Elementary Chemistry

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Tóm tắt nội dung

[EN] This text is a collection of problems, from easy to advanced, about *hydrogen & air*. This text is also a supplementary material for my lecture note on Elementary Chemistry grade 8, which is stored & downloadable at the following link: GitHub/NQBH/hobby/elementary chemistry/grade 8/lecture¹. The latest version of this text has been stored & downloadable at the following link: GitHub/NQBH/hobby/elementary chemistry/grade 8/hydrogen².

[VI] Tài liệu này là 1 bộ sưu tập các bài tập chọn lọc từ cơ bản đến nâng cao về oxi & không khí. Tài liệu này là phần bài tập bổ sung cho tài liệu chính – bài giảng GitHub/NQBH/hobby/elementary chemistry/grade 8/lecture của tác giả viết cho Hóa Sơ Cấp lớp 8. Phiên bản mới nhất của tài liệu này được lưu trữ & có thể tải xuống ở link sau: GitHub/NQBH/hobby/elementary chemistry/grade 8/hydrogen.

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¹URL: https://github.com/NQBH/hobby/blob/master/elementary_chemistry/grade_8/NQBH_elementary_chemistry_grade_8.pdf.

 $^{^2 \}text{URL: https://github.com/NQBH/hobby/blob/master/elementary_chemistry/grade_8/hydrogen/NQBH_hydrogen.pdf.}$

1 Wikipedia/Hydrogen

"Hydrogen is the chemical element with the symbol H & atomic number 1. Hydrogen is the lightest element. At standard conditions hydrogen is a gas of diatomic molecules having the formula H₂. It is colorless, odorless, tasteless, non-toxic, & highly combustible. Hydrogen is the most abundant chemical substance in the universe, constituting roughly 75% of all normal matter. Stars such as the Sun are mainly composed of hydrogen in the plasma state. Most of the hydrogen on Earth exists in molecular forms such as water & organic compounds. For the most common isotope of hydrogen (symbol ¹H) each atom has 1 proton, 1 electron, & no neutrons.

In the early universe, the formation of protons, the nuclei of hydrogen, occurred during the 1st second after the Big Bang. The emergence of neutral hydrogen atoms throughout the universe occurred about 370000 years later during the recombination epoch, when the plasma had cooled enough for electrons to remain bound to protons.

Hydrogen is nonmetallic (except it becomes metallic at extremely high pressures) & readily forms a single covalent bond with most nonmetallic elements, forming compounds such as water & nearly all organic compounds. Hydrogen plays a particularly important role in acid-base reactions because these reactions usually involve the exchange of protons between soluble molecules. In ionic compounds, hydrogen can take the form of a negative charge (i.e., anion) where it is known as a hydride, or as a positively charged (i.e., cation) species denoted by the symbol H⁺. The H⁺ cation is simply a proton (symbol p) but its behavior in aqueous solutions & in ionic compounds involves screeing of its electric charge by nearby polar molecules or anions. Because hydrogen is the only neutral atom for which the Schrödinger equation can be solved analytically, the study of its energetics & chemical bonding has played a key role in the development of quantum mechanics.

Hydrogen gas was 1st artificially produced in the early 16th century by the reaction of acids on metals. In 1766–1781, Henry Cavendish was the 1st to recognize that hydrogen gas was a discrete substance, & that it produces water when burned, the property for which it was later named: in Greek, hydrogen means "water-former".

Industrial production is mainly from steam reforming of natural gas, oil reforming, or coal gasification. A small percentage is also produced using more energy-intensive methods such as the electrolysis of water. Most hydrogen is used near the site of its production, the 2 largest uses being fossil fuel processing (e.g., hydrocracking) & ammonia production, mostly for the fertilizer market. It can be burned to produce heat or combined with oxygen in fuel cells to generate electricity directly, with water being the only emissions at the point of usage. Hydrogen atoms (but not gaseous molecules) are problematic in metallurgy because they can embrittle many metals." – Wikipedia/hydrogen

- 1.1 Properties
- 1.2 History
- 1.3 Cosmic Prevalence & Distribution
- 1.4 Production
- 1.5 Applications
- 1.6 Biological Reactions
- 1.7 Safety & Precautions