ds$ . Now I assert that the curve thus described by the body to be the curve (from among all other curves connecting the same endpoints) that minimizes  $\int Mv ds$  or, provided that  $M$  is constant along the path,  $M \int v ds$ .” – Leonhard Euler

As Euler states,  $\int Mv ds$  is the integral of the momentum over distance traveled, which, in modern notation, equals the abbreviated or **reduced action** **Euler’s principle**  $\delta \int p dq = 0$ . Thus, Euler made an equivalent & (apparently) independent statement of the variational principle in the same year as Maupertuis, albeit slightly later. Curiously, Euler did not claim any priority, as the following episode shows.

4. **Disputed priority.** Maupertuis’ priority was disputed in 1751 by the mathematician **Samuel König**, who claimed that it had been invented by **Gottfried Leibniz** in 1707. Although similar to many of Leibniz’s arguments, the principle itself has not been documented in Leibniz’s works. König himself showed a *copy* of a 1707 letter from Leibniz to **Jacob Hermann** with the principle, but the *original* letter has been lost. In contentious proceedings, König was accused of forgery<sup>2</sup>, & even the **King of Prussia** entered the debate, defending Maupertuis (the head of his Academy), while **Voltaire** defended König.

Euler, rather than claiming priority, was a staunch defender of Maupertuis, & Euler himself prosecuted König for forgery before the Berlin Academy on Apr 13, 1752. The claims of forgery were re-examined 150 years later, & archival work by C.I. Gerhardt in 1898 & W. Kabitz in 1913 uncovered other copies of the letter, & 3 others cited by König, in the **Bernoulli archives**.” – [Wikipedia/stationary-action principle/origins, statements, & controversy](#)

### 1.13.3 Further development

“Euler continued to write on the topic; in his *Réflexions sur quelques loix générales de la nature* (1748), he called action “effort”. His expression corresponds to modern **potential energy**, & his statement of least action says that the total potential energy of a system of bodies at rest is minimized, a principle of modern statics.

1. LAGRANGE & HAMILTON. Main article: [Wikipedia/Hamilton’s principle](#). Much of the calculus of variations was stated by **Joseph-Louis Lagrange** in 1760 & he proceeded to apply this to problems in dynamics. In *Mécanique analytique* (1788) Lagrange derived the general **equations of motion** of a mechanical body. **William Rowan Hamilton** in 1834 & 1835 applied the variational principle to the classical **Lagrangian function**  $L = T - V$  to obtain the **Euler–Lagrange equations** in their present form.
2. JACOBI, MORSE, & CARATHEODORY. In 1842, **Carl Gustav Jacobi** tackled the problem of whether the variational principle always found minima as opposed to other **stationary points** (maxima or stationary **saddle points**); most of his work focused on **geodesics** on 2D surfaces. The 1st clear general statements were given by **Marston Morse** in the 1920s & 1930s, leading to what is now known as **Morse theory**. E.g., Morse showed that the number of **conjugate points** in a trajectory equaled the number of negative eigenvalues in the 2nd variation of the Lagrangian. A particularly elegant derivation of the Euler–Lagrange equation was formulated by **Constantin Caratheodory** & published by him in 1935.
3. GAUSS & HERTZ. Other extremal principles of **classical mechanics** have been formulated, e.g. **Gauss’s principle of least constraint** & its corollary, **Hertz’s principle of least curvature**.” – [Wikipedia/stationary-action principle/further development](#)

### 1.13.4 Disputes about possible teleological aspects

“The mathematical equivalence of the **differential equations of motion** & their **integral** counterpart has important philosophical implications. The differential equations are statements about quantities localized to a single point in space or single moment of time. E.g., **Newton’s 2nd law**  $\mathbf{F} = m\mathbf{a}$  states that the *instantaneous* force  $\mathbf{F}$  applied to a mass  $m$  produces an acceleration  $\mathbf{a}$  at the same *instant*. By contrast, the action principle is not localized to a point; rather, it involves integrals over an interval of time & (for fields) an extended region of space. Moreover, in the usual formulation of **classical** action principles, the initial & final states of the system are fixed, e.g.,

*Given that the particle begins at position  $x_1$  at time  $t_1$  & ends at position  $x_2$  at time  $t_2$ , the physical trajectory that connects these 2 endpoints is an **extremum** of the action integral.*

In particular, the fixing of the *final* state has been interpreted as giving the action principle a **teleological character** which has been controversial historically. However, according to W. Yourgrau & S. Mandelstam, *the teleological approach ... presupposes that the variational principles themselves have mathematical characteristics which they de facto do not possess*. In addition, some critics maintain this apparent **teleology** occurs because of the way in which the question was asked. By specifying some but not all aspects of both the initial & final conditions (the positions but not the velocities) we are making some inferences about the initial conditions from the final conditions, & it is this “backward” inference that can be seen as a teleological explanation.

<sup>2</sup>**forgery** [n] (plural **forgeries**) **1.** [uncountable] the crime of copying money, documents, etc. in order to cheat people; **2.** [countable] something, e.g. a document, piece of paper money, etc., that has been copied in order to cheat people, SYNONYM: **fake**.

Teleology can also be overcome if we consider the classical description as a limiting case of the **quantum** formalism of **path integration**, in which stationary paths are obtained as a result of interference of amplitudes along all possible paths.

The short story **Story of Your Life** by the speculative fiction writer **Ted Chiang** contains visual depictions of **Fermat's Principle** along with a discussion of its teleological dimension. **Keith Devlin's** *The Math Instinct* contains a chapter, "Elvis the Welsh Corgi Who Can Do Calculus" that discusses the calculus "embedded" in some animals as they solve the "least time" problem in actual situations." – **Wikipedia/stationary-action principle**

## 1.14 Wikipedia/statistical mechanics

"In physics, *statistical mechanics* is a mathematical framework that applies **statistical methods** & **probability theory** to large assemblies of microscopic entities. Sometimes called *statistical physics* or *statistical thermodynamics*, its applications include many problems in the fields of physics, biology, chemistry, neuroscience, computer science, information theory, & sociology. Its main purpose is to clarify the properties of matter in aggregate, in terms of physical laws governing atomic motion.

Statistical mechanics arose out of the development of **classical thermodynamics**, a field for which it was successful in explaining macroscopic physical properties – e.g. **temperature**, **pressure**, & **heat capacity** – in terms of microscopic parameters that fluctuate about average values & are characterized by **probability distributions**.

While classical thermodynamics is primarily concerned with **thermodynamic equilibrium**, statistical mechanics has been applied in **non-equilibrium statistical mechanics** to the issues of microscopically modeling the speed of **irreversible processes** that are driven by imbalances. Examples of such processes include **chemical reactions** & flows of particles & heat. The **fluctuation-dissipation theorem** is the basic knowledge obtained from applying non-equilibrium statistical mechanics to study the simplest non-equilibrium situation of a steady state current flow in a system of many particles.

### 1.14.1 History

In 1738, Swiss physicist & mathematician **DANIEL BERNOULLI** published *Hydrodynamica* which laid the basis for the **kinetic theory of gases**. In this work, BERNOULLI posited the argument, still used to this day, that gases consist of great numbers of molecules moving in all directions, that their impact on a surface causes the gas pressure that we feel, & that what we experience as **heat** is simply the kinetic energy of their motion.

The founding of the field of statistical mechanics is generally credited to 3 physicists:

- **LUDWIG BOLTZMANN**, who developed the fundamental interpretation of **entropy** in terms of a collection of microstates
- **JAMES CLERK MAXWELL**, who developed models of probability distribution of such states
- **JOSIAH WILLARD GIBBS**, who coined the name of the field in 1884

In 1859, after reading a paper on the diffusion of molecules by **RUDOLF CLAUSIUS**, Scottish physicist **JAMES CLERK MAXWELL** formulated the **Maxwell distribution** of molecular velocities, which gave the proportion of molecules having a certain velocity in a specific range. This was the 1st-ever statistical law in physics. MAXWELL also gave the 1st mechanical argument that molecular collisions entail an equalization of temperatures & hence a tendency towards equilibrium. 5 years later, in 1864, **LUDWIG BOLTZMANN**, a young student in Vienna, came across MAXWELL's paper & spent much of his life developing the subject further.

Statistical mechanics was initiated in the 1870s with the work of BOLTZMANN, much of which was collectively published in his 1896 *Lectures on Gas Theory*. BOLTZMANN's original papers on the statistical interpretation of thermodynamics, the **H-theorem**, **transport theory**, **thermal equilibrium**, the **equation of state** gases, & similar subjects, occupy  $\approx 2000$  pages in the proceedings of the Vienna Academy & other societies. BOLTZMANN introduced the concept of an equilibrium statistical ensemble & also investigated for the 1st time non-equilibrium statistical mechanics, with his H-theorem.

The term "statistical mechanics" was coined by the American mathematical physicist **J. WILLARD GIBBS** in 1884. According to GIBBS, the term "statistical", in the context of mechanics, i.e., statistical mechanics, was 1st used by the Scottish physicist **JAMES CLERK MAXWELL** in 1871:

"In dealing with masses of matter, while we do not perceive the individual molecules, we are compelled to adopt what I have described as the statistical method of calculation, & to abandon the strict dynamical method, in which we follow every motion by the calculus." – J. CLERK MAXWELL

"Probabilistic mechanics" might today seem a more appropriate term, but "statistical mechanics" is firmly entrenched. Shortly before his death, GIBBS published in 1902 *Elementary Principles in Statistical Mechanics*, a book which formalized statistical mechanics as a fully general approach to address all mechanical systems – macroscopic or microscopic, gaseous or non-gaseous. GIBBS's methods were initially derived in the framework **classical mechanics**, however they were of such generality that they were found to adapt easily to the later **quantum mechanics**, & still form the foundation of statistical mechanics to this day.

### 1.14.2 Principles: mechanics & ensembles

Main articles: **Wikipedia/mechanics**, **Wikipedia/statistical ensemble**. In physics, 2 types of mechanics are usually examined: **classical mechanics** & **quantum mechanics**. For both types of mechanics, the standard mathematical approach is to consider 2 concepts:

- The complete state of the mathematical system at a given time, mathematically encoded as a **phase point** (classical mechanics) or a pure **quantum state vector** (quantum mechanics).
- An equation of motion which carries the state forward in time: **Hamilton's equations** (classical mechanics) or the **Schrödinger equation** (quantum mechanics).

Using these 2 concepts, the state at any other time, past or future, can in principle be calculated. There is however a disconnect between these laws & everyday life experiences, as we do not find it necessary (nor even theoretically possible) to know exactly at a microscopic level the simultaneous positions & velocities of each molecule while carrying out processes at the human scale (e.g., when performing a chemical reaction). Statistical mechanics fills this disconnection between the laws of mechanics & the practical experience of incomplete knowledge, by adding some uncertainty about which state the system is in.

Whereas ordinary mechanics only considers the behavior of a single state, statistical mechanics introduces the **statistical ensemble**, which is a large collection of virtual, independent copies of the system in various states. The statistical ensemble is a **probability distribution** over all possible states of the system. In classical statistical mechanics, the ensemble is a probability distribution over phase points (as opposed to a single phase point in ordinary mechanics), usually represented as a distribution in a **phase space** with **canonical coordinate** axes. In quantum statistical mechanics, the ensemble is a probability distribution over pure states & can be compactly summarized as a **density matrix**.

As is usual for probabilities, the ensemble can be interpreted in different ways:

- an ensemble can be taken to represent the various possible states that a *single system* could be in (**epistemic probability**, a form of knowledge), or
- the members of the ensemble can be understood as the states of the systems in experiments repeated on independent systems which have been prepared in a similar but imperfectly controlled manner (**empirical probability**), in the limit of an infinite number of trials.

These 2 meanings are equivalent for many purposes, & will be used interchangeably in this article.

However the probability is interpreted, each state in the ensemble evolves over time according to the equation of motion. Thus, the ensemble itself (the probability distribution over states) also evolves, as the virtual systems in the ensemble continually leave 1 state & enter another. The ensemble evolution is given by the **Liouville equation** (classical mechanics) or the **von Neumann equation** (quantum mechanics). These equations are simply derived by the application of the mechanical equation of motion separately to each virtual system contained in the ensemble, with the probability of the virtual system being conserved over time as it evolves from state to state.

1 special class of ensemble is those ensembles that do not evolve over time. These ensembles are known as *equilibrium ensembles* & their condition is known as *statistical equilibrium*. Statistical equilibrium occurs if, for each state in the ensemble, the ensemble also contains all of its future & past states with probabilities = probability of being in that state. (By contrast, **mechanical equilibrium** is a state with a balance of forces that has ceased to evolve.) The study of equilibrium ensembles of isolated systems is the focus of statistical thermodynamics. Non-equilibrium statistical mechanics addresses the more general case of ensembles that change over time, &/or ensembles of non-isolated systems.

### 1.14.3 Statistical thermodynamics

The primary goal of statistical thermodynamics (also known as equilibrium statistical mechanics) is to derive the **classical thermodynamics** of materials in terms of the properties of their constituent particles & the interactions between them. I.e., statistical thermodynamics provides a connection between the macroscopic properties of materials in **thermodynamic equilibrium**, & the microscopic behaviors & motions occurring inside the material.

Whereas statistical mechanics proper involves dynamics, here the attention is focused on *statistical equilibrium* (steady state). Statistical equilibrium does not mean that the particles have stopped moving (**mechanical equilibrium**), rather, only that the ensemble is not evolving.

1. **Fundamental postulate.**
2. **3 thermodynamics ensembles**
3. **Calculation methods.**

### 1.14.4 Non-equilibrium statistical mechanics

### 1.14.5 Applications

The ensemble formalism can be used to analyze general mechanical systems with uncertainty in knowledge about the state of a system. Ensembles are also used in:

- **propagation of uncertainty** over time,
- **regression analysis** of gravitational **orbits**
- **ensemble forecasting** of weather

- dynamics of **neural networks**
- bounded-rational **potential games** in **game theory** & economics.

Statistical physics explains & quantitatively describes **superconductivity**, **superfluidity**, **turbulence**, collective phenomena in **solids** & **plasma**, & the structural features of **liquid**. It underlies the modern **astrophysics**. In solid state physics, statistical physics aids the study of **liquid crystals**, **phase transitions**, & **critical phenomena**. many experimental studies of matter are entirely based on the statistical description of a system. These include the scattering of cold **neutrons**, **X-ray**, **visible light**, & more. Statistical physics also plays a role in materials science, nuclear physics, astrophysics, chemistry, biology, & medicine (e.g., study of the spread of infectious diseases).

Analytical & computational techniques derived from statistical physics of disordered systems, can be extended to large-scale problems, including ML, e.g., to analyze the weight space of deep **neural networks**. Statistical physics is thus finding applications in the area of **medical diagnostics**.

**Quantum statistical mechanics.** **Quantum statistical mechanics** is *statistical mechanics* applied to **quantum mechanical systems**. In quantum mechanics, a **statistical ensemble** (probability distribution over possible **quantum states**) is described by a **density operator**  $S$ , which is a nonnegative, **self-adjoint**, **trace-class** operator of trace 1 on the **Hilbert space**  $H$  describing the quantum system. This can be shown under various **mathematical formalisms for quantum mechanics**. 1 such formalism is provided by **quantum logic**.” – Wikipedia/statistical mechanics

## 1.15 Wikipedia/universal (metaphysics)

“In **metaphysics**, a *universal* is what particular things have in common, namely characteristics or qualities. In other words, universals are repeatable or recurrent entities that can be instantiated or exemplified by many particular things. E.g., suppose there are 2 chairs in a room, each of which is green. These 2 chairs share the quality “**chairness**”, as well as “greenness” or the quality of being green; in other words, they share 2 “universals”. There are 3 major kinds of qualities or characteristics: **types or kinds** (e.g. mammal), **properties** (e.g., short, strong), & **relations** (e.g., father of, next to). These are all different types of universals.

Paradigmatically, universals are **abstract** (e.g., humanity), whereas **particulars** are **concrete** (e.g., the personhood of Socrates). However, universals are not necessarily abstract & particulars are not necessarily concrete. E.g., one might hold that numbers are particular yet abstract objects. Likewise, some philosophers, e.g., **D. M. ARMSTRONG**, consider universals to be concrete.

Most do not consider **classes** to be universals, although some prominent philosophers do, e.g. JOHN BIGELOW.

### 1.15.1 Problem of universals

Main article: **Wikipedia/problem of universals**. The **Wikipedia/problem of universals** is an ancient problem in metaphysics on the existence of universals. The problem arises from attempts to account for the phenomenon of **similarity** or attribute agreement among things. E.g., **grass** & **GRANNY SMITH apples** are similar or agree in attribute, namely in having the attribute of greenness. The issue is how to account for this sort of agreement in attribute among things.

There are many philosophical positions regarding universals. Taking “**beauty**” as an example, 4 positions are:

- **Idealism**: beauty is a property constructed in the mind, so it exists only in descriptions of things.
- **Platonic extreme realism**: beauty is a property that exists in an ideal form independently of any mind or thing.
- **Aristotelian moderate realism** or **conceptualism**: beauty is a property of things (*fundamentum in re*) that the mind abstracts from these beautiful things.
- **Nominalism**: there are no universals, only individuals.

Taking a broader view, the main positions are generally considered classifiable as: **extreme realism**, **nominalism** (sometimes simply named “anti-realism” with regard to universals), **moderate realism**, & **idealism**. Extreme Realists posit the existence of independent, abstract universals to account for attribute agreement. Nominalists deny that universals exist, claiming that they are not necessary to explain attribute agreement. Conceptualists posit that universals exist only in the **mind**, or when conceptualized, denying the independent existence of universals, but accepting they have a *fundamentum in re*. Complications which arise include the implications of language use & the complexity of relating language to **ontology**.

### 1.15.2 Particular

Main article: **Wikipedia/particular**. A universal may have instances, known as its *particulars*. E.g., the type *dog* (or *doghood*) is a universal, as are the property *red* (or *redness*) & the relation *betweenness* (or *being between*). Any particular dog, red thing, or object that is between other things is not a universal, however, but is an *instance* of a universal. I.e., a universal type (*doghood*), property (*redness*), or relation (*betweenness*) **inheres** in a particular object (a specific dog, red thing, or object between other things).



### 1.15.3 Platonic realism

**Platonic realism** holds universals to be the **referents** of general items, e.g. the **abstract**, nonphysical, non-mental entities to which words e.g. “sameness”, “circularity”, & “beauty” refer. Particulars are the referents of proper names, e.g. “Phaedo”, or of definite descriptions that identify single objects, e.g. the phrase, “that person over there”. Other metaphysical theories may use the terminology of universals to describe physical entities.

Plato’s examples of what we might today call universals included mathematical & geometrical ideas such as a circle & natural numbers as universals. Plato’s views on universals did, however, vary across several different discussions. In some cases, PLATO spoke as if the perfect circle functioned as the **form** or blueprint for all copies & for the word definition of *circle*. In other discussions, PLATO describes particulars as “participating” in the associated universal.

Contemporary realists agree with the thesis that universals are multiply-exemplifiable entities. E.g., **D. M. ARMSTRONG**, **NICHOLAS WOLTERSTORFF**, **REINHARDT GROSSMANN**, **MICHAEL LOUX**.

### 1.15.4 Nominalism

Nominalists hold that universals are not real mind-independent entities but either merely concepts (sometimes called “conceptualism”) or merely names. Nominalists typically argue that properties are abstract particulars (like tropes) rather than universals. **JP MORELAND** distinguishes between “extreme” & “moderate” nominalism. Examples of nominalists include Buddhist logicians & **apoha** theorists, the medieval philosophers **ROSCELIN OF COMPIÈGNE** & **William of Ockham** & contemporary philosophers **W. V. O. QUINE**, **WILFRED SELLARS**, **D. C. Williams**, & **Keith Campbell**.

### 1.15.5 Ness-ity-hood principle

The *ness-ity-hood principle* is used mainly by English-speaking philosophers to generate convenient, concise names for universals or **properties**. According to the Ness-Itty-Hood Principle, a name for any universal may be formed by taking the name of the **predicate** & adding the suffix “ness”, “ity”, or “hood”. E.g., the universal that is distinctive of left-handers may be formed by taking the predicate “left-handed” & adding “ness”, which yields the name “left-handedness”. The principle is most helpful in cases where there is not an established or standard name of the universal in ordinary English usage: What is the name of the universal distinctive of chairs? “Chair” in English is used not only as a subject (as in “The chair is broken”), but also as a predicate (as in “That is a chair”). So to generate a name for the universal distinctive of chairs, take the predicate “chair” & add “ness”, which yields “chairness”. – [Wikipedia/universal \(metaphysics\)](#)

## 2 RICHARD P. FEYNMAN, ROBERT B. LEIGHTON, MATTHEW SANDS. 6 Easy Pieces: Essentials of Physics Explained By Its Most Brilliant Teacher

### Resources – Tài nguyên.

1. [\[FLS11a\]](#). RICHARD P. FEYNMAN. *Six Easy Pieces: Essentials of Physics Explained by Its Most Brilliant Teacher*. [3332 Amazon ratings][28807 Goodreads ratings]

Amazon review. Learn how to think like a physicist from a Nobel laureate & “1 of the greatest minds of the 20th century” (*New York Review of Books*) with these 6 classic & beloved lessons.

It was RICHARD P. FEYNMAN’s outrageous & scintillating method of teaching that earned him legendary status among students & professors of physics. From 1961 to 1963, FEYNMAN delivered a series of lectures at the California Institute of Technology that revolutionized the teaching of physics around the world. *6 Easy Pieces*, taken from these famous Lectures on Physics, represent the most accessible material from the series.

In these classic lessons, FEYNMAN introduces the general reader to the following topics: atoms, basic physics, energy, gravitation, quantum mechanics, & the relationship of physics to other topics. With his dazzling & inimitable wit, FEYNMAN presents each discussion with a minimum of jargon. Filled with wonderful examples & clever illustrations, *6 Easy Pieces* is the ideal introduction to the fundamentals of physics by 1 of the most admired & accessible physicists of modern times.

### Editorial reviews.

- “1 of the greatest minds of the 20th century.” – *New York Review of Books*
- “The essence of physics & FEYNMAN. No jargon<sup>3</sup>, just ideas, excitement, & the straight dope. & real answers, like ‘we don’t know.’” – STEPHEN WOLFRAM
- “The most original mind of his generation.” – FREEMAN DYSON
- “If 1 book was all that could be passed onto the next generation of scientists it would undoubtedly have to be *6 Easy Pieces*.” – JOHN GRIBBIN, *New Scientist*

About the Author. RICHARD P. FEYNMAN (1918–1988) was the Richard Chace Tolman Professor of Theoretical Physics at the California Institute of Technology. He was awarded the 1965 Nobel Prize for his work on the development of quantum field theory. He was also 1 of the most famous & beloved figures of the 20th century, both in physics & in the public arena.

<sup>3</sup>words or expressions that are used by a particular profession or group of people, & are difficult for others to understand. tiếng lóng.

**Publisher's Note.** *6 Easy Pieces* grew out of the need to bring to as wide an audience as possible a substantial yet nontechnical physics primer based on the science of RICHARD FEYNMAN. Chose 6 easiest chapters from FEYNMAN's celebrated & landmark text, *The Feynman Lectures on Physics*, originally published in 1963, which remains his most famous publication. General readers are fortunate that FEYNMAN chose to present certain key topics in largely qualitative terms without formal mathematics, & these are brought together for *6 Easy Pieces*.

## Introduction

"There is a popular misconception that science is an impersonal, dispassionate, & thoroughly objective enterprise. Whereas most other human activities are dominated by fashions, fads, & personalities, science is supposed to be constrained by agreed rules of procedure & rigorous tests. It is the results that count, not the people who produce them.

– Có một quan niệm sai lầm phổ biến rằng khoa học là một công việc khách quan, vô tư và hoàn toàn khách quan. Trong khi hầu hết các hoạt động khác của con người bị chi phối bởi thời trang, một nhất thời và tính cách, thì khoa học được cho là bị hạn chế bởi các quy tắc thủ tục đã thống nhất và các bài kiểm tra nghiêm ngặt. Kết quả mới là điều quan trọng chứ không phải con người tạo ra chúng.

This is, of course, manifest nonsense. Science is a people-driven activity like all human endeavor, & just as subject to fashion & whim. In this case fashion is set not so much by choice of subject matter, but by the way scientists think about the world. Each age adopts its particular approach to scientific problems, usually following the trail blazed by certain dominant figures who both set the agenda & define the best methods to tackle it. Occasionally scientists attain sufficient stature that they become noticed by the general public, & when endowed with outstanding flair a scientist may become an icon for the entire scientific community. In earlier centuries ISAAC NEWTON was an icon. NEWTON personified the gentleman scientist – well connected, devoutly religious, unhurried, & methodical in his work. His style of doing science set the standard for 200 years. In the 1st half of the 20th century ALBERT EINSTEIN replaced NEWTON as the popular scientist icon. Eccentric, disheveled, Germanic, absent-minded, utterly absorbed in his work, & an archetypal abstract thinker, EINSTEIN changed the way that physics is done by questioning the very concepts that define the subject.

RICHARD FEYNMAN has become an icon for late 20th-century physics – the 1st American to achieve this status. Born in New York in 1918 & educated on the East Coast, he was too late to participate in the Golden Age of physics, which, in the 1st 3 decades of this century, transformed our worldview with the twin revolutions of the theory of relativity & quantum mechanics. These sweeping developments laid the foundations of the edifice we now call the New Physics. FEYNMAN started with those foundations & helped build the ground floor of the New Physics. His contributions touched almost every corner of the subject & have had a deep & abiding influence over the way that physicists think about the physical universe.

FEYNMAN was a theoretical physicist par excellence. NEWTON had been both experimentalist & theorist in equal measure. EINSTEIN was quite simply contemptuous of experiment, preferring to put his faith in pure thought. FEYNMAN was driven to develop a deep theoretical understanding of nature, but he always reminded close to the real & often grubby world of experimental results. Nobody who watched the elderly FEYNMAN elucidate the cause of the Challenger space shuttle disaster by dipping an elastic band in ice water could doubt that here was both a showman & a very practical thinker.

Initially, FEYNMAN made a name for himself from his work on the theory of subatomic particles, specifically the topic known as quantum electrodynamics or QED. Quantum theory began with this topic. In 1900, the German physicist MAX PLANCK proposed that light & other electromagnetic radiation, which had hitherto been regarded as waves, paradoxically behaved like tiny packets of energy, or "quanta," when interacting with matter. These particular quanta became known as photons. By the early 1930s the architects of the new quantum mechanics had worked out a mathematical scheme to describe the emission & absorption of photons by electrically charged particles e.g. electrons. Although this early formulation of QED enjoyed some limited success, the theory was clearly flawed. In many cases calculations gave inconsistent & even infinite answers to well-posed physical questions. It was to the problem of constructing a consistent theory of QED that the young FEYNMAN turned his attention in the late 1940s.

To place QED on a sound basis it was necessary to make the theory consistent not only with the principles of quantum mechanics but with those of the special theory of relativity too. These 2 theories come with their own distinctive mathematical machinery, complicated systems of equations that can indeed be combined & reconciled to yield a satisfactory description of QED. Doing this was a tough undertaking, requiring a high degree of mathematical skill, & was the approach followed by FEYNMAN's contemporaries. FEYNMAN himself, however, took a radically different route – so radical, in fact, that he was more or less able to write down the answers straightaway without using any mathematics!

To aid this extraordinary feat of intuition, FEYNMAN invented a simple system of eponymous diagrams. FEYNMAN diagrams are a symbolic but powerfully heuristic way of picturing what is going on when electrons, photons, & other particles interact with each other. These days FEYNMAN diagrams are a routine aid to calculation, but in the early 1950s they marked a startling departure from the traditional way of doing theoretical physics.

The particular problem of constructing a consistent theory of quantum electrodynamics, although it was a milestone in the development of physics, was just the start. It was to define a distinctive FEYNMAN style, a style destined to produce a string of important results from a broad range of topics in physical science. The FEYNMAN style can best be described as a mixture of reverence & disrespect for received wisdom.

Physics is an exact science, & the existing body of knowledge, while incomplete, can't simply be shrugged aside. FEYNMAN acquired a formidable grasp of the accepted principles of physics at a very young age, & he chose to work almost entirely on conventional problems. He was not the sort of genius to beaver away in isolation in a backwater of the discipline & to stumble across the profoundly new. His special talent was to approach essentially mainstream topics in an idiosyncratic way. This meant eschewing existing formalisms & developing his own highly intuitive approach. Whereas most theoretical physicists rely on careful



mathematical calculation to provide a guide & a crutch to take them into unfamiliar territory, FEYNMAN's attitude was almost cavalier. You yet the impression that he could read nature like a book & simply report on what he found, without the tedium of complex analysis.

Indeed, in pursuing his interests in this manner FEYNMAN displayed a healthy contempt for rigorous formalisms. It is hard to convey the depth of genius that is necessary to work like this. Theoretical physics is 1 of the toughest intellectual exercises, combining abstract concepts that defy visualization with extreme mathematical complexity. Only by adopting the highest standards of mental discipline can most physicists make progress. Yet FEYNMAN appeared to ride roughshod over this strict code of practice & pluck new results like ready-made fruit from the Tree of Knowledge.

The FEYNMAN style owed a great deal to the personality of the man. In his professional & private life he seemed to treat the world as a hugely entertaining game. The physical universe presented him with a fascinating series of puzzles & challenges, & so did his social environment. A lifelong prankster, he treated authority & the academic establishment with the same sort of disrespect he showed for stuffy mathematical formalism. Never one to suffer fools gladly, he broke the rules whenever he found them arbitrary or absurd. His autobiographical writings contain amusing stories of FEYNMAN outwitting the atom-bomb security services during the war, FEYNMAN cracking safes, FEYNMAN disarming women with outrageously bold behavior. He treated his Nobel Prize, awarded for his work on QED, in a similar take-it-or-leave-it manner.

Alongside this distaste for formality, FEYNMAN had a fascination with the quirky & obscure. Many will remember his obsession with the long-lost country of Tuva in Central Asia, captured so delightfully in a documentary film made near the time of his death. His other passions included playing the bongo drums, painting, frequenting strip clubs, & deciphering Mayan texts.

FEYNMAN himself did much to cultivate his distinctive persona. Although reluctant to put pen to paper, he was voluble in conversation, & loved to tell stories about his ideas & escapades. These anecdotes, accumulated over the years, helped add to his mystique & made him a proverbial legend in his own lifetime. His engaging manner endeared him greatly to students, especially the younger ones, many of whom idolized him. When FEYNMAN died of cancer in 1988 the students at Caltech, where he had worked for most of his career, unfurled a banner with the simple message: "We love you Dick."

It was FEYNMAN's happy-go-lucky approach to life in general & physics in particular that made him such a superb communicator. He had little time for formal lecturing or even for supervising Ph.D. students. Nevertheless he could give brilliant lectures when it suited him, deploying all the sparkling wit, penetrating insight, & irreverence that he brought to bear on his research work.

In the early 1960s FEYNMAN was persuaded to teach an introductory physics course to Caltech freshman & sophomores. He did so with characteristic panache & his inimitable blend of informality, zest, & offbeat humor. Fortunately, these priceless lectures were saved for posterity in book form. Though far removed in style & presentation from more conventional teaching texts, *The Feynman Lectures on Physics* were a huge success, & they excited & inspired a generation of students across the world. 3 decades on, these volumes have lost nothing of their sparkle & lucidity. *6 Easy Pieces* is culled directly from *The Feynman Lectures on Physics*. It is intended to give general readers a substantive taste of FEYNMAN the Educator by drawing on the early, nontechnical chapters from that landmark work. The result is a delightful volume – it serves both as a primer on physics for nonscientists & as a primer on FEYNMAN himself.

What is most impressive about FEYNMAN's carefully crafted exposition is the way that he can develop far-reaching physical notions from the most slender investment in concepts, & a minimum in the way of mathematics & technical jargon. He has the knack of finding just the right analogy or everyday illustration to bring out the essence of a deep principle, without obscuring it in incidental or irrelevant details.

The selection of topics contained in this volume is not intended as a comprehensive survey of modern physics, but as a tantalizing taste of the FEYNMAN approach. We soon discover how he can illuminate even mundane topics like force & motion with new insights. Key concepts are illustrated by examples drawn from daily life or antiquity. Physics is continually linked to other sciences while leaving the reader in no doubt about which is the fundamental discipline.

Right at the beginning of *6 Easy Pieces* we learn how all physics is rooted in the notion of law – the existence of an ordered universe that can be understood by the application of rational reasoning. However, the laws of physics are not transparent to us in our direct observations of nature. They are frustratingly hidden, subtly encoded in the phenomena we study. The arcane procedures of the physicist – a mixture of carefully designed experimentation & mathematical theorizing – are needed to unveil the underlying law-like reality.

Possibly the best-known law of physics is NEWTON's inverse square law of gravitation, discussed here in Chap. 5. The topic is introduced in the context of the solar system & KEPLER's laws of planetary motion. But gravitation is universal, applying across the cosmos, enabling FEYNMAN to spice his account with examples from astronomy & cosmology. Commenting on a picture of a globular cluster somehow held together by unseen forces, he waxes lyrical: "If one cannot see gravitational acting here, he has no soul."

Other laws are known that refer to the various nongravitational forces of nature that describe how particles of matter interact with each other. There is but a handful of these forces, & FEYNMAN himself holds the considerable distinction of being 1 of the few scientists in history to discover a new law of physics, pertaining to the way that a weak nuclear force affects the behavior of certain subatomic particles.

High-energy particle physics was the jewel in the crown of postwar science, at once awesome & glamorous, with its huge accelerator machines & seemingly unending list of newly discovered subatomic particles. FEYNMAN's research was directed mostly toward making sense of the results of this enterprise. A great unifying theme among particle physicists has been the role of symmetry & conservation laws in bringing order to the subatomic zoo.

As it happens, many of the symmetries known to particle physicists were familiar already in classical physics. Chief among these are the symmetries that arise from the homogeneity of space & time. Take time: apart from cosmology, where the big bang marked the beginning of time, there is nothing in physics to distinguish 1 moment of time from the next. Physicists say

that the world is “invariant under time translations,” i.e., whether you take midnight or midday to be the zero of time in your measurements, it makes no difference to the description of physical phenomena. Physical processes do not depend on an absolute zero of time. It turns out that this symmetry under time translation directly implies 1 of the most basic, & also most useful, laws of physics: the law of conservation of energy. This law says that you can move energy around & change its form but you can’t create or destroy it. FEYNMAN makes this law crystal clear with his amusing story of DENNIS the Menace who is always mischievously hiding his toy building blocks from his mother (Chap. 4).

The most challenging lecture in this volume is the last, which is an exposition on quantum physics. It is no exaggeration to say that quantum mechanics had dominated 20th-century physics & is far & away the most successful scientific theory in existence. It is indispensable for understanding subatomic particles, atoms, & nuclei, molecules, & chemical bonding, the structure of solids, superconductors & superfluids, the electrical & thermal conductivity of metals & semiconductors, the structure of stars, & much else. It has practical applications ranging from the laser to the microchip. All this from a theory that at 1st sight – & 2nd sight – looks absolutely crazy! NIELS BOHR, 1 of the founders of quantum mechanics, once remarked that anybody who is not shocked by the theory hasn’t understood it.

The problem is that quantum ideas strike at the very heart of what we might call commonsense reality. In particular, the idea that physical objects e.g. electrons or atoms enjoy an independent existence, with a complete set of physical properties at all times, is called into question. E.g., an electron cannot have a position in space & a well-defined speed at the same moment. If you look for where an electron is located, you will find it at a place, & if you measure its speed you will obtain a definite answer, but you cannot make both observations at once. Nor is it meaningful to attribute definite yet unknown values for the position & speed to an electron in the absence of a complete set of observations.

This indeterminism in the very nature of atomic particles is encapsulated by HEISENBERG’s celebrated uncertainty principle. This puts strict limits on the precision with which properties e.g. position & speed can be simultaneously known. A sharp value for position smears the range of possible values of speed & vice versa. Quantum fuzziness shows up in the way electrons, photons, & other particles move. Certain experiments can reveal them taking definite paths through space, after the fashion of bullets following trajectories toward a target. But other experimental arrangements reveal that these entities can also behave like waves, showing characteristic patterns of diffraction & interference.

FEYNMAN’s masterly analysis of the famous “2-slit” experiment, which teases out the “shocking” wave-particle duality in its starkest form, has become a classic in the history of scientific exposition. With a few very simple ideas, FEYNMAN manages to take the reader to the very heart of the quantum mystery, & leaves us dazzled by the paradoxical nature of reality that it exposes.

Although quantum mechanics had made the textbooks by the early 1930s, it is typical of FEYNMAN that, as a young man, he preferred to refashion the theory for himself in an entirely new guise. The FEYNMAN method has the virtue that it provides us with a vivid picture of nature’s quantum trickery at work. The idea is that the path of a particle through space is not generally well defined in quantum mechanics. We can imagine a freely moving electron, say, not merely traveling in a straight line between A & B as common sense would suggest, but taking a variety of wiggly routes. FEYNMAN invites us to imagine that somehow the electron explores all possible routes, & in the absence of an observation about which path is taken we must suppose that all these alternative paths somehow contribute to the reality. So when an electron arrives at a point in space – say a target screen – many different histories must be integrated together to create this 1 event.

FEYNMAN’s so-called *path-integral*, or sum-over-histories approach to quantum mechanics, set this remarkable concept out as a mathematical procedure. It remained more or less a curiosity for many years, but as physicists pushed quantum mechanics to its limits – applying it to gravitation & even cosmology – so the FEYNMAN approach turned out to offer the best calculational tool for describing a quantum universe. History may well judge that, among his many outstanding contributions to physics, the path-integral formulation of quantum mechanics is the most significant.

Many of the ideas discussed in this volume are deeply philosophical. Yet FEYNMAN had an abiding suspicion of philosophers. I once had occasion to tackle him about the nature of mathematics & the laws of physics, & whether abstract mathematical laws could be considered to enjoy an independent Platonic existence. He gave a spirited & skillful description of why this indeed appears so but soon backed off when I pressed him to take a specific philosophical position. He was similarly wary when I attempted to draw him out on the subject of reductionism. With hindsight, I believe that FEYNMAN was not, after all, contemptuous of philosophical problems. But, just as he was able to do fine mathematical physics without systematic mathematics, so he produced some fine philosophical insights without systematic philosophy. It was formalism he disliked, not content.

It is unlikely that the world will see another RICHARD FEYNMAN. He was very much a man of his time. The FEYNMAN style worked well for a subject that was in the process of consolidating a revolution & embarking on the far-reaching exploration of its consequences. Postwar physics was secure in its foundations, mature in its theoretical structures, yet wide open for kibitzing exploitation. FEYNMAN entered a wonderland of abstract concepts & imprinted his personal brand of thinking upon many of them. This book provides a unique glimpse into the mind of a remarkable human being.” – Sep 1994, PAUL DAVIES

## Special Preface from The Feynman Lectures on Physics

Toward the end of his life, RICHARD FEYNMAN’s fame had transcended the confines of the scientific community. His exploits as a member of the commission investigating the space shuttle *Challenger* disaster gave him widespread exposure; similarly, a best-selling book about his picaresque adventures made him a folk hero almost of the proportions of ALBERT EINSTEIN. But back in 1961, even before his Nobel Prize increased his visibility to the general public, FEYNMAN was more than merely famous among members of the scientific community – he was legendary. Undoubtedly, the extraordinary power of his teaching helped spread & enrich the legend of RICHARD FEYNMAN.

He was a truly great teacher, perhaps the greatest of his era & ours. For FEYNMAN, the lecture hall was a theater, & the lecturer a performer, responsible for providing drama & fireworks as well as facts & figures. He would prowl about the front of a classroom, arms waving, “the impossible combination of theoretical physicist & circus barker, all body motion & sound effects,” wrote *The New York Times*. Whether he addressed an audience of students, colleagues, or the general public, for those lucky enough to see FEYNMAN lecture in person, the experience was usually unconventional & always unforgettable, like the man himself.

He was the master of high drama, adept at riveting the attention of every lecture-hall audience. Many years ago, he taught a course in Advanced Quantum Mechanics, a large class comprised of a few registered graduate students & most of the Caltech physics faculty. During 1 lecture, FEYNMAN started explaining how to represent certain complicated integrals diagrammatically: time on this axis, space on that axis, wiggly line for this straight line, etc. Having described what is known to the world of physics as a FEYNMAN diagram, he turned around to face the class, grinning wickedly, “& this is caled *THE* diagram!” FEYNMAN had reached the denouement, the lecture hall erupted with spontaneous applause.

For many years after the lectures that make up this book were given, FEYNMAN was an occasional guest lecturer for Caltech’s freshman physics course. Naturally, his appearances had to be kept secret so there would be room left in the hall for the registered students. At 1 such lecture the subject was curved-space time, & FEYNMAN was characteristically brilliant. But the unforgettable moment came at the beginning of the lecture. The supernova of 1987 had just been discovered, & FEYNMAN was very excited about it. He said, “Tycho Brahe had his supernova, & Kepler had his. Then there weren’t any for 400 years. But now I have mine.” The class fell silent, & FEYNMAN continued on. “There are  $10^{11}$  stars in the galaxy. That used to be a *huge* number. But it’s only  $10^{11}$ . It’s < the national deficit! We used to call them astronomical numbers. Now we should call them economical numbers.” The class dissolved in laughter, & FEYNMAN, having captured his audience, went on with his lecture.

Showmanship aside, FEYNMAN’s pedagogical technique was simple. A summation of his teaching philosophy was found among his papers in the Caltech archives, in a note he had scribbled to himself while in Brazil in 1952:

1st figure out why you want the students to learn the subject & what you want them to know, & the method will result more or less by common sense.

What came to FEYNMAN by “common sense” were often brilliant twists that perfectly captured the essence of his point. Once, during a public lecture, he was trying to explain why one must not verify an idea using the same data that suggested the idea in the 1st place. Seeming to wander off the subject, FEYNMAN began talking about license plates. “You know, the most amazing thing happened to me tonight. I was coming here, on the way to the lecture, & I came in through the parking lot. & you won’t believe what happened. I saw a car with the license plate ARW 357. Can you imagine? Of all the millions of license plates in the state, what was the chance that I would see that particular one tonight? Amazing!” A point that even many scientists fail to grasp was made clear through FEYNMAN’s remarkable “common sense.”

In 35 years at Caltech (1952–1987), FEYNMAN was listed as teacher of record for 34 courses. 25 of them were advanced graduate courses, strictly limited to graduate students, unless undergraduates asked permission to take them (they often did, & permission was nearly always granted). The rest were mainly introductory graduate courses. Only once did FEYNMAN teach courses purely for undergraduates, & that was the celebrated occasion in the academic years 1961–1962 & 1962–1963, with a brief reprise in 1964, when he gave the lectures that were to become *The Feynman Lectures on Physics*.

At the time there was a consensus at Caltech that freshman & sophomore students were getting turned off rather than spurred on by their 2 years of compulsory physics. To remedy the situation, FEYNMAN was asked to design a series of lectures to be given to the students over the course of 2 years, 1st to freshman, & then to the same class as sophomores. When he agreed, it was immediately decided that the lectures should be transcribed for publication. That job turned out to be far more difficult than anyone had imagined. Turning out publishable books require a tremendous amount of work on the part of his colleagues, as well as FEYNMAN himself, who did the final editing of every chapter.

& the nuts & bolts of running a course had to be addressed. This task was greatly complicated by the fact that FEYNMAN had only a vague outline of what he wanted to cover. I.e., no one knew what FEYNMAN would say until he stood in front of a lecture hall filled with students & said it. The Caltech professors who assisted him would then scramble as best they could to handle mundane details, e.g., making up homework problems.

Why did FEYNMAN devote > 2 years to revolutionize the way beginning physics was taught? One can only speculate, but there were probably 3 basic reasons. One is that he loved to have an audience, & this gave him a bigger theater than he usually had in graduate courses. The 2nd was that he genuinely cared about students, & he simply thought that teaching freshmen was an important thing to do. The 3rd & perhaps most important reason was the sheer challenge of reformulating physics, as he understood it, so that it could be presented to young students. This was his specialty, & was the standard by which he measured whether something was really understood. FEYNMAN was once asked by a Caltech faculty member to explain why spin  $\frac{1}{2}$  particles obey Fermi-Dirac statistics. He gauged his audience perfectly & said, “I’ll prepare a freshman lecture on it.” But a few days later he returned & said, “You know, I couldn’t do it. I couldn’t reduce it to the freshman level. That means we really don’t understand it.”

This specialty of reducing deep ideas to simple, understandable terms is evident throughout *The Feynman Lectures on Physics*, but nowhere more so than in his treatment of quantum mechanics. To aficionados, what he has done is clear. He has presented, to beginning students, the path integral method, the technique of his own devising that allowed him to solve some of the most profound problems in physics. His own work using path integrals, among other achievements, led to the 1965 Nobel Prize that he shared with Julian Schwinger & Sin-Itaro Tomonaga.

Through the distant veil of memory, many of the students & faculty attending the lectures have said that having 2 years of physics with FEYNMAN was the experience of a lifetime. But that’s not how it seemed at the time. Many of the students dreaded the class, & as the course wore on, attendance by the registered students started dropping alarmingly. But at the same time,

more & more faculty & graduate students started attending. The room stayed full, & FEYNMAN may never have known he was losing some of his intended audience. But even in FEYNMAN's view, his pedagogical endeavor did not succeed. He wrote in the 1963 preface to the *Lectures*: "I don't think I did very well by the students." Rereading the books, one sometimes seems to catch FEYNMAN looking over his shoulder, not at his young audience, but directly at his colleagues, saying, "Look at that! Look how I finessed that point! Wasn't that clever?" But even when he thought he was explaining things lucidly to freshmen or sophomores, it was not really they who were able to benefit most from what he was doing. It was his peers – scientists, physicists, & professors – who would be the main beneficiaries of his magnificent achievement, which was nothing less than to see physics through the fresh & dynamic perspective of RICHARD FEYNMAN.

FEYNMAN was more than a great teacher. His gift was that he was an extraordinary teacher of teachers. If the purpose in giving *The Feynman Lectures on Physics* was to prepare a roomful of undergraduate students to solve examination problems in physics, he cannot be said to have succeeded particularly well. Moreover, if the intent was for the books to serve as introductory college textbooks, he cannot be said to have achieved his goal. Nevertheless, the books have been translated into 10 foreign languages & are available in 4 bilingual editions. FEYNMAN himself believed that his most important contribution to physics would not be QED, or the theory of superfluid helium, or polarons, or partons. His foremost contribution would be the 3 red books of *They Feynman Lectures on Physics*. That belief fully justifies this commemorative issue of these celebrated books." – Apr 1989, DAVID L. GOODSTEIN, GERRY NEUGEBAUER, California Institute of Technology

## 2.1 FEYNMAN's Preface from The Feynman Lectures on Physics

"These are the lectures in physics that I gave last year & the year before to the freshman & sophomore classes at Caltech. The lectures are, of course, not verbatim – they have been edited, sometimes extensively & sometimes less so. The lectures form only part of the complete course. The whole group of 180 students gathered in a big lecture room twice a week to hear these lectures & then they broke up into small groups of 15–20 students in recitation sections under the guidance of a teaching assistant. In addition, there was a laboratory session once a week.

The special problem we tried to get at with these lectures was to maintain the interest of the very enthusiastic & rather smart students coming out of the high schools & into Caltech. They have heard a lot about how interesting & exciting physics is – the theory of relativity, quantum mechanics, & other modern ideas. By the end of 2 years of our previous course, many would be very discouraged because they were really very few grand, new, modern ideas presented to them. They were made to study inclined planes, electrostatics, & so forth, & after 2 years it was quite stultifying. The problem was whether or not we could make a course which would save the more advanced & excited student by maintaining his enthusiasm.

The lectures here are not in any way meant to be a survey course, but are very serious. I thought to address them to the most intelligent in the class & to make sure, if possible, that even the most intelligent student was unable to completely encompass everything that was in the lectures – by putting in suggestions of applications of the ideas & concepts in various directions outside the main line of attack. For this reason, though, I tried very hard to make all the statements as accurate as possible, to point out in every case where the equations & ideas fitted into the body of physics, & how – when they learned more – things would be modified. I also felt that for such students it is important to indicate what it is that they should – if they are sufficiently clever – be able to understand by deduction from what has been said before, & what is being put in as something new. When new ideas came in, I would try either to deduce them if they were deducible, or to explain that it *was* a new idea which hadn't any basis in terms of things they had already learned & which was not supposed to be provable – but was just added in.

At the start of these lectures, I assumed that the students knew something when they came out of high school – such things as geometrical optics, simple chemistry ideas, & so on. I also didn't see that there was any reason to make the lectures in a definite order, in the sense that I would not be allowed to mention something until I was ready to discuss it in detail. There was a great deal of mention of things to come, without complete discussions. These more complete discussions would come later when the preparation became more advanced. Examples are the discussions of inductance, & of energy levels, which are at 1st brought in in a very qualitative way & are later developed more completely.

At the same time that I was aiming at the more active student, I also wanted to take care of the fellow for whom the extra fireworks & side applications are merely disquieting & who cannot be expected to learn most of the material in the lecture at all. For such students, I wanted there to be at least a central core or backbone of material which he *could* get. Even if he didn't understand everything in a lecture, I hoped he wouldn't get nervous. I didn't expect him to understand everything, but only the central & most direct features. It takes, of course, a certain intelligence on his part to see which are the central theorems & central ideas, & which are the more advanced side issues & applications which he may understand only in later years.

In giving these lectures there was 1 serious difficulty: in the way the course was given, there wasn't any feedback from the students to the lecturer to indicate how well the lectures were going over. This is indeed a very serious difficulty, & I don't know how good the lectures really are. The whole thing was essentially an experiment. & if I did it again I wouldn't do it the same way – I hope I *don't* have to do it again! I think, though, that things worked out – so far as the physics is concerned – quite satisfactorily in the 1st year.

In the 2nd year I was not so satisfied. In the 1st part of the course, dealing with electricity & magnetism, I couldn't think of any really unique or different way of doing it – of any way that would be particularly more exciting than the usual way of presenting it. So I don't think I did very much in the lectures on electricity & magnetism. At the end of the 2nd year I had originally intended to go on, after the electricity & magnetism, by giving some more lectures on the properties of materials, but mainly to take up things like fundamental modes, solutions of the diffusion equation, vibrating systems, orthogonal functions, ... developing the 1st stages of what are usually called "the mathematical methods of physics." In retrospect, I think that if I were doing it again I would go back to that original idea. But since it was not planned that I would be giving these lectures

again, it was suggested that it might be a good idea to try to give an introduction to the quantum mechanics – what you will find in Vol. III.

It is perfectly clear that students who will major in physics can wait until their 3rd year for quantum mechanics. On the other hand, the argument was made that many of the students in our course study physics as a background for their primary interest in other fields. & the usual way of dealing with quantum mechanics makes that subject almost unavailable for the great majority of students because they have to take so long to learn it. Yet, in its real applications – especially in its more complex applications, e.g., in electrical engineering & chemistry – the full machinery of the differential equation approach is not actually used. So I tried to describe the principles of quantum mechanics in a way which wouldn't require that one 1st know the mathematics of PDEs. Even for a physicist I think that is an interesting thing to try to do – to present quantum mechanics in this reverse fashion – for several reasons which may be apparent in the lectures themselves. However, I think that the experiment in the quantum mechanics part was not completely successful – in large part because I really did not have enough time at the end (I should, e.g., have had 3 or 4 more lectures in order to deal more completely with such matters as energy bands & the spatial dependence of amplitudes). Also, I had never presented the subject this way before, so the lack of feedback was particularly serious. I now believe the quantum mechanics should be given at a later time. Maybe I'll have a chance to do it again someday. Then I'll do it right.

The reason there are no lectures on how to solve problems is because there were recitation sections. Although I did put in 3 lectures in the 1st year on how to solve problems, they are not included here. Also there as a lecture on inertial guidance which certainly belongs after the lecture on rotating systems, but which was, unfortunately, omitted. The 5th & 6th lectures are actually due to MATTHEW SANDS, as I was out of town.

The question, of course, is how well this experiment has succeeded. My own point of view – which, however, does not seem to be shared by most of the people who worked with the students – is pessimistic. I don't think I did very well by the students. When I look at the way the majority of the students handled the problems on the examinations, I think that the system is a failure. Of course, my friends point out to me that there were 1 or 2 dozen students who – very surprisingly – understood almost everything in all of the lectures, & who were quite active in working with the material & worrying about the many points in an excited & interested way. These people have now, I believe, a 1st-rate background in physics – & they are, after all, the ones I was trying to get at. But then, “The power of instruction is seldom of much efficacy except in those happy dispositions where it is almost superfluous.” (Gibbon)

Still, I didn't want to leave any student completely behind, as perhaps I did. I think 1 way we could help the students more would be by putting more hard work into developing a set of problems which would elucidate some of the ideas in the lectures. Problems give a good opportunity to fill out the material of the lectures & make more realistic, more complete, & more settled in the mind the ideas that have been exposed.

I think, however, that there isn't any solution to this problem of education other than to realize that the best teaching can be done only when there is a direct individual relationship between a student & a good teacher – a situation in which the student discusses the ideas, thinks about the things, & talks about the things. It's impossible to learn very much by simply sitting in a lecture, or even by simply doing problems that are assigned. But in our modern times we have so many students to teach that we have to try to find some substitute for the ideal. Perhaps my lectures can make some contribution. Perhaps in some small place where there are individual teachers & students, they may get some inspiration or some ideas from the lectures. Perhaps they will have fun thinking them through – or going on to develop some of the ideas further.” Jun 1963, RICHARD P. FEYNMAN

## 2.2 Atoms in Motion

### 2.2.1 Introduction

This 2-year course in physics is presented from the point of view that you, the reader, are going to be a physicist. This is not necessarily the case of course, but that is what every professor in every subject assumes! If you are going to be a physicist, you will have a lot to study: 200 years of the most rapidly developing field of knowledge that there is. So much knowledge, in fact, that you might think that you cannot learn all of it in 4 years, & truly you cannot; you will have to go to graduate school too!

Surprisingly enough, in spite of the tremendous amount of work that has been done for all this time it is possible to condense the enormous mass of results to a large extent – i.e., to find *laws* which summarize all our knowledge. Even so, the laws are so hard to grasp that it is unfair to you to start exploring this tremendous subject without some kind of map or outline of the relationship of 1 part of the subject of science to another. Following these preliminary remarks, the 1st 3 chapters will therefore outline the relation of physics to the rest of the sciences, the relations of the sciences to each other, & the meaning of science, to help us develop a “feel” for the subject.

You might ask why we cannot teach physics by just giving the basic laws on page 1 & then showing how they work in all possible circumstances, as we do in Euclidean geometry, where we state the axioms & then make all sorts of deductions. (So, not satisfied to learn physics in 4 years, you want to learn it in 4 minutes?) We cannot do it in this way for 2 reasons. 1st, we do not yet *know* all the basic laws: there is an expanding frontier of ignorance. 2nd, the correct statement of the laws of physics involves some very unfamiliar ideas which require advanced mathematics for their description. Therefore, one needs a considerable amount of preparatory training even to learn what the *words* mean. No, it is not possible to do it that way. We can only do it piece by piece.

Each piece, or part, of the whole of nature is always merely an *approximation* to the complete truth, or the complete truth so far as we know it. In fact, everything we know is only some kind of approximation, because *we know that we do not know all the laws* as yet. Therefore, things must be learned only to be unlearned again or, more likely, to be corrected.

The principle of science, the definition, almost, is: *The test of all knowledge is experiment*. Experiment is the *sole judge* of scientific “truth.” But what is the source of knowledge? Where do the laws that are to be tested come from? Experiment, itself, helps to produce these laws, in the sense that it gives us hints. But also needed is *imagination* to create from these hints the great generalizations – to guess at the wonderful, simple, but very strange patterns beneath them all, & then to experiment to check again whether we have made the right guess. This imagining process is so difficult that there is a division of labor in physics: there are *theoretical* physicists who imagine, deduce, & guess at new laws, but do not experiment; & then there are *experimental* physicists who experiment, imagine, deduce, & guess.

We said that the laws of nature & approximate: that we 1st find the “wrong” ones, & then we find the “right” ones. Now, how can an experiment be “wrong”? 1st, in a trivial way: if something is wrong with the apparatus that you did not notice. But these things are easily fixed, & checked back & forth. So without snatching at such minor things, how *can* the results of an experiment be wrong? Only by being inaccurate. E.g., the mass of an object never seems to change: a spinning top has the same weight as a still one. So a “law” was invented: mass is constant, independent of speed. That “law” is now found to be incorrect. Mass is found to increase with velocity, but appreciable increases require velocities near that of light. A *true* law is: if an object moves with a speed of  $< 100$  miles a second the mass is constant to within 1 part in a million. In some such approximate form this is a correct law. So in practice one might think that the new law makes no significant difference. Well, yes & no. For ordinary speeds we can certainly forget it & use the simple constant-mass law as a good approximation. But for high speeds we are wrong, & the higher the speed, the more wrong we are.

Finally, & most interesting, *philosophically we are completely wrong* with the approximate law. Our entire picture of the world has to be altered even though the mass changes only by a little bit. This is a very peculiar thing about the philosophy, or the ideas, behind the laws. Even a very small effect sometimes requires profound changes in our ideas.

Now, what should we teach 1st? Should we teach the *correct* but unfamiliar law with its strange & difficult conceptual ideas, e.g., theory of relativity, 4-dimensional space-time, & so on? Or should we 1st teach the simple “constant-mass” law, which is only approximate, but does not involve such difficult ideas? The 1st is more exciting, more wonderful, & more fun, but the 2nd is easier to get at 1st, & is a 1st step to a real understanding of the 1st idea. This point arises again & again in teaching physics. At different times we shall have to resolve it in different ways, but at each stage it is worth learning what is now known, how accurate it is, how it fits into everything else, & how it may be changed when we learn more.

Let us now proceed with our outline, or general map, of our understanding of science today (in particular, physics, but also of other sciences on the periphery), on that when we later concentrate on some particular point we will have some idea of the background, why that particular point is interesting, & how it fits into the big structure. So, what *is* our overall picture of the world?” – [FLS11a, pp. 1–4]

## 2.2.2 Matter is made of atoms

“If, in some cataclysm, all of scientific knowledge were to be destroyed, & only 1 sentence passed on to the next generations of creatures, what statement would contain the most information in the fewest words? I believe it is the *atomic hypothesis* (or the *atomic fact*, or whatever you wish to call it) that *all things are made of atoms – little particles that move around in perpetual motion, attracting each other when they are a little distance apart, but repelling upon being squeezed into one another*. In that 1 sentence, you will see, there is an *enormous* amount of information about the world, if just a little imagination & thinking are applied.

To illustrate the power of the atomic idea, suppose that we have a drop of water a quarter of an inch on the side. If we look at it very closely we see nothing but water – smooth, continuous water. Even if we magnify it with the best optical microscope available – roughly 2000 times – then the water drop will be roughly 40 feet across, about as big as a large room, & if we looked rather closely, we would *still* see relatively smooth water – but here & there small football-shaped things swimming back & forth. Very interesting. These are paramecia. You may stop at this point & get so curious about the paramecia with their wiggling cilia & twisting bodies that you go no further, except perhaps to magnify the paramecia still more & see inside. This, of course, is a subject for biology, but for the present we pass on & look still more closely at the water material itself, magnifying it 2000 times again. Now the drop of water extends about 15 miles across, & if we look very closely at it we see a kind of teeming, something which no longer has a smooth appearance – it looks something like a crowd at a football game as seen from a very great distance. In order to see what this teeming is about, we will magnify it another 2050 times & we will see something similar to Fig. 1.1 – a picture of water magnified  $10^9$  times, but idealized in several ways. In the 1st place, the particles are drawn in a simple manner with sharp edges, which is inaccurate. Secondly, for simplicity, they are sketched almost schematically in a 2D arrangement, but of course they are moving around in 3D. Notice that there are 2 kinds of “blobs” or circles to represent the atoms of oxygen (black) & hydrogen (white), & that each oxygen has 2 hydrogens tied to it. (Each little group of an oxygen with its 2 hydrogens is called a molecule.) The picture is idealized further in that the real particles in nature are continually jiggling & bouncing, turning & twisting around one another. You will have to imagine this as a dynamic rather than a static picture. Another thing that cannot be illustrated in a drawing is the fact that the particles are “stuck together” – that they attract each other, this one pulled by that one, etc. The whole group is “glued together,” so to speak. On the other hand, the particles do not squeeze through each other. If you try to squeeze 2 of them too close together, they repel.

The atoms are  $1$  or  $2 \cdot 10^{-8}$  cm in radius. Now  $10^{-8}$  cm is called an *angstrom* (just as another name), so we say they are 1 or 2 angstroms  $\text{\AA}$  in radius. Another way to remember their size is this: if an apple is magnified to the size of the earth, then the atoms in the apple are approximately the size of the original apple.

Now imagine this great drop of water with all of these jiggling particles stuck together & tagging along with each other. The water keeps its volume; it does not fall apart, because of the attraction of the molecules for each other. If the drop is on a slope, where it can move from 1 place to another, the water will flow, but it does not just disappear – things do not just fly apart –



because of the molecular attraction. Now the jiggling motion is what we represent as *heat*: when we increase the temperature, we increase the motion. If we heat the water, the jiggling increases & the volume between the atoms increases, & if the heating continues there comes a time when the pull between the molecules is not enough to hold them together & they *do* fly apart & become separated from one another. Of course, this is how we manufacture steam out of water – by increasing the temperature; the particles fly apart because of the increased motion.

” – [FLS11a, pp. 4–4]

## 2.3 Basic Physics

## 2.4 The Relation of Physics to Other Sciences

## 2.5 Conservation of Energy

## 2.6 The Theory of Gravitation

## 2.7 Quantum Behavior

# 3 RICHARD FEYNMAN. The Feynman Lectures on Physics

### Resources – Tài nguyên.

1. [FLS11b]. RICHARD P. FEYNMAN, ROBERT B. LEIGHTON, MATTHEW SANDS. *The Feynman Lectures on Physics*. [1078 Amazon ratings][7848 Goodreads ratings]

Amazon review. The legendary introduction to physics from the subject’s greatest teacher. “The whole thing was basically an experiment,” RICHARD FEYNMAN said late in his career, looking back on the origins of his lectures. The experiment turned out to be hugely successful, spawning a book that has remained a definitive introduction to physics for decades. Ranging from the most basic principles of Newtonian physics through such formidable theories as general relativity & quantum mechanics, FEYNMAN’s lectures stand as a monument of clear exposition & deep insight. Now, we are reintroducing the printed books to the trade, fully corrected, for the 1st time ever, & in collaboration with Caltech. Timeless & collectible, the lectures are essential reading, not just for students of physics but for anyone seeking an introduction to the field from the inimitable FEYNMAN.

# 4 Miscellaneous

## Tài liệu

- [FLS11a] Richard P. Feynman, Robert B. Leighton, and Matthew Sands. *Six Easy Pieces: Essentials of Physics Explained by Its Most Brilliant Teacher*. 4th edition. Basic Books Publisher, 2011, p. 176.
- [FLS11b] Richard P. Feynman, Robert B. Leighton, and Matthew Sands. *The Feynman Lectures on Physics*. The New Millennium Edition. Basic Books Publisher, 2011, p. 1552.